The cylinder at spatial infinity and asymptotic charges

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Introduction

- The Newman-Penrose (NP) constants serve as conserved quantities at null infinity in asymptotically flat gravitational fields.
- These constants present a comprehensive conservation system for various spins: spin-1 fields and spin-2 fields, with our research focusing on spin-0 fields linked to wave equation solutions.
- In the detailed context, while an infinite series of conserved quantities is identified in the linear theory, the non-linear General Relativity theory conserves only ten.

Conservation laws

• These charges are computed as 2-surface integrals at cuts $C \approx \mathbb{S}^2$ of null infinity \mathscr{I} .

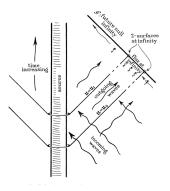


Figure: Visual representation of the behavior of the Newman-Penrose constants at null infinity.

The i^0 cylinder representation in Minkowski spacetime

- Consider spherical polar coordinates $(\tilde{t}, \tilde{\rho}, \vartheta^A)$ with A = 1, 2.
- ullet The metric of physical Minkowski spacetime in this coordinate system is given by $ilde{\eta}$:

$$\tilde{\boldsymbol{\eta}} = -\mathbf{d}\tilde{t} \otimes \mathbf{d}\tilde{t} + \mathbf{d}\tilde{\rho} \otimes \mathbf{d}\tilde{\rho} + \tilde{\rho}^2 \sigma. \tag{1}$$

• Introduce unphysical spherical polar coordinates (t, ρ, ϑ^A) as an intermediate step:

$$t = \frac{\tilde{t}}{\tilde{\rho}^2 - \tilde{t}^2}, \quad \rho = \frac{\tilde{\rho}}{\tilde{\rho}^2 - \tilde{t}^2}.$$
 (2)

ullet The conformal metric in unphysical coordinates, $oldsymbol{\eta}=\Xi^2oldsymbol{ ilde{\eta}}$:

$$\eta = -\frac{1}{\tilde{\rho}^2 - \tilde{t}^2} (-\mathbf{d}\tilde{t} \otimes \mathbf{d}\tilde{t} + \mathbf{d}\tilde{\rho} \otimes \mathbf{d}\tilde{\rho} + \tilde{\rho}^2 \sigma). \tag{3}$$

i⁰ Representation in Unphysical Coordinates

- In this conformal representation $(t \in (-\infty, \infty), \rho \in [0, \infty))$, spatial infinity and the origin interchange.
- i^0 is represented by the point $(t=0, \rho=0)$ in (\mathbb{R}^4, η) .
- Introduce coordinates $(\tau, \rho, \vartheta^A)$ with $t = \rho \tau$.
- Consider the conformal metric $\mathbf{g} = \rho^{-2} \boldsymbol{\eta}$.
- ullet Express the unphysical metric $oldsymbol{g}$ in F-coordinates:

$$\mathbf{g} = -\mathbf{d}\tau \otimes \mathbf{d}\tau + \frac{1 - \tau^2}{\rho^2} \mathbf{d}\rho \otimes \mathbf{d}\rho - \frac{\tau}{\rho} \left(\mathbf{d}\rho \otimes \mathbf{d}\tau + \mathbf{d}\rho \otimes \mathbf{d}\tau \right) + \sigma. \tag{4}$$

Conformal Factor and Lorentz Transformation

• The conformal factor Θ in F-coordinates and physical coordinates:

$$\Theta := \rho(1 - \tau^2) = \frac{1}{\tilde{\rho}}.\tag{5}$$

• The boost parameter κ :

$$\kappa := \frac{1+\tau}{1-\tau} = -\frac{\tilde{v}}{\tilde{u}}.\tag{6}$$

• The Lorentz transformation that connects the NP and F-frames:

$$(\Lambda_+)^2 := \Theta^{-1} \kappa^{-1}, \quad (\Lambda_-)^2 := \Theta^{-1} \kappa.$$
 (7)

i⁰ Cylinder and Null Frames

Identify future and past null infinity in the conformal representation:

$$\begin{split} \mathscr{I}^+ &\equiv \{ p \in \mathcal{M} \mid \tau(p) = 1 \}, \\ \mathscr{I}^- &\equiv \{ p \in \mathcal{M} \mid \tau(p) = -1 \}. \end{split}$$

