

A Storage Model for Supporting Figures and Other Artefacts in Scientific Libraries: the Case Study of Invenio

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Berlin, 29.09.2011

Outline

- The Invenio and Inspire projects
- Old data model
- New use cases of storin data
 - Figures
 - Data preservation
- New data model
- Uploading data into the repository

Invenio

CERN Document Server

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- Created to be a basis for CERN Document Server
- Meta-data represented in MARC

Application layer

Main file(s):

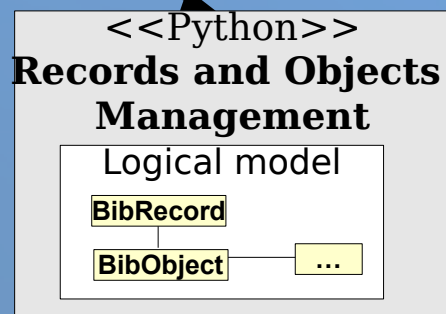
0101431
version 1
0101431.pdf [782.53 KB] 28 Jun 2011, 14:21
0101431.ps.gz [1.91 KB] 28 Jun 2011, 14:21

Additional file(s):

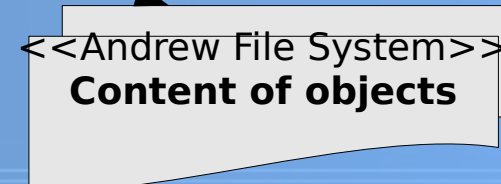
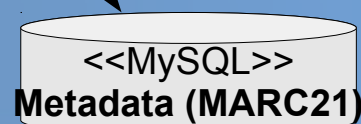
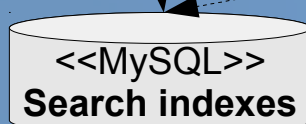
0101431.fig1
version 1
0101431.fig1.ps.gz [2.31 KB] 28 Jun 2011, 14:21
0101431.fig2
version 1
0101431.fig2.ps.gz [231.35 KB] 28 Jun 2011, 14:21



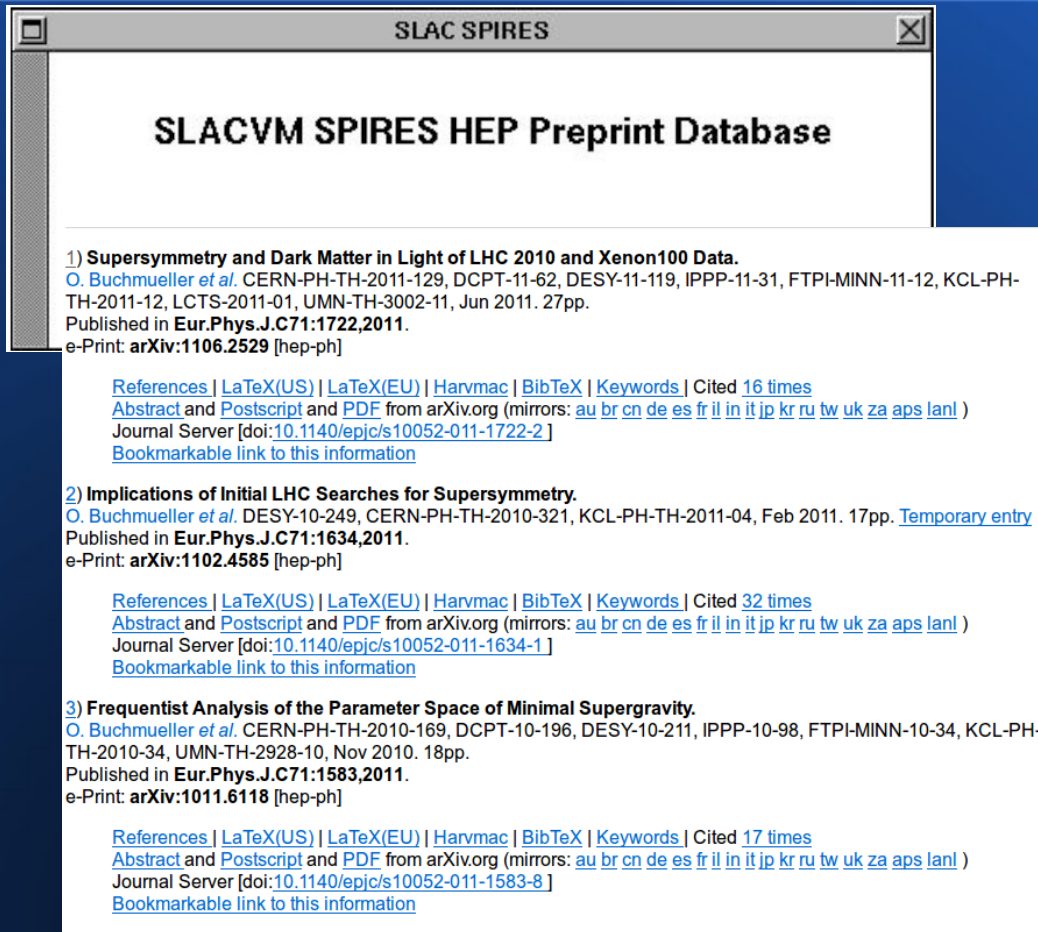
Middleware layer



Storage layer




SPIRES



- Database of preprints started in 70s
- In 90s the first WEB page in USA
- Very difficult to maintain, extremely slow

