```
(*starting date: 04 May 2021*)
(*analytical model of Alfven wings when the background field has a flow-
 aligned component, B0=(B0x, 0, B0z), where B0x > 0 and B0z < 0*)
(*This model generates output in TWO different coordinate
 systems: (x,y,z) where z is antiparallel to B0,
and (xcor, ycor=y, zcor), where xcor is along the incident flow direction*)
(*05/07/21: update: equations for magnetic field corrected*)
(*17 June 2021: Adjusted to Triton,*)
(*Basic Input Parameters*)
(*-----*)
RE = 1353 * 10^3; (*moon radius in m*)
B0 = 5.1369 * 10^{-9}; (*background field magnitude in T*)
theta = 42.94896 * \frac{2 \pi}{360}; (*inclination angle in rd,
as defined in Figure 3 of Neubauer1980,
theta=0 means u0 perpendicular to B0. !TODO: Theta needs to become a free parameter!*)
B0x = 0;
B0y = 0;
B0z = -B0; (*background magnetic field in system (x,y,z)*)
B0xcor = B0x Cos[theta] - B0z Sin[theta];
B0ycor = 0;
B0zcor = B0x Sin[theta] + B0z Cos[theta];
(*background magnetic field in system (xcor,ycor,zcor)*)
n0 = 0.11 * 10^6; (*upstream number density in 1/m^3*)
u0 = 43 * 10^3; (*magnitude of upstream flow velocity in m/s*)
u0xcor = u0;
u0ycor = 0;
u0zcor = 0; (*upstream flow velocity in system (xcor,ycor,zcor)*)
u0x = u0 Cos[theta];
u0y = 0;
u\theta z = -u\theta \sin[theta];(*upstream flow velocity in system (x,y,z)*)
E0 = u0 * B0 * Cos[theta]; (*MAGNITUDE OF convective electric field*)
E0norm = u0 * B0; (*may be needed for normalization purposes*)
mp = 1.6726231 * 10^{-27}; (*proton mass*)
m = 7.5 * mp; (*upstream ion mass*)
mu0 = 4 * 3.14159265359 * 10^{-7}; (*magnetic permeability of vacuum*)
(*-----*)
(*Derived parameters for Alfven wings*)
vA = \frac{B0}{\sqrt{mu0 * n0 * m}}; (*Alfven velocity*)
MA = u0 / vA (*Alfvenic Mach number*)
                                          ; (*Alfven conductance, northern wing*)
         mu0 * vA * \sqrt{1 + MA^2 - 2 * MA * Sin[theta]}
                                          ; (*Alfven conductance, southern wing*)
         mu0 * vA * \sqrt{1 + MA^2 + 2 * MA * Sin[theta]}
```

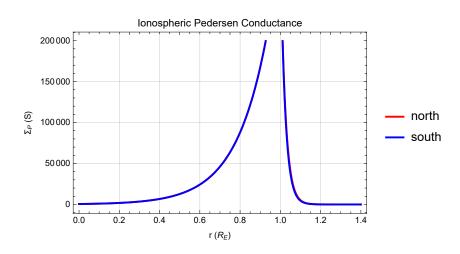
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ThetaN = ArcTan \left[\frac{\text{MA Cos}[\text{theta}]}{1 - \text{MA Sin}[\text{theta}]}\right] * \frac{360}{2 \pi}
(*inclination angle against B0 for NORTHERN Alfven wing*)
ThetaS = ArcTan \left[\frac{\text{MA Cos}[\text{theta}]}{1 + \text{MA Sin}[\text{theta}]}\right] * \frac{360}{2 \pi}
(*inclination angle against B0 for NORTHERN Alfven wing*)
SinThetaN = \frac{MA Cos[theta]}{\sqrt{1 + MA^2 - 2 * MA * Sin[theta]}};
CosThetaN = \frac{1 - MA \sin[theta]}{\sqrt{1 + MA^2 - 2 * MA * Sin[theta]}};
\mbox{SinThetaS} = \frac{\mbox{MA Cos[theta]}}{\sqrt{\mbox{1 + MA}^2 + 2 * MA * Sin[theta]}};
CosThetaS = 1 + MA Sin[theta];
              \sqrt{1 + MA^2 + 2 * MA * Sin[theta]}
0.348577
18.5011
11.6498
(*-----*)
(*Ionosphere model, mostly taken from Simon2021 analytical model,
in coordinates (x,y,z)*)
(*Important: need SAME conductance profile in both hemispheres,
otherwise boundary condition from Saur2007 needs to be applied*)
(*-----*)
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SigmaP = 500; (*exospheric Pedersen conductance,