• The i^0 -cylinder represents spatial infinity as an extended set $\mathcal{I} \approx \mathbb{R} \times \mathbb{S}^2$:

$$\mathcal{I} \equiv \{ p \in \mathcal{M} \mid |\tau(p)| \ 1, \ \rho(p) = 0 \},$$

$$\mathcal{I}^0 \equiv \{ p \in \mathcal{M} \mid \tau(p) = 0, \ \rho(p) = 0 \}.$$

F-Frame and Null Frames

• Introduce the *F*-frame:

$$m{e} = (1+\tau) m{\partial}_{\tau} - \rho m{\partial}_{\rho}, \quad \underline{\pmb{e}} = (1-\tau) m{\partial}_{\tau} + \rho m{\partial}_{\rho}, \quad \pmb{e}_{\mathbf{A}} \quad \text{with}$$
 $m{A} = \{\uparrow, \downarrow\}.$ (8)

• The NP-frame hinged at \mathscr{I}^{\pm} :

NP hinged at
$$\mathscr{I}^+$$
: $e^+, \underline{e}^+, e_A^+$,
NP hinged at \mathscr{I}^- : $e^-, \underline{e}^-, e_A^-$.

Transformation between NP and F-frames:

$$\begin{array}{lll} \textit{NP hinged at } \mathscr{I}^+: & \pmb{e}^+ = \Theta^{-2} \textit{L}, & \underline{\pmb{e}}^+ = \underline{\textit{L}}, & \pmb{e}_{\pmb{A}}^+ = \pmb{e}_{\pmb{A}} = \Theta^{-1} \tilde{\pmb{e}}_{\pmb{A}} \\ & \textit{NP hinged at } \mathscr{I}^-: & \pmb{e}^- = \textit{L}, & \underline{\pmb{e}}^- = \Theta^{-2} \underline{\textit{L}}, & \pmb{e}_{\pmb{A}}^- = \pmb{e}_{\pmb{A}} = \Theta^{-1} \tilde{\pmb{e}}_{\pmb{A}}. \\ \end{array}$$

• For $(\tilde{M}, \tilde{\mathbf{g}})$ and (M, \mathbf{g}) , the transformation of the D'Alembertian operator under conformal transformations is,

$$\Box \phi - \frac{1}{6} \phi R = \Omega^{-3} \left(\tilde{\Box} \tilde{\phi} - \frac{1}{6} \tilde{\phi} \tilde{R} \right). \tag{9}$$

• Using F-coordinates, the wave equation is represented by

$$(\tau^2 - 1) \partial_{\tau}^2 \phi - 2\rho \tau \partial_{\tau} \partial_{\rho} \phi + \rho^2 \partial_{\rho}^2 \phi + 2\tau \partial_{\tau} \phi + \Delta_{S^2} \phi = 0.$$
 (10)

We consider the Ansatz

$$\phi = \sum_{p=0}^{\infty} \sum_{\ell=0}^{p} \sum_{m=-\ell}^{m=\ell} \frac{1}{p!} a_{p;\ell,m}(\tau) \rho^p Y_{\ell m}.$$
 (11)

• Solving (10) simplifies to solving the following (ODE) for every p, ℓ , and m:

$$(1-\tau^2)\ddot{a}_{p;\ell,m} + 2\tau(p-1)\dot{a}_{p,\ell,m} + (\ell+p)(\ell-p+1)a_{p;\ell,m} = 0.$$
 (12)

Lemma

The solution to equation (12) is given by:

