

In the beginning of the Dark Ages, electrically neutral hydrogen gas filled the universe. As stars formed, they ionized the regions immediately around them, creating bubbles here and there. Eventually these bubbles merged together, and intergalactic gas became entirely ionized.

# Cosmic Dawn & Square Kilometer Array

Jonathan Pritchard  
(Astrophysics)

RADIO QUIET ZONE

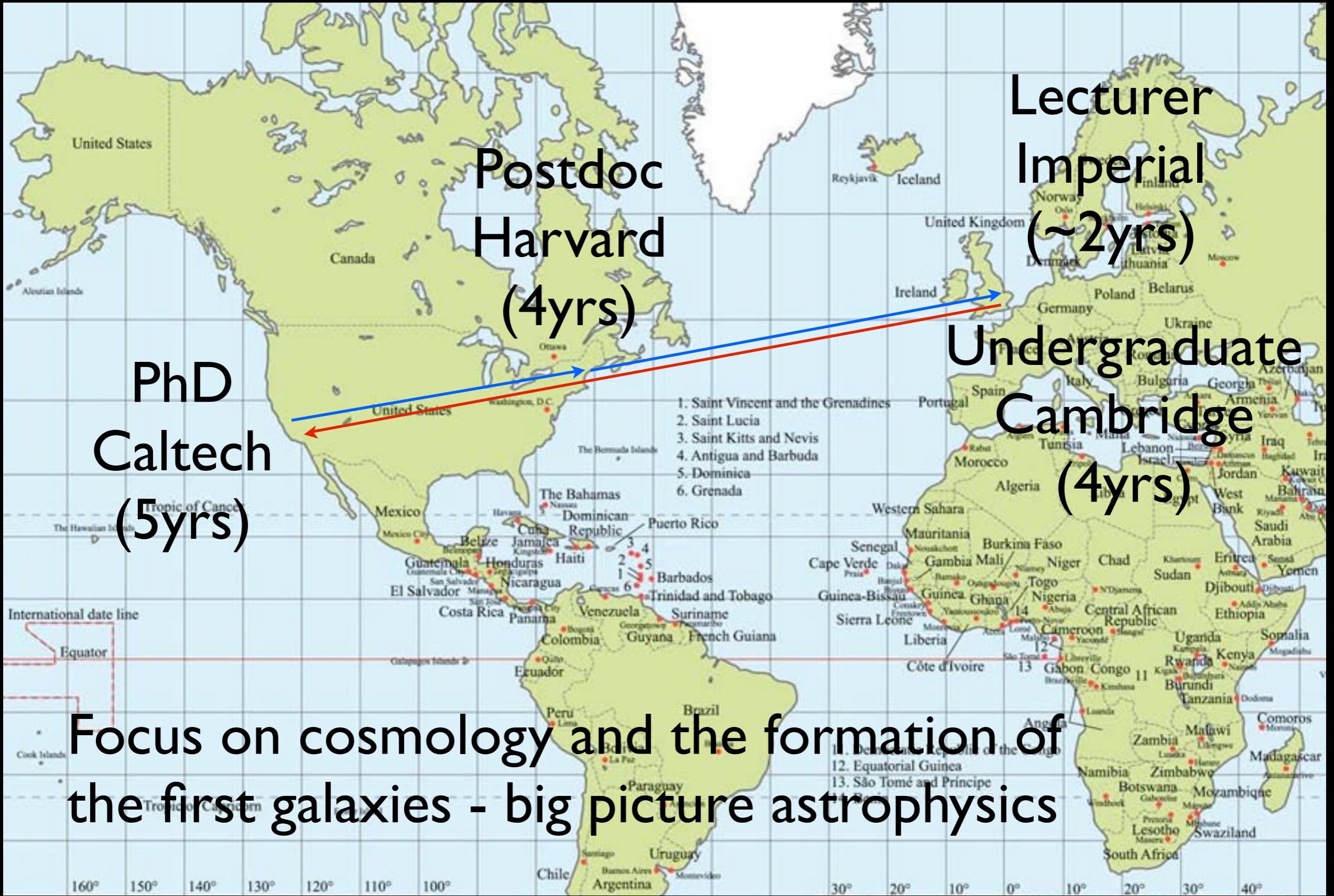


SWITCH OFF  
MOBILE PHONES

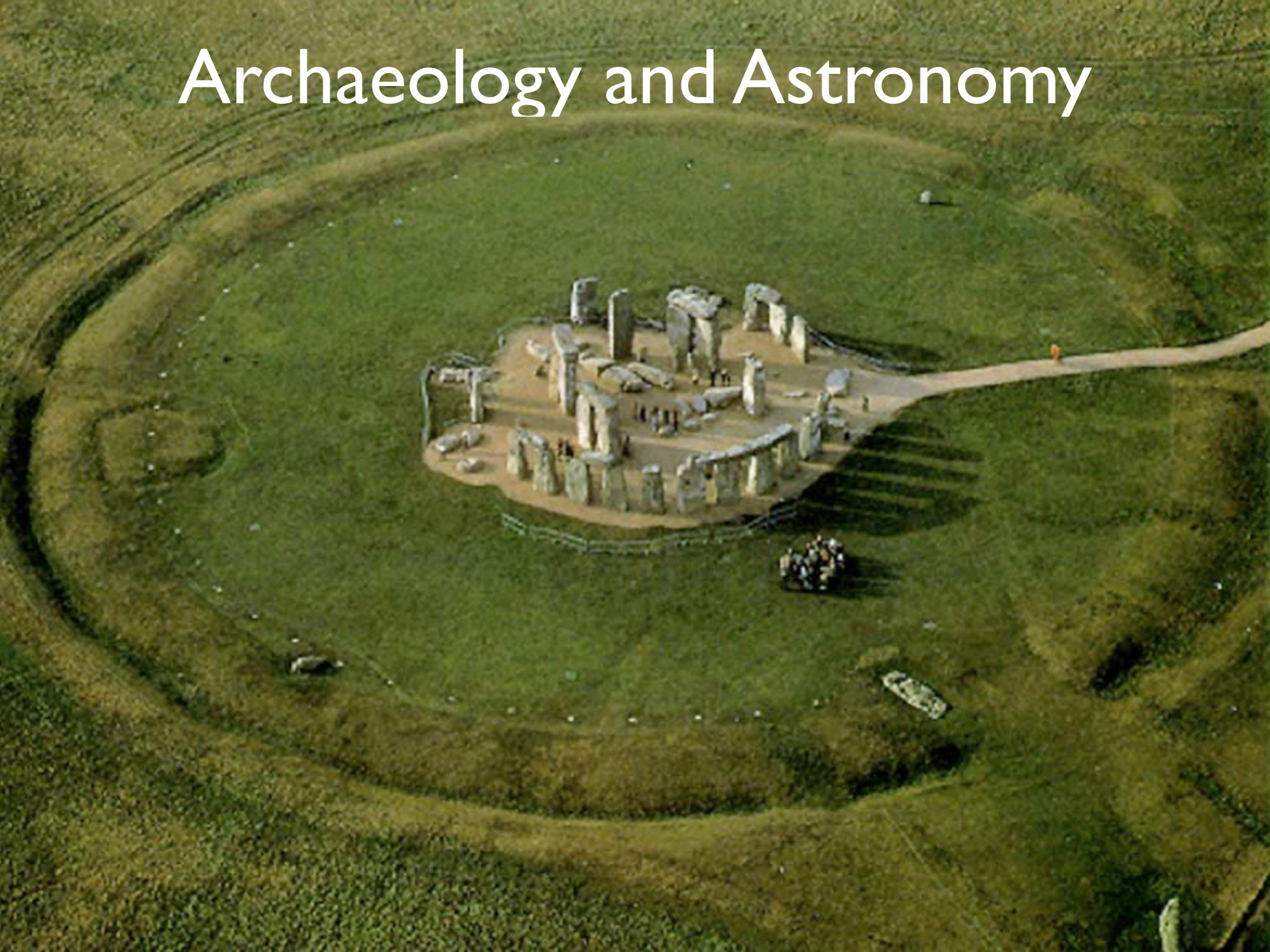
10



# Who am I?



# Archaeology and Astronomy

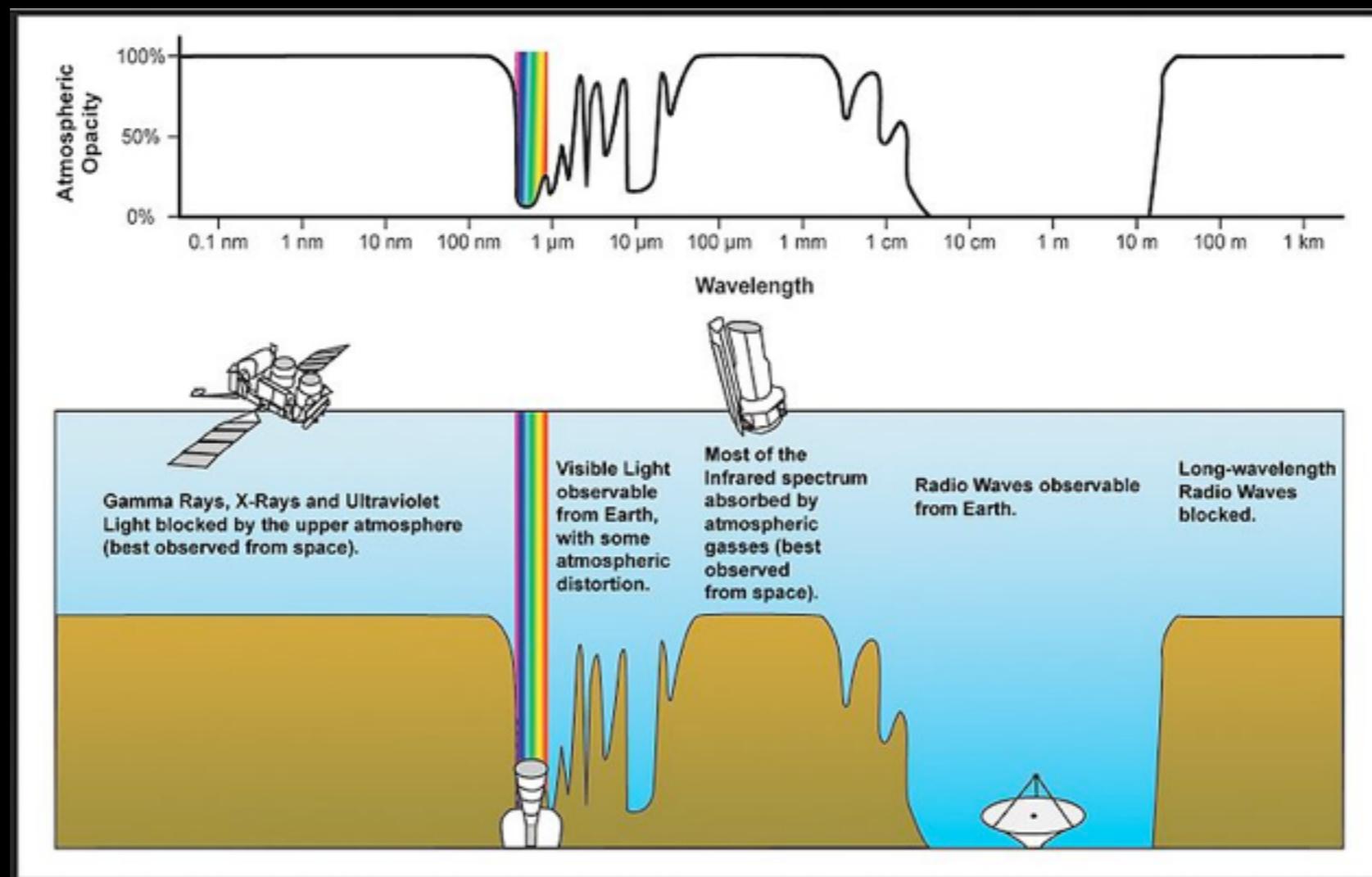


# Seeing the past

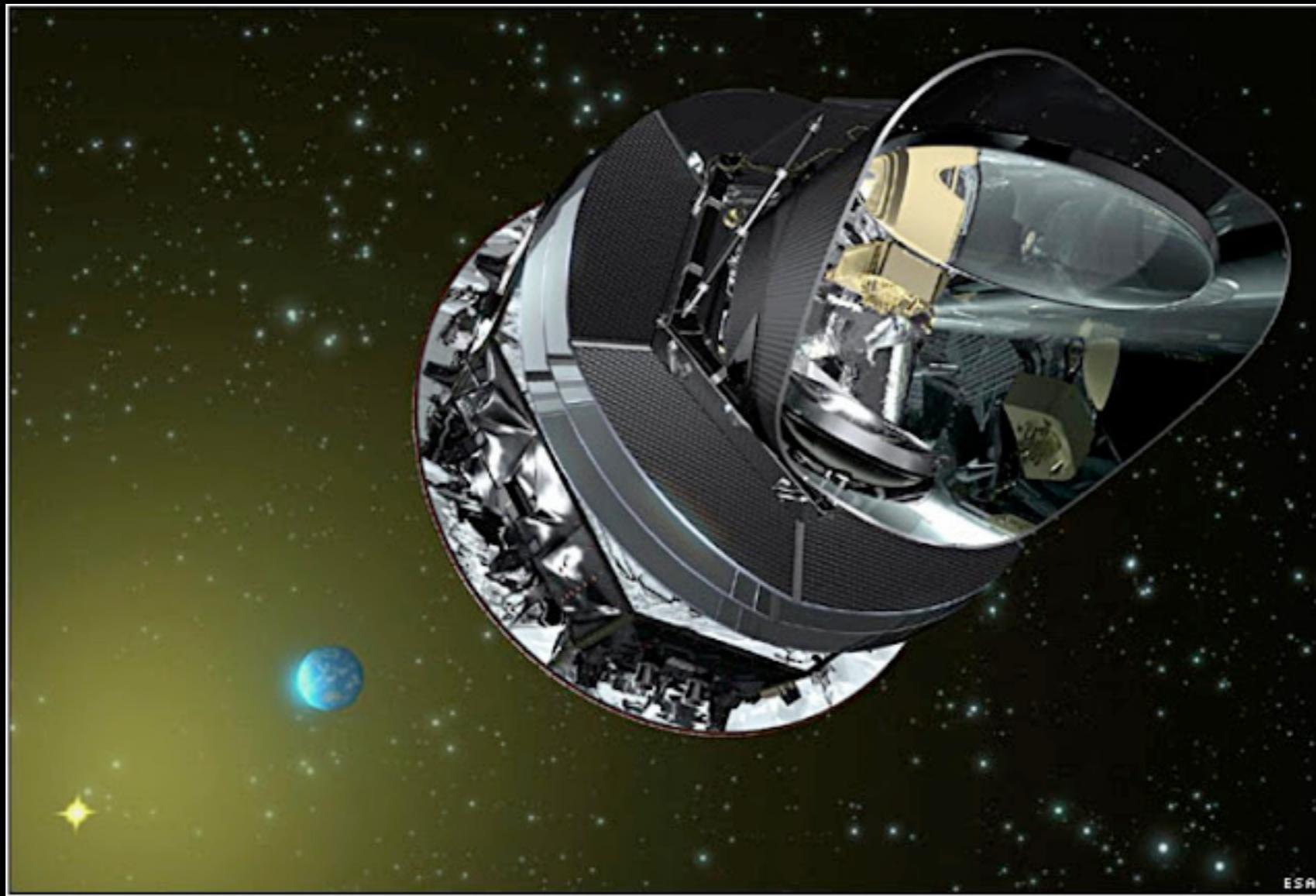
- Light travels at 300,000 km/s. Fast, but finite.
- When we see distant objects, we're seeing them as they were when that light was emitted
- We're seeing the past
- Moon is ~1 light seconds away  
Sun is ~8 light minutes away
- We don't get to choose when we see. Each distance corresponds to a specific time in cosmic history

# What do we see?

- Different frequencies show us different objects at different distances.
- Limits set by technology and absorption along the line of sight

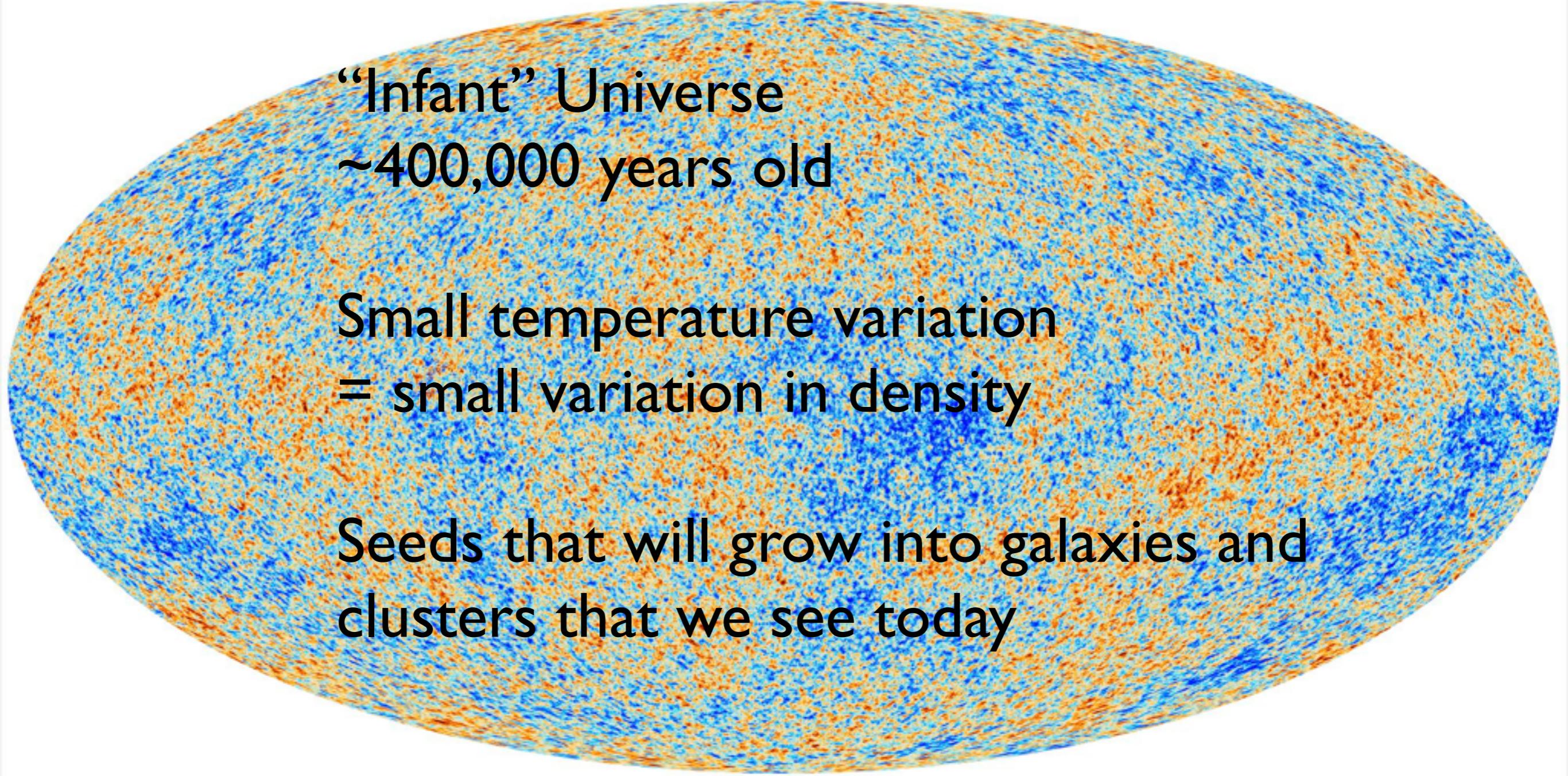


# Cosmic Microwave Background



Planck satellite

# Cosmic Microwave Background



“Infant” Universe  
~400,000 years old

Small temperature variation  
= small variation in density

Seeds that will grow into galaxies and  
clusters that we see today

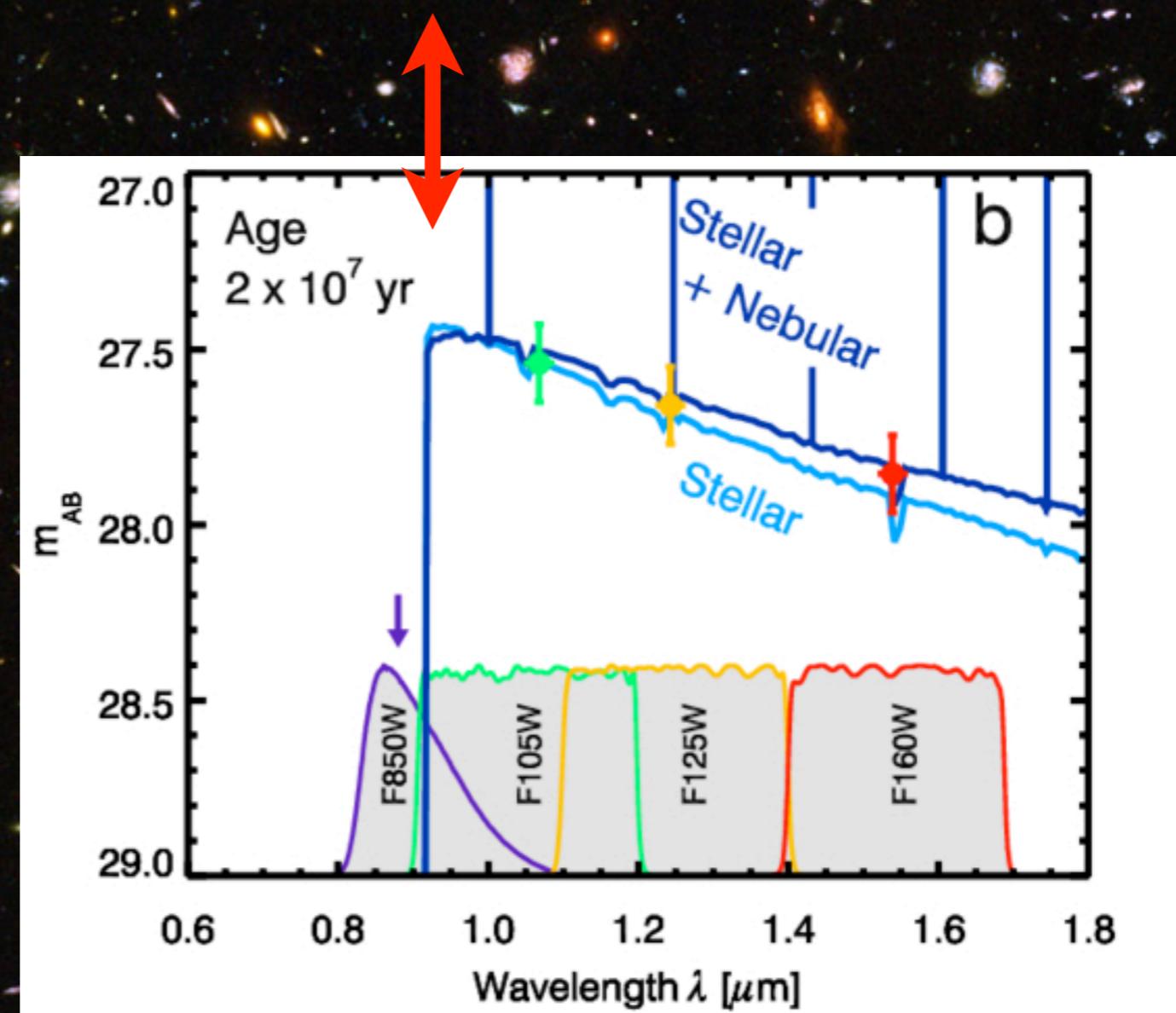
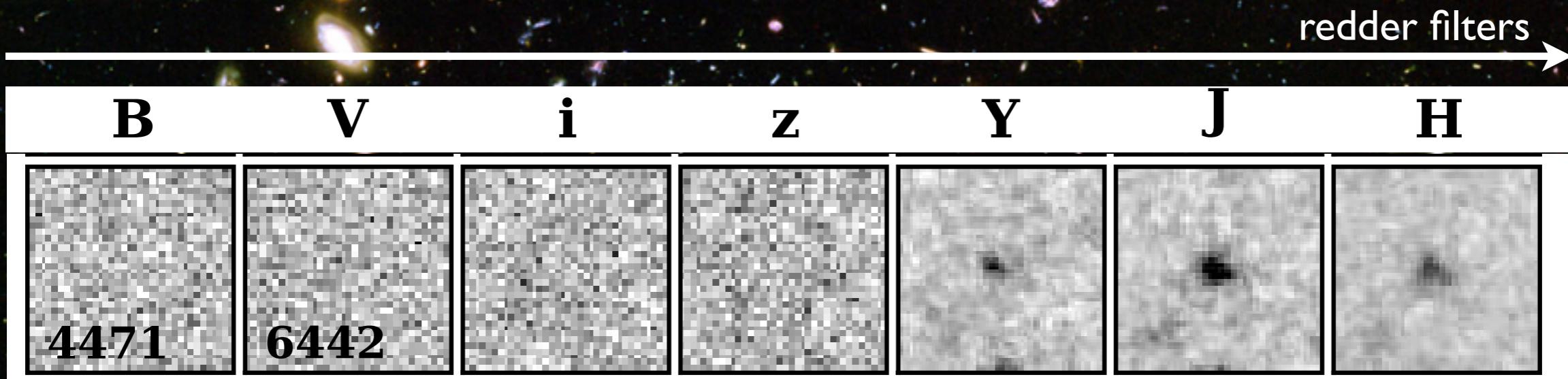
# Hubble Space Telescope