called SigmaP0 in the 2021 paper, STILL SAME SIGMAP0 IN BOTH HEMISPHERES*)
H = 70 * 10<sup>3</sup>; (*exospheric scale height, from Arnold2020b*)
r[x, y] = \sqrt{x^2 + y^2}; (*radial distance from z axis in cylindrical coordinates,
NOT radial distance to Europa's surface,
measured perpendicular to the background magnetic field or z axis,
(x,y) dependence NOT needed for merely plotting the exospheric conductance profile*)
R1 = RE; (*radial boundary of the Europa fluxtube, x^2+y^2 = RE^2)
R2 = RE + 5H; (*outer boundary of region with non-zero transverse conductance,
assumming the atmosphere "ends" at a distance of 5 scale heights. exp(-5)=0.00674*)
betaN = 1/(3 \text{ H}); (*NEW 05/2021: north only*)
betaS =
  1 Log[((SigmaP + SigmaAn) Exp[betaN * R1] + (SigmaAs - SigmaAn)) / (SigmaP + SigmaAs)];
(*given input parameter for conductance profile,
using some multiple of the exospheric scale height H here,
NEW; BETA IS DIFFERENT FOR NORTHERN AND SOUTHERN HEMISPHERES*)
Sigma1n[r_] = (SigmaP + SigmaAn) * Exp[betaN * r] - SigmaAn;
Sigma1s[r_] = (SigmaP + SigmaAs) * Exp[betaS * r] - SigmaAs;
(*Pedersen conductance within the Europa fluxtube r < R1,
05/2021: NOW DIFFERENT FOR NORTH AND SOUTH!!*)
         Log \left[ \frac{SigmaP + SigmaAn}{cigmaAn} \right] + betaN * R1
deltaN = -
         Log [ SigmaAn Exp[deltaN* (R2-R1)]+SigmaAs-SigmaAn]
                     SigmaAs
deltaS = -
Sigma2n[r_] = SigmaAn * Exp[deltaN * (R2 - r)] - SigmaAn;
Sigma2s[r_] = SigmaAs * Exp[deltaS * (R2 - r)] - SigmaAs;
(*Pedersen conductance outside the Europa fluxtube,
but within the exosphere R1 < r < R2, NEW 05/2021: NOW DIFFERENT FOR NORTH AND SOUTH!!*)
Sigma3[r ] = 0; (*Pedersen conductance outside of Europa fluxtube,
R2 < r, always the same in both hemispheres, no N/S splitting required*)
SigmaExon[r] =
  Piecewise[{Sigma1n[r], r \leq R1}, {Sigma2n[r], R1 < r \leq R2}, {Sigma3[r], R2 < r}];
SigmaExos[r_] = Piecewise[{{Sigma1s[r], r ≤ R1},
    \{Sigma2s[r], R1 < r \le R2\}, \{Sigma3[r], R2 < r\}\}\};
(*putting all three segments together, 05/2021: NOW DIFFERENT EQUATIONS FOR NORTH
  AND SOUTH BUT PROFILES SHOULD STILL LOOK N E A R L Y identical in regions 1 and 2!!*)
```

```
(*Ionosphere model, plots and derived parameters*)
AvSigmaPn = \frac{1}{R2} \int_{0}^{R2} SigmaExon[r] dr
  (*average Pedersen conductivity, northern hemisphere*)
AvSigmaPs = \frac{1}{R2} \int_{0}^{R2} SigmaExos[r] dr
  (*average Pedersen conductivity, southern hemisphere*)
AvSigmaPn / AvSigmaPs (*should be as close as possible to 1*)
 alphaN = -

    (*average interaction strength,

                                    AvSigmaPn + 2 * SigmaAn
 north !this is only an estimation!*)
                                                          AvSigmaPs
                                                                                                                       (*average interaction strength,
 alphaS = -
                                    AvSigmaPs + 2 * SigmaAs
 north !this is only an estimation!*)