• For $p \ge 1$ and $0 \le \ell \le p-1$

$$a(\tau)_{\rho;\ell,m} = A_{\rho,\ell,m} \left(\frac{1-\tau}{2}\right)^{\rho} P_{\ell}^{(\rho,-\rho)}(\tau) + B_{\rho,\ell,m} \left(\frac{1+\tau}{2}\right)^{\rho} P_{\ell}^{(-\rho,\rho)}(\tau)$$
(13)

② For $p \ge 0$ and $\ell = p$:

$$a_{p;p,m}(\tau) = \left(\frac{1-\tau}{2}\right)^p \left(\frac{1+\tau}{2}\right)^p \left(C_{p,p,m} + D_{p,p,m} \int_0^\tau \frac{ds}{(1-s^2)^{p+1}}\right) \tag{14}$$

• When analyzing the hypergeometric function presented in Equation (14) for different values of p we can see the emergence of logarithmic terms. Let's consider the cases of p=0 and p=1, which yield the following expressions:

$$\begin{aligned} a_{0;0,0}(\tau) &= C_{000} + \frac{1}{2} D_{000}(\log(1+\tau) - \log(1-\tau)) \\ a_{1;1,m}(\tau) &= \frac{1}{4} (1-\tau)(1+\tau) \left(C_{11m} + \frac{1}{4} D_{11m}(\log(1+\tau) - \log(1-\tau) + 2\tau(1-\tau^2)) \right). \end{aligned} \tag{15}$$

 These logarithmic terms have implications for the linear version of the associated peeling property.

Remark

(Regularity condition). Lemma 1 implies that expanding the integral in (14) results in logarithmic terms, hence $D_{p,p,m}=0$ is called the regularity condition. The solutions for $a(\tau)$ are polynomic in τ , except for $\ell=p$ where one needs to impose the regularity condition to only have polynomic solutions.

• We can see how peeling is violated for the spin-0 field by expanding $\tilde{\phi}$ in terms of the *F*-frame. Making use of eqs (2.34), (2.35) and *Lemma* 1, we truncate (11) up tp p=1 and get the following expression:

$$\tilde{\phi} = \frac{C_{000}}{\tilde{\rho}} + \frac{1}{2\tilde{\rho}} D_{000} \log \left(\frac{\tilde{\rho} + \tilde{t}}{\tilde{\rho} - \tilde{t}} \right) Y_{00} + \frac{1}{8\tilde{\rho}} \left(1 - \frac{\tilde{t}^2}{\tilde{\rho}^2} \right) \\
\left[D_{11-1} \log \left(\frac{\tilde{\rho} + \tilde{t}}{\tilde{\rho} - \tilde{t}} \right) Y_{1-1} + D_{110} \log \left(\frac{\tilde{\rho} + \tilde{t}}{\tilde{\rho} - \tilde{t}} \right) Y_{10} \right] \\
+ \left[D_{111} \log \left(\frac{\tilde{\rho} + \tilde{t}}{\tilde{\rho} - \tilde{t}} \right) Y_{11} \right] + \frac{1}{2\tilde{\rho}} \left(\tilde{t}^2 - \tilde{\rho}^2 \right) (A_{100} + B_{100}) Y_{00} + \\
\frac{1}{4\tilde{\rho}} \left(1 - \frac{\tilde{t}^2}{\tilde{\rho}^2} \right) (C_{11-1} Y_{1-1} + C_{110} Y_{10} + C_{111} Y_{11}) . \tag{17}$$

• The expantion of the physical field $\tilde{\phi}$ helps to understand the concept of peeling in the spin-0 case. In the spin-0 case, the peeling property is violated by the logarithmic terms that appear in the expansion of $\tilde{\phi}$.

