## The cylinder at spatial infinity and asymptotic charges

#### Rafael Pinto

Instituto Superior Técnico



September 26, 2023



Advisors: Dr. Edgar Gasperín and Dr. Alex Vañó Viñuales

## Table of Contents

- Newman-Penrose constants
- 2 Cylinder at  $i^0$
- 3 Spin-0 fields close to  $i^0$  and  $\mathscr{I}$
- 4 The NP-constants for the spin-0 fields close to  $i^0$  &  $\mathscr I$
- 5 The classical NP constants at  $\mathscr{I}^+$
- **6** The  $i^0$  cylinder logarithmic NP constants at  $\mathscr{I}^-$
- The general case: Induction proof
- 8 The NP constants in terms of initial data
- Onclusions

### Introduction

- The Newman-Penrose (NP) constants serve as conserved quantities at null infinity in asymptotically flat gravitational fields.
- These constants present a comprehensive conservation system for various spins: spin-1 fields and spin-2 fields, with our research focusing on spin-0 fields linked to wave equation solutions.
- In the detailed context, while an infinite series of conserved quantities is identified in the linear theory, the non-linear General Relativity theory conserves only ten.

### Conservation laws

• These charges are computed as 2-surface integrals at cuts  $C \approx \mathbb{S}^2$  of null infinity  $\mathscr{I}$ .

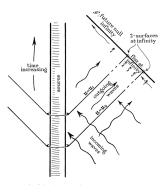


Figure: Visual representation of the behavior of the Newman-Penrose constants at null infinity.

## The $i^0$ cylinder representation in Minkowski spacetime

- Consider spherical polar coordinates  $(\tilde{t}, \tilde{\rho}, \vartheta^A)$  with A = 1, 2.
- ullet The metric of physical Minkowski spacetime in this coordinate system is given by  $ilde{\eta}$ :

$$\tilde{\boldsymbol{\eta}} = -\mathbf{d}\tilde{t} \otimes \mathbf{d}\tilde{t} + \mathbf{d}\tilde{\rho} \otimes \mathbf{d}\tilde{\rho} + \tilde{\rho}^2 \sigma. \tag{1}$$

• Introduce unphysical spherical polar coordinates  $(t, \rho, \vartheta^A)$  as an intermediate step:

$$t = \frac{\tilde{t}}{\tilde{\rho}^2 - \tilde{t}^2}, \quad \rho = \frac{\tilde{\rho}}{\tilde{\rho}^2 - \tilde{t}^2}.$$
 (2)

ullet The conformal metric in unphysical coordinates,  $oldsymbol{\eta}=\Xi^2oldsymbol{ ilde{\eta}}$ :

$$\eta = -\frac{1}{\tilde{\rho}^2 - \tilde{t}^2} (-\mathbf{d}\tilde{t} \otimes \mathbf{d}\tilde{t} + \mathbf{d}\tilde{\rho} \otimes \mathbf{d}\tilde{\rho} + \tilde{\rho}^2 \sigma). \tag{3}$$

# i<sup>0</sup> Representation in Unphysical Coordinates

- In this conformal representation  $(t \in (-\infty, \infty), \rho \in [0, \infty))$ , spatial infinity and the origin interchange.
- $i^0$  is represented by the point  $(t=0, \rho=0)$  in  $(\mathbb{R}^4, \eta)$ .
- Introduce coordinates  $(\tau, \rho, \vartheta^A)$  with  $t = \rho \tau$ .
- Consider the conformal metric  $\mathbf{g} = \rho^{-2} \boldsymbol{\eta}$ .
- ullet Express the unphysical metric  ${m g}$  in  ${m F}$ -coordinates:

$$\mathbf{g} = -\mathbf{d}\tau \otimes \mathbf{d}\tau + \frac{1 - \tau^2}{\rho^2} \mathbf{d}\rho \otimes \mathbf{d}\rho - \frac{\tau}{\rho} \left( \mathbf{d}\rho \otimes \mathbf{d}\tau + \mathbf{d}\rho \otimes \mathbf{d}\tau \right) + \sigma. \tag{4}$$

