```
In [153]: #Stars and stuff
          from astroquery.gaia import Gaia
          from astropy.io.votable import parse, parse_single_table
          from astropy.timeseries import LombScargle
          import pandas as pd
          #Math
          import numpy as np
          from scipy.linalg import lstsq
          #Machine Learning
          from sklearn.model_selection import train_test_split
          #Matplotlib
          import matplotlib.pyplot as plt
          #Caching
          from joblib import Memory
          location = "./cachedir"
          memory = Memory(location, verbose=0)
          #File manipulation
          from urllib.request import urlopen
          import io
          from glob import glob
          import os
  In [2]: def get gaia query rrlyrae(num stars = 100, num clean epochs = 40, conds
          =None, verbose=False):
              add = ""
              if conds is not None:
                  add = f"AND {conds}"
              query = f'''
                  SELECT TOP {num stars} *
                  FROM gaiadr2.gaia source as gaia
                  JOIN gaiadr2.vari rrlyrae using (source id)
                  WHERE
                       num_clean_epochs_g > {num_clean_epochs}
               ''' + add
              if verbose:
                  print(query)
              job = Gaia.launch job async(query)
              return job.get_results()
          get gaia query rrlyrae cached = memory.cache(get gaia query rrlyrae)
In [117]: | def get_gaia_query_general(query):
              job = Gaia.launch job async(query)
              return job.get results()
          get_gaia_query_general_cached = memory.cache(get_gaia_query_general)
  In [3]: def rmse(a, b):
              return np.sqrt(np.mean(np.square(a-b)))
```

```
In [4]: def mse(a,b):
              return np.mean(np.square(a-b))
 In [63]: def get Gband(data,index=None,source id=None):
              assert index!=None or source id!=None, "Must pass in either index or
          source id"
              if index!=None:
                   if source id!=None:
                       print(f"Using index: {index} instead of source id: {source i
          d}")
                   selected row=data[index]
                   source_id=selected_row['source_id']
                  print(f"Analyzing star with source_id: {source_id}")
              lc f = f"lightcurve xmls/{source id}.xml"
              if os.path.exists(lc f):
                   votable = parse_single_table(lc_f)
              else:
                  url = selected_row['epoch_photometry_url']
                   if type(url) == bytes:
                       url = url.decode('utf-8')
                  votable = parse(url)
                  votable.format = 'binary'
                  with open(lc f, 'w') as d:
                       votable.to_xml(d)
                   votable = parse_single_table(lc_f)
              Gstring = "G" if os.name == 'nt' else b'G'
              Gband = votable.array[votable.array['band'] == Gstring]
              return Gband
In [105]: type(rrlyrae rrc['pf'][0]) == np.ma.core.MaskedConstant
Out[105]: True
In [106]: def get maxmin freqs(data):
              if any([type(x)==np.ma.core.MaskedConstant for x in data['pf']]):
                   col name = 'p1 o'
              else:
                   col name = 'pf'
              max_freq, min_freq = 1/np.min(data[col_name]), 1/np.max(data[col_nam
          e1)
              return max freq, min freq
```

```
In [73]: def estimate period(Gband, max freq=2.465, min_freq=1.044, p=False):
             mag = Gband['mag']
             flux = Gband['flux']
             time = Gband['time']
             flux_err = Gband['flux_error']
             freq, power = LombScargle(time, flux, flux_err).autopower(maximum_fr
         equency=max freq,
                                                                        minimum fre
         quency=min_freq,
                                                                        nyquist_fac
         tor=100)
             period = 1/freq[np.argmax(power)]
             phase = time % period
             if p:
                 plt.figure(figsize=(8,5))
                 plt.plot(freq, power, '-k')
                 plt.xlabel("Frequency")
                 plt.ylabel("Spectral Power")
                 fig, (ax1,ax2) = plt.subplots(1,2, figsize=(16,5))
                 ax1.scatter(phase, flux)
                 ax1.set(xlabel="Phase", ylabel="Flux")
                 ax1.grid()
                 ax2.scatter(phase, mag)
                 ax2.set(xlabel="Magnitude", ylabel="Flux")
                 ax2.grid()
                 plt.show()
                 print(f"Estimated period: 1/{freq[np.argmax(power)]:.5f} = {peri
         od:.5f}")
                   print(f"Period as reported by vari rrlyrae: {recorded period:.
         5f}")
                   print(f"RMSE: {np.sqrt(np.mean(np.square(period - recorded per
         iod)))}")
             return period
```

```
In [7]: def plot_magnitude(Gband):
    time = Gband['time']
    mag = Gband['mag']
    mag_uncertainty = 1.09/Gband['flux_over_error']

    plt.figure(figsize=(8,5))
    plt.title("Magnitude w/ Magnitude Uncertainty")
    plt.fill_between(time, mag+mag_uncertainty/2,mag-mag_uncertainty/2,
    color='blue', alpha=0.5)
    plt.xlabel("Time")
    plt.ylabel("Magnitude")
    plt.grid()
    plt.show()
    print(f"Estimated mean: {np.log(np.average(np.exp(mag)))}")
```

```
In [8]: def setup_fit(flux, time, omega, k):
    num_samps = len(time)
    k_s = np.arange(1,k+1)
    tk_tiling = np.outer(time, k_s)*omega
    X = np.zeros((num_samps, 2*k+1))
    X[:,0] = np.ones(num_samps)
    X[:,1:k+1] = np.sin(tk_tiling)
    X[:,k+1:] = np.cos(tk_tiling)
    return X
```

```
In [9]: def pseudo_fourier(omega, t, A0, a, b):
    assert len(a) == len(b), f"Length of a and length of b must be the s
ame"

K = len(a)
s,c = np.zeros(len(t)), np.zeros(len(t))
for k in range(1, K+1):
    s += a[k-1]*np.sin(k*omega*t)
    c += b[k-1]*np.cos(k*omega*t)
return A0 + s + c
```

```
In [99]: def get_prediction(Gband, max_freq, min_freq, k, p=False):
             flux,time=Gband['flux'], Gband['time']
             source_id=Gband['source_id'][0]
             period=estimate period(Gband, max freq=max freq, min freq=min freq)
             omega=2*np.pi/period
             X = setup fit(flux, time, omega, k)
             beta = lstsq(X, flux.data)[0]
             A0=beta[0]
             a=beta[1:k+1]
             b=beta[k+1:]
             time interp=np.arange(np.min(time), np.max(time), 1)
             f_interp = pseudo_fourier(omega=2*np.pi/period, t=time_interp, A0=A0
         , a=a, b=b)
             phase interp, flux interp = sort by phase(time_interp%period, f_inte
         rp)
             if p:
                  fig, ((ax1, ax2), (ax3, ax4)) = plt.subplots(2,2, figsize=(16,7))
         ), gridspec_kw={'height_ratios':[3,1]})
                 ax1.scatter(time%period, flux, label="Original",color='black')
                 ax1.scatter(time%period, X.dot(beta), label="Predicted", color
         ='green')
                 ax1.set(title=f"SourceID: {source id}\nRMSE: {rmse(flux, X.dot(b))
         eta)):.5f}",ylabel="Flux")
                 ax1.legend()
                 ax1.grid()
                 ax3.scatter(time%period, flux-X.dot(beta), label="Residual",colo
         r='red')
                 ax3.set(xlabel="Phase")
                 ax3.grid()
                 ax2.scatter(time%period, flux, label="Original f")
                 ax2.plot(phase interp, flux interp, label="Predicted f", color=
         "red")
                 ax2.set(title="Interpolated Values", xlabel="Phase", ylabel="Flu
         x")
                 ax2.legend()
                 ax2.grid()
                 fig.delaxes(ax4)
                 plt.show()
```