The NP-constants for the spin-0 fields close to $i^0 \& \mathscr{I}$

The NP constants emerge from a set of asymptotic conservation laws.
 For the spin-0 field in flat spacetime one has:

$$\underline{L}(\tilde{\rho}^{-2\ell}L(e^+)^{\ell+1}\phi_{\ell m}) = 0, \qquad L(\tilde{\rho}^{-2\ell}\underline{L}(e^-)^{\ell+1}\phi_{\ell m}) = 0 \qquad (18)$$

here, $\phi_{\ell m} = \int_{\mathbb{S}^2} \phi \, Y_{\ell m} d\sigma$ represents the integral of ϕ multiplied by the spherical harmonics $Y_{\ell m}$ over the surface \mathbb{S}^2 , where $d\sigma$ denotes the area element on \mathbb{S}^2 .

• One can introduce the $f(\tilde{\rho})$ -modified NP constants in the following manner:

$${}^{f}\mathcal{N}_{\ell,m}^{+} := f(\tilde{\rho})L(\boldsymbol{e}^{+})^{\ell}\phi_{\ell m}|_{C^{+}}, \tag{19}$$

$${}^{f}\mathcal{N}_{\ell,m}^{-} := f(\tilde{\rho})\underline{L}(\underline{e}^{-})^{\ell}\phi_{\ell m}|_{C^{-}}.$$
 (20)

here, $C^{\pm} \approx \mathbb{S}^2$ represents a cut of \mathscr{I}^{\pm} .

• In the specific case of $f(\tilde{\rho}) = \tilde{\rho}^2$, these quantities are referred to as the "classical NP-constants" and are denoted as $\mathcal{N}_{\ell,m}^{\pm}$. They can be succinctly expressed as follows:

$$\mathcal{N}_{\ell,m}^{+} := (\boldsymbol{e}^{+})^{\ell+1} \phi_{\ell m} |_{C^{+}}, \tag{21}$$

$$\mathcal{N}_{\ell,m}^{-} := (\underline{\underline{e}}^{-})^{\ell+1} \phi_{\ell m} |_{C^{-}}. \tag{22}$$

In the notation for ${}^f\mathcal{N}^+_{\ell,m}$ presented below, it will be implicitly assumed that m takes values from $-\ell$ to ℓ .

The classical NP constants at \mathscr{I}^+

• This analysis is facilitated by the expression,

$$\phi_{\ell m} = \sum_{p=\ell}^{\infty} \frac{1}{p!} a_{p;\ell,m}(\tau) \rho^p. \tag{23}$$

• Considering $\ell=0$, the computation of ${\boldsymbol e}^+(\phi_{00})$ is sufficient.

$$\mathbf{e}^{+}(\phi_{\ell m}) = 4\rho^{-1}(1+\tau)^{-2} \sum_{p=0}^{\infty} \frac{1}{p!} \rho^{p}((1+\tau)\dot{a}_{p;\ell,m} - pa_{p;\ell,m}). \quad (24)$$

With

$$Q_{p;\ell,m}^{0}(\tau) := (1+\tau)\dot{a}_{p;\ell,m} - pa_{p;\ell,m}. \tag{25}$$

ullet With this definition in place, we can express $m{e}^+(\phi_{\ell m})$ as follows:

$$\mathbf{e}^{+}(\phi_{\ell m}) = 4(\Lambda_{+})^{2} \sum_{p=0}^{\infty} \frac{1}{p!} \rho^{p} Q_{p,\ell,m}^{0}(\tau).$$
 (26)

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• To compute the $\ell = 0$ NP constant at \mathscr{I}^+ , it is necessary to evaluate $e^+(\phi_{00})$ at a specific cut C^+ of \mathscr{I}^+ . By utilizing equation (26) we obtain the following expression:

$$\mathcal{N}_{0,0}^{+} = \lim_{\substack{\rho \to \rho_{\star} \\ \tau \to 1}} \mathbf{e}^{+}(\phi_{00}) = \sum_{p=0}^{\infty} \frac{1}{p!} \rho_{\star}^{p-1} Q_{p,0,0}^{0}|_{\mathscr{I}^{+}} = -A_{100}.$$
 (27)

