# A short manual for 2d FLI - two-dimensional Fast Laplace Inversion

## **Schlumberger-Doll Research**

## **April 16, 2003**

2D FLI is an algorithm to perform two-dimensional numercial inverse Laplace transformation, developed at Schlumberger-Doll Research. This algorithm solves the integral equation:

$$M(\underline{\square},\underline{\square}2) = \underline{\square}dT1dT2 \cdot F(T1,T2) \exp(\underline{\square}\square/T1) \exp(\underline{\square}\square2/T2) + Noise$$

where M is the signal as a function of  $\square 1$  and  $\square 2$  and the algorithm solves for F(T1,T2) subject to the non-negativity constraint,  $F \ge 0$ .

Details were published in the following papers:

- 1. L. Venkataramanan et. al., IEEE Tran. SP. 50, 1017-1026 (May, 2002).
- 2. Y.-Q. Song, et al., J. Magn. Reson. 154, 261-268(2002).
- 3. M D Hurlimann and L Venkataramanan, J. Magn. Reson. 157, 31 (2002).
- 4. US Patent 6,462,542 B1, October 8, 2002

This document describes the use of the 2d FLI program.

### How to use the program

## **Requirement:**

Matlab5.3 and above, optimization toolbox

#### Files in the distribution:

FLIv1.0:

example:

example1.mat example1.matRe.mat

FLI.m

FLIdrive.m

FLIEstimate.m

FLIminfun.m

misc:

CalcT1T2r m

ComputeProjections.m

FileRead.m

ReadT1T2.m

## plotting:

PlotData2.m PlotFEstimate.m

PlotFig.m PlotT1T2.m

ReadME.pdf

### Main files:

FLI

This is the main program for the algorithm.

- 1. read the data file with 2d data and parameters,
- 2. Set kernels, range of T1 and T2, digitization, choose regularization methods,
- 3. perform singular value decomposition (SVD) and compress the data,
- 4. optimize for the best F. The resulting density function (F) is stored in matlab variable FEst.

### FLIEstimate.m

Core routine to optimize for F.

On the distribution CD, the compiled versions for Solaris (FLIEstimate.mexsol) and Window (FLIEstimate.dll) are included, instead of the source file. They are compiled for MATLAB 6 release 13.

#### FLIminfun.m

Minimization function, used in FLIEstimate.m and internally by matlab.

# **Example driver file:**

FLIdrive.m

Set the path before running FLI.

# Input data and parameters

**Data**: A matrix of the 2 dimensional input data **NoiseStd**: A variable holding the variance of the noise

**Tau\_1**: A column vector holding the □1 values, with the same length as width of

the matrix **Data**.

**Tau 2**: A column vector holding the t2 values, with the same length as the length

of the matrix **Data**.

Example of the example 1.mat (Output from matlab):

>> wnos		
Name	Size	Bytes Class
Data	2046x20	327360 double array
NoiseStd	1x1	8 double array
Tau_1	20x1	160 double array
Tau 2	2046x1	16368 double array

# **Execution of the program**

In matlab, once the path is set to include the FLI.m file, command FLI will execute the program.

The command will first prompt for an input file. Pressing the return key without an input file will allow the program to analyze the data in the example folder.

The result of the inversion is stored in a file whose name is the input file name plus "Re.mat".

# Main output variables

# The resulting 2D spectrum

FEst	100x101	80800 double array		
The resulting 2D fit to the data				
Fitdata	2046x20	327360 double array		
Kernel functions				
K_1	20x101	16160 double array		
K_2	2046x100	1636800 double array		
Kernel_1	1x1	984 inline object		
Kernel_2	1x1	974 inline object		

# The resulting T1/T2 ratio spectrum

T1T2r	1x100	800 double array
T1T2rdist	100x1	800 double array

# The projected spectra of T1 and T2

T1 1x100 800 double array

T1_dist	1x100	800 double array
T2	1x100	800 double array
T2 dist	1x100	800 double array

