

1 Digital Elevation Model (DEM) uncertainty and
2 hazard analysis using a geophysical flow model

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14 **Abstract**

15 To come soon...

16 **1 Introduction**

17 Numerous workers have explored the use of computational fluid dynamics to
18 produce volcanic hazards maps for a variety of phenomena at a number of vol-
19 canoes [Baxter and Gresham, 1997, Calvache et al., 1997, Sheridan et al., 2005].
20 Hazards maps for ground-hugging flows, such as pyroclastic density currents and
21 lava flows are constructed using a digital representation of the terrain [Takahashi
22 and Tsujimoto, 2000, Dalbey et al., 2008]. Usually these terrain representations
23 are digital elevation models (DEMs). For this type of study, elevation is rightly
24 recognized as the most essential and fundamental of variables in geographic
25 information analysis [Atkinson, 2002].

26 However, a digital representation of a terrain surface is an approximation of
27 reality and is subject to some degree of error. The error usually is not known
28 in terms of both magnitude and spatial distribution. There are in fact large
29 uncertainties associated with the construction of DEMs. In [Wechsler and Kroll,
30 2006] it was shown that DEMs contain errors derived from a variety of sources:
31 sampling, measurement and interpolation, and these errors cannot always be
32 well estimated.

33 The most important part of DEM error propagation analysis is the appro-
34 priate characterization of the error within the DEM itself, including information
35 about its distribution and spatial structure [Shortridge, 2001]. DEM vendors
36 generally provide users with a measure of vertical accuracy in the form of the
37 root mean squared error (RMSE) statistic. However many papers have reported
38 on the limitations of a single value of accuracy, stressing that DEM error is
39 spatially variable and autocorrelated [Wechsler and Kroll, 2006, Darnell et al.,
40 2008]. Also the magnitude of the DEM error is closely related to the charac-
41 teristics of the terrain surface. For example, slope will influence interpolation
42 procedures.

43 DEM error propagation analysis was introduced to the GIS community in
44 the early 1990s. In the work of Heuvelink et al. [1990], error propagation in
45 calculating slope and aspect was represented using Monte Carlo simulation and
46 Taylor series approximation. It was shown that standard deviations of slope
47 and aspect were higher than expected. The effect of error in the DEMs on the
48 erosion models was emphasized. A method used by Weng [2002] in quantifi-
49 cation of the uncertainty of DEMs was to create various DEMs using different
50 interpolation methods and to examine the RMSE from the source map, sam-
51 pling and measurement error, and the interpolation process. It was concluded
52 that RMSE can be used as a general indicator of DEM uncertainty. In recent
53 literature, DEM error without spatial autocorrelation was considered to be a
54 worst-case scenario [Heuvelink et al., 1989, Van Niel et al., 2004, Oksanen, 2006],
55 but no analysis based on terrain morphology and the effect of different DEMs
56 was done. Wechsler and Kroll [2006] developed four different methods for rep-
57 resenting the spatial dependence of error through random fields to assess the
58 effect on topographic parameters of the DEM uncertainty. The study showed
59 that uncertainty in the DEM is manifested at higher elevations in local steeper
60 slopes, on both slope and elevation maps. Florinsky [1998] showed that the ef-
61 fect of DEM uncertainty on the accuracy of slope and aspect estimation cannot
62 be determined by using data from topographic maps or field surveys, because
63 accurate derivatives cannot be determined.

One key feature of spatial data is the autocorrelation of observations in space. Generally, spatial autocorrelation refers to the correlation between the same attribute at two locations. Observations in close spatial proximity tend to be more related than are observations at larger distance or separation. Errors in spatial data (such as incorrect elevation values assigned to a point) are spatially autocorrelated. The effect of correlated DEM error has been investigated in the literature [Fisher, 1991, Goodchild et al., 1992]. It was shown that not only is error spatially variable throughout a DEM, but within the elevation model the error value of an individual grid cell is related to the error in neighboring cells. Unfortunately, DEM providers do not include information regarding the spatial dependence or spatial relationship of errors.

Stochastic modeling uses stochastic conditional simulation to generate multiple equally likely representations of an actual terrain surface. Hunter and Goodchild [1997], Ehlschlaeger and Shortridge [1996] computed a normal distribution of maps or realizations to reproduce the spatial autocorrelation encountered in the original error surface, filtered using a Gaussian convolution filter, with kernel sizes derived from autocorrelation analysis of the original error surfaces.

Various researchers have applied stochastic techniques to evaluate uncertainty in DEM data. Ehlschlaeger and Shortridge [1996] stochastically simulated error in a DEM to evaluate the impact of DEM uncertainty on a least-cost-path application. Hunter and Goodchild [1997] investigated the effect of simulated changes in elevation at different levels of spatial autocorrelation on slope and aspect calculations. Hebel and Purves [2008] produced uncertainty surfaces to show the impact of DEM uncertainty on an ice sheet model. Darnell et al. [2008] developed a fuzzy framework to examine the probable and possible uncertainties in classifying landslide hazard.

The aim of this paper is to quantify the variation in the output of a computational flow model for block and ash flows, when the model inputs are given as a range of possible uncertainties/values. In particular, we focus on assessing the influence of DEM uncertainties, along with uncertainties in initial size and location of the avalanche, and the internal and bed friction angles. There is uncertainty in all of these inputs and it can be represented using either field data or stochastic methods. The distribution or the range of the parameters can be obtained from laboratory and field instruments for friction angles, and historical records of flow frequency and magnitude for size of the initial failure. Stochastic methods are used to assess the uncertainties in the DEMs. In particular, a perturbation of the elevation based on the measured error model, and also an unconditional stochastic simulation are used [Ehlschlaeger and Short-

102 ridge, 1996]. Both methods generate multiple likely representations of the actual
103 terrain, while the second one accounts for the spatial autocorrelation between
104 elevation points. The effect of DEM uncertainty and its impact on the output
105 model is analyzed by constructing a hazards map and performing a "proba-
106 bility analysis" for two volcanoes with different morphology: Galeras Volcano,
107 Colombia, and Mammoth Mountain, CA, USA.

108 We adapt here an approach based largely on the method of Ehlschlaeger and
109 Shortridge [1996], which uses the difference between two independent DEMs to
110 train a Gaussian model of error.

111 2 Methodology

112 In previous work [Stefanescu et al., 2010], the effect of DEMs on the output of
113 TITAN2D was investigated by comparing the output (maximum flow depth over
114 the entire simulated time) from different DEMs. These DEMs were obtained
115 from different techniques at different resolution. Two types of analysis were
116 performed: a qualitative analysis and a statistical analysis. The qualitative
117 analysis consisted of a comparison of the footprint of the flow, extended to a pixel
118 based classification. The pixels were classified into inundated and non-inundated
119 classes. For the statistical analysis we performed a Kolmogorov – Smirnov test
120 to check if two output datasets differed significantly. The conclusion was that
121 for moderate and smaller scale flows, use of different DEMs affects computation
122 of accurate footprints of the flow.

123 This conclusion motivated us in examining the effect of DEM uncertainty by
124 creating a model of the error and sampling it to create an ensemble of possible
125 terrains. The flow simulation is then run on every member of this ensemble.

126 Naive, cell-by-cell approaches to treating DEM uncertainty quickly lead to
127 the use of thousands if not millions of random variables, resulting in a computa-
128 tionally infeasible problem. On the other hand, the error model described above
129 can be parameterized with one or two random variables. The parametrization
130 methods are based on the assumption that the available DEM is a representation
131 of the terrain to which errors have been added because of instrumental uncer-
132 tainty. Therefore, the DEM can be assumed to be one of an infinite number of
133 elevation realizations.