• To calculate the $\ell=1$ NP constants, we must evaluate $(e^+)^2(\phi_{1m})$.

$$(e^{+})^{2}(\phi_{\ell m}) = 4^{2}(\Lambda_{+})^{2} e((\Lambda_{+})^{2} \sum_{p=\ell}^{\infty} \frac{1}{p!} \rho^{p} Q_{p,\ell,m}^{0}(\tau)),$$
 (28)

where

$$\boldsymbol{e}\left(\Lambda_{+}^{2}\right) = -\Lambda_{+}^{2}.\tag{29}$$

Utilizing the aforementioned equations yields:

$$(\mathbf{e}^{+})^{2}(\phi_{\ell m}) = 4^{2}(\Lambda_{+})^{4} \sum_{p=\ell}^{\infty} \frac{1}{p!} \rho^{p} ((1+\tau) \dot{Q}_{p;\ell,m}^{0}(\tau) - (p+1) Q_{p;\ell,m}^{0}(\tau))$$

Where

$$Q^1_{\rho,\ell,m}(\tau) := \left((1+\tau) \, \dot{Q}^0_{\rho,\ell,m}(\tau) - (\rho+1) \, Q^0_{\rho,\ell,m}(\tau) \right). \tag{31}$$

Finally, one has:

$$\mathcal{N}_{1,m}^{+} = \lim_{\substack{\rho \to 0 \\ \tau \to 1}} 2^{-4} \frac{1}{2!} Q_{2,1,m}^{1} \cdot (\tau) = 3A_{21m}. \tag{32}$$

• The NP constants for \mathscr{I}^- can be calculated in a similar manner, where the time reversed version of the F-frame is used.

The i^0 cylinder logarithmic NP constants at \mathscr{I}^-

- There exists a choice of $f(\tilde{\rho})$ that allows the calculation of the associated modified NP constants without imposing the regularity condition.
- The constants $D_{\rho;\ell,m}$ that define the regularity condition precisely correspond to the $f(\tilde{\rho})=\tilde{\rho}$ -modified NP constants.
- ullet We will compute the $\ell=0$ and $\ell=1$ modified NP constants ${}^f\mathcal{N}_{\ell,m}^-.$

$$\tilde{\rho}\underline{L}(\phi_{\ell m}) = \frac{(1+\tau)}{(1-\tau)} \sum_{\rho=0}^{\infty} \underline{Q}_{\rho;\ell,m}^{0}(\tau) \rho^{\rho}. \tag{33}$$

Therefore, for $\ell = 0$, we have:

$$\tilde{\rho} \mathcal{N}_{0,0}^{-} = \lim_{\substack{\rho \to \rho_{\star} \\ \tau \to -1}} \kappa(\underline{\mathbf{e}}^{-})(\phi_{00}) = \sum_{\rho=0}^{\infty} \rho_{\star}^{\rho} \left[\frac{(1+\tau)}{(1-\tau)} \underline{Q}_{\rho;0,0}^{0} \right] |_{\mathscr{I}^{-}}.$$
(34)

• Evaluating at the critical set \mathcal{I}^- , we obtain:

$$\tilde{\rho} \mathcal{N}_{0,0}^{-} = \lim_{\substack{\rho \to \rho_{\star} \\ \tau \to -1}} \sum_{p=0}^{\infty} \rho_{\star}^{p} \left[\frac{(1+\tau)}{(1-\tau)} \underline{Q}_{p;0,0}^{0} \right] = \frac{1}{2} D_{000}.$$
 (35)

• Similarly, for $\ell = 1$, the relevant quantity to evaluate is:

$$\tilde{\rho}\underline{L}(\underline{e}^{-})\phi_{1m} = 4\kappa(\Lambda_{-})^{2} \sum_{p=1}^{\infty} \frac{1}{p!} \rho^{p} \underline{Q}_{p;\ell,m}^{1}.$$
(36)

