

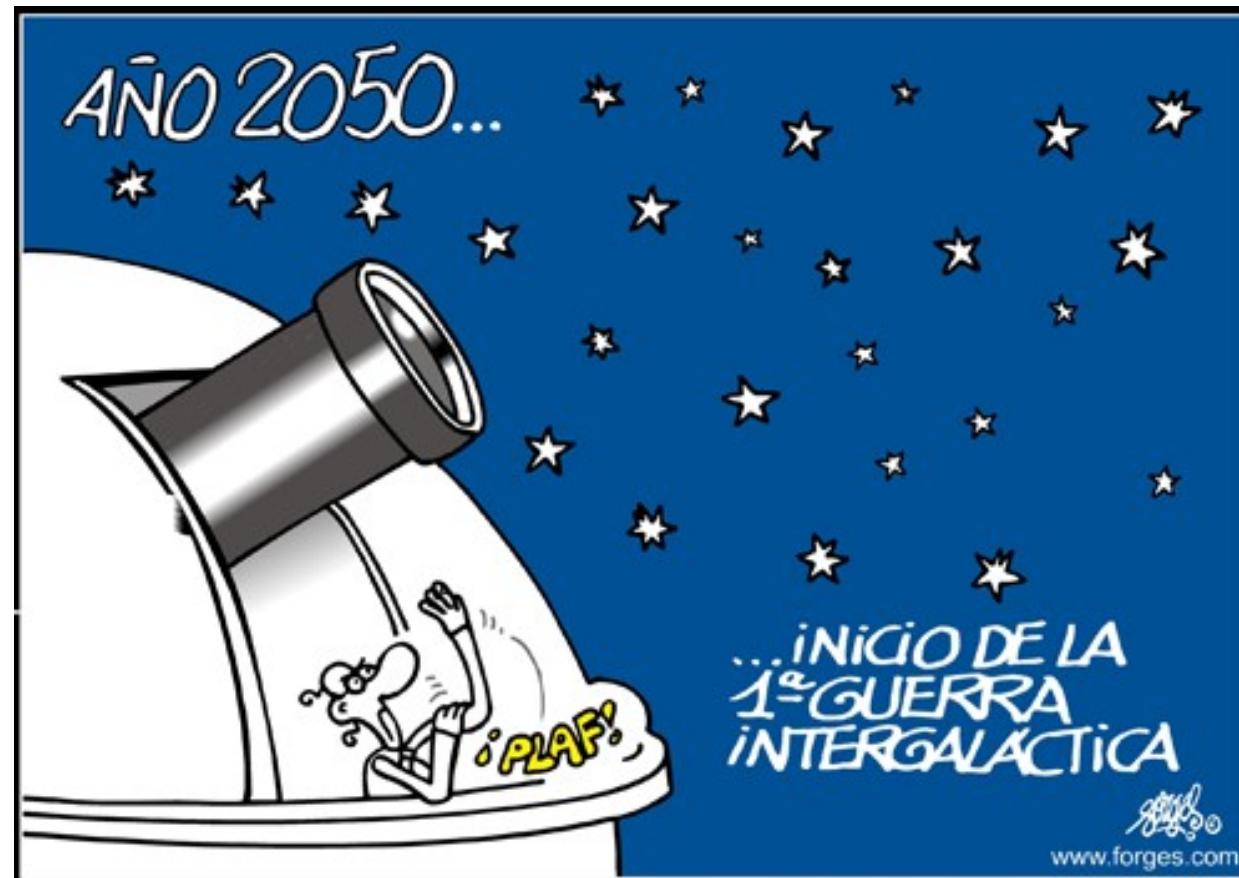
The Big Bang of astronomical data. How to use Python to survive the data flood.



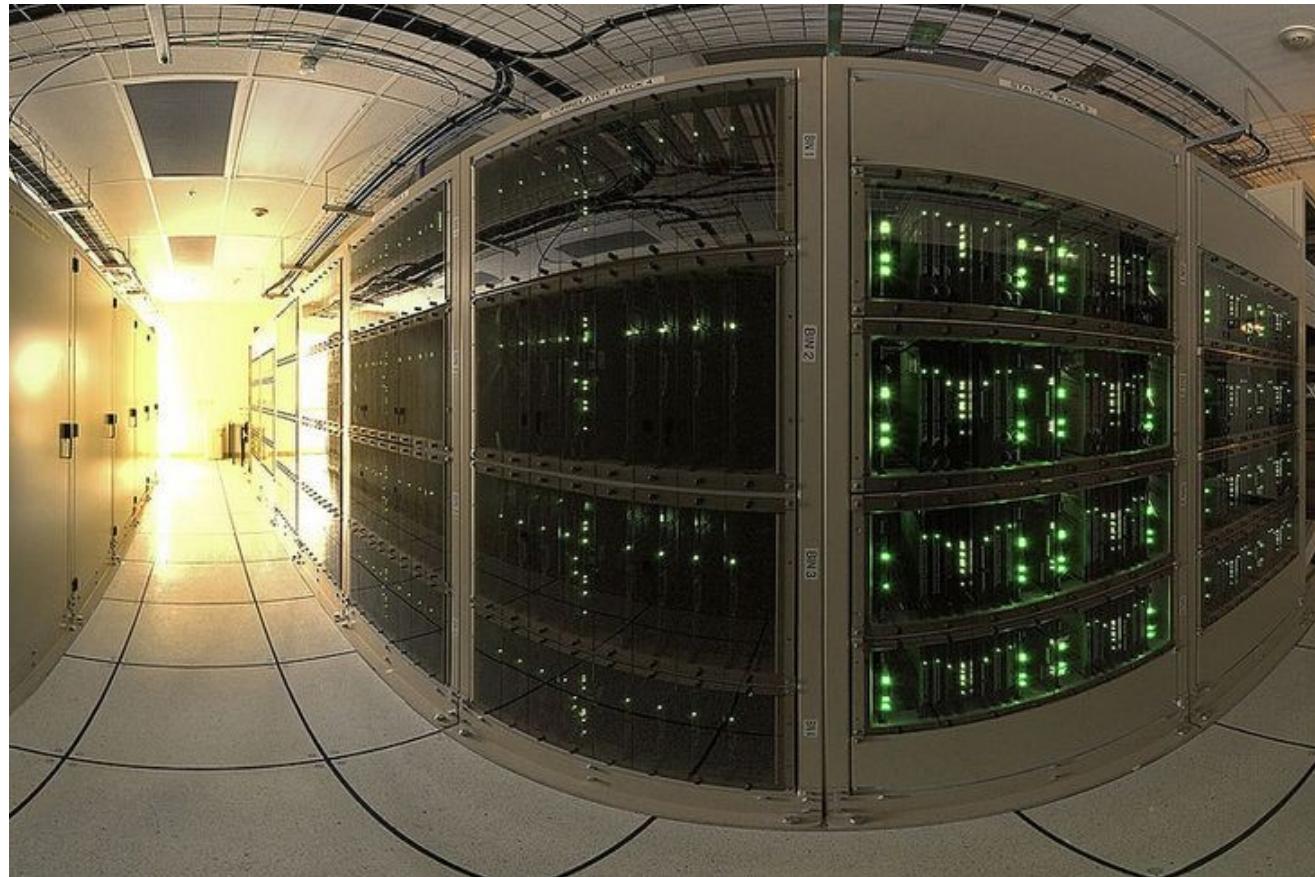
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the LOFAR surveys team

The new astronomy



The new astronomy



ALMA correlator

Astronomy and Python

- Python is currently the main language used for astronomy
- General Python computing libraries: numpy, scipy, matplotlib, pandas, emcee...
- Specific astronomical libraries (see
<http://www.astropython.org/packages/>)
 - astroML: machine learning and data mining
 - astropy: main general library for astronomy
 - etc.

