

Image Restoration

Image Restoration

- Image Enhancement
 - A subjective process using mostly heuristics.
- Image Restoration
 - An objective process assuming a-priori knowledge of the degradation process.

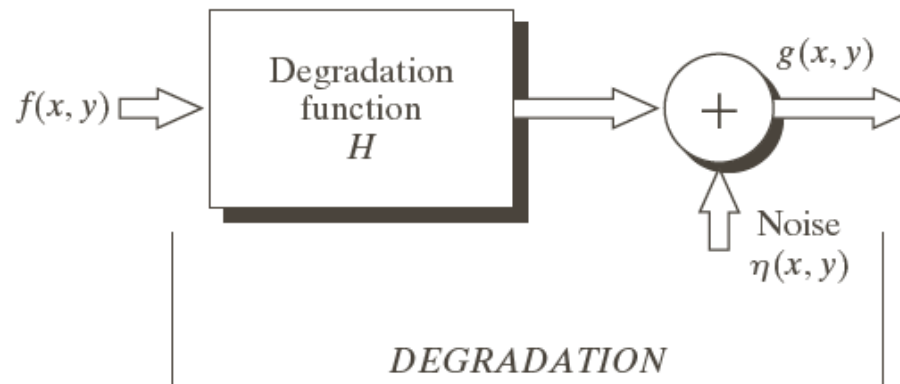
blur due to motion



Modeling Image Degradation

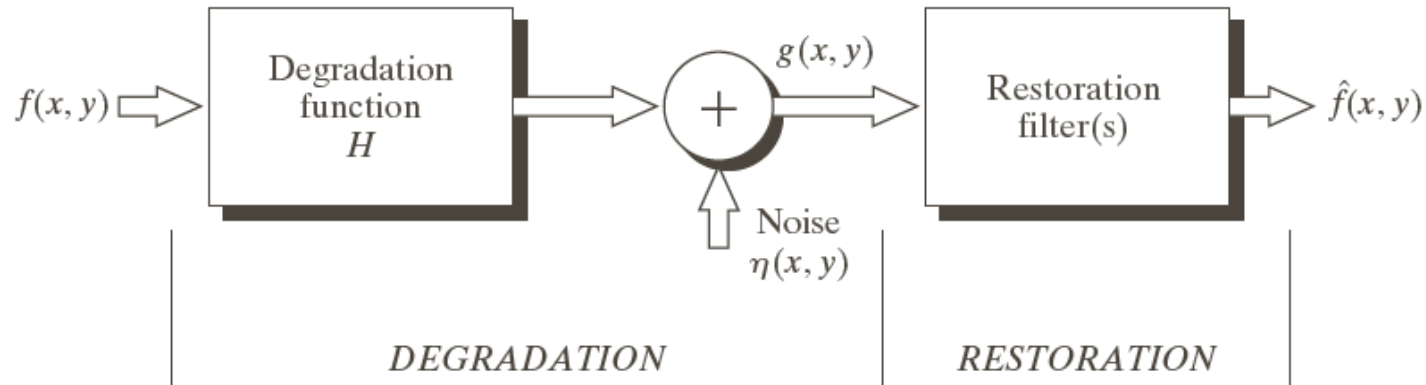
- Degradation process can be modeled through a **degradation** function **H** and additive noise **$\eta(\mathbf{x}, \mathbf{y})$** .

$$\underline{g(x, y) = H[f(x, y)] + n(x, y)}$$



Goal of Image Restoration

- Given some knowledge of \mathbf{H} and noise $\boldsymbol{\eta}(\mathbf{x}, \mathbf{y})$, the objective of image restoration is to obtain an estimate of the original image.



Performance Characterization

- How do we characterize the performance of different image restoration algorithms?

Mean Square Error (MSE)

$$\text{MSE} = \frac{1}{MN} \sum_{x=0}^{M-1} \sum_{y=0}^{N-1} [f(x, y) - \hat{f}(x, y)]^2$$

Warning: do not necessarily
imply
“best” in the visual sense.

Signal to Noise Ratio (SNR)

$$\text{SNR} = \frac{\sum_{u=0}^{M-1} \sum_{v=0}^{N-1} \hat{f}(x, y)^2}{\sum_{u=0}^{M-1} \sum_{v=0}^{N-1} [f(x, y) - \hat{f}(x, y)]^2}$$

Assumptions about degradation function H

- H is **linear**:

$$H[f_1 + f_2] = H[f_1] + H[f_2], \quad (1)$$

$$H[kf] = kH[f] \quad (2)$$

- H is **shift invariant**:

$$\text{If } H[f(x, y)] = g(x, y) \text{ then } H[f(x - a, y - b)] = g(x - a, y - b) \quad (3)$$

(i.e., shifting the input merely shifts the output by the same amount)

Degradation Model

(assuming linearity and shift invariance)

- Any function $f(x,y)$ can be written as follows:

$$f(x, y) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} f(a, b) \delta(x - a, y - b) da db$$

- Then, $H[f(x,y)]$ is:

from (1)

$$H[f(x, y)] = H\left[\int_{-\infty}^{\infty} \int_{-\infty}^{\infty} f(a, b) \delta(x - a, y - b) da db\right] =$$

from (2)

$$\int_{-\infty}^{\infty} \int_{-\infty}^{\infty} H[f(a, b) \delta(x - a, y - b)] da db = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} f(a, b) H[\delta(x - a, y - b)] da db$$

Degradation Model - Continuous Case (cont'd)

- Suppose

$$H[\delta(x, y)] = h(x, y)$$



impulse response

- Since $H(x, y)$ is shift invariant:

from (3)

$$H[\delta(x - a, y - b)] = h(x - a, y - b)$$

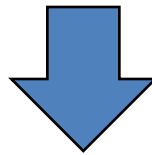
- The

$$H[f(x, y)] = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} f(a, b) h(x - a, y - b) da db = f(x, y) * h(x, y)$$

Degradation Model - Continuous Case (cont'd)

- Under the assumptions of linearity and shift invariance:

$$\underline{g(x, y) = H[f(x, y)] + n(x, y)}$$



simplifies to:

$$\text{Degradation Model: } g(x, y) = f(x, y) * h(x, y) + n(x, y)$$

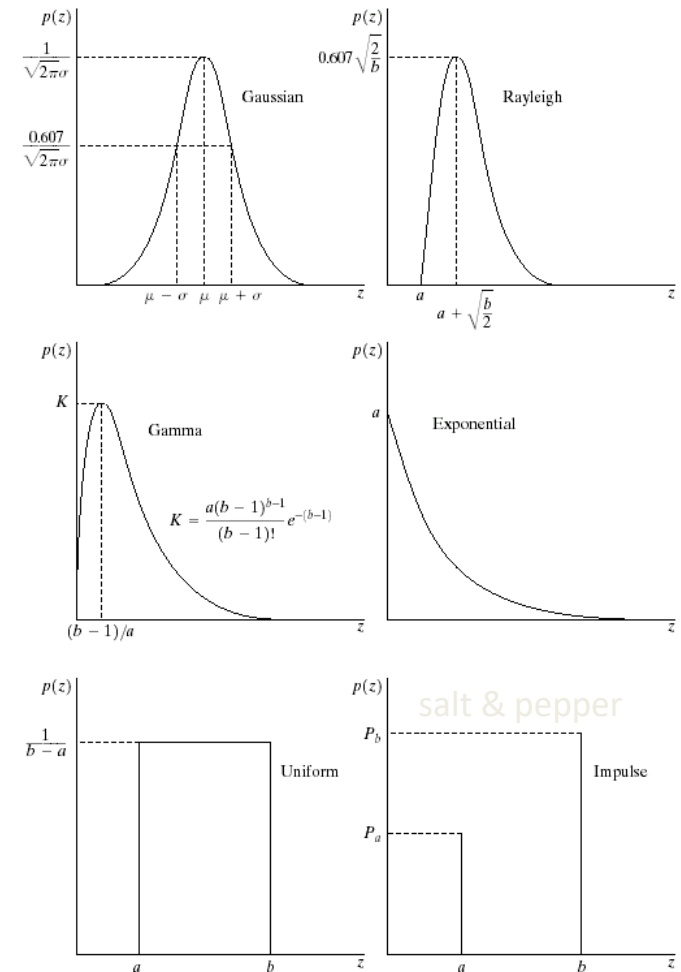
$$\text{or in freq. domain: } G(u, v) = H(u, v)F(u, v) + N(u, v)$$

Noise Properties

- Noise arises typically during image acquisition and/or transmission.
- For simplicity, we will assume that:
 - (1) noise is **independent** of spatial coordinates
 - (2) there is **no correlation** between pixel values and noise values (not true for periodic noise)

Noise Models

- Many types of noise can be modeled using a **probability density** function.
- Model is typically chosen based on some understanding of the noise source.