 Plot[\{SigmaExon[r*RE], SigmaExos[r*RE]\}, \{r, 0, 1.4\}, FrameLabel \rightarrow \{"r (R_E)", "\Sigma_P (S)"\}, \{r, 0, 1.4\}, FrameLabel \rightarrow \{"r (R_E)", "\Sigma_P (S)"\}, \{r, 0, 1.4\}, FrameLabel \rightarrow \{"r (R_E)", "\Sigma_P (S)"\}, \{r, 0, 1.4\}, FrameLabel \rightarrow \{"r (R_E)", "\Sigma_P (S)"\}, \{r, 0, 1.4\}, FrameLabel \rightarrow \{"r (R_E)", "\Sigma_P (S)"\}, \{r, 0, 1.4\}, FrameLabel \rightarrow \{"r (R_E)", "\Sigma_P (S)"\}, \{r, 0, 1.4\}, FrameLabel \rightarrow \{"r (R_E)", "\Sigma_P (S)"\}, \{r, 0, 1.4\}, \{r
     GridLines → Automatic, Frame → True, PlotLabel → "Ionospheric Pedersen Conductance",
     PlotStyle → {{Red, Thick}, {Blue, Thick}}, PlotLegends -> {"north", "south"}]
45475.7
45190.2
1.00632
0.999647
0.999774
```



```
(*05 May 2021: Potentials, magnetic field and flow velocity, all in system
  (*linear system, northern wing*)
 \left\{ \frac{1 - \left( \text{betaN} * \text{R1} + 1 \right) * \text{Exp[-betaN} * \text{R1}]}{\text{R1}^2}, -\frac{1}{\text{R1}^2}, -\frac{1}{\text{R1}^2} \text{Exp[deltaN} * \text{R1}] \left( \text{deltaN} * \text{R1} - 1 \right), \right. \\ \left. \theta \right\}, \left\{ \theta, \frac{1}{\text{R2}^2}, \frac{\text{Exp[deltaN} * \text{R2}] * \left( \text{deltaN} * \text{R2} - 1 \right)}{\text{R2}^2}, \frac{1}{\text{R2}^2} \right\} \right\}; 
 c = \{\{0\}, \{E0 * R2\}, \{0\}, \{E0\}\}\}
 MatrixForm[Mn];
  {{K1n}, {K3n}, {K4n}, {K5n}} = Inverse[Mn].c;
  (*linear system, southern wing*)
 Ms = \left\{ \left\{ \frac{\text{betaS} * R1 - 1 + \text{Exp}[-\text{betaS} * R1]}{\text{R1}}, \frac{\text{deltaS} * R1 + 1}{\text{R1}}, \frac{-\text{Exp}[\text{deltaS} * R1]}{\text{R1}}, 0 \right\},
             \{\theta, -\frac{\text{deltaS} * R2 + 1}{R2}, \frac{\text{Exp[deltaS} * R2]}{R2}, -\frac{1}{R2}\},
              \{ \frac{1 - \left( \text{betaS} * \text{R1} + 1 \right) * \text{Exp[-betaS} * \text{R1}]}{\text{R1}^2}, -\frac{1}{\text{R1}^2}, -\frac{1}{\text{R1}^2} \text{Exp[deltaS} * \text{R1}] \left( \text{deltaS} * \text{R1} - 1 \right), \\ \emptyset \}, \left\{ \emptyset, \frac{1}{\text{R2}^2}, \frac{\text{Exp[deltaS} * \text{R2}] * \left( \text{deltaS} * \text{R2} - 1 \right)}{\text{R2}^2}, \frac{1}{\text{R2}^2} \right\} \}; 
  (*same c vector as for northern wing*
 MatrixForm[Ms];
  {{K1s}, {K3s}, {K4s}, {K5s}} = Inverse[Ms].c;
  Inverse: Result for Inverse of badly conditioned matrix
                  \{ \{4.02398 \times 10^{-6}, 0.0000309993, -4.46292 \times 10^{11}, 0.\}, \{0., -\ll 20\%, \ll 23\%, -5.87199 \times 10^{-7}\}, \{\ll 1\%\}, \{0., 3.44803 \times 10^{-13}, 4.18514\}, \{0., 3.48514\}, \{0., 3.48514\}, \{0., 3.48514\}, \{0., 3.48514\}, \{0., 3.48514\}, \{0., 3.48514\}, \{0., 3.48514\}, \{0., 3.4
                                 \times 10<sup>11</sup>, 3.44803 \times 10<sup>-13</sup>}} may contain significant numerical errors.
  Inverse: Result for Inverse of badly conditioned matrix
                 \{\{4.02823 \times 10^{-6}, 0.0000322907, -2.56121 \times 10^{12}, 0.\}, \{0., -\infty24 \gg, \infty22 \gg, -5.87199 \times 10^{-7}\}, \{\infty1 \gg\}, \{0., 3.44803 \times 10^{-13}, 3.93852\}\}
                                 \times 10<sup>12</sup>, 3.44803 \times 10<sup>-13</sup>}} may contain significant numerical errors.