The future of astronomy

- New state of the art astronomical infrastructures that produce an overwhelming amount of data
- Examples:
 - ESA Gaia
 - Large Synoptic Survey Telescope
 - ESA Euclid
 - The Square Kilometre Array and its pathfinders (LOFAR, ASKAP, Meerkat...)
 - Etc.

Large Synoptic Survey Telescope



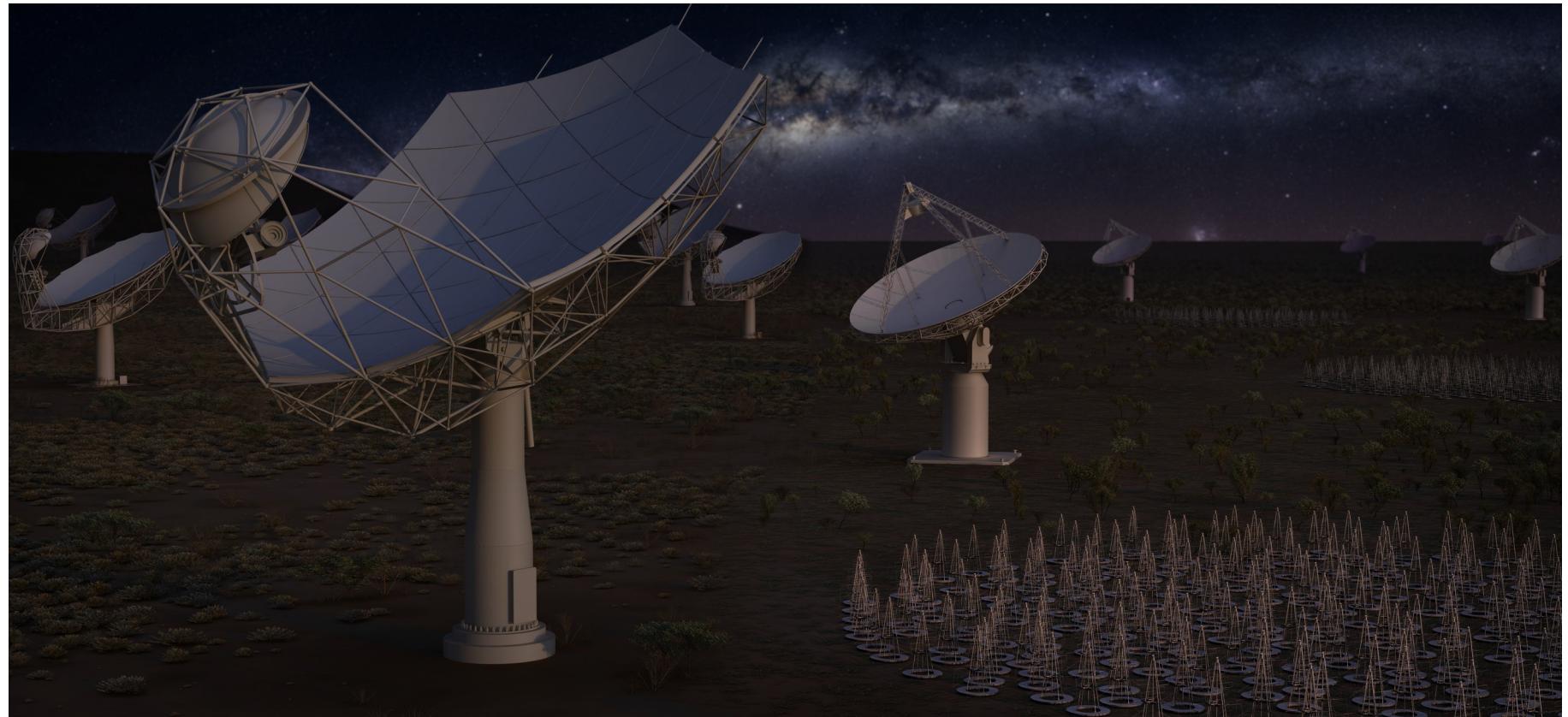
- 8.4 m mirror
- Covers the full visible sky every two nights
- Under construction - operational in 2022

Large Synoptic Survey Telescope



- Camera 189x16 Mpix
- Pipeline preprocessing: 3GB/s
- 30 TB per night during 10 years
- 2 M events triggered per night

Square Kilometre Array (SKA)



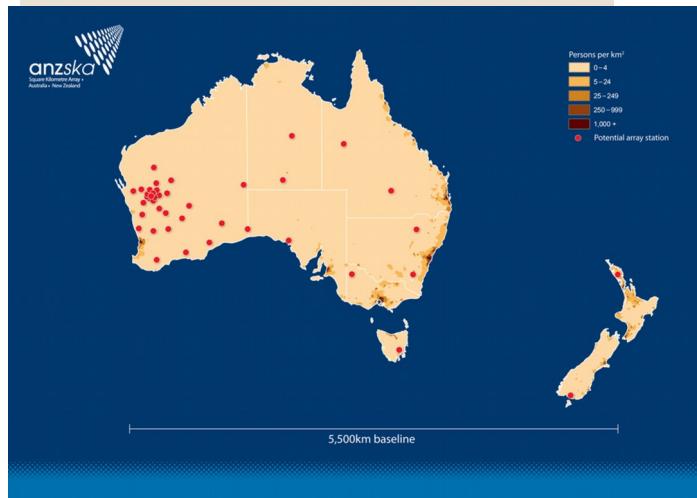
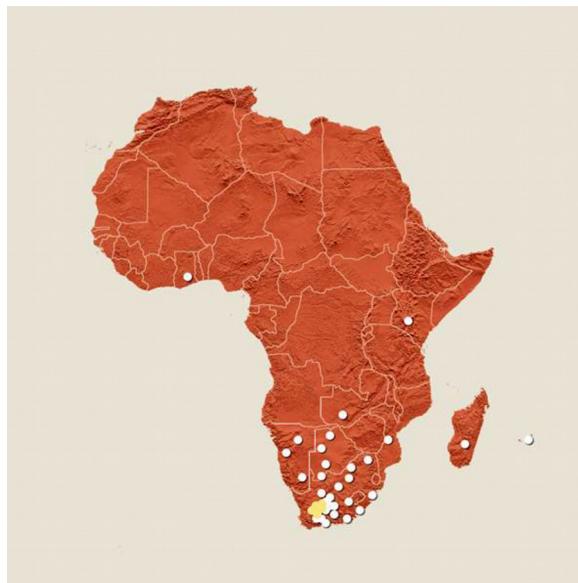
- Radio telescope with 1 km^2 of collecting area
- Phase 1 - 2020

SKA data



- Phase 1:
 - 10 TB/s from the antennas to the correlator
 - 40 GB/s of data → 70 PB per year
 - 1 MW infrastructure and 10 MW processing