# Hubble Ultra Deep Field

- Patch smaller than the moon on sky
- Furthest galaxies are located at distances when Universe was 1 billion years old
- “Mature” Universe

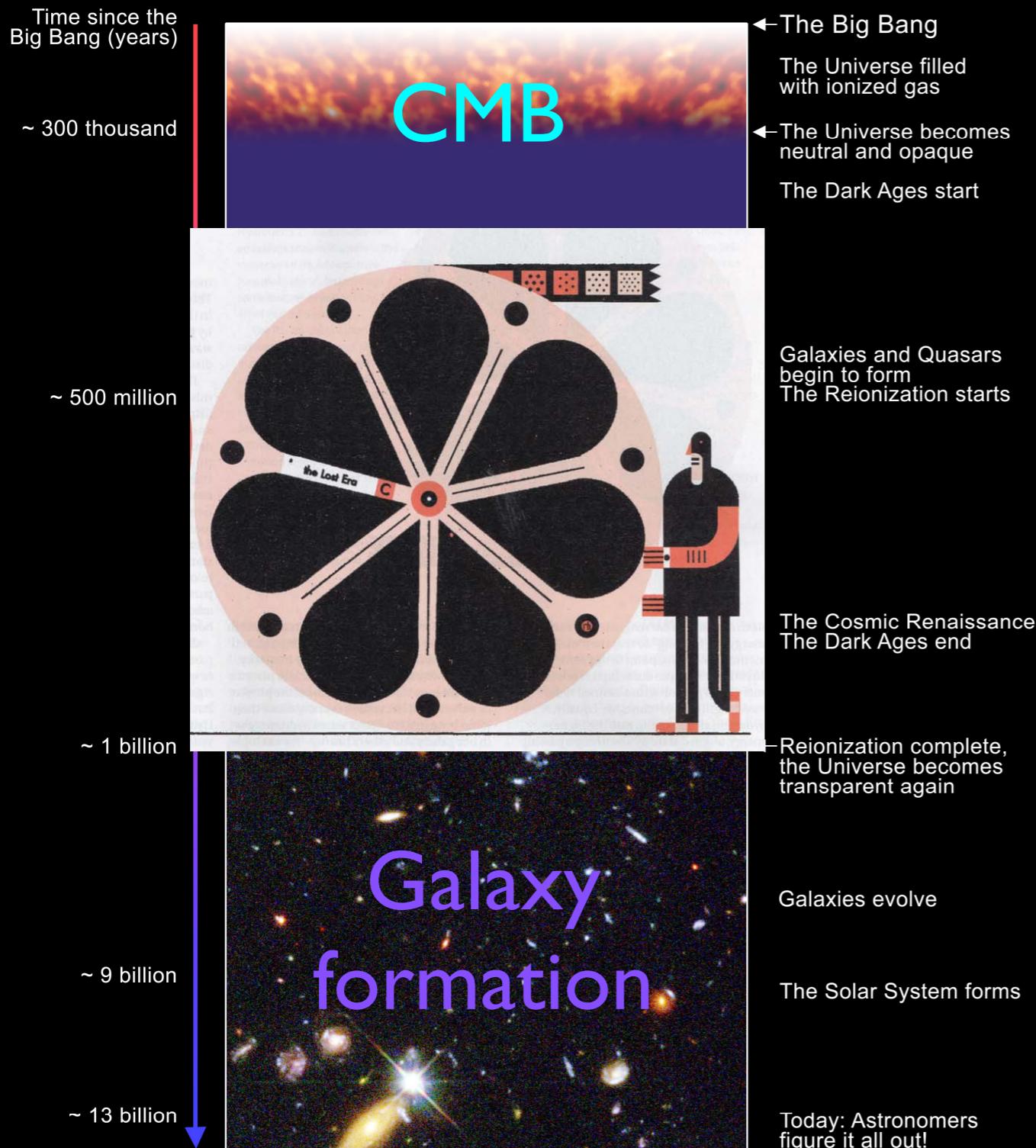
Hubble (and soon JWST) identify high redshift galaxies as  
“drop outs”



# The Cosmic Timeline

## What is the Reionization Era?

A Schematic Outline of the Cosmic History



- CMB shows initial conditions
- Deep galaxy surveys show later galaxies
- What went on in between?
- Currently NO observations!



# The Missing Reel

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DIRECTOR

T. Universe

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CAMERAMAN

S.K.A.

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DATE

13.75 GYrABB

TAKE

1

SCENE

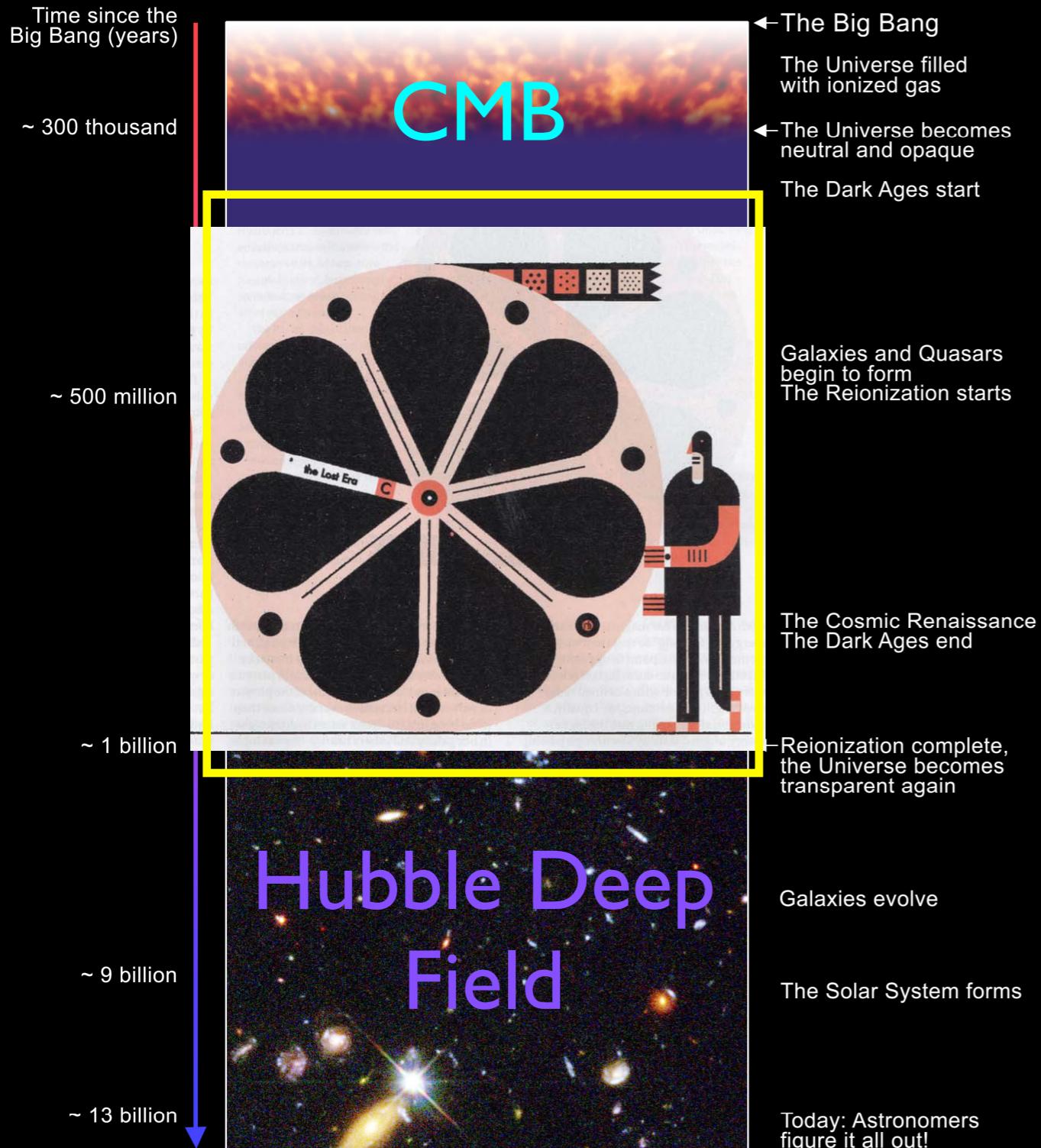
10<sup>51</sup>

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# The Missing Reel

## What is the Reionization Era?

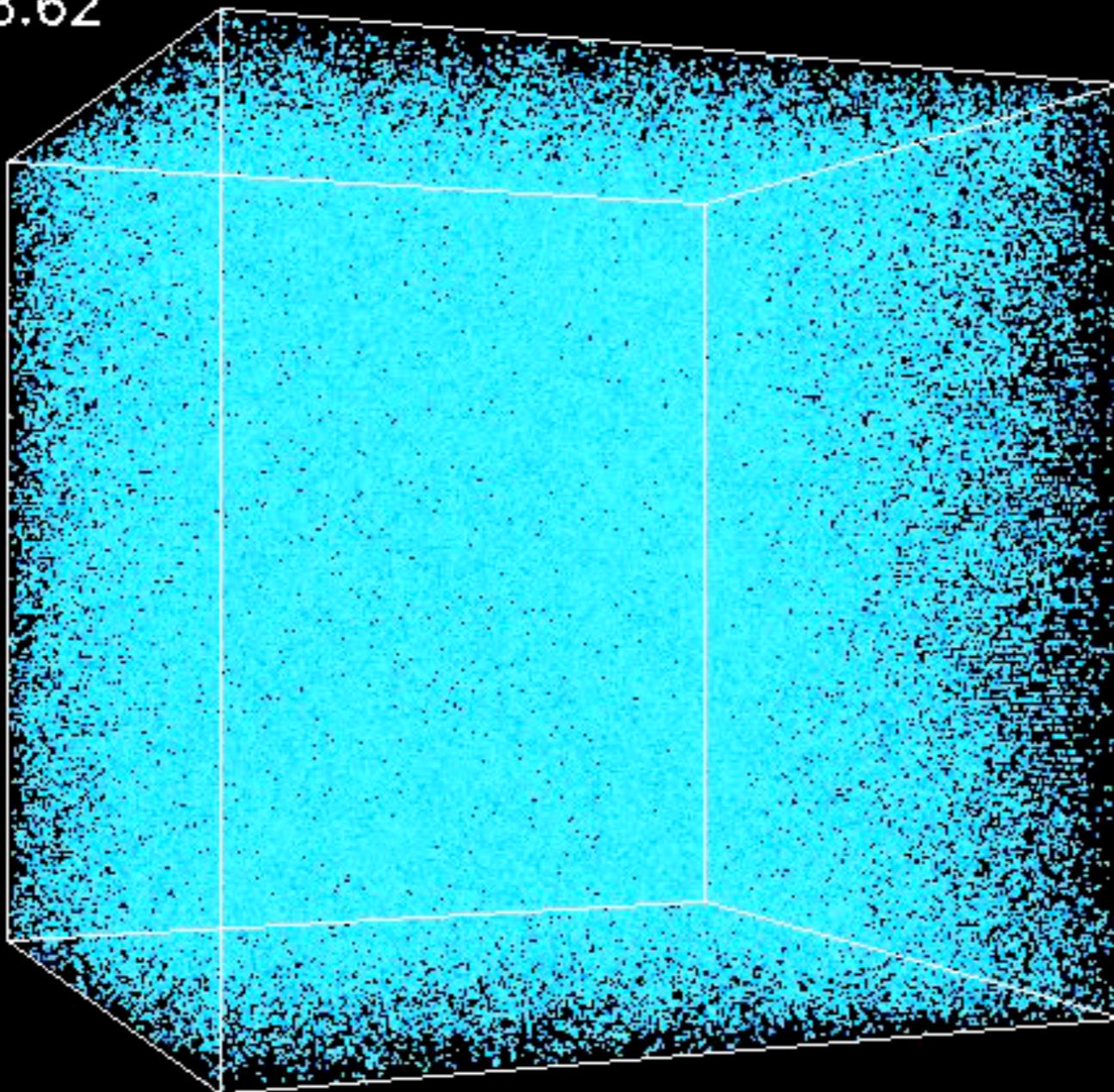
A Schematic Outline of the Cosmic History



- Dark Ages
- Cosmic Dawn
- Reionization

# Dark sure, but interesting

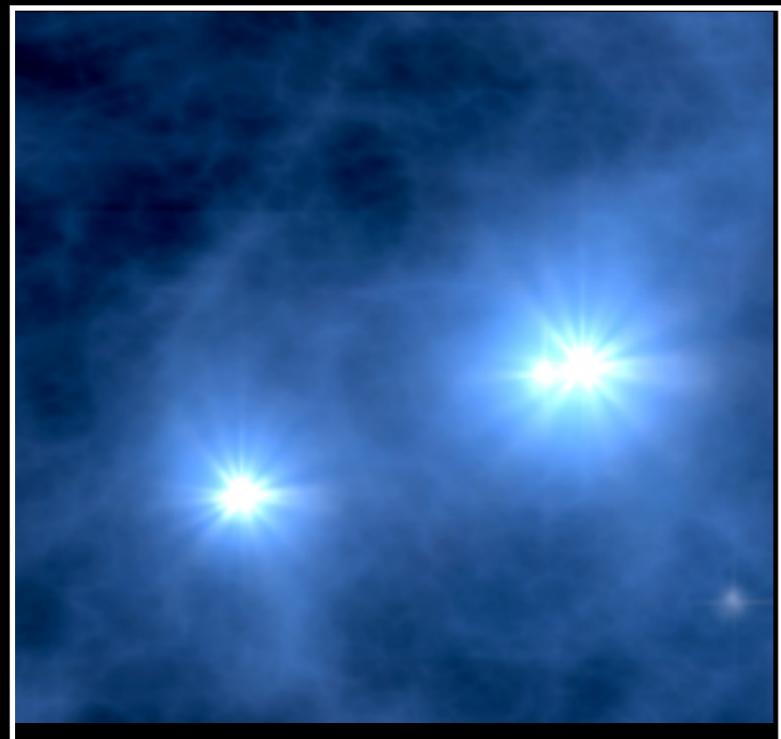
Z=28.62



Structures grow as gravity causes collapse

# Fire breathing monsters?

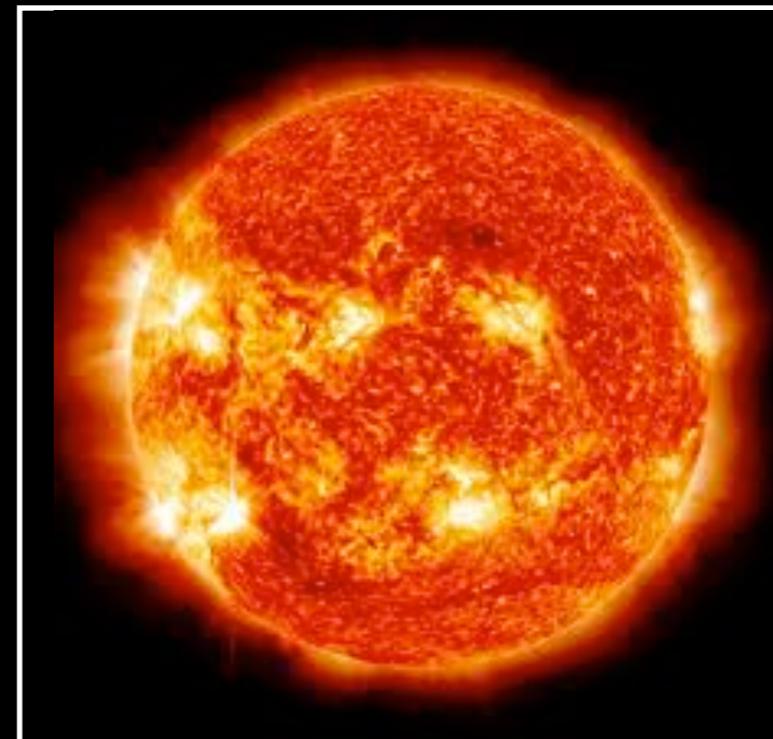
- First stars formed from primordial gas - hydrogen and helium
- Later stars form from gas enriched with metals - carbon, oxygen, iron,...
- First stars may have been larger, hotter, brighter  
=> live fast, die young (Pop III lifetime  $\sim$  1 million years)



Population III  
Metal-free     $\sim 100 \text{ Msol}$   
                   $\sim 20 \text{ Lsol}$



Population II  
Metal-poor  
(galactic bulge/globular clusters)



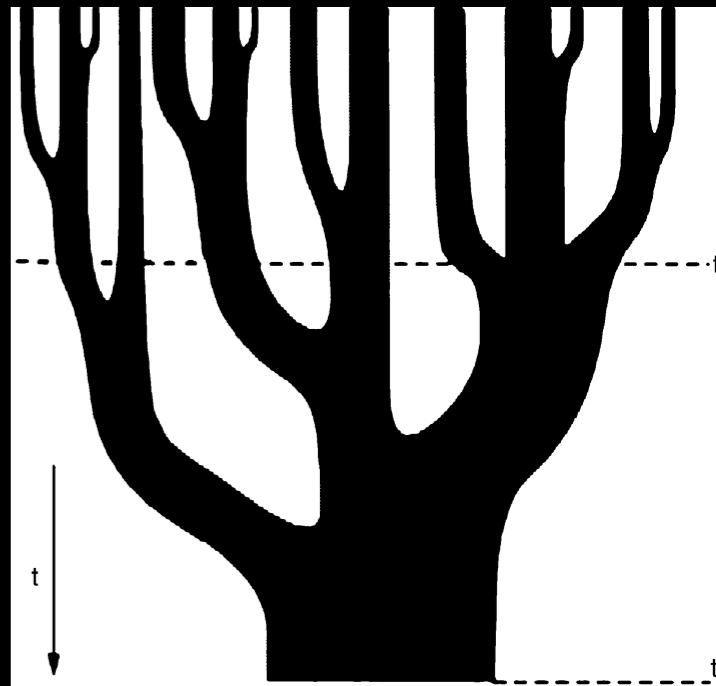
Population I  
Metal-rich  
(spiral arms of Milky way)  
Our sun

# First stars



# Galactic archaeology

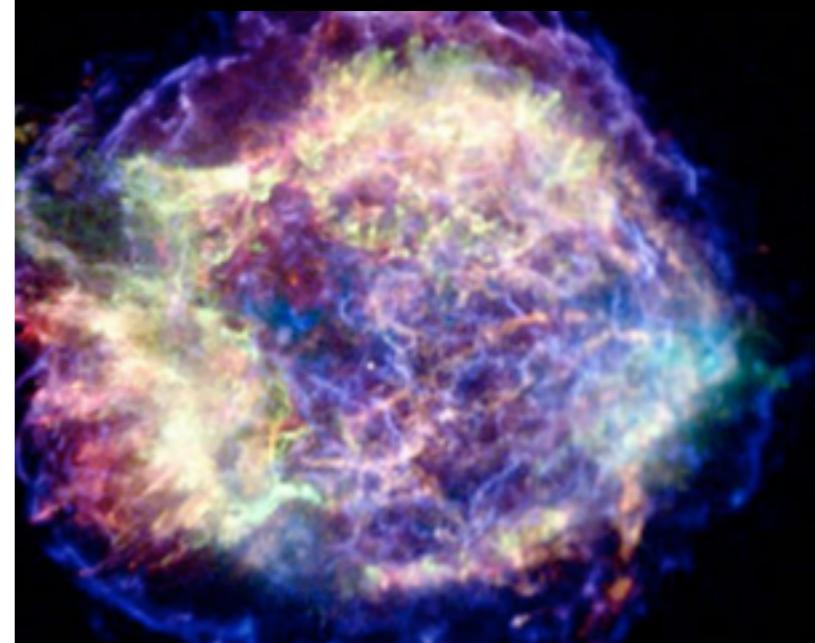
- “hierarchical structure formation”
  - small galaxies merge to form larger ones
- Can search for oldest stars in our galaxy as **low-mass metal-poor stars**
- Less fossils, more **very old citizens of Universe**



## LETTER

doi:10.1038/nature12990

A single low-energy, iron-poor supernova  
as the source of metals in the star SMSS  
J031300.36–670839.3



S. C. Keller<sup>1</sup>, M. S. Bessell<sup>1</sup>, A. Frebel<sup>2</sup>, A. R. Casey<sup>1</sup>, M. Asplund<sup>1</sup>, H. R. Jacobson<sup>2</sup>, K. Lind<sup>3</sup>, J. E. Norris<sup>1</sup>, D. Yong<sup>1</sup>, A. Heger<sup>4</sup>, Z. Magic<sup>1,5</sup>, G. S. Da Costa<sup>1</sup>, B. P. Schmidt<sup>1</sup> & P. Tisserand<sup>1</sup>

# First galaxies

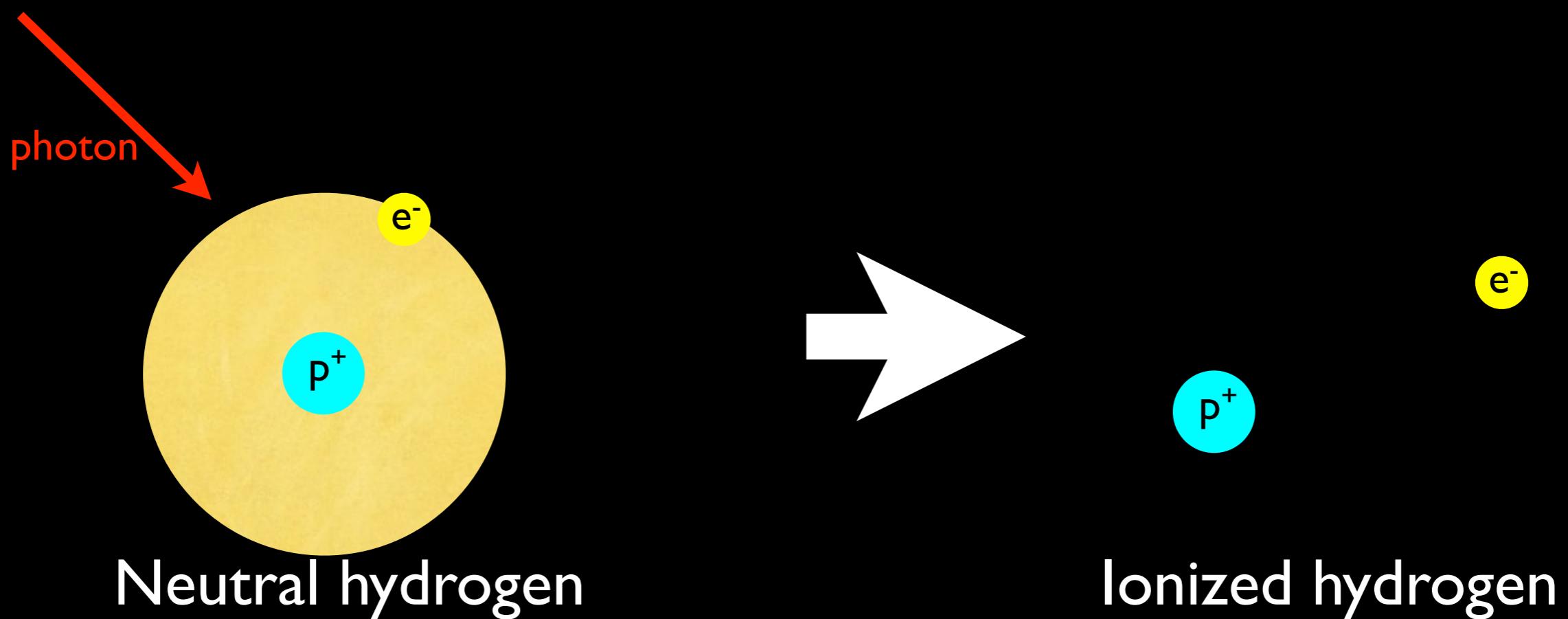
- Galaxy is a collection of gravitationally bound stars
- First star to form altered surroundings
  - heating, ionization, metal enrichment, winds,...
- IF feedback prevents star formation for a while then galaxies are “bursty” - 1 star at a time
- IF feedback mild then may get more continuous star formation
- Prima dona vs team players...