134 2.1 Method 1

In this paper, we have available two "types" of DEMs of each mountain, which are used in creating DEM-to-DEM difference maps. Different realizations of the terrain were constructed by adding to one DEM – considered to represent the "true" elevation – a "random" perturbation. Since any two types of DEMs are obtained using different techniques, the difference between them can be added to that which is assumed to be the "true" DEM to give us a set of possible DEMs. Thus, the resulting realizations are consistent with the available set of DEMs. Randomness in the perturbations is created by multiplying the difference map with a normally distributed factor between 0 and 1.

$$R = M + \epsilon \cdot Diff \quad (1)$$

135 where R is a realization of the terrain, M is the DEM that best represents
 136 the terrain (the "true" DEM), $Diff$ is the difference map and ϵ is a random
 137 variable. In this way we can define a set of DEM realizations using only one
 138 random variable.

139 2.2 Method 2

The main idea of Method 2 derives from geostatistics, where vectors of a random variable $Z(\mathcal{U})$ are used to characterize unknown values, generated by probability distributions which characterize the uncertainty of the random variables. Z represents a continuous random field and \mathcal{U} represents the area covered by the random field. The random field function used is implemented in the function *r.random.surface* [Ehlschlaeger and Goodchild, 1994] of GRASS GIS [Mitasova et al., 1996], and generates fields containing cells with a normal distribution (mean of 0.0 and variance of 1.0). The random field function derives its spatial dependence by the use of a filter function. The following equation is used to generate the random fields:

$$Z(\mathcal{U}) = \frac{\sum_v w_{u,v} \epsilon_v}{\sqrt{\sum_v w_{u,v}^2}}, \quad u \in \mathcal{U}, v \in \mathcal{V} \quad (2)$$

$$w_{u,v} = \begin{cases} 1 & : d_{u,v} \leq F \\ \left(1 - \frac{d_{u,v} - F}{D - F}\right)^E & F < d_{u,v} \leq D, u \in \mathcal{U}, v \in \mathcal{V} \\ 0 & : d_{u,v} > D \end{cases} \quad (3)$$

140 where $Z(\mathcal{U})$ is the random field, \mathcal{V} is the set of points affecting the area \mathcal{U} ,
 141 $w_{u,v}$ is the spatial autocorrelative effect between points u and v , v is a random

variable with a mean of 0 and variance of 1, $d_{u,v}$ is the distance between u and v , D is the minimum distance of spatial independence, E is the distance decay exponent, and F is a parameter that helps to capture the terrain features.

Random fields are calibrated to the spatial continuity of the field being simulated using a correlogram function. The method used to fit the correlogram and to choose the best descriptive parameters of the random field (the minimum distance of spatial independence, the correlated distance decay exponent and the filter parameter) is a weighted least-square estimator implemented in GRASS's *r.lags.difference*. After running hundreds of tests with multiple combinations of D , E and F , we found the best random field to fit the error map characteristics based on the minimum sum of least squares difference between the error field's correlogram and a target correlogram. *xxFigure 1 shows the error map correlogram compared to several correlograms closely fitting the error correlogram. It can be seen the dependence of the correlation coefficient with the distance. xx*

The correlogram model was then used with sequential Gaussian simulation to generate N error map realizations. Each error realization was added to the "true" DEM to generate equally probable realizations of the topography for the error structure of a DEM under consideration:

$$R(U) = m(U) + m(m(T)) + (m(s^2(T)) \cdot \epsilon) \cdot Z(U) \quad (4)$$

where $R(U)$ is a realization of an elevation dataset $m(U)$, T is a group of sets of spatially uncorrelated sample points, ϵ is a random variable with mean 0.0 and variance 1.0, and $Z(U)$ is the random field which captures the autocorrelation between points. The mean and the standard deviation are determined from randomly drawn, spatially independent points scattered across the error map. The error map was generated by subtracting the lower quality DEM from the "true" DEM, and characterizes the error of the lower quality DEM at each point.

2.3 DEM realizations

Many DEM users are aware that DEM uncertainty affects the results of their application, however, in most cases the DEM is accepted as the true representation of the earth's surface. In this section, the two methods for generating multiple realizations of the terrain are presented for both Galeras Volcano and Mammoth Mountain to test whether it is safe to assume that the representation of topography is acceptable as it is.

The motivation for creating realizations of the DEM was to be able to use the DEM along with other uncertain parameters as uncertain inputs for the

173 calculation of a hazard map using the computational fluid dynamical model
174 TITAN2D. One of our working hypotheses is that the DEM contributes a sig-
175 nificant proportion of the variance in simulated flow, hence hazard map output.
176 For sampling the input parameter space, a Latin Hypercube Sampling (LHS)
177 was implemented.

178 For Galeras Volcano, two test DEMs at 30 m spacing were considered for
179 our analysis. The SRTM (Shuttle Radar Topography Mission) 30m DEM was
180 derived by spline interpolation from a 90m DEM of southern Colombia using
181 radar data collected in 2000, while the ASTER (Advanced Spaceborne Thermal
182 Emission and Reflection Radiometer) DEM was calculated at the Jet Propulsion
183 Laboratory using orthorectified imagery from 12 January 2010 (Fig. 2 a). The
184 ASTER dataset was used as a surrogate for the “true” elevation while the SRTM
185 dataset was used in creating the error model.

186 Two 30-m resolution DEMs derived from independent techniques were used
187 for Mammoth Mountain. A TOPSAR dataset was considered to be the “true”
188 elevation, while an SRTM dataset was used in creating the error map. A rect-
189 angular area of approximately 42 kilometers² was defined within the TOPSAR
190 and SRTM DEMs (Fig. 2 b).

191 For Method 1, sixty-four (64) DEM realizations were created and used as
192 input parameters for the TITAN2D simulator along with uncertain parameters
193 presented in ???. The input space is defined by seven parameters.

194 For Method 2, realizations of the terrain surface were created by taking
195 into consideration the spatial autocorrelation of the error. The error map was
196 obtained by subtracting the elevation of a given DEM from the “true” elevation
197 at each location. The correlogram for the difference map was calculated to
198 determine the range of spatial dependence of elevation points. We found that
199 spatial dependence persisted above a threshold value of the correlogram cross-
200 correlation coefficient of 0.4 to a distance of 2500 meters for Galeras and 2100
201 meters for Mammoth. To determine the probability distribution function (pdf)
202 for the stochastic simulation, 91 sets of spot locations were selected from the
203 map, each set containing 91 points, all pairs of points were separated by more
204 than 2500 meters or 2100 meters, respectively. For each DEM, pdf statistics
205 were derived. The random field parameters were chosen after testing more than
206 400 random field parameters for the smallest difference between the error model
207 correlogram and the random field. This occurs when the minimum distance of
208 spatial independence, $D = 2500$; the distance decay, $E = 0.8$, and the filter
209 parameter, $F = 400$ for Galeras and $D = 2100$, $E = 0.7$, and $F = 350$ for
210 Mammoth. A total of 64 equally probable potential elevation surfaces of the

211 area having a 30-m resolution were generated.

212 2.4 Hazard map construction

213 There are numerous ways to create a volcanic hazards map based on compu-
214 tational fluid dynamics modeling. The traditional Monte Carlo method can
215 be used if we assume that uncertainty in model input parameters is the main
216 restriction to our knowledge of future events at a given volcano. This is the
217 case, for example, if we know that block and ash flows are common at a given
218 volcano, but it is difficult to know the size or volume of potential future events.
219 Although Monte Carlo is relatively simple to implement, it converges slowly
220 and is unaffordable computationally because of the number of time-consuming
221 simulations. A single TITAN2D run might take 20 minutes on a single proces-
222 sor. To obtain three-digit accuracy in the expected value of a specified function
223 would require a million runs. One million runs of 20-min calculations running
224 non-stop on 64 processor would take 217 days [Dalbey et al., 2008].