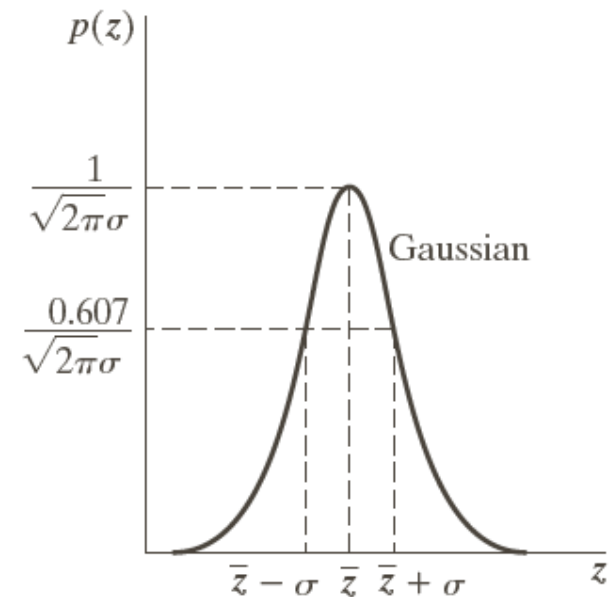
Gaussian noise

$$p(z) = \frac{1}{\sqrt{2\pi}\sigma} e^{-(z-\bar{z})^2 / 2\sigma^2}$$

- Gaussian noise arises in an image due to factors such as:

- (1) Electronic circuit noise

- (2) Sensor noise due to poor illumination and/or high temperature

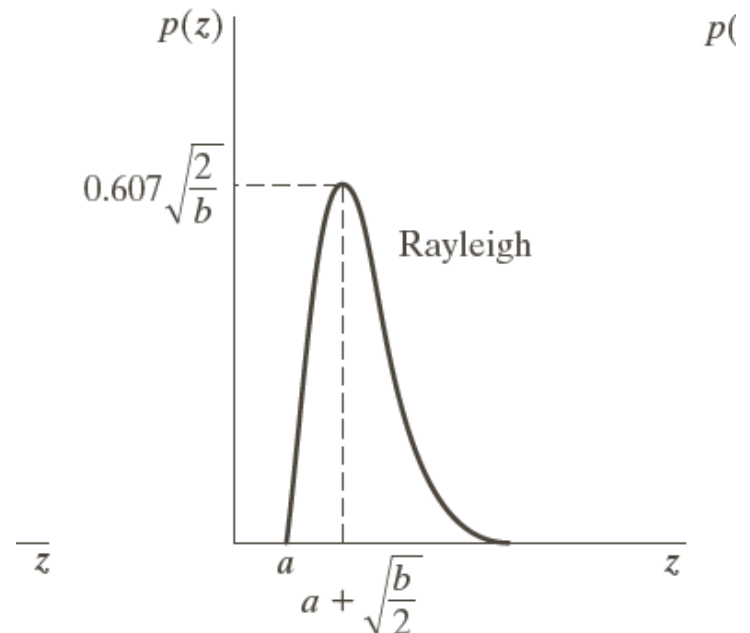


Rayleigh noise

$$p(z) = \begin{cases} \frac{2}{b}(z-a)e^{-(z-a)^2/b} & z \geq a \\ 0 & z < a \end{cases}$$

$$\bar{z} = a + \sqrt{\pi b / 4} \quad \sigma^2 = \frac{b(4 - \pi)}{4}$$

- Typically used to characterize noise in **range** imaging.

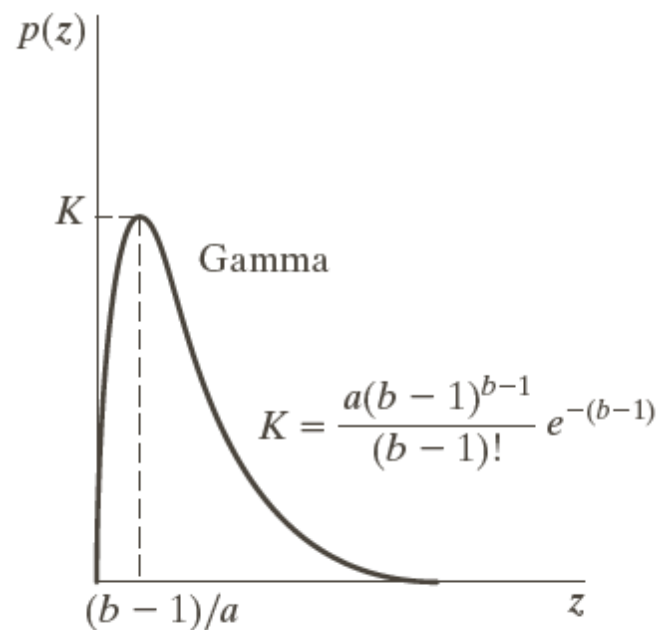


Gamma (Erlang) noise

$$p(z) = \begin{cases} \frac{a^b z^{b-1}}{(b-1)!} e^{-az} & z \geq 0 \\ 0 & z < 0 \end{cases}$$

$$\bar{z} = \frac{b}{a} \quad \sigma^2 = \frac{b}{a^2}$$

- Typically used to characterize noise in laser imaging.

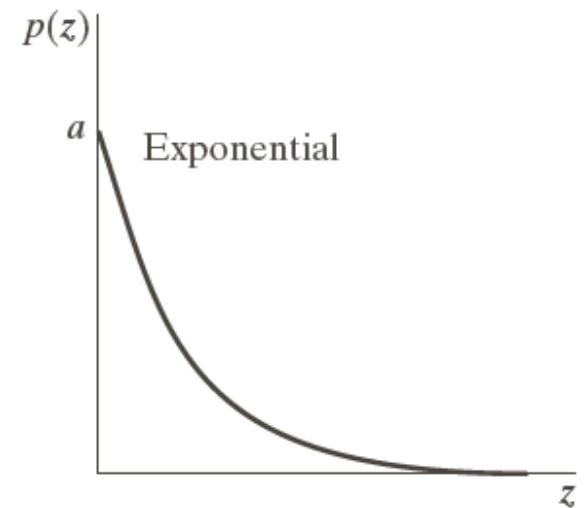


Exponential noise

$$p(z) = \begin{cases} ae^{-az} & z \geq 0 \\ 0 & z < 0 \end{cases}$$

$$\bar{z} = \frac{1}{a} \quad \sigma^2 = \frac{1}{a^2}$$

- Typically used to characterize noise in laser imaging.

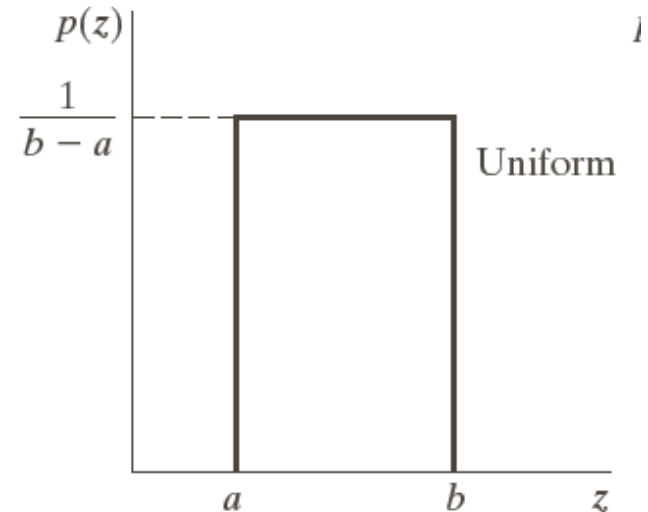


Uniform noise

$$p(z) = \begin{cases} \frac{1}{b-a} & a \leq z \leq b \\ 0 & \text{otherwise} \end{cases}$$

$$\bar{z} = \frac{a+b}{2} \quad \sigma^2 = \frac{(b-a)^2}{12}$$

- Least used in practice.
- Useful as the basis for random number generators.

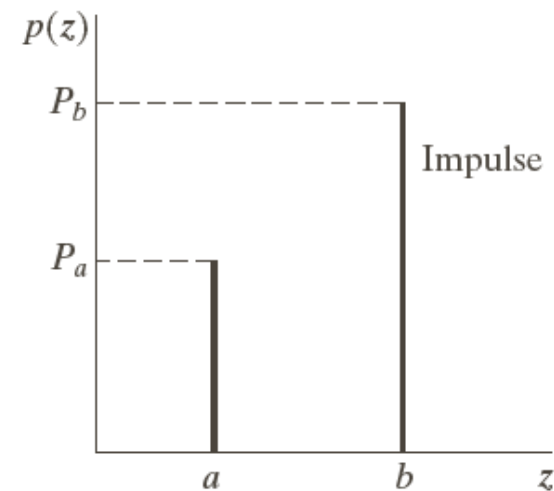


Impulse noise

$$p(z) = \begin{cases} P_a & z = a \\ P_b & z = b \\ 0 & \text{otherwise} \end{cases}$$

- If $P_a = P_b$, $a=0$, and $b=255$, then this is salt and pepper noise.

