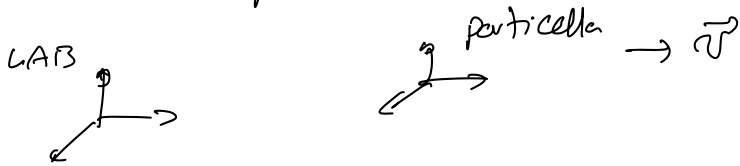


G-impulso:  $\underline{P} = (E, \vec{P})$   $E = \gamma m$   $\vec{P} = \gamma m \vec{v}$

$$u^\mu = \underline{u} = \frac{dx^\mu}{d\tau} = \left( \frac{dt}{d\tau}, \frac{d\vec{x}}{d\tau} \right) \quad \gamma = \frac{1}{\sqrt{1-\beta^2}} \quad \beta = \frac{v}{c}$$

$\tau$ : tempo nel rif. solido con la particella.



$$t_{LAB} \equiv t = \gamma \tau \Rightarrow dt = \gamma d\tau$$

$$\underline{u} = \left( \gamma, \gamma \frac{d\vec{x}}{dt} \right) = \left( \gamma, \gamma \vec{v} \right) = \left( \gamma, \gamma \vec{\beta} \right)$$

$$\vec{\beta} = \left( \frac{v_x}{c}, \frac{v_y}{c}, \frac{v_z}{c} \right)$$

$$E = \gamma m \quad \vec{P} = \gamma m \vec{v} \quad (\text{dalla lagrangiana}).$$

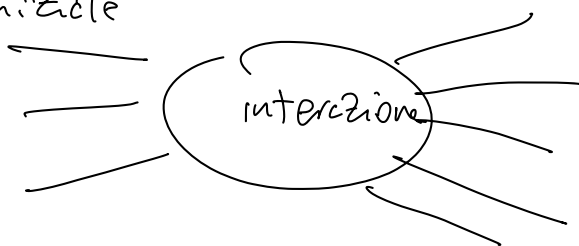
$$\underline{P} = m \underline{u}$$

legge di conservazione di G-impulso:

$$|i\rangle \longrightarrow |f\rangle$$

$$\underline{P}_{in} = \sum_{j=1}^{N_{in}} \underline{P}_j^{in} = \sum_{K=1}^{N_{fn}} \underline{P}_K^{fn} = \underline{P}_f$$

stato iniziale finale



$$E^2 = p^2 + m^2 \Rightarrow E = \pm \sqrt{p^2 + m^2}$$

Eq. Schrödinger:  $H\psi = E\psi$

$$p \rightarrow -i\vec{\nabla}$$

$$E \rightarrow i\frac{d}{dt}$$

$$i\frac{d}{dt}\psi = H\psi$$

mescola  $\frac{d}{dt}$  con  $\nabla^2$

$$E^2 = p^2 + m^2 \implies \left(i \frac{d}{dt}\right)^2 = (-i \vec{\nabla})^2 + m^2$$

$$-\partial_t^2 \psi = (-\nabla^2 + m^2) \psi$$

$$(\partial_t^2 - \nabla^2 + m^2) \psi(x,t) = 0$$

Equazione di

Klein-Gordon

per particelle senza

spin.

$$(\square + m^2) \psi(x,t) = 0$$

$$\square = \partial_t^2 - \nabla^2$$

$$\psi(x,t) = A e^{-i \underline{p} \cdot \underline{r}} = A e^{-i E t} e^{+i \vec{p} \cdot \vec{x}}$$

$$E = \pm \sqrt{p^2 + m^2}$$

Equazione di Dirac 1928

$$(i \gamma^\mu \partial_\mu - m) \psi(x,t) = 0$$

$\psi$ : spinore

$\gamma^\mu$ : matrici  $4 \times 4$  di Dirac

$\mu = 0, 1, 2, 3$

$$\partial_\mu: \frac{\partial}{\partial x^\mu}$$

$$\underline{\mathbb{I}} = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}$$

$$\gamma^0 = \gamma_0 = \begin{pmatrix} \mathbb{I} & 0 \\ 0 & -\mathbb{I} \end{pmatrix}; \quad \gamma^i = -\gamma_i = \begin{pmatrix} 0 & \sigma_i \\ -\sigma_i & 0 \end{pmatrix}$$

