Emmy's \mathcal{M} -Theory Recipes

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From non-perturbative string emerged \mathcal{M} -Theory.

What are its complete degree of freedom?



Quantum deformations by $\alpha_M' \propto (\ell_P)^2$ allowed by:

- b the reparametrisation invariance (GR)
- □ unitarity of the theory (QGR)
- $\triangleright \mathcal{U}$ -duality symmetries $(\mathcal{M} \cdots)$



 \mathcal{M} -Theory is equivalent to a

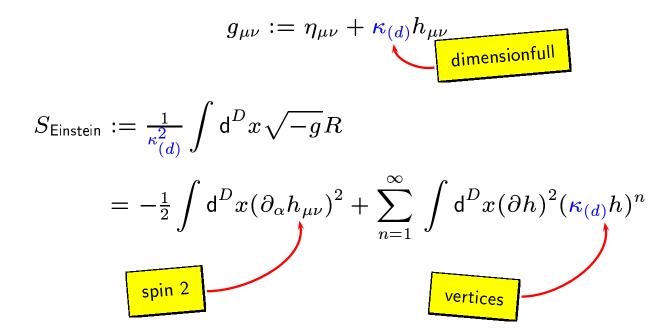
local quantum field theory of gravitation.

Quantum Gravity

Classical General Relativity exists for every $D \geq 4$.

Topologically trivial in D=2 and non-dynamical in D=3.

Defined perturbatively by expanding around a vacuum



Quantum General Relativity receives loop corrections:

▶ Infinite set of counter-terms (non-renormalizability)

't Hooft, Veltman

$$\Delta S_{D=4} = \infty \times \sum_{n \geq 0} \int d^4 x \left(\kappa_{(d)}^2 R \right)^n R^2$$

- Corrections do not violate unitarity nor introduce tachyons or ghosts

[Goroff, Sagnotti], [van de Ven]

▶ Ultra-violet behaviour is background independent

Eight Spoonfuls of SUGRA

Supergravity = GR + fermions with local grassmanian invariance.

$$S \sim \frac{1}{\kappa_{(d)}^2} \int \mathrm{d}^4 x \, \sqrt{-g} \left(R - \frac{i}{2} \bar{\psi}^\alpha{}_\mu \, (\Gamma^{\mu\nu\rho})_{\alpha\beta} \, D_\nu(\omega) \psi^\beta{}_\rho \right)$$

$$\delta^{(0)} e_\mu{}^r = \bar{\epsilon} \Gamma^r \psi_\mu, \quad \delta^{(0)} \psi_\mu = D_\mu(\omega) \epsilon$$
 [Deser, Zumino], [Ferrara et al]

- $ho \quad lpha = 1, \cdots, 4N$ with $N \leq 8$ otherwise exist massless spin higher than 2 and more than one particle with spin 2. [Nahm]
- → A "better" Ultra-violet behaviour [Grisaru], [Deser et al.], [Gates et al.], [Kallosh]

$$\Delta S_{D=4} = 0 \times \int d^4x \, R^2 + 0 \times \int d^4x \, \kappa_{(d)}^2 R^3 + \int d^4x \, \kappa_{(d)}^4 R^4 + \cdots$$
 Field Redefinition Supersymmetry Quantum correction

Boil Everything to D=11

Classical N = 1 D = 11 is

[Cremmer, Julia, Scherk]

$$S = \frac{1}{\kappa_{(d)}^2} \int \mathrm{d}^{11} x e \left(R + \bar{\psi}_{\mu} \Gamma^{\mu\nu\rho} D_{\nu} \psi_{\rho} + G_{(4)}^2 \right) + \frac{\lambda}{6\kappa_{(d)}^2} \int C_{(3)} \wedge G_{(4)} \wedge G_{(4)}$$

ightharpoonup Maximal: no supergravity theories in D>11

Nahm

ightharpoonup Minimal: in its field content: $(e_{\mu}{}^{r}, \psi_{\mu}{}^{\alpha}, C_{(3)})$.

Nahm

▶ Unique: no cosmological constant allowed

Bautier et al

ightharpoonup Non-trivial: has exact maximally supersymmetric vacua: $AdS_{4|7} imes S^{7|4}$

[Kallosh, Rajaraman]

ho Finite ? $\kappa^2_{(11)}=(\ell_P)^9$ does not pair with powers of Riemann tensor R^n

Naive Argument

Incorporate Higher-order Derivative

Finite higher-derivative corrections are necessary:

▶ Anomaly considerations [Duff, Liu, Minasian], [Vafa, Witten], [Green, Schwarz]

$$S = (\alpha'_M)^3 \int C_{(3)} \wedge t_8 R^4.$$

versus

ightharpoonup Strong coupling limit of string theory amplitudes gives finite R^4 , $\square^2 R^4$, . . .