# The error in the 180 degree pulse

beta 1x1 8 double array

# Plotting example

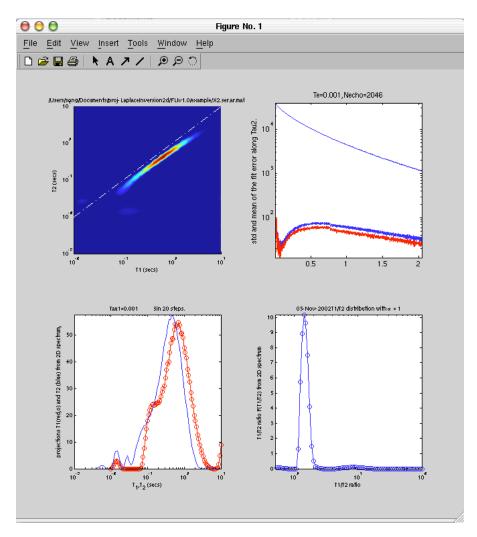
Command: PlotT1T2

Panel 1: the 2D spectrum, FEst, the line is for T1=T2.

Panel 2: fit error estimate: top blue line is the data for the shortest tau\_1; the lower blue line variance of the fit error; the lower red line is the mean of the fit error.

Panel 3: projected spectra of T1 and T2. Title shows the range of Tau\_1.

Panel 4: projected spectrum of T1/T2 ratio.



### **User Modifications of FLI**

#### **Modification to FLI.m**

## 1. Modification for different regularization parameter and schemes

Regularization (alpha) can affect the final inversion result and its value depends on the signal-to-noise ratio of the data. For example, a small alpha can be used with data of high S/N to obtain details of the spectrum. However, when S/N is low, a larger alpha should be used often resulting in a smoother spectrum. In practice, one should perform the inversion with several values of alpha in order to gain confidence in the features observed in the inversion results.

The regularization is set by the following in FLI.m:

Alpha\_Auto = 0: fixed regularization, and the value is AlphaStart. Alpha\_Auto = 1: automatic regularization using BRD method. For this

> method to work well, the Data should be free from systematic errors and NoiseStd reflect the variance of the

random noise.

Alpha Auto=2: A series of regularization will be used to calculate the

inversion. Then the heel of the fit error (c2) will be found

to be the optimal regularization.

### 2. Modifications for different range of T1 and T2

These following variables should be set to reflect the range of T1 and T2 for the specific data set for optimal inversion result. For example, the minimum T1 (InitTime\_T1) should not be smaller than the smallest Tau\_1 and the maximum T1 (FinalTime\_T1) should not be much larger than the maximum Tau\_1. The similar rule should apply to T2.

The number of points for T1 and T2 should not be much smaller than 100. For example, Number\_T1=20 might be too small to allow a good fit for some data.

The ranges of T1 and T2 are set by the following in FLI.m:

InitTime\_T1 = .01; % The initial value of T1 (in seconds)
FinalTime\_T1 = 10; % The final value of T1 (in seconds)
Number\_T1 = 100; % The number of T1's

InitTime\_T2 = .001; % The initial value of T2 (in seconds)
FinalTime\_T2 = 10; % The final value of T2(in seconds)
Number T2 = 100; % The number of T2's

### 3. Modifications for DC offset

The non-ideal inversion of the 180 degree pulse and the rf inhomogeneity would result in a DC offset. This offset can be estimated by adding a column in K\_1. This feature is enabled by:

```
AllowDCOffset = 1; % allow a dc offset % AllowDCOffset = 0; % no offset
```

### 4. Modifications for different kernels

The inversion kernels can be modified to suit the specific data. The kernels are defined as inline functions and calculated later.

```
%exponential decay
%Kernel_1 = inline('exp(- Tau * (1./ TimeConst))','Tau','TimeConst');
%inv recovery
Kernel_1 = inline('1- 2*exp( - Tau * (1./ TimeConst))','Tau','TimeConst');
```

## 5. Modification to neglect the non-negativity constraint

It is not implemented in the current software.

The principle of the modification is discussed in the reference 1 (equation 36).

## Modification to FLIestimate.m, FLIminfun.m

User should not modify these files.

## Modification to files in the directory misc

User should not modify these files.

## Modification to files in the directory plotting

The files in those directories are provided as examples. Users may use them as is or as templates for their own routines to display the inversion results.

### **Contact information**

L. Venkataramanan, lvenkataramanan@ridgefield.oilfield.slb.com M. D. Hurlimann, mhurlimann@ridgefield.oilfield.slb.com

Y.-Q. Song, ysong@slb.com

Schlumberger-Doll Research, Old quarry road, Ridgefield, CT 06877