INFO: Query finished. [astroquery.utils.tap.core]

Out[11]: Table length=10

solution_id	designation	random_index	ref_epoch	ra	
			yr	deg	
int64	object	int64	float64	float64	
1635721458409799680	Gaia DR2 5866125710834119808	841033097	2015.5	212.93756378519396	1.81
1635721458409799680	Gaia DR2 5978435871487788288	1394969254	2015.5	256.52946118354015	0.207
1635721458409799680	Gaia DR2 5704736782734774528	122877659	2015.5	132.81229544277778	0.148
1635721458409799680	Gaia DR2 5816755332315333888	259791940	2015.5	254.94654730343584	0.055
1635721458409799680	Gaia DR2 5821611776409134976	299065082	2015.5	246.3262948830741	0.0259
1635721458409799680	Gaia DR2 5642603243216872576	1311594757	2015.5	132.9564341215911	0.047
1635721458409799680	Gaia DR2 5813181197970338560	1546661016	2015.5	263.35817148226175	0.011
1635721458409799680	Gaia DR2 5630421856972980224	923739124	2015.5	140.99364763000955	0.0560
1635721458409799680	Gaia DR2 5810405553887250432	191264318	2015.5	268.50110356278077	0.0131
1635721458409799680	Gaia DR2 5821156028840408576	1453327615	2015.5	244.35278635111487	0.0163

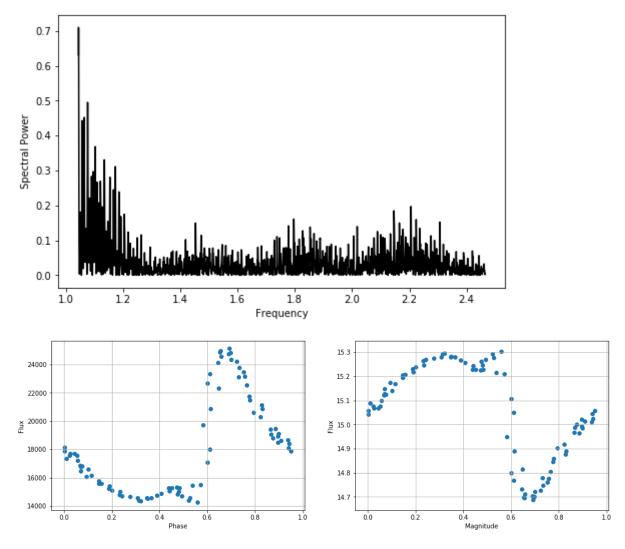
```
In [22]: for row in rrlyrae_100:
    url = row['epoch_photometry_url']
    source_id = row['source_id']
    dest = f"lightcurve_xmls/{source_id}.xml"
    if os.path.exists(dest):
        print("It exists!")
        continue
    votable = parse(url.decode('utf-8'))
    votable.format = 'binary'
    with open(dest, 'w') as d:
        votable.to_xml(d)
```