# What sort of light?

- As stars live and die they produce remnants - neutron stars, black holes
- Black holes can coalesce at center of galaxy and grow rapidly => supermassive black hole
- Accretion of gas onto black hole releases energy as radiation - can out shine galaxy = Active Galactic Nuclei (AGN)
- Jet of non-thermal emission + hot X-rays

# Blowing bubbles

- Light more energetic than 13.6 eV can ionize hydrogen
- Ionized bubbles form around galaxies and grow, merging over time.
- Ultimately, the space between galaxies becomes entirely ionized



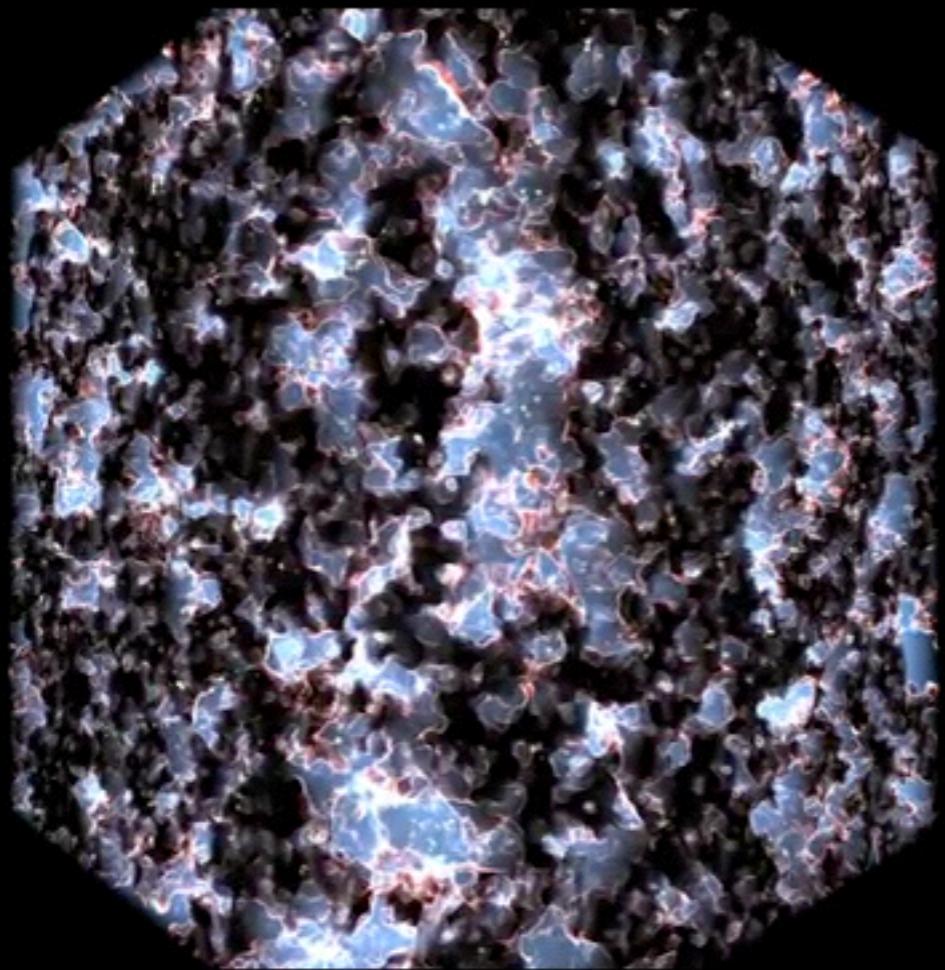
# Reionization



# The last phase transition???

- Reionization is the last major phase transition in the Universe from space being filled with cold, neutral gas to hot, ionized gas.
- When and how long did reionization take place?
- What were the sources that drove it?
  - Massive metal-free Population III stars
  - Many metal-poor Population II stars
  - Accretion onto supermassive black holes (AGN)
- How did ionized bubbles grow and merge?
  - small bubbles around individual galaxies
  - larger bubbles around groups of galaxies

# What do we know?



Reionization



# Reionization

# Casablanca

Cast:

dark matter halos  
Population III stars  
Population II stars  
Galaxies  
Black holes  
Dark stars  
Miniquasar

...

Cast:

Taxi driver  
Police man  
Rick  
Souk merchant  
Lazlo  
Ilsa  
Man with camel

...

Know the cast, but  
who are the leads and who the bit players?

# Reionization

Plot:

Universe was neutral

Luminous sources form

Universe became ionized

# Casablanca

Plot:

Boy meets girl

Boy falls for girl

Boy loses girl

Know the plot highlights,  
but what are the details?



# Cosmic movie making

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DIRECTOR

T. Universe

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CAMERAMAN

S.K.A.

---

DATE

13.75 GYrABB

TAKE

1

SCENE

$10^{51}$

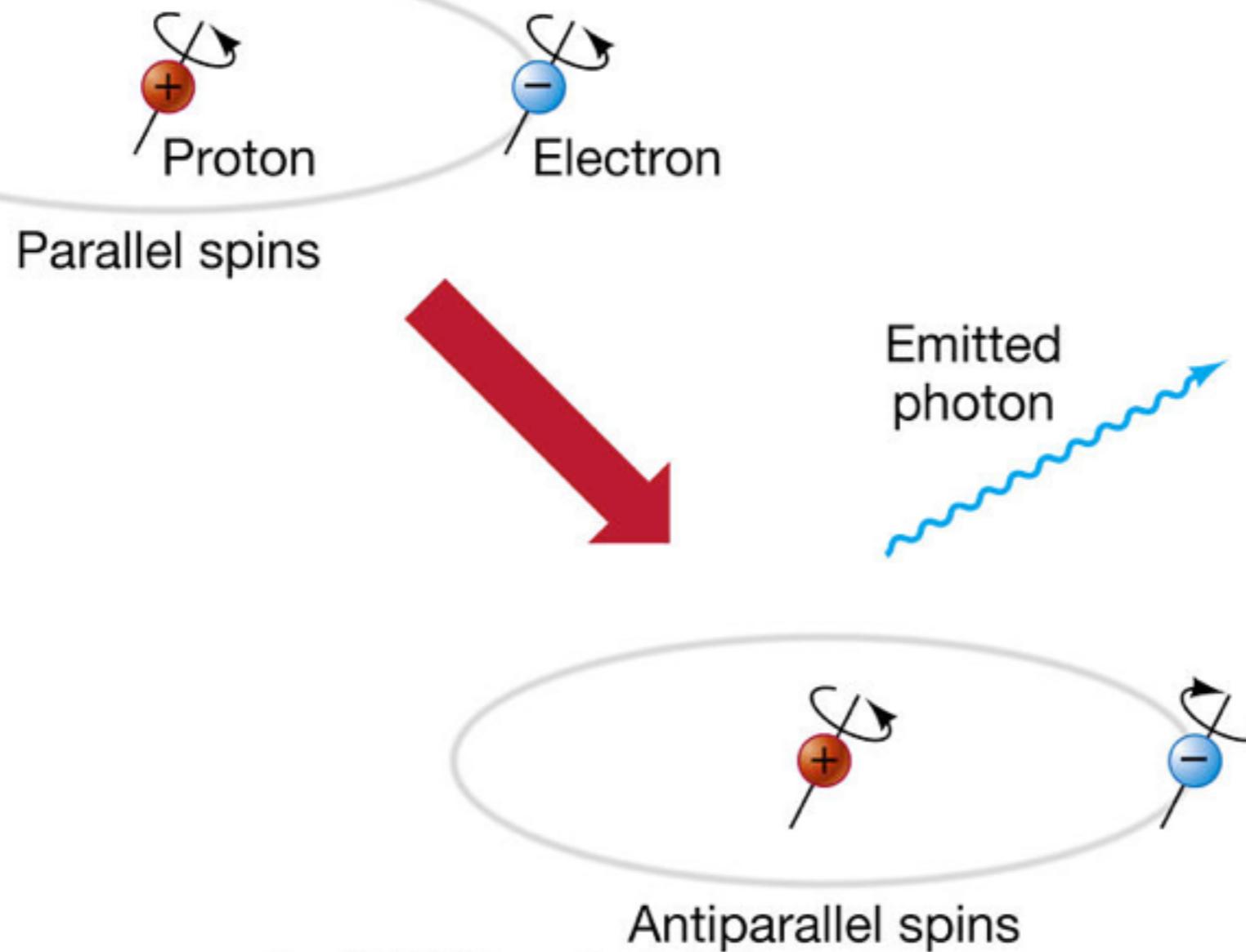
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# How to observe reionization?

- Should be able to see some of the bright galaxies with the James Webb Space Telescope  
**BUT**  
faint galaxies are too faint. Sad since they're numerous
- Really want to see the hydrogen gas that gets ionized  
**BUT**  
light from ionization and recombination of hydrogen gets blocked before it gets to us
- Need a new technique to map the hydrogen

# The 21 cm line

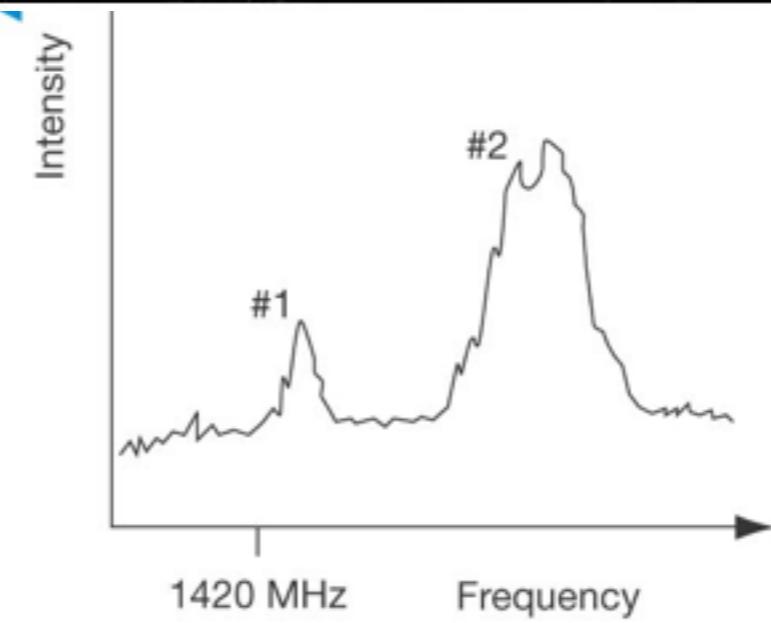
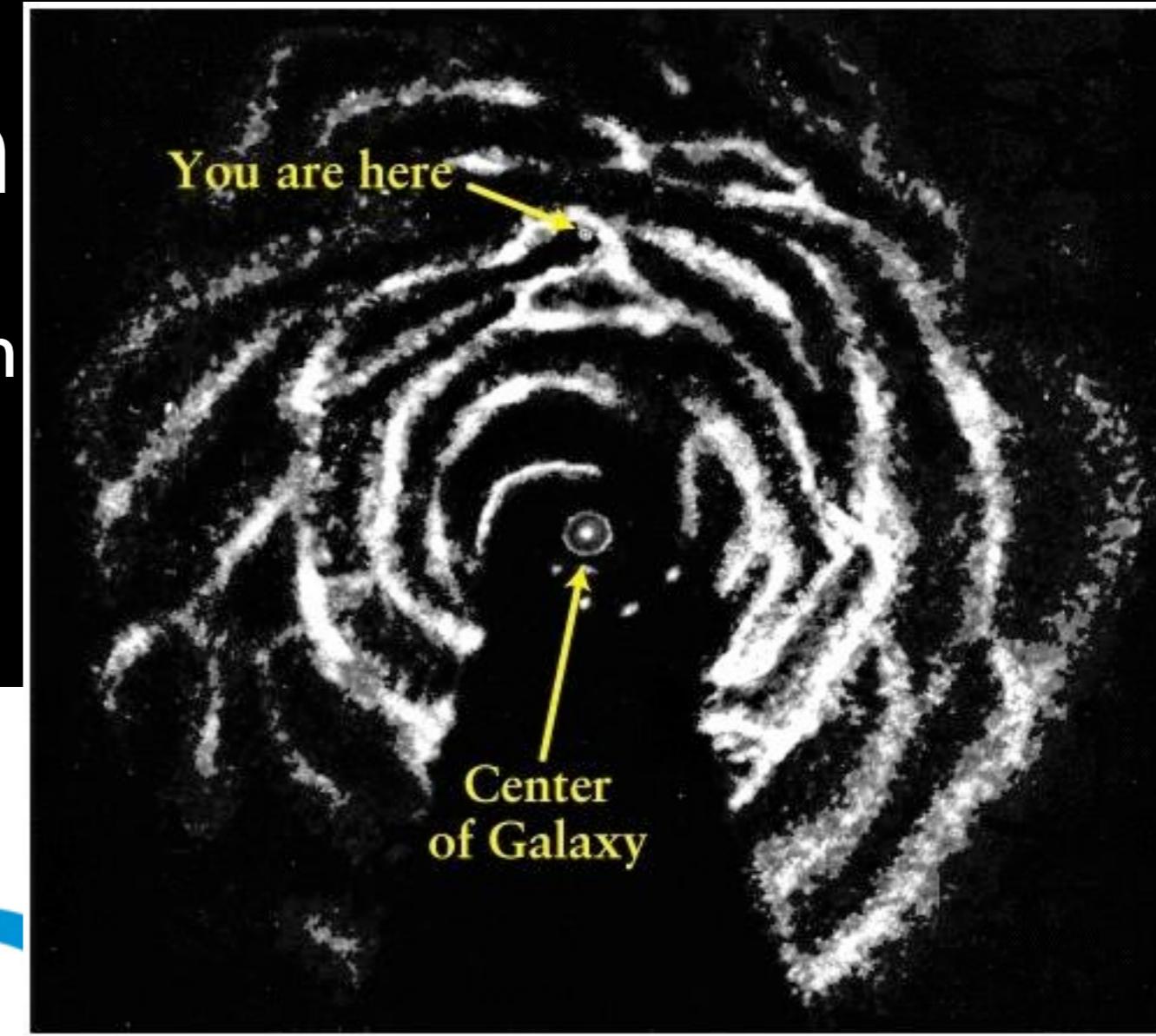
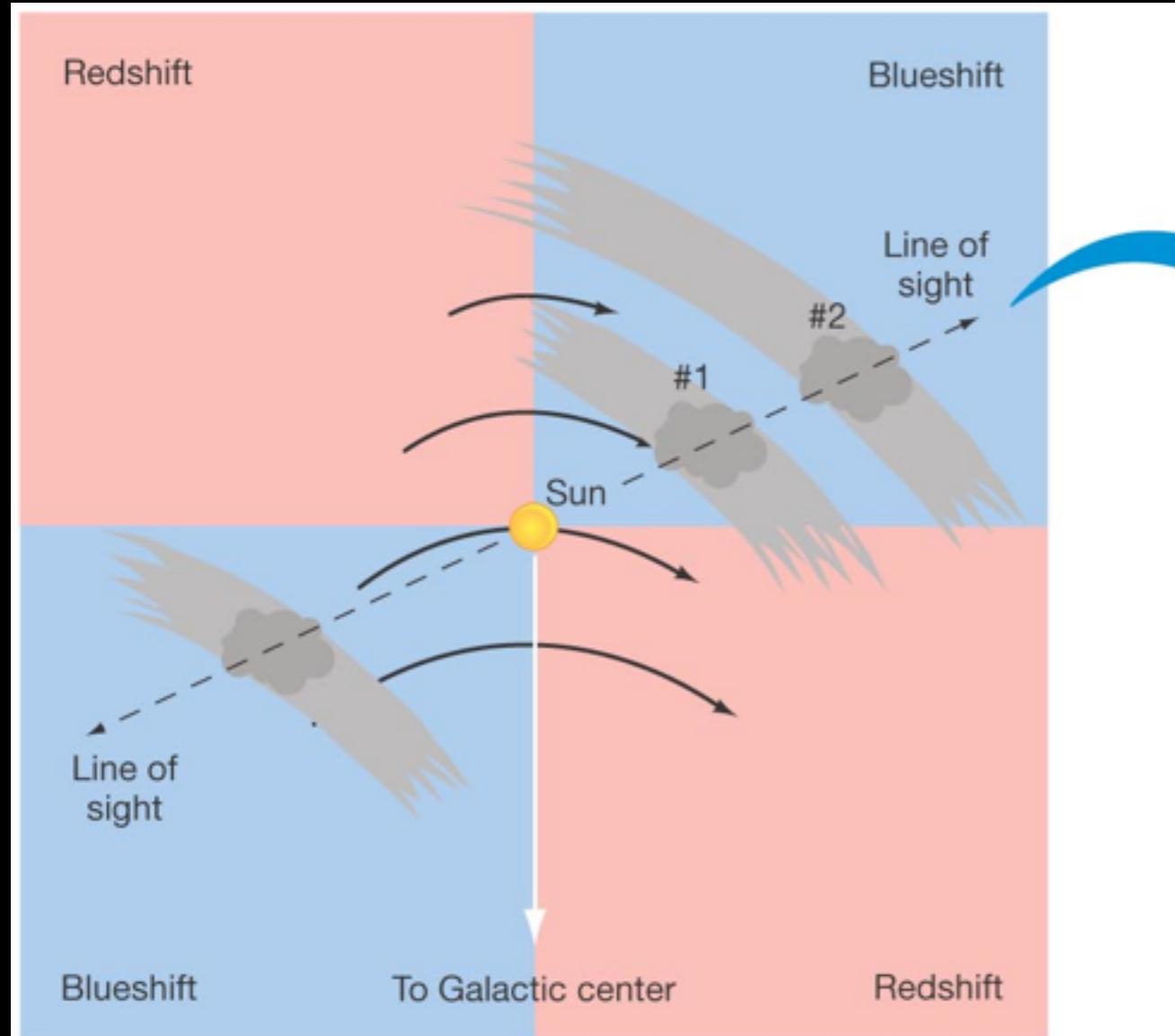
- Hydrogen can emit or absorb radio waves with wavelength of 21cm from spin flips
- These radio waves can propagate to us today



$$21\text{cm} = 1.4 \text{ GHz}$$

# Nearby 21 cm

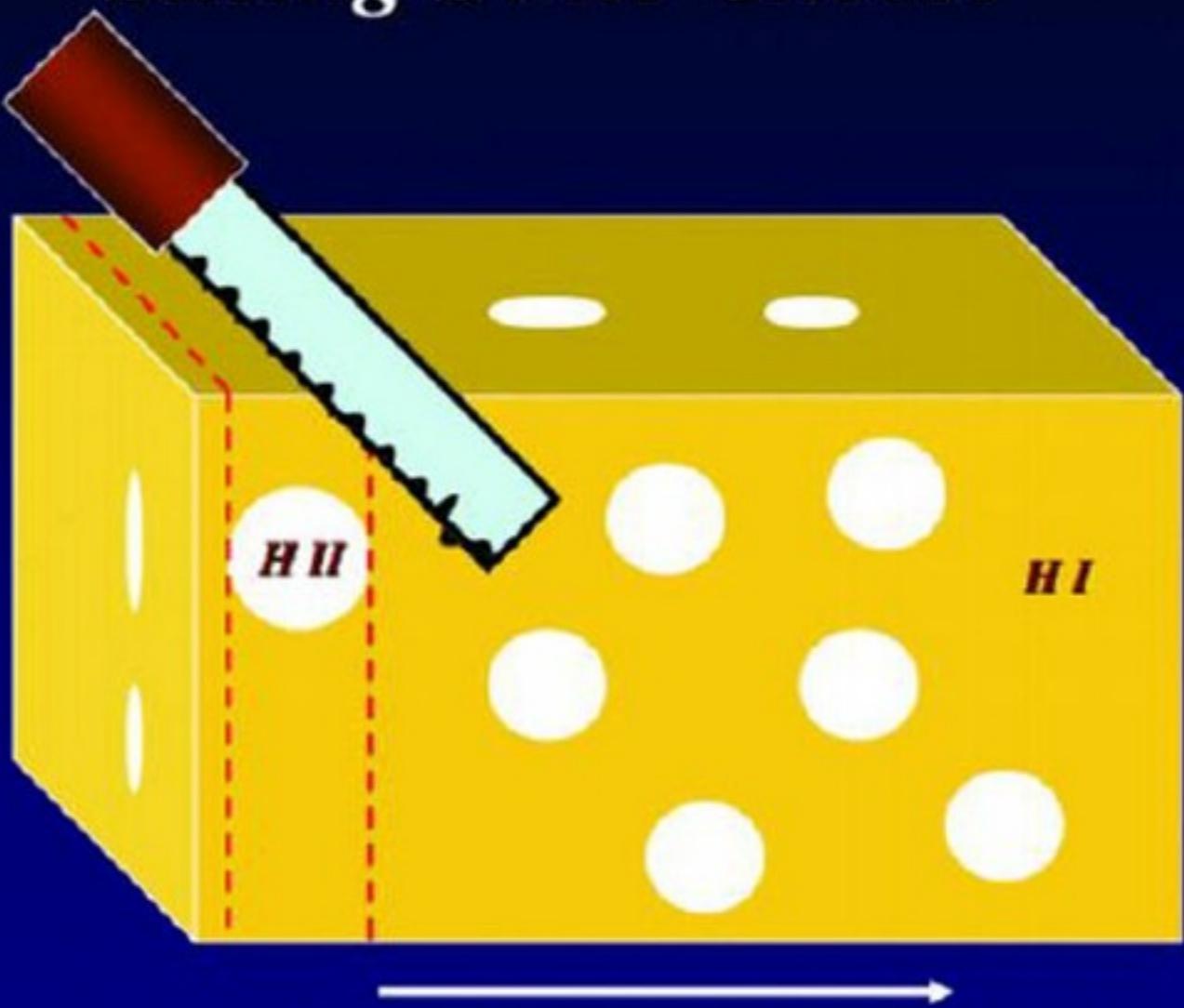
- 21 cm light will be shifted in
- Used in nearby Universe to



# Cutting Swiss Cheese

- At cosmic distances, expansion of Universe produces Doppler shift = redshifting

*21cm Tomography of Ionized Bubbles During Reionization is like  
Slicing Swiss Cheese*



$$\text{Observed wavelength} \leftrightarrow \text{distance}$$
$$21\text{cm} \times (1 + z)$$

# Where should we look?

- Frequencies **40-200MHz** correspond to **100million - 1billion years** after the Big Bang

1420 MHz =>  $z=0$

$t_{\text{Age}}(z=0) \sim 13.7 \text{ Gyr}$

200 MHz  $\rightarrow z = 6$

$t_{\text{Age}}(z = 6) \approx 1 \text{ Gyr}$

100 MHz  $\rightarrow z = 13$

$t_{\text{Age}}(z = 10) \approx 500 \text{ Myr}$

70 MHz  $\rightarrow z \approx 20$

$t_{\text{Age}}(z = 20) \approx 150 \text{ Myr}$

50 MHz =>  $z \sim 27$

$t_{\text{Age}}(z=27) \sim 100 \text{ Myr}$

# Radio telescopes of the past

## NEW RADIO WAVES TRACED TO CENTRE OF THE MILKY WAY

Mysterious Static, Reported  
by K. G. Jansky, Held to  
Differ From Cosmic Ray.

