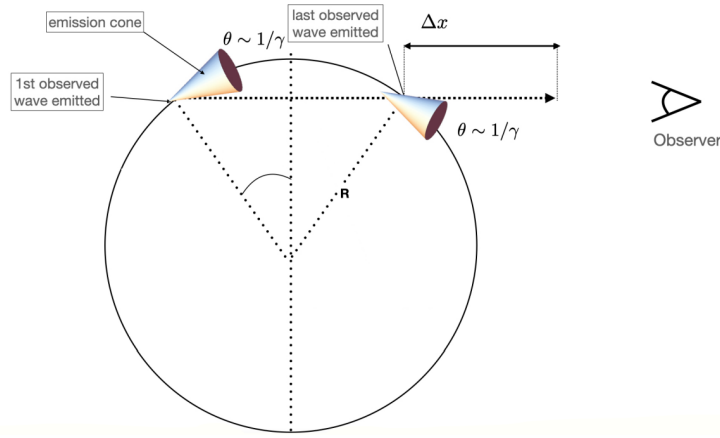


Homework 9

Problem 1

A particle is moving relativistically with Lorentz factor $\gamma \gg 1$ along a circle of radius R , as shown in the figure below. By aberration, a particle emits mostly within an angle $\sim 1/\gamma$ with respect to its instantaneous velocity. Estimate a distance Δx and corresponding time $\Delta x/c$ between first and last moment that the observer sees the particle. (An electromagnetic signal is first emitted at the "first observed wave emitted", and propagates with speed c . The particle nearly catches up with its own radiation, and so at the point "last observed wave emitted", it is behind the first emitted wave by Δx .)



Solution. Consider a particle moving in a circle at relativistic speed with Lorentz factor is $\gamma \gg 1$. The radiation is beamed within angle $\theta \sim 1/\gamma$.

The arc length $\Delta \ell$ taken by the particle is given by

$$\Delta \ell = 2\theta R \sim \frac{2R}{\gamma}.$$

and the time Δt that passes for the particle after taking the arc length path is

$$\Delta t = \frac{\Delta \ell}{v} \sim \frac{2R}{\gamma v}.$$

Thus, the distance between the first and last moment the observer sees the particle is

$$\begin{aligned} \Delta x &= c\Delta t - v\Delta t \\ &= \frac{2Rc}{\gamma v} - \frac{2R}{\gamma} \\ &= \frac{2Rc}{\gamma v} \left(1 - \frac{v}{c}\right) \end{aligned}$$

and the time between the first and last moment the observer sees the particle is

$$\Delta \tau = \frac{\Delta x}{c} = \frac{2R}{\gamma v} \left(1 - \frac{v}{c}\right).$$

Problem 2 - Magneto-dipolar Emission

A sphere of radius R carries dipolar magnetic field B_0 and rotates with spin frequency Ω . Estimate the emitted power. [Hint: Check Landau & Lifshitz, vol 2, eq. (71.5)]

Solution. Landau and Lifshitz, Volume 2, Eq. 71.5 states that the total radiation I is given by

$$I = \frac{2}{3c^3} \ddot{d}^2 + \frac{1}{180c^5} \ddot{D}_{\alpha\beta}^2 + \frac{2}{3c^3} \ddot{m}^2,$$

where the terms corresponds to dipole radiation, quadrupole radiation, and magnetic dipole radiation, respectively.

To get the power, we have to integrate the radiation. Since we are considering a dipolar magnetic field, we only consider the last term. Additionally, we know that the magnetic field has the form

$$|\mathbf{B}_0| = \frac{\mu_0}{2} \frac{IR^2}{(z^2 + R^2)^{\frac{3}{2}}}.$$

For $r \gg R$, we have

$$\begin{aligned} |\mathbf{B}_0| &= \frac{\mu_0}{2} \frac{IR^2}{z^3} \\ &= \frac{\mu_0}{4\pi} \frac{2I\pi R^2}{r^3} \\ &= \frac{\mu_0}{4\pi} \frac{2IA}{r^3} \\ &= \frac{\mu_0}{4\pi} \frac{2}{r^3} m, \end{aligned}$$

where A is the area of the loop. From the previous, we can write $\mathbf{m} \sim r^3 \mathbf{B}_0$.

The magnetic dipole moment can be expressed in terms of its angular position as

$$m = m_0 \cos(\Omega t) \hat{\mathbf{n}},$$

where m_0 is the magnitude of the dipole, Ω is the spin frequency, and $\hat{\mathbf{n}}$ is along \hat{z} . Taking the second time derivative, we have

$$\ddot{m} = -m_0 \Omega^2 \cos(\Omega t) \hat{z} = -\Omega^2 m.$$

To calculate power, we only need the magnitude of the magnetic dipole radiation; hence, $|\ddot{m}| = \Omega^2 m$. Thus, we have

$$\begin{aligned} P &= \frac{2}{3c^3} |\ddot{m}^2| \\ &= \frac{2}{3c^3} (\Omega^2 m)^2 \\ &= \frac{2}{3c^3} (\Omega^2 r^3 \mathbf{B}_0)^2 \\ &= \frac{2}{3c^3} \Omega^4 r^6 \mathbf{B}_0^2. \end{aligned}$$

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Problem 3

In the problem above, the mass of the sphere is M (so that its moment of inertia is $I \approx (2/5)MR^2$). An observer can measure spin Ω and the rate of change $\dot{\Omega}$. Find the magnetic field on the surface in terms of M, R, Ω and $\dot{\Omega}$.

Solution. The rotational energy is given by

$$E = \frac{1}{2}I\Omega^2.$$

We know that $\frac{dE}{dt} = -P$, so we calculate the time derivative of the energy. We have

$$\begin{aligned}\frac{dE}{dt} &= I\Omega\dot{\Omega} = -P \\ &= \frac{2}{3c^3}\Omega^4r^6\mathbf{B}_0^2.\end{aligned}$$

Thus, the magnetic field on the surface is

$$\begin{aligned}|\mathbf{B}_0| &= \sqrt{\frac{3c^3I\Omega\dot{\Omega}}{2\Omega^4r^6}} \\ &= \frac{c}{r^3}\sqrt{\frac{3cI\dot{\Omega}}{2\Omega^3}} \\ &= \frac{c}{\Omega r^3}\sqrt{\frac{3cI\dot{\Omega}}{2\Omega}}.\end{aligned}$$

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