### Conformal Factor and Lorentz Transformation

• The conformal factor  $\Theta$  in F-coordinates and physical coordinates:

$$\Theta := \rho(1 - \tau^2) = \frac{1}{\tilde{\rho}}.\tag{5}$$

• The boost parameter  $\kappa$ :

$$\kappa := \frac{1+\tau}{1-\tau} = -\frac{\tilde{v}}{\tilde{u}}.\tag{6}$$

• The Lorentz transformation that connects the NP and F-frames:

$$(\Lambda_+)^2 := \Theta^{-1} \kappa^{-1}, \quad (\Lambda_-)^2 := \Theta^{-1} \kappa.$$
 (7)

## i<sup>0</sup> Cylinder and Null Frames

Identify future and past null infinity in the conformal representation:

$$\begin{split} \mathscr{I}^+ &\equiv \{ p \in \mathcal{M} \mid \tau(p) = 1 \}, \\ \mathscr{I}^- &\equiv \{ p \in \mathcal{M} \mid \tau(p) = -1 \}. \end{split}$$

• The  $i^0$ -cylinder represents spatial infinity as an extended set  $\mathcal{I} \approx \mathbb{R} \times \mathbb{S}^2$ :

$$\mathcal{I} \equiv \{ p \in \mathcal{M} \mid |\tau(p)| \ 1, \ \rho(p) = 0 \}, \tag{8}$$

$$\mathcal{I}^0 \equiv \{ p \in \mathcal{M} \mid \tau(p) = 0, \, \rho(p) = 0 \}. \tag{9}$$

## F-Frame and Null Frames

• Introduce the *F*-frame:

$$m{e} = (1+\tau) m{\partial}_{\tau} - \rho m{\partial}_{\rho}, \quad \underline{\pmb{e}} = (1-\tau) m{\partial}_{\tau} + \rho m{\partial}_{\rho}, \quad \pmb{e}_{\pmb{A}} \quad \text{ with } \ m{A} = \{\uparrow, \downarrow\}.$$
 (10)

• The NP-frame hinged at  $\mathscr{I}^{\pm}$ :

NP hinged at 
$$\mathscr{I}^+$$
:  $e^+, \underline{e}^+, e_A^+$ ,  
NP hinged at  $\mathscr{I}^-$ :  $e^-, \underline{e}^-, e_A^-$ .

Transformation between NP and F-frames:

$$\begin{array}{lll} \textit{NP hinged at } \mathscr{I}^+: & \pmb{e}^+ = \Theta^{-2} \textit{L}, & \underline{\pmb{e}}^+ = \underline{\textit{L}}, & \pmb{e}_{\pmb{A}}^+ = \pmb{e}_{\pmb{A}} = \Theta^{-1} \tilde{\pmb{e}}_{\pmb{A}} \\ & \textit{NP hinged at } \mathscr{I}^-: & \pmb{e}^- = \textit{L}, & \underline{\pmb{e}}^- = \Theta^{-2} \underline{\textit{L}}, & \pmb{e}_{\pmb{A}}^- = \pmb{e}_{\pmb{A}} = \Theta^{-1} \tilde{\pmb{e}}_{\pmb{A}}. \\ \end{array}$$

• For  $(\tilde{M}, \tilde{\mathbf{g}})$  and  $(M, \mathbf{g})$ , the transformation of the D'Alembertian operator under conformal transformations is,

$$\Box \phi - \frac{1}{6} \phi R = \Omega^{-3} \left( \tilde{\Box} \tilde{\phi} - \frac{1}{6} \tilde{\phi} \tilde{R} \right). \tag{11}$$

• Using F-coordinates, the wave equation is represented by

$$(\tau^2 - 1) \partial_{\tau}^2 \phi - 2\rho \tau \partial_{\tau} \partial_{\rho} \phi + \rho^2 \partial_{\rho}^2 \phi + 2\tau \partial_{\tau} \phi + \Delta_{S^2} \phi = 0.$$
 (12)