Invenio + SPIRES = INSPIRE

 Welcome to **INSPIRE** β : the upgrade of SPIRES
We now recommend that you use this site instead of SPIRES
Please send feedback on INSPIRE to feedback@inspirebeta.net

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Information References (10) Citations (0) Files Plots

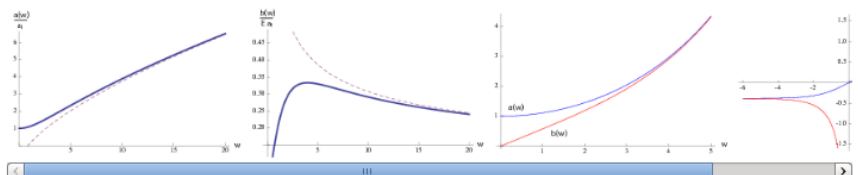
Classical bouncing Universes from vector fields.

Michal Artymowski, Zygmunt Lalak.

Sep 2011
11 pp. e-Print: [arXiv:1109.5901](https://arxiv.org/abs/1109.5901) [hep-th]

Abstract: For the anisotropic Universe filled with massless vector field in the General Relativity frame we obtain bouncing solution for one of scale factors. We obtain the Universe with finite maximal energy density, finite value of $R, R^{\mu\nu} R_{\mu\nu}, R^{\mu\nu\alpha\beta} R_{\mu\nu\alpha\beta}$ and non-zero value of a scale factor for directions transverse to a vector field. Such a bounce can be also obtained for a massive vector field with kinetic initial conditions, which gives isotropic low energy limit. We discuss the existence of a bounce for a massless vector field with additional matter fields, such as cosmological constant or dust. We also discuss bouncing solution for massless vector field domination in $n + 2$ dimensional space-time.

Note: * Temporary entry *



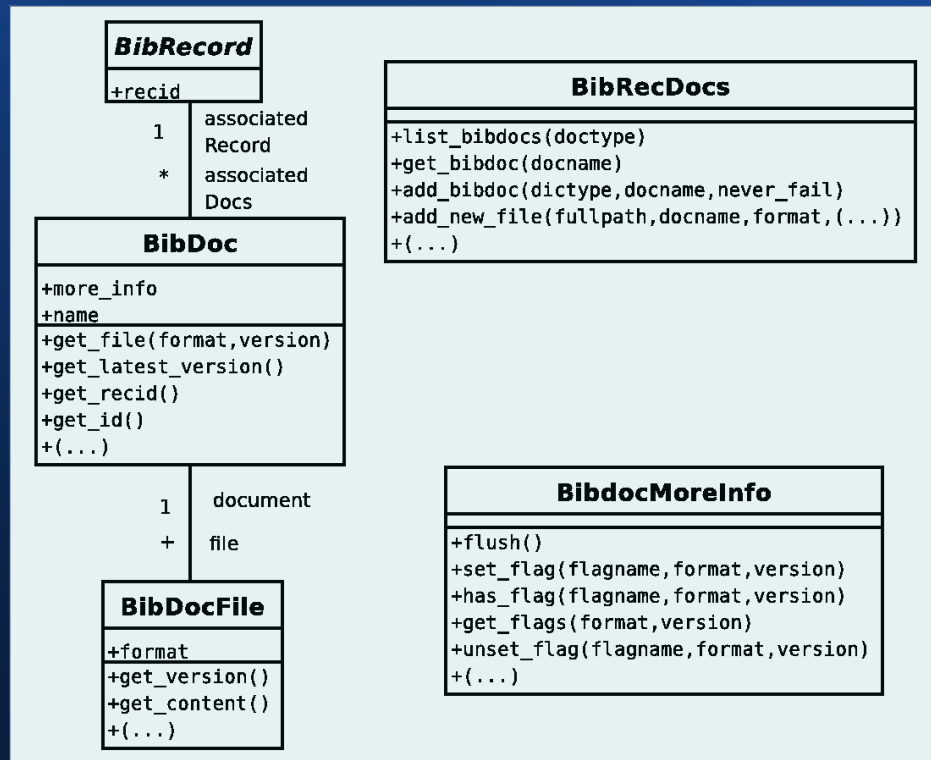
Record created 2011-09-28, last modified 2011-09-28 [Similar records](#)

- Large community of users
- Multiple sources of data (SPIRES, arXiv, direct submissions, publishers)
- Nearly 10000000 records