  (*Potentials, their derivatives, magnetic field, and bulk velocity*)
  sinphi[x_, y_] = \frac{y}{\sqrt{x^2 + v^2}};
 cosphi[x_, y_] = \frac{x}{\sqrt{x^2 + y^2}};
  (*potential psi, northern wing*)
         sinphi[x, y] * \frac{1}{r[x, y]} (K1n * (betaN * r[x, y] - 1) + K1n * Exp[-betaN * r[x, y]]);
 psi2n[x_, y_] = sinphi[x, y] *
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\frac{1}{r[x, y]} \left( -K3n * \left( deltaN * r[x, y] + 1 \right) + K4n * Exp[deltaN * r[x, y]] \right);
psi3n[x_, y_] = sinphi[x, y] * \left(E0 * r[x, y] + \frac{K5n}{r[x, y]}\right);
psin[x_{y}] = Piecewise[{0, r[x, y] = 0}, {psi1n[x, y], 0 < r[x, y] \le R1},
      \{psi2n[x, y], R1 < r[x, y] \le R2\}, \{psi3n[x, y], R2 < r[x, y]\}\}\}
(*potential psi, southern wing*)
psi1s[x_, y_] =
   sinphi[x, y] * \frac{1}{r[x, y]} (K1s * (betaS * r[x, y] - 1) + K1s * Exp[-betaS * r[x, y]]);
psi2s[x_, y_] = sinphi[x, y] *
    \frac{1}{r[x, y]} \left( -K3s * \left( deltaS * r[x, y] + 1 \right) + K4s * Exp[deltaS * r[x, y]] \right);
psi3s[x_, y_] = sinphi[x, y] * \left(E0 * r[x, y] + \frac{K5s}{r[x, y]}\right);
psis[x_y] = Piecewise[{{0, r[x, y] = 0}, {psi1s[x, y], 0 < r[x, y] \le R1},
      \{psi2s[x, y], R1 < r[x, y] \le R2\}, \{psi3s[x, y], R2 < r[x, y]\}\}\}
(*derivatives of psi, magnetic field, and flow pattern u within the Alfven wings*)
(*derivatives, northern wing*)
dxpsi1n[x_, y_] = cosphi[x, y] * sinphi[x, y] * K1n *
      ((1 - Exp[-betaN * r[x, y]] * (betaN * r[x, y] + 1)) / (r[x, y] * r[x, y])) - \frac{sinphi[x, y]}{r[x, y]} * (betaN * r[x, y]) / (r[x, y] * r[x, y])
      cosphi[x, y] * \frac{1}{r[x, y]} (K1n * (betaN * r[x, y] - 1) + K1n * Exp[-betaN * r[x, y]]);
dxpsi2n[x_, y_] = cosphi[x, y] * sinphi[x, y] *
      ((K3n + K4n * Exp[deltaN * r[x, y]] * (deltaN * r[x, y] - 1)) / (r[x, y] * r[x, y])) -
     \frac{\text{sinphi}[x, y]}{r[x, y]} * \text{cosphi}[x, y] *
      \frac{1}{r[x, y]} \left( -K3n * \left( deltaN * r[x, y] + 1 \right) + K4n * Exp[deltaN * r[x, y]] \right);
dxpsi3n[x_, y_] = cosphi[x, y] * sinphi[x, y] * \left(E0 - \frac{K5n}{r[x, y] * r[x, y]}\right) -
     \frac{\text{sinphi}[x, y]}{r[x, y]} * \text{cosphi}[x, y] * \left(E0 * r[x, y] + \frac{K5n}{r[x, y]}\right);
dypsiln[x\_, y\_] = sinphi[x, y] * sinphi[x, y] * K1n *
      ((1 - Exp[-betaN * r[x, y]] * (betaN * r[x, y] + 1)) / (r[x, y] * r[x, y])) + \frac{cosphi[x, y]}{r[x, y]} *
      cosphi[x, y] * \frac{1}{r[x. v]} (K1n * (betaN * r[x, y] - 1) + K1n * Exp[-betaN * r[x, y]]);
dypsi2n[x_, y_] = sinphi[x, y] * sinphi[x, y] *
      \left(\left(\mathsf{K3n} + \mathsf{K4n} * \mathsf{Exp}[\mathsf{deltaN} * r[x, y]] * \left(\mathsf{deltaN} * r[x, y] - 1\right)\right) / (r[x, y] * r[x, y])\right) +
     cosphi[x, y]
r[x v] * cosphi[x, y] *
      \frac{1}{r[x,y]} \left( -K3n * \left( deltaN * r[x,y] + 1 \right) + K4n * Exp[deltaN * r[x,y]] \right);