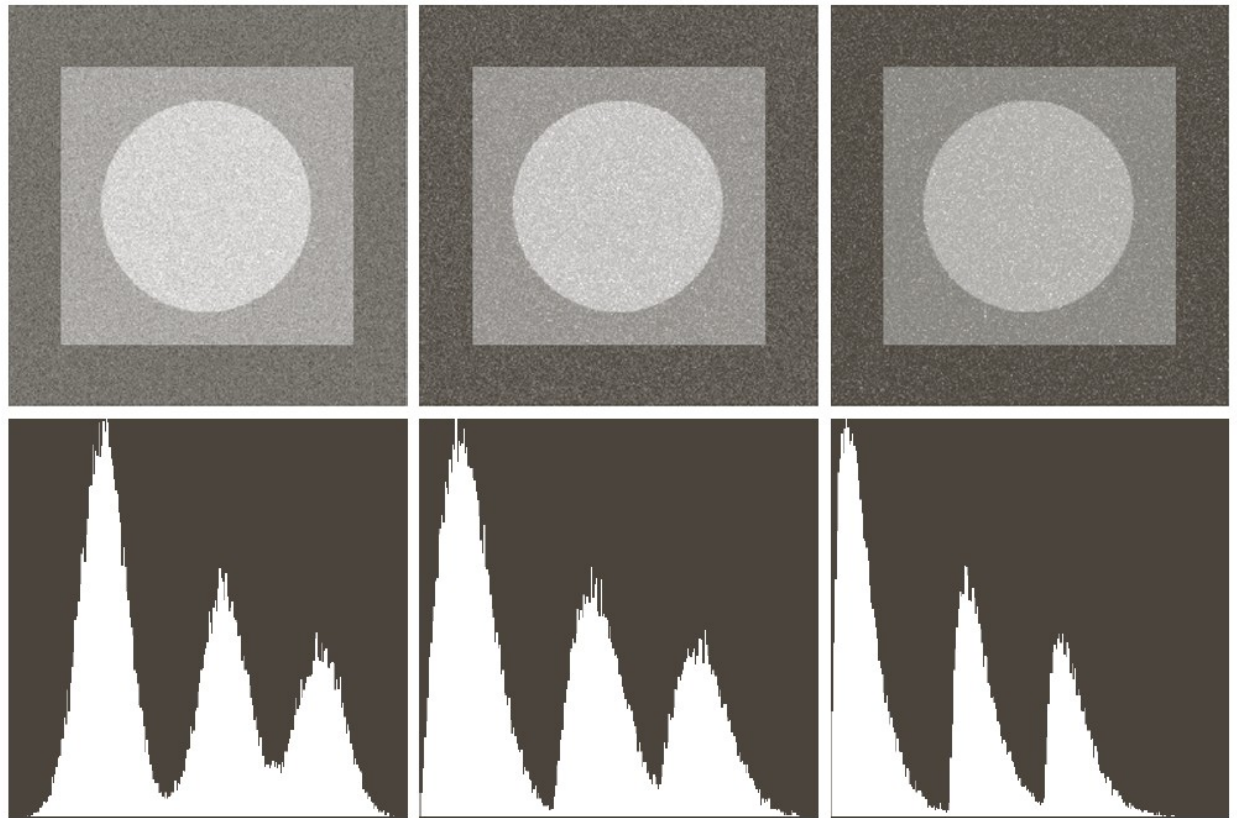
- Common in situations where quick transients (e.g., faulty switching), takes place during imaging.



Examples

noise corrupted images and their histograms

test pattern



Gaussian

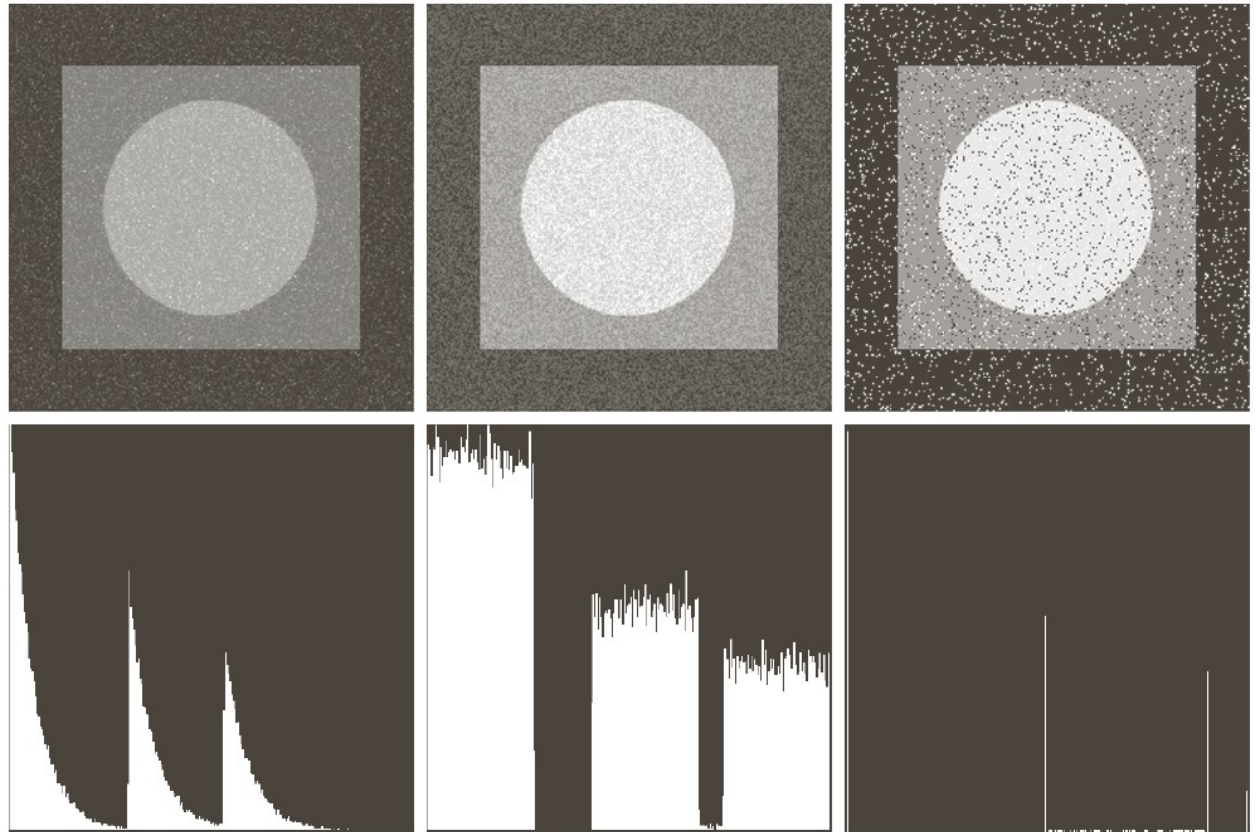
Rayleigh

Gamma

Examples (cont'd)

noise corrupted images and their histograms

test pattern



Exponential

Uniform

Salt & Pepper

Estimation of noise parameters

- Estimate the noise model parameters from a small patch of reasonably **constant** background intensity.
- For impulse noise, estimate probability of black/white pixels (i.e., choose a mid-gray region).

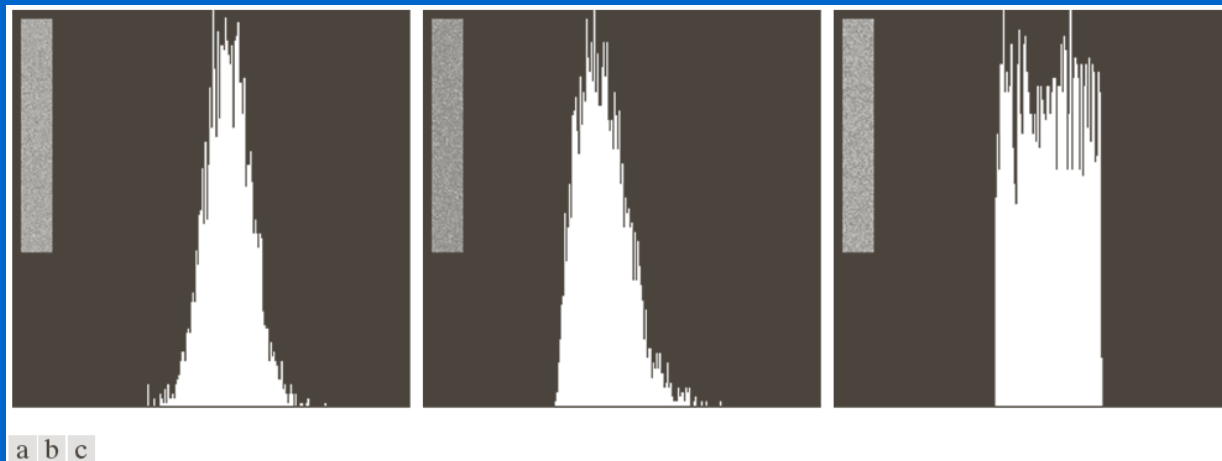


FIGURE 5.6 Histograms computed using small strips (shown as inserts) from (a) the Gaussian, (b) the Rayleigh, and (c) the uniform noisy images in Fig. 5.4.

