

$\Gamma(n \rightarrow p + \bar{\nu}_e)$: prob. di decad. / unità di tempo.

$$N(t) = N_0 e^{-\Gamma t} \quad N(t) \propto |\Psi|^2 \propto e^{-\Gamma t}$$

$$H = H_0 + \frac{H_I}{2}$$

$$\Psi(\vec{r}, t) = \psi(\vec{r}) e^{-iE_0 t} e^{-\frac{\Gamma}{2}t}$$

Fourier $\psi(\vec{r}) \rightarrow \psi(E)$

$$A(E) = \frac{1}{2\pi} \int_{-\infty}^{\infty} dt e^{iEt} (e^{-iE_0 t} e^{-\frac{\Gamma}{2}t})$$

$t > 0$ per le particelle create a $t = 0$.

che decadono nel tempo medio

$$\langle t \rangle = T = \frac{1}{\Gamma}$$

$$A(E) = \frac{1}{2\pi} \frac{1}{i(E - E_0) - \frac{\Gamma}{2}}$$

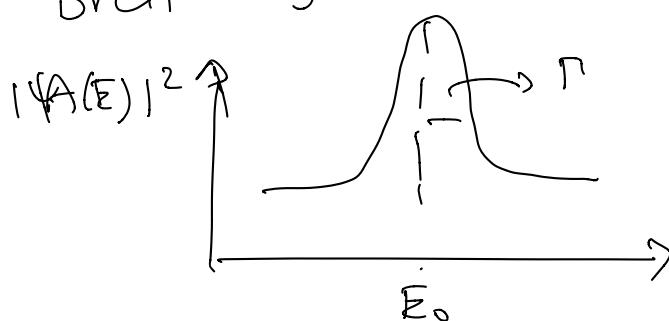
$$|A(E)|^2 \propto \left(\frac{1}{2\pi}\right)^2 \frac{1}{(E - E_0)^2 + \left(\frac{\Gamma}{2}\right)^2}$$

Lorentziana

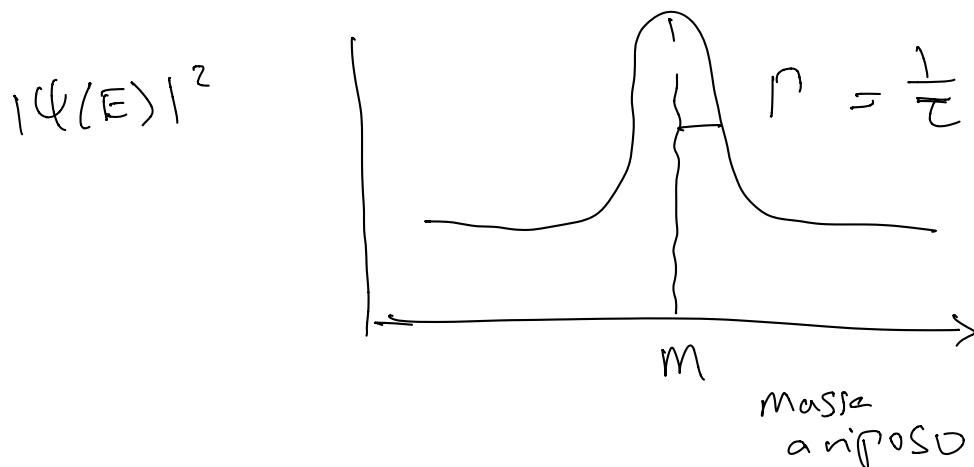
Cauchy

Breit-Wigner

Risposta



nelle nⁱf solide le can le particelle instabili



$\tau \rightarrow$ grande. Limite particelle stabile \Rightarrow massa ben determinata.

$$n \cdot \tau = t_1 = 1$$

$$\mu \rightarrow e^- \bar{\nu}_e \nu_\mu \quad \tau \approx 10^{-6} \text{ s.}$$

$$\bar{u} \rightarrow j^- \bar{\nu}_j \quad \tau \approx 10^{-8} \text{ s}$$

$$\pi^- = \begin{pmatrix} \bar{u} \\ d \end{pmatrix} \quad \text{bosone sclerare} \quad S=0 \quad m \approx 140 \text{ MeV}$$

$$\rho^- = \begin{pmatrix} \bar{u} \\ d \end{pmatrix} \quad S=1$$

$$m_\rho = 770 \text{ MeV} \quad \eta \approx 150 \text{ MeV}$$

$$f_{20} \rightarrow g_{20}$$

$$\text{Se } \Gamma \rightarrow 0$$

$$B-W \rightarrow$$



nelle nⁱf reale Distrib. spazio = $F(x)$ \bigcirc Risol.