We consider the Ansatz

$$\phi = \sum_{p=0}^{\infty} \sum_{\ell=0}^{p} \sum_{m=-\ell}^{m=\ell} \frac{1}{p!} a_{p;\ell,m}(\tau) \rho^p Y_{\ell m}.$$
 (13)

• Solving (12) simplifies to solving the following (ODE) for every p,  $\ell$ , and m:

$$(1-\tau^2)\ddot{a}_{p;\ell,m} + 2\tau(p-1)\dot{a}_{p,\ell,m} + (\ell+p)(\ell-p+1)a_{p;\ell,m} = 0.$$
 (14)

#### Lemma

The solution to equation (14) is given by:

• For  $p \ge 1$  and  $0 \le \ell \le p-1$ 

$$a(\tau)_{p;\ell,m} = A_{p,\ell,m} \left(\frac{1-\tau}{2}\right)^{p} P_{\ell}^{(p,-p)}(\tau) + B_{p,\ell,m} \left(\frac{1+\tau}{2}\right)^{p} P_{\ell}^{(-p,p)}(\tau)$$
(15)

② For  $p \ge 0$  and  $\ell = p$ :

$$a_{p;p,m}(\tau) = \left(\frac{1-\tau}{2}\right)^p \left(\frac{1+\tau}{2}\right)^p \left(C_{p,p,m} + D_{p,p,m} \int_0^{\tau} \frac{ds}{(1-s^2)^{p+1}}\right)$$
(16)

• When analyzing the hypergeometric function presented in Equation (16) for different values of p we can see the emergence of logarithmic terms. Let's consider the cases of p=0 and p=1, which yield the following expressions:

$$\begin{aligned} a_{0;0,0}(\tau) &= C_{000} + \frac{1}{2} D_{000}(\log(1+\tau) - \log(1-\tau)) \\ a_{1;1,m}(\tau) &= \frac{1}{4} (1-\tau)(1+\tau) \left( C_{11m} + \frac{1}{4} D_{11m}(\log(1+\tau) - \log(1-\tau) + 2\tau(1-\tau^2)) \right). \end{aligned} \tag{17}$$

 These logarithmic terms have implications for the linear version of the associated peeling property.

#### Remark

(Regularity condition). Lemma 1 implies that expanding the integral in (16) results in logarithmic terms, hence  $D_{p,p,m}=0$  is called the regularity condition. The solutions for  $a(\tau)$  are polynomic in  $\tau$ , except for  $\ell=p$  where one needs to impose the regularity condition to only have polynomic solutions.

• We can see how peeling is violated for the spin-0 field by expanding  $\tilde{\phi}$  in terms of the *F*-frame. Making use of eqs (2.34), (2.35) and *Lemma* 1, we truncate (13) up tp p=1 and get the following expression:

$$\tilde{\phi} = \frac{C_{000}}{\tilde{\rho}} + \frac{1}{2\tilde{\rho}} D_{000} \log \left( \frac{\tilde{\rho} + \tilde{t}}{\tilde{\rho} - \tilde{t}} \right) Y_{00} + \frac{1}{8\tilde{\rho}} \left( 1 - \frac{\tilde{t}^2}{\tilde{\rho}^2} \right) 
\left[ D_{11-1} \log \left( \frac{\tilde{\rho} + \tilde{t}}{\tilde{\rho} - \tilde{t}} \right) Y_{1-1} + D_{110} \log \left( \frac{\tilde{\rho} + \tilde{t}}{\tilde{\rho} - \tilde{t}} \right) Y_{10} \right] 
+ \left[ D_{111} \log \left( \frac{\tilde{\rho} + \tilde{t}}{\tilde{\rho} - \tilde{t}} \right) Y_{11} \right] + \frac{1}{2\tilde{\rho}} \left( \tilde{t}^2 - \tilde{\rho}^2 \right) (A_{100} + B_{100}) Y_{00} + 
\frac{1}{4\tilde{\rho}} \left( 1 - \frac{\tilde{t}^2}{\tilde{\rho}^2} \right) (C_{11-1} Y_{1-1} + C_{110} Y_{10} + C_{111} Y_{11}).$$
(19)

• The expantion of the physical field  $\tilde{\phi}$  helps to understand the concept of peeling in the spin-0 case. In the spin-0 case, the peeling property is violated by the logarithmic terms that appear in the expansion of  $\tilde{\phi}$ .