225 Here, we briefly described the use of an hierarchical emulator that signifi-
226 cantly reduces computational cost; a detailed discussion of the methodology can
227 be found in Dalbey [2009], Dalbey et al. [2008]. An emulator can be thought of
228 as a fast surrogate for a single numerical model simulation (a simulator). We
229 describe the process of computing a hazard map for block and ash flows with
230 uncertain model inputs introduced by Dalbey [2009]. Two-level construction of
231 a group or ensemble of emulators is used to include a separation of uncertain
232 inputs and geographic coordinates. The process starts by identifying the model
233 inputs whose uncertainties will drive the process. In our case, the uncertain
234 flow inputs we use are volume and shape, starting location, basal and internal
235 friction angles, and finally topography, as given by the DEM. For the result-
236 ing eight-dimensional parameter input space, a Latin Hypercube Sampling was
237 performed to determine parameter values at which simulations were to be run
238 [Bayarri et al., 2010]. As priors for the emulator, simulation outputs for each of
239 these input parameter vectors were stored at 64 grid points.

240 The variable of interest for our application is the field of maximum flow depth
241 over time for each spatial position, at each of the downsampled input parameter
242 gridpoints. Tessellations of the geographic coordinate space and the parameter
243 input space are constructed (we use Delaunay triangulation). At a designated
244 location, \mathbf{x}^* , of the input parameter plus spatial coordinate space at which the
245 hazard is to be computed, the covering simplex $S_{\mathbf{x}}^*$ of the parameter space is
246 identified, and all nodes of that simplex are enumerated, as are all nodes within

247 a neighborhood (two hops in the tessellation) of the covering simplex nodes. For
 248 each such two-hop node, we tessellate in the spatial coordinates and evaluate
 249 all emulators constructed over these nodes. We average these coordinate space
 250 emulators to (the coordinate components of) \mathbf{x}^* by barycentric weighting; notice
 251 there will be an emulator for each parameter input sample point. Now in the
 252 input parameter space, construct a tessellation of the two-hop nodes and average
 253 the emulators to \mathbf{x}^* by barycentric weighting of the fine-scale emulator. The
 254 emulator is now readily and quickly evaluated for each evaluation. The hazard
 255 map construction can now proceed by treating the emulator as a surrogate for
 256 the simulator in the classical Monte Carlo procedure. For any point in the
 257 domain we can now exercise the emulator to get potential flows and hence
 258 exceedance probabilities.

259 **2.5 TITAN2D and flow simulations**

260 TITAN2D was developed for modeling dry geophysical granular flows, such as
 261 debris avalanches and block and ash flows. Given a digital elevation map speci-
 262 fying the topography of a volcano and the values of input parameters, including
 263 the initial volume of erupted material and the friction angles, TITAN2D calcu-
 264 lates the flow depth and velocity at any location throughout the duration of an
 265 event. The TITAN2D code combines numerical simulations of a natural granu-
 266 lar flow with digital terrain data. It is based on a depth-averaged model for an
 267 incompressible granular materia governed by Coulomb-type friction interactions
 268 [Savage and Hutter, 1989]. The governing equations are obtained by applying
 269 conservation laws to the incompressible continuum, providing appropriate con-
 270 stitutive modeling assumptions, and then taking advantage of the shallowness
 271 of the flows (flows are much longer and wider than they are deep) to obtain
 272 simpler depth-averaged representations [Patra et al., 2005]. The motion of the
 273 material is considered to be gravitationally driven and resisted by both internal
 274 and bed friction. The stress boundary conditions are: no stress at the upper
 275 free-surface and a Coulomb-like friction law imposed at the interface between
 276 the material and the basal surface.

277 The primary factor driving the flow is the component of gravity tangential
 278 to the surface, which depends on a local slope computed from the elevation
 279 data, hence, the criticality of the DEM to the flow computations. The resulting
 280 hyperbolic system of equations was solved using a finite-volume scheme with a
 281 second-order Godunov solver. Although many real geophysical flows — such as
 282 debris flows — are fluidized, in this study we deal only with granular material

that has not been fluidized, such as dome-collapse block and ash flows or rock avalanches initiated by slope instability. The program runs in parallel, using the Message Passing Interface (MPI) to allow communication between multiple processors, increasing computational power, decreasing computational time and allowing use of large datasets. The algorithm uses local adaptive mesh refinement for shock capturing, and dynamic load balancing for the efficient use of computational resources. Topographic data are included in the simulation through a preprocessing routine in which the digital elevation data are imported. TITAN2D performs flow simulations on a DEM of a desired region, the simulation accuracy being highly dependent on the level of the DEM resolution and quality.

Inputs to the code are the size and location of the initial volume, the internal and bed friction and the DEM. Dalbey et al. [2008] presented several methods for characterizing the effect of input data uncertainty on model output. At that time, efficient methods for representing the uncertainty associated with spatial parameters like terrain elevation were not well understood.

2.6 Bayes Linear Method

The straightforward way to account for uncertain inputs and stochastic forcing is a Monte Carlo approach — run many simulations and ‘average’ the results in some fashion. If simulations are expensive to run, this approach is not feasible. To circumvent this difficulty, the statistics community has developed the idea of an emulator. In essence, the emulator is a regression surface based on a representative sample of simulations at selected inputs, accompanied by statistical error bounds. Equipped with this surface, output values at new (untested) input values need not be run. Instead output results can be determined by evaluating the emulator. There are indeed many methods – kriging, metamodels, support vector machines, etc., by which such surrogates may be constructed and there exists a body of literature on the topic [Simpson et al., 2001, Clarke et al., 2005]. One often-used emulator is the GAuSSian Process (GASP) emulator, which assumes the regression has the form of a trend plus a Gaussian [Kennedy and O’Hagan, 2001, Conti and O’Hagan, 2007, O’Hagan, 2006, Bayarri et al., 2010]. To construct a GASP emulator, the covariance structure of the Gaussian must be assumed and parameters determined by Bayesian or partially Bayesian methodology. A fully Bayesian determination of the emulator can be costly, especially if the input data is high-dimensional. Here we use the Bayes Linear method (BLM) [Goldstein, 2007] to construct an emulator. Given prior beliefs

319 (B) of mean and variance, the BLM updates these beliefs conditioned on the
 320 data (D). Note that “data” generally here refers to the output of computation-
 321 ally expensive physics based simulators. Because only the first two moments
 322 of a distribution are determined, the BLM is exact only for Gaussian distri-
 323 butions. As an emulator construction, the BLM update is simpler than a full
 324 GASP construction, but the resulting emulator is comparable. Given the prior
 325 expectation $E[B]$ and variance $var(B)$, the BLM updates are

$$\begin{aligned}
 E_D(B) &= E[B] + cov(B, D)(var(D))^{-1}[D - E[D]] \\
 var_D(B) &= var(B) - cov(B, D)(var(D))^{-1}cov(D, B)
 \end{aligned}
 \tag{5}$$

326 These update formulae can be derived by minimizing the mean square error
 327 $(B - a^T D)^2$ between B and some linear combination of the data. Thus the
 328 BLM update can be viewed as the projection of the set of prior beliefs onto the
 329 span of the data.

330 **3 Implementation**

331 **3.1 Case study I: Galeras Volcano**

332 Galeras Volcano (elevation 4,276 meters), located in southwestern Colombia
 333 ($1^\circ 13.31'$ N and $77^\circ 21.68'$ W), is one of the most active volcanoes on the world
 334 [Hurtado and Cortes, 1997]. Nearly 400,000 people currently live near the vol-
 335 cano; 10,000 of them reside within the zone of high volcanic hazard. Pyroclastic
 336 flows pose a major hazard for this population. The current period of activity
 337 that began in 2004 presents a serious problem for all stakeholders: decision mak-
 338 ers, scientists, public safety officials, and the general population. Computational
 339 modeling has the potential to provide useful information for hazard assessment
 340 and risk mitigation. However, there is a need to evaluate the validity of the
 341 modeling and the quality of the DEMs available for use in such modeling.

342 Galeras is an important volcano for computational flow modeling from both
 343 risk management and scientific perspectives [Calvache et al., 1997]. Forecasts
 344 of volcanic explosions using various geophysical tools [Narvaez et al., 1997] have
 345 occasionally brought public warnings to a high level of alert during the past 20
 346 years. When the alert reaches the highest level, the public are urged to evacuate
 347 some local areas; this occurred as recently as January, 2010. The worst event at
 348 Galeras occurred in 1993, when an eruption killed 9 scientists and journalists
 349 [Baxter and Gresham, 1997].