Se $\Gamma \neq 0 \Rightarrow$ Gaussiana detta dalla nⁱcⁱ. spazio

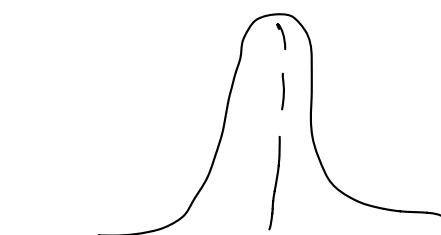
$\mathcal{P} \neq 0$

$$\mathcal{B}-\mathcal{W}(E|_{M,F}) \otimes \mathcal{G}$$

$$\frac{\Gamma}{M}$$

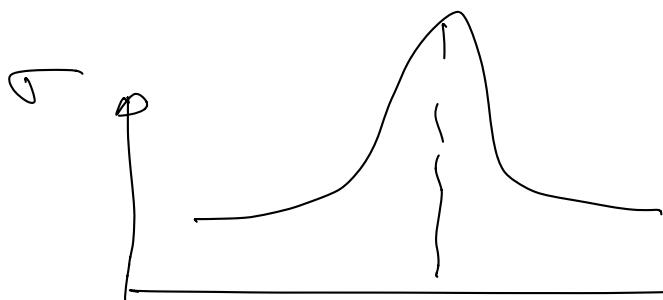
$$\pi^- \rightarrow \mu^- \bar{\nu}_\mu \quad \bar{B}^0 \rightarrow K^- \pi^+$$

$$K^0 \rightarrow \pi^+ \pi^-$$



$$m_{inv,1} = \sqrt{|\underline{P}_{\pi^+} + \underline{P}_{\pi^-}|^2}$$

$$\pi^+ \rightarrow \bar{\nu}^- \quad \text{misimo } \sigma$$



$$\sqrt{s}: \text{en. rel. C.d.m}$$

$$\pi^+ \pi^-$$

$$\pi^+ \pi^- \rightarrow X$$

$$\pi^+ \pi^-$$

$$\pi^+ \pi^- \pi^0$$

$$\pi^+ \pi^- \pi^+ \pi^-$$

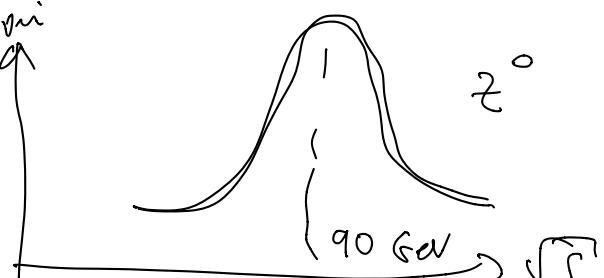
$$\pi^+ \pi^- \pi^0 \pi^0$$

$$\pi^+ \pi^- p \bar{p}$$

\sqrt{s} aumenta

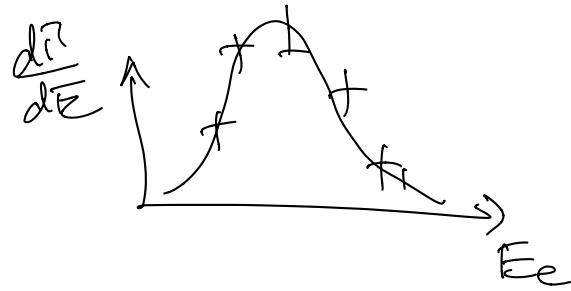
reac^{2,ini}

$$e^+ e^- \rightarrow X$$



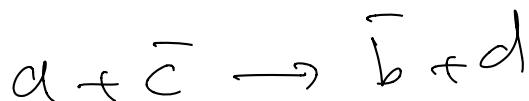
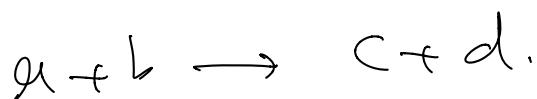
$$n \rightarrow p e^- \bar{\nu}_e$$