dypsi3n[x_{,} y_{]} = sinphi[x, y] * sinphi[x, y] * \left(E0 - \frac{K5n}{r[x, y] * r[x, y]}\right) +
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\frac{\mathsf{cosphi}[x,y]}{\mathsf{r}[x,y]} * \mathsf{cosphi}[x,y] * \left(\mathsf{E0} * \mathsf{r}[x,y] + \frac{\mathsf{K5n}}{\mathsf{r}[x,y]}\right);
dxpsin[x_, y_] = Piecewise[{{0, r[x, y] == 0}, {dxpsi1n[x, y], 0 < r[x, y] \le R1},
      \{dxpsi2n[x, y], R1 < r[x, y] \le R2\}, \{dxpsi3n[x, y], R2 < r[x, y]\}\}\}
dypsin[x_{,}y_{]} = Piecewise\left[\left\{\left\{K1n * \frac{betaN^{2}}{2}, r[x, y] = 0\right\}, \left\{dypsiln[x, y], 0 < r[x, y] \le R1\right\}\right\}
      \{dypsi2n[x, y], R1 < r[x, y] \le R2\}, \{dypsi3n[x, y], R2 < r[x, y]\}\};
(*derivaties, southern wing*)
dxpsi1s[x_, y_] = cosphi[x, y] * sinphi[x, y] * K1s *
      ((1 - Exp[-betaS * r[x, y]] * (betaS * r[x, y] + 1)) / (r[x, y] * r[x, y])) - \frac{sinpni[x, y]}{r[x, y]} *
      cosphi[x, y] * \frac{1}{r[x, y]} (K1s * (betaS * r[x, y] - 1) + K1s * Exp[-betaS * r[x, y]]);
dxpsi2s[x_, y_] = cosphi[x, y] * sinphi[x, y] *
      \left(\left(\mathsf{K3s} + \mathsf{K4s} * \mathsf{Exp}[\mathsf{deltaS} * r[x, y]] * \left(\mathsf{deltaS} * r[x, y] - 1\right)\right) / (r[x, y] * r[x, y])\right) -
     sinphi[x, y]
 * cosphi[x, y] *
      \frac{1}{r[x, y]} \left( -K3s * \left( deltaS * r[x, y] + 1 \right) + K4s * Exp[deltaS * r[x, y]] \right);
dxpsi3s[x_, y_] = cosphi[x, y] * sinphi[x, y] * \left(E0 - \frac{K5s}{r[x, y] * r[x, y]}\right) - \frac{K5s}{r[x, y] * r[x, y]}
     \frac{\sinh[x, y]}{r[x, y]} * \cosh[x, y] * \left(E0 * r[x, y] + \frac{K5s}{r[x, y]}\right);
dypsils[x_, y_] = sinphi[x, y] * sinphi[x, y] * K1s *
      ((1 - Exp[-betaS * r[x, y]] * (betaS * r[x, y] + 1)) / (r[x, y] * r[x, y])) + \frac{cosphi[x, y]}{r[x, y]} *
      cosphi[x, y] * \frac{1}{r[x, y]} (K1s * (betaS * r[x, y] - 1) + K1s * Exp[-betaS * r[x, y]]);
dypsi2s[x_, y_] = sinphi[x, y] * sinphi[x, y] *
      ((K3s + K4s * Exp[deltaS * r[x, y]] * (deltaS * r[x, y] - 1)) / (r[x, y] * r[x, y])) +
     cosphi[x, y]
* cosphi[x, y] *
      \frac{1}{r[x, v]} \left(-K3s * \left(deltaS * r[x, y] + 1\right) + K4s * Exp[deltaS * r[x, y]]\right);
dypsi3s[x_, y_] = sinphi[x, y] * sinphi[x, y] * \left(E0 - \frac{K5s}{r[x, y] * r[x, y]}\right) + \frac{K5s}{r[x, y] * r[x, y]}
     \frac{\cosh[x, y]}{r[x, y]} * \cosh[x, y] * \left( E0 * r[x, y] + \frac{K5s}{r[x, y]} \right);
dxpsis[x_{y}] = Piecewise[{0, r[x, y] = 0}, {dxpsi1s[x, y], 0 < r[x, y] \le R1},
      \{dxpsi2s[x, y], R1 < r[x, y] \le R2\}, \{dxpsi3s[x, y], R2 < r[x, y]\}\}\}
dypsis[x_{,y_{]}} = Piecewise\left[\left\{\left\{K1s * \frac{betaS^{2}}{2}, r[x, y] = 0\right\}, \left\{dypsi1s[x, y], 0 < r[x, y] \le R1\right\},\right\}
      \{dypsi2s[x, y], R1 < r[x, y] \le R2\}, \{dypsi3s[x, y], R2 < r[x, y]\}\}\}
(*magnetic field, northern wing*)
Bxnorth[x_, y_] = -SinThetaN *
      \sqrt{(B0^2 - mu0^2 * SigmaAn^2 * ((CosThetaN)^2 (dxpsin[x, y])^2 + (dypsin[x, y])^2)} +