Invenio/INSPIRE

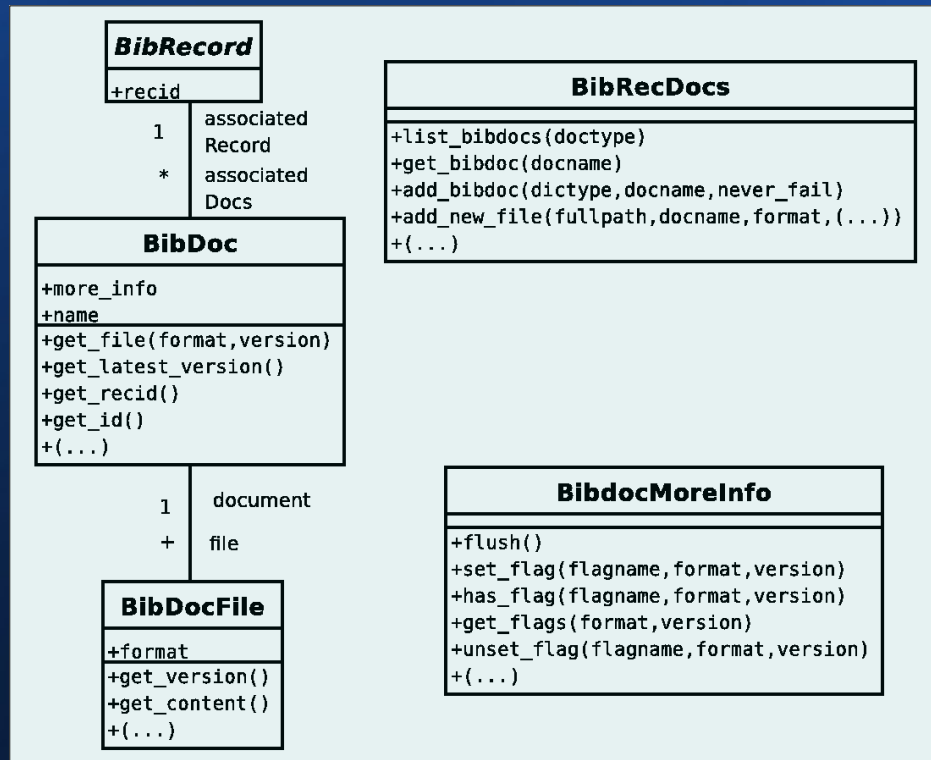
- Invenio – digital library software developed at CERN to manage the repository of documents created in the institution
- SPIRES – The digital library of preprints created at SLAC.
- **In**venio + **SPIRES** = INSPIRE

Non-bibliographical data in Invenio



- Documents represented as BibDoc instances
- Document supports versions and different formats
- Internal data stored in a BibdocMoreInfo instance

Non-bibliographical data in Invenio



- Internal meta-data stored in a MoreInfo instance
- Link between a MARC record and the document
 - Every document must belong to exactly one record

Figures from scientific publications

Automatic extraction

Manual extraction

C. The DSR-compatible cases

As mentioned, the Coleman-Glashow scenario has only been studied and known to produce acceptable physics as a scenario for a full breakdown of special-relativistic Lorentz symmetry. Taking at face value the available data the only special-relativistic superluminal option, the tachyon, is "ruled out" (if we trust the data). From the previous subsection we do have a viable candidate, the Coleman-Glashow case but requires a preferred frame (an "ether frame"). Next let us explore another question: if one takes at face value the presently available data is it then automatic that one is forced to violate the Relativity Principle and formulate the theory in an "ether frame"?

We here explore this by considering the DSR-compatible cases of Eq. (4) and the case of Eq. (5) for which it is established that the modification of the speed law can be implemented relativistically.

In Figs. 5 and 6 we show the results of fitting respectively the case of Eq. (4) and the case of Eq. (5) on the OPERA+FERMILAB79+FERMILAB79 data already shown in our Fig. 2. The results are not encouraging: those fits come with a reduced χ^2 of 2.01 for the case of Eq. (4), in Fig. 5, and of 2.55 for the case of Eq. (5).

Before "rejecting" both these DSR pictures we took the additional step already described above for the case of the rejection of the special-relativistic tachyon: we investigated whether the discouraging reduced χ^2 of the fits were affected very significantly by the worrisome large bias correction applied in Ref. [3] on the FERMILAB79 data. For the case of the quadratic scenario of Eq. (5) we found that this generous attitude was not of much help: the minimum reduced χ^2 obtained by a readjustment of the bias parameter, was still a discouraging 2.29. So, in the same sense in which we confidently rejected the special-relativistic tachyon, we also disfavor the quadratic scenario of Eq. (5) as a plausible description of the OPERA+FERMILAB79+FERMILAB79 (if they can be taken at face value).

For the linear case of Eq. (4) we did find a large sensitivity on the specification of the bias parameter introduced in Ref. [3]. And found that by readjusting the bias parameter one easily finds values smaller than 2 for the reduced χ^2 , all the way down to a value of reduced χ^2 of 1.08. In light of this we reserve judgement on this linear case of Eq. (4), sure that forthcoming additional data on the energy dependence speed of μ neutrinos will allow to draw a more robust conclusion.

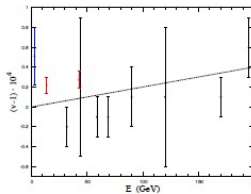


FIG. 5. fit with the DSR-linear case

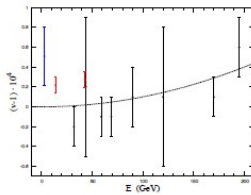


FIG. 6. fit with the DSR-quadratic case

V. SN1987A, OTHER NEUTRINO SPECIES AND LARGE EXTRA DIMENSIONS

We have focused so far on a range of energies which is considerably wider than the OPERA range, but still very narrow in absolute terms. And we focused on data which apply (or can be interpreted as applying) exclusively to μ neutrinos. We shall soon argue that there are some advantages to this approach in a situation such as the one raised by the data recently reported by OPERA.