SKA data



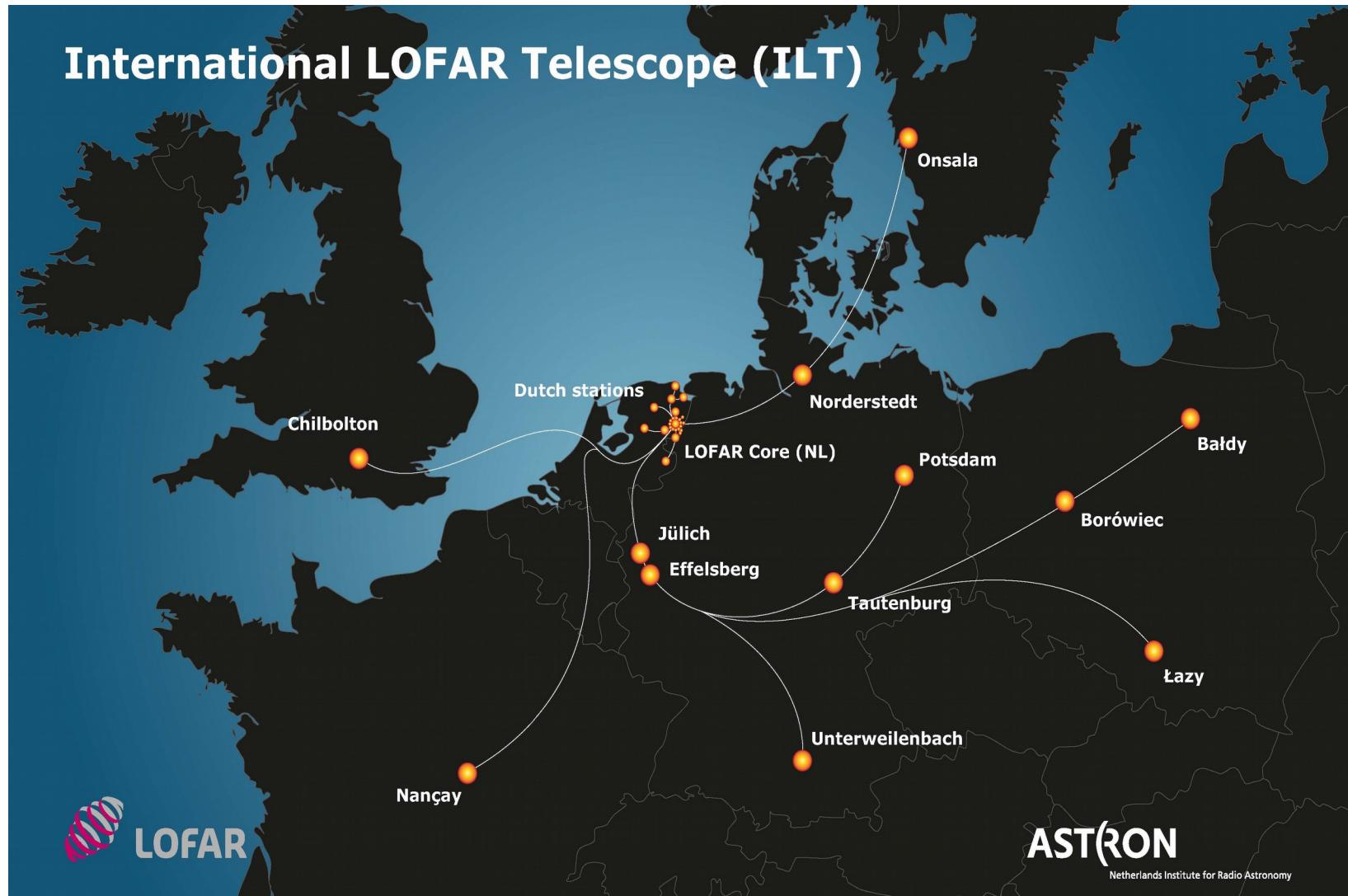
- Phase 2:
 - 160 TB/s from the antennas to the correlator
 - > 100 GB/s of data
→ 4.6 EB per year
 - 200 to 2000 dishes
 - 130K to 1M antennas

LOFAR

- Low Frequency Array
- Software defined radio-interferometer working at low frequencies (30 to 240 MHz)
- One of the Square Kilometre Array pathfinders



