

Distributed Congestion Control in LEO Satellite Networks

Abstract—Satellite communication in LEO constellations has become an emerging topic of interest. Due to the high number of LEO satellites in a typical constellation, a centralized algorithm for minimum-delay packet routing would incur significant signaling and computational overhead. We can exploit the deterministic topology of the satellite constellation to calculate the minimum-delay path between any two nodes in the satellite network, but that does not take into account the traffic information at the nodes along this minimum-delay path.

We propose a distributed probabilistic congestion control scheme to minimise end-to-end delay. In the scheme, each satellite, while sending a packet to its neighbour, adds a header with a simple metric indicating its own congestion level. The decision to route packets is taken based on the latest traffic information received from the neighbours. We build this algorithm onto the Datagram Routing Algorithm, which provides the minimum delay path, and the decision for the next hop is taken by the congestion control algorithm. We compare the proposed congestion control mechanism with the existing congestion control used by the DRA, and show improvements over the same.

I. INTRODUCTION

What is the problem?

With the advent of cost-effective space launch systems, the feasibility of space-based communication networks has turned into a reality. Dense Low Earth Orbit constellations such as Starlink and OneWeb have joined sparse constellations like Iridium in orbit and are operational. Communication using LEO satellite constellations is

favoured over GEO satellites due to the much lower ground-to-satellite propagation delay.

The challenges faced by satellite constellations are very different from those encountered by terrestrial networks. The nodes in a satellite constellation are constantly moving relative to the ground, so association and handover in a ground-to-satellite link are non-trivial problems. The satellites typically deployed in a constellation are small in size (about 150 kg), which results in limited on-board processing and storage capacity. In addition, the small size of the satellites leads to difficulties in antenna pointing. The inter-satellite links are also characterized by high propagation and transmission delays and high BER. Limited on-board storage capacity leads to packet drops when the nodes get congested, thus degrading the flow of packets. The network has two types of inter-satellite links (ISLs), namely *intra-plane* ISLs, which are the ISLs between two neighbouring satellites in the same orbital plane and *inter-plane* ISLs, which are the ISLs between two neighbouring satellites in different orbital planes. The inter-plane ISLs are difficult to maintain in the polar regions due to the rapid movement of satellites and switching of relative positions.

Due to the dynamic nature of the satellite constellation, paths computed at a central location and sent to the nodes in the network would need a lot of transmissions and computations involving a dense network. Thus, a distributed routing and congestion control algorithm is

preferred. The DRA takes advantage of the spherical geometry of the network and calculates the optimum minimum delay path using the relative positions of the nodes. After that, it is the job of the congestion control algorithm to pick the next hop for a packet to reach its destination with the minimum queueing and propagation delay. The problem of choosing the next hop for the packet in the presence of congestion is the problem that this work focuses on. The choice has to be made locally, and without knowledge of the congestion level of every node along the minimum delay path.

Why is it interesting and important?

A distributed congestion control algorithm that can deal with uneven node congestion levels to route packets from source to destination with low packet drops and end-to-end delay would be easy to implement on-board, and would offer better QoS.

Why is it hard?

The problem of choosing an optimum schedule which minimizes the total end-to-end delay for a given set of packets has been shown to be NP-hard in [1]. Thus, heuristics-based approaches have been used to tackle this problem. In particular, [2] [3] use a basic threshold on the outgoing buffer to determine whether the link is congested. The DRA does not use local congestion information to reroute packets. We address this problem by using the packet headers as a way of conveying traffic information in the form of a single metric indicating the congestion level, and then probabilistically choosing the next hop for a packet. We simulate a typical LEO satellite constellation, and compare the performance of the DRA and our own algorithm in terms of end-to-end delay and packet drops. **Add figures of improvement.**
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II. RELATED WORKS

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III. SYSTEM MODEL

The Low Earth Orbit satellite constellation considered is a Walker star constellation, with satellites in polar orbits. The setup and terminology used is based on the terminology used in [2]. There are N orbital planes, with M satellites per plane. Thus, the angular spacing between the planes is $360^\circ/2N$. **(Add figure of satellite constellation)**. All satellites are at a fixed altitude h from the ground, thus forming an orbital shell. We consider the network to be comprised of *virtual nodes*, as in [2], with different satellites occupying the virtual nodes at different instances of time. The virtual nodes are fixed in location with respect to the ground, and are filled up by the nearest satellite. Each virtual node location can be expressed in terms of the plane number p and the satellite number s , $0 \leq p \leq N - 1$, $0 \leq s \leq M - 1$. Each node in the network can be represented as a tuple (p, s) . This model does not deal with the mobility of satellites, and can be used to perform routing based entirely on the position of the virtual nodes.

As this is a Walker star constellation, each node has two neighbours in the same orbital plane, i.e., *up* and *down*, as well as two neighbours in different orbital planes, i.e., *left* and *right*. The direction *up* is the direction of movement of the satellites in the orbits, while the direction *right* is the direction East from the prime meridian (longitude 0°). **(Add figure denoting neighbours)**

Due to the circular arrangement of orbital planes, there exist counter-rotating seams where a satellite moving north has have a neighbour to its left who is moving south. Some works operate with these inter-plane ISLs disabled, with high Doppler effects and difficulties in antenna pointing as justifications, but we have considered advances in beamforming and antenna tracking, and have decided to use these links across seams as operational links. Inter-plane ISLs in the polar regions, however, are

considered to be shut off due to the change in the orientation of neighbours. (Add figure denoting switching of left and right neighbours). Polar regions are defined using a latitude threshold θ_{polar} , with $\theta_{\text{polar}} = 75^\circ$ taken as the default boundary. The latitudes above θ_{polar} are considered to be in the polar regions.

The lengths of the intra-plane ISLs are the same, as nodes are equally spaced in the orbital planes. For an orbital radius of R , the length of an intra-plane ISL L_v is given by 1, while that of an inter-plane ISL situated at latitude θ is given by 2

$$L_v = R\sqrt{2(1 - \cos(360^\circ/M))} \quad (1)$$

$$L_h = R\cos\theta\sqrt{2(1 - \cos(360^\circ/2N))} \quad (2)$$

As seen in constellation figure, the inter-plane ISLs are shorter towards the poles and longer towards the equator.

IV. PROBLEM FORMULATION

V. ALGORITHMS

VI. SIMULATION SETUP

VII. RESULTS

VIII. CONCLUSIONS AND FUTURE WORK

REFERENCES

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