But before we do that let us instead widen our horizons, considering other neutrino species and other ranges of energy. From this perspective one should immediately consider the observations of neutrinos from the supernova 1987a (see, e.g., Refs. [7, 27]). It is useful from this perspective to consider the Coleman-Glashow picture, which did very well in our peculiar test on the OPERA+FERMILAB79+FERMILAB79 data, and also the linear DSR-compatible case of Eq. (4), which actually did not perform that well in our analysis of

FIG. 5. fit with the DSR-linear case

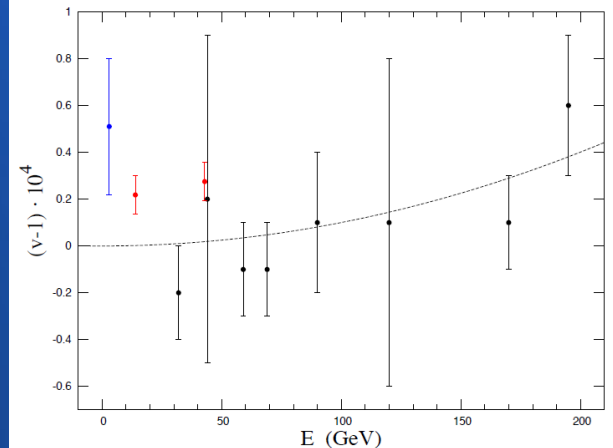


FIG. 6. fit with the DSR-quadratic case

⁵ For completeness we note here (even though it is irrelevant for the narrow scopes of the present exploratory study) that just the deforming Galilean boosts into Lorentz boosts requires the introduction of relativity of simultaneity; we recently understood (Ref. [7, 22]) that in turn deforming Lorentz boosts into DSR-Lorentz boosts requires the introduction of relativity of locality.

C. The DSR-compatible cases

As mentioned, the Coleman-Glashow scenario has only been studied and known to produce acceptable physics as a scenario for a full breakdown of special-relativistic Lorentz symmetry. Taking at face value the available data the only special-relativistic superluminal option, the tachyon, is "ruled out" (if we trust the data). From the previous subsection we do have a viable candidate, the Coleman-Glashow case but requires a preferred frame (an "ether frame"). Next let us explore another question: if one takes at face value the presently-available data is it then automatic that one is forced to violate the Relativity Principle and formulate the theory in an "ether

We here explore this by considering the DSR-compatible cases of Eq. (6) and the case of Eq. (7) for which it is established that the modification of the speed law can be imple-

In Figs. 5 and 6 we show the results of fitting respectively the case of Eq. (6) and the case of Eq. (7) on the OPERA+FERMILAB07+FERMILAB79 data already shown in our Fig. 2. The results are not encouraging: those fits come with a reduced χ^2 of 2.01 for the case of Eq. (6), in Fig. 5, and

Before “rejecting” both these DSR pictures we took the additional step already described above for the case of the rejection of the special-relativistic tachyon: we investigated whether the discouraging reduced χ^2 of the fits were affected very significantly by the worrisome large bias correction applied in Ref. [6] on the FERMIAB79 data. For the case of the quadratic scenario of Fig. 6 we found that this generous attitude was not of much help: the minimum reduced χ^2 obtained by a readjustment of the bias parameter, was still a discouraging 2.29. So, in the same sense in which we confidently rejected the special-relativistic tachyon, we also disfavor the quadratic scenario of Fig. 6 as a plausible description of the OPEA+FERMILAB07+FERMIAB79 (if they can

For the linear case of Fig. 4, we did find a large sensitivity on the specification of the bias parameter introduced in Ref. [3]. And found that by readjusting the bias parameter one easily finds values smaller than 2 for the reduced χ^2 , all the way down to a value of reduced χ^2 of 1.08. In light of this we reserve judgement on this linear case of Fig. 4, sure that forthcoming additional data on the energy dependence speed of a neutrinos will allow to draw a more robust conclusion.

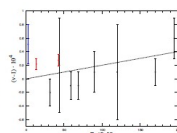


FIG. 5. fit with the DSR linear case.

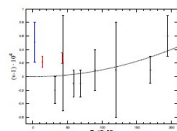


FIG. 6. fit with the DSR-quadratic case

圖 2-1 不同類型之建築工程之估價方法

We have focused so far on a range of energies which is considerably wider than the OPERA range, but still very narrow in absolute terms. And we focused on data which apply (it can be interpreted as applying) exclusively to μ neutrinos. We shall soon argue that there are some advantages to this approach in a situation such as the one raised by the data record.