$\sigma_i$ : di Pauli

Rapp. di Pauli-Dirac

$$\gamma^0 = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} \quad \gamma^1 = \begin{pmatrix} 0 & 0 & 0 & 1 \\ 0 & 0 & 1 & 0 \\ 0 & -1 & 0 & 0 \\ -1 & 0 & 0 & 0 \end{pmatrix}$$

$$\gamma^2 = \begin{pmatrix} 0 & 0 & 0 & -i \\ 0 & 0 & i & 0 \\ 0 & i & 0 & 0 \\ -i & 0 & 0 & 0 \end{pmatrix} \quad \gamma^3 = \begin{pmatrix} 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & -1 \\ -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \end{pmatrix}$$

$$+ \sqrt{p^2 + m^2}$$

$$\psi(x^\mu) = u^{(r)} e^{-i p^\mu x_\mu} \quad [\text{positive energy solutions}] \rightarrow$$

$$\psi(x^\mu) = v^{(r)} e^{i p^\mu x_\mu} \quad [\text{negative energy solutions}] \rightarrow - \sqrt{p^2 + m^2}$$

$$u^{(r)}(\mathbf{p}) = \frac{1}{\sqrt{E_p + m}} \begin{pmatrix} (E_p + m)\chi^{(r)} \\ (\boldsymbol{\sigma} \cdot \mathbf{p})\chi^{(r)} \end{pmatrix}$$

$$\chi^{(1)} = \begin{pmatrix} 1 \\ 0 \end{pmatrix} ; \quad \chi^{(2)} = \begin{pmatrix} 0 \\ 1 \end{pmatrix}$$

$$v^{(r)}(\mathbf{p}) = \frac{1}{\sqrt{E_p + m}} \begin{pmatrix} (\boldsymbol{\sigma} \cdot \mathbf{p})\chi^{(r)} \\ (E_p + m)\chi^{(r)} \end{pmatrix}$$

$$e^{-iEt}$$

$$E > 0$$

particelle

$$e^{-i(-E)(-t)}$$

$$E < 0$$

anti-particelle

(propagate indietro nel tempo)

$$e^{-} \quad p$$



esp.  
di Thompson  
1897

$$e^{+}$$

positrone  
Anderson.

$$\bar{p}$$

anti-protone

Segrè - Chamber pin.

Interazioni: EM, forte, debole

particelle elementari:

$$\begin{matrix} q = -1 \\ q = 0 \end{matrix} \begin{pmatrix} e^{-} \\ \nu_e \end{pmatrix} \quad \begin{matrix} \text{muone} \\ \mu^{-} \\ \nu_{\mu} \end{matrix} \quad \begin{pmatrix} \tau^{-} \\ \nu_{\tau} \end{pmatrix}$$

neutrino  
elettronico

muonico

leptoni (fermioni)  $S = 1/2$

interagiscono

solo con EM

e debole

$p$   
 $n$  : neutrone  
barioni  
fatti da 3 quark.

adroni : anche interazione  
forte.

ma non elementari.

$q = +\frac{2}{3}e$   
 $q = -\frac{1}{3}$

$\begin{pmatrix} \text{up} \\ \text{down} \end{pmatrix}$	$\begin{pmatrix} \text{charm} \\ \text{strange} \end{pmatrix}$	$\begin{pmatrix} \text{top} \\ \text{bottom/beauty} \end{pmatrix}$	quark non esistono liberi ma solo confinati negli adroni
			(Fermioni) $s = 1/2$

protoni:  $uud$

$$q_p = \frac{2}{3} + \frac{2}{3} - \frac{1}{3} = +\frac{2}{3} = 1$$

neutrone:  $udd$

$$q_n = +\frac{2}{3} - \frac{1}{3} - \frac{1}{3} = 0$$

Flavor dei quark: 6

3 colori per ciascuno quark  $r, g, b$

mesoni:  $q, \bar{q}$

anti-quark:  $\bar{q}$

$$\begin{array}{l}
 q = +\frac{1}{3} \\
 q = -\frac{2}{3}
 \end{array}
 \begin{pmatrix} \text{anti-down} & \bar{d} \\ \bar{u} \end{pmatrix}
 \begin{pmatrix} \bar{s} \\ \bar{c} \end{pmatrix}
 \begin{pmatrix} \bar{b} \\ \bar{t} \end{pmatrix}$$