## The NP-constants for the spin-0 fields close to $i^0 \& \mathscr{I}$

The NP constants emerge from a set of asymptotic conservation laws.
 For the spin-0 field in flat spacetime one has:

$$\underline{L}(\tilde{\rho}^{-2\ell}L(e^+)^{\ell+1}\phi_{\ell m}) = 0, \qquad L(\tilde{\rho}^{-2\ell}\underline{L}(e^-)^{\ell+1}\phi_{\ell m}) = 0 \qquad (20)$$

here,  $\phi_{\ell m} = \int_{\mathbb{S}^2} \phi \, Y_{\ell m} d\sigma$  represents the integral of  $\phi$  multiplied by the spherical harmonics  $Y_{\ell m}$  over the surface  $\mathbb{S}^2$ , where  $d\sigma$  denotes the area element on  $\mathbb{S}^2$ .

• One can introduce the  $f(\tilde{\rho})$ -modified NP constants in the following manner:

$${}^{f}\mathcal{N}_{\ell,m}^{+} := f(\tilde{\rho})L(\boldsymbol{e}^{+})^{\ell}\phi_{\ell m}|_{C^{+}}, \tag{21}$$

$${}^{f}\mathcal{N}_{\ell,m}^{-} := f(\tilde{\rho})\underline{L}(\underline{e}^{-})^{\ell}\phi_{\ell m}|_{C^{-}}.$$
 (22)

here,  $C^{\pm} \approx \mathbb{S}^2$  represents a cut of  $\mathscr{I}^{\pm}$ .

• In the specific case of  $f(\tilde{\rho}) = \tilde{\rho}^2$ , these quantities are referred to as the "classical NP-constants" and are denoted as  $\mathcal{N}_{\ell,m}^{\pm}$ . They can be succinctly expressed as follows:

$$\mathcal{N}_{\ell,m}^{+} := (\boldsymbol{e}^{+})^{\ell+1} \phi_{\ell m} |_{C^{+}}, \tag{23}$$

$$\mathcal{N}_{\ell,m}^{-} := (\underline{\underline{e}}^{-})^{\ell+1} \phi_{\ell m} |_{C^{-}}. \tag{24}$$

In the notation for  ${}^f\mathcal{N}^+_{\ell,m}$  presented below, it will be implicitly assumed that m takes values from  $-\ell$  to  $\ell$ .

## The classical NP constants at $\mathscr{I}^+$

• This analysis is facilitated by the expression,

$$\phi_{\ell m} = \sum_{p=\ell}^{\infty} \frac{1}{p!} a_{p;\ell,m}(\tau) \rho^p. \tag{25}$$

• Considering  $\ell=0$ , the computation of  ${\boldsymbol e}^+(\phi_{00})$  is sufficient.

$$\mathbf{e}^{+}(\phi_{\ell m}) = 4\rho^{-1}(1+\tau)^{-2} \sum_{p=0}^{\infty} \frac{1}{p!} \rho^{p}((1+\tau)\dot{a}_{p;\ell,m} - pa_{p;\ell,m}). \quad (26)$$

With

$$Q_{p;\ell,m}^{0}(\tau) := (1+\tau)\dot{a}_{p;\ell,m} - pa_{p;\ell,m}.$$
 (27)

ullet With this definition in place, we can express  $m{e}^+(\phi_{\ell m})$  as follows:

$$\mathbf{e}^{+}(\phi_{\ell m}) = 4(\Lambda_{+})^{2} \sum_{p=0}^{\infty} \frac{1}{p!} \rho^{p} Q_{p,\ell,m}^{0}(\tau). \tag{28}$$