But before we do that let us instead widen our horizon considering other neutrino species and other ranges of energy. From this perspective one should immediately consider the observations of neutrinos from the supernova 1987A (see, e.g., Refs. [7, 8]). It is useful from this perspective to consider the Coleman-Glashow picture, which did very well in an oscillation test on the OPERA+FERMILAB07+FERMILAB07 data, and also the linear DSR-compatible case of Eq. (4), which actually did not perform that well in an analogous

³ For completeness we note here (even though it is irrelevant for the narrow scopes of the present exploratory study) that just like deforming the Galilean boosts into Lorentz boosts requires the introduction of relativity of simultaneity, we recently understood (Järv & 2022) that in turn deforming Lorentz boosts into DSR-Lorentz boosts requires the introduction of relativity of locality.

- Figures should exist independently from articles

extraction

- Some type of meta data should not be presented to users directly

FIG. 5. fit with the DSR-linear case

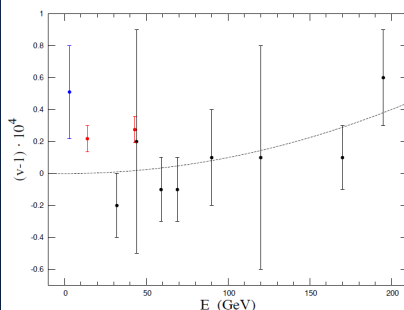
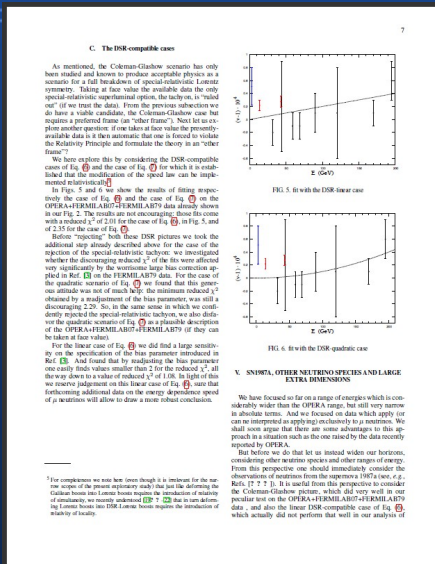
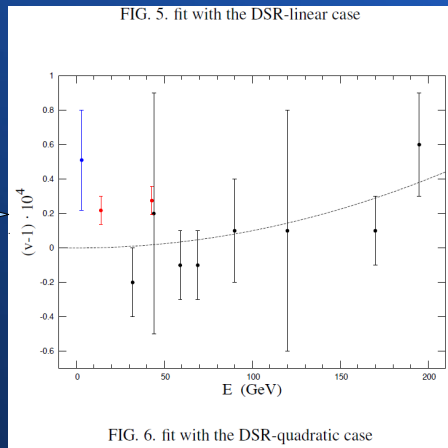


FIG. 6. fit with the DSR-quadratic case

Figures from scientific publications



Extracted from



Extracted from

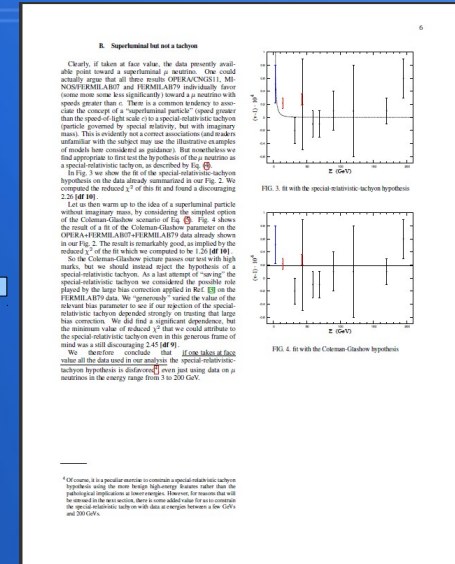


Figure 1 \updownarrow Describes the same data

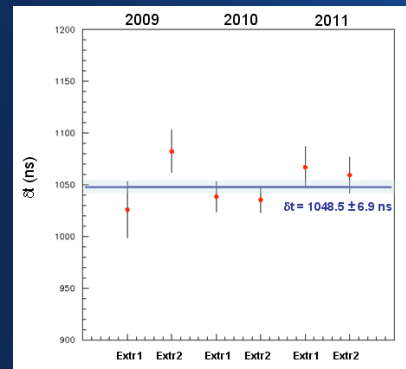


Figure 2 (extracted from different publication)