Adroni: interazione forte + EM + debole

Baroni:  
( $q_i q_j q_k$ )

$q_i \bar{q}_j$   
mesoni

$i, j, k: u, d, c, s, b$

$$|\underline{p}|^2 = \underline{E}^2 - p^2 = m^2$$

invariante di Lorentz  
in tutti i riferimenti

$$|\underline{p}|^2 = \underline{p} \cdot \underline{p}$$

$\phi: s\bar{s}$

$$\begin{array}{l}
 \pi^+ : u\bar{d} \\
 \pi^- : \bar{u}d \\
 \pi^0 = \frac{\bar{u}u + d\bar{d}}{\sqrt{2}}
 \end{array}
 \left. \vphantom{\begin{array}{l} \pi^+ \\ \pi^- \\ \pi^0 \end{array}} \right\} \text{pioni}$$

$$\begin{array}{l}
 \text{Kaon: } s\bar{d} \quad K^0 \\
 \quad \quad s\bar{u} \quad K^- \\
 \quad \quad \bar{s}u \quad K^+ \\
 \quad \quad \bar{s}d \quad \bar{K}^0
 \end{array}$$

Bosoni :

colore  $\gamma$   $S=1$   $m=0$ .

debole bosoni:  $W^\pm, Z^0$

$q_c = 0, \pm 1$   $S=1$

gluoni: 8 gluoni  $S=1, q=0$   
mediatori di interazione forte  
 $q_{\text{colore}} \neq 0$ .

Bosone di Higgs:  $S=0$

Ruolo di H è generare massa per tutte le particelle.

Decadimento delle particelle instabili

$e^-$ ,  $p$  : stabili

neutrone instabile

$$n \rightarrow p + e^- + \bar{\nu}_e$$

$$\pi^+ \rightarrow \mu^+ \nu_\mu$$

$$e^+ \nu_e$$

$$\mu^- \rightarrow e^- \bar{\nu}_e \nu_\mu$$

Decadimento:  $1 \rightarrow 2 + 3 + \dots + N$

$t=0$   $N=N_0$  particelle instabili identiche

probabilità di decadimento per unità di tempo  $\lambda$

-  $\lambda$  non dipende dal numero di particelle

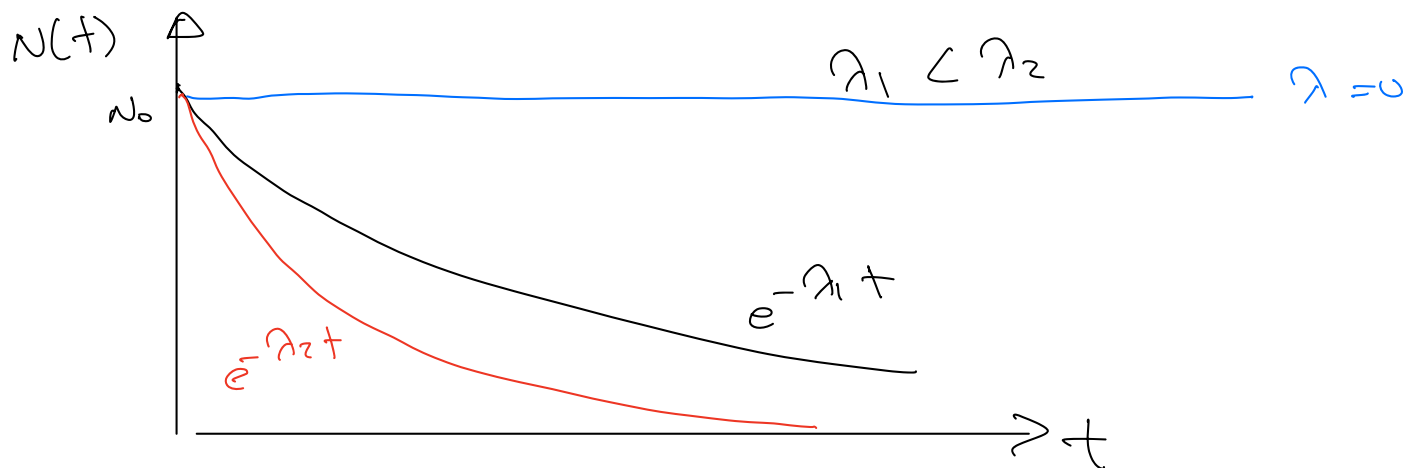
-  $\lambda$  non dipende dal tempo

Caratteristica intrinseca delle particelle.