(1) Estimate mean and variance

(2) Compute a and b (i.e., parameters of noise distribution)

Restoration in the presence of noise only

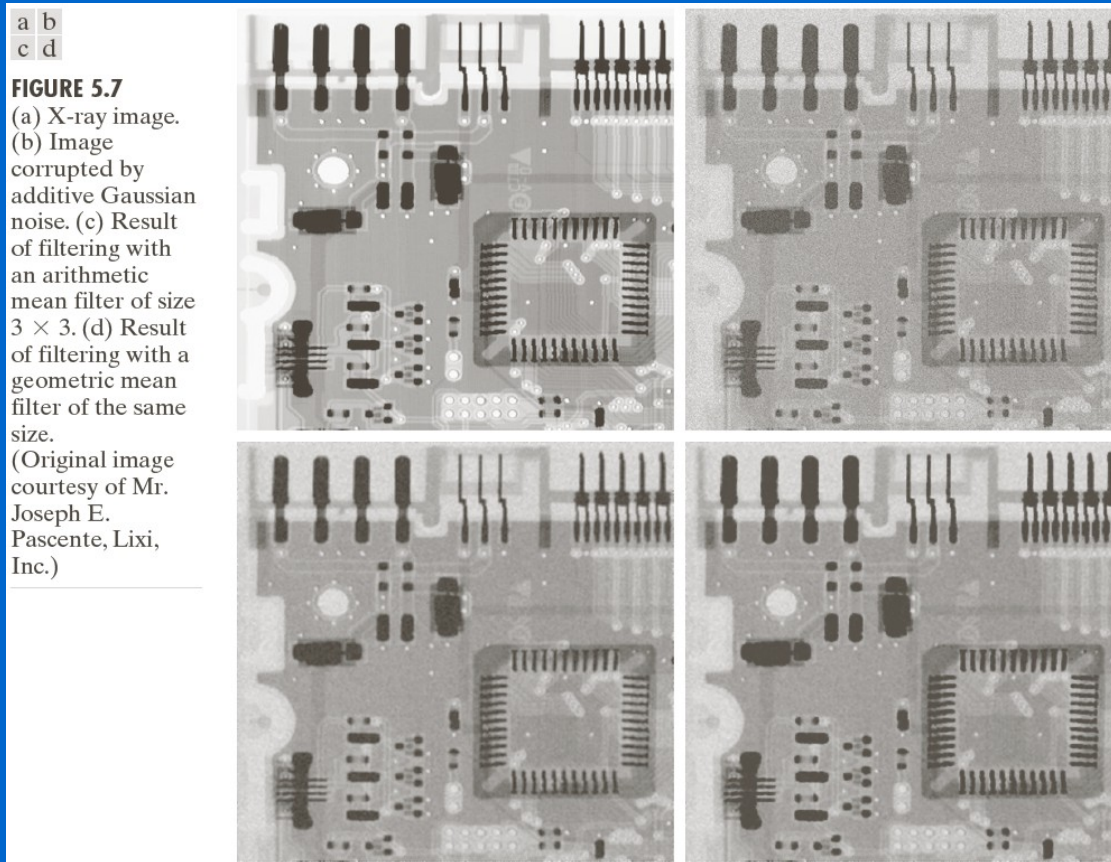
- If noise can be estimated, then subtract it from the input image:

estimate $n(x,y)$

$$g(x,y) = f(x,y) + n(x,y) \rightarrow \hat{f}(x,y) = g(x,y) - \hat{n}(x,y)$$

– Frequency-domain filters (band-reject, band-pass, etc.)

Arithmetic/Geometric mean filters



Geometric mean filters tend to preserve more details (i.e., less blurring)

m x n
mask

$$\hat{f}(x, y) = \frac{1}{mn} \sum_{(s,t) \in S_{x,y}} g(s, t)$$

$$\hat{f}(x, y) = \left[\prod_{(s,t) \in S_{x,y}} g(s, t) \right]^{\frac{1}{mn}}$$

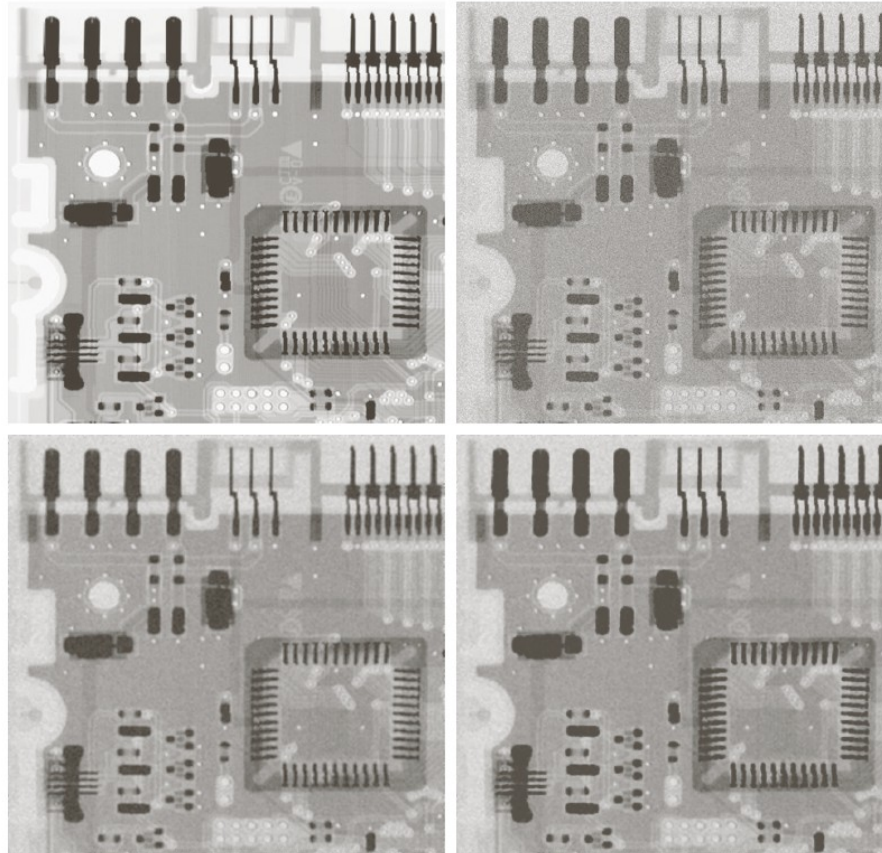
Arithmetic/Geometric mean filters

a b
c d

FIGURE 5.7

(a) X-ray image.
(b) Image corrupted by additive Gaussian noise. (c) Result of filtering with an arithmetic mean filter of size 3×3 . (d) Result of filtering with a geometric mean filter of the same size.

(Original image courtesy of Mr. Joseph E. Pascente, Lixi, Inc.)



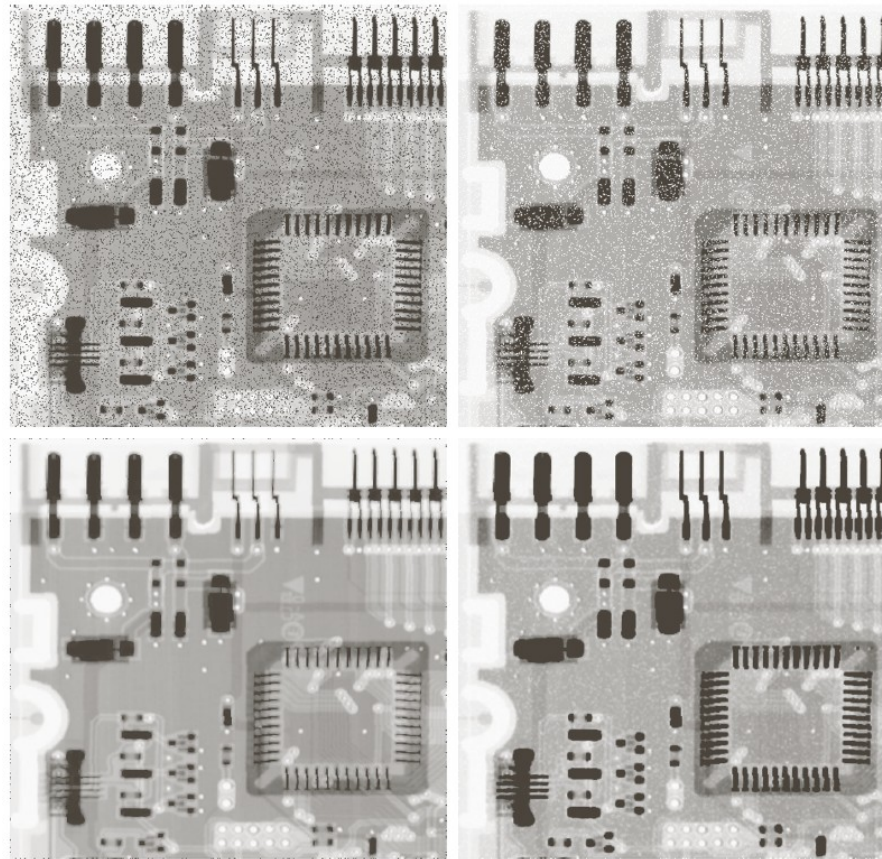
Geometric mean filters tend to preserve more details (i.e., less blurring)

$m \times n$
mask

$$\hat{f}(x, y) = \frac{1}{mn} \sum_{(s,t) \in S_{x,y}} g(s, t)$$

$$\hat{f}(x, y) = \left[\prod_{(s,t) \in S_{x,y}} g(s, t) \right]^{\frac{1}{mn}}$$

Contra-Harmonic filters



a b
c d

FIGURE 5.8

(a) Image corrupted by pepper noise with a probability of 0.1. (b) Image corrupted by salt noise with the same probability. (c) Result of filtering (a) with a 3×3 contra-harmonic filter of order 1.5. (d) Result of filtering (b) with $Q = -1.5$.