LOFAR Stations



LOFAR Stations



LOFAR frequencies

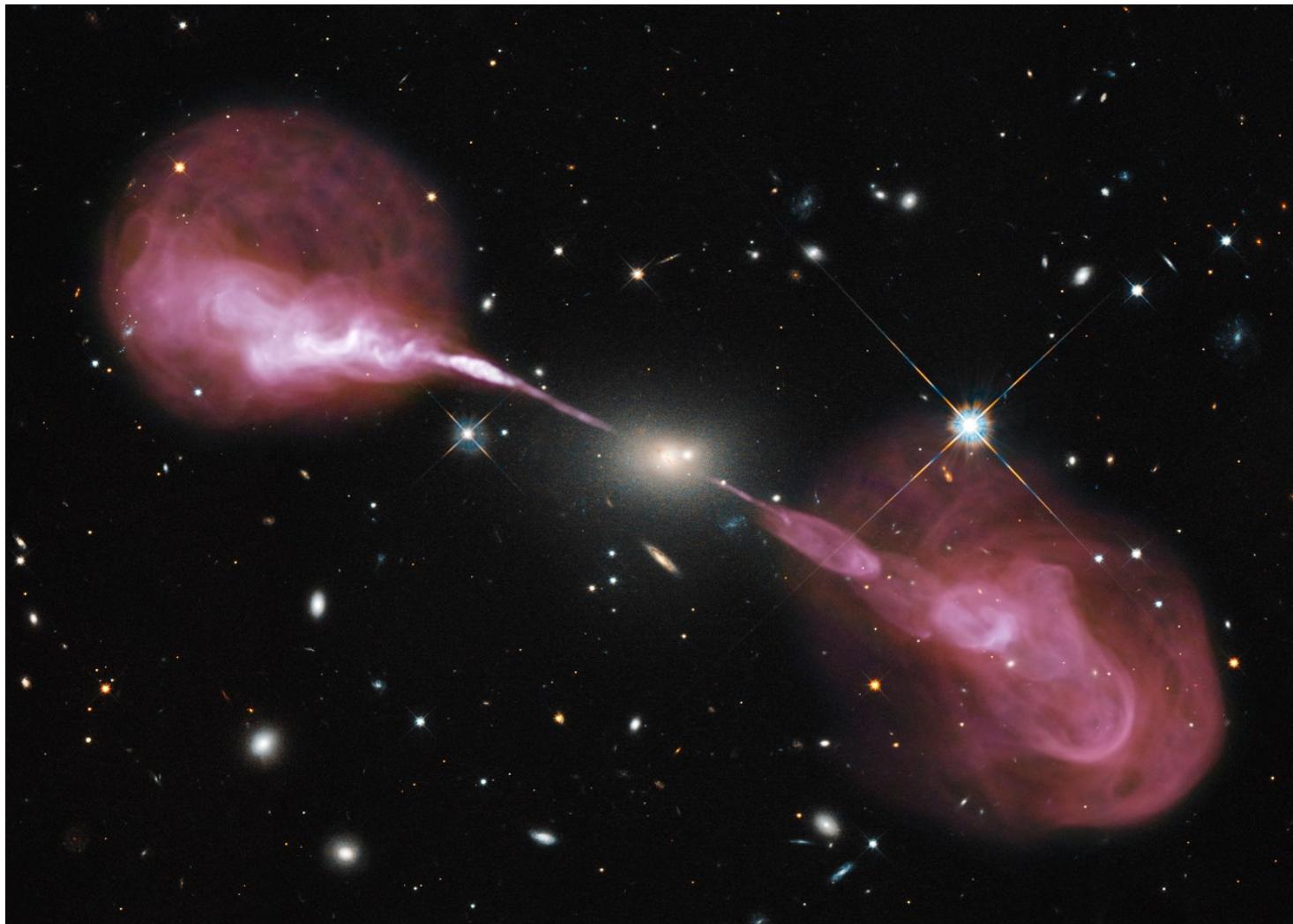
- LBA 30-80 MHz
- HBA 120-240 MHz



LOFAR science

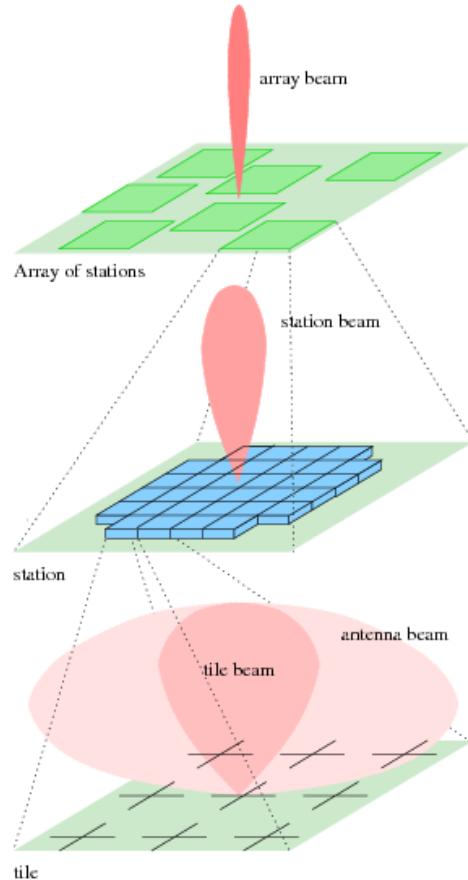
- Origin and evolution of galaxies and supermassive black holes
- Epoch of reionization
- Solar science and space weather
- Transients
- Map the galaxy using pulsars
- Exoplanets, SETI

Radio galaxies



Hercules A. Credits: NASA and the NRAO

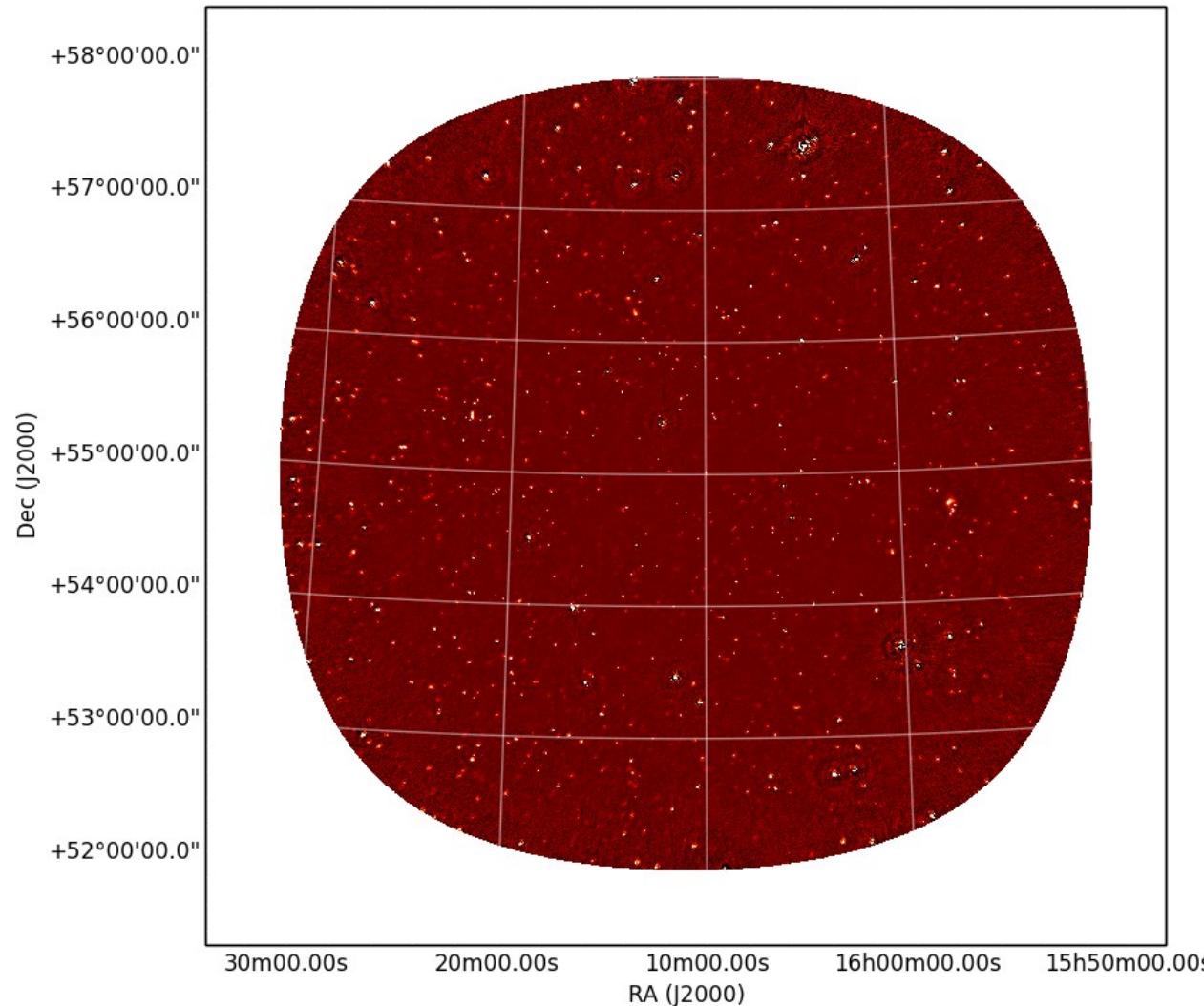
LOFAR aperture synthesis



- field of view diameter of ~5 deg at 150 MHz
- resolution < 5 arcsec (up to 0.1 arcsec)