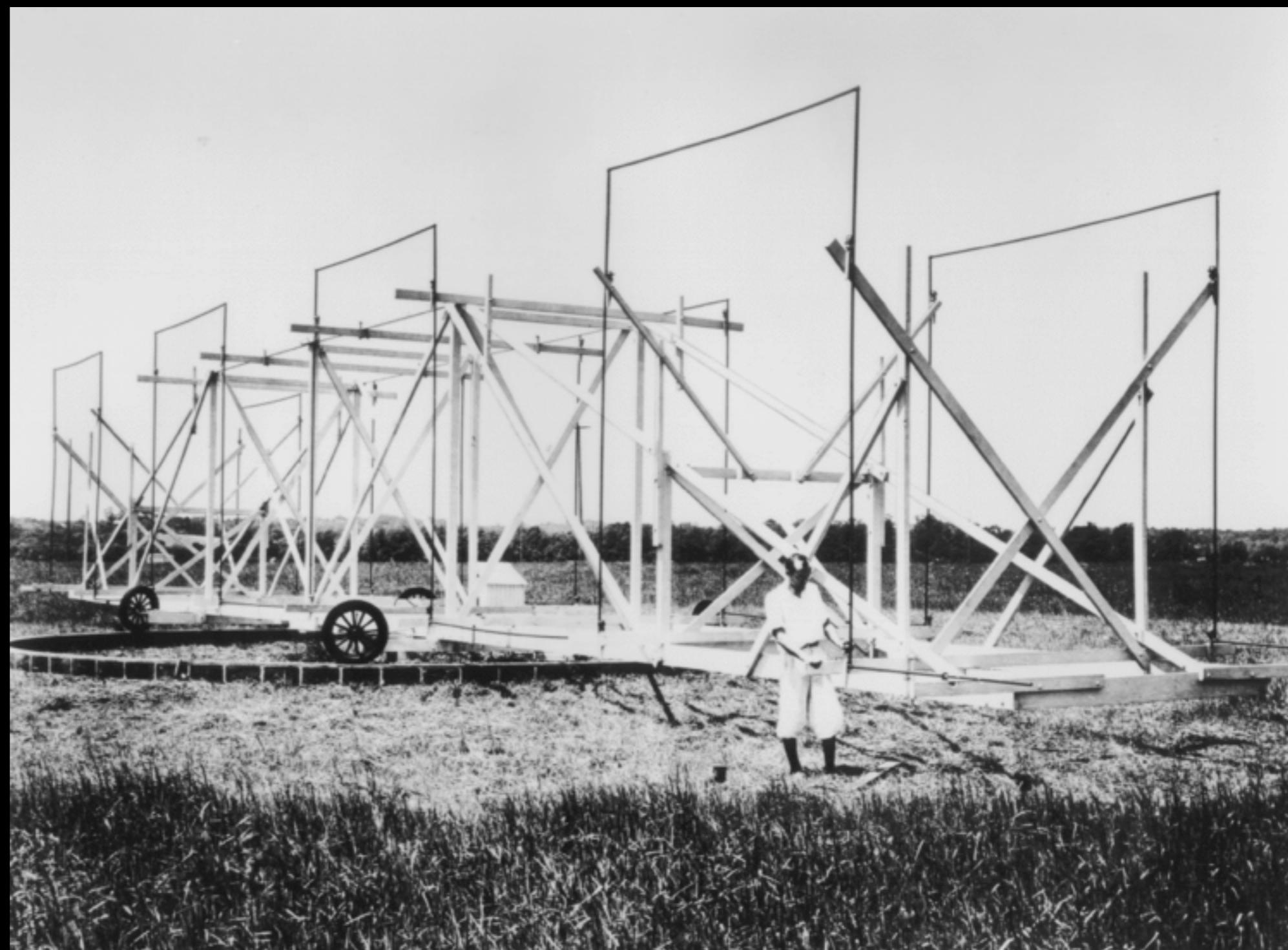
### DIRECTION IS UNCHANGING

Recorded and Tested for More  
Than Year to Identify It as  
From Earth's Galaxy.

### ITS INTENSITY IS LOW

Only Delicate Receiver Is Able to  
Register—No Evidence of  
Interstellar Signaling.

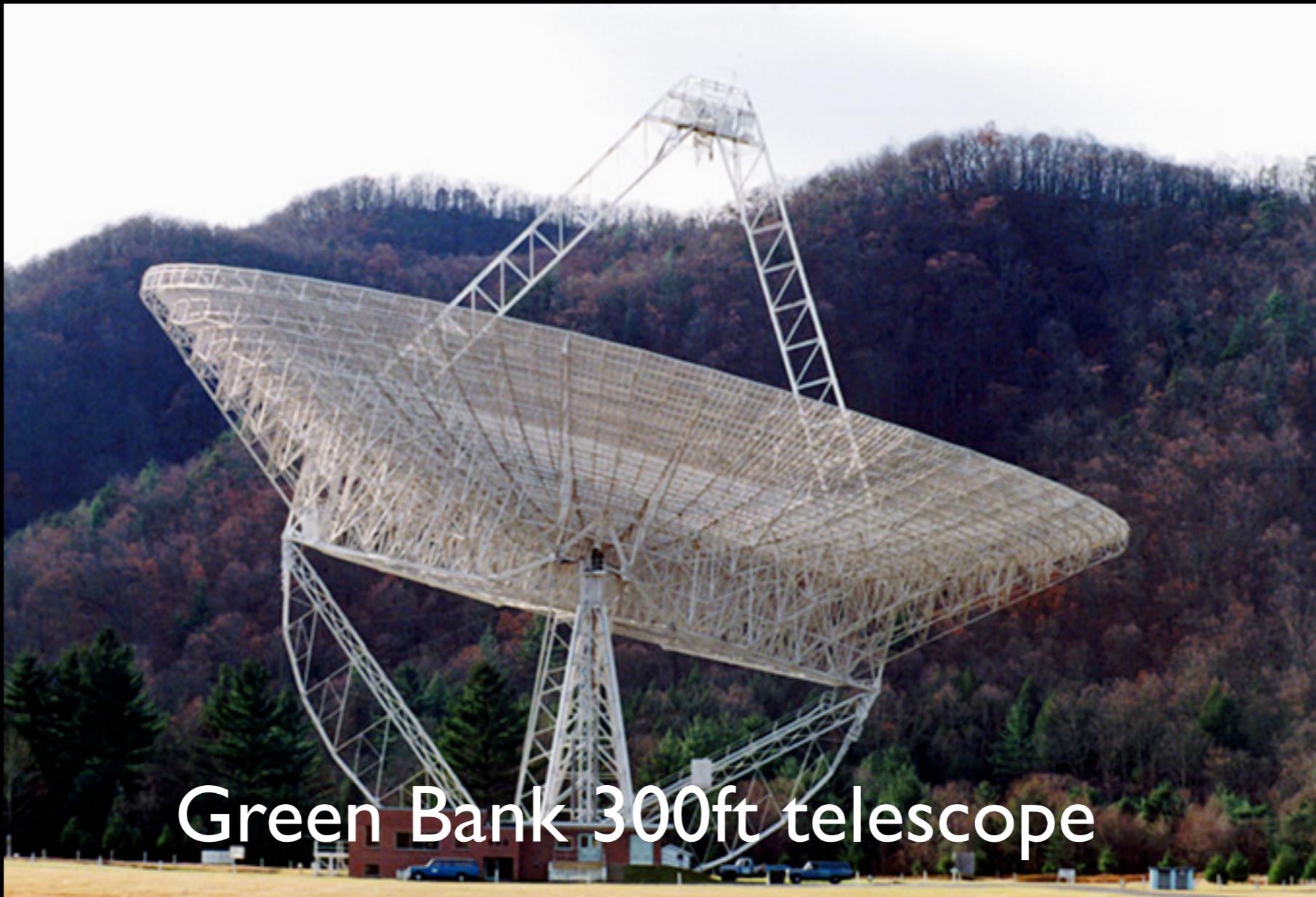
Discovery of mysterious radio  
waves which appear to come from  
the centre of the Milky Way galaxy  
was announced yesterday by the  
Bell Telephone Laboratories. The  
discovery was made during re-  
search studies on static by Karl G.  
Jansky of the radio research de-  
partment at Holmdel, N. J., and  
was described by him in a paper  
delivered before the International  
Scientific Radio Union in Wash-  
ington.



Front page  
New York Times  
May 3, 1933

Jansky with his 20.5 MHz telescope near Bell Labs

# Bigger is better!



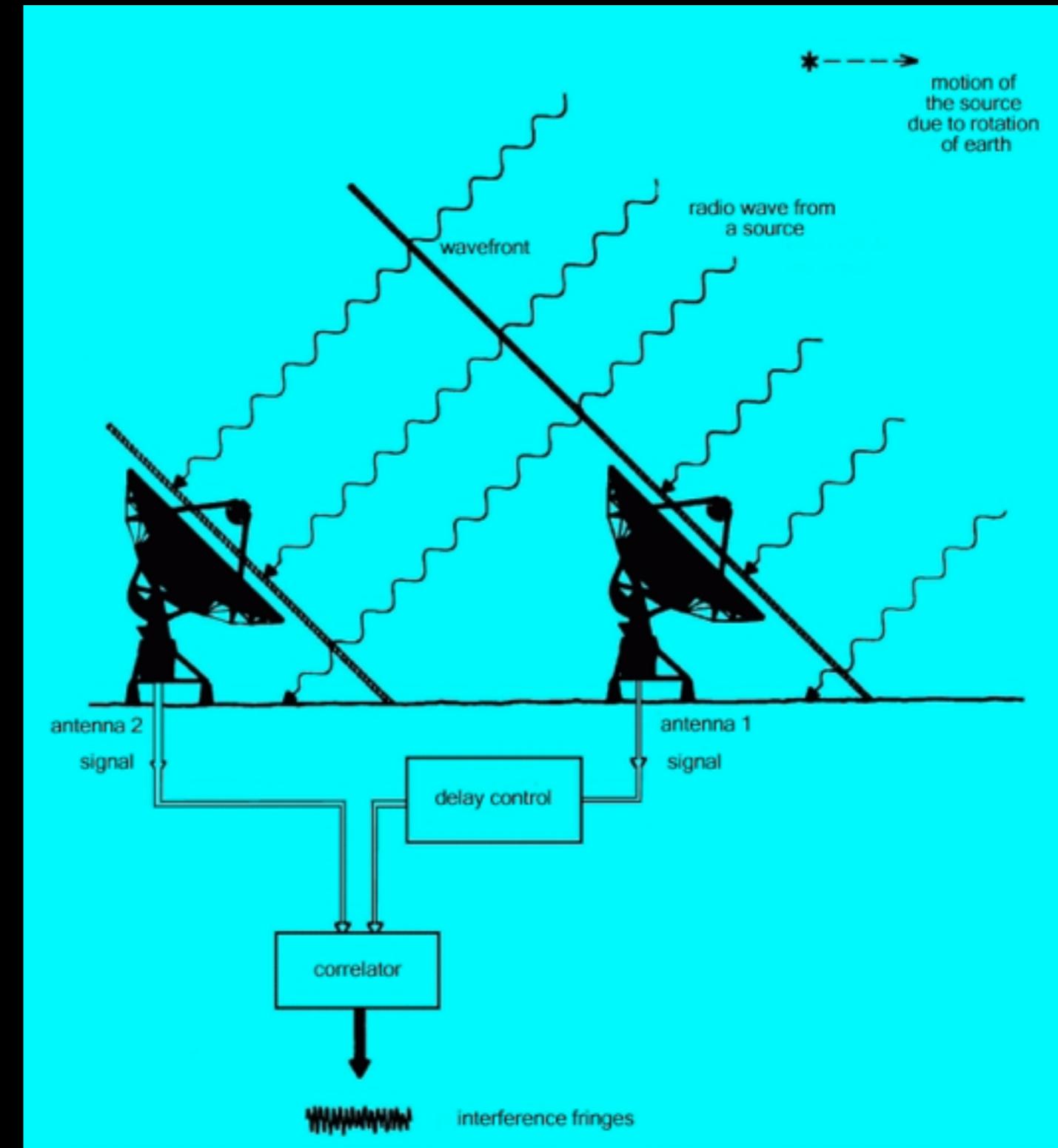
# Building big dishes is hard



collapsed in 1988 - structural failure

# Interferometry

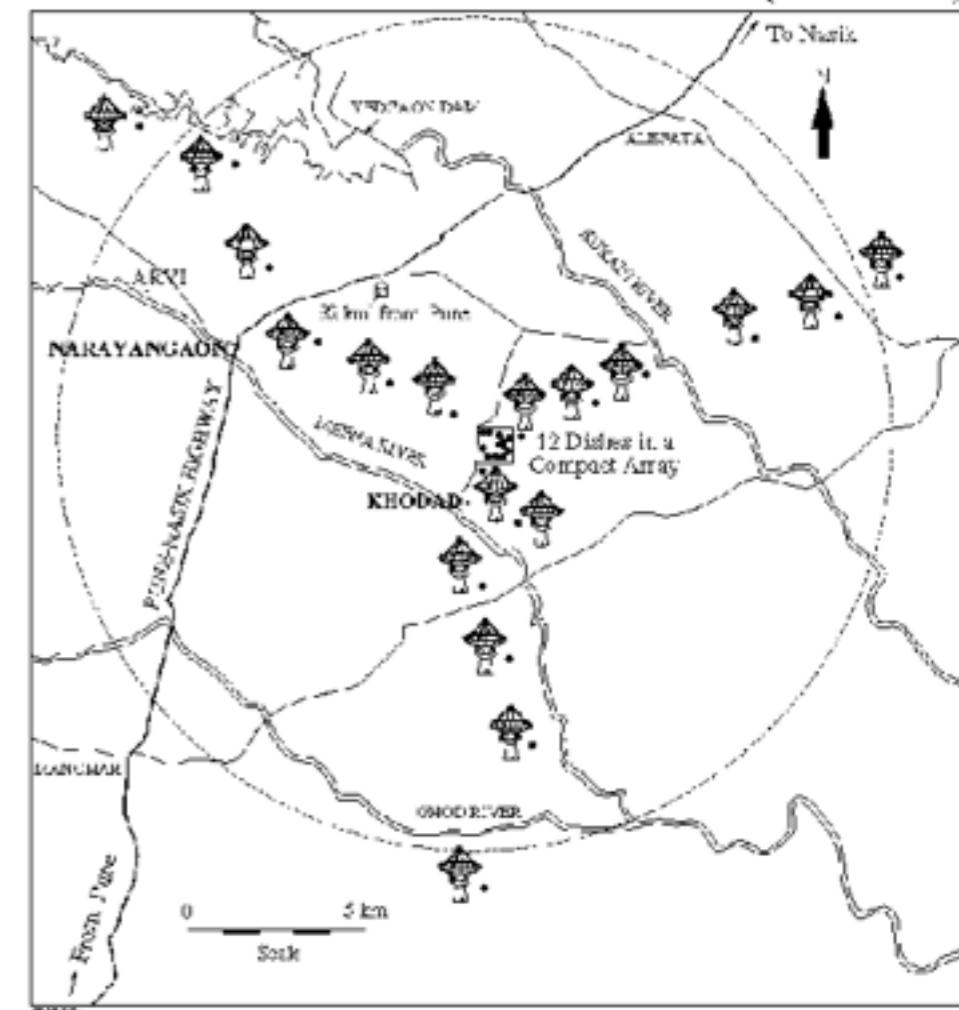
- Radio telescope samples phase and amplitude of electromagnetic wave
- Combine signal from many elements to synthesise a bigger telescope  
= ‘interferometry’
- More elements  
= more collecting area & better quality image



# Giant Meter Radio Telescope



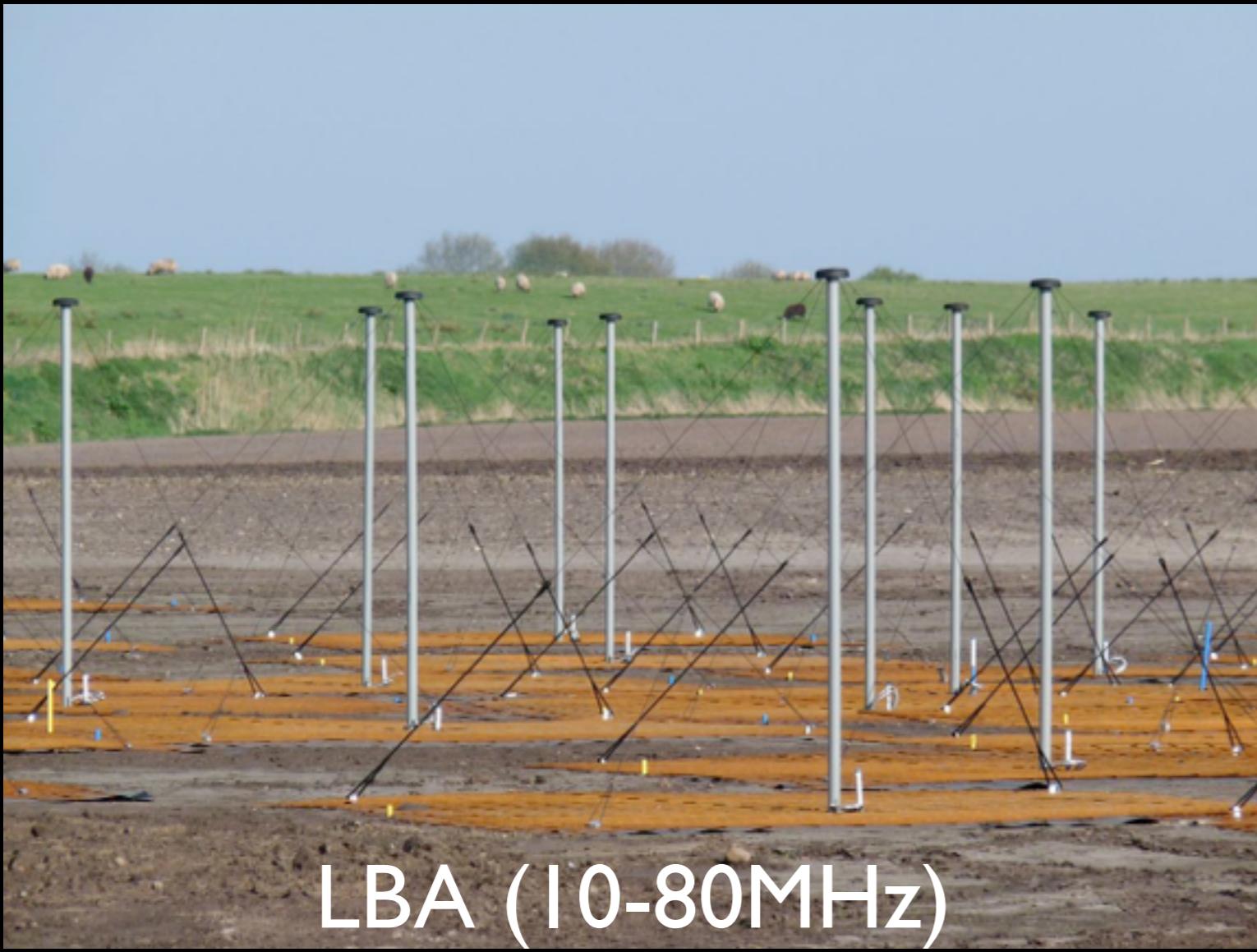
LOCATIONS OF GMRT ANTENNAS ( 30 dishes )



Built in 1995, 80 km from Pune, India  
30 x 45m dishes. 50MHz-1500MHz

# LOw Frequency ARray (LOFAR)

- First of modern arrays built to target cosmic 21cm signal
- Individual elements are simple antennae collected into ‘tiles’
- Plug into supercomputer and correlate = ‘Digital astronomy’



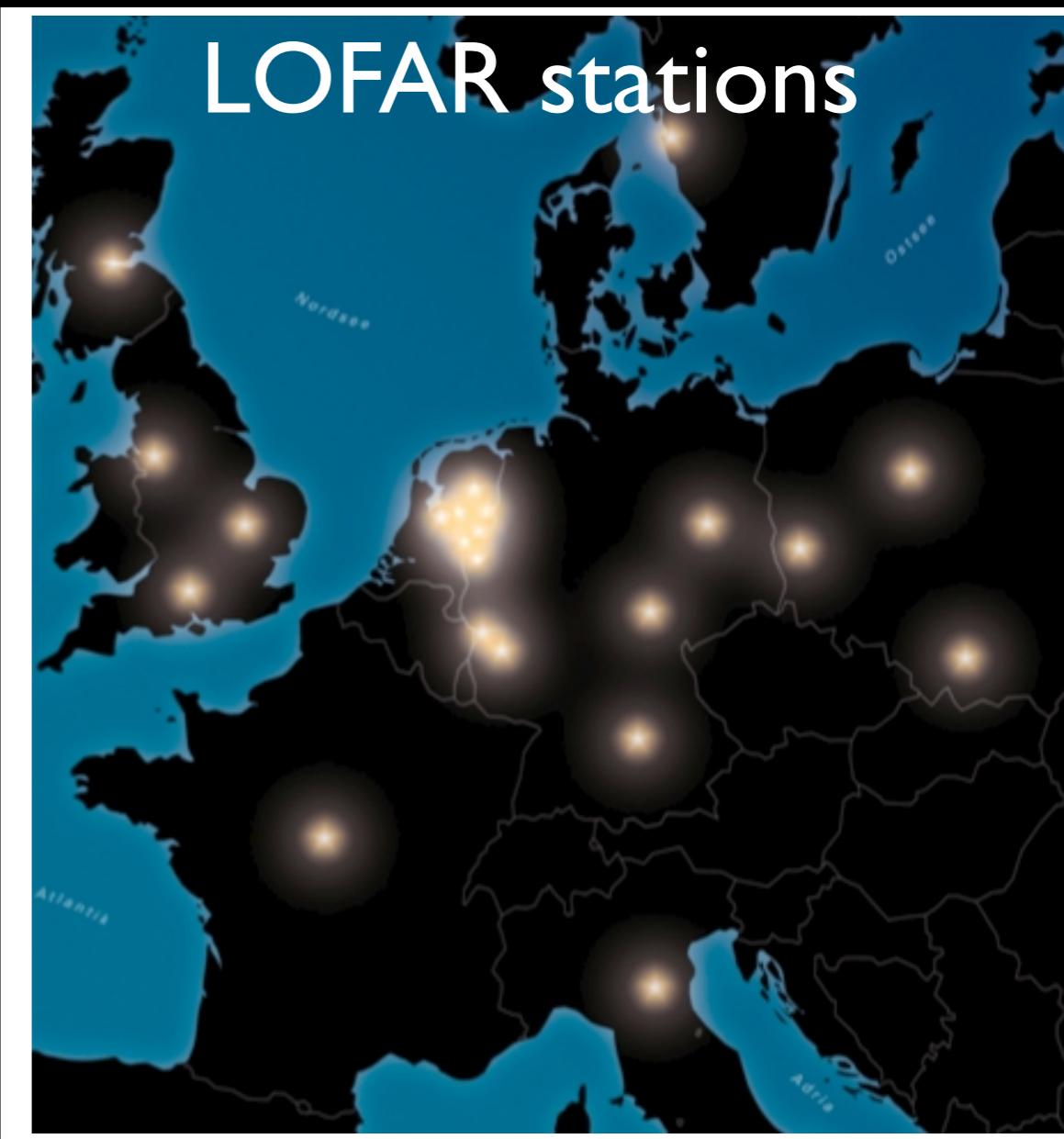
LBA (10-80MHz)