Pauli suppose ν_e
 Fermi calculate $\frac{d\Gamma}{dE}$



$$n \rightarrow p e^- \bar{\nu}_e$$

$$V_I = G_F$$



$$E_S = 1.8 \text{ MeV}$$

1956 Reines - Cowan.

Vicino a un reattore $P = 1 \text{ GW}$

Fissione dell'urano $Q = 200 \text{ MeV}$
 $\langle N_V \rangle = 6$

$$P = 10^9 \frac{J}{s} \quad 1 \text{ eV} = 1.6 \times 10^{-19} \frac{\text{C} \cdot \text{V}}{\text{J}}$$

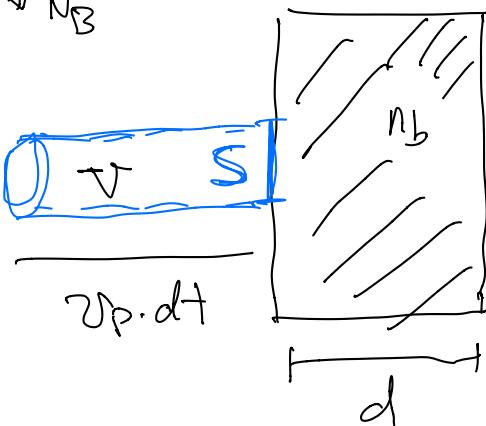
$$= \frac{10^9}{1.6 \times 10^{-19}} \frac{\text{eV}}{\text{s}} =$$

$$\# \frac{\text{reazioni}}{\text{sec}} = \frac{P}{Q} \Rightarrow 3 \times 10^{19} \text{ Hz}$$

$$\# \nu_e = 6 \times 3 \times 10^{19} = 2 \times 10^{20} \text{ v/s. } \frac{dN_p}{dt}$$



$$\frac{dNr}{dt} = \sigma \cdot \frac{dN_p}{dt} s \downarrow n_b \cdot dS$$



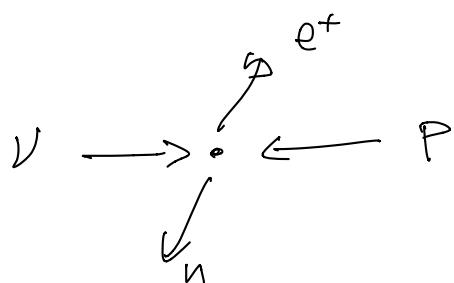
$$\frac{dNr}{dt} = \sigma \cdot \frac{dN_p}{dt} \cdot S \quad N_B$$

$$\frac{dN_p}{dt} = \phi_p = \mathcal{V}_p \cdot n_p$$

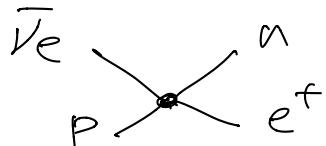
$$= \mathcal{V}_p \cdot \frac{n_p}{V}$$

$$\underbrace{\frac{1}{N_B} \frac{1}{n_p} \frac{dNr}{dt}}_{\sigma} = \sigma \cdot \frac{\mathcal{V}_p}{V}$$

$$\begin{aligned} P(i \rightarrow f) &= P(\bar{\nu}_e + p \rightarrow n + e^+) \\ &= 2\pi |\mu_{f|i}|^2 \rho(E) \end{aligned}$$



$$\mu_{f|i} = -i \int d^3r \psi_f^* H_i \psi_i$$



$$M_{fi} = -i \frac{G_F}{\sqrt{V}} N \Rightarrow |M_{fi}|^2 = \frac{G_F^2}{V^2} |N|^2$$

$$\int \psi_n^* \psi_p d^3r$$

$$\rho(E) = \int \delta(E_f - E_i) dN$$

$$\vec{P}_e + \vec{P}_n = 0 \quad |\vec{P}_e| = |\vec{P}_n| = P^*$$

$$dN_e = \frac{1}{8} \frac{V}{(2\pi)^3} C \pi P^*^2 dP^*$$

$$E_f = E_e^* + E_n^* = \sqrt{m_e^2 + P^*^2} + \sqrt{m_n^2 + P^*^2}$$

$$\delta(E_f - E_i) dP^* = \delta(E_f - E_i) \underbrace{\frac{dP^*}{dE_f}}_{dE_f} dE_f$$

$$\frac{dE_f}{dP^*} = \frac{RP^*}{2E_e^*} + \frac{RP^*}{2E_n^*} = \beta_e^* + \beta_n^* = \beta_f^*$$

$$\rho(E) = (-) P^*^2 \frac{1}{\beta_f^*}$$

$$\Gamma \frac{dp}{dt} = (2\pi) \frac{G_F^2}{\chi^2} |N|^2 \Gamma (-) P^*^2 \frac{1}{\beta_f^*}$$