• To compute the  $\ell=0$  NP constant at  $\mathscr{I}^+$ , it is necessary to evaluate  $\boldsymbol{e}^+(\phi_{00})$  at a specific cut  $C^+$  of  $\mathscr{I}^+$ . By utilizing equation (28) we obtain the following expression:

$$\mathcal{N}_{0,0}^{+} = \lim_{\substack{\rho \to \rho_{\star} \\ \tau \to 1}} \mathbf{e}^{+}(\phi_{00}) = \sum_{p=0}^{\infty} \frac{1}{p!} \rho_{\star}^{p-1} Q_{p,0,0}^{0}|_{\mathscr{I}^{+}} = -A_{100}.$$
 (29)

ullet To calculate the  $\ell=1$  NP constants, we must evaluate  $(oldsymbol{e}^+)^2(\phi_{1m}).$ 

$$(e^{+})^{2}(\phi_{\ell m}) = 4^{2}(\Lambda_{+})^{2} e((\Lambda_{+})^{2} \sum_{p=\ell}^{\infty} \frac{1}{p!} \rho^{p} Q_{p,\ell,m}^{0}(\tau)),$$
 (30)

where

$$\boldsymbol{e}\left(\Lambda_{+}^{2}\right) = -\Lambda_{+}^{2}.\tag{31}$$

• Utilizing the aforementioned equations yields:

$$(\mathbf{e}^{+})^{2}(\phi_{\ell m}) = 4^{2}(\Lambda_{+})^{4} \sum_{p=\ell}^{\infty} \frac{1}{p!} \rho^{p} ((1+\tau) \dot{Q}_{p;\ell,m}^{0}(\tau) - (p+1) Q_{p;\ell,m}^{0}(\tau))$$

(32)

Where

$$Q^1_{p,\ell,m}(\tau) := \left( (1+\tau) \dot{Q}^0_{p,\ell,m}(\tau) - (p+1) Q^0_{p,\ell,m}(\tau) \right). \tag{33}$$

Finally, one has:

$$\mathcal{N}_{1,m}^{+} = \lim_{\substack{\rho \to 0 \\ \tau \to 1}} 2^{-4} \frac{1}{2!} Q_{2,1,m}^{1} \cdot (\tau) = 3A_{21m}. \tag{34}$$

• The NP constants for  $\mathscr{I}^-$  can be calculated in a similar manner, where the time reversed version of the F-frame is used.

# The $i^0$ cylinder logarithmic NP constants at $\mathscr{I}^-$

- There exists a choice of  $f(\tilde{\rho})$  that allows the calculation of the associated modified NP constants without imposing the regularity condition.
- The constants  $D_{p;\ell,m}$  that define the regularity condition precisely correspond to the  $f(\tilde{\rho})=\tilde{\rho}$ -modified NP constants.
- $\bullet$  We will compute the  $\ell=0$  and  $\ell=1$  modified NP constants  ${}^f\mathcal{N}_{\ell,\textit{m}}^-.$

$$\tilde{\rho}\underline{L}(\phi_{\ell m}) = \frac{(1+\tau)}{(1-\tau)} \sum_{\rho=0}^{\infty} \underline{Q}_{\rho;\ell,m}^{0}(\tau) \rho^{\rho}. \tag{35}$$

Therefore, for  $\ell = 0$ , we have:

$$\tilde{\rho} \mathcal{N}_{0,0}^{-} = \lim_{\substack{\rho \to \rho_{\star} \\ \tau \to -1}} \kappa(\underline{\mathbf{e}}^{-})(\phi_{00}) = \sum_{\rho=0}^{\infty} \rho_{\star}^{\rho} \left[ \frac{(1+\tau)}{(1-\tau)} \underline{Q}_{\rho;0,0}^{0} \right] |_{\mathscr{I}^{-}}.$$
 (36)

• Evaluating at the critical set  $\mathcal{I}^-$ , we obtain:

$${}^{\tilde{\rho}}\mathcal{N}_{0,0}^{-} = \lim_{\substack{\rho \to \rho_{\star} \\ \tau \to -1}} \sum_{p=0}^{\infty} \rho_{\star}^{p} \left[ \frac{(1+\tau)}{(1-\tau)} \underline{Q}_{p;0,0}^{0} \right] = \frac{1}{2} D_{000}.$$
 (37)