Q: order of filter

$$\hat{f}(x, y) = \frac{\sum_{(s,t) \in S_{x,y}} [g(s, t)]^{Q+1}}{\sum_{(s,t) \in S_{x,y}} [g(s, t)]^Q}$$

Good for removing salt **or** pepper noise (but **not** both simultaneously)

$Q > 0 \rightarrow$ pepper noise

$Q < 0 \rightarrow$ salt noise

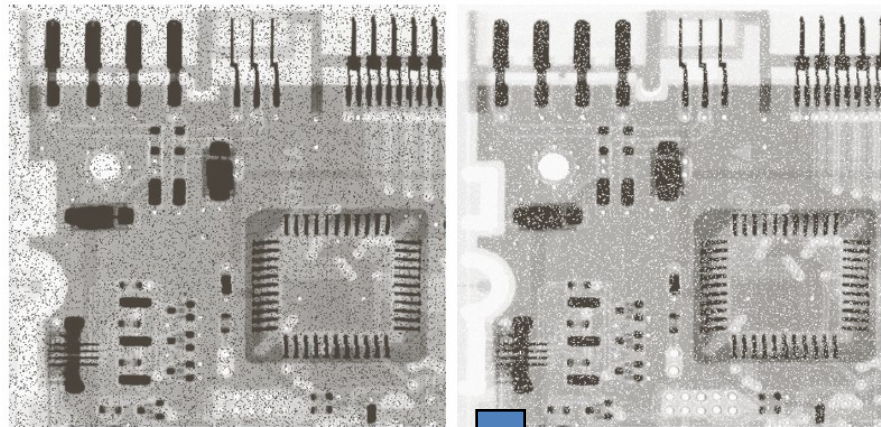
Q=1.5

Q= -1.5

If $Q=0$, same as arithmetic filter!

Contra-Harmonic filters (cont'd)

Example of selecting the wrong sign :



a	b
c	d

FIGURE 5.8

(a) Image corrupted by pepper noise with a probability of 0.1. (b) Image corrupted by salt noise with the same probability.

$Q = -1.5$

$Q = 1.5$

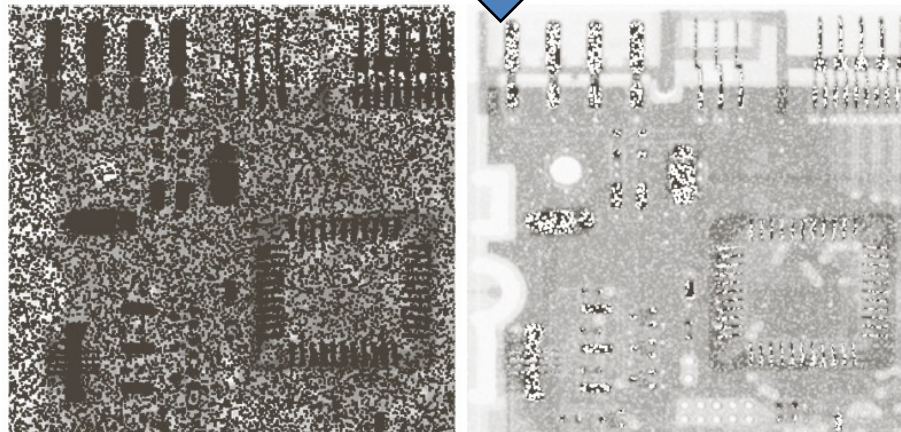


a	b
---	---

FIGURE 5.9

Results of selecting the wrong sign in contraharmonic filtering.

(a) Result of filtering Fig. 5.8(a) with a contraharmonic filter of size 3×3 and $Q = -1.5$.
(b) Result of filtering 5.8(b) with $Q = 1.5$.



Order statistics filters

- Their response is based on the ordering (ranking) of the pixels contained in an area covered by the filter.
- Order statistics filters are nonlinear spatial filters.

Median filter

Multiple passes can improve results

$$\hat{f}(x, y) = \underset{(s, t) \in S_{x, y}}{\text{median}} \{g(s, t)\}$$

Less blurring compared
to arithmetic filters

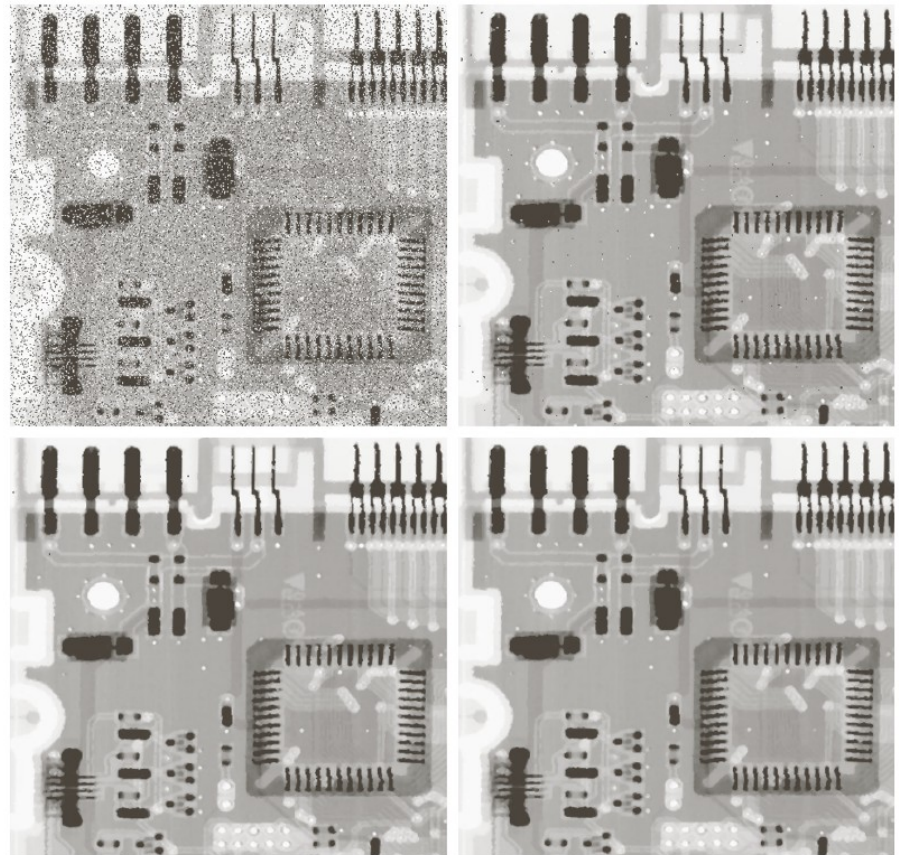
a b
c d

FIGURE 5.10

(a) Image corrupted by salt-and-pepper noise with probabilities $P_a = P_b = 0.1$.

(b) Result of one pass with a median filter of size 3×3 .