HBA (120-240 MHz)



# LOFAR

- Core located in Netherlands
- Stations scattered over Europe
- Now looking for cosmic 21 cm signal



# Square Kilometer Array

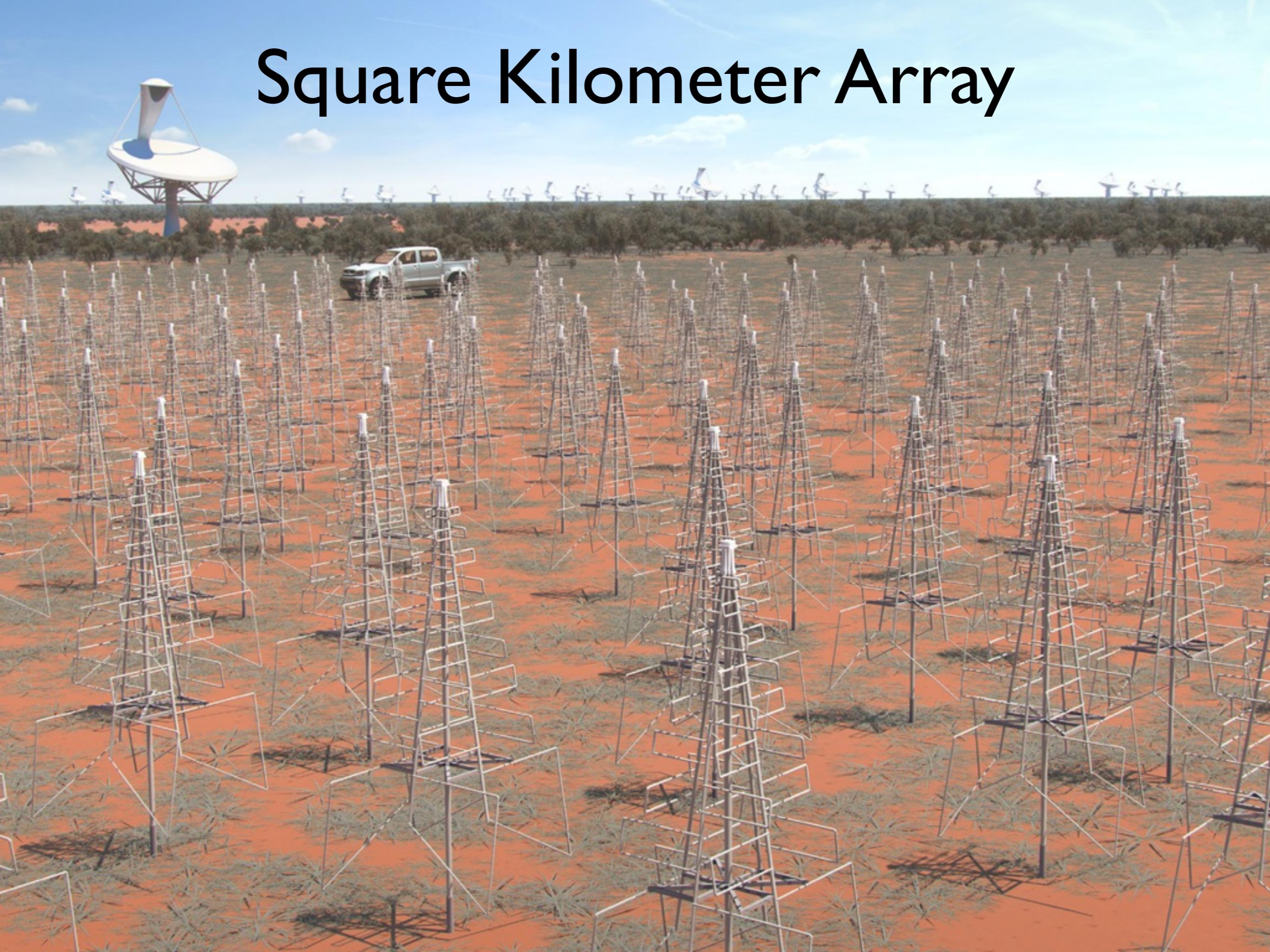
- Will have ~ 1 million antennae and 1 km<sup>2</sup> collecting area
- Distributed over core of ~ 1 km to outlying stations ~ 100km away
- Located in South Africa and Australia
- First light ~2020
- International HQ at Jodrell Bank near Manchester
- Cost: €650 million for Phase I
- BIG DATA!!! - SKA Phase I will produce today's internet per night

# Square Kilometer Array



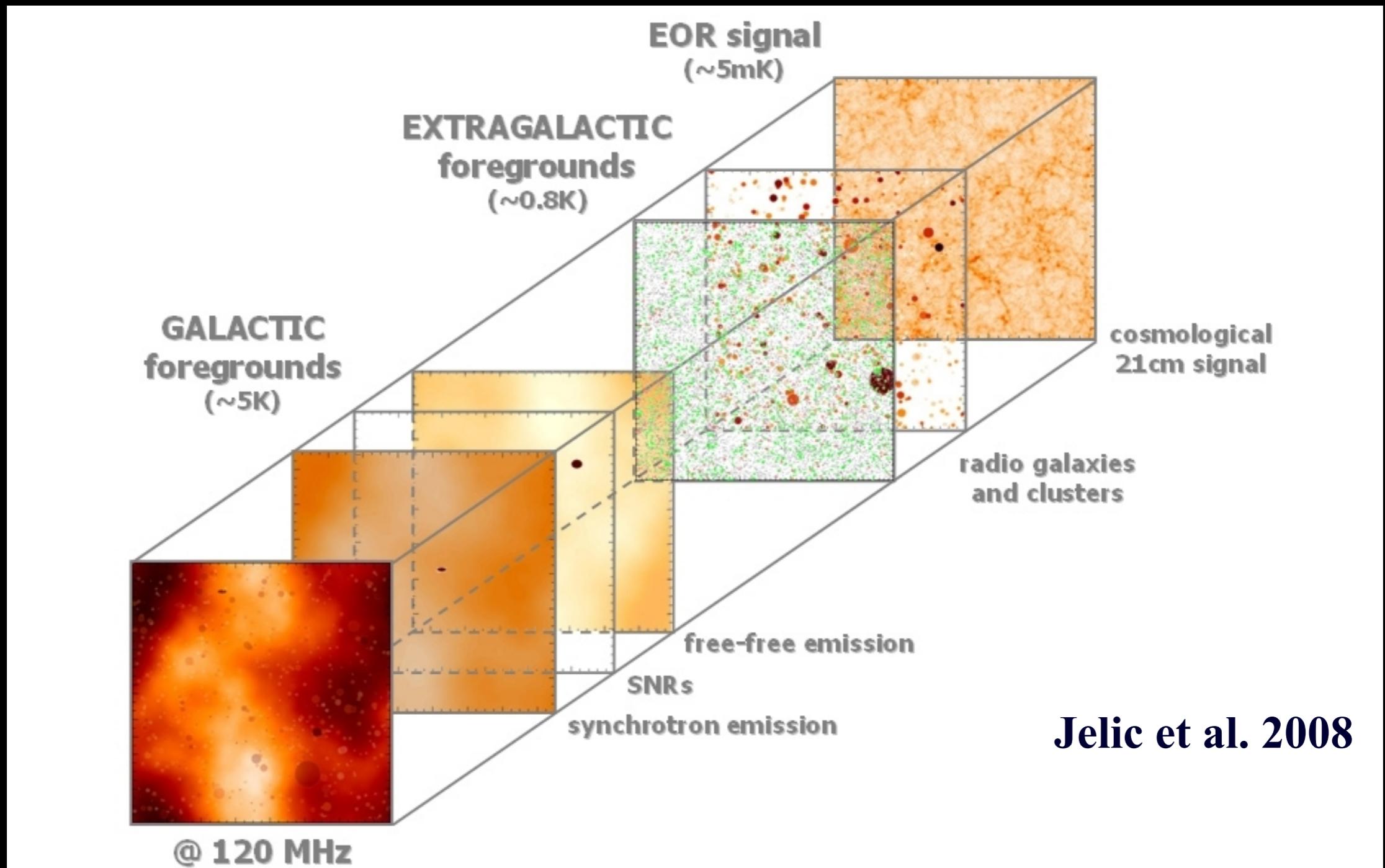
The Square Kilometre Array

# Square Kilometer Array



# Key challenges

- Man made radio interference
- Astronomical foregrounds - 1000x larger than signal



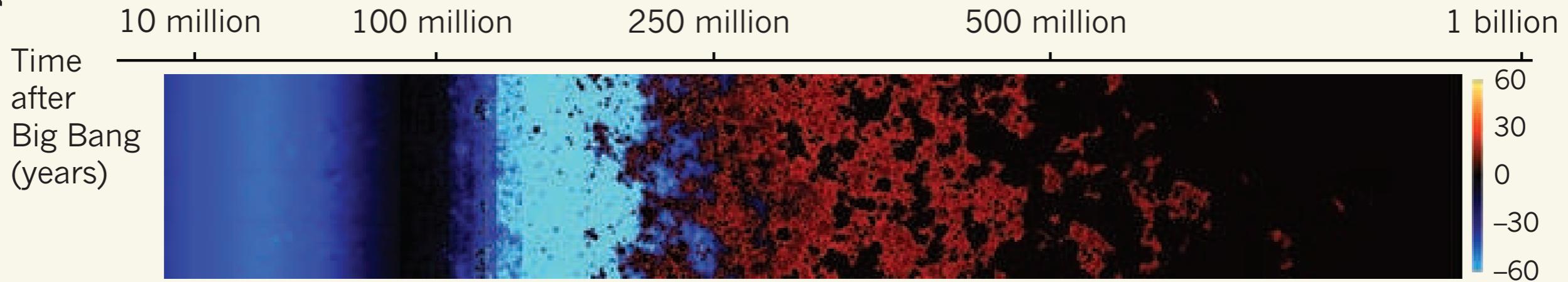
# 21 cm signal

Moon?

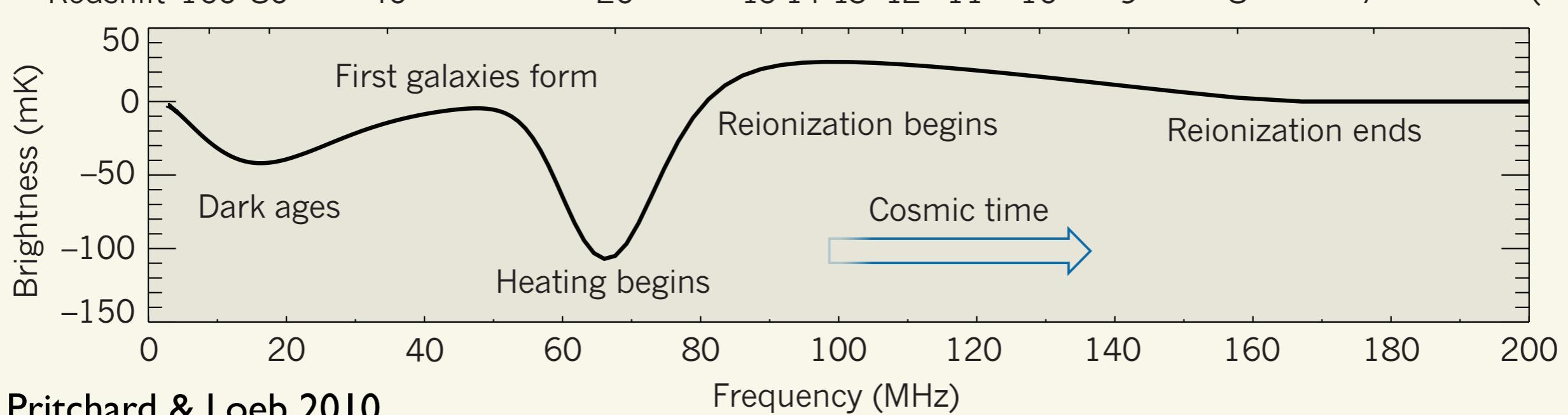
SKA

LOFAR/MWA/PAPER

a



b

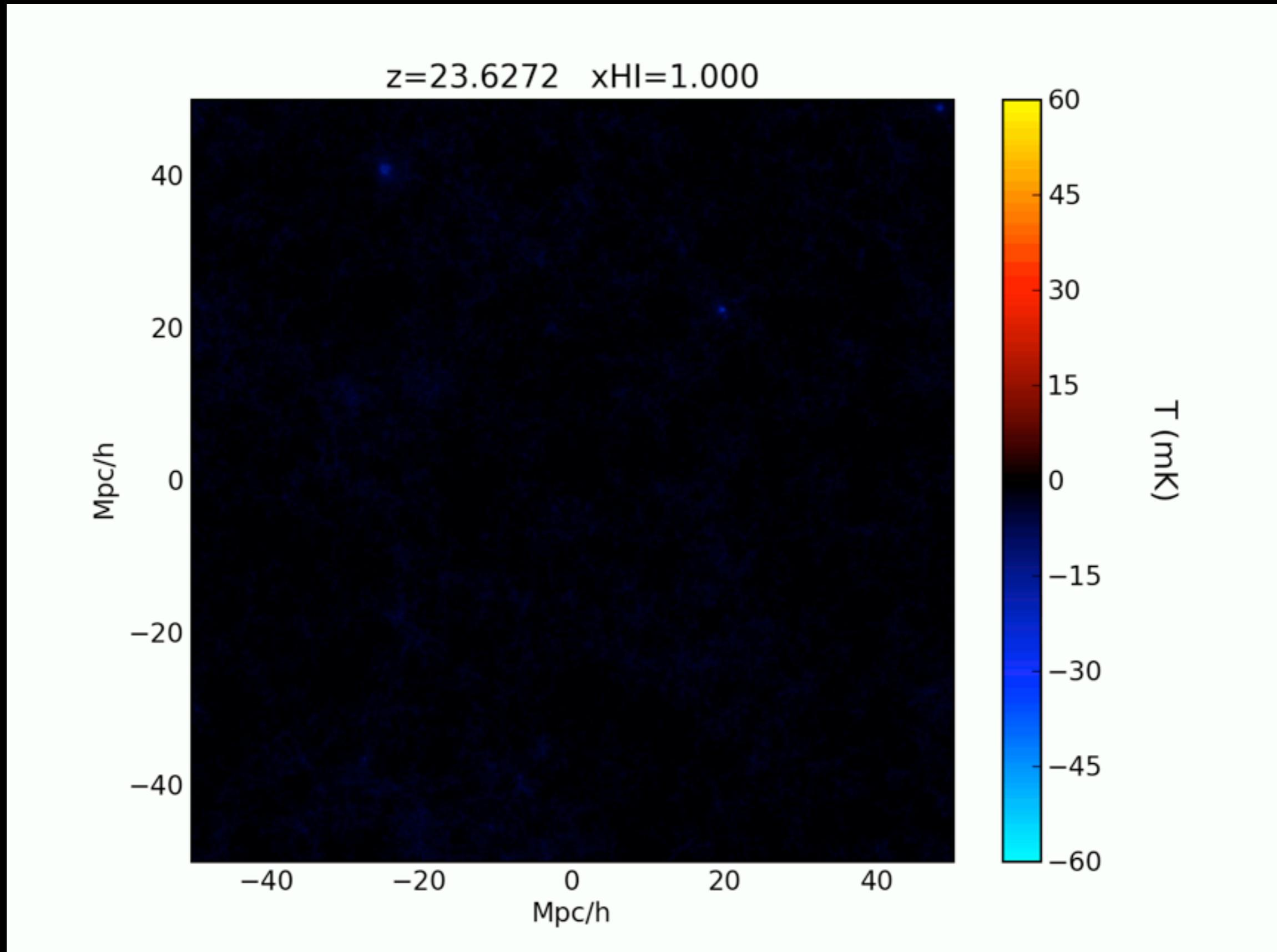


Pritchard & Loeb 2010

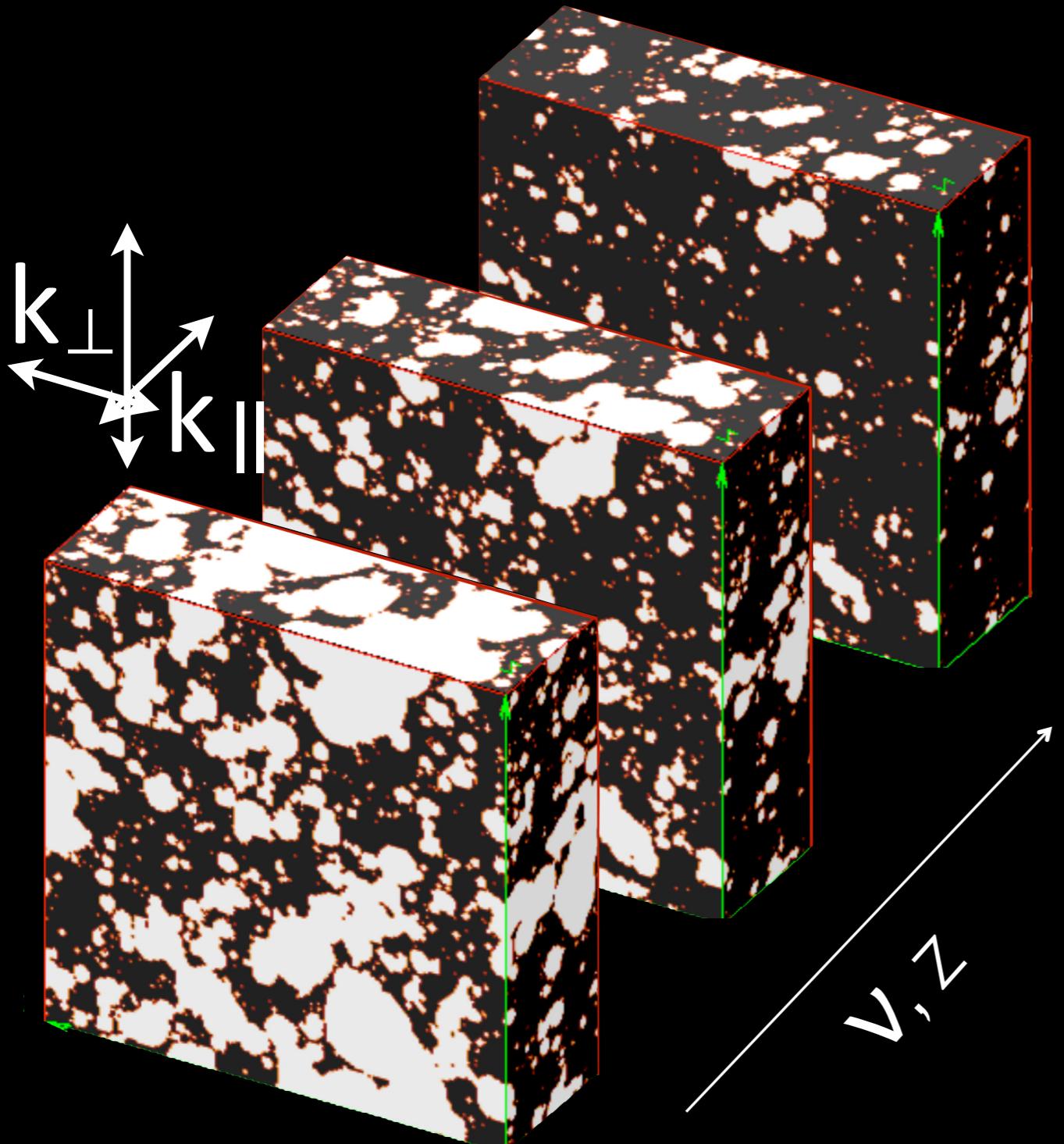
Where CMB is a photo; 21cm signal is a movie

Sensitive to key moments in cosmic history

# Cosmic movies



# What might we learn?



- When did first galaxies begin heating and ionizing the Universe.
- Sizes & number of bubbles linked to brightness and abundance of galaxies
- Temperature of hydrogen linked to X-ray emission from early black holes
- Map distribution of matter throughout Universe

# Things to watch out for

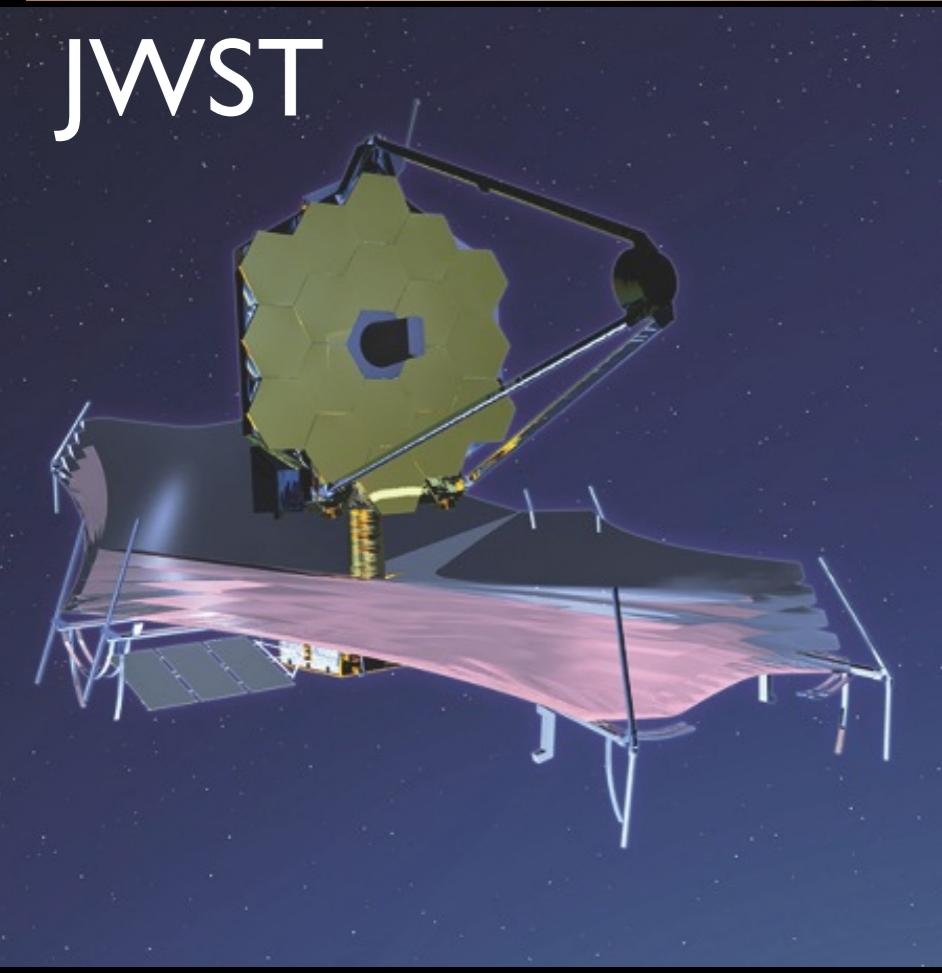
- Radio astronomy being transformed by computing power
- New instruments; new capabilities; new frontiers
- Big data; big computing challenges  
=> new algorithms needed
- 21 cm studies will reveal Universe from 100 million years after the Big Bang through to today

# Exciting times ahead!

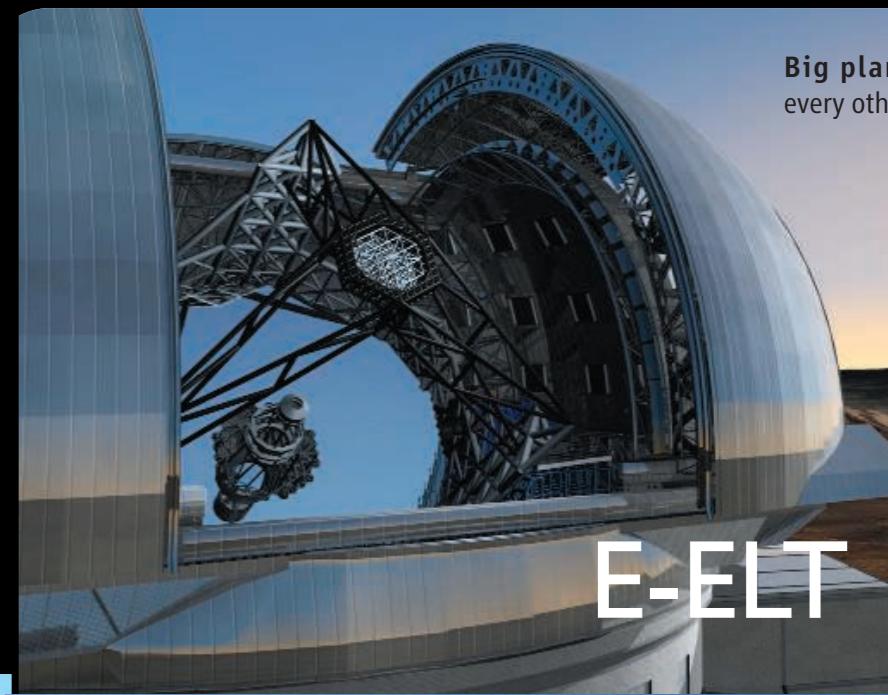
ALMA



JWST



TMT



GMT



Big plan  
every othe

## We are listening - Diane Ackerman

As our metal eyes wake  
to absolute night,  
where whispers fly  
from the beginning of time,  
we cup our ears to the heavens.