• Similarly, for  $\ell = 1$ , the relevant quantity to evaluate is:

$$\tilde{\rho}\underline{L}(\underline{e}^{-})\phi_{1m} = 4\kappa(\Lambda_{-})^{2} \sum_{p=1}^{\infty} \frac{1}{p!} \rho^{p} \underline{Q}_{p;\ell,m}^{1}.$$
(38)

• Therefore, for  $\ell = 1$ , we have:

$$\tilde{\rho} \mathcal{N}_{1,m}^{-} = \lim_{\substack{\rho \to \rho_{\star} \\ \tau \to -1}} \kappa \underline{e}(\underline{e}^{-}) \phi_{1m} = \sum_{p=1}^{\infty} \frac{1}{p!} \rho_{\star}^{p-1} (\kappa \underline{Q}_{p;1,m}^{1}) |_{\mathscr{I}^{-}}.$$
(39)

• Evaluating at the critical set  $\mathcal{I}^-$ , we obtain:

$$\tilde{\rho} \mathcal{N}_{1,m}^{-} = -\frac{1}{4} D_{11m}. \tag{40}$$

## The general case: Induction proof

### Proposition

$$(\mathbf{e}^{+})^{n}\phi_{\ell m} = (\Lambda_{+})^{2n} \sum_{p=\ell}^{\infty} \frac{1}{p!} Q_{p;\ell,m}^{n-1}(\tau) \rho^{p}$$
(41)

with

$$Q_{p;\ell,m}^{n-1}(\tau) = \sum_{q=0}^{n} (-1)^q p^{[q]} \binom{n}{q} (1+\tau)^{n-q} a_{p;\ell,m}^{(n-q)}(\tau)$$
 (42)

where  $p^{[q]}$  denotes the rising factorial.

• Assuming that equations (41) and (42) are valid (induction hypothesis), we can apply  $e^+$  to equation (41), which yields the following expression:

$$(\mathbf{e}^{+})^{n+1}\phi_{\ell m} = (\Lambda_{+})^{2}\mathbf{e}\Big((\Lambda_{+})^{2n}\sum_{p=\ell}^{\infty}\rho^{p}Q_{p;\ell,m}^{n-1}(\tau)\Big).$$
 (43)

Expanding expression (43), we can get the following result:

$$(\mathbf{e}^{+})^{n+1}\phi_{\ell m} = (\Lambda_{+})^{2(n+1)} \sum_{p=\ell}^{\infty} \rho^{p} \left[ (1+\tau) \dot{Q}_{p;\ell,m}^{n-1} - (p+n) Q_{p;\ell,m}^{n-1} \right].$$
(44)

Thus we have shown that:

$$(e^+)^{n+1}\phi_{\ell m} = (\Lambda_+)^{2(n+1)} \sum_{p=\ell}^{\infty} \rho^p R_{p;\ell,m}^n$$
 (45)

where

$$R_{p;\ell,m}^{n} = (1+\tau)\dot{Q}_{p;\ell,m}^{n-1} - (p+n)Q_{p;\ell,m}^{n-1}.$$
 (46)

• Using the induction hypothesis (42), we can compute all the pieces to construct  $R_{n\cdot\ell}^n$ :

$$\dot{Q}_{p;\ell,m}^{n-1} = \frac{d}{d\tau} \left[ \sum_{q=0}^{n} (-1)^{q} p^{[q]} \binom{n}{q} (1+\tau)^{n-q} a_{p;\ell,m}^{(n-q)}(\tau) \right] = 
= \sum_{q=0}^{n} (-1)^{q} p^{[q]} \binom{n}{q} 
\left[ (1+\tau)^{n-q} a_{p;\ell,m}^{(n-q+1)} + (n-q)(1+\tau)^{n-q-1} a_{p;\ell,m}^{(n-q)} \right].$$
(47)

 By substituting equation (42) (induction hypothesis) and rearranging, we obtain the following:

$$R_{p;\ell,m}^{n} = \sum_{q=0}^{n} (-1)^{n} p^{[q]} \binom{n}{q}$$

$$\{ (1+\tau)^{n-q+1} a^{(n-q+1)} - (p+q)(1+\tau)^{n-q} a^{(n-q)} \}.$$
 (48)