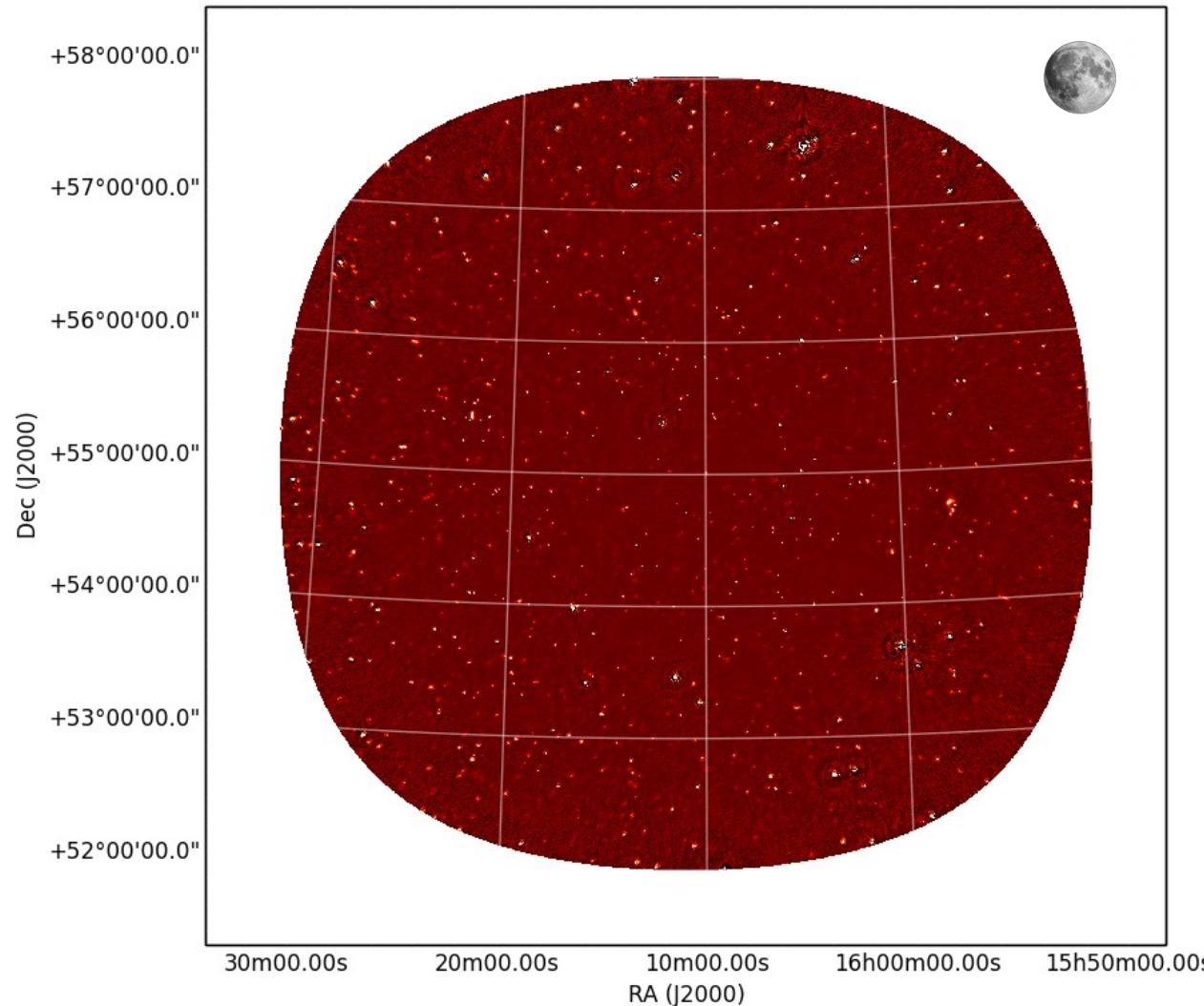
LOFAR imaging



In 8 hours
~40 sq. deg.
5000 sources

Calibration on
IAA (Granada) cluster

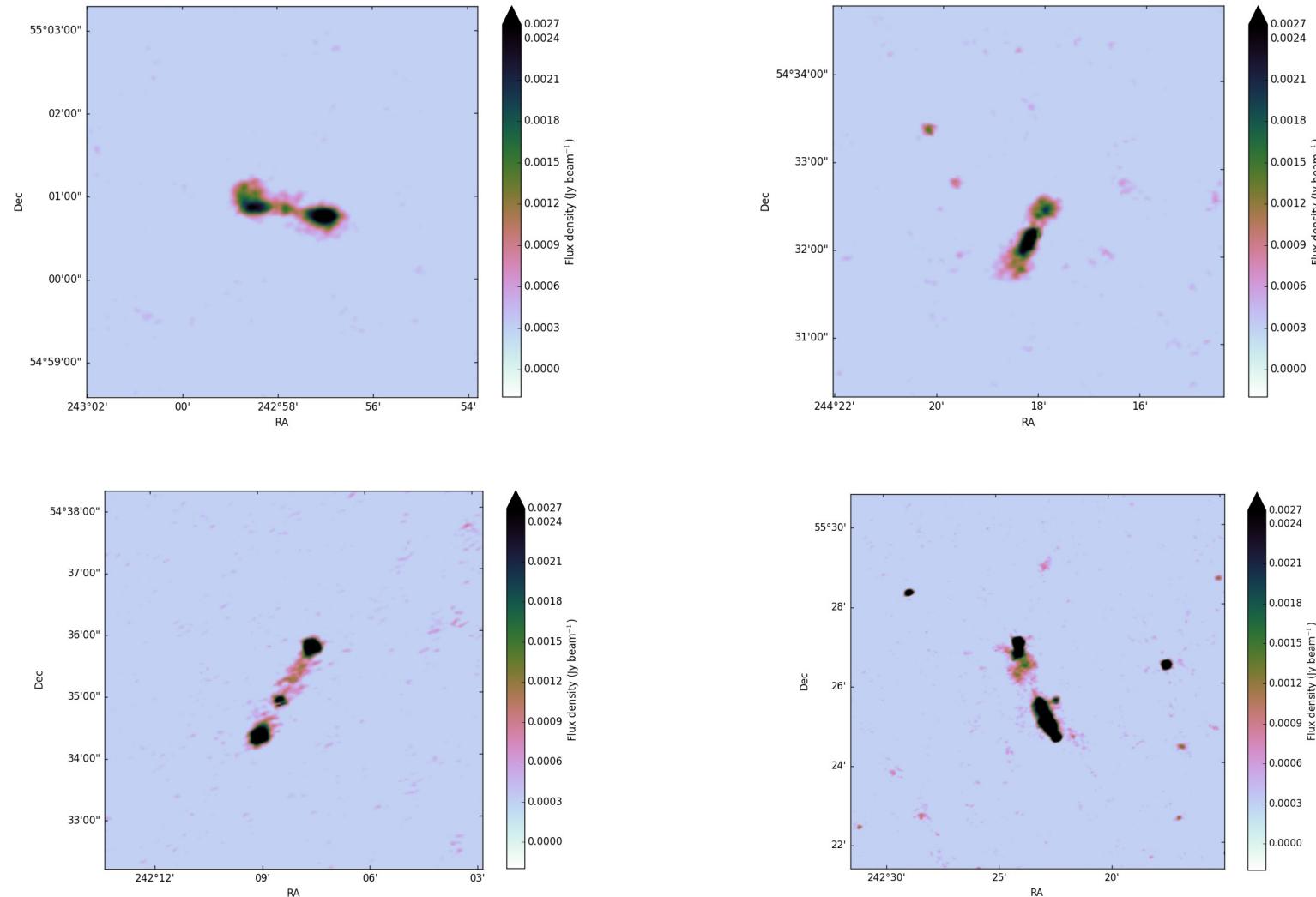
LOFAR imaging



In 8 hours
~40 sq. deg.
5000 sources

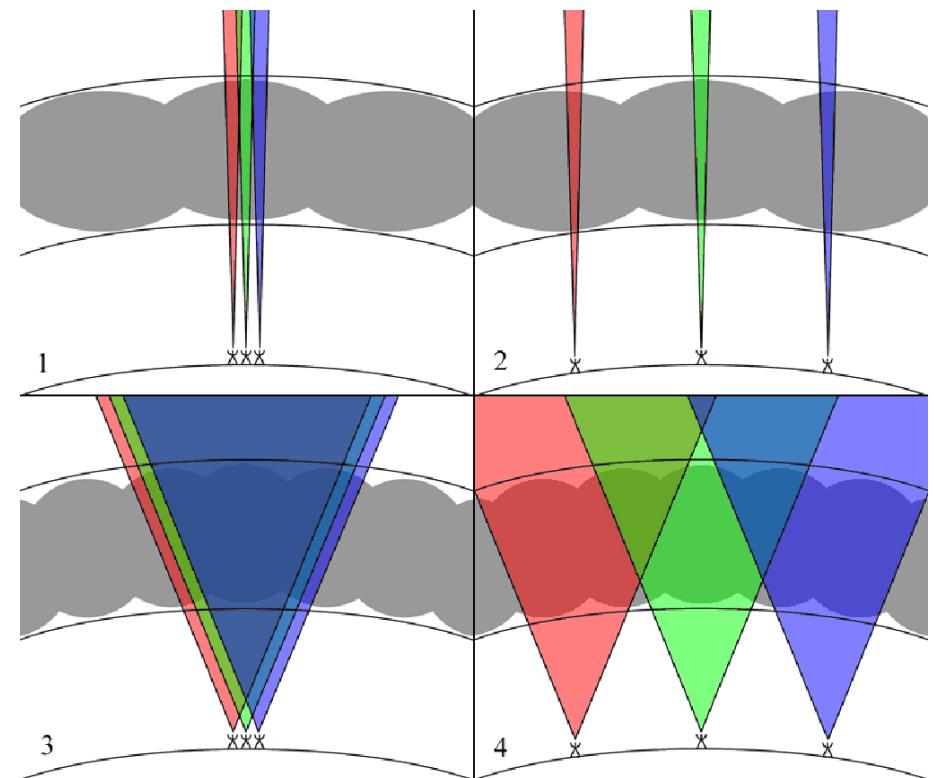
Calibration on
IAA (Granada) cluster

Extended sources



Ionosphere

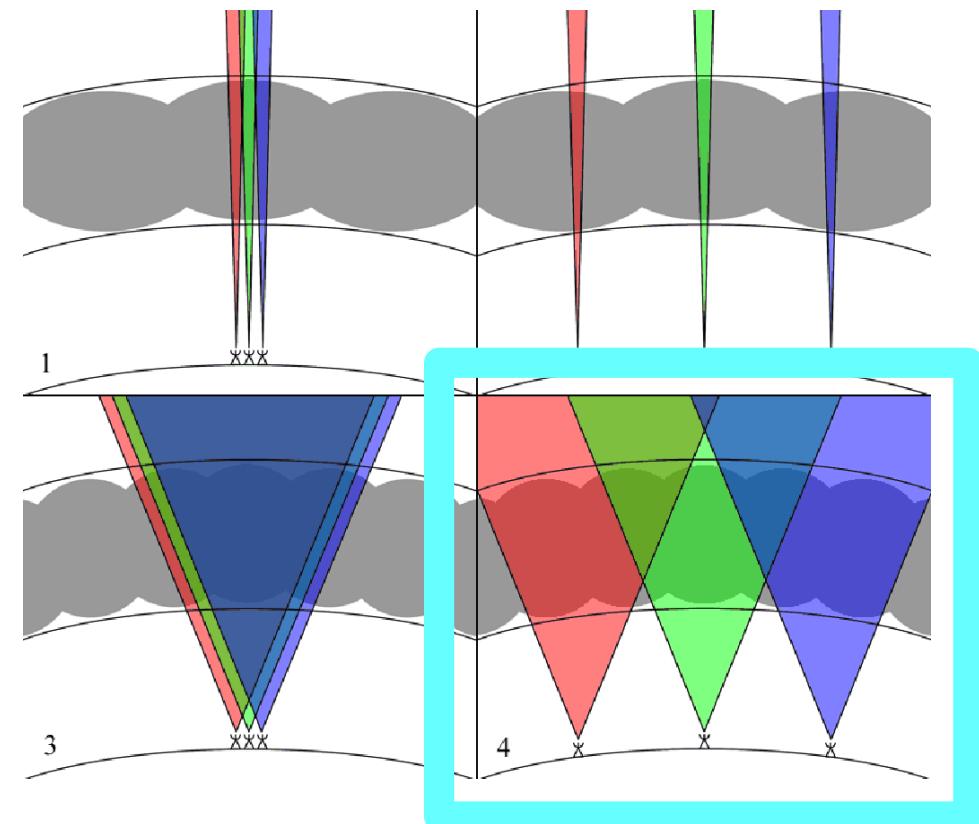
- Effect depends on frequency, length of the baselines and f.o.v.
- LOFAR, worst case:
 - Wide field of view