350 The topography of the volcano presents a problem for creation of a good
 351 DEM. The irregular morphology on a small scale, with steep slopes, narrow
 352 channels, deep gorges and abrupt cliffs poses problems for the creation of ac-
 353 curate topographic models [Ordoñez Villota and Jentzsch, 2000]. In addition,
 354 the current flow hazard map at Galeras is mainly based on the sparse geological
 355 record [Calvache, 1990]. Dense vegetation, deep erosion, successive deposits of
 356 lava and pyroclastic flows hinder the tracing of specific deposits in the field.
 357 The diverse effects of this landscape, as reflected in DEMs created by different
 358 processes and of different scales, must be examined and quantified to determine
 359 the level of confidence that can be placed in model results. Galeras provides a
 360 wide range of topographic features that challenge the use of computational flow
 361 models.

362 **3.2 Case study II : Mammoth Mountain**

363 Mammoth Mountain is a large, geologically young, composite dome volcano lo-
 364 cated on the southwestern rim of Long Valley Caldera, California [Bailey, 1989].
 365 There are many active hazards issues for Mammoth Mountain, including snow
 366 avalanches, rock avalanches and debris flows. In addition, it is intersected by
 367 the Mono-Inyo Craters volcanic chain, which is the most active volcanic region
 368 in the southwestern U.S. If Mono-Inyo type activity occurs on Mammoth Moun-
 369 tain, then domes may form. These new domes would be growing atop a steep
 370 edifice, and therefore could become gravitationally unstable. Given that block
 371 and ash flows occurred at Mammoth Mountain during its older dome growth
 372 stage, there is reason to believe that renewed dome formation would result in
 373 block and ash flow activity. If this is so, then parts of Mammoth Lakes, CA, are
 374 at risk from block and ash flows. Our previous work on Mammoth Mountain
 375 (Stefanescu et al., submitted) was the testing of the hypothesis that different
 376 DEMs result in different model outputs of block and ash flow inundation.

377 **3.3 Model Set-up**

378 In our process to quantify the DEM uncertainties using TITAN2D, a set of
 379 parameters was drawn on which to set the bounds of the input domain: in-
 380 ternal friction angle, basal friction angle, flow volume, location and DEM. The
 381 numerical values for these parameters were chosen to bracket the range of flow
 382 volumes and initial locations, and to be representative of the friction angles that
 383 have been used by other researchers in their computational models. The same
 384 reasonable parameter values were used for both volcanoes, so they do not nec-

essarily represent any optimization for a particular case. The internal friction angle has little effect on the output of the flow models [Dalbey et al., 2008, Sheridan et al., 2005]. Many TITAN users have chosen values of internal friction that range between 15 and 37 degrees with values between 30 and 35 being the most frequent values used [Patra et al., 2005, Murcia et al., 2010]. For our study we use an internal friction angle uniformly distributed between 20 and 25 degrees.

The value of the basal friction angle has a large effect on flow dynamics in the TITAN2D simulations [Patra et al., 2005, Stinton et al., 2006]. Factors that could affect the choice of basal friction angle include the volume of the flow, the type of the pyroclastic flow, the nature of the substrate and the amount of channelization. Murcia et al. [2010] listed the basal friction values chosen by TITAN2D users; they range between 5 and 28 degrees; the mean value being about 15 degrees. We are using a basal friction angle uniformly distributed between 15 and 20 degrees.

Volumes of pyroclastic flows at stratovolcanoes typically cover a few orders of magnitude. The volume values in this study bracket the range of possible pyroclastic flows for both Mammoth and Galeras. According to Calvache [1990], Galeras volcano produced 5 large pyroclastic flow eruptive episodes; an historic eruption in 1866, and prehistoric events in 1100, 2300, 2900, and 4500 yBP. The total deposit volumes of these episodes range from $O(10^6 - 9 \times 10^6)m^3$. Block and ash flows on Mammoth Mountain might contain $O(10^5 - 10^7)m^3$ of material [Patra et al., 2005, Burkett, 2007]. Thus, our choice of volumes ranges from 1.9×10^5 to $5 \times 10^6 m^3$. The shape of the initial failure region is approximated as a paraboloid for which the volume is calculated as follows:

$$V = \frac{\pi}{2} \cdot r_{min} \cdot r_{max} \cdot h_{max} \quad (6)$$

For a good match of the volume range, the radius values were uniformly distributed between 25 and 500 m, while the initial height followed the same distribution with values between 10 and 150 m.

Initiation locations were taken from previous mapping of vent sites, coupled with knowledge of known weak areas within the volcano as indicated by hydrothermal alteration. Around the centers of the separate initiation locations, different starting positions were uniformly distributed in a circle of radius 200 m. A rectangular area of approximative $40 km^2$ was defined around the vent within the available DEMs as the potential run-out area.

4 Results and Conclusions

xxherexx

One of the goals of our analysis was to understand the effect of the spatial structure of available DEMs on hazard maps. Figure 2 c,d shows the correlograms for the ASTER DEM and the TOPSAR DEM, which are the DEMs considered to best represent the real topography for Galeras Volcano and Mammoth Mountain, respectively. It is apparent that data processing resulted in a smoothing and filtering of the TOPSAR DEM which causes the correlation coefficient to vary smoothly as a function of distance and any two elevation values. Using a distance between two points of 2000m, for the ASTER DEM the correlation coefficient is 0.6, whereas for the TOPSAR DEM the correlation coefficient is 0.4. This means that elevation values within the ASTER DEM are more highly correlated.

Starting from these premises, we can explain the hazard map output for the cases when the DEM is considered to be an input parameter for the TITAND2D model. Figures 3 (a) and (c) display maps of Galeras of the probability that the flow depth will exceed 0.5 m in the next ten years using Method 1 or Method 2, respectively to create the terrain realizations.

In Figures 3 and 4 show maps at Galeras and Mammoth of the spatially varying non-confidence in the probability-of-hazard map depicted in Figure 4 (a) and (b), respectively. The non-confidence is defined as the computed standard deviation of probability-of-hazard σ_P divided by the probability-of-hazard P . When calculated by standard means, as was done here, the ratio σ_P/P only measures the non-confidence in the statistics due to insufficient re-sampling of the input parameter space.

After a visual inspection of the figures we can conclude that the difference between hazard map outputs is more obvious for Galeras than for Mammoth Mountain. From the non-confidence figure it is observed that the error is concentrated at flow margins.

For Mammoth Mountain the differences are less obvious, but with important differences again concentrated at the edge of the flow. A better illustration of how the probabilities are varying for the case when Method 1 is used compared to Method 2 is showed in Figure 5. We observe that comparing every point where there is a probability of having a flow depth greater than 0.5 m for the two methods, the results for Galeras show a greater dispersion than Mountain Mountain. When the flow is deep, the probability is high and tends to cluster near 1 for both mountains. As the probability decreases dispersion becomes

greater for Galeras. We can conclude that as the topography becomes more highly correlated, one should use a more complex creation method as stochastic Method 2. It appears that the spatial autocorrelation of the elevation points influences the hazard map output and a "controlled" perturbation of the elevation as Method 1 will not capture this effect.

One of the conclusions of the previous work was that for moderate and smaller scale flows, different representations of the terrain more profoundly affect computation of an accurate footprint of the flow. We built a new set of hazard maps for the case when the volume is low, with a range between $10^4 - 5 \times 10^4$ and for a high volume between $9 \times 10^6 - 5 \times 10^7$. Since only 517 pixels for Mammoth and 872 pixels for Galeras were generated, the low flow hazard map gives extreme results, which means that we have a hazard (flow greater than 0.5m) with either probability ~ 1 or ~ 0 . In Figure 6 (a), (b) we observe that there is a significant mismatch of prediction (left upper corner and right lower corner) for both volcanoes that can be critical in case of a risk assessment. For high volume, for Galeras, we observe that the area of probability greater than zero is much smaller when we are using Method 2 compared to Method 1. We don't know the reason for this counterintuitive effect.