We are listening

on the volcanic rim of Flagstaff  
and in the fields beyond Boston,  
in a great array that blooms  
like coral from the desert floor,  
on highwire webs patrolled  
by computer spiders in Puerto Rico.

We are listening for a sound  
beyond us, beyond sound,

searching for a lighthouse  
in the breakwaters of our uncertainty,  
an electronic murmur,  
a bright, fragile *I am*.

Small as tree frogs  
staking out one end  
of an endless swamp,  
we are listening  
through the longest night  
we imagine, which dawns  
between the life and times of stars.

Our voice trembles  
with its own electric,  
we who mood like iguanas,  
we who breathe sleep  
for a third of our lives,  
we who heat food  
to the steaminess of fresh prey,  
then feast with such  
good manners it grows cold.

In mind gardens  
and on real verandas  
we are listening,  
rapt among the Persian lilacs  
and the crickets,  
while radio telescopes  
roll their heads, as if in anguish.

With our scurrying minds  
and our lidless will  
and our lank, floppy bodies  
and our galloping yens  
and our deep, cosmic loneliness  
and our starboard hearts  
where love careens,  
we are listening,  
the small bipeds  
with the giant dreams.



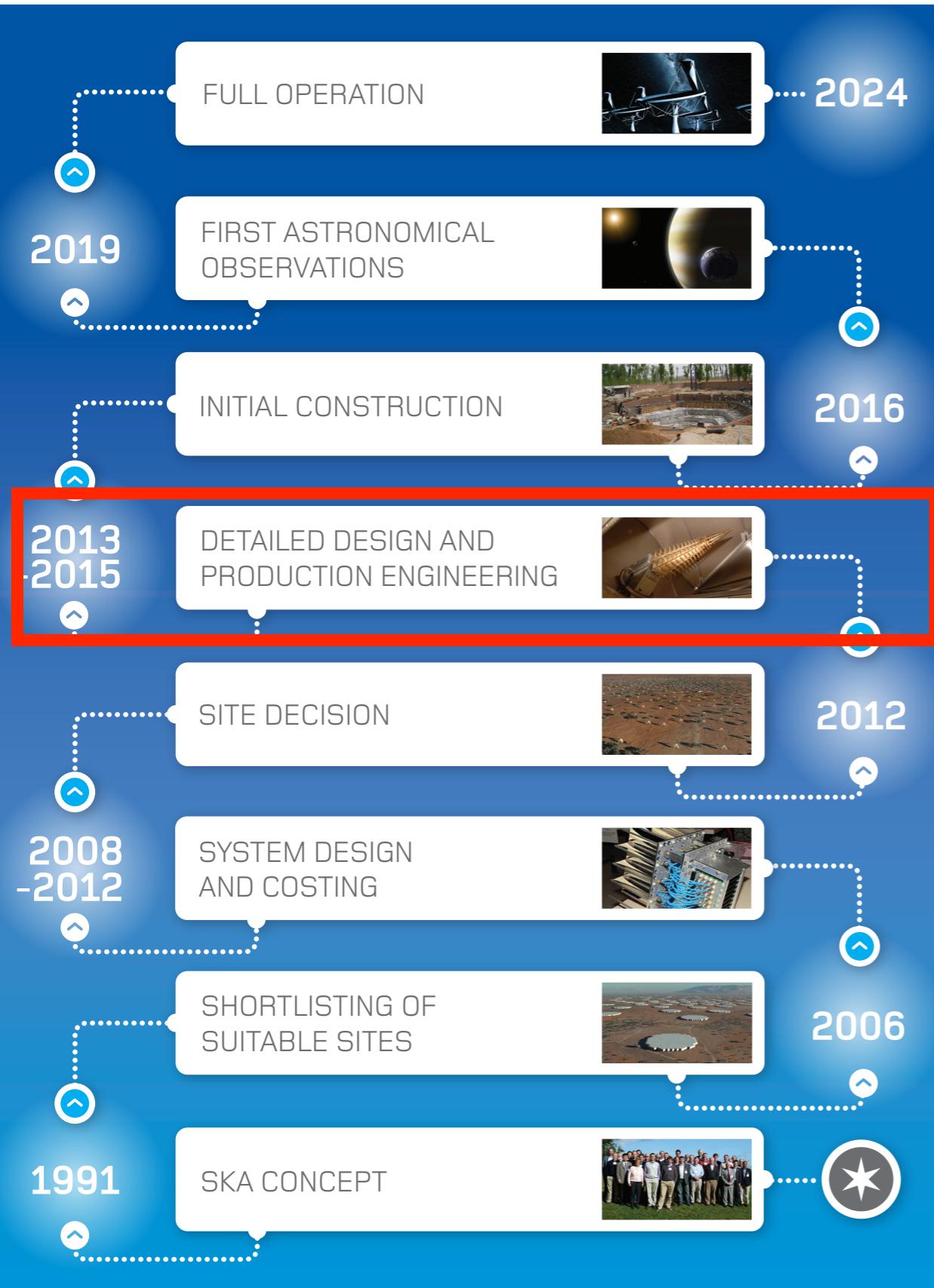
In the beginning of the Dark Ages, electrically neutral hydrogen gas filled the universe. As stars formed, they ionized the regions immediately around them, creating bubbles here and there. Eventually these bubbles merged together, and intergalactic gas became entirely ionized.

# Thank you for your consideration





# SKA status



SKA director general - Phil Diamond

SKA sites: South Africa - Karoo  
Australia - Western Outback

SKA-low: 50-350 MHz (Australia)  
SKA-mid: 350MHz-3GHz (SA)

Baseline design currently under discussion

## Current



SKA project office - Jodrell Bank, Manchester





# MWA



# Arecibo



305m dish

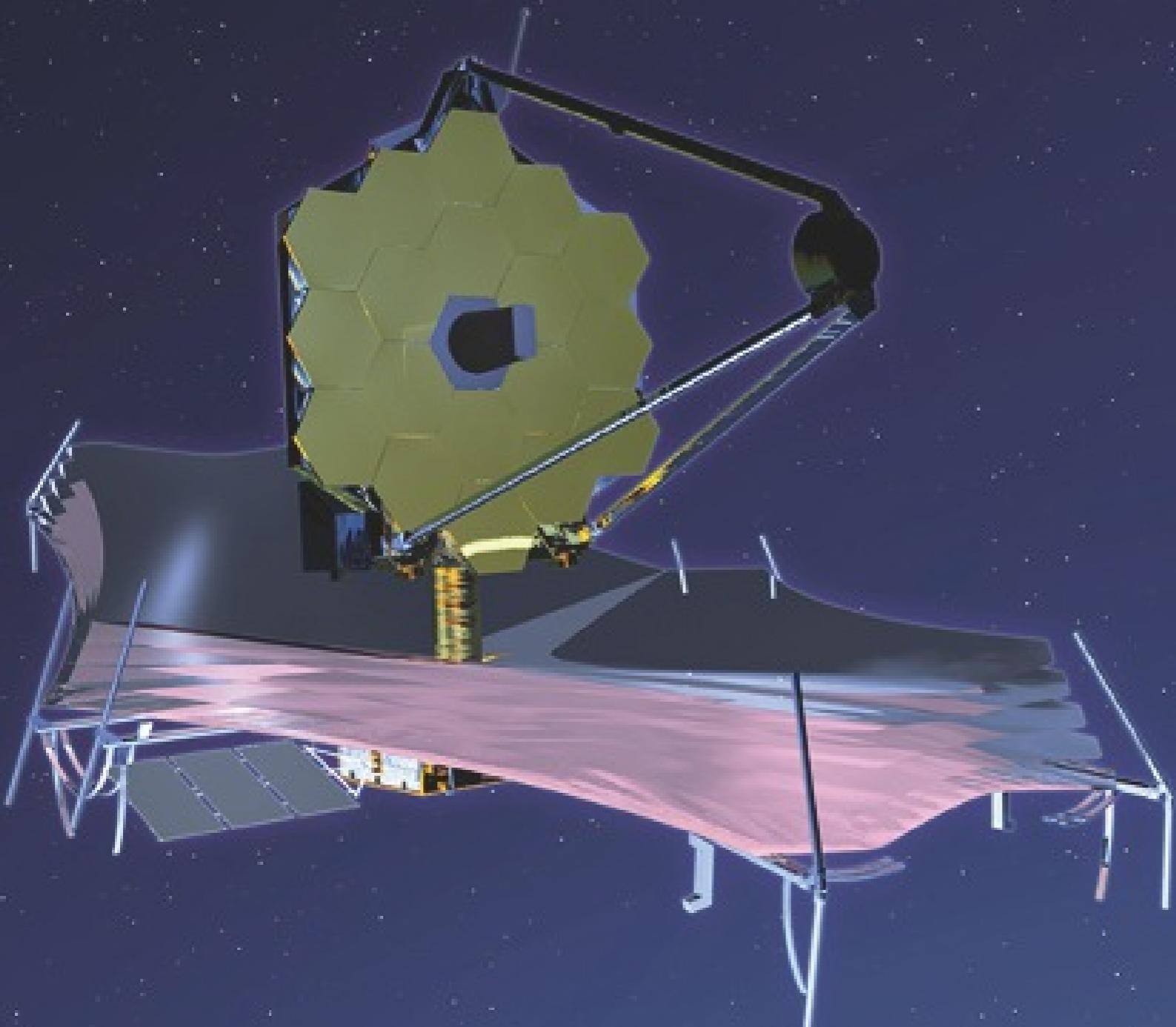


# VLA





# JWST





# ALMA





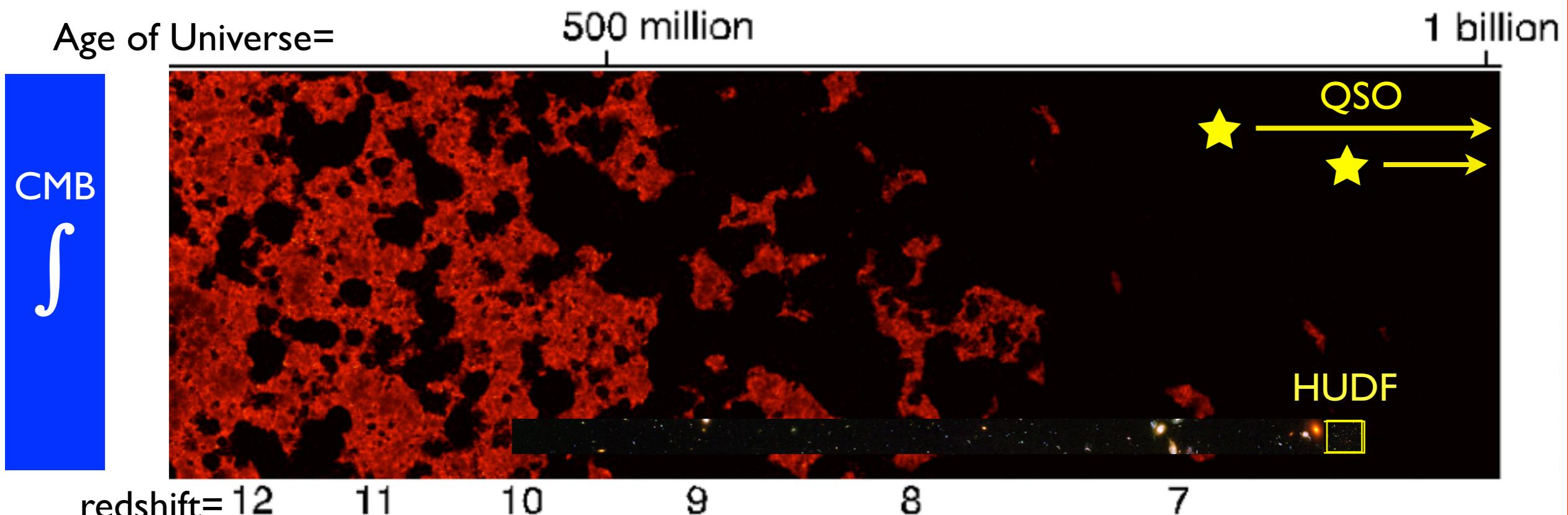
# DARE

Imperial College  
London





# More needed...



Existing observations leaves much unanswered:

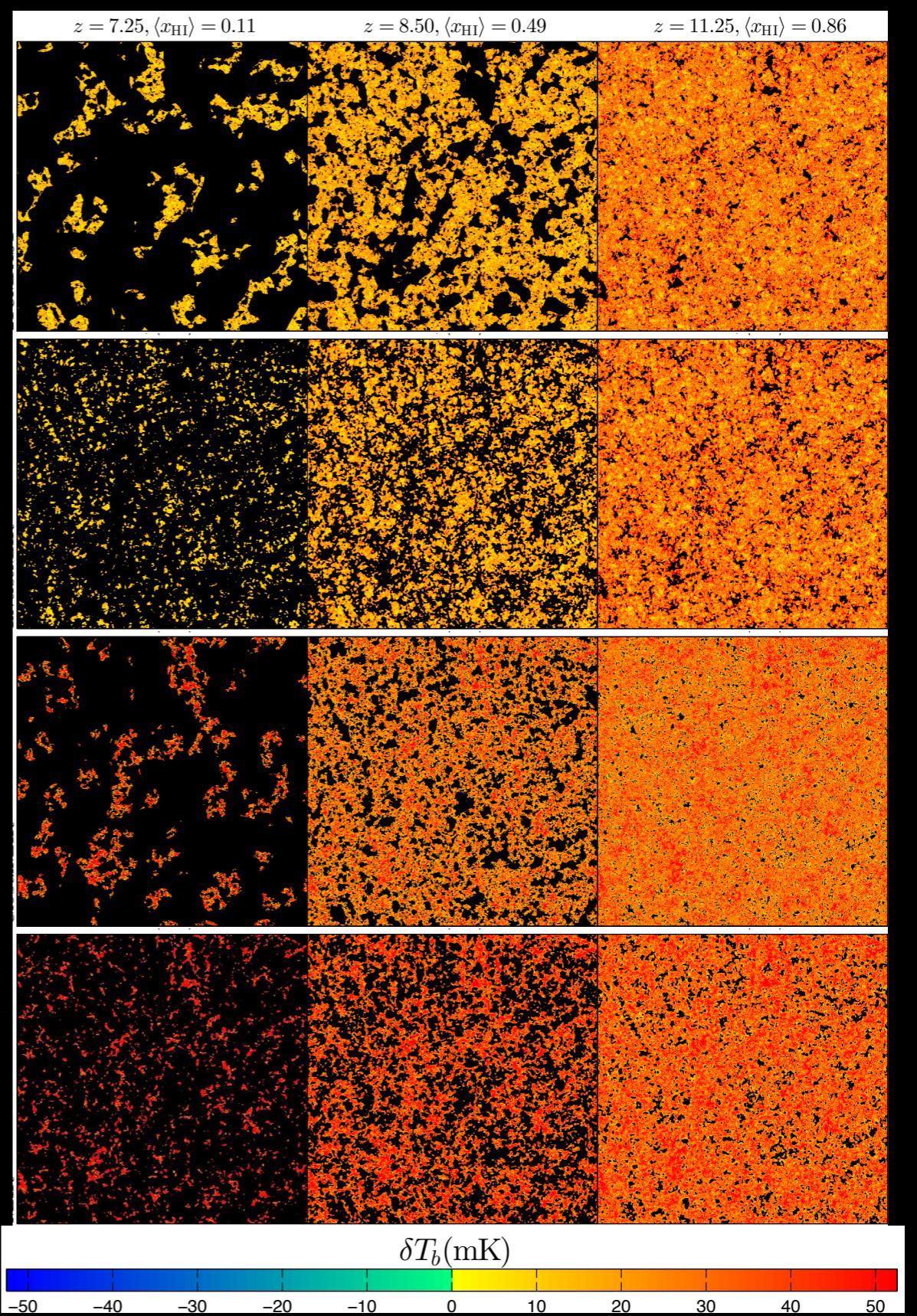
- 1) Lyman-alpha forest: end point  $z > 6.5$
- 2) CMB optical depth: mid point  $z \sim 11$
- 3) kSZ amplitude: duration  $z < 4.4$  ?

HST probes skewer much smaller than scale of ionized regions + only brightest sources

Large galaxy samples with LAE surveys or Euclid possible to  $z \sim 8$

Fundamental need for new types of observation to understand details of reionization

# Simulation of galaxy formation



We could imagine that light moved more slowly  
say 1km/year

Then 18000km to Wellington,NZ would take 18000years  
we'd see back to the beginning of human history