 - Long distance baselines
 - Low frequency



H. Intema

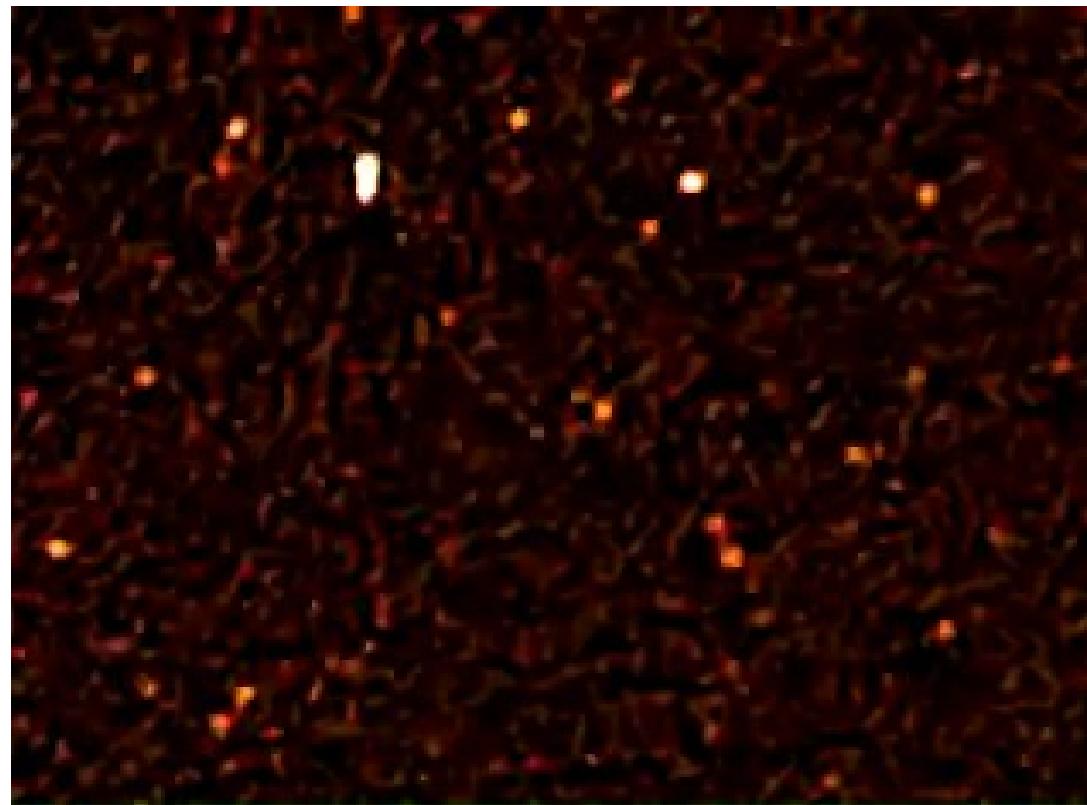
Ionosphere

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H. Intema

Ionosphere



Challenges for the astronomer

- User data calibration (remove the effect of the ionosphere and the RFI)
 - 8 hours full resolution → ~20 TB
 - Minimum of 2 CPU years to run the calibration
 - Experimental pipeline
- LOFAR calibration software
 - Difficult to install
 - Continuous development