The question that we tried to answer next is "What is the effect of the DEM in constructing a hazard map?". A quantitative and qualitative analysis is performed for the case when the "original" deterministic DEM (ASTER 30m for Galeras and TOPSAR 30m for Mammoth as in Figure 7 is used as input parameter for the hierarchical emulator, or when the input are a set of terrain realizations obtained following the methodology presented above.

We compare the hazard maps produced when DEM uncertainty is included against maps produced when DEM uncertainty is included. Figures 8 and 9 show that for Galeras the probability that the flow was deeper than 0.5m varies considerably from the case of no DEM uncertainty. Hence, the DEM is an important input parameter and the effect of DEM it is not diminished by other uncertain parameters or the method used. We observe that capturing the probability of having flow greater than 0.5 m at the edge of flow is done with error in both cases. For Mammoth Mountain the effect of DEM is much more obvious when Method 2 is used, one of the causes might be the early termination of the flow compared to the original DEM.

Perturbing the DEM is more important as autocorrelation increases, as the scatter in comparative probabilities is greater for Galeras than for Mammoth.

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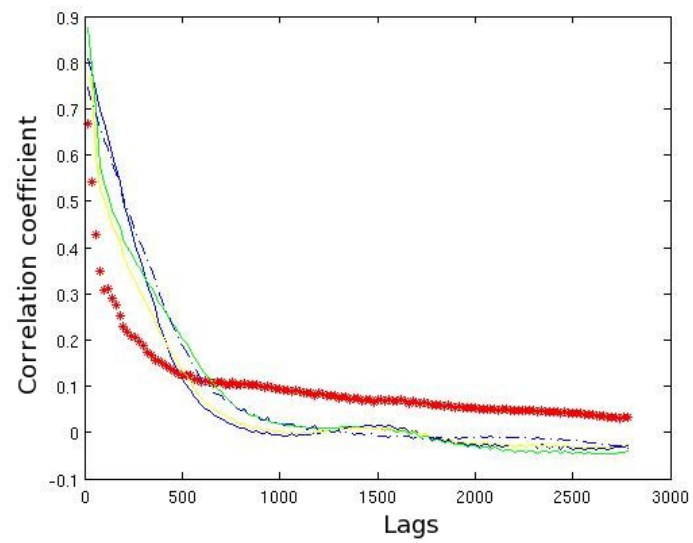
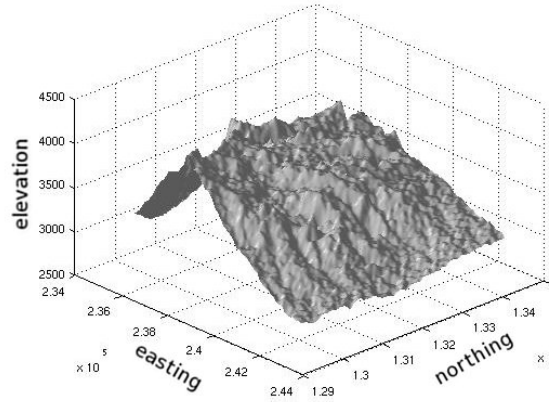
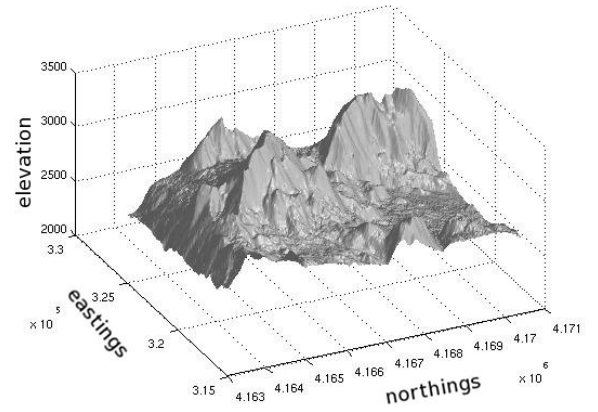


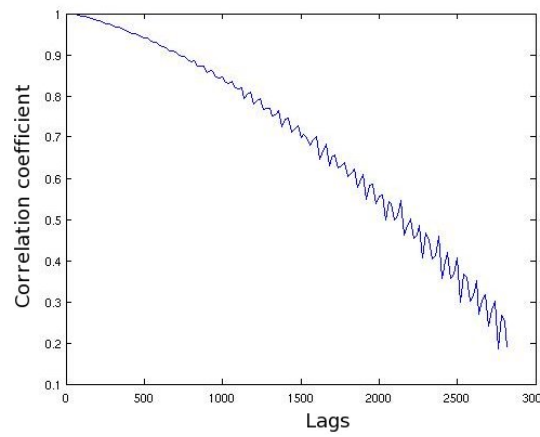
Figure 1: Error map correlogram (red) and various random fields



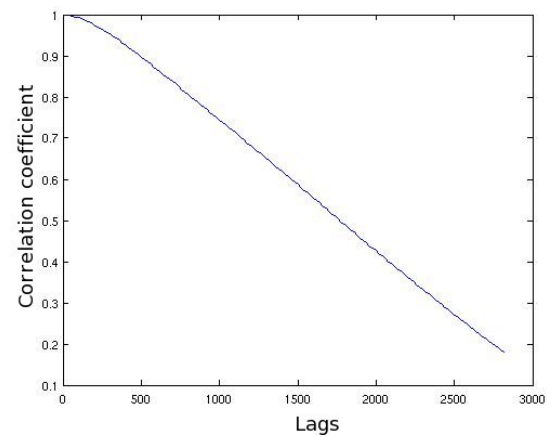
(a)



(b)



(c)



(d)

Figure 2: a) The Galeras ASTER 30m DEM terrain surface (Easting, Northing and elevation coordinates) (b) The Mammoth NED 30m DEM terrain surface (Easting Northing and elevation coordinates) (c) Galeras Volcano ASTER DEM correlogram (d) Mammoth Mountain TOPSAR DEM correlogram