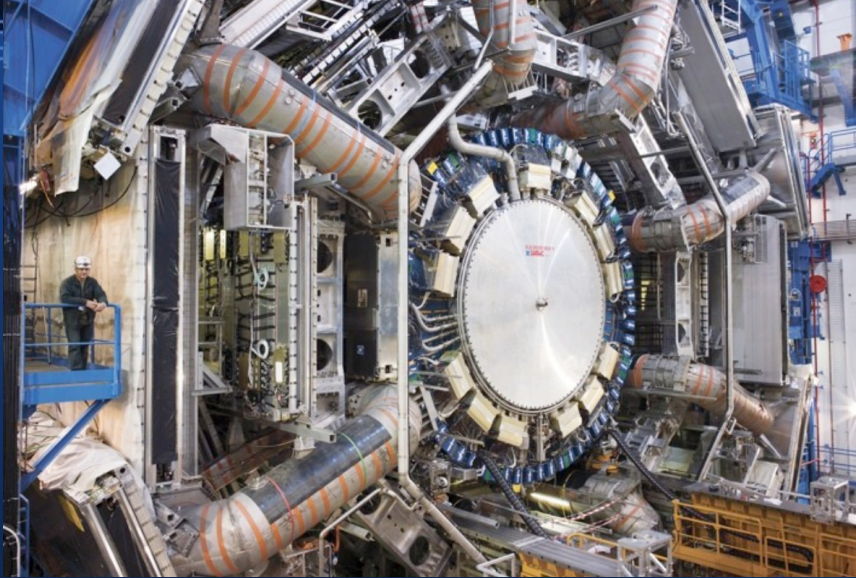
Data-model requirements

- Different types of relations between figures
(illustrates the same data, is subfigure of...)
- Relation of being extracted from a document
- Meta-data of figures, different versions of figure, relations between figures, links between figure and document it is extracted from
- Storage of more complicated data-types

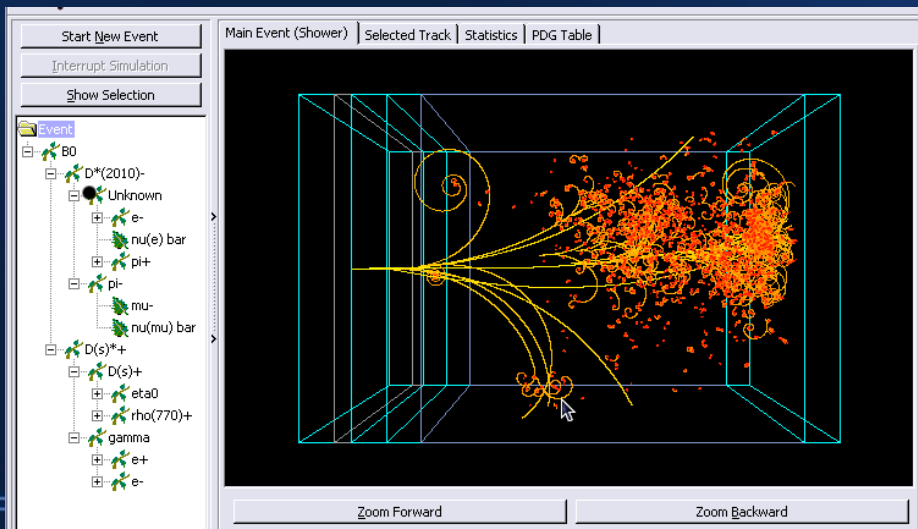
Examples of meta-data associated with different entities

Figure <i>(the most general meta-data)</i>	Figure version <i>(Appearance related meta-data)</i>	Figure storage format	Relation between figure and document	Relation between two figures
Type of figure	Semantics of a figure	Access permissions	Position of a figure within the original document	Type of relation
Quantities presented on axis (in the case of plots)	Data extracted from figure		Caption of a figure within a particular document	<additional type-dependent fields>
Units and scales of axis (in the case of plots)			References from within text	
...			Figure identifier within a document	