$$t: \quad dN = -(\lambda dt) N$$

$\lambda$  prob. di decad. in tempo  $dt$

$$\frac{dN}{N} = -\lambda dt \Rightarrow N(t) = N(0) e^{-\lambda t}$$



$$\frac{N(t)}{N(0)} = e^{-\lambda t}$$

vita media propria  
Definizione

$$\langle t \rangle = \frac{\int_0^{\infty} t e^{-\lambda t} dt}{\int_0^{\infty} e^{-\lambda t} dt} = \frac{1}{\lambda}$$

$$\langle t \rangle = \frac{1}{\lambda} \equiv \tau \quad \text{caratteristica della particella.}$$

$$N(t) = N_0 e^{-t/\tau}$$

$$N(t=\tau) = N_0 e^{-\tau/\tau} = \frac{N_0}{e} = 37\% N_0 \quad \text{Sopravvissuti}$$

$$N(t=10\tau) = \frac{N_0}{e^{10}}$$

$$T_{1/2} = N(t') = \frac{N_0}{2} = N_0 e^{-t'/\tau}$$

$$\Rightarrow \frac{1}{2} = e^{-t'/\tau}$$

$$\ln \frac{1}{2} = \ln(e^{-t'/\tau})$$

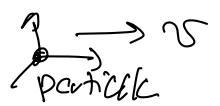
$$-\ln 2 = -t'/\tau$$

$$\boxed{T_{1/2} = \tau \cdot \ln 2}$$

tempo di dimezzamento

$\tau$ : tempo medio nel n°. solide con le particelle.

$$N(t) = N_0 e^{-t^*/\tau} \quad \text{not solidale.}$$



$$LAB: N_0 e^{-\frac{t_{LAB}}{\tau}}$$

$$\text{muoni } \mu^-: \tau = 2.2 \mu\text{s} = 2.2 \times 10^{-6} \text{ s}$$

$$p = 100 \text{ GeV} = 100 \times 10^9 \text{ eV}$$

$$m_\mu = 106 \text{ MeV} = 100 \times 10^6 \text{ eV}$$

$$\gamma = \frac{E}{m}$$

$$E^2 = p^2 + m^2 = (10^{11} \text{ eV})^2 + (\cancel{10^8})^2$$

$$t_{LAB} = \frac{X_{LAB}}{v} = \frac{X_{LAB}}{\beta c}$$

$$e^{-\frac{t_{LAB}}{\tau}} = e^{-\frac{X_{LAB}}{\beta \gamma c \tau}} = e^{-\frac{X_{LAB}}{L}}$$

$$\beta \gamma = \frac{|\vec{p}|}{m} = \frac{100 \text{ GeV}}{100 \text{ MeV}} = 1000$$

$$\begin{aligned} \beta \gamma c \tau &= 10^3 \times 3 \times 10^8 \frac{\text{m}}{\text{s}} \times 2.2 \times 10^{-6} \text{ s} \\ &= 6.6 \times 10^5 \text{ m} = 660 \text{ km} \end{aligned}$$

$\beta \gamma c \tau = \langle l \rangle$  lunghezza media di Soprow.  
in Laboratorio.

$$N_\mu(15 \text{ km}) = N_0 e^{-15/660}$$

$$200 \text{ MeV fm} = 1.$$

$$1 \text{ fm} = \frac{1}{200} \frac{1}{\text{MeV}}$$

Spessore dell'atmosfera in MeV

$$\begin{aligned} 15 \text{ km} &= 15 \times 10^3 \text{ m} \\ &= 15 \times 10^3 \times 10^{15} \text{ fm} \end{aligned}$$