• It follows from the definition of the rising factorial that  $p^{[q+1]}=(p+q)p^{[q]}$ , in particular  $p^{[n+1]}=(p+n+1)p^{[n]}$ . Using this fact and relabeling according to  $i=q-1 \rightarrow q=i+1$  and i(q=1)=0 & i(q=n+1)=n, we have:

$$R_{p;\ell,m}^{n} = \rho^{[0]} (1+\tau)^{n+1} a^{(n+1)} +$$

$$+ \sum_{i=0}^{n-1} (-1)^{i+1} \rho^{[i+1]} \left[ \binom{n}{i+1} + \binom{n}{i} \right] (1+\tau)^{n-i} a^{n-i} +$$

$$+ (-1)^{n+1} \rho^{[n+1]} a.$$
(49)

 Using the recursive identity of the binomial coefficients, the latter can be rewritten more compactly as

$$R_{p;\ell,m}^{n} = \sum_{q=0}^{n+1} (-1)^{q} p^{[q]} \binom{n+1}{q} (1+\tau)^{n-q+1} a^{(n-q+1)}.$$
 (50)

• And recalling equation (42) then one concludes that

$$R_{p;\ell,m}^n = Q_{p;\ell,m}^n. \tag{51}$$

### The NP constants in terms of initial data

Classical NP Constants at \$\mathcal{I}^+\$:

$$\mathcal{N}_{\ell,m}^+ = q^+(\ell) A_{\ell+1,\ell,m}$$

• Modified NP Constants at  $\mathscr{I}^+$  (with  $f(\tilde{\rho}) = \tilde{\rho}$ ):

$$^{ ilde{
ho}}\mathcal{N}_{\ell,m}^{+}=q^{+}(\ell)\;D_{\ell,\ell,m}$$

Classical NP Constants at \$\mathcal{I}^-\$:

$$\mathcal{N}_{\ell,m}^- = q^-(\ell) B_{\ell+1,\ell,m}$$

• Modified NP Constants at  $\mathscr{I}^-$  (with  $f(\tilde{\rho}) = \tilde{\rho}$ ):

$$^{ ilde{
ho}}\mathcal{N}_{\ell,m}^{-}=q^{-}(\ell)\;D_{\ell,\ell,m}$$

- Key Observations for Spin-0 Field Near *i*<sup>0</sup>:
  - Regularity condition is essential for well-defined NP constants.
  - 2 Classical NP constants differ between  $\mathscr{I}^{\pm}$ .
  - **3** Logarithmic NP constants at  $\mathscr{I}^{\pm}$  are equivalent (up to a numerical constant) with the same initial data.

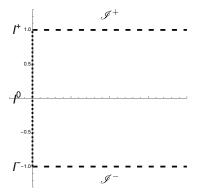


Figure: Representation of the Friedrich Cylinder. Past null infinity is represented by  $\mathscr{I}^-$ , Future null infinity  $\mathscr{I}^+$ , spatial infinity as  $i^0$  and the critical sets as  $\mathcal{I}^+$  and  $\mathcal{I}^-$ .

### Conclusions

- Computed the NP constants for a spin-0 field near spatial and null infinity using the  $i^0$  cylinder framework, exploring the link between NP constants at future and past null infinity based on initial data.
- Analytic initial data prescribed near  $i^0$  results in solution irregularity at the convergence sets  $\mathcal{I}^{\pm}$ , controlled by a constant  $D_{p;p,m}$ .
- When the regularity condition,  $D_{p;p,m}=0$ , is unmet, the classical NP constants become undefined, allowing a polyhomogeneous expansion for the physical field  $\tilde{\phi}$  around null infinities.
- Upon satisfying the regularity condition, the classical NP constants at  $\mathscr{I}^\pm$  arise from unique initial data components.
- Conclusively, while classical NP constants at  $\mathscr{I}^{\pm}$  usually lack correspondence, the  $i^0$  cylinder NP constants at  $\mathscr{I}^{\pm}$  consistently align due to shared origins in the initial data.