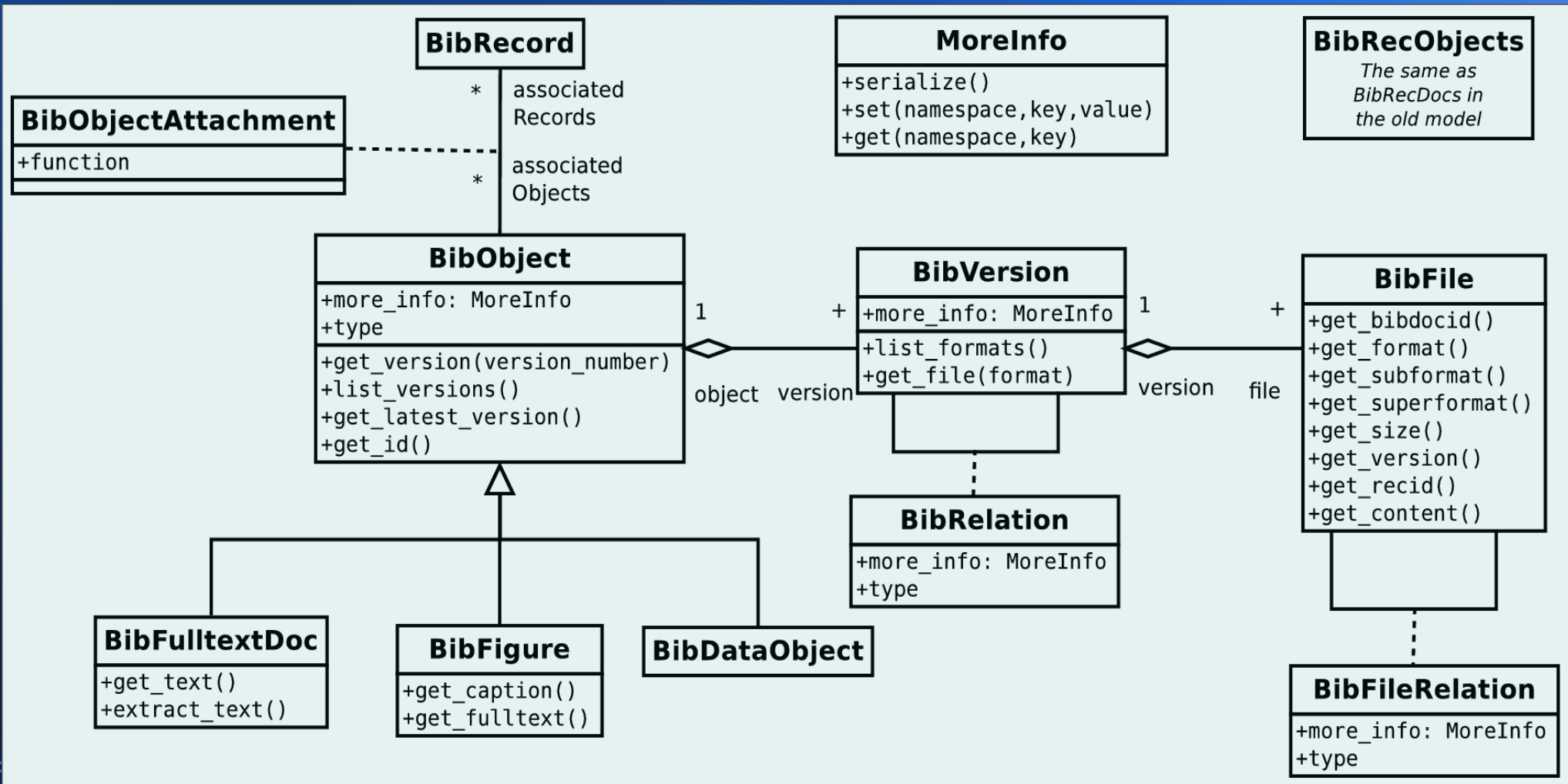
HEP data preservation



- Storing raw data
- Storing intermediate analysis
- Storing additional documentation
- Assigning Digital Object Identifiers



New architecture for storing non-bibliographical objects



BibObject

- Abstract representation of a document (not
- Document-type specific functionalities are implemented by subclasses (defined by modules of Invenio and loaded dynamically)
- Identified by a globally-unique identifier and by a name unique in the scope of a bibliographical record

BibVersion, BibFile

- Represent increasing specialisation of a document
- BibVersion represents a particular revision of an object (corresponding for example to correction of mistakes)
- BibFile describes a particular encoding of a version of an object (encapsulates the real file, remembers the format of a file)

BibRelation – link between entities

- Allows to describe dependencies and connections between different entities of the data model
- Allows specifying an arbitrary type of the relation (for example „is extracted from“, „is the same as“ etc...)