and you'd have to wait ~18 hours after hitting  
a lightswitch to see the light

London  
0km



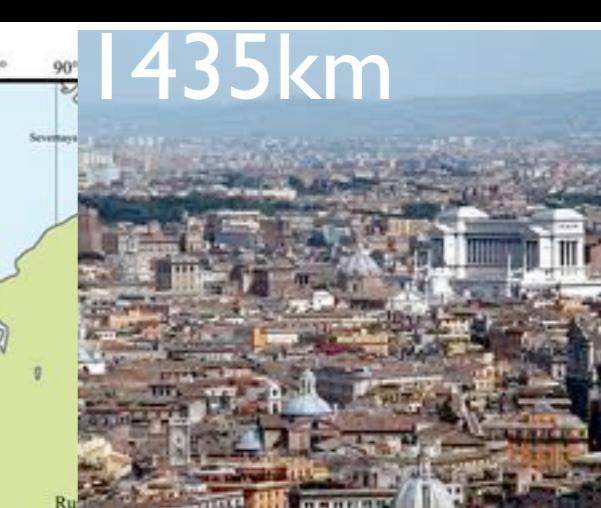
Paris  
344km



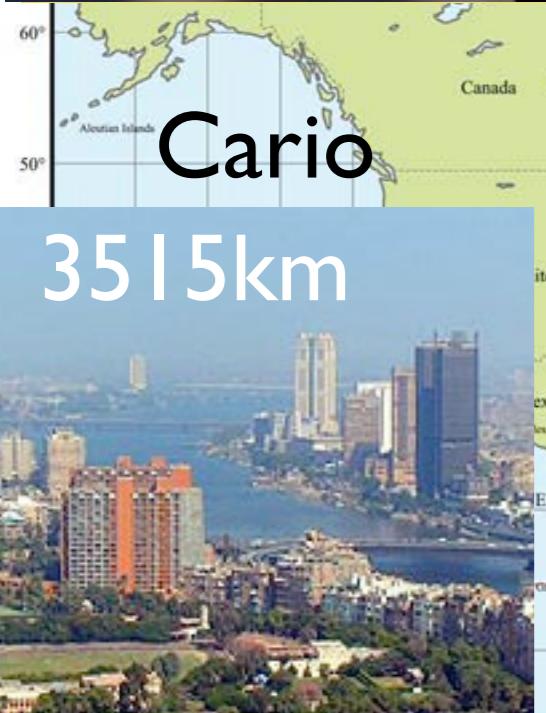
Berlin  
932km



Rome  
1435km



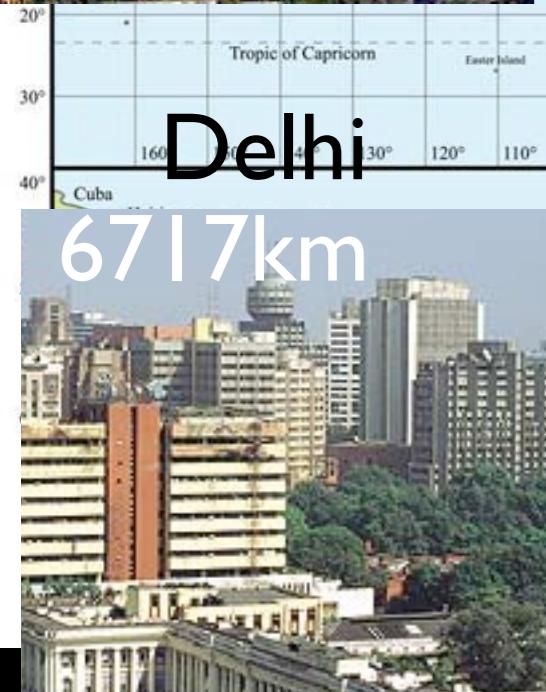
Cario  
3515km



Baghdad  
4097km



Delhi  
6717km



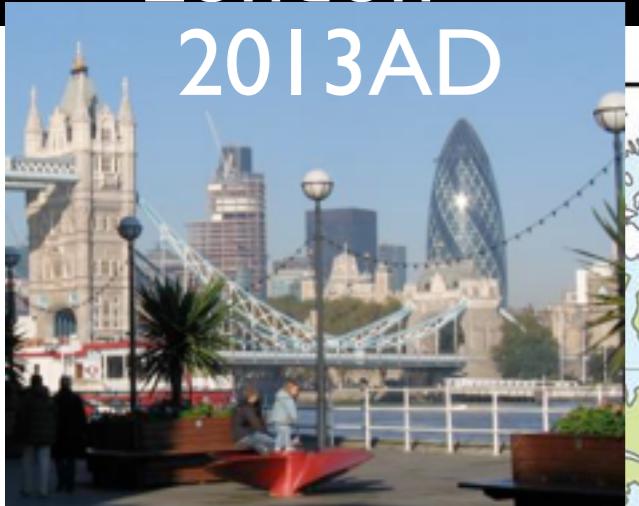
Sydney  
17012km



Wellington  
18834km



London  
2013AD



Paris

1669AD



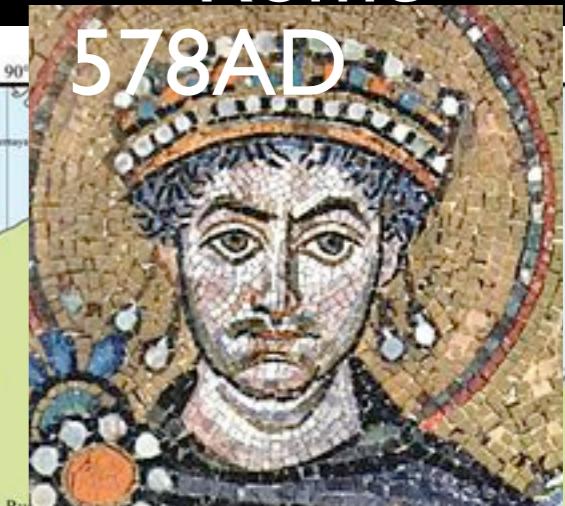
Berlin

1081AD



Rome

578AD



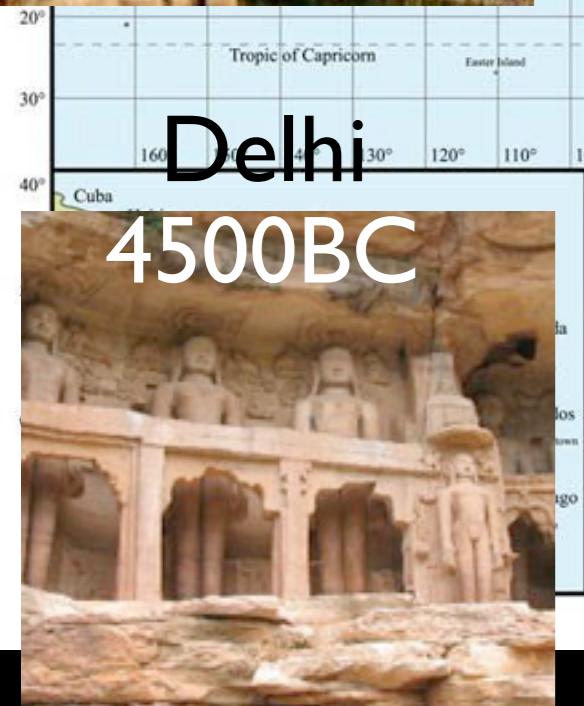
Cario



Baghdad



1502BC

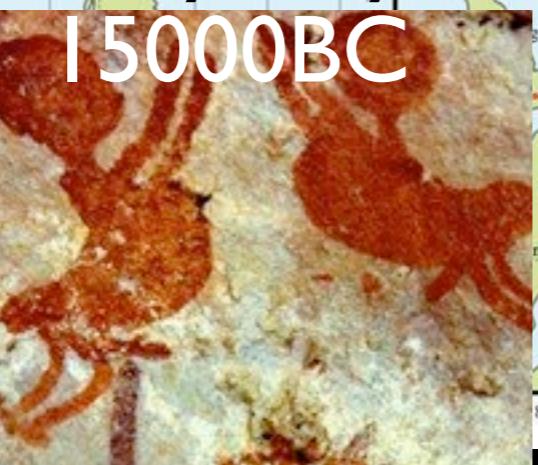


Delhi

4500BC

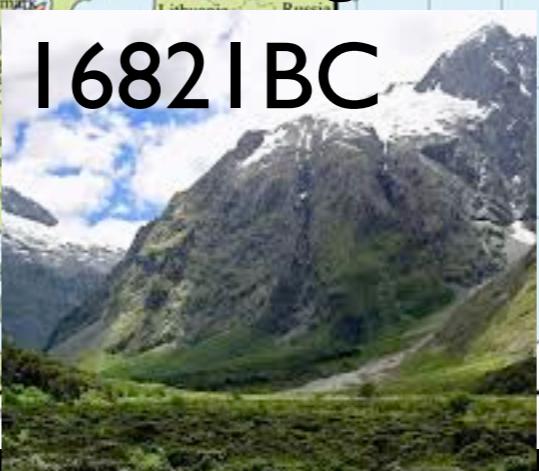
Sydney

15000BC



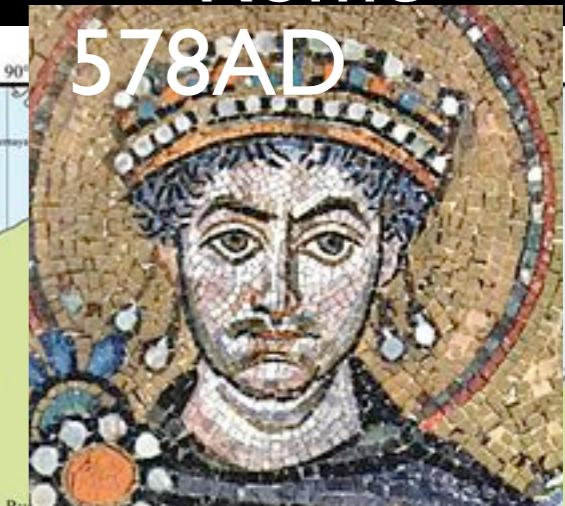
Wellington

1682BC



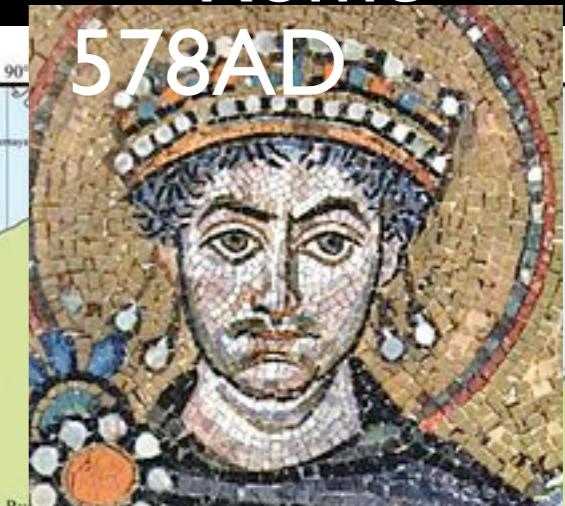
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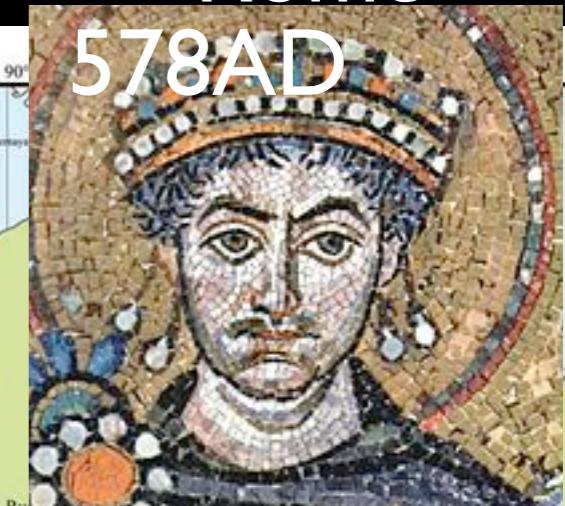
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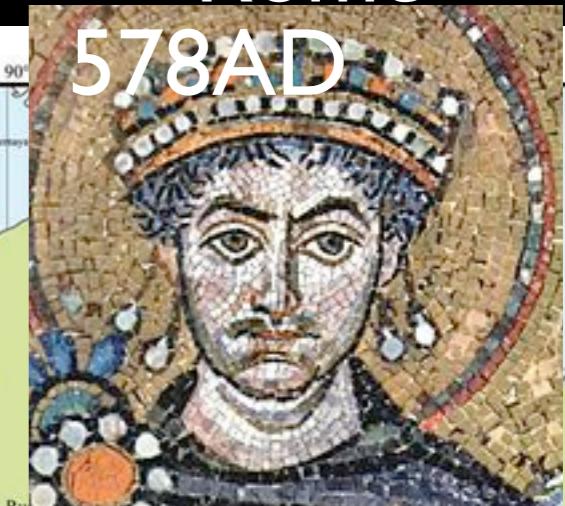
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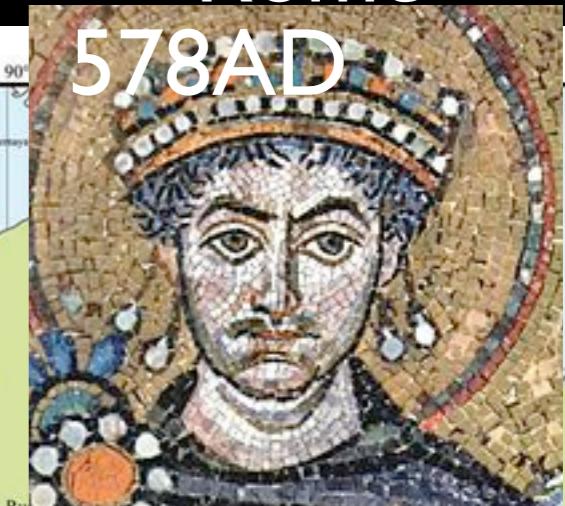
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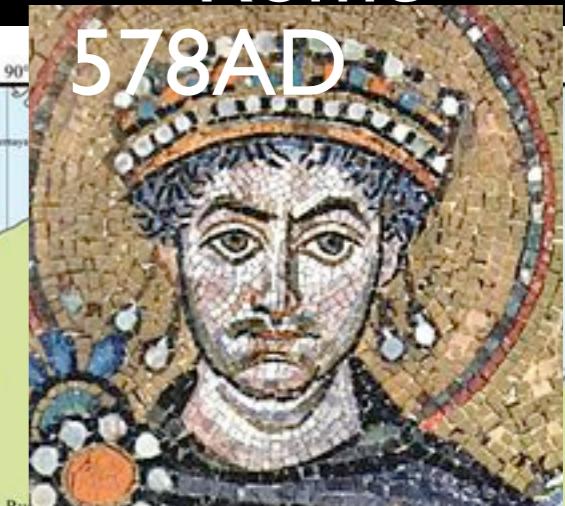
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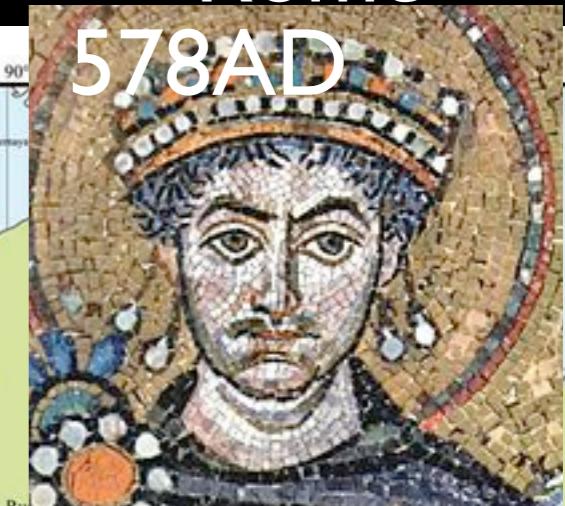
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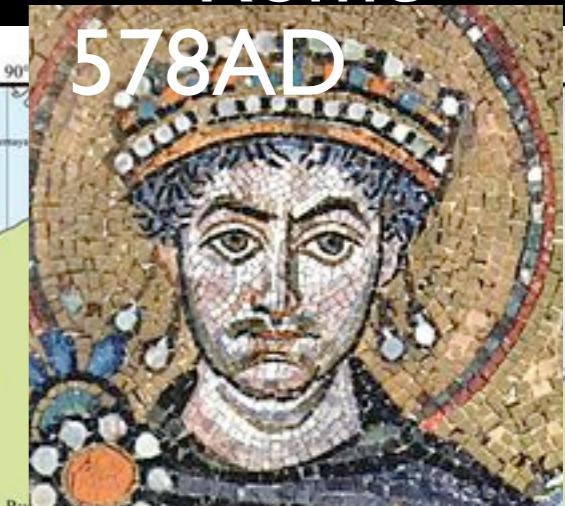
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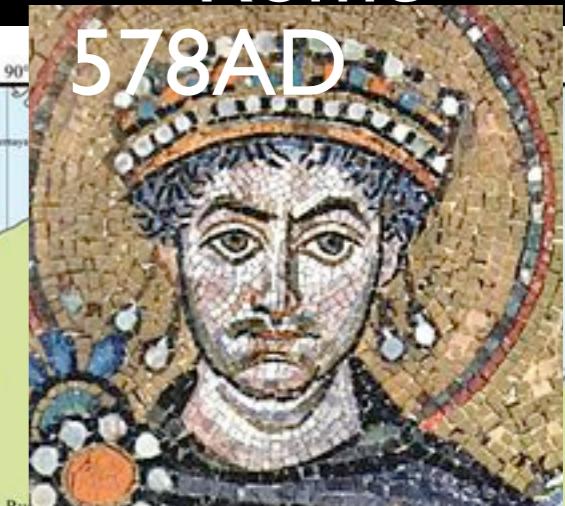
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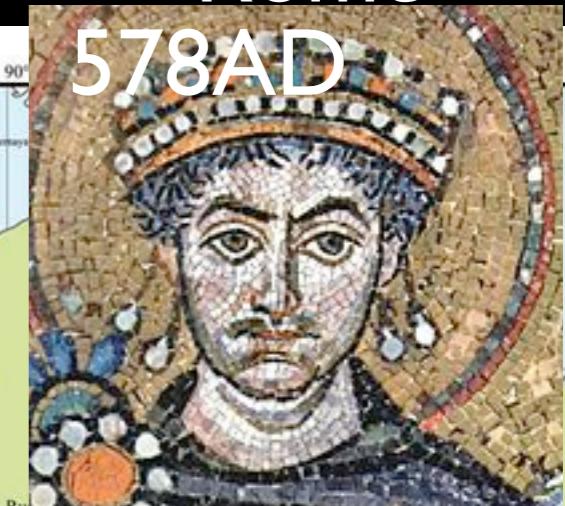
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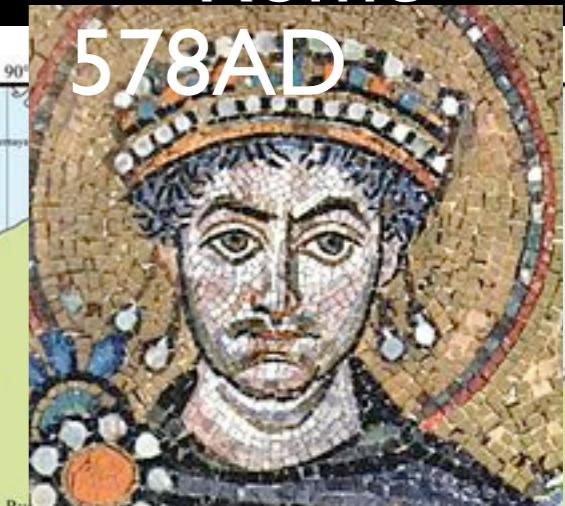
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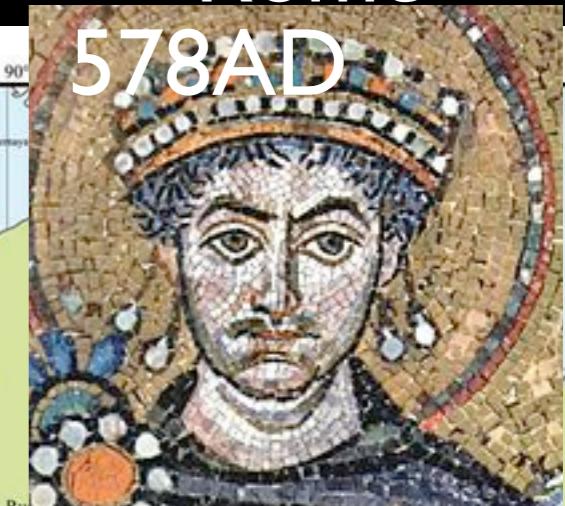
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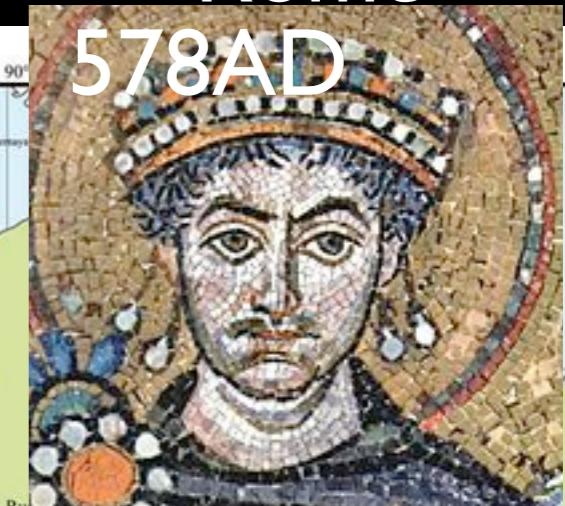
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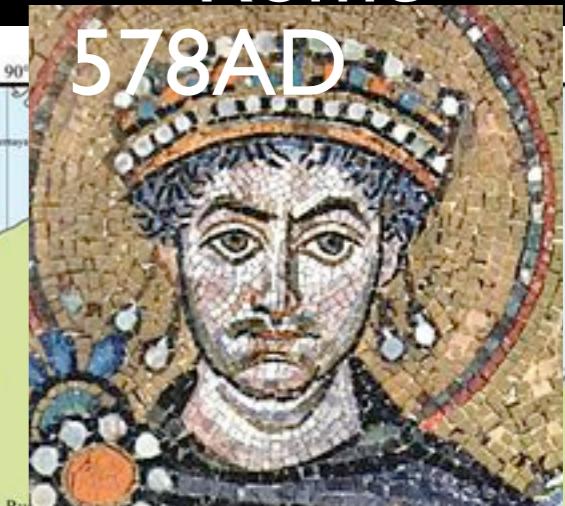
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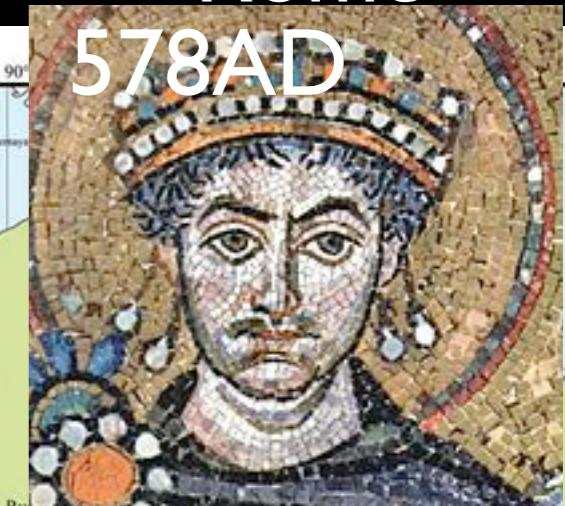
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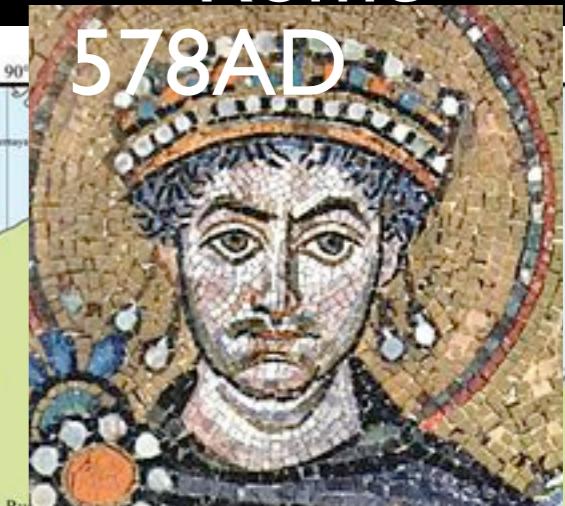
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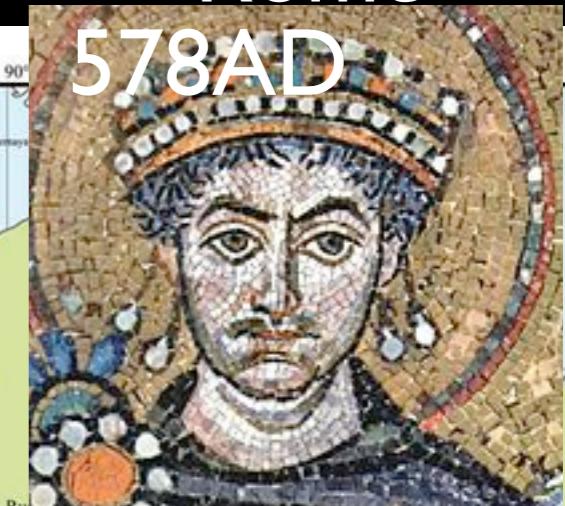
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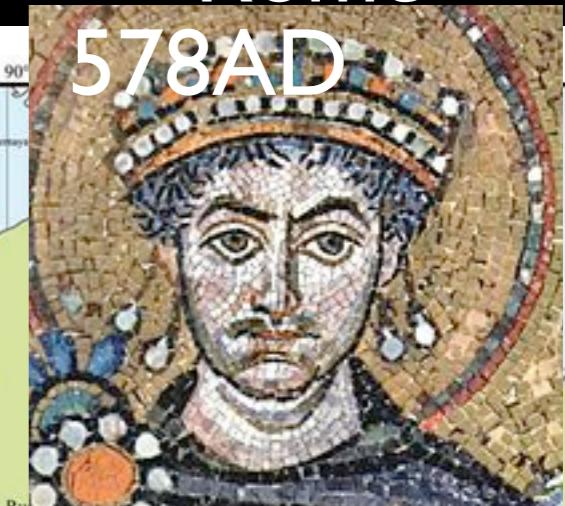
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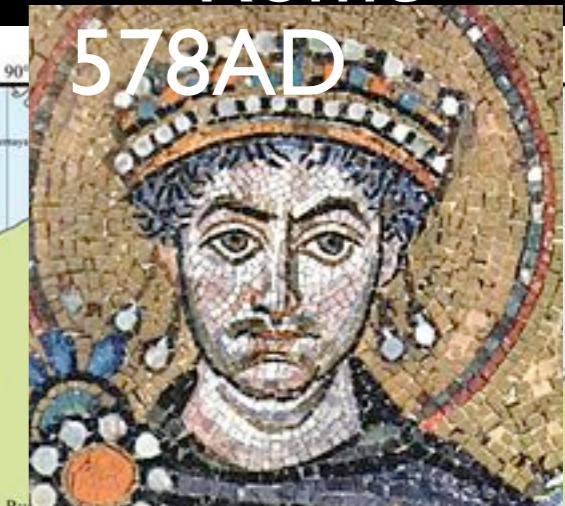
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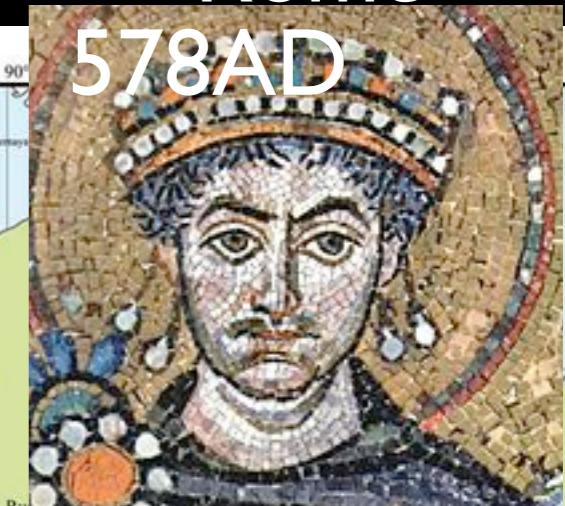
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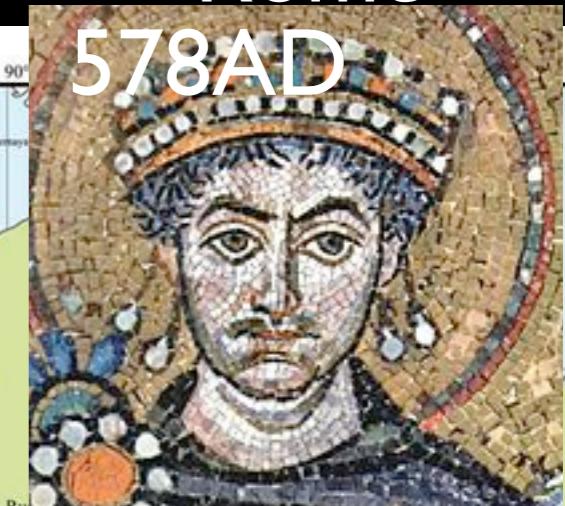