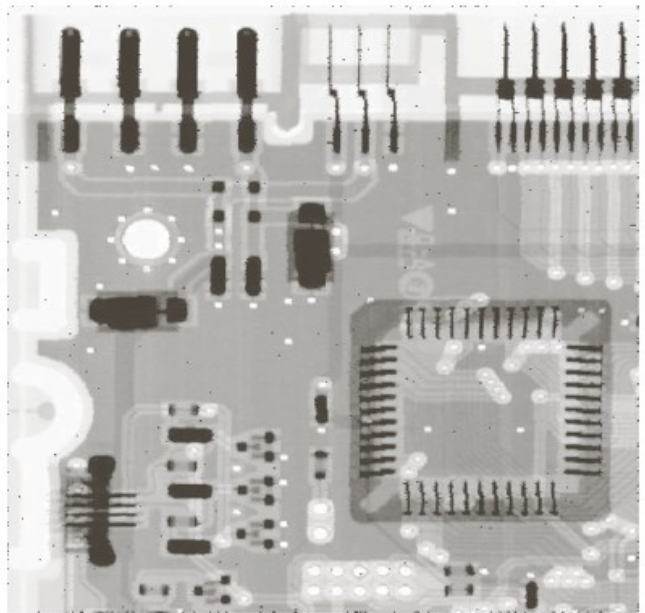
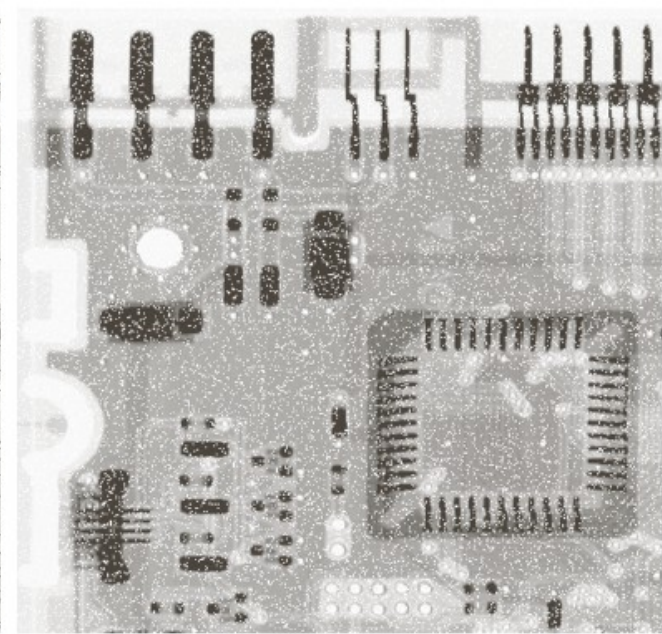
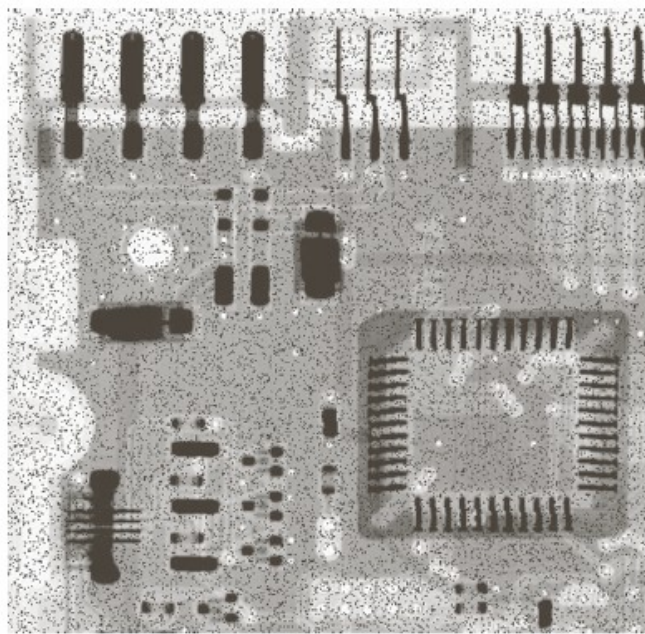
(c) Result of processing (b) with this filter.
(d) Result of processing (c) with the same filter.



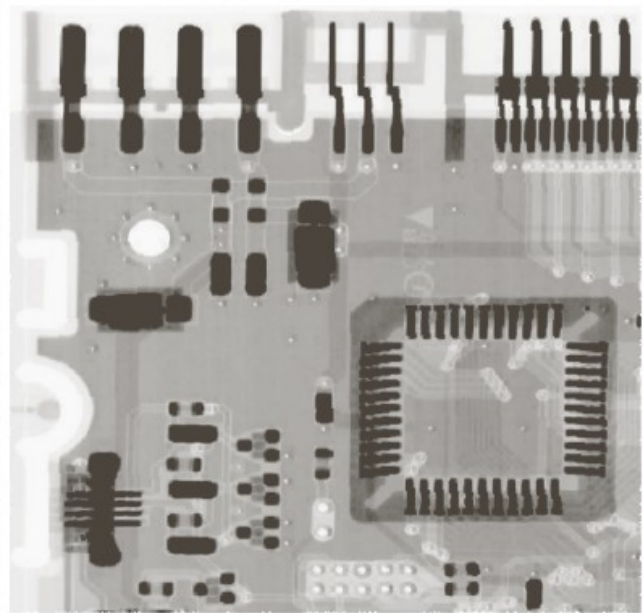
Max/Min filters

$$\hat{f}(x, y) = \min_{(s, t) \in S_{xy}} \{g(s, t)\}$$

$$\hat{f}(x, y) = \max_{(s, t) \in S_{xy}} \{g(s, t)\}$$



Max filter is good
for pepper noise



Min filter is good
for salt noise

Alpha-trimmed mean filter

Good for **multiple** types of noise (e.g., Uniform and salt-and-pepper noise)

Assume an $m \times n$ neighborhood:

$$0 \leq d \leq mn - 1$$

- (1) Disregard $d/2$ lowest and $d/2$ highest values
- (2) Average the remaining values

Special cases:

$d=0 \rightarrow$ arithmetic mean

$d=mn-1 \rightarrow$ median

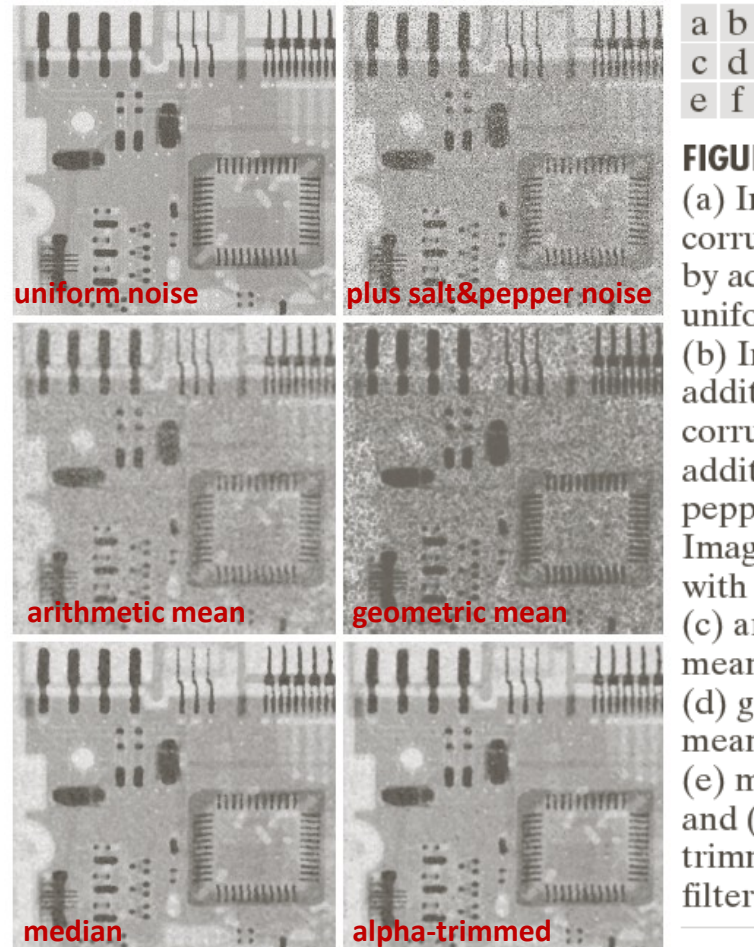


FIGURE 5.12

(a) Image corrupted by additive uniform noise.
(b) Image additionally corrupted by additive salt-and-pepper noise. Image (b) filtered with a 5×5 ;
(c) arithmetic mean filter;
(d) geometric mean filter;
(e) median filter; and (f) alpha-trimmed mean filter with $d = 5$.

Adaptive Filters

- Adaptive filters have superior performance compared to non-adaptive filters.
 - Non-fixed (i.e., adaptive) parameters
 - Have higher complexity.

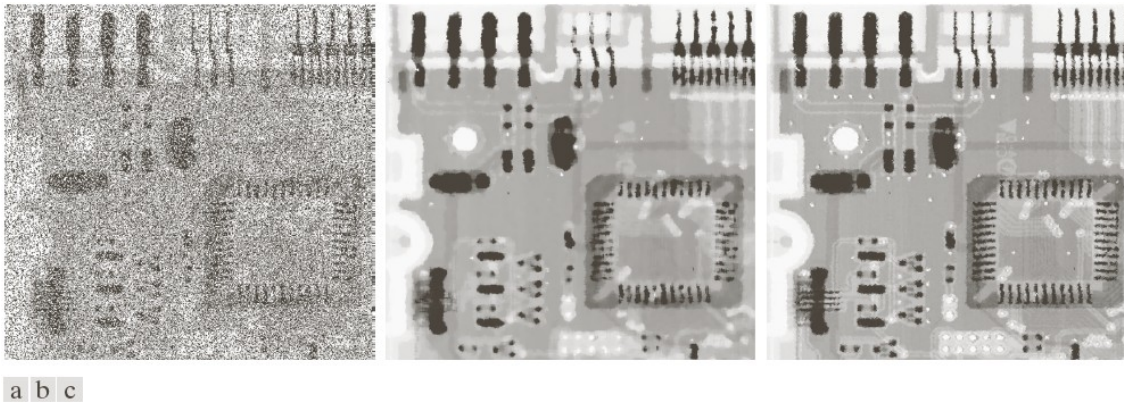


FIGURE 5.14 (a) Image corrupted by salt-and-pepper noise with probabilities $P_a = P_b = 0.25$. (b) Result of filtering with a 7×7 median filter. (c) Result of adaptive median filtering with $S_{\max} = 7$.