5813181197970338560

```
In [47]: selected_index = 6
selected_Gband = get_Gband(rrlyrae_100, selected_index)
```

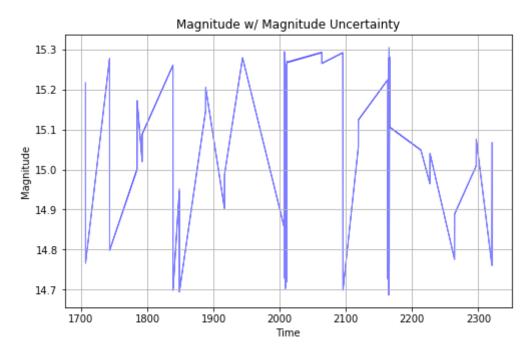
Analyzing star with source_id: 5813181197970338560

```
In [48]: period = estimate_period(selected_Gband, p=True)
    recorded_period = rrlyrae_100['pf'][selected_index]
    print(f"Period as reported by vari_rrlyrae: {recorded_period:.5f}")
    print(f"RMSE: {np.sqrt(np.mean(np.square(period - recorded_period)))}")
    plot_magnitude(selected_Gband)
```



Estimated period: 1/1.04433 = 0.95756
Period as reported by vari_rrlyrae: 0.95765

RMSE: 9.488197693585665e-05

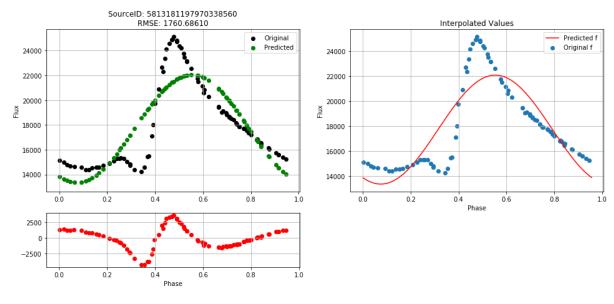


Estimated mean: 15.070074092017139

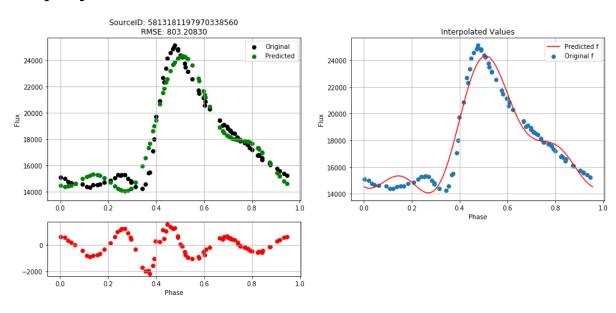
```
In [108]: k_vals = [1,3,5,7,9]
    max_freq, min_freq = get_maxmin_freqs(rrlyrae_100)
    for k in k_vals:
        print(f"Analyzing fit for k={k}...")

        get_prediction(selected_Gband, max_freq=max_freq, min_freq=min_freq, k=k, p=True)
```

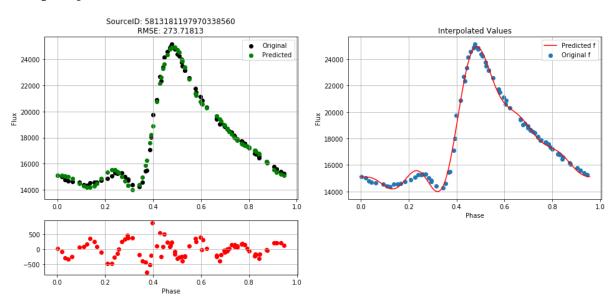
Analyzing fit for k=1...



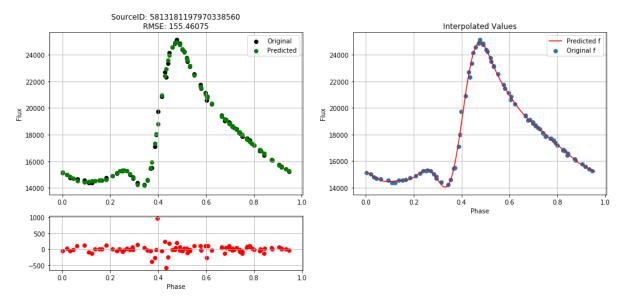
Analyzing fit for k=3...



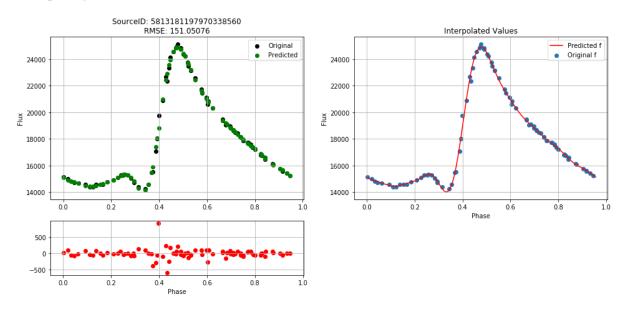
Analyzing fit for k=5...