    mu0 * SigmaAn * CosThetaN * dypsin[x, y];
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Bynorth[x_, y_] = -CosThetaN * mu0 * SigmaAn * dxpsin[<math>x, y];
Bznorth[x_, y_] = -CosThetaN *
      \sqrt{(B0^2 - mu0^2 * SigmaAn^2 * ((CosThetaN)^2 (dxpsin[x, y])^2 + (dypsin[x, y])^2)} -
    mu0 * SigmaAn * SinThetaN * dypsin[x, y];
(*magnetic field, southern wing*)
Bxsouth[x_, y_] = SinThetaS *
      \sqrt{(B0^2 - mu0^2 * SigmaAs^2 * ((CosThetaS)^2 (dxpsis[x, y])^2 + (dypsis[x, y])^2)} -
    mu0 * SigmaAs * CosThetaS * dypsis[x, y];
Bysouth[x_, y_] = CosThetaS * mu0 * SigmaAs * dxpsis[x, y];
Bzsouth[x_, y_] = -CosThetaS *
      \sqrt{(B0^2 - mu0^2 * SigmaAs^2 * ((CosThetaS)^2 (dxpsis[x, y])^2 + (dypsis[x, y])^2)} -
    SinThetaS * mu0 * SigmaAs * dypsis[x, y];
(*velocity field, northern wing*)
uxnorth[x_, y_] = u0x + \frac{Bxnorth[x, y] - B0x}{\sqrt{mu0 * n0 * m}};
uynorth[x_, y_] = u\theta y + \frac{Bynorth[x, y] - B\theta y}{---}:
                                \sqrt{\text{mu0} * \text{n0} * \text{m}}
uznorth[x_, y_] = u0z + \frac{Bznorth[x, y] - B0z}{\sqrt{mu0 * n0 * m}};
(*velocity field, southern wing*)
uxsouth[x_, y_] = u0x - \frac{Bxsouth[x, y] - B0x}{\sqrt{mu0 * n0 * m}};
uysouth[x_, y_] = u\theta y - \frac{Bysouth[x, y] - B\theta y}{2}
                                 \sqrt{\text{mu0} * \text{n0} * \text{m}}
uzsouth[x_, y_] = u0z - \frac{Bzsouth[x, y] - B0z}{\sqrt{mu0 * n0 * m}};
```

```
(*a few plots from Simon2021, to make sure that this all works*)
(*magnetic field, see figure 6 from Simon2021 northern and southern wing
 should be somewhat different now*)
Plotmagnorth = Plot[\{Bxnorth[0.4RE, y*RE]/(B0),\}
    Bynorth [0.4 RE, y * RE] / (B0), Bznorth [0.4 RE, y * RE] / (B0) + B0 / B0},
  \{y, -3, 3\}, FrameLabel \rightarrow {"y (R<sub>E</sub>)", "B<sub>x/y/z</sub> / B<sub>0</sub>"}, GridLines \rightarrow Automatic,
  Frame → True, PlotLabel → "Magnetic Field in Northern Wing, x = 0.4 R<sub>F</sub> Line",
  PlotStyle → {{Red, Thick}, {Green, Thick}, {Blue, Thick}},
  Epilog \rightarrow {Directive[Dashed, Orange], InfiniteLine[{{-0.9165, -0.3}, {-0.9165, 0.3}}],
     InfiniteLine[{{0.9165, -0.3}, {0.9165, 0.3}}],
     InfiniteLine[{{1.258, -0.3}, {1.258, 0.3}}], InfiniteLine[
       {{-1.258, -0.3}, {-1.258, 0.3}}]} (*,PlotLegends ->{"B<sub>x</sub>","B<sub>v</sub>","B<sub>z</sub>"}*)
Plotmagsouth = Plot[{Bxsouth[0.4 RE, y * RE] / (B0), Bysouth[0.4 RE, y * RE] / (B0),
    Bzsouth [0.4 \text{ RE}, y * \text{RE}] / (B0) + B0 / B0, \{y, -3, 3\},
  FrameLabel \rightarrow {"y (R<sub>E</sub>)", "B<sub>x/y/z</sub> / B<sub>0</sub>"}, GridLines \rightarrow Automatic, Frame \rightarrow True,
  PlotLabel → "Magnetic Field in Southern Wing, x = 0.4 R<sub>E</sub> Line",
  PlotStyle → {{Red, Thick}, {Green, Thick}, {Blue, Thick}},
  Epilog \rightarrow {Directive [Dashed, Orange], InfiniteLine [{{-0.9165, -0.3}, {-0.9165, 0.3}}],