$$\Gamma(\bar{\nu}_e + p \rightarrow e^+ \tau^-) \propto \frac{1}{8\pi} G_F^2 |N|^2 \frac{1}{\beta_P} \frac{1}{\beta_f^*} P^*^2$$

$$|N|^2 \approx 1 \quad \beta_P = \frac{P_V}{E_V} = 1 \quad \beta_f^* \approx 1$$

$$\Gamma \approx G_F^2 P^*^2 \quad G_F \approx 10^{-5} \text{ GeV}^{-2}$$

$$[G_F^2] \approx E^{-4} = L^4 \quad [P^*^2] \equiv E^2 = L^{-2}$$

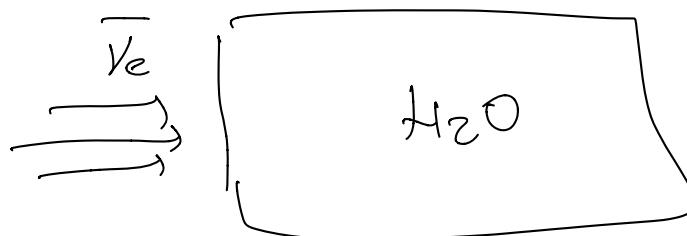
$$[\sigma] \approx [L]^2$$

$$G_F \approx 1.6 \times 10^{-50} \text{ GeV}^{-4}$$

$$1 = 200 \text{ MeV} \times f_{\text{mu}} = 0.2 \text{ GeV} \times f_{\text{mu}}$$

$$1 \text{ fm} = \frac{1}{0.2} \text{ GeV}^{-1} = 5 \text{ GeV}^{-1}$$

$$\Gamma \approx 10^{-44} \text{ cm}^2 \left(\frac{P^*}{\text{MeV}} \right)^2$$



$$\frac{dN_r}{dt} = \sigma \cdot n \cdot dx \cdot \frac{dN_p}{dt}.$$

$$\frac{\frac{dN_r}{dt}}{\frac{dN_p}{dt}} = P(\text{cut}) = \sigma \cdot n \cdot dx.$$

$$P(x) = e^{-\underbrace{\sigma \cdot n \cdot x}_{\lambda}}$$

$\frac{1}{\lambda} = \sigma \cdot n$ length mean free interaction

$$H_2O : n_b = \frac{\rho}{A} N_A \tau = \frac{1}{18} \times 6 \times 10^{23} \text{ cm}^{-3}$$

$$\approx 3.2 \times 10^{22} \text{ cm}^{-3}$$

$$\frac{1}{\lambda} \propto \sigma \cdot n = 10^{-44} \text{ cm}^2 \times 3 \times 10^{22} \text{ cm}^{-3} \left(\frac{P}{\mu\text{ev}} \right)^2$$

$$= 10^{-22} \text{ cm}^{-1}$$

$$\lambda \approx 10^{22} \text{ cm.}$$

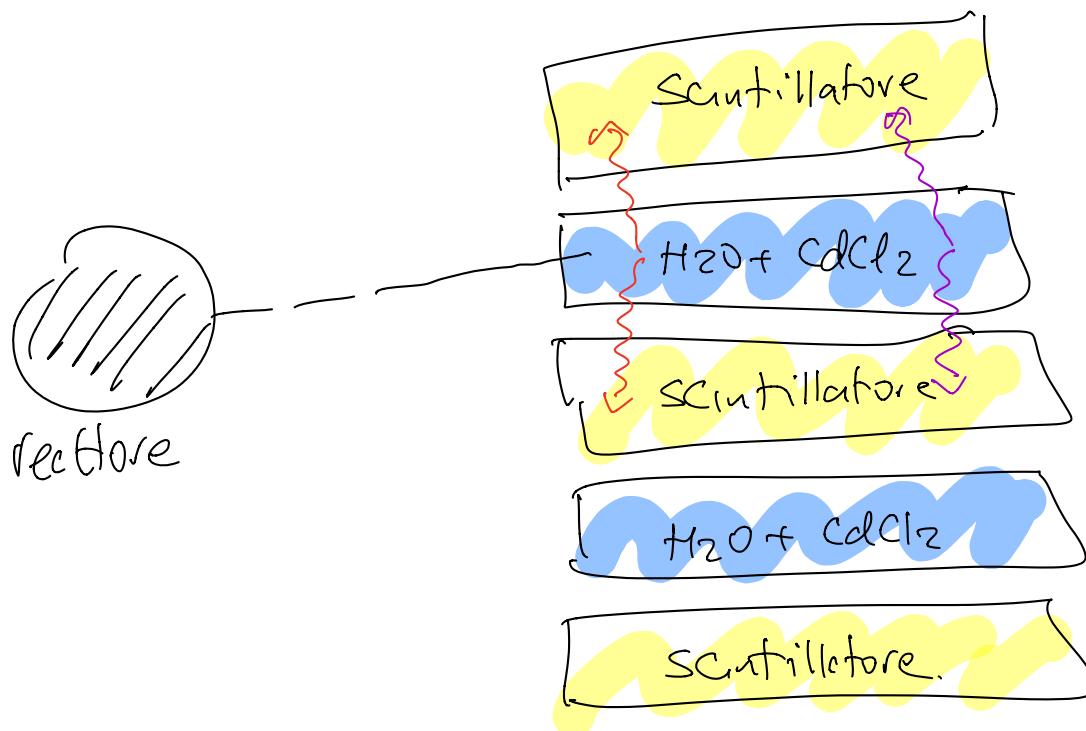
$$R_{\text{Mars-Milano}} \quad 10^3 \text{ km} = 10^6 \text{ cm}$$