Adaptive median
filtering:

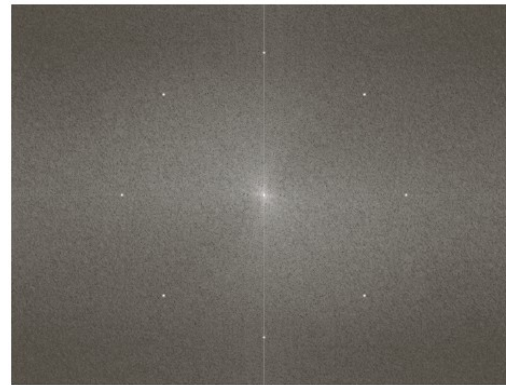
Periodic noise

- Arises from interferences (e.g., electrical or electromechanical) during image acquisition.
- Can be analyzed and filtered quite effectively in the frequency domain using a **band-reject** filter.

image corrupted by sinusoidal noise



spectrum of noisy image



Band-reject filters

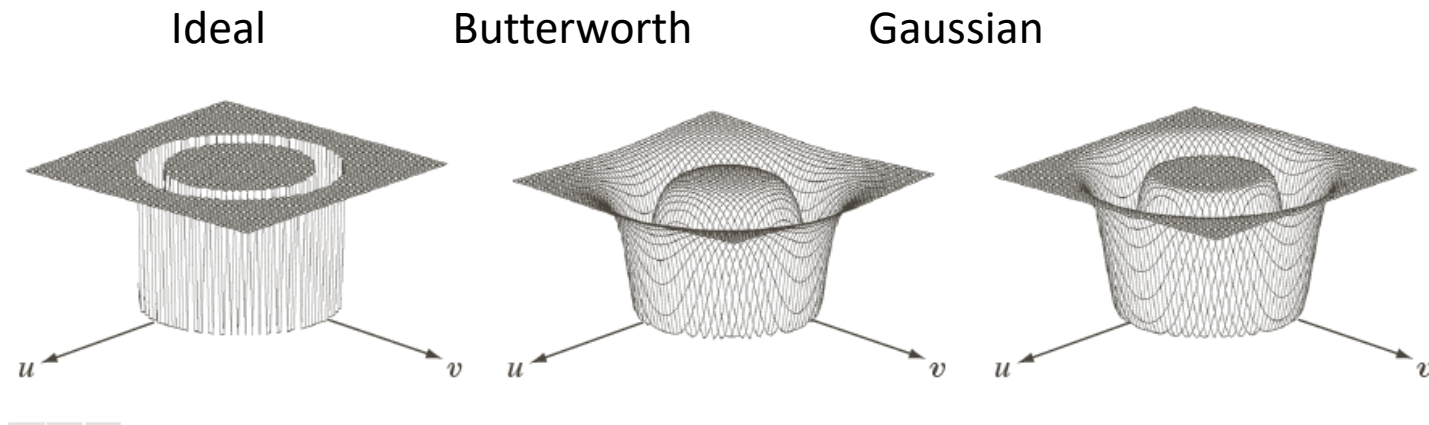


TABLE 4.6

Bandreject filters. W is the width of the band, D is the distance $D(u, v)$ from the center of the filter, D_0 is the cutoff frequency, and n is the order of the Butterworth filter. We show D instead of $D(u, v)$ to simplify the notation in the table.

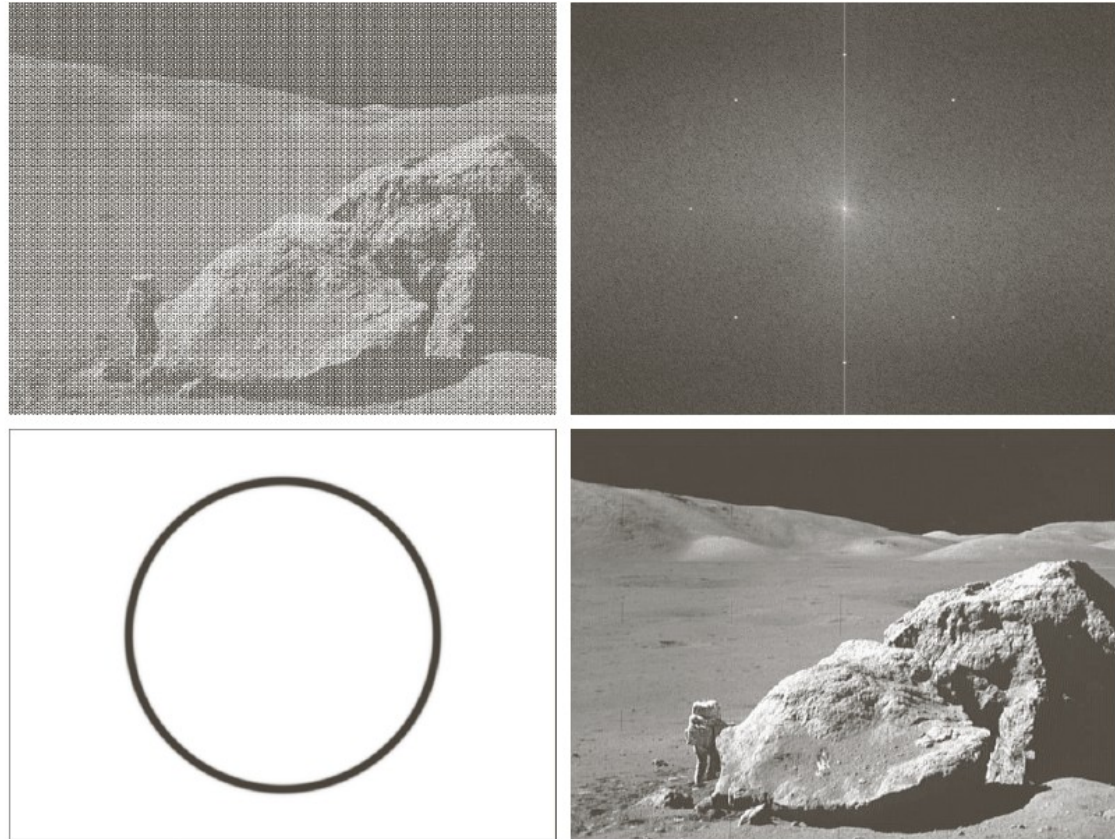
Ideal	Butterworth	Gaussian
$H(u, v) = \begin{cases} 0 & \text{if } D_0 - \frac{W}{2} \leq D \leq D_0 + \frac{W}{2} \\ 1 & \text{otherwise} \end{cases}$	$H(u, v) = \frac{1}{1 + \left[\frac{DW}{D^2 - D_0^2} \right]^{2n}}$	$H(u, v) = 1 - e^{-\left[\frac{D^2 - D_0^2}{DW} \right]^2}$

Band-reject filters (cont'd)

a	b
c	d

FIGURE 5.16

(a) Image corrupted by sinusoidal noise.
(b) Spectrum of (a).
(c) Butterworth bandreject filter (white represents 1).
(d) Result of filtering.
(Original image courtesy of NASA.)



Bandpass filters

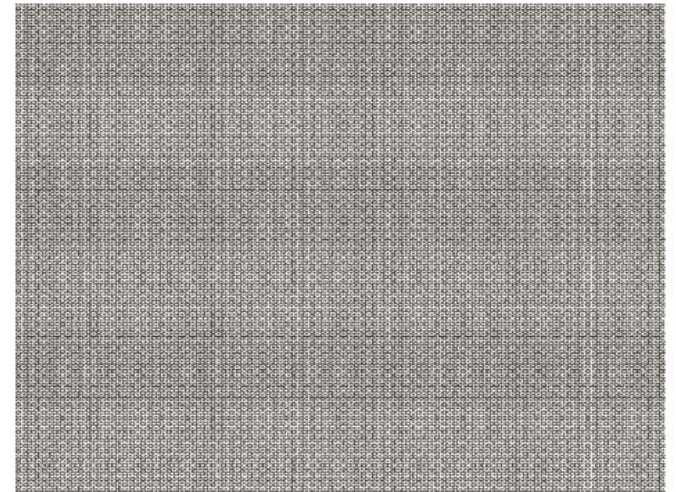
- Performs the opposite operation of a band-reject filter.

$$H_{BP}(u,v) = 1 - H_{BR}(u,v)$$

- Useful in isolating the effect of specific frequencies in an image.