• Therefore, for $\ell = 1$, we have:

$$\tilde{\rho} \mathcal{N}_{1,m}^{-} = \lim_{\substack{\rho \to \rho_{\star} \\ \tau \to -1}} \kappa \underline{e}(\underline{e}^{-}) \phi_{1m} = \sum_{p=1}^{\infty} \frac{1}{p!} \rho_{\star}^{p-1} (\kappa \underline{Q}_{p;1,m}^{1}) |_{\mathscr{I}^{-}}.$$
(37)

• Evaluating at the critical set \mathcal{I}^- , we obtain:

$$\tilde{\rho} \mathcal{N}_{1,m}^{-} = -\frac{1}{4} D_{11m}. \tag{38}$$

The general case: Induction proof

Proposition

$$(\mathbf{e}^{+})^{n}\phi_{\ell m} = (\Lambda_{+})^{2n} \sum_{\rho=\ell}^{\infty} \frac{1}{\rho!} Q_{\rho;\ell,m}^{n-1}(\tau) \rho^{\rho}$$
(39)

with

$$Q_{p;\ell,m}^{n-1}(\tau) = \sum_{q=0}^{n} (-1)^q p^{[q]} \binom{n}{q} (1+\tau)^{n-q} a_{p;\ell,m}^{(n-q)}(\tau)$$
 (40)

where $p^{[q]}$ denotes the rising factorial.

• Assuming that equations (39) and (40) are valid (induction hypothesis), we can apply e^+ to equation (39), which yields the following expression:

$$(e^{+})^{n+1}\phi_{\ell m} = (\Lambda_{+})^{2}e\Big((\Lambda_{+})^{2n}\sum_{p=\ell}^{\infty}\rho^{p}Q_{p;\ell,m}^{n-1}(\tau)\Big).$$
 (41)

Expanding expression (41), we can get the following result:

$$(\mathbf{e}^{+})^{n+1}\phi_{\ell m} = (\Lambda_{+})^{2(n+1)} \sum_{p=\ell}^{\infty} \rho^{p} \left[(1+\tau) \dot{Q}_{p;\ell,m}^{n-1} - (p+n) Q_{p;\ell,m}^{n-1} \right].$$
(42)

Thus we have shown that:

$$(\mathbf{e}^{+})^{n+1}\phi_{\ell m} = (\Lambda_{+})^{2(n+1)} \sum_{p=\ell}^{\infty} \rho^{p} R_{p;\ell,m}^{n}$$
 (43)

where

$$R_{p;\ell,m}^{n} = (1+\tau)\dot{Q}_{p;\ell,m}^{n-1} - (p+n)Q_{p;\ell,m}^{n-1}.$$
 (44)

 Using the induction hypothesis (40), we can compute all the pieces to construct Rⁿ_{p:ℓ m}:

$$\dot{Q}_{p;\ell,m}^{n-1} = \frac{d}{d\tau} \left[\sum_{q=0}^{n} (-1)^{q} p^{[q]} \binom{n}{q} (1+\tau)^{n-q} a_{p;\ell,m}^{(n-q)}(\tau) \right] =
= \sum_{q=0}^{n} (-1)^{q} p^{[q]} \binom{n}{q}
\left[(1+\tau)^{n-q} a_{p;\ell,m}^{(n-q+1)} + (n-q)(1+\tau)^{n-q-1} a_{p;\ell,m}^{(n-q)} \right].$$
(45)

 By substituting equation (40) (induction hypothesis) and rearranging, we obtain the following:

$$R_{p;\ell,m}^{n} = \sum_{q=0}^{n} (-1)^{n} p^{[q]} \binom{n}{q}$$

$$\{ (1+\tau)^{n-q+1} a^{(n-q+1)} - (p+q)(1+\tau)^{n-q} a^{(n-q)} \}.$$
 (46)