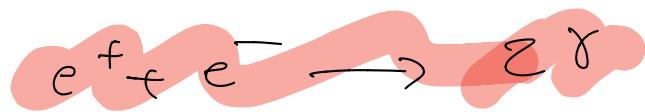
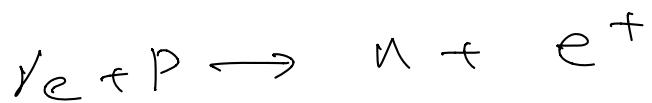
$$= 10^8 \text{ cm.}$$

$$T_{\text{Terre-Sole}} = 150 \text{ km} \times 10^6 = 10^8 \times 10^3 \times 10^2 \text{ cm}$$

$$= 10^{13} \text{ cm.}$$

$$\frac{dN_r}{dt} = \sigma \cdot n \frac{dP}{dt} d = \sigma \cdot N_B \phi_P$$

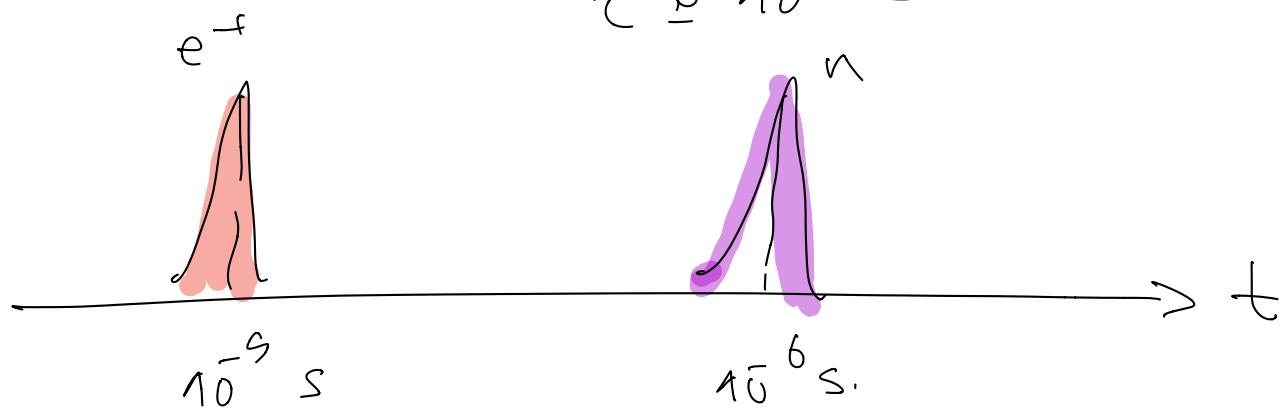




$$\tau \approx 10^{-9} \text{ s.}$$



$$\tau \approx 10^{-6} \text{ s}$$



$$\Delta T \quad N_{\text{tot}} \leq 567 \quad \text{Reaction accs.}$$

$$N_{\text{tot}} = N_S + N_B$$

$$\Delta T \quad N_{\text{tot}} = 20^9 \quad \text{reaction spnts.}$$

$$N_{\text{tot}} = N_B$$

$$N_S = \frac{N_{\text{tot}} - N_B}{SN_B} = \frac{567 - 20^9}{\sqrt{20}}$$