     InfiniteLine[{{0.9165, -0.3}, {0.9165, 0.3}}],
     InfiniteLine[{{1.258, -0.3}, {1.258, 0.3}}],
     InfiniteLine[\{\{-1.258, -0.3\}, \{-1.258, 0.3\}\}\}\} (*,PlotLegends ->\{"B_x","B_y","B_z"\}*)
              Magnetic Field in Northern Wing, x = 0.4 R_E Line
   0.2
   0.1
   0.0
   -0.1
   -0.2
```

-0.3

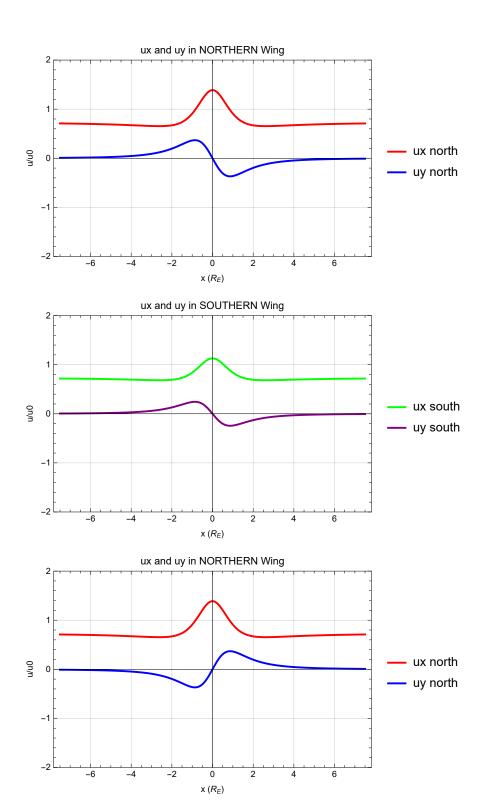
 $y(R_E)$

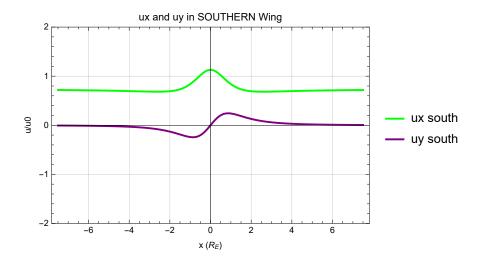
-0.10 -0.15

_1

 $y(R_E)$

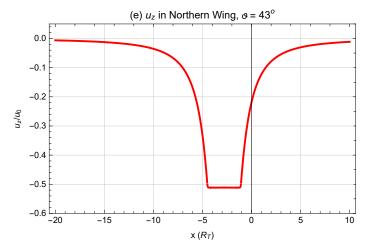
```
(*velocity field, figure 2 of Simon2021*)
uxnorthshiftpos[x_] = uxnorth[x, 1.5 RE] / u0;
(*cut at y=1.5 RE, i.e., displaced TOWARD Jupiter*)
uynorthshiftpos[x] = uynorth[x, 1.5 RE] / u0;
uxsouthshiftpos[x_] = uxsouth[x, 1.5 RE] / u0;
uysouthshiftpos[x_] = uysouth[x, 1.5 RE] /u0;
Plot[{uxnorthshiftpos[x * RE], uynorthshiftpos[x * RE]}, {x, -7.5, 7.5},
FrameLabel → {"x (R<sub>E</sub>)", "u/u0"}, GridLines → Automatic, Frame → True,
PlotLabel → "ux and uy in NORTHERN Wing", PlotStyle → {{Red, Thick}, {Blue, Thick}},
PlotLegends -> {"ux north", "uy north"}, PlotRange -> {-2, 2}]
Plot[{uxsouthshiftpos[x * RE], uysouthshiftpos[x * RE]}, {x, -7.5, 7.5},
 FrameLabel \rightarrow {"x (R<sub>E</sub>)", "u/u0"}, GridLines \rightarrow Automatic, Frame \rightarrow True,
PlotLabel → "ux and uy in SOUTHERN Wing", PlotStyle → {{Green, Thick}, {Purple, Thick}},
PlotLegends -> {"ux south", "uy south"}, PlotRange -> {-2, 2}]
uxnorthshiftneg[x_] = uxnorth[x, -1.5 RE] / u0; (*cut at y=-1.5 RE,
i.e., displaced AWAY FROM Jupiter*)
uynorthshiftneg[x_] = uynorth[x, -1.5 RE] / u0;
uxsouthshiftneg[x] = uxsouth[x, -1.5 RE] / u0;
uysouthshiftneg[x] = uysouth[x, -1.5 RE] / u0;
Plot[{uxnorthshiftneg[x * RE], uynorthshiftneg[x * RE]}, {x, -7.5, 7.5},
FrameLabel \rightarrow {"x (R<sub>E</sub>)", "u/u0"}, GridLines \rightarrow Automatic, Frame \rightarrow True,
PlotLabel → "ux and uy in NORTHERN Wing", PlotStyle → {{Red, Thick}, {Blue, Thick}},
PlotLegends -> {"ux north", "uy north"}, PlotRange -> {-2, 2}]
Plot[{uxsouthshiftneg[x * RE], uysouthshiftneg[x * RE]}, {x, -7.5, 7.5},
 FrameLabel \rightarrow {"x (R<sub>E</sub>)", "u/u0"}, GridLines \rightarrow Automatic, Frame \rightarrow True,
PlotLabel → "ux and uy in SOUTHERN Wing", PlotStyle → {{Green, Thick}, {Purple, Thick}},
PlotLegends -> {"ux south", "uy south"}, PlotRange -> {-2, 2}]
```

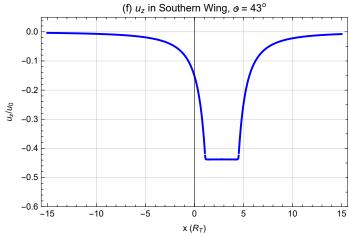




```
(*so far, everything is given in coordinate system (x,y,z) where z is
 antiparallel to the background field. Now we need to transform into
 system (xcor,ycor,zcor) where xcor is along the coritation direction*)
(*-----
(*position vectors*)
x[xcor_, ycor_, zcor_] = xcor * Cos[theta] + zcor * Sin[theta];
y[xcor_, ycor_, zcor_] = ycor;
z[xcor_, ycor_, zcor_] = -xcor * Sin[theta] + zcor * Cos[theta];
(*z is actually not even needed,
since in the (x,y,z) system the solution is constant in z direction,