noise
pattern



Estimating degradation H

$$G(u,v)=H(u,v)F(u,v) + N(u,v)$$

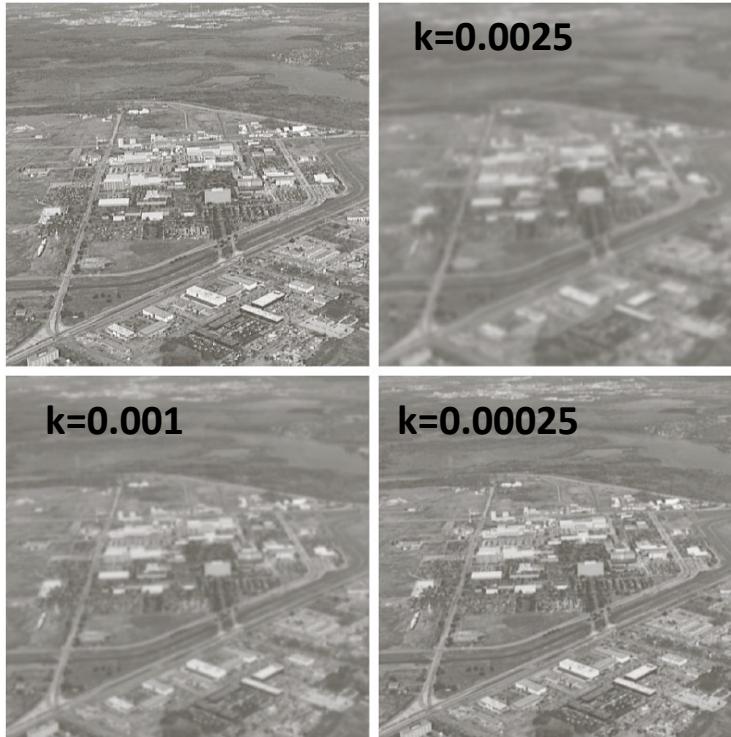
- Typically, H is modeled mathematically; let's look at
two examples:
 - Atmospheric turbulence
 - Motion blurring

Degradation due to environmental conditions

- Atmospheric turbulence (Hufnagel and Stanley [1964])

a b
c d

FIGURE 5.25
Illustration of the
atmospheric
turbulence model.
(a) Negligible
turbulence.
(b) Severe
turbulence,
 $k = 0.0025$.
(c) Mild
turbulence,
 $k = 0.001$.
(d) Low
turbulence,
 $k = 0.00025$.
(Original image
courtesy of
NASA.)



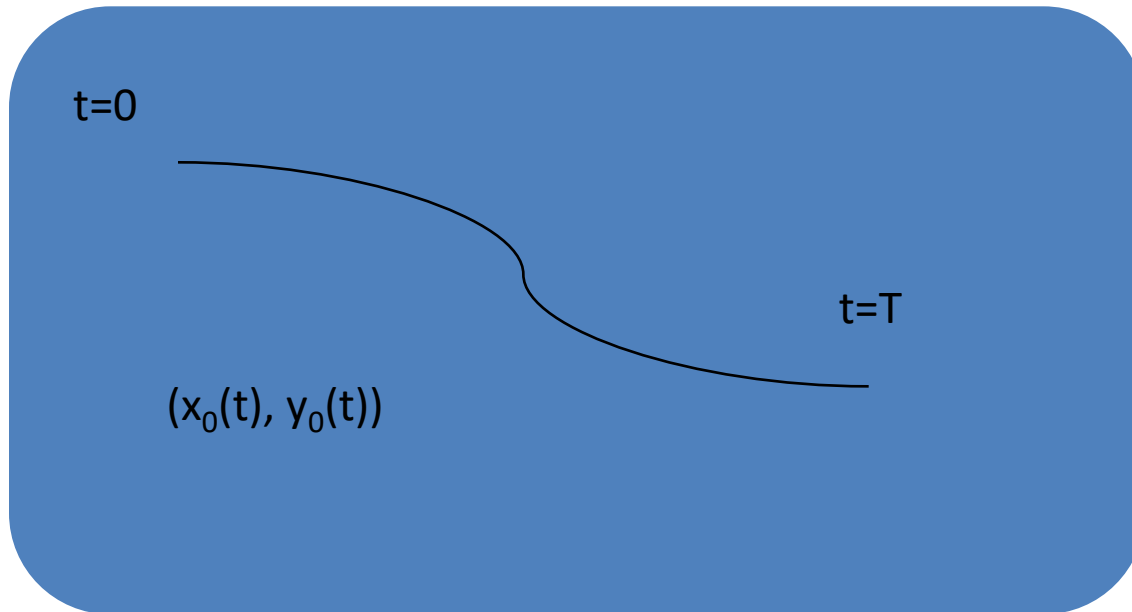
$$G(u,v) = H(u,v)F(u,v) + N(u,v)$$

where:

$$H(u,v) = e^{-k(u^2+v^2)^{5/6}}$$

Degradation due to uniform linear motion

- Consider the case of camera/object **planar motion** (2D) where $x_0(t)$ and $y_0(t)$ is the motion trajectory.



e.g., camera motion



Degradation due to uniform linear motion

- If **T** is the exposure time (i.e., time interval during which the camera shutter is open), then the output image $g(x,y)$ is:

$$g(x, y) = \int_0^T f(x - x_0(t), y - y_0(t)) dt$$

- Taking the FT: $G(u,v) = H(u,v)F(u,v)$

where

$$H(u, v) = \int_0^T e^{-j2\pi(ux_0(t) + vy_0(t))} dt$$

Degradation due to uniform linear motion (cont'd)

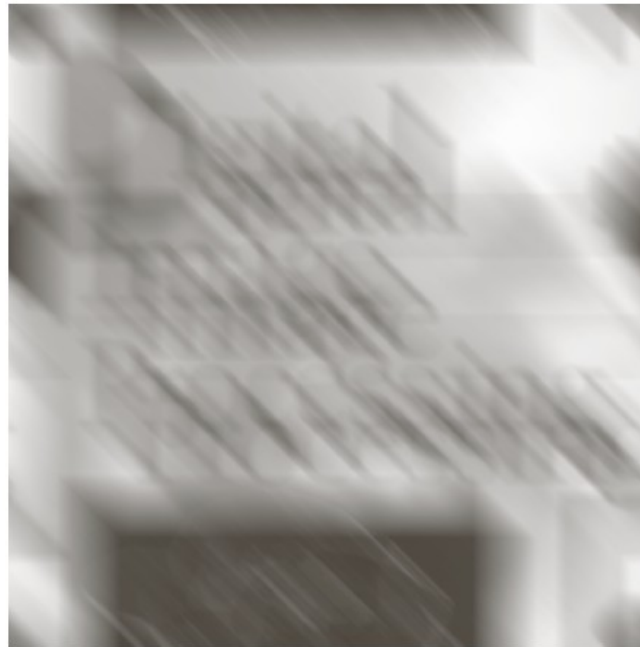
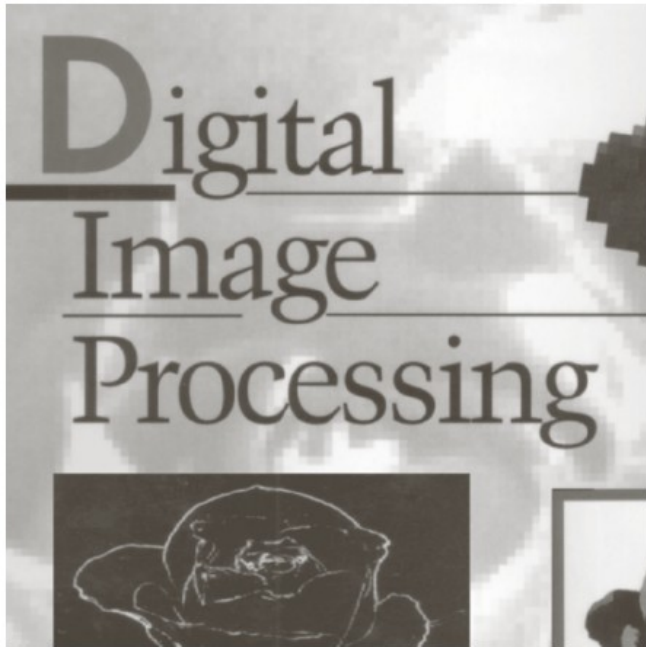
- If $x_0(t)=\alpha t/T$ and $y_0(t)=0$, then:

$$H(u, v) = \int_0^T e^{-j2\pi(ux_0(t))} dt = \int_0^T e^{-j2\pi u \alpha t / T} dt = \frac{T}{\pi u \alpha} \sin(\pi u \alpha) e^{-j\pi u \alpha}$$

- If $x_0(t)=\alpha t/T$ and $y_0(t)=bt/T$, then:

$$H(u, v) = \frac{T}{\pi(u\alpha + vb)} \sin(\pi(u\alpha + vb)) e^{-j\pi(u\alpha + vb)}$$

Degradation due to uniform linear motion (cont'd)



a b

FIGURE 5.26

(a) Original image.

(b) Result of blurring using the function in Eq. (5.6-11) with

$a = b = 0.1$ and $T = 1$.

$G(u,v)=H(u,v)F(u,v)$ where:

$$H(u,v) = \frac{T}{\pi(ua + vb)} \sin(\pi(ua + vb)) e^{-j\pi(ua + vb)}$$