Analyzing fit for k=7...

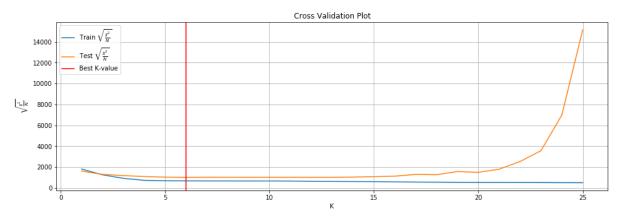


Analyzing fit for k=9...



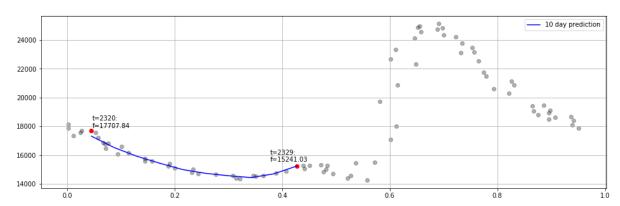
```
In [56]: k_vals = np.arange(1,25+1)
    train_mses = np.zeros(len(k_vals))
    test_mses = np.zeros(len(k_vals))
    time_train, time_test, flux_train, flux_test = train_test_split(time, flux, test_size=0.2)
```

```
In [57]: p = False
         for i,k in enumerate(k vals):
               print(f"Analyzing fit for k=\{k\}...")
             X = setup_fit(flux_train, time_train, omega, k)
             beta train = lstsq(X, flux train.data)[0]
             A0=beta_train[0]
             a=beta train[1:k+1]
             b=beta train[k+1:]
             pred_train = pseudo_fourier(omega, time_train, A0, a, b)
             pred_test = pseudo_fourier(omega, time_test, A0, a, b)
             train_mse = rmse(flux_train, pred_train)
             train mses[i] = train mse
             test_mse = rmse(flux_test, pred_test)
             test mses[i] = test mse
             if p and k \ge 20:
                 fig, (ax1, ax2) = plt.subplots(1,2, figsize=(16,5))
                 ax1.scatter(time_train%period,flux_train, label="Truth")
                 ax1.scatter(time_train%period,pred_train,label="Pred")
                 ax1.set(title=f"K={k}: $\frac{{\chi^2}}{{N}} = {train_mse:.3f}
         $")
                 ax1.legend()
                 ax2.scatter(time test%period,flux test, label="Truth")
                 ax2.scatter(time test%period, pred test, label="Pred")
                 ax2.set(title=f"$\frac{{\chi^2}}{{N}} = {test mse:.3f}$")
                 ax2.legend()
                 plt.show()
         # print(test mses)
         fig,ax = plt.subplots(figsize=(16,5))
         ax.plot(k_vals, train_mses, label="Train $\\sqrt{{\\frac{{\\chi^2}}}
         {{N}}}};")
         ax.plot(k vals, test mses, label="Test $\\sqrt{{\\frac{{\\chi^2}}}{{N}}}}
         $")
         ax.axvline(np.argmin(test mses)+1,color='red', label="Best K-value")
         ax.legend()
         ax.grid()
         ax.set(xlabel="K", ylabel="$\sqrt{{\frac{{\chi^2}}{{N}}}}, title="C
         ross Validation Plot")
         plt.show()
```



```
In [58]: k best = np.argmin(test mses)+1
         X = setup fit(flux, time, omega, k best)
         beta_best = lstsq(X, flux.data)[0]
         A0 best=beta best[0]
         a_best=beta_best[1:k_best+1]
         b_best=beta_best[k_best+1:]
         last flux, last time = flux[-1], time[-1]
         print(last flux, last time)
         final_time = last_time+10
         time extrap = np.arange(last time, final time)
         pred extrap = pseudo fourier(omega, time extrap, A0 best, a best, b best
         phase extrap, flux extrap = sort by phase(time extrap%period, pred extra
         fig,ax = plt.subplots(figsize=(16,5))
         ax.scatter(time%period, flux,color='black', alpha=0.3)
         ax.scatter(last time%period, last flux,color='red')
         ax.annotate(f"t={last_time:.0f}:\nf={last_flux:.2f}", (last_time%period+
         0.002, last flux+200))
         ax.plot(phase extrap, flux extrap, color='blue', label="10 day predictio
         n")
         ax.scatter(phase extrap[-1], flux extrap[-1], color='red')
         ax.annotate(f"t={time extrap[-1]:.0f}:\nf={flux extrap[-1]:.2f}", (phase
         extrap[-1]-0.05, flux extrap[-1]+300)
         ax.grid()
         ax.legend()
         plt.show()
```