i.e., there are no z dependencies in any quantity*)
(*magnetic field components in northern wing*)
Bxnorthcor[xcor_, ycor_, zcor_] =
  Bxnorth[x[xcor, ycor, zcor], y[xcor, ycor, zcor]] * Cos[theta] -
   Bznorth[x[xcor, ycor, zcor], y[xcor, ycor, zcor]] * Sin[theta];
Bynorthcor[xcor_, ycor_, zcor_] = Bynorth[x[xcor, ycor, zcor], y[xcor, ycor, zcor]];
Bznorthcor[xcor_, ycor_, zcor_] =
  Bxnorth[x[xcor, ycor, zcor], y[xcor, ycor, zcor]] * Sin[theta] +
   Bznorth[x[xcor, ycor, zcor], y[xcor, ycor, zcor]] * Cos[theta];
(*magnetic field components in southern wing*)
Bxsouthcor[xcor_, ycor_, zcor_] =
  Bxsouth[x[xcor, ycor, zcor], y[xcor, ycor, zcor]] * Cos[theta] -
   Bzsouth[x[xcor, ycor, zcor], y[xcor, ycor, zcor]] * Sin[theta];
Bysouthcor[xcor_, ycor_, zcor_] = Bysouth[x[xcor, ycor, zcor], y[xcor, ycor, zcor]];
Bzsouthcor[xcor_, ycor_, zcor_] =
  Bxsouth[x[xcor, ycor, zcor], y[xcor, ycor, zcor]] * Sin[theta] +
   Bzsouth[x[xcor, ycor, zcor], y[xcor, ycor, zcor]] * Cos[theta];
(*bulk velocity in northern wing*)
uxnorthcor[xcor_, ycor_, zcor_] =
  uxnorth[x[xcor, ycor, zcor], y[xcor, ycor, zcor]] * Cos[theta] -
   uznorth[x[xcor, ycor, zcor], y[xcor, ycor, zcor]] * Sin[theta];
uynorthcor[xcor_, ycor_, zcor_] = uynorth[x[xcor, ycor, zcor], y[xcor, ycor, zcor]];
uznorthcor[xcor_, ycor_, zcor_] =
  uxnorth[x[xcor, ycor, zcor], y[xcor, ycor, zcor]] * Sin[theta] +
   uznorth[x[xcor, ycor, zcor], y[xcor, ycor, zcor]] * Cos[theta];
(*bulk velocity in southern wing*)
uxsouthcor[xcor_, ycor_, zcor_] =
  uxsouth[x[xcor, ycor, zcor], y[xcor, ycor, zcor]] * Cos[theta] -
   uzsouth[x[xcor, ycor, zcor], y[xcor, ycor, zcor]] * Sin[theta];
uysouthcor[xcor_, ycor_, zcor_] = uysouth[x[xcor, ycor, zcor], y[xcor, ycor, zcor]];
uzsouthcor[xcor_, ycor_, zcor_] =
  uxsouth[x[xcor, ycor, zcor], y[xcor, ycor, zcor]] * Sin[theta] +
   uzsouth[x[xcor, ycor, zcor], y[xcor, ycor, zcor]] * Cos[theta];
```

```
(*now: plotting the results in system (xcor,ycor, zcor)*)
Plot1 = Plot[uznorthcor[xcor * RE, 0, 3 * RE] /u0, {xcor, -20, 10},
   FrameLabel \rightarrow {"x (R<sub>T</sub>)", "u<sub>z</sub>/u<sub>0</sub>"}, GridLines \rightarrow Automatic, Frame \rightarrow True,
   PlotLabel \rightarrow "(e) u_z in Northern Wing, oetage = 43^{\circ}", PlotStyle oetage \in \{\text{Red, Thick}\}\
   (*,PlotLegends ->{"uz north"}*), PlotRange → {-0.6, 0.05}]
Plot2 = Plot \left[ uzsouthcor[xcor * RE, 0, -3 * RE] / u0, \{xcor, -15, 15\} \right]
   FrameLabel \rightarrow {"x (R<sub>T</sub>)", "u<sub>z</sub>/u<sub>0</sub>"}, GridLines \rightarrow Automatic, Frame \rightarrow True,
   PlotLabel \rightarrow "(f) u_z in Southern Wing, oldsymbol{o} = 43^{\circ}", PlotStyle \rightarrow {{Blue, Thick}},
   PlotRange \rightarrow \{-0.6, 0.05\} (*,PlotLegends -> \{"uz south"\}*)
Export["uz_north43.pdf", Plot1]
Export["uz_south43.pdf", Plot2]
```





uz_north43.pdf

uz_south43.pdf