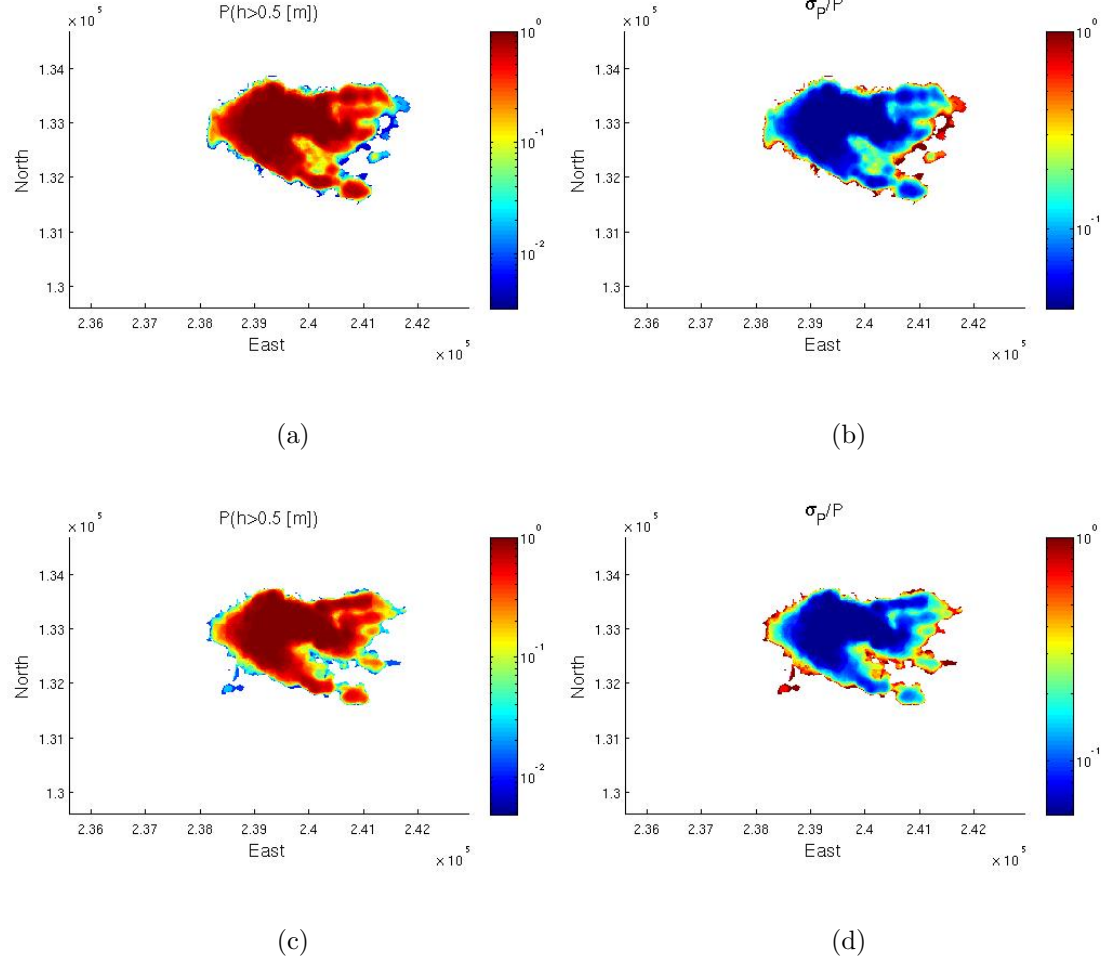


Figure 3: a) Probability that a flow will exceed 0.5 m in depth as a function of position on Galeras Volcano, Columbia, given the uncertainties in DEM and input parameters using Method 1 to create DEM realizations (b) Standard deviation in the estimate that the flow will exceed 0.5 m in depth – Method 1 (c) Probability that a flow will exceed 0.5 m in depth as a function of position on Galeras Volcano, Columbia, given the uncertainties in DEM and input parameters using Method 2 to create DEM realizations (d) Standard deviation in the estimate that the flow will exceed 0.5 m in depth – Method 2

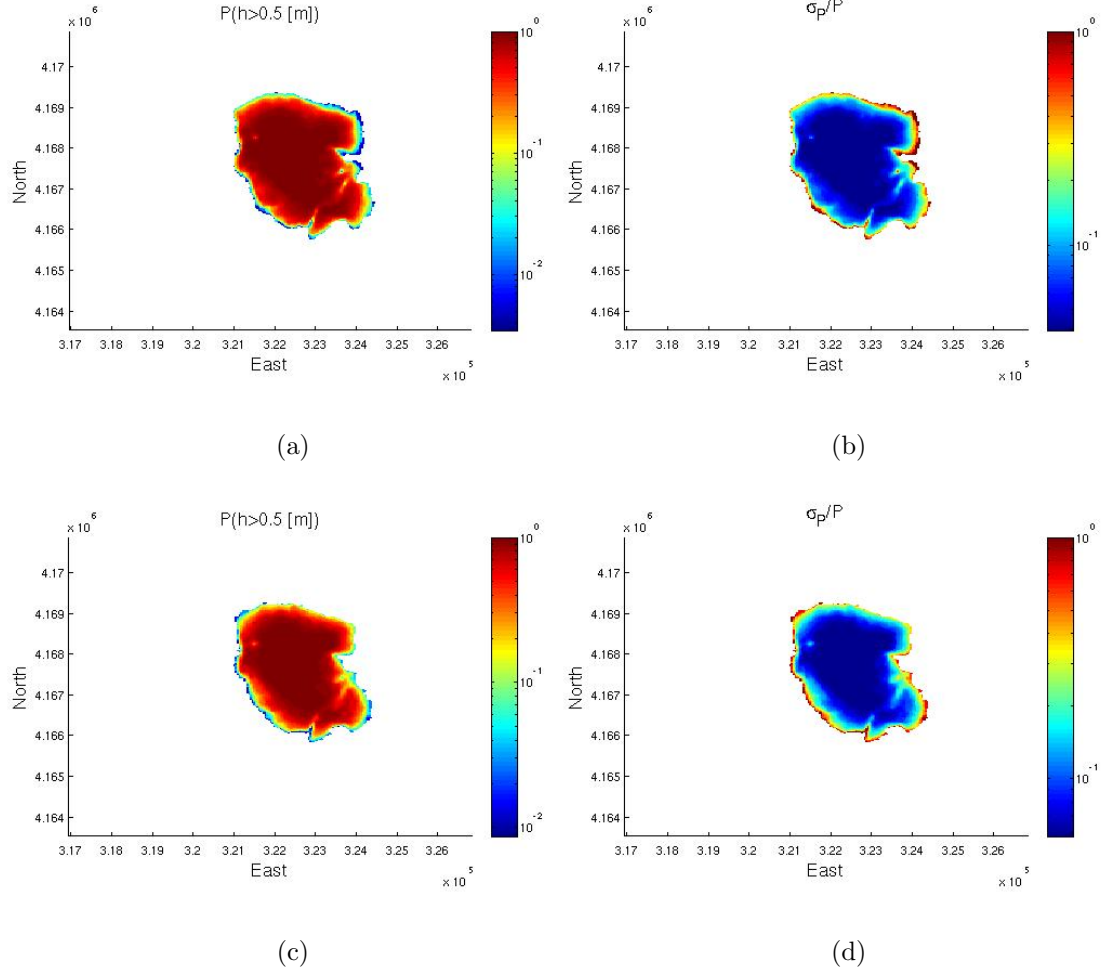


Figure 4: a) Probability that a flow will exceed 0.5 m in depth as a function of position on Mammoth Mountain, CA, given the uncertainties in DEM and input parameters using Method 1 to create DEM realizations. (b) Standard deviation in the estimate that the flow will exceed 0.5 m in depth – Method 1 (c) Probability that a flow will exceed 0.5 m in depth as a function of position on Mammoth Mountain, CA, given the uncertainties in DEM and input parameters using Method 2 to create DEM realizations. (b) Standard deviation in the estimate that the flow will exceed 0.5 m in depth – Method 2

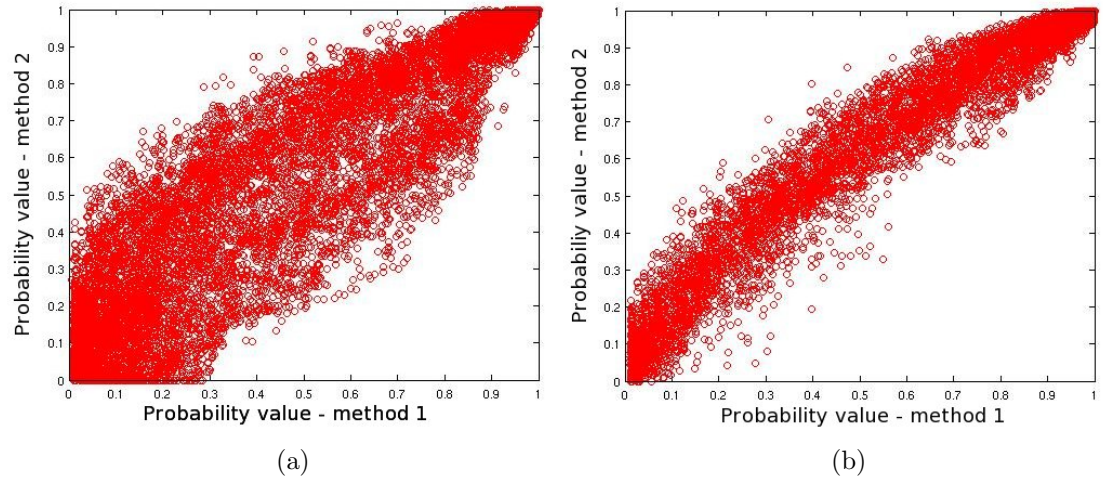


Figure 5: a) The probability that flow will exceed 0.5 m Method 1 versus Method 2 for Galeras Volcano, Colombia (b) The probability that flow will exceed 0.5 m Method 1 versus Method 2 for Mammoth Mountain, CA