Computational solution needed

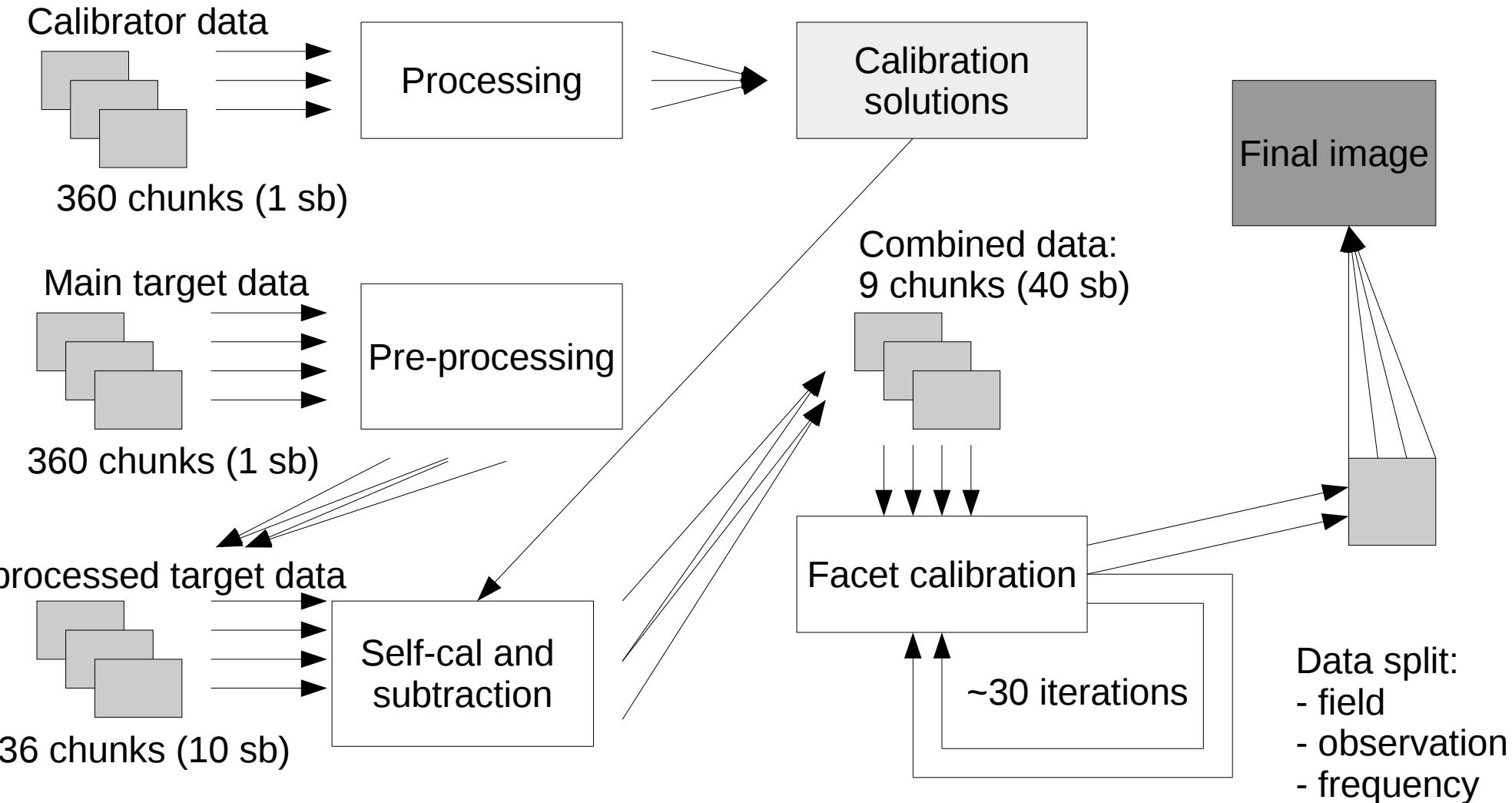
- Parallelizable:
 - Deal with a large amount of data in a reasonable time.
- Flexible:
 - Adapt the infrastructure (“hardware”) to different calibration strategies
 - Deal with quickly changing temperamental software
 - On-demand (optional but very useful)

HPC, HTC and cloud computing

- Tests in different infrastructures: clusters, GRID, cloud, etcetera.
- SKA-AWS astrocompute proposal
 - Preparation of the base infrastructure (virtual machine images, check provisioning of spot instances, etc.)
 - Data transfer: 50 TB
 - Adapt calibration pipeline and run