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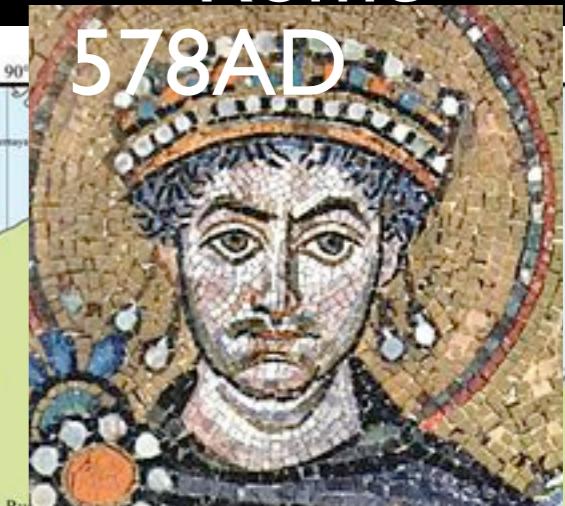
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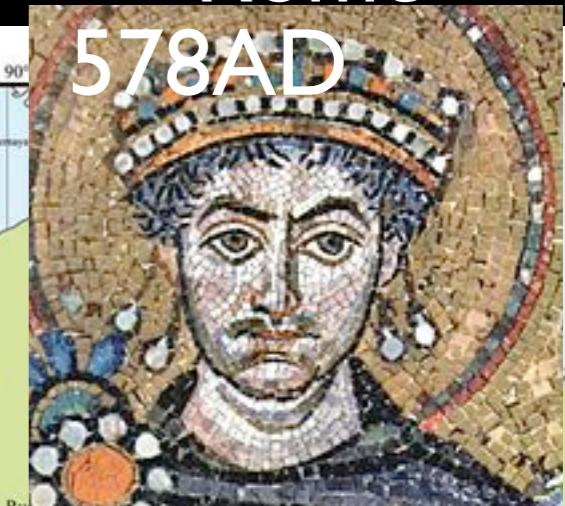
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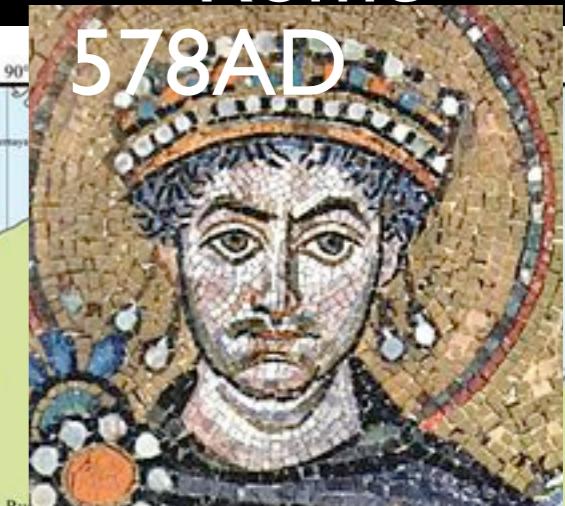
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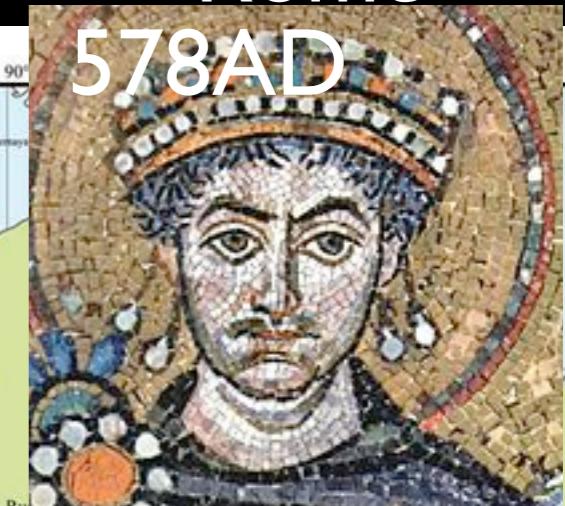
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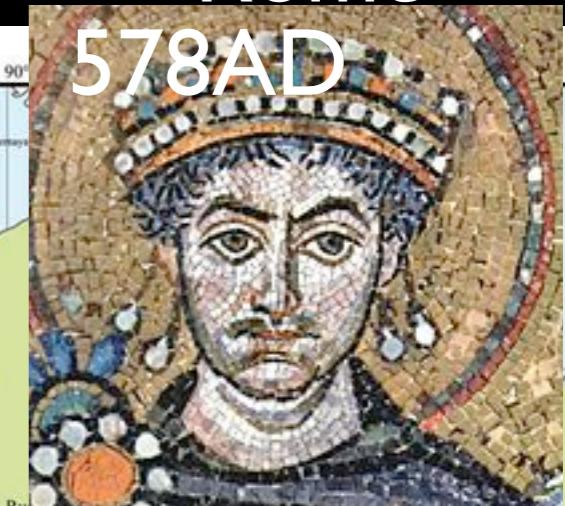
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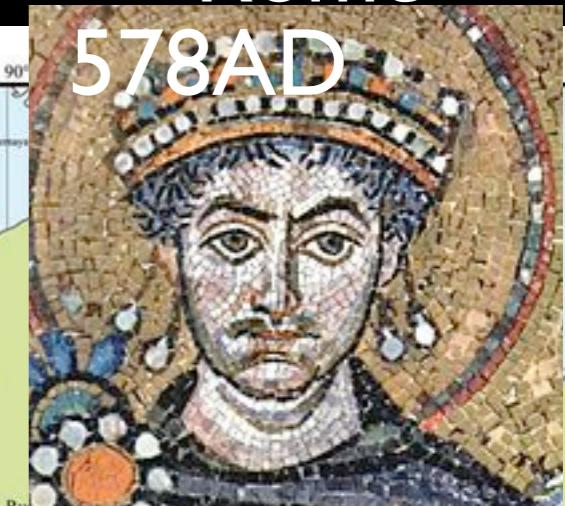
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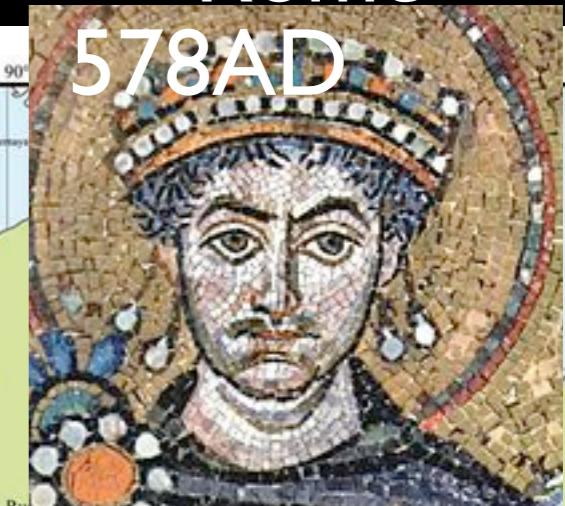
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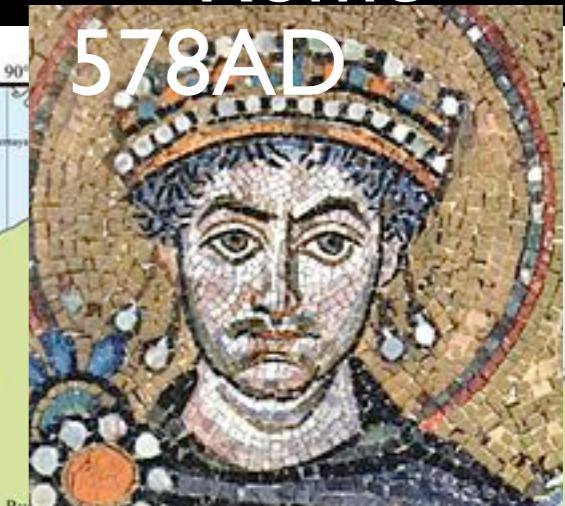
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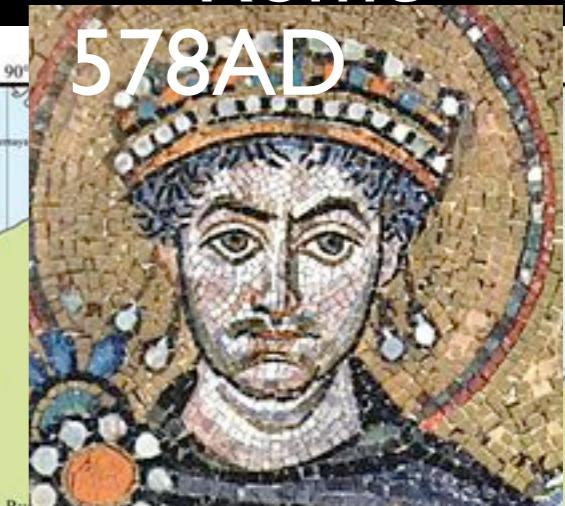
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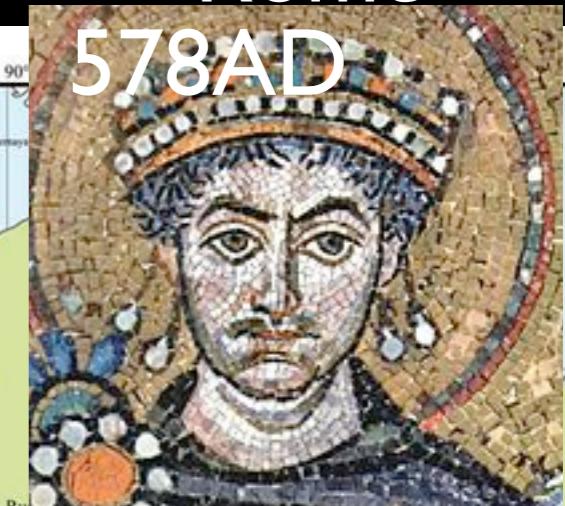
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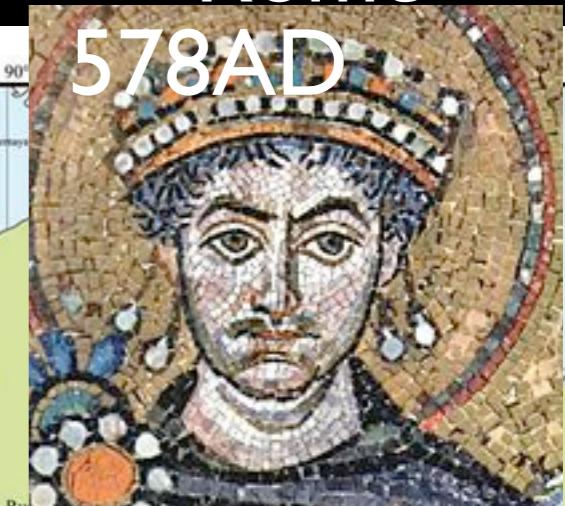
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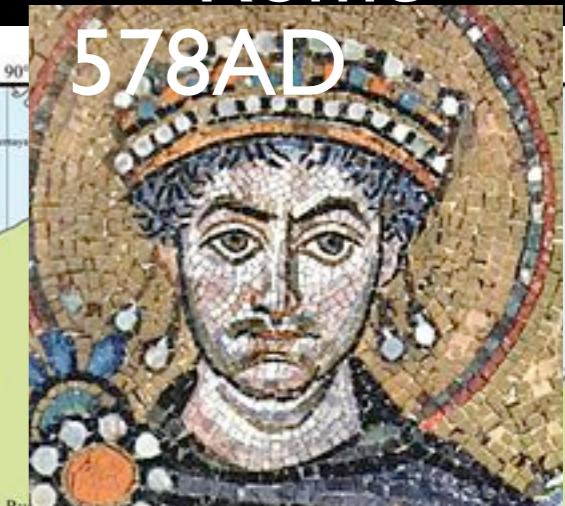
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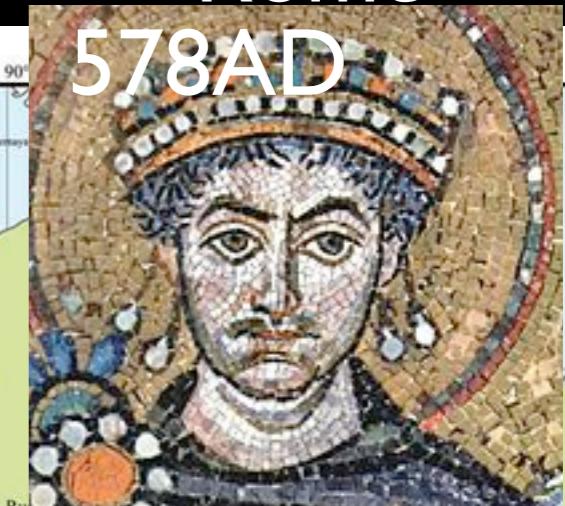
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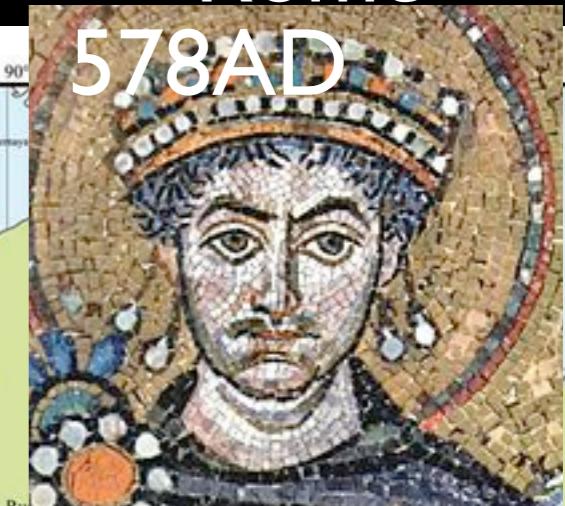
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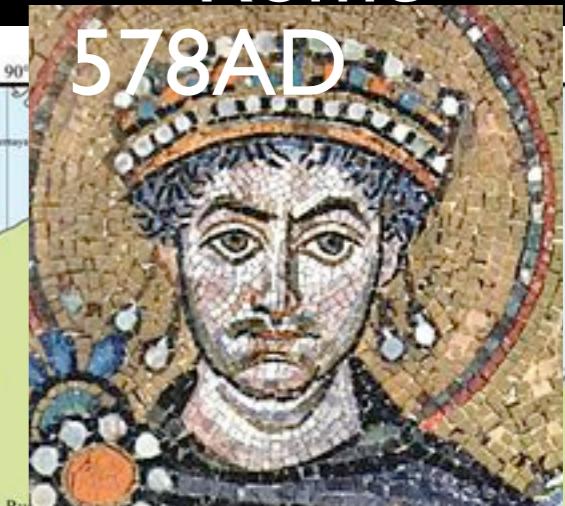
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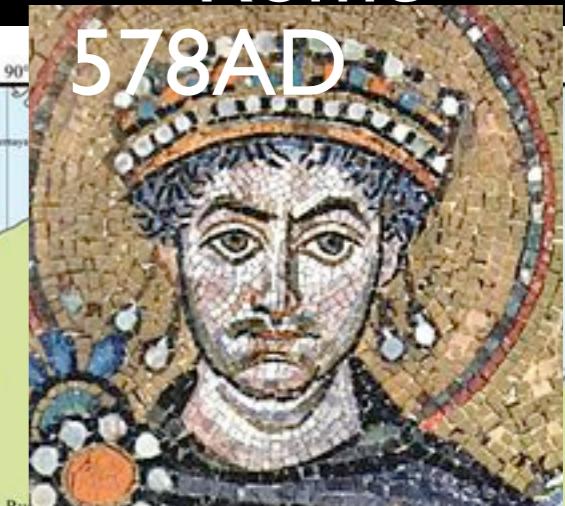
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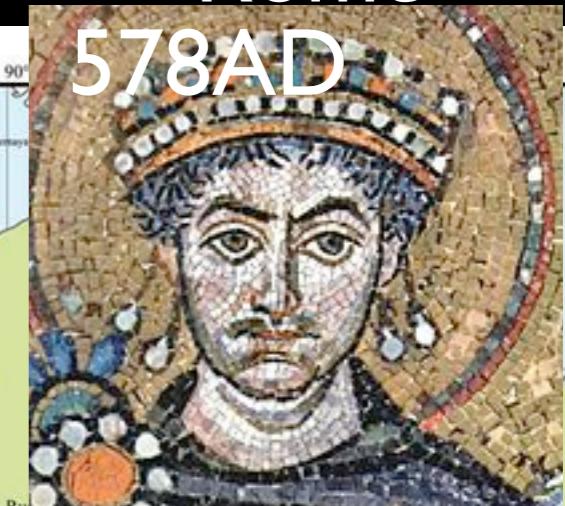
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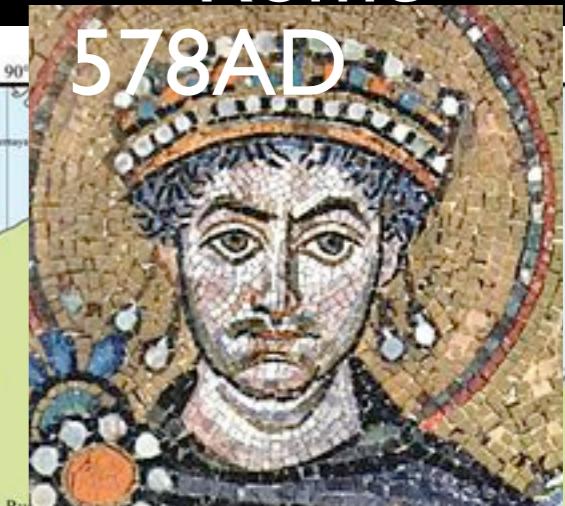
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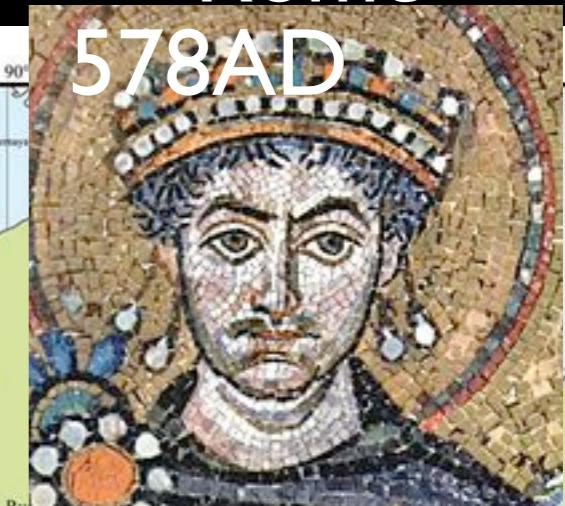
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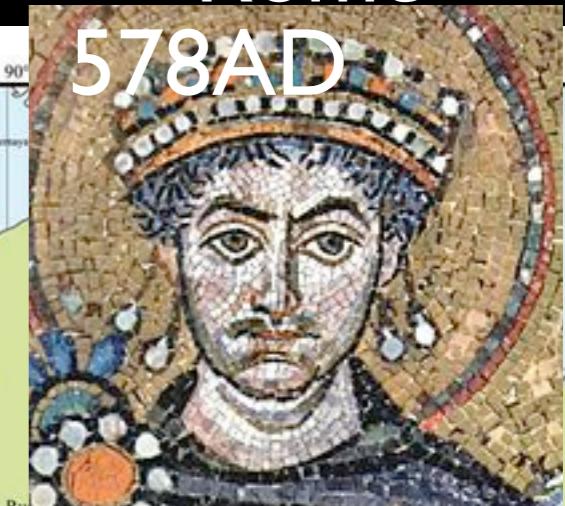
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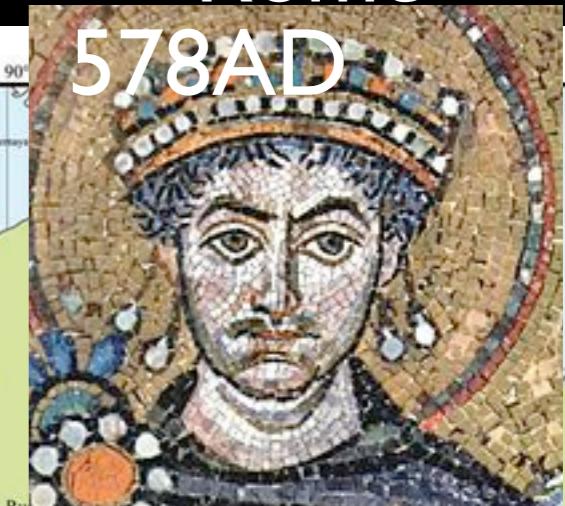
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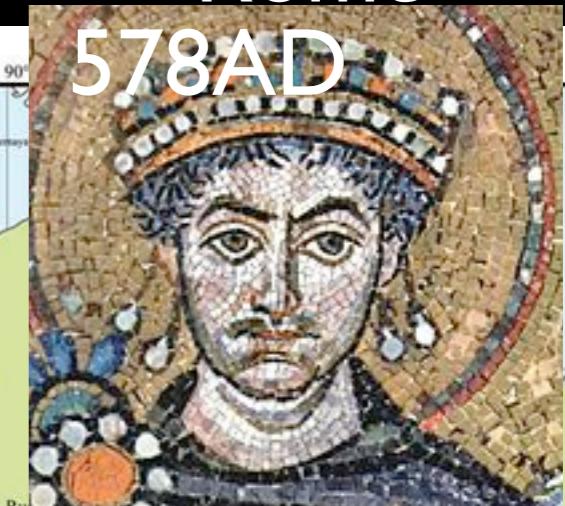
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# Stellar Archeology

- Our galaxy has formed via mergers - “hierarchical structure formation” - smaller units form first then merger to form larger ones
- Some components were galaxies that formed long ago
- Low mass stars have lifetime > age universe so may be still around today
- Search for metal-poor and metal-free stars in our Milky way

LETTER

doi:10.1038/nature12990

A single low-energy, iron-poor supernova  
as the source of metals in the star SMSS  
J031300.36–670839.3