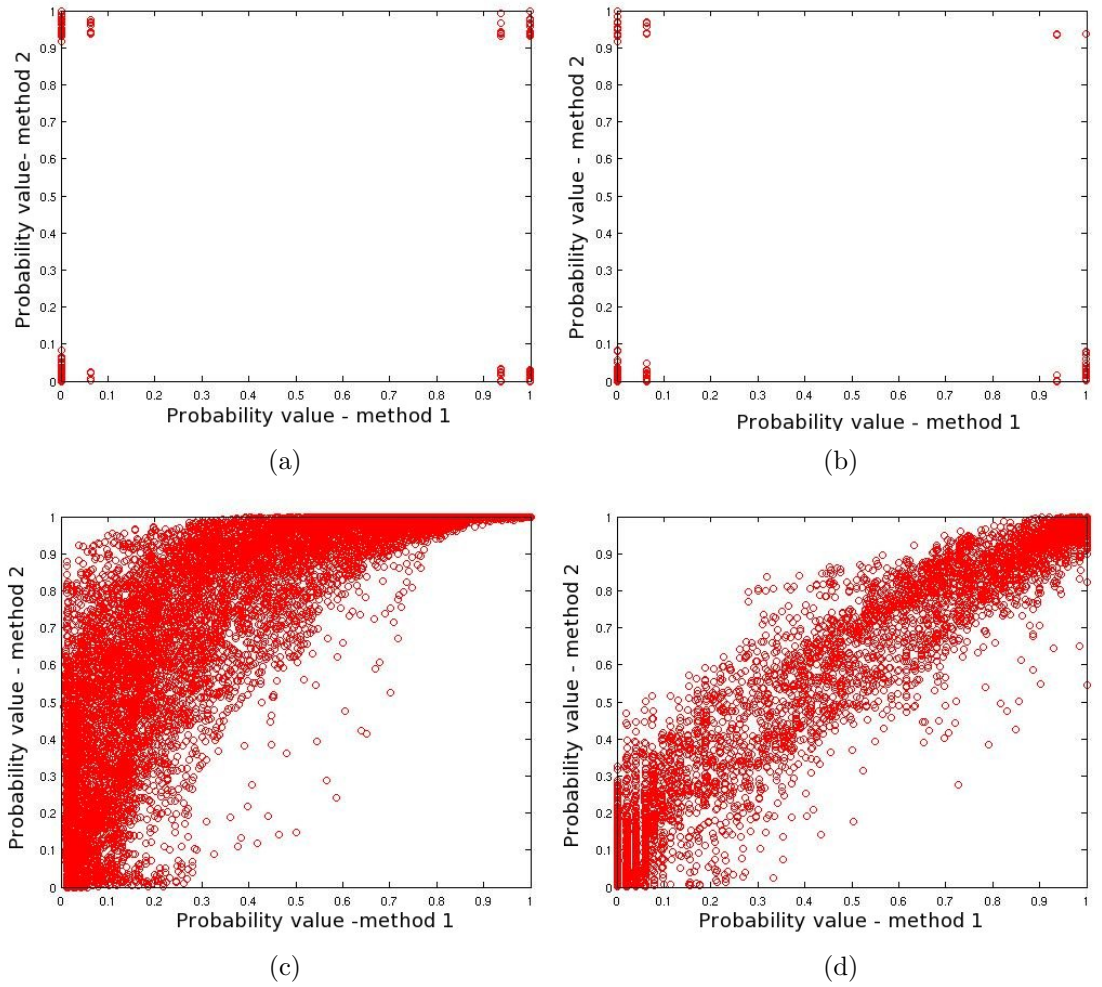


Figure 6: (a) The probability that flow will exceed 0.5 m Method 1 versus Method 2 for: (a) Galeras Volcano, Colombia for low flow (b) Mammoth Mountain, CA for low flow (c) Galeras Volcano, Colombia for high flow (d) Mammoth Mountain, CA for high flow

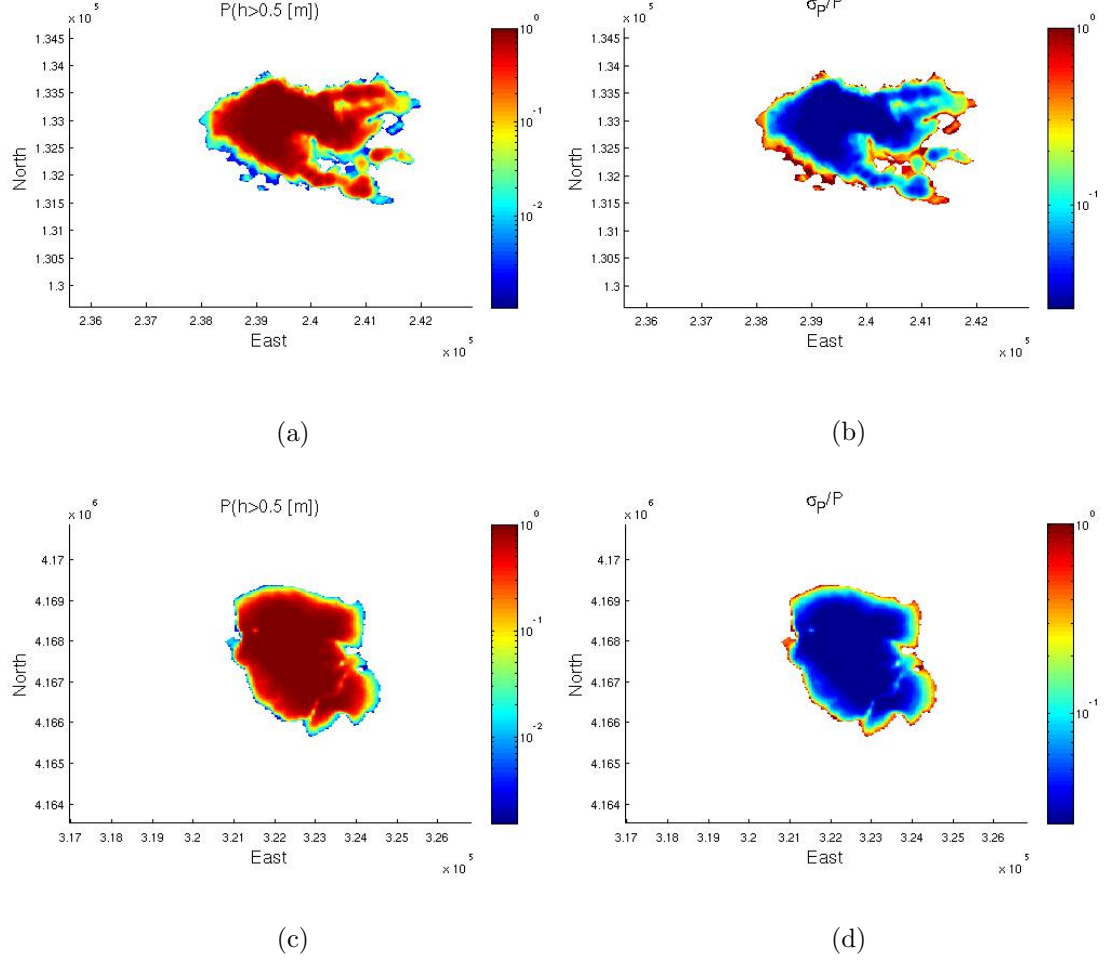


Figure 7: (a) The probability that flow will exceed 0.5 m for Galeras ASTER (b) Standard deviation in the estimate that the flow will exceed 0.5 m in depth for Galeras ASTER (c) The probability that flow will exceed 0.5 m for Mammoth TOPSAR (d) Standard deviation in the estimate that the flow will exceed 0.5 m in depth for Mammoth TOPSAR