• It follows from the definition of the rising factorial that $p^{[q+1]}=(p+q)p^{[q]}$, in particular $p^{[n+1]}=(p+n+1)p^{[n]}$. Using this fact and relabeling according to $i=q-1 \rightarrow q=i+1$ and i(q=1)=0 & i(q=n+1)=n, we have:

$$R_{p;\ell,m}^{n} = p^{[0]} (1+\tau)^{n+1} a^{(n+1)} +$$

$$+ \sum_{i=0}^{n-1} (-1)^{i+1} p^{[i+1]} \left[\binom{n}{i+1} + \binom{n}{i} \right] (1+\tau)^{n-i} a^{n-i} +$$

$$+ (-1)^{n+1} p^{[n+1]} a.$$
(47)

 Using the recursive identity of the binomial coefficients, the latter can be rewritten more compactly as

$$R_{p;\ell,m}^{n} = \sum_{q=0}^{n+1} (-1)^{q} p^{[q]} \binom{n+1}{q} (1+\tau)^{n-q+1} a^{(n-q+1)}. \tag{48}$$

• And recalling equation (40) then one concludes that

$$R_{p;\ell,m}^n = Q_{p;\ell,m}^n. \tag{49}$$

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The NP constants in terms of initial data

Classical NP Constants at \$\mathcal{I}^+\$:

$$\mathcal{N}_{\ell,m}^+ = q^+(\ell) A_{\ell+1,\ell,m}$$

• Modified NP Constants at \mathscr{I}^+ (with $f(\tilde{\rho}) = \tilde{\rho}$):

$$^{ ilde{
ho}}\mathcal{N}_{\ell,m}^{+}=q^{+}(\ell)\;D_{\ell,\ell,m}$$

Classical NP Constants at \$\mathcal{I}^-\$:

$$\mathcal{N}_{\ell,m}^- = q^-(\ell) B_{\ell+1,\ell,m}$$

• Modified NP Constants at \mathscr{I}^- (with $f(\tilde{\rho}) = \tilde{\rho}$):

$$\tilde{\rho}\mathcal{N}_{\ell,m}^-=q^-(\ell)\;D_{\ell,\ell,m}$$

- Key Observations for Spin-0 Field Near *i*⁰:
 - Regularity condition is essential for well-defined NP constants.
 - 2 Classical NP constants differ between \mathscr{I}^{\pm} .
 - § Logarithmic NP constants at \mathscr{I}^{\pm} are equivalent (up to a numerical constant) with the same initial data.

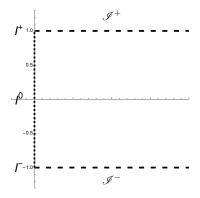


Figure: Representation of the Friedrich Cylinder. Past null infinity is represented by \mathscr{I}^- , Future null infinity \mathscr{I}^+ , spatial infinity as i^0 and the critical sets as \mathcal{I}^+ and \mathcal{I}^- .

Conclusions

- Computed the NP constants for a spin-0 field near spatial and null infinity using the i^0 cylinder framework, exploring the link between NP constants at future and past null infinity based on initial data.
- Analytic initial data prescribed near i^0 results in solution irregularity at the convergence sets \mathcal{I}^{\pm} , controlled by a constant $D_{p;p,m}$.
- When the regularity condition, $D_{p;p,m}=0$, is unmet, the classical NP constants become undefined, allowing a polyhomogeneous expansion for the physical field $\tilde{\phi}$ around null infinities.
- Upon satisfying the regularity condition, the classical NP constants at \mathscr{I}^\pm arise from unique initial data components.
- Conclusively, while classical NP constants at \mathscr{I}^{\pm} usually lack correspondence, the i^0 cylinder NP constants at \mathscr{I}^{\pm} consistently align due to shared origins in the initial data.