

P 8

$$r = a(t) R$$

↓ actual distance which increases with time      ↓ scale factor      ↓ Co moving distance (independent of time)

$$\text{i) } v = \dot{r} = \dot{a}(t) R$$
$$d = r = a(t) R$$

$$\Rightarrow H(t) = \frac{\dot{a}}{a}$$

$$\text{ii) } \dot{a} = a H, \quad H \text{ is independent of time}$$
$$\ddot{r} = a H^2 R = H^2 d, \quad a = a(t), \quad d = aR$$

$$\frac{c}{2} > \frac{Gm}{d^2}$$

$$\Rightarrow \boxed{d > \sqrt[3]{\frac{Gm}{H^2}}}$$

iii) Try substituting the values and comparing the two

iv) Note:  $1+z = \frac{f_e}{f_0} = \frac{\lambda_0}{\lambda_e} = \frac{a(t_0)}{a(t_e)}$

$$\underline{a(t_e) = \frac{a(t_0)}{1+z} = \frac{1}{1+z}}$$

v) All the wavelengths will be redshifted and radiance

$$P_{\text{obs}} = P_{\text{em}} \left( \frac{\lambda_{\text{obs}}}{\lambda_{\text{em}}} \right)^4$$

power will fall by a factor  
of  $(1+z)^4$ . Hence,  $T_{\text{obs}} = \frac{T}{1+z}$

v.i)  $4\pi d^2$ !

v.ii) Let  $d_e$  be the distance  
b/w Earth and the body  
at the time of emission  
then flux observed at  
Earth will be:

$$F = \frac{1}{4\pi d_e^2} \frac{L}{(1+z)^4} = \frac{L}{4\pi d^2 (1+z)^2}$$

[Analyse the situation in the frame  
of comoving coordinates]

v.iii)



$$\theta = \frac{l}{d\sqrt{1+z}} = \frac{l(1+z)}{d}$$

ix) Eliminate  $d$  :

$$\frac{F}{\theta^2} = \frac{L}{4\pi(1+z)^4 l^2}$$

$$x) \quad 1+z = \left( \frac{L\theta^2}{4\pi l^2 F} \right)^{1/4}$$

$$d = \frac{l}{\theta} \left( \frac{L\theta^2}{4\pi l^2 F} \right)^{1/4} = \sqrt{\frac{l}{\theta}} \sqrt{\frac{L}{4\pi F}}$$