17707.83890462764 2320.202939518489

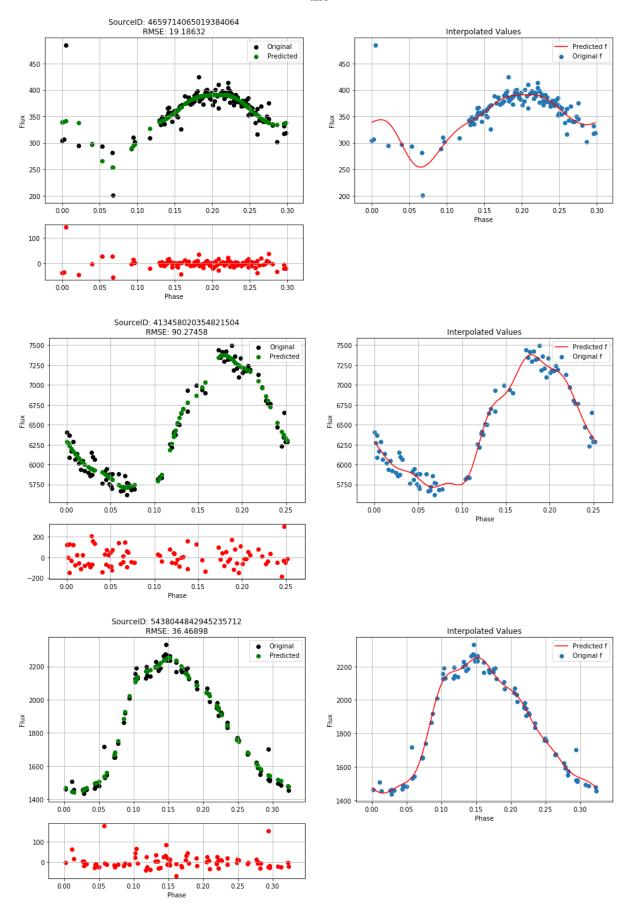


Out[110]: Table length=3

solution_id	designation	random_index	ref_epoch	ra	
			yr	deg	
int64	object	int64	float64	float64	
1635721458409799680	Gaia DR2 4659714065019384064	1255299599	2015.5	89.71078260193757	0.3040
1635721458409799680	Gaia DR2 413458020354821504	1251053435	2015.5	19.111954139381382	0.0238
1635721458409799680	Gaia DR2 5438044842945235712	865958970	2015.5	142.96098504851844	0.0451

Analyzing star with source_id: 4659714065019384064 Analyzing star with source_id: 413458020354821504 Analyzing star with source_id: 5438044842945235712

```
In [112]: max_freq, min_freq = get_maxmin_freqs(rrlyrae_rrc)
    for Gband in rrc_Gbands:
        get_prediction(Gband,max_freq=max_freq, min_freq=min_freq, k=k_best,
        p=True)
```

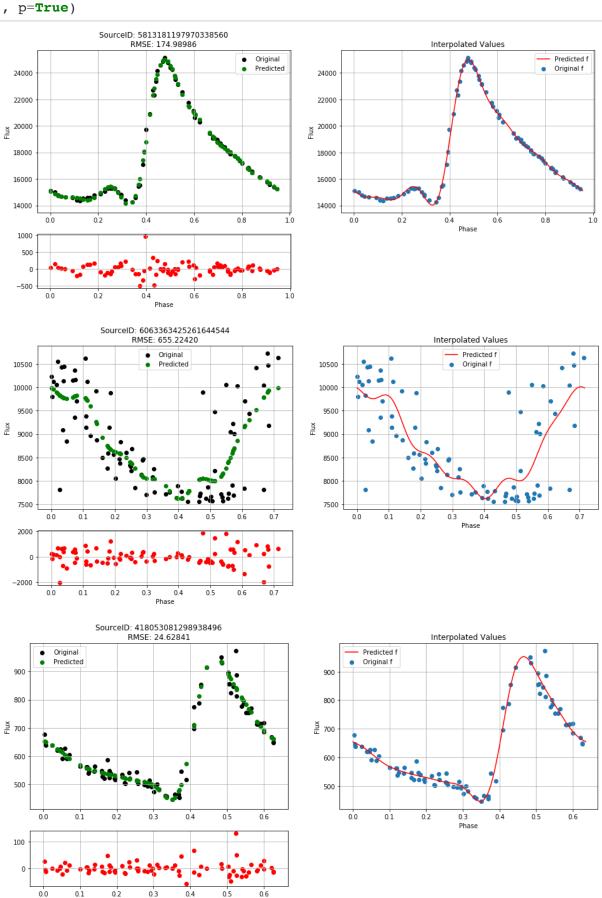


Out[113]: Table length=3

solution_id	designation	random_index	ref_epoch	ra	
			yr	deg	
int64	object	int64	float64	float64	
1635721458409799680	Gaia DR2 5813181197970338560	1546661016	2015.5	263.35817148226175	0.011
1635721458409799680	Gaia DR2 6063363425261644544	307736288	2015.5	197.52702038202472	0.0206
1635721458409799680	Gaia DR2 418053081298938496	516077856	2015.5	8.774802583044732	0.114

```
In [114]: rrab_Gbands = [get_Gband(rrlyrae_rrab, i) for i, row in enumerate(rrlyra
e_rrab)]
```