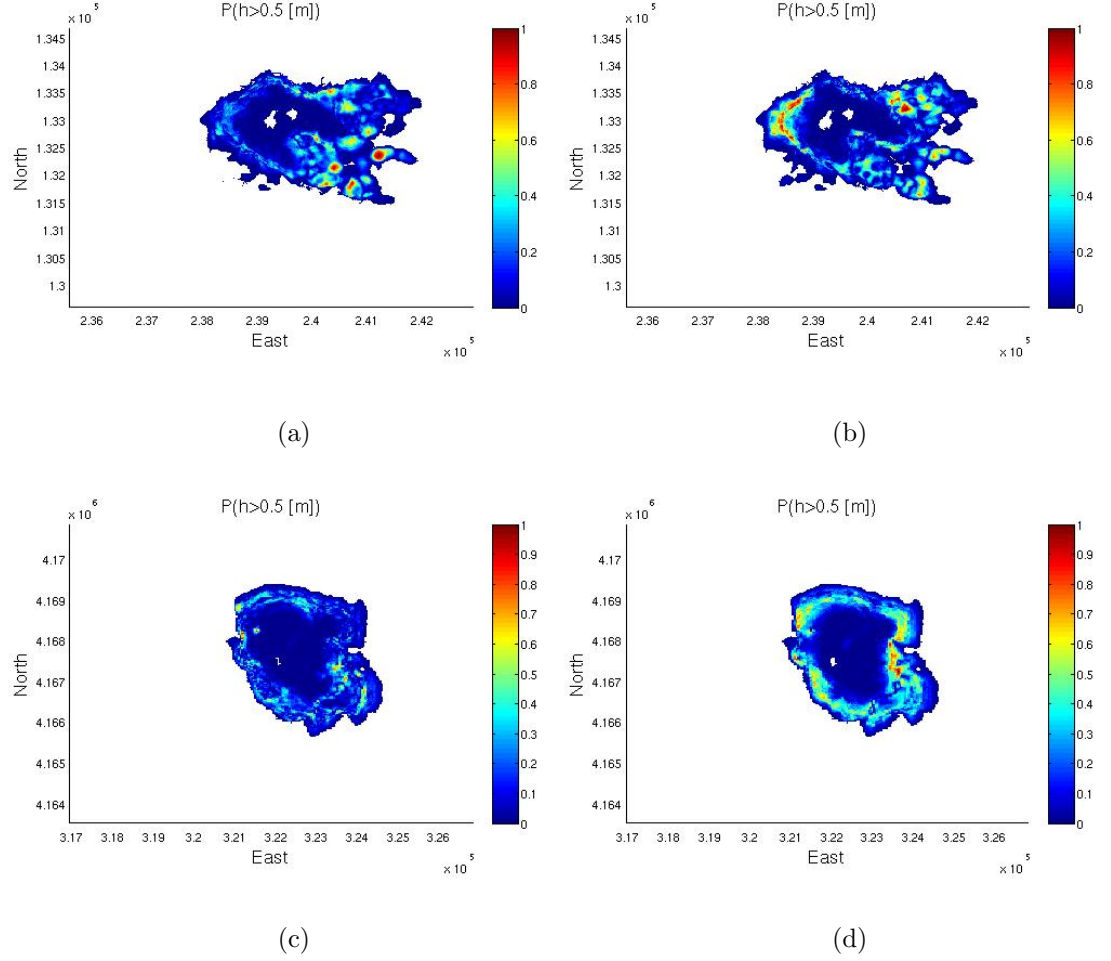


Figure 8: Probability difference map (absolute value) between: (a) Mammoth TOPSAR hazard map and the Method 1 hazard map (b) Mammoth TOPSAR hazard map and the Method 2 hazard map (c) Galeras ASTER hazard map and the Method 1 hazard map (d) Galeras ASTER hazard map and the Method 2 hazard map

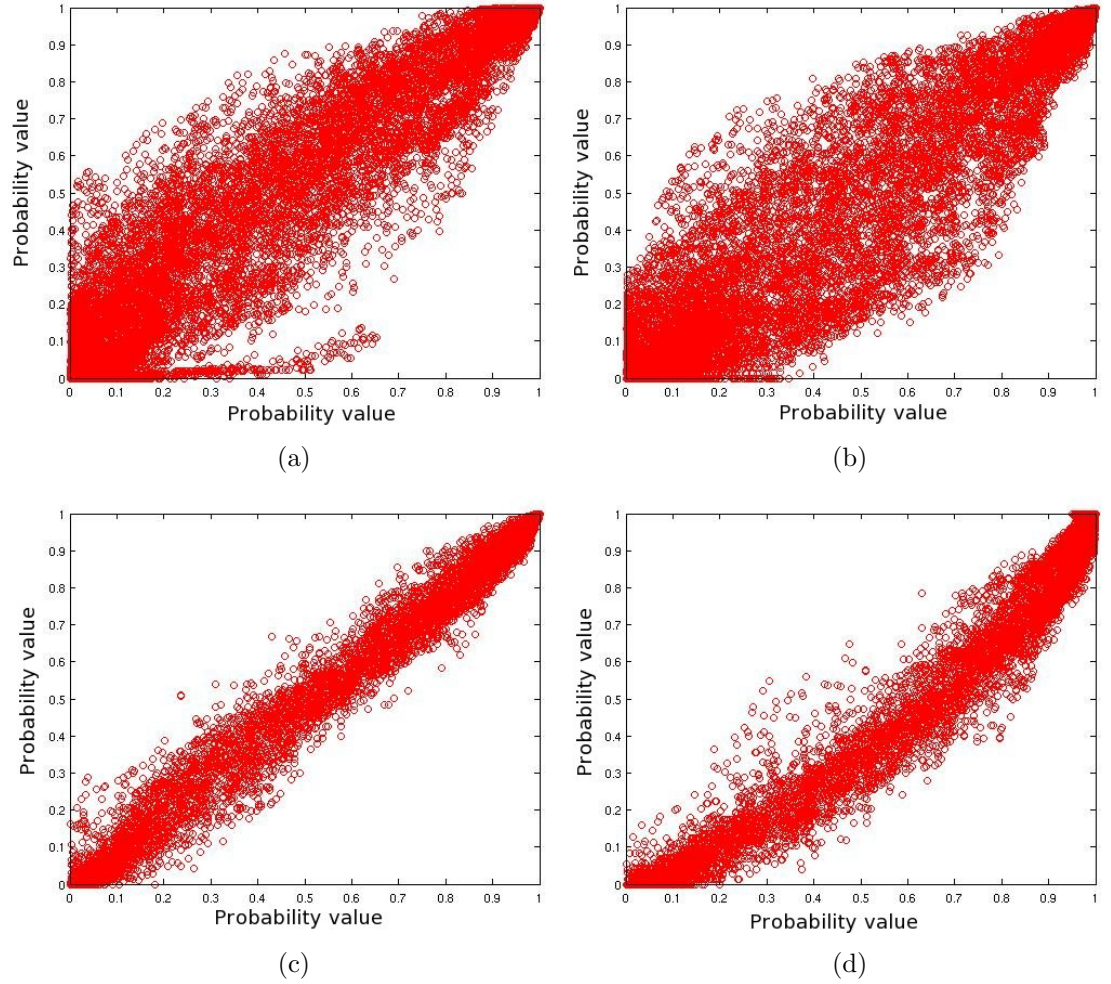


Figure 9: (a) The probability that flow will exceed 0.5 m for Galeras ASTER hazard map versus Method 1 (b) The probability that flow will exceed 0.5 m for Galeras ASTER hazard map versus Method 2 (c) The probability that flow will exceed 0.5 m Mammoth TOPSAR hazard map versus Method 1 (d) The probability that flow will exceed 0.5 m Mammoth TOPSAR hazard map versus Method 2