[Green, Kwon, Vanhove], [Green, Gutperle, Vanhove], [Green, Vanhove]

This is an effective theory of a microscopic quantum theory of gravitation.

Add some Brane

Branes are solitonic extended solutions of supergravity theories

$$S = -\int \mathrm{d}^{p+1} \xi \, \sqrt{-\det(g + \mathrm{d}A - B)} + \int C \wedge e^{\mathrm{d}A - B}$$
background fields

Consistency of the world-sheet theory constraints the background fields

 \triangleright higher-loop σ -model β -functions [Grisaru et al]

$$S = \frac{1}{\alpha'} \int d^2 \sigma \, \partial_i X^{\mu} \partial^i X^{\nu} G_{\mu\nu}(X)$$

 \triangleright κ -symmetry invariance

$$\delta_{\kappa}(X^{\mu}(\xi), \theta^{\alpha}(\xi)) = \kappa^{a} E_{a}^{M}, \quad \delta_{\kappa} A = i_{\kappa} B, \quad \kappa^{\alpha}(\xi) = P_{+} \zeta$$

only if the backgrounds $E_M{}^A$, B_{MN} and $C_{M_1...M_k}$ satisfy the on-shell constraints that imply their equation-of-motions

Quantum corrections to the effective brane world-volume theory \iff corrections to \mathcal{M} -Theory.

Serve with Strings

Noether's deformations of the (D=11) supersymmetry algebra

$$\left(\delta_{\epsilon}^{(0)} + \sum_{n\geq 3} \mathbf{a_n} \left(\alpha_M'\right)^n \delta_{\epsilon}^{(n)}\right) \left(S^{(0)} + \sum_{n\geq 3} \mathbf{a_n} \left(\alpha_M'\right)^n S^{(n)}\right) = 0$$

The super-algebra closes on-shell:

 $egin{align} [\delta_{1, ext{sg}},\delta_{2, ext{sg}}] &= \delta^{ ext{translation}}(\xi^{
u}) \ &+ \delta^{ ext{susy}}(-\xi^{
u}\psi_{
u}) \ &+ \delta^{ ext{gauge}}(-\xi^{\sigma}C_{\sigma
u
ho} - 2ar{\epsilon}_2\Gamma_{
u
ho}\epsilon_1) \ &+ \delta^{ ext{Lorentz}}(\xi^{
u}\hat{\omega}_{
u}{}^r{}_t + 4ar{\epsilon}_2S^{r\
u
ho\sigma\kappa}{}_t\epsilon_1\hat{F}_{
u
ho\sigma\kappa}) \,, \end{aligned}$

 ξ^{μ} receives quantum corrections

$$\xi^{\mu} = \bar{\epsilon}_1 \Gamma^{\mu} \epsilon_2 + \frac{a_3}{3} \times 0 \times (\alpha'_M)^3 t_8^{\mu t s_1 \cdots s_6} R_{[r_1 r_2}^{s_1 s_2} R_{r_3 r_4}^{s_3 s_4} R_{r_5]t}^{s_5 s_6}$$

[Peeters, Vanhove, Westerberg]

field dependent parameters

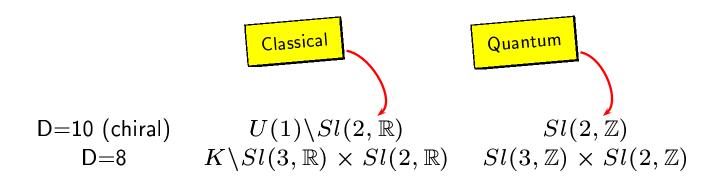
Improve with $\mathcal{U}-$ duality toppings

Is the procedure unique? Field theory reasonning are not enough!

 $\triangleright a_n$ not fixed by supersymmetry alone

All extended supergravity theories in $D \leq 10$ presents non-linear symmetries

 \mathcal{U} — duality groups



- \triangleright Invariance under the quantum symmetry fixes the a_n by constraining the
 - Supersymmetry algebra

Green, Sethi

S-matrix elements

[Green, Kwon, Vanhove], [Green, Gutperle, Vanhove], [Green, Vanhove]

More recipes:

how to accomodate the Membrane with these higher-derivative corrections?