Analyzing star with source_id: 5813181197970338560 Analyzing star with source_id: 6063363425261644544 Analyzing star with source_id: 418053081298938496



```
In [151]: query="""
          SELECT *
          FROM gaiaedr3.gaia_source as gaia3
          JOIN gaiaedr3.dr2_neighbourhood as gaia2
              ON gaia2.dr3_source_id = gaia3.source_id
          JOIN gaiadr2.vari_rrlyrae as rrlyrae
              ON gaia2.dr2_source_id = rrlyrae.source_id
          JOIN external.gaiaedr3_distance as dists
              ON gaia3.source id = dists.source id
          WHERE
              abs(b)>30
              AND parallax>0.25
              AND parallax_over_error>5
              AND pf IS NOT NULL
          rrlyrae gaia3 = get gaia query general cached(query)
          print(f"Got {len(rrlyrae gaia3)} stars")
          rrlyrae gaia3[:10]
```

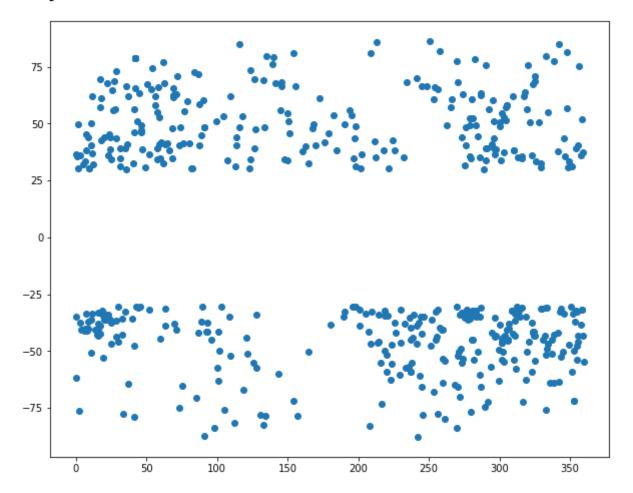
INFO: Query finished. [astroquery.utils.tap.core]
Got 548 stars

Out[151]: Table length=10

solution_id	designation	source_id	random_index	ref_epoch	
				yr	
int64	object	int64	int64	float64	
1636042515805110273	Gaia EDR3 4657904848780392832	4657904848780392832	1297321964	2016.0	82. ⁻
1636042515805110273	Gaia EDR3 4798586678570579072	4798586678570579072	279121884	2016.0	83.4
1636042515805110273	Gaia EDR3 4685757887726594816	4685757887726594816	1480418546	2016.0	11.20
1636042515805110273	Gaia EDR3 4689637956899105792	4689637956899105792	1084076130	2016.0	5.919
1636042515805110273	Gaia EDR3 4689637961175354240	4689637961175354240	47042557	2016.0	5.9 ⁻
1636042515805110273	Gaia EDR3 4700326863447344768	4700326863447344768	267955142	2016.0	39.2
1636042515805110273	Gaia EDR3 4690886594062571008	4690886594062571008	1401140051	2016.0	15.77
1636042515805110273	Gaia EDR3 4658012051209225984	4658012051209225984	847616141	2016.0	81.9
1636042515805110273	Gaia EDR3 4658145092114070144	4658145092114070144	1681799523	2016.0	79.6
1636042515805110273	Gaia EDR3 4685834544315524224	4685834544315524224	398366609	2016.0	11.38

```
In [152]: plt.figure(figsize=(10,8))
    plt.scatter(rrlyrae_gaia3['l'], rrlyrae_gaia3['b'])
```

Out[152]: <matplotlib.collections.PathCollection at 0x1a26c48250>



```
In [159]:
          rrlyrae gaia3 df = rrlyrae gaia3.to pandas()
          dists = pd.concat(
                   [rrlyrae gaia3 df[["r med geo", "r hi geo", "r lo geo"]], 1000/r
          rlyrae_gaia3_df['parallax']],
                  axis=1
              ).rename(columns={"parallax":"distance"}).sort_values("distance")
          n = np.arange(dists.shape[0])
          plt.figure(figsize=(20,10))
          plt.plot(n,dists["distance"], label=r"$\frac{1}{\mathrm{parallax}}$")
          plt.plot(n, dists["r_med_geo"], label="Geometric Distances")
          plt.plot(n, dists["r_hi_geo"], "--", color="red", alpha=0.4, label="geo
           err")
          plt.plot(n, dists["r_lo_geo"], "--", color="red", alpha=0.4)
          plt.ylabel("Parsecs")
          plt.xlabel("Star Index (sorted by distance)")
          plt.legend()
          plt.show()
```

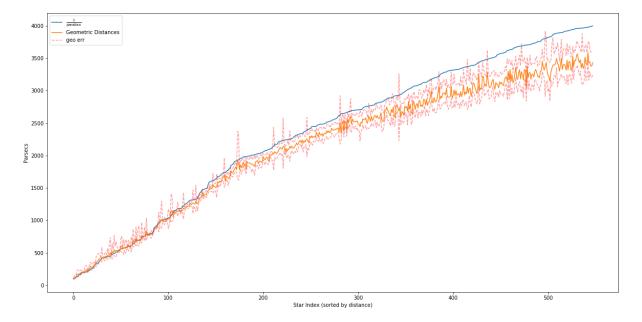
eractiveshell.py:3331: TableReplaceWarning: converted column 'phot_bp_n
_contaminated_transits' from integer to float
 exec(code_obj, self.user_global_ns, self.user_ns)
/Users/Rafferino/anaconda3/lib/python3.7/site-packages/IPython/core/int
eractiveshell.py:3331: TableReplaceWarning: converted column 'phot_bp_n
_blended_transits' from integer to float
 exec(code_obj, self.user_global_ns, self.user_ns)
/Users/Rafferino/anaconda3/lib/python3.7/site-packages/IPython/core/int
eractiveshell.py:3331: TableReplaceWarning: converted column 'phot_rp_n
 contaminated transits' from integer to float