<http://www.lofarcloud.uk>

Experimental calibration pipeline

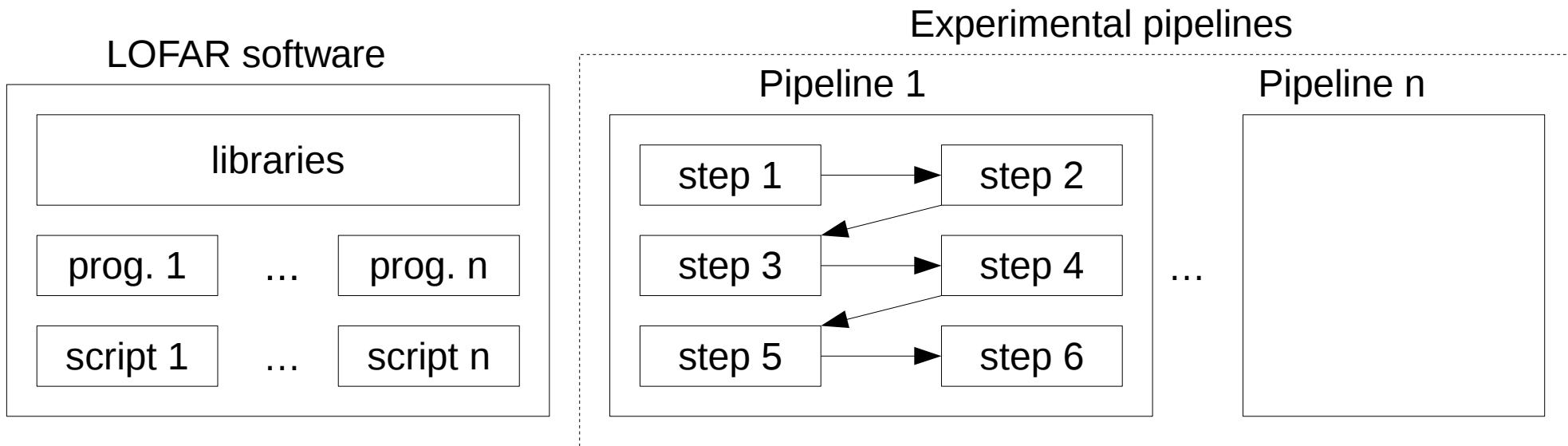


The role of Python

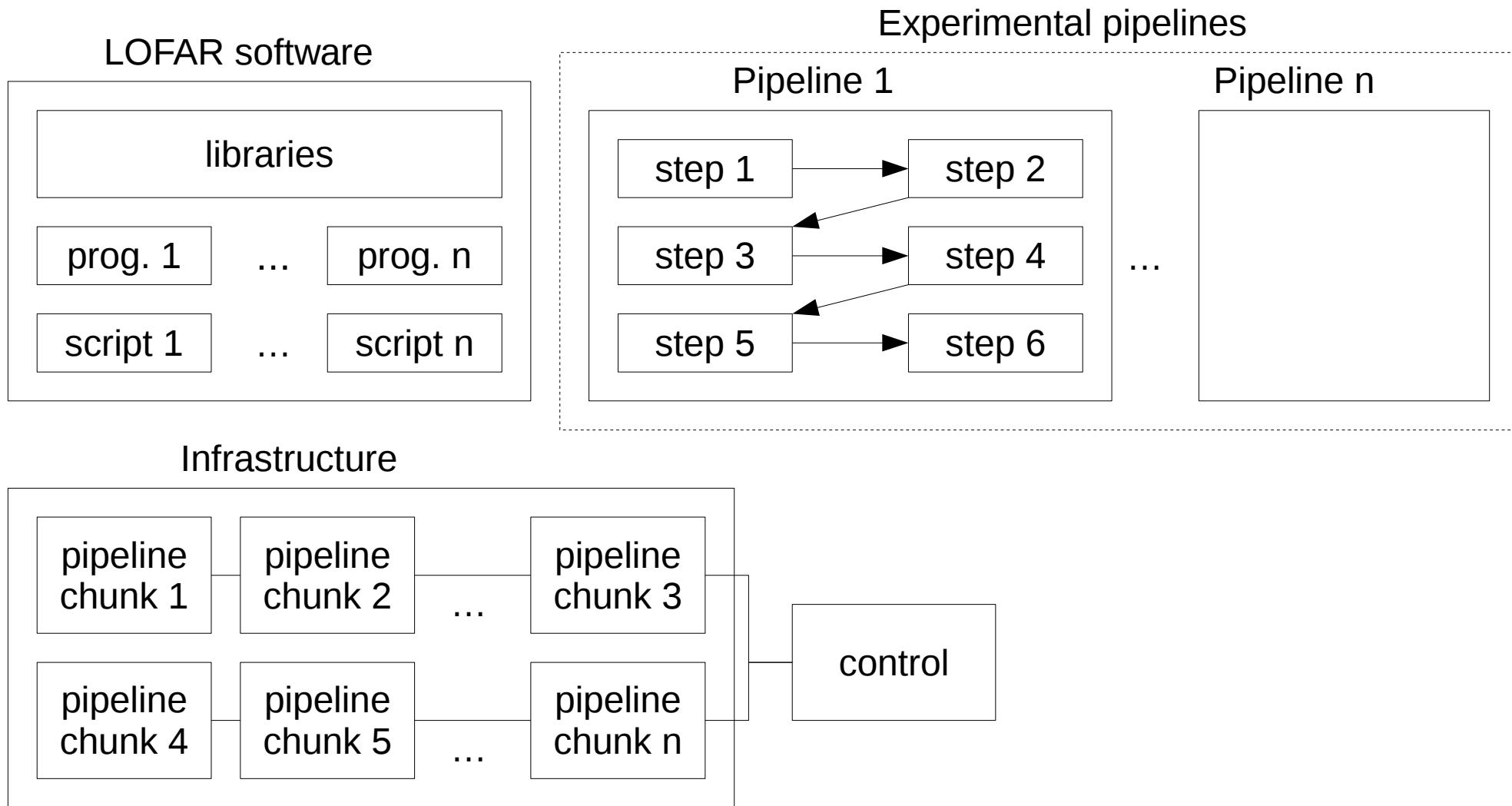
LOFAR software



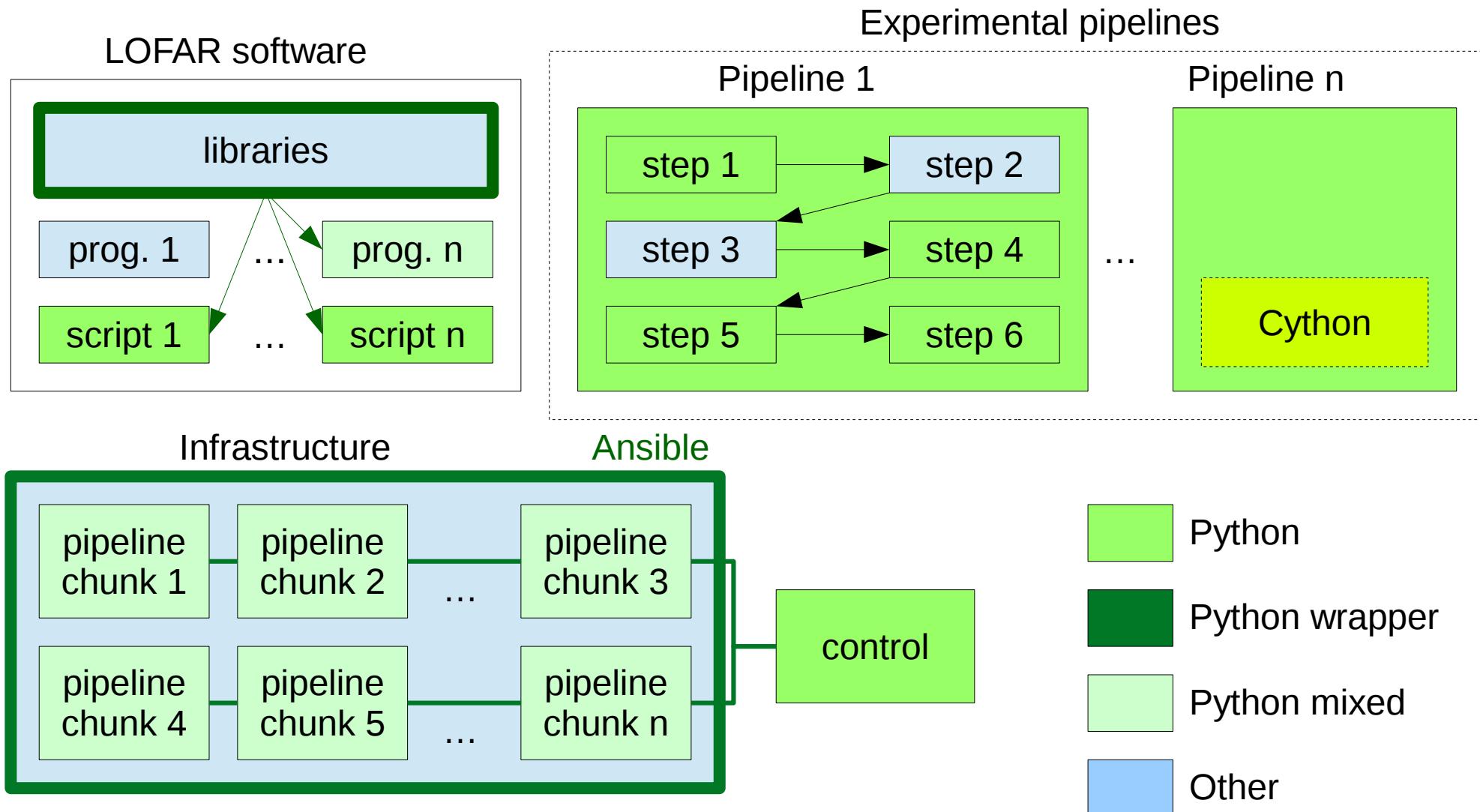
The role of Python



The role of Python



The role of Python



Summary

- Big software and data managing challenges associated to new astronomical infrastructures, even for final users.
- The role of Python:
 - Quick prototyping - fundamental for experimental pipelines and testing.
 - Multi-domain - Can be used for a wide range of problems.
 - Robust - Enough to write “real” efficient software.
 - Unifying tool - that holds all together.