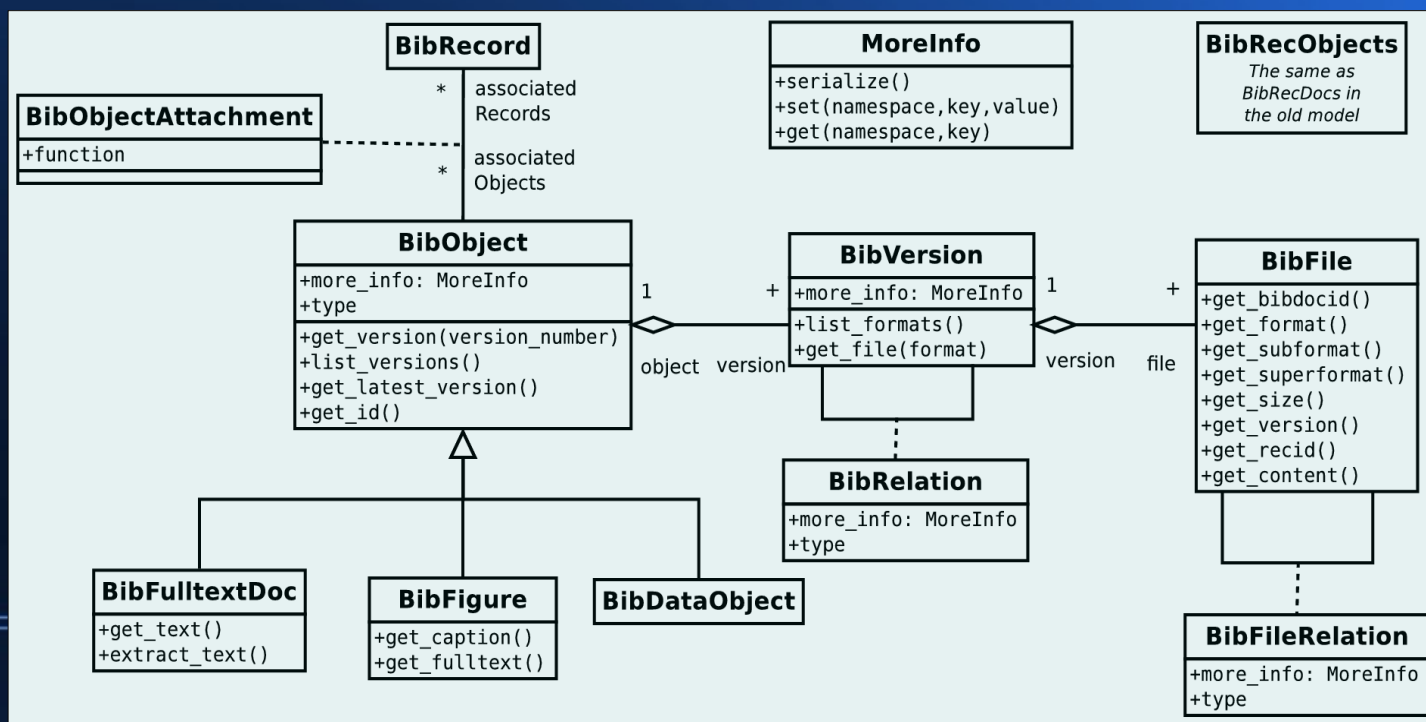
MoreInfo: custom meta-data container

Namespace → key → value

- Can be attached to any entity (BibObject, BibVersion, BibFile, BibRelation)
- Persistently stores a generic dictionaries (every module has their own identified by the namespace)

Data model and figures storage

Data model	Figures
BibObject	Figure
BibObjectVersion	Figure version
BibFile	Particular encoding of a figure
BibRelation	Relation between figure and original document
BibRelation	Relation between two figures



Comparison between data models

	Old model	New model
Attaching document to a bibliographic record	YES	YES
Attaching the same document to many records	NO	YES
Storing custom data keys in MoreInfo dictionaries	NO	YES
Creating documents not attached to any records	NO	YES

Uploading data to Invenio

```
<?xml version="1.0" encoding="UTF-8"?>
<collection xmlns="http://www.loc.gov/MARC21/slim">
  <record>
    <controlfield tag="001">929725</controlfield>
    <datafield tag="970" ind1=" " ind2=" ">
      <subfield code="a">SPIRES-9208755</subfield>
    </datafield>
    <datafield tag="100" ind1=" " ind2=" ">
      <subfield code="a">Artymowski, Michal</subfield>
    </datafield>
    <datafield tag="700" ind1=" " ind2=" ">
      <subfield code="a">Lalak, Zygmunt</subfield>
    </datafield>
    <datafield tag="856" ind1="4" ind2=" ">
      <subfield code="u">http://inspirebeta.net/record/929725/files/arXiv:1109.5901.pdf</subfield>
    </datafield>
    <datafield tag="856" ind1="4" ind2=" ">
      <subfield code="u">http://inspirebeta.net/record/929725/files/BBH2hnkanon.png</subfield>
      <subfield code="y">00003 The left panel shows the evolution of scalar factors for the domination of a massless vector field with non canonical kinetic term and  $f \propto a^{-4}$ . The right panel presents the evolution of Hubble parameters in the same model. One can see, that after  $\sqrt{\rho_I} \sim 5$  the Universe becomes isotropic and enters the era of exponential expansion.</subfield>
    </datafield>
    <datafield tag="856" ind1="4" ind2=" ">
      <subfield code="u">http://inspirebeta.net/record/929725/files/BBH2hnkanon.png</subfield>
    </datafield>
  </record>
</collection>
```

sample.xml Top L7 (nXML Valid)

- New record is encoded in MARC XML
- BibUpload is executed adding uploading task to the BibSched queue
- BibSched uploads data to the main database

FFT = Fulltext File Transfer

```
<?xml version="1.0" encoding="UTF-8"?>
<collection xmlns="http://www.loc.gov/MARC21/slim">
  <record>
    <datafield tag="FFT" ind1=" " ind2=" ">
      <subfield code="a">http://invenio-software.org/download
d/invenio-demo-site-files/0106015_01.jpg</subfield>
      <subfield code="r">restricted_picture</subfield>
    </datafield>
    <datafield tag="FFT" ind1=" " ind2=" ">
      <subfield code="a">http://invenio-software.org/download
d/invenio-demo-site-files/0106015_01.gif</subfield>
      <subfield code="f">.gif;icon</subfield>
      <subfield code="r">restricted_picture</subfield>
    </datafield>
  </record>
</collection>
```

- Artificial, interpreted and removed during the BibUpload phase
- One entry represents one file
- Enforces documents to be attached to a record

subfield	explanation
a	URL of the file to upload
t	Function of the document within the record
f	Format of the file
...	

Uploading data in new format

- New artificial MARC XML fields:
 - BRT (Uploading and modifying relations between documents)
 - MIT (Uploading MoreInfo fields)
 - BDR (Attaching existing objects to records)
- Open for extension supporting METS

Uploading MoreInfo

- Externally (MIT field) or internally (from within FFT/BRT)
- Values encoded in Json or serialised Python objects
- Semantics completely decoupled from BibUpload modes (insert/replace/correct/...)

Conclusions & outlook

- The proposed a model is a flexible approach that facilitates the support of custom objects (figures, data files, software, ...)
 - Based on it, new applications for searching and accessing digital objects can be developed
- Issues to address in the future
 - Integration of extended objects within the INVENIO platform
 - Search and display information about custom objects
 - Need of assigning Digital Object Identifiers (DOI) to stored data objects
 - To store the persistent state of a data object (management of versions)

Thank you !



<http://invenio-software.org/>
<http://www.projectthepinspire.net/>

piotr.praczyk@cern.ch