/Users/Rafferino/anaconda3/lib/python3.7/site-packages/IPython/core/int

exec(code_obj, self.user_global_ns, self.user_ns)
/Users/Rafferino/anaconda3/lib/python3.7/site-packages/IPython/core/int
eractiveshell.py:3331: TableReplaceWarning: converted column 'phot_rp_n

exec(code obj, self.user global ns, self.user ns)

blended transits' from integer to float



```
In [160]: | query = """
               SELECT *
               FROM gaiadr2.gaia_source as gaia2
               JOIN gaiaedr3.dr2_neighbourhood as link
                   ON gaia2.source_id = link.dr2_source_id
               JOIN gaiadr2.vari_rrlyrae as rrlyrae
                   ON rrlyrae.source_id = gaia2.source_id
               WHERE
                   parallax over error > 5
                   AND abs(b) > 30
                   AND parallax > 0.25
                   AND pf IS NOT NULL
           0.00
          rrlyrae_gaia2 = get_gaia_query_general_cached(query)
          print(f"Got {len(rrlyrae_gaia2)} stars")
          rrlyrae_gaia2[:10]
```

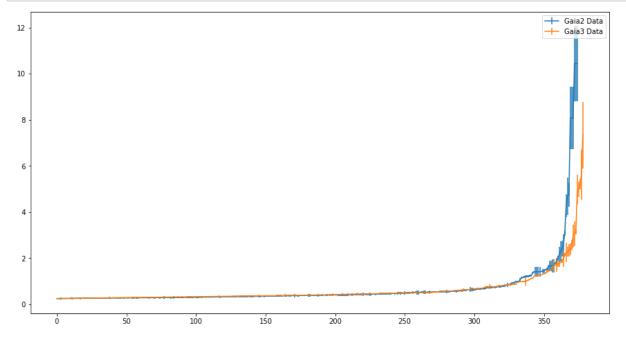
INFO: Query finished. [astroquery.utils.tap.core]
Got 518 stars

Out[160]: Table length=10

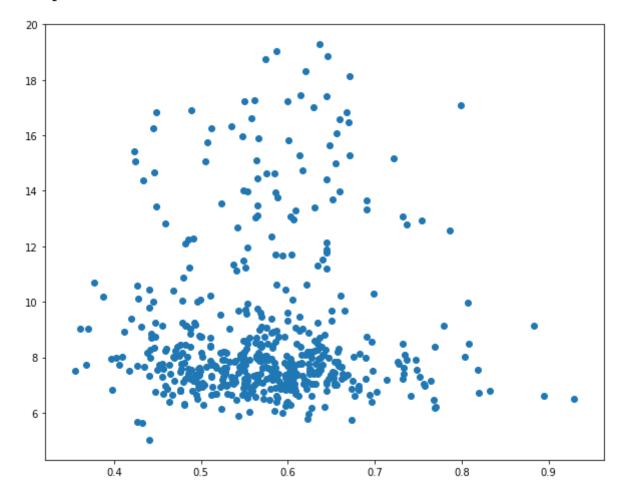
solution_id	designation	source_id	random_index	ref_epoch	
				yr	
int64	object	int64	int64	float64	
1635721458409799680	Gaia DR2 4654631533876789632	4654631533876789632	502718668	2015.5	74
1635721458409799680	Gaia DR2 4654631533876789632	4654631533876789632	502718668	2015.5	74
1635721458409799680	Gaia DR2 4685757887726594816	4685757887726594816	812747682	2015.5	11.20
1635721458409799680	Gaia DR2 4689637956899105792	4689637956899105792	1530397305	2015.5	5.9 ⁻
1635721458409799680	Gaia DR2 4689637956899105792	4689637956899105792	1530397305	2015.5	5.9 ⁻
1635721458409799680	Gaia DR2 4700326863447344768	4700326863447344768	642465858	2015.5	39.2
1635721458409799680	Gaia DR2 4658186602866639872	4658186602866639872	1491505400	2015.5	79.(
1635721458409799680	Gaia DR2 4658186602866639872	4658186602866639872	1491505400	2015.5	79.(
1635721458409799680	Gaia DR2 4632502526615290624	4632502526615290624	1511122055	2015.5	32.40
1635721458409799680	Gaia DR2 4632502526615290624	4632502526615290624	1511122055	2015.5	32.40

```
mask = [(source_id in rrlyrae_gaia2["dr3_source_id"]) for source_id in r
          rlyrae gaia3["source id"]]
           gaia3_data = rrlyrae_gaia3[mask].to_pandas().sort_values("parallax")
          gaia3_data["dr3_source_id"]
Out[173]: 271
                  3186033648144893696
          340
                  6910756005950942336
          93
                  3945610468551775616
          171
                  612591803903942016
          183
                  1540009731422069504
                         . . .
          61
                  4943063090574101376
          63
                  4955988365155969792
          3
                  4700326863447344768
          279
                  4998160202357316096
          179
                  1328057321618952064
          Name: dr3_source_id, Length: 379, dtype: int64
In [176]: mask2 = [(source id in rrlyrae gaia3["dr3 source id"]) for source id in
          rrlyrae gaia2["source id"]]
           gaia2_data = rrlyrae_gaia2[mask2].to_pandas().sort_values("<mark>parallax"</mark>)
          len(gaia2 data)
          /Users/Rafferino/anaconda3/lib/python3.7/site-packages/IPython/core/int
          eractiveshell.py:3331: TableReplaceWarning: converted column 'priam fla
          gs' from integer to float
            exec(code obj, self.user global ns, self.user ns)
          /Users/Rafferino/anaconda3/lib/python3.7/site-packages/IPython/core/int
          eractiveshell.py:3331: TableReplaceWarning: converted column 'flame fla
          gs' from integer to float
            exec(code obj, self.user global ns, self.user ns)
Out[176]: 375
```

```
In [177]: plt.figure(figsize=(15,8))
    plt.errorbar(np.arange(len(gaia2_data)), gaia2_data["parallax"], yerr=ga
    ia2_data["parallax_error"], label="Gaia2 Data")
    plt.errorbar(np.arange(len(gaia3_data)), gaia3_data["parallax"], yerr=ga
    ia3_data["parallax_error"], label="Gaia3 Data")
    plt.legend()
    plt.show()
```



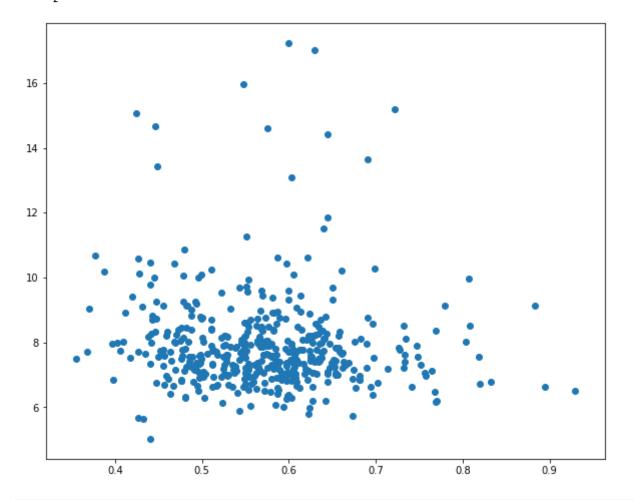
Out[179]: <matplotlib.collections.PathCollection at 0x1a27edeb50>



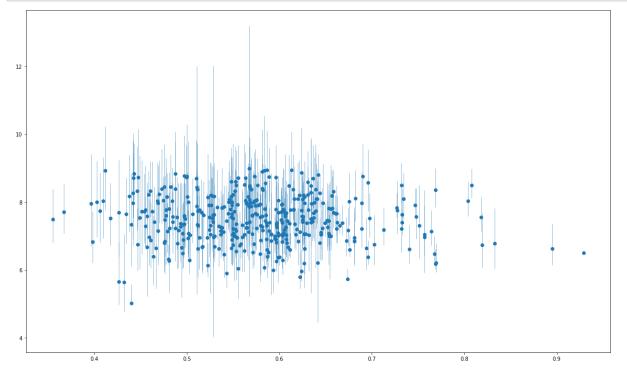
```
In [ ]: | query = """
            SELECT *
            FROM gaiaedr3.gaia_source as gaia
            JOIN gaiaedr3.dr2 neighbourhood as link
                ON gaia.source id = link.dr3 source id
            JOIN gaiadr2.vari_rrlyrae as lyrae
                ON lyrae.source id = link.dr2 source id
            JOIN external.gaiaedr3_distance as dists
                ON gaia.source id = dists.source id
            WHERE
                parallax over error > 5
                AND abs(b) > 30
                AND parallax > 0.25
                AND pf IS NOT NULL
                AND astrometric excess noise < 1
                AND 1 + 0.015*power(bp rp, 2) < phot bp rp excess factor
                AND phot bp rp excess factor < 1.3 + 0.06*power(bp rp, 2)
        0.00
        rrlyrae_gaia3_filt = get_gaia_query_general_cached(query)
```

```
In [181]: Mg = rrlyrae_gaia3_filt["phot_g_mean_mag"]-5*np.log10((rrlyrae_gaia3_fil
t["r_hi_geo"]-rrlyrae_gaia3_filt["r_lo_geo"])/2)+5
plt.figure(figsize=(10,8))
plt.scatter(rrlyrae_gaia3_filt["pf"], Mg)
```

Out[181]: <matplotlib.collections.PathCollection at 0x1a25617d90>



```
In [191]: plt.figure(figsize=(20,12))
    plt.errorbar(rrlyrae_gaia3_filt_df["pf"], rrlyrae_gaia3_filt_df["Mg"], y
    err=error, fmt="o", elinewidth=0.5)
    plt.show()
```



In []: