



The MeqTree module working prototype for LOFAR calibration

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The LOFAR challenge

- ◆ Pathological ionosphere
- ◆ Very crowded fields (confusion, flops)
- ◆ Very bright sources (dynamic range)
- ◆ Variable (f,t,k) station beamshapes
- ◆ Instrumental polarisation (X/Y voltage beams)
- ◆ High station beam sidelobes (all-sky imaging)
- ◆ New types of instrumental effects (e.g. BSR)
- ◆ Huge data volume (Tbytes/hr)

Ionosphere movie (GMRT)

The bottom line

- ◆ The new generation of large radio telescopes (LOFAR, ALMA, SKA) needs software around a more complex **Measurement Equation (M.E.)** than the existing packages can offer
- ◆ A user should not depend on package builders to implement a particular M.E.
- ◆ MeqTree is an attempt to do this (the only one!)
- ◆ The existing telescopes need something also...

The Measurement Equation

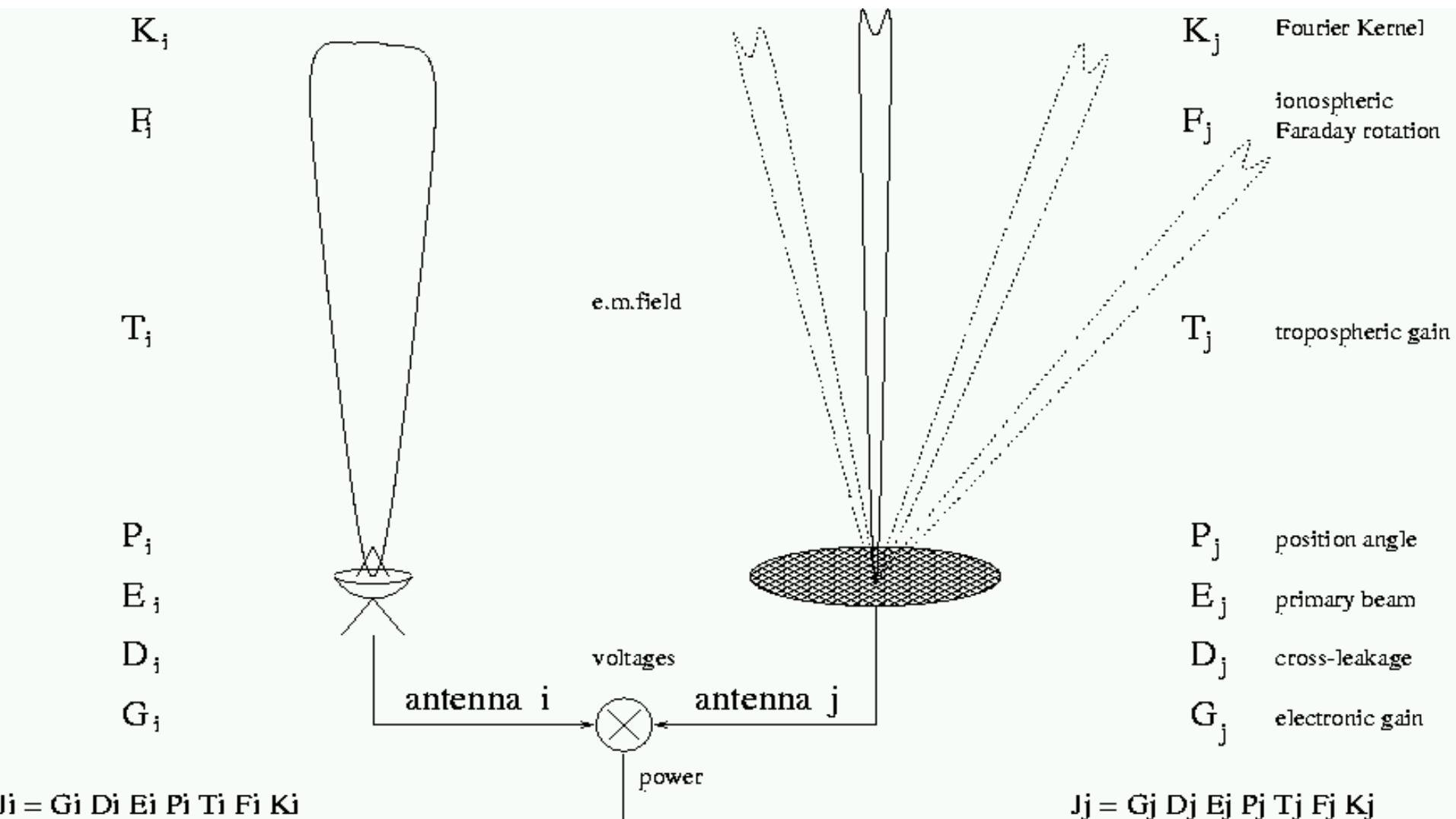
$$\vec{V}_{ij}(u, v) = \sum_k \int df \int dt \left(J_{ik} \otimes J_{jk}^* \right) S \vec{I}_k$$

See also:
AIPS++
note 185

Hamaker-Bregman-Sault formalism
Station-pair (i,j): Vector of 4 ‘correlations’
Matrix equation: full polarization
Instrumental Jones matrices J (2x2)
Stokes matrix S (4x4)
Sum over sources (k): Flux [I,Q,U,V]
Integral ^{time} over time-freq domain



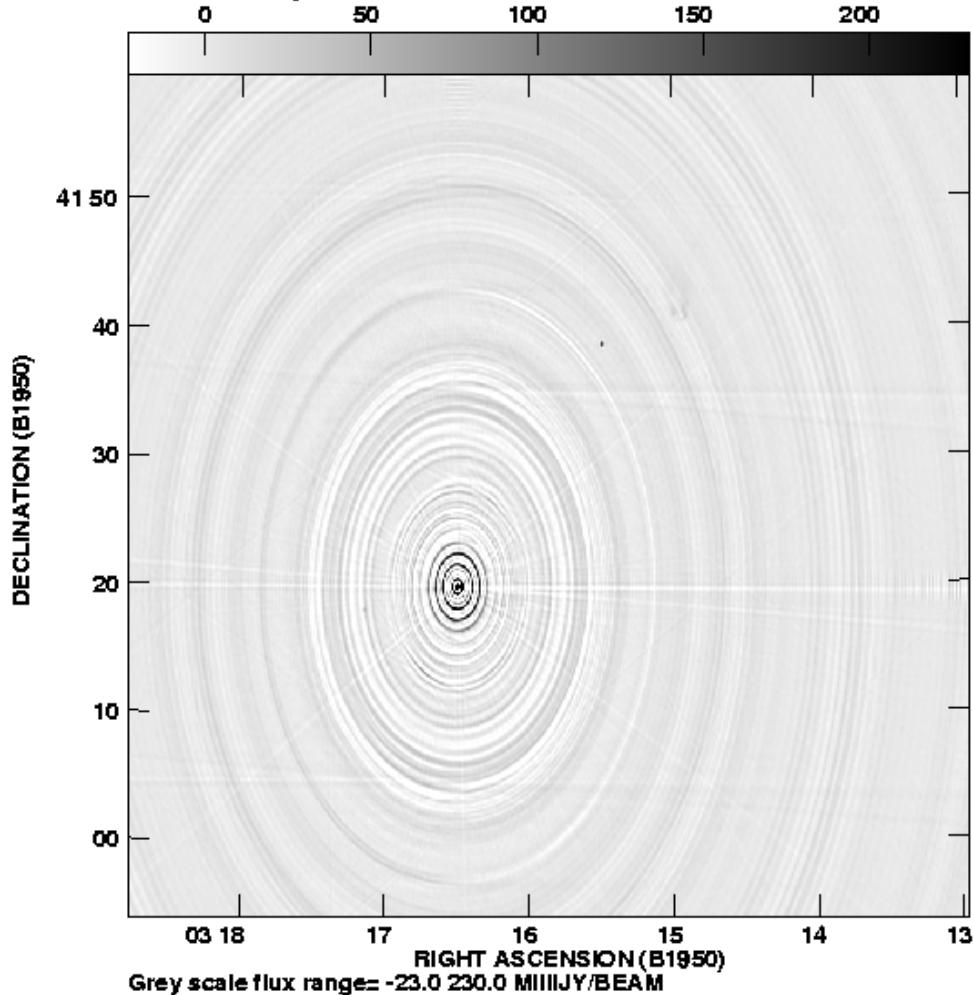
M.E. instrumental effects



M_{ij} multiplicative
interferometer-based effects
 A_{ij} additive

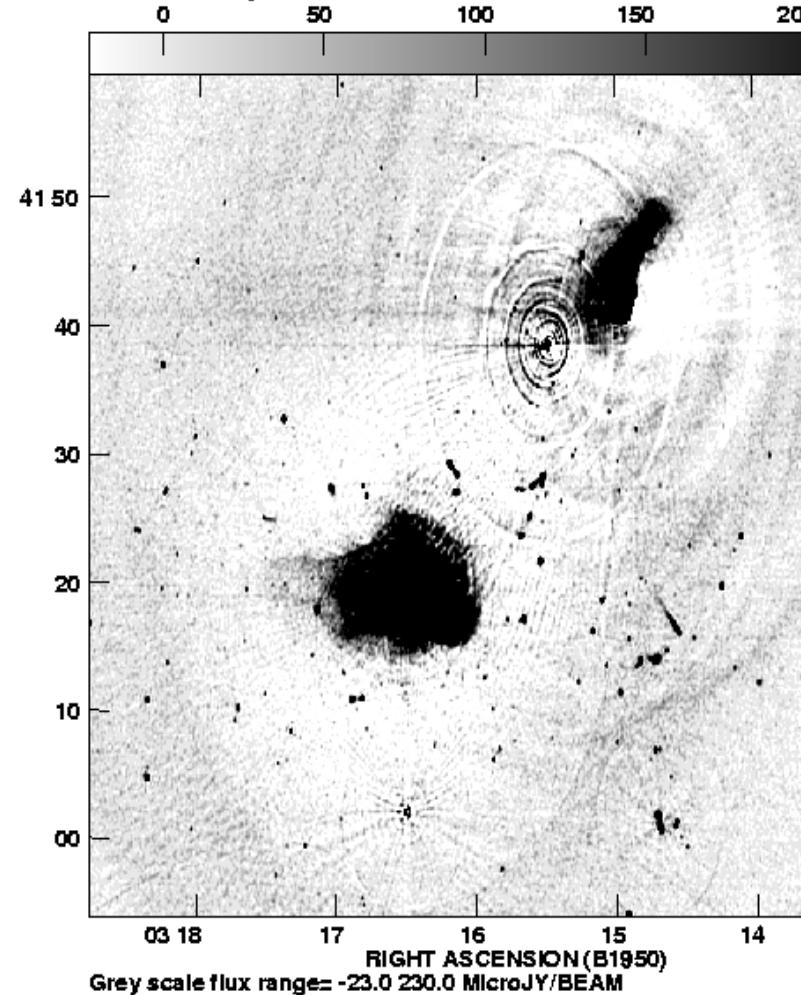
3C84, WSRT, 21cm

Plot file version 1 created 10-OCT-2002 12:05:38
*PERSEUS 1397.875 MHZ 21INOSC.IMAP.1



Plot file version 2 created 27-SEP-2002 17:53:30

*PERSEUS 1397.875 MHZ 21PERSI.IMAP.1



Dynamic range: 100 : 1

Dynamic range: 500.000

The big issues

- ◆ Image-plane effects
 - ⊕ Extra processing
 - ⊕ Apply to extended sources
- ◆ Minimise the number of parameters
 - ⊕ Minimum Ionospheric Model (MIM)
 - ⊕ Beamshapes
- ◆ Data volume (performance)

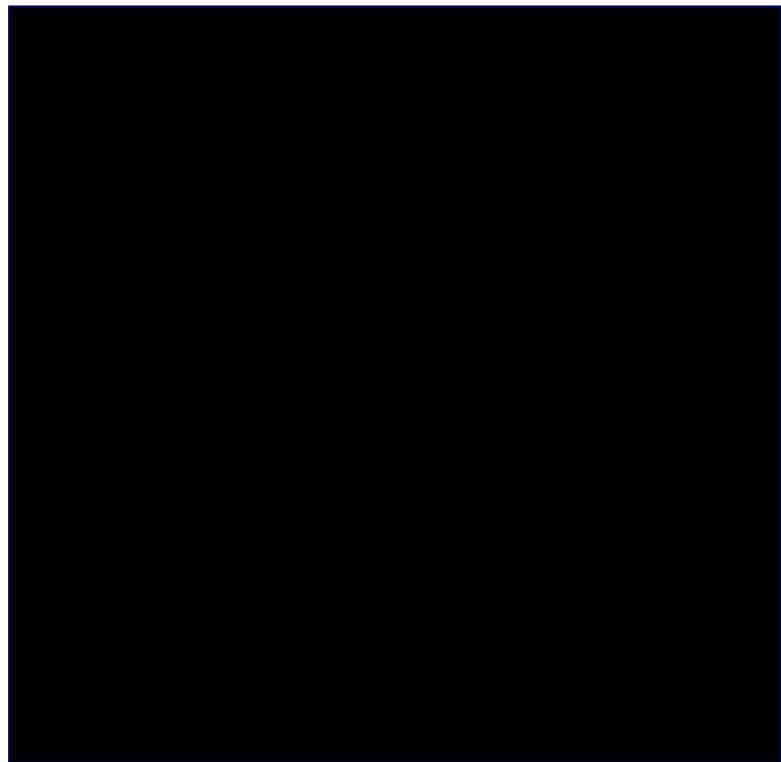
Path to LOFAR calibration software

- ◆ Two separate teams:
 - ❖ Development: MeqTree module and AIPS++
 - ❖ Implementation: Black Board Selfcal (BBS)
- ◆ MeqTree is a tool for M.E. development
 - ❖ Reduce the number of parameters
 - ❖ Low threshold
 - ❖ Exercise on real data (WSRT/WHAT, VLA74 Hz ...)
- ◆ Involve lots of people (especially next generation)
- ◆ Continue with MeqTree module afterwards

Alternate between two domains

- ◆ Instrumental errors are in Fourier domain
 - ❖ They vary in time and frequency
 - ❖ They vary over the Field of View (!)
- ◆ Bright sources are subtracted in Fourier domain
- ◆ Faint sources are found in image plane
 - ❖ But subtracted in Fourier plane (PSF)
- ◆ Residual images must be deconvolved

The result of successful calibration



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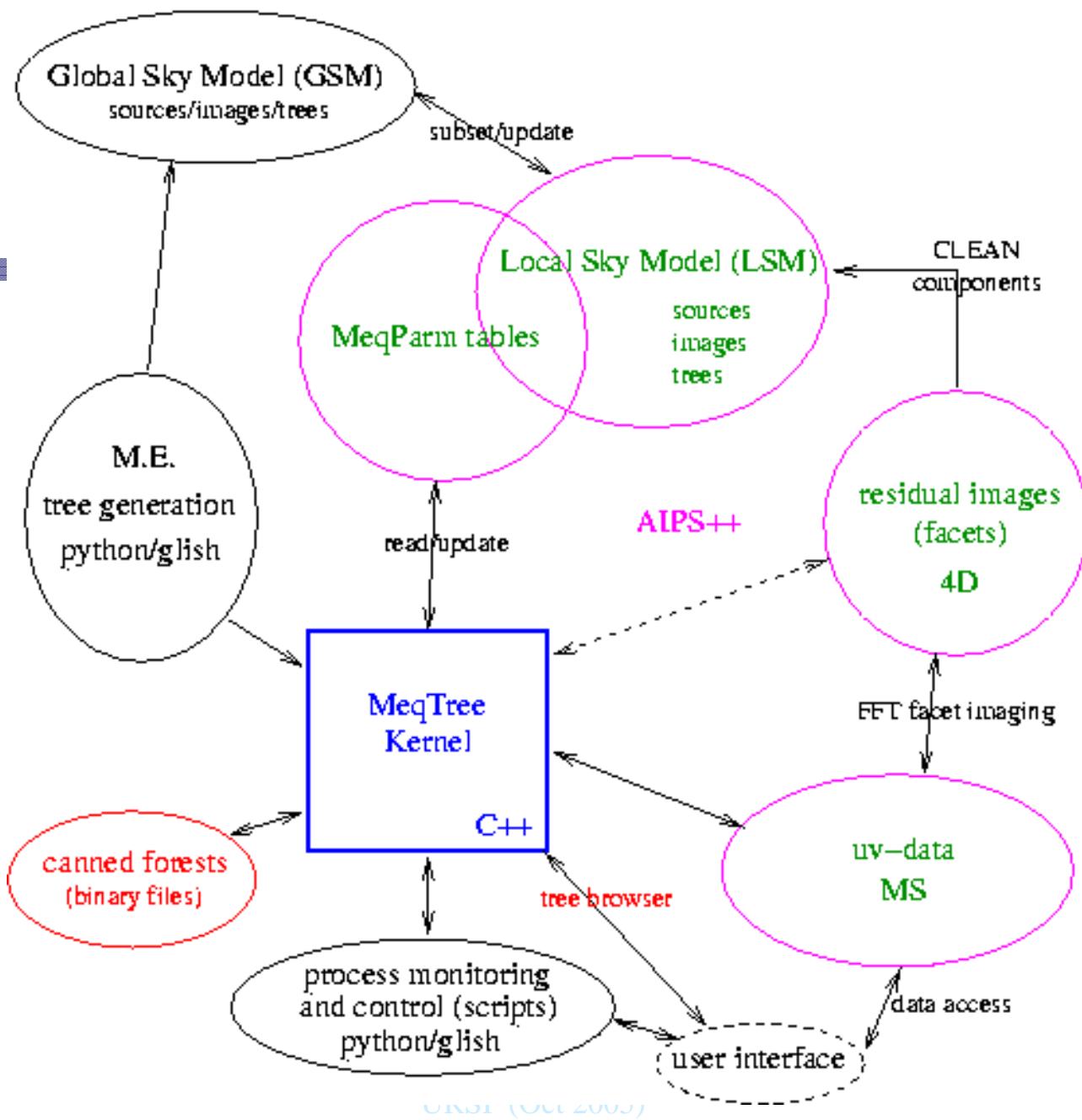
PSS_5 JEN/2004/25/11





Part II: MeqTree Module Overview

MeqTree Environment

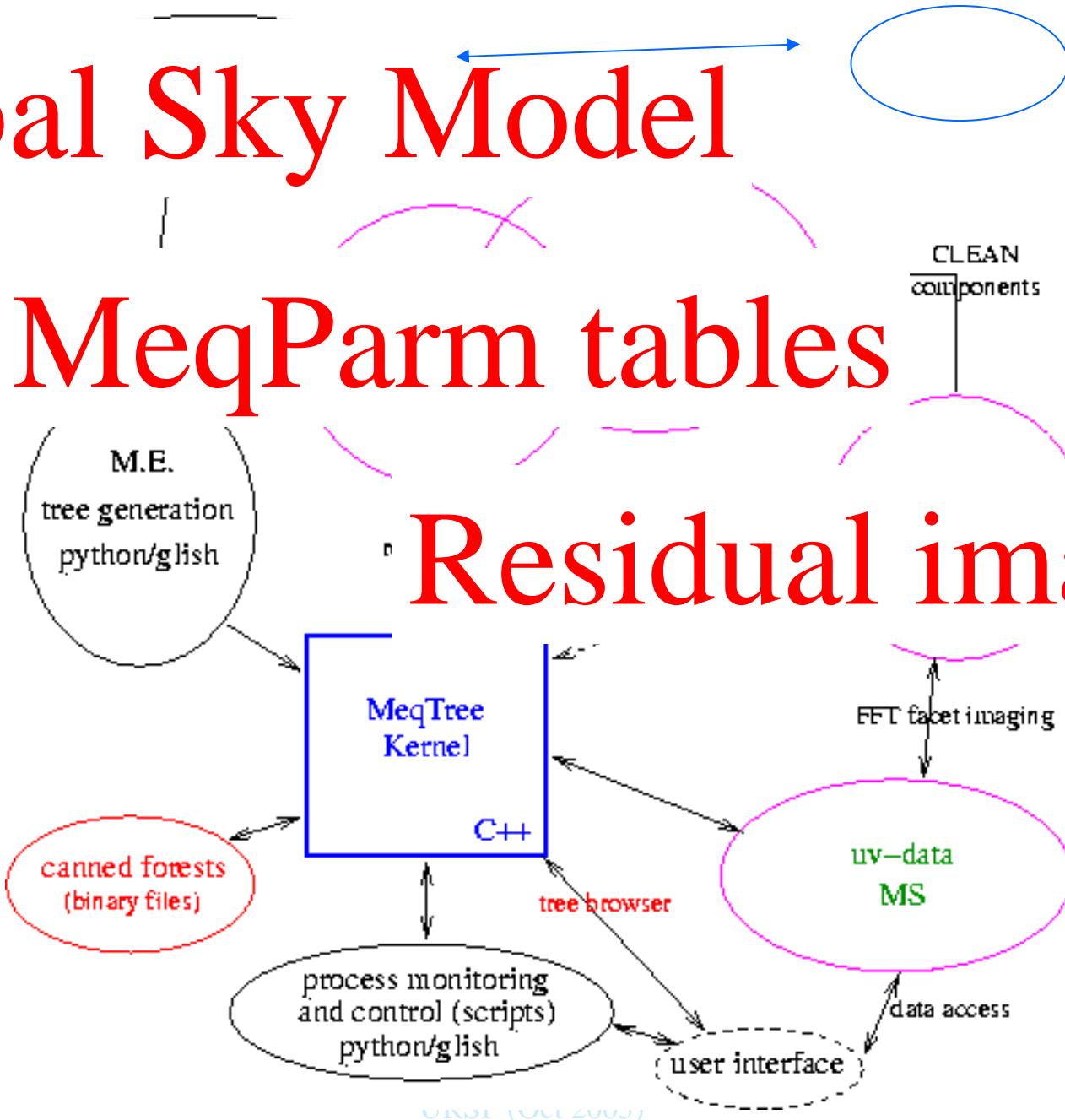


Global Sky Model

LOFAR deliverables

MeqParm tables

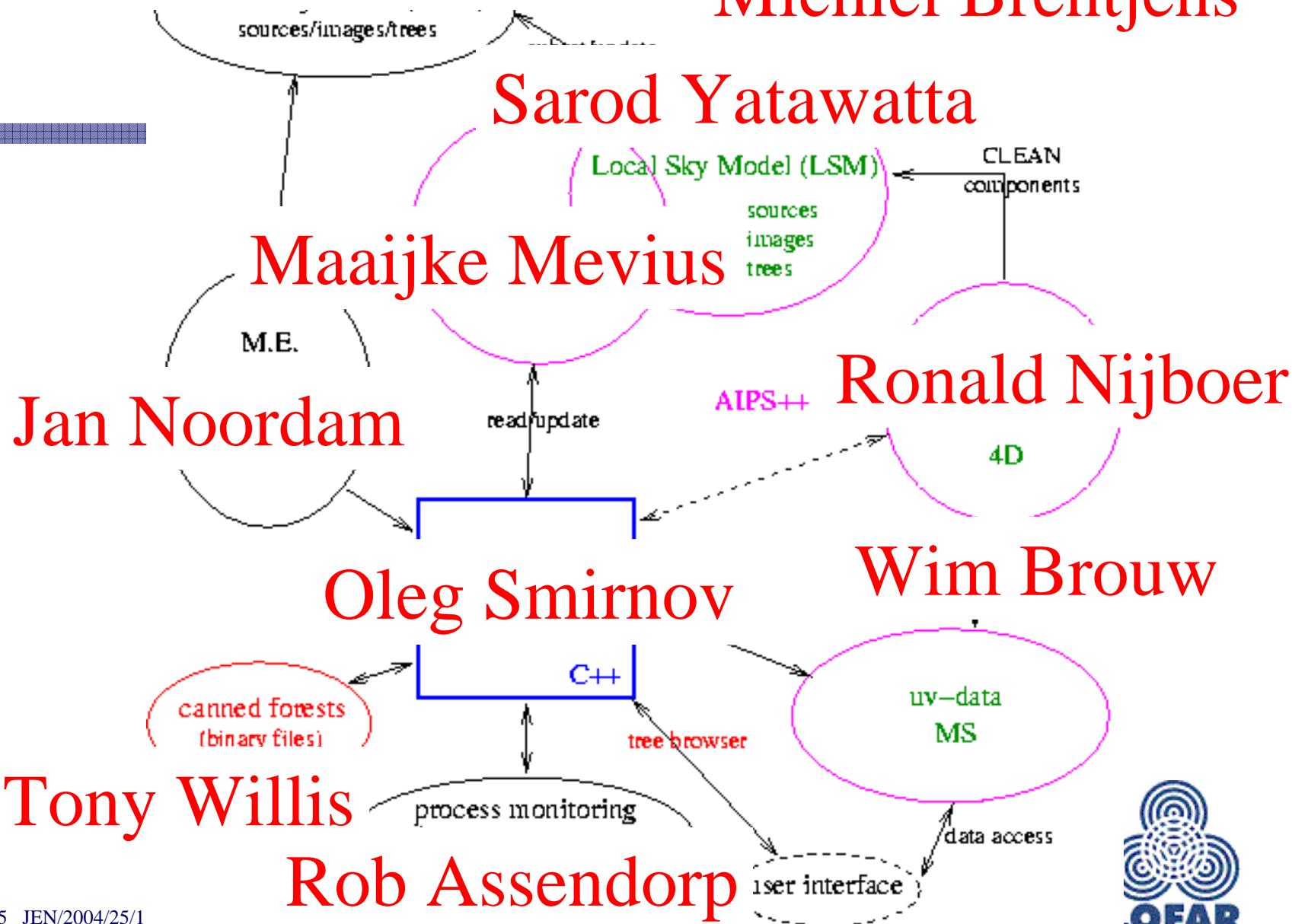
Residual images



The MeqTree Team

Tom Oosterloo

Michiel Brentjens

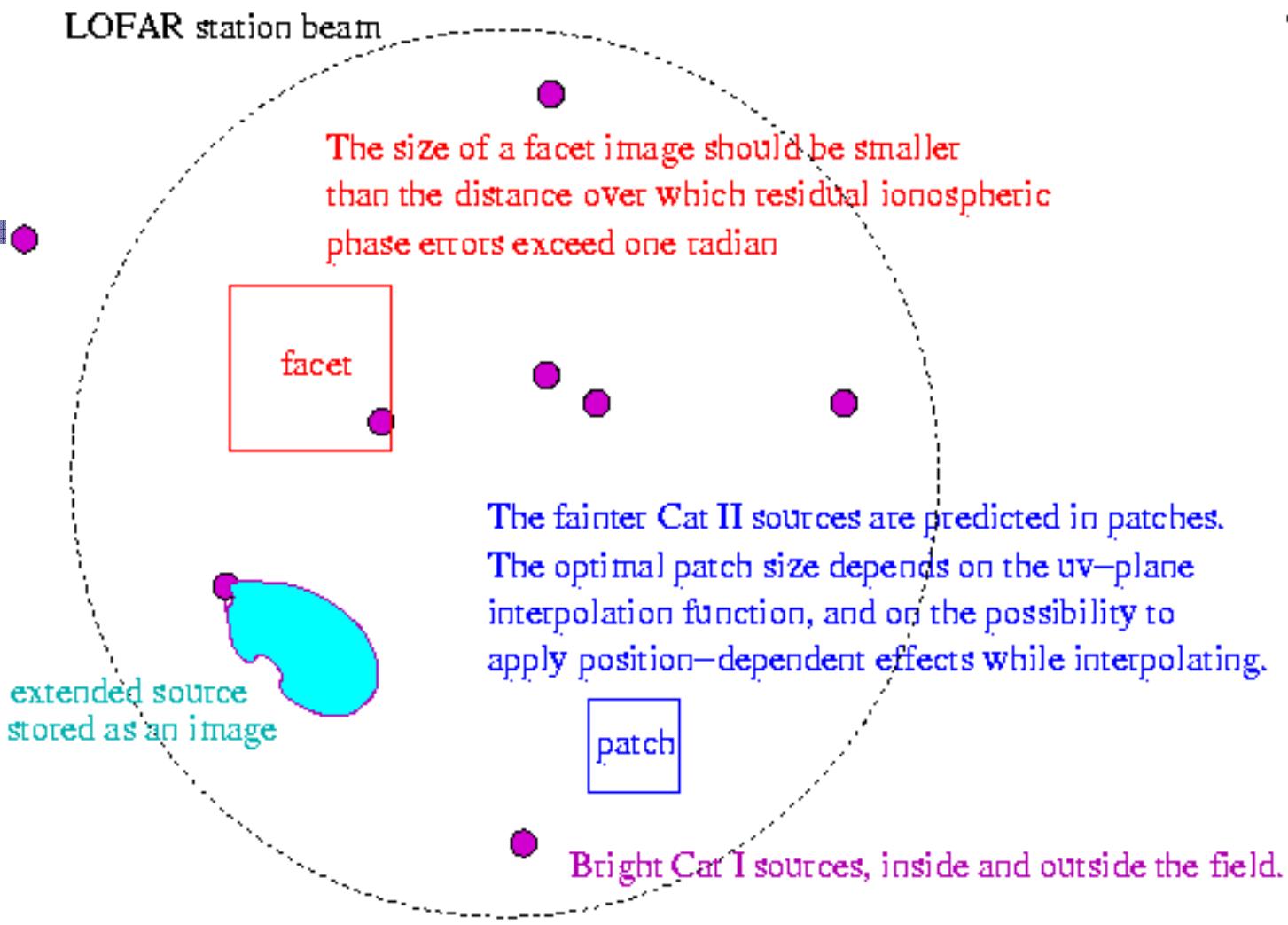


The MeqTree Kernel

- ◆ Implements arbitrary M.E.
- ◆ Solves for subsets of M.E. parameters
- ◆ Policy Free (!)
 - ❖ Will be applied to other areas
- ◆ Data interface (MS)
- ◆ MeqParm interface (AIPS++ tables)
- ◆ Tree Definition Language (TDL)
- ◆ User interaction via the MeqBrowser

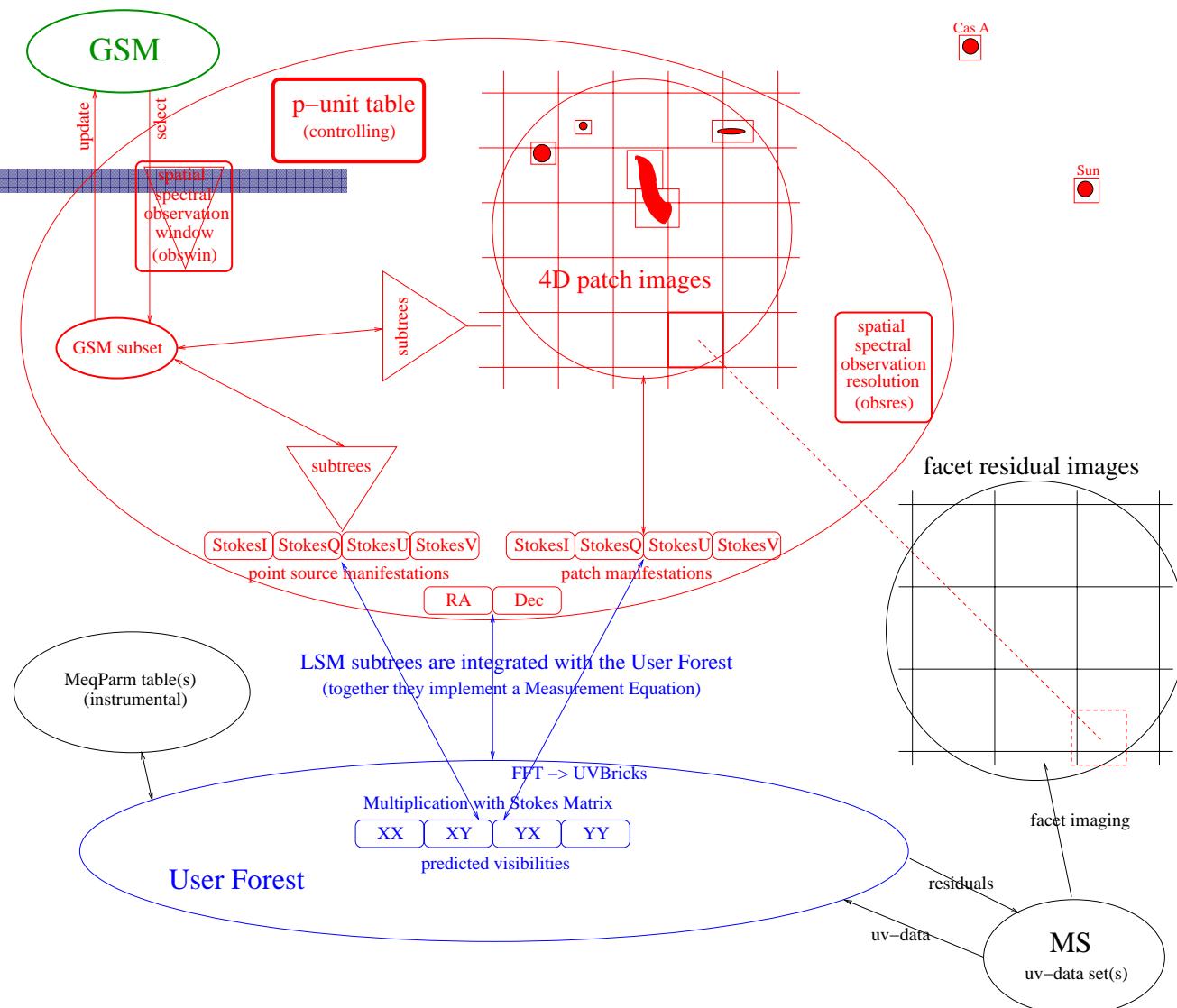
Local Sky Model (LSM)

Contents of a Local Sky Model



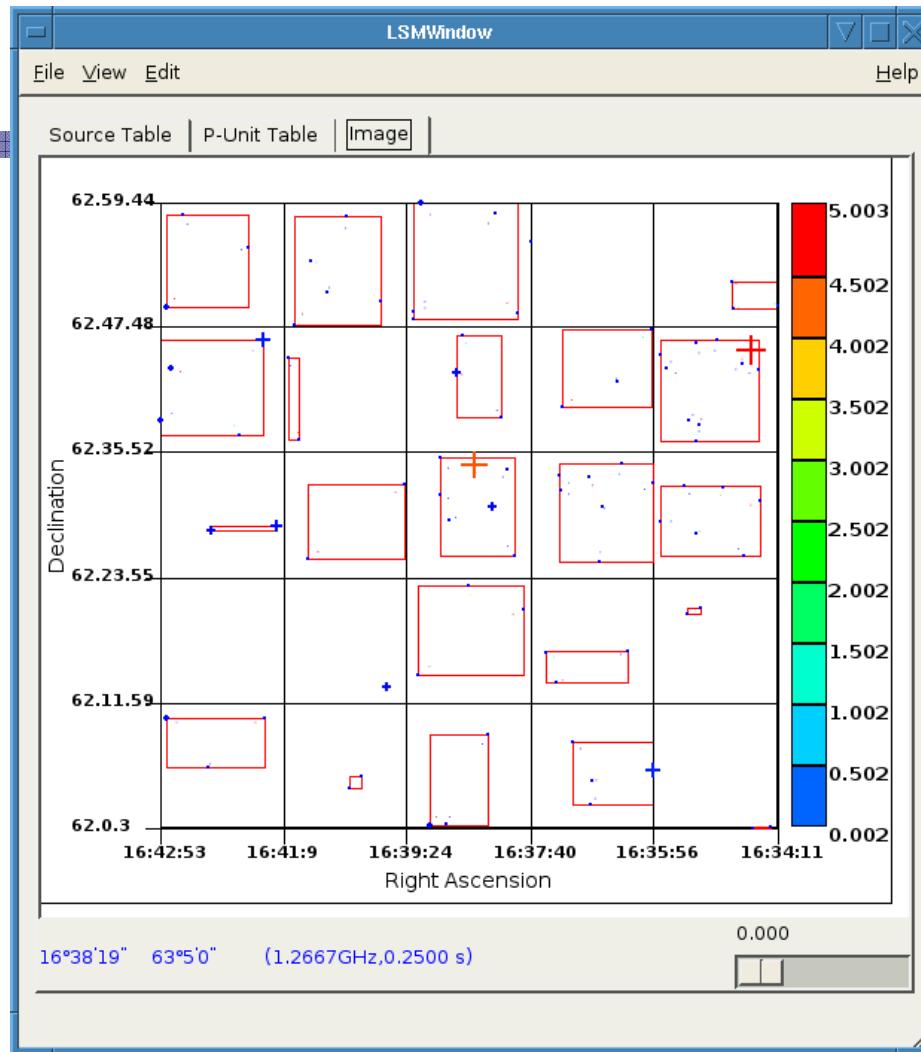
URSI (Oct 2005)

Local Sky Model (red) and user tree (blue)



URSI (Oct 2005)

LSM Interface GUI



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Three categories of sources

- ◆ Cat I: Bright enough to be used for Selfcal
 - ⊕ Subtracted individually from the uv-data
- ◆ Cat II: Bright enough to be identified in a residual image, and put into the LSM
 - ⊕ Subtracted in groups (patches) from the uv-data
- ◆ Too faint to be in the LSM, so not subtracted in the uv-plane
 - ⊕ Imperfectly imaged...

Peeling

- ◆ Deal with sources/patches sequentially
 - ⊕ Large potential saving in processing
- ◆ Selfcal: Solve for instrumental parameters in the direction of bright sources, in order of brightness
 - ⊕ Potential problem: contamination of fainter sources
- ◆ Subtraction: All LSM sources (Cat I and Cat II)
- ◆ Each time, the phase centre is shifted to the position of the peeling source or patch

Bottleneck: Source subtraction

- ◆ All LSM sources are subtracted from the uv-data
- ◆ Requires uv-model derived from LSM
- ◆ But: We are not interested in the details (flux, position) of individual foreground sources
- ◆ Are there ways to ‘filter’ them out?
- ◆ The PAST team may have the beginning of such an algorithm...

field-of view

Residual (facet) imaging

- ◆ Split the field-of-view up in facets
- ◆ Correct residual uv-data for facet centre
 - ⊕ Errors increase towards the edge
- ◆ Integrate (f,t) to decrease field-of-view
- ◆ FFT: Make a residual image of each facet
- ◆ Analyse residual images:
 - ⊕ Find/update Cat II GSM sources
 - ⊕ Deconvolve Cat III sources (if possible)
- ◆ Combine with GSM: LOFAR deliverables

facet



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Source extraction from residual images

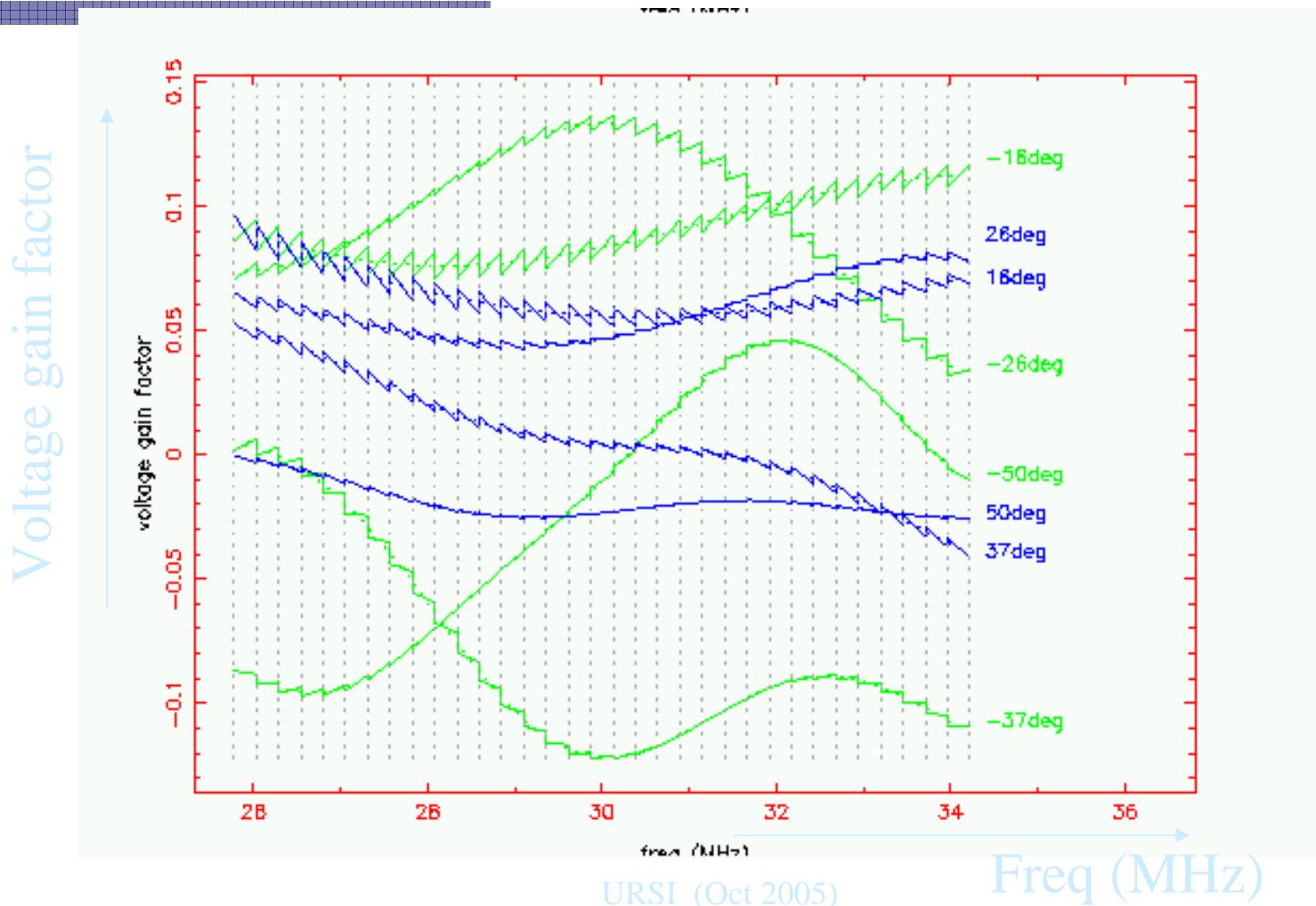
- ◆ Need a PSF
 - ⊕ Strictly valid for centre of field only
 - ⊕ But may degrade benignly towards the edge
- ◆ First check positions where Cat II sources have been subtracted: update LSM source parameters
- ◆ Then look for new (Cat III) sources for LSM
- ◆ Iterate until diminishing returns
- ◆ Ultimately, use the LSM to update the GSM

Some imaging artefacts



- ◆ Remnants of Cat I subtraction (peeling)
- ◆ Remnants of Radio Frequency Interference
- ◆ BSR effect (far from bright sources...?)
- ◆ Pattern of facet edges (cat III imaging)
- ◆ Sidelobe confusion (PSF sidelobe level)
- ◆ Pattern of *p-unit* edges (cat II subtraction)
- ◆ Etc, etc

Bandpass Sawtooth Ripple (BSR effect)



Global Sky Model (GSM)

- ◆ Very large number of LOFAR sources
- ◆ Inhomogeneous parametrisation
 - ⊕ Includes shapelets, pixons and actual images
 - ⊕ **Use MeqTrees in GSM/LSM**
- ◆ Freq/time dependencies
 - ⊕ **Use funklets in GSM/LSM**
- ◆ Should be compatible with Virtual Observatory



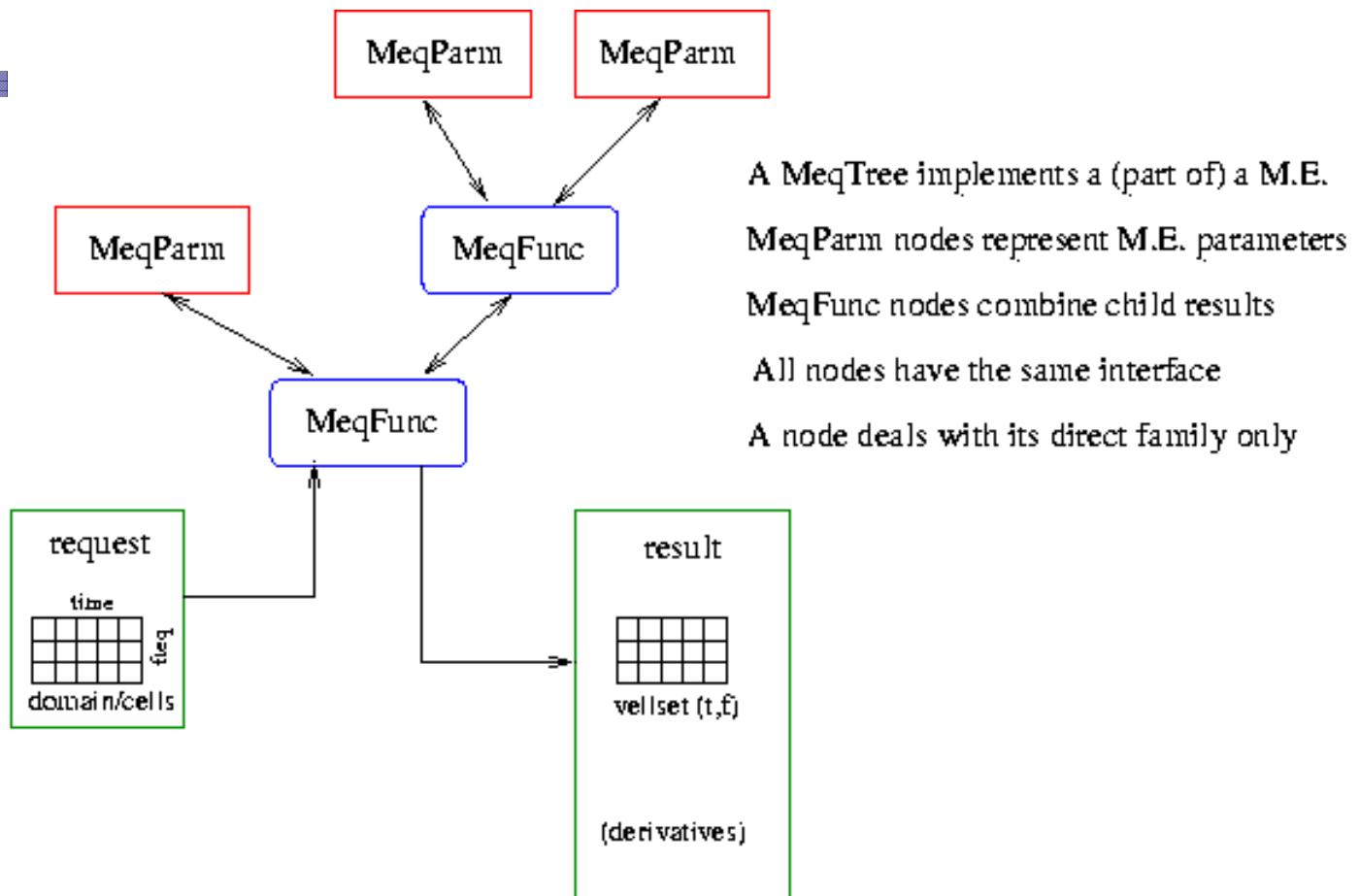
PART III: MeqTree details

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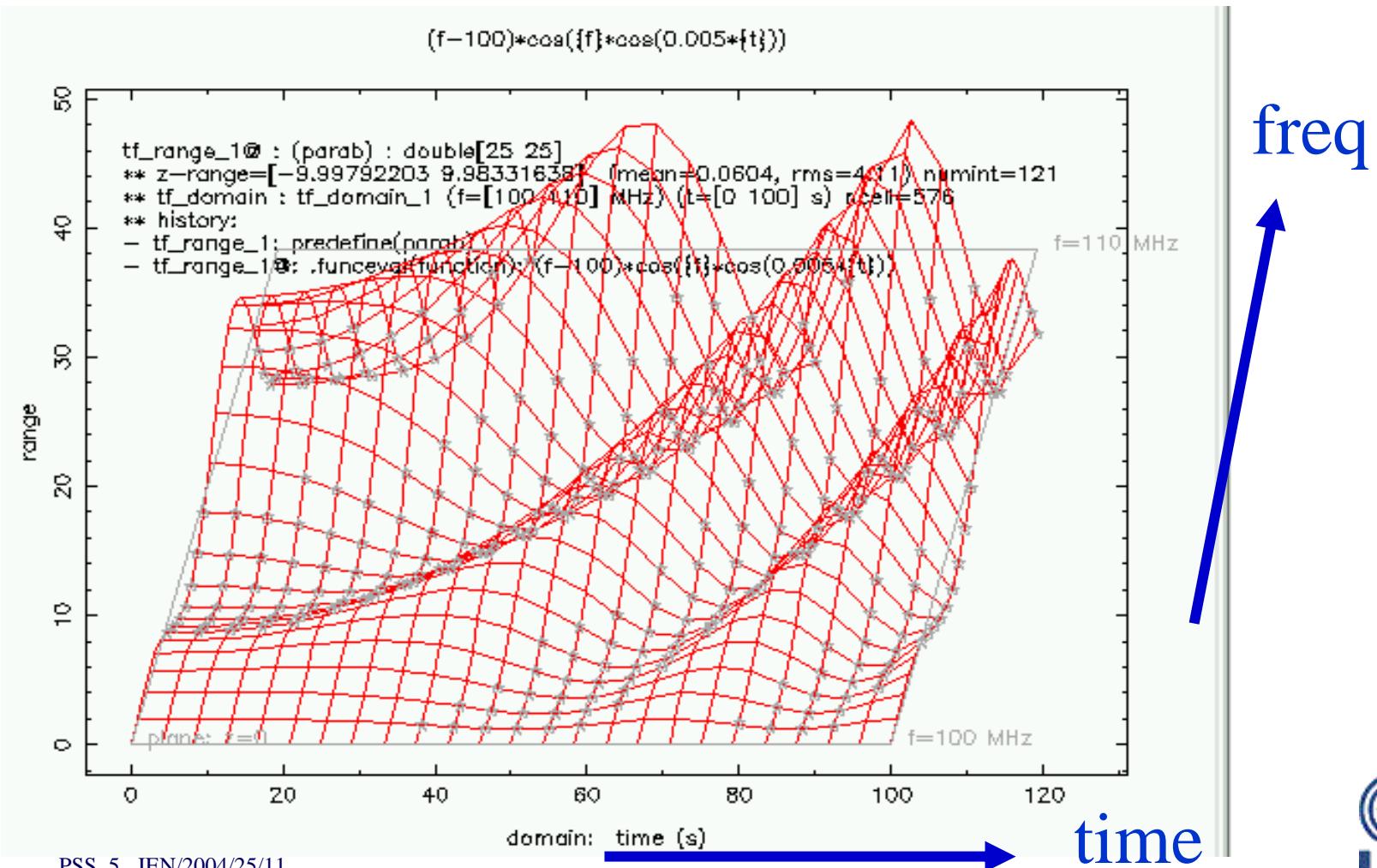


My first MeqTree



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A node result contains vells, i.e. an array of values for the cells of a given domain (t, f):

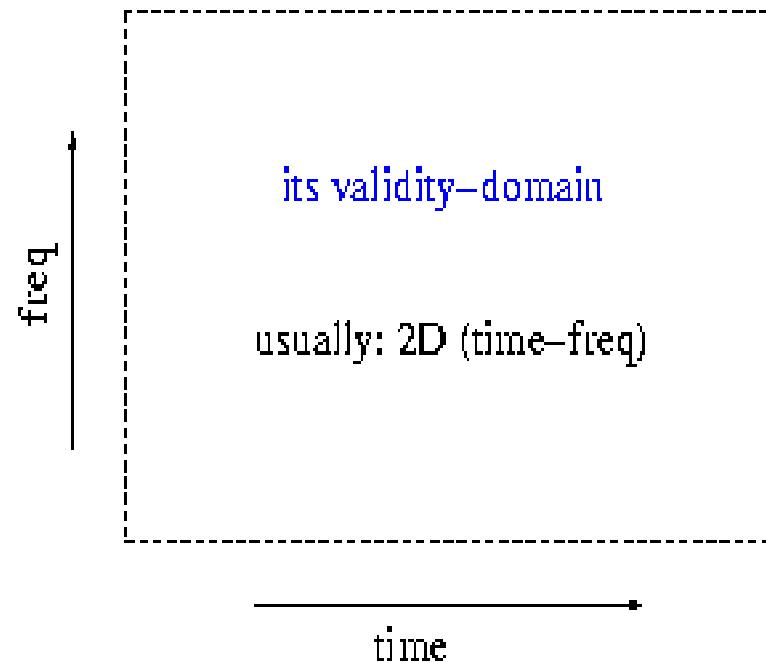


MeqParm values are 'funklets'

| | | | |
|-----|-----|-----|-----|
| c00 | c10 | c20 | c30 |
| c01 | c11 | c21 | ... |
| c02 | c12 | ... | ... |

A polc: an N-dim array of polynomial coefficients

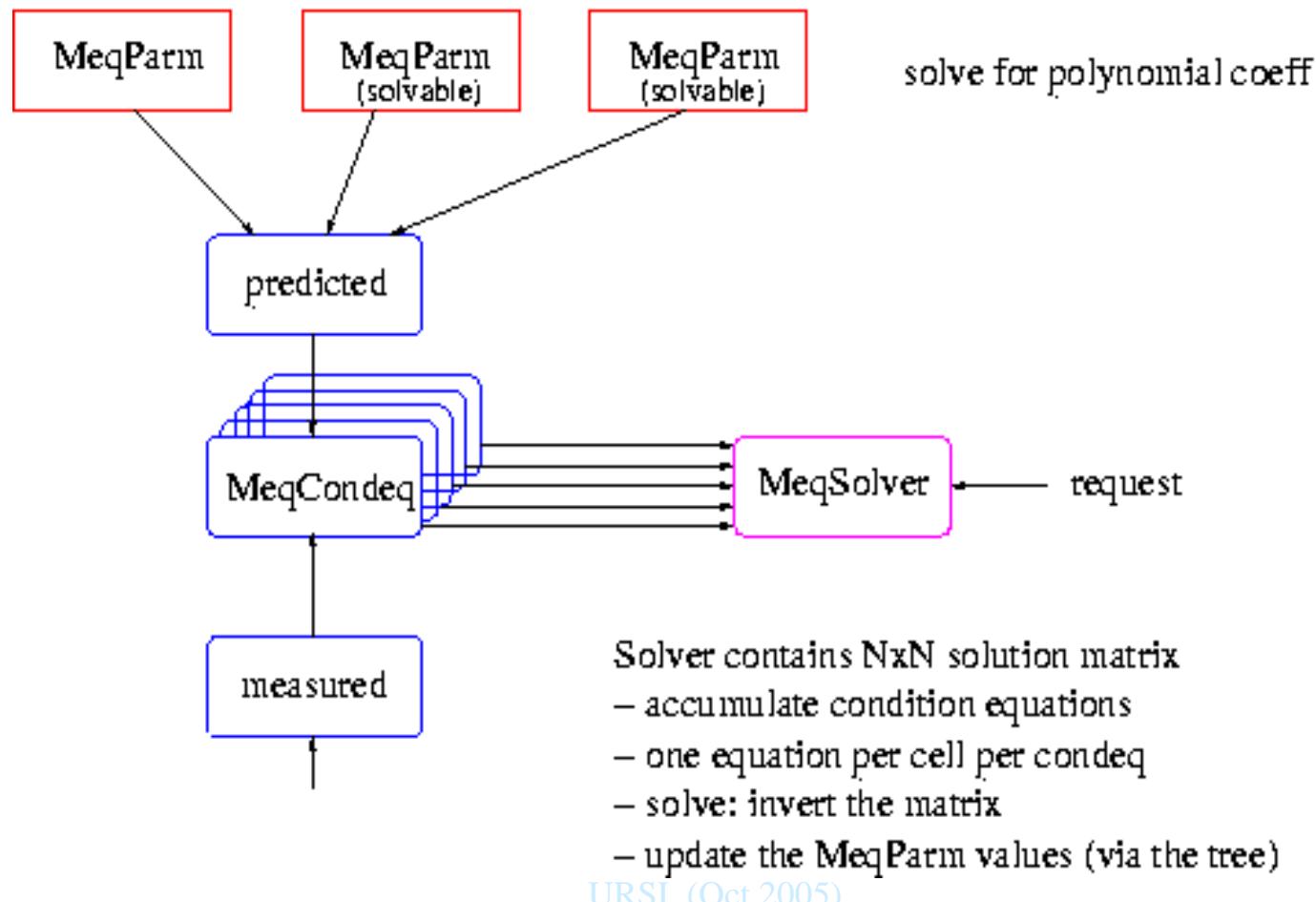
(or coeff of any other smooth function)



Funklets (cont'd)

- ◆ Most M.E. parameters have 2D (f,t) polynomial funklets
- ◆ But: funklets can have any number of dimensions, with different base ‘functionals’ along each axis
- ◆ Examples of complex funklets:
 - ⊕ The Minimum Ionospheric Model (MIM)
 - ⊕ Station voltage beam shapes

A solving tree



MeqSolver issues

- ◆ Based on AIPS++ fitting (WNB)
- ◆ SVD (always inverts)
- ◆ Non-linear (requires iteration)
- ◆ Use solver metrics to monitor quality
- ◆ Local intelligence to minimise iteration
- ◆ Resample data to minimise cells (=equations)
- ◆ Use tiled solutions to minimise node overhead
- ◆ Optimize inversion where possible
- ◆ Alternatives?

The many windows into MeqTrees

- ◆ Every node may be inspected individually
 - ❖ Contents of its state record
 - ❖ Plot of its result
- ◆ Built-in views of the results of groups of nodes
- ◆ History nodes collect intermediate results
- ◆ Debugger functionality (pause, resume, brktpnts)
- ◆ The meqbrower itself is a new kind of user interface (incl. editing of script hierarchy)

MeqBrowser with tree, script and views

meqserver.0.26742.9 - idle

MeqTimba IDL Bookmarks Debug View Help

Trees Tabs Grid

Tabs

2 errors Snapshots MG_JEN_cps_GDJones.py

Jobs L:76 C:43 MG_JEN_cps_GDJones.py

```
# punit = 'SItest'

MG = MG_JEN_exec.MG_init('MG_JEN_cps_GDJones.py',
    last_changed = 'd28sep2005',
    punit= 'QU',           # name of calibrator source
    stations=range(4),     # specify the (subset of)
    parmtable=None,        # name of MeqParm tab

    fdeg_Gampl=2,          # degree of freq polynomial
    fdeg_Gphase='fdeg_Gampl',
    tdeg_Gampl=1,          # degree of time polynomial
    tdeg_Gphase='tdeg_Gampl',
    tile_size_Gampl=None,   # used in tiled solution
    tile_size_Gphase='tile_size_Gampl',

    fdeg_dang=2,            # degree of freq polynomial
    fdeg_dell='fdeg_dang',
    tdeg_dang=1,             # degree of time polynomial
    tdeg_dell='tdeg_dang',
    tile_size_dang=None,      # used in tiled solution
    tile_size_dell='tile_size_dang',

    num_iter=10,             # number of solver iterations
    epsilon=1e-4,            # iteration stop criterion
    flag_spigots=False,       # If True, insert a flagger
    flag_sinks=False,         # If True, insert a flagger
    visu_spigots=True,        # If True, insert built-in viewer
    visu_solver=False,         # If True, insert built-in viewer
    visu_sinks=True,          # If True, insert built-in viewer
    trace=False)              # If True, produce progress messages

MG.stream_control = record(ms_name='/home/rnijboer/D1.H5',
    data_column_name='DATA',
    tile_size=10,                  # input tile-size
    channel_start_index=10,
    channel_end_index=50,           # -10 should indicate 10
    output_col='RESIDUALS')

MG = MG_JEN_exec.MG_check(MG)

#-----
```

Script executed successfully. 1822 node definitions (of which 5 are root nodes) sent to the kernel. 1 predefined function(s) available.

Gridded Viewers

allcorrs dconc_spigots_uvp_allcor > (MeDataCollect) rq 2.1.2.1spigots_uvp_allcorrs dconc_spigots_uvp_allcorrs: > (MeDataCollect) rq 2.3.2.1spigots_uvp_allcorrs

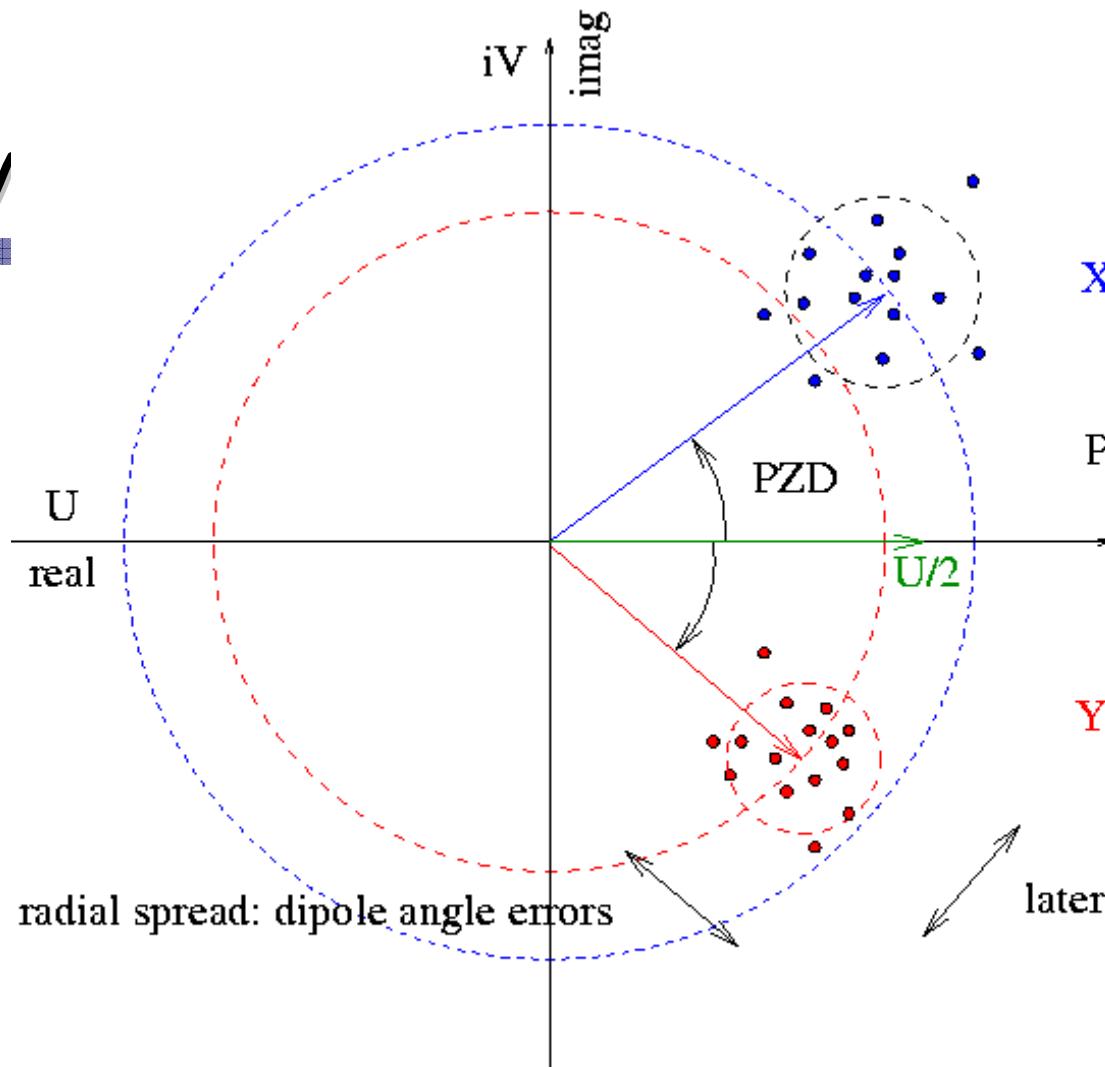
dconc_corrected_JGD_QU_allcor > (MeDataCollect) rq 2.3.2.0corrected_JGD_QU_allcorrs dconc_corrected_JGD_QU_allcor > (MeDataCollect) rq 2.3.2.0corrected_JGD_QU_allcorrs

VSZ:263/264M RSS:189/190M DS:226/227M

LUTAR

WSRT polarization calibration

XY



$$XY = (U/2) \exp(iPZD) + eI$$

PZD: X-Y phase zero diff

$$YX = (U/2) \exp(-iPZD) + eI$$

WSRT XY and YX data from a polarised (but: V=0) central point source, e.g. 3c286

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MeqBrowser with parm fiddler

meqserver.0.3306.10 - idle

MeqTimba View Bookmarks Debugger Help

node

- Forest state
- All nodes (1374)
- By class (22)
- Root nodes (3)
 - _ROOT_rqq_solver
 - _ROOT_dataCollect
 - noMS_root**
- 09: noMS_root[s1=4][s2=5]
- 08: noMS_root[s1=3][s2=5]
- 07: noMS_root[s1=3][s2=4]
- 06: noMS_root[s1=2][s2=5]
- 05: noMS_root[s1=2][s2=4]
- 04: noMS_root[s1=2][s2=3]
- 03: noMS_root[s1=1][s2=5]
- 02: noMS_root[s1=1][s2=4]
- 01: noMS_root[s1=1][s2=3]
- 00: noMS_root[s1=1][s2=2]
- 0: sc_25_dconc_result[s1=1][s2=2]
 - (dconc_9(result|cross))
 - (dconc_8(result|parallel))
 - (dconc_7(result|allcorrs))
- 0: visu_3_result[s1=1][s2=2]
- 0: rqq_solver_GD
 - 3: sc_17_dconc_corrected_GD[s1=1][s2=1]
 - 2: dcoll_6(realvsiimag:(d)parms_of_solver)
 - 1: dcoll_5(realvsiimag:condeqs_of_solver)
 - 0: solver_GD
- 39: condeq_4[s1=4][s2=5][c=YY]

Messages Trees Snapshots Events

idle

MeqTimba View Bookmarks Debugger Help

node

star solver_GD star uvdata_allcorrs

dconc_1(rawdata|allcorrs) (MeqDataCollect)

dconc 1 (rawdata|allcorrs):

Imaginary Axis

Real Axis

star uvdata_cross

dconc_4(corrected_GD|allcorrs) (MeqDataCollect)

dconc 4 (corrected_GD|allcorrs):

Imaginary Axis

Real Axis

star dconc_7(result|allcorrs)

(MeqDataCollect)

dconc 7 (result|allcorrs):

Imaginary Axis

Real Axis

Fiddle from noMS_root

| name | no parm selected |
|------|------------------|
| 53 | gain[s=2][p= |
| 54 | gain[s=3][p= |
| 55 | gain[s=3][p= |
| 56 | gain[s=4][p= |
| 57 | gain[s=4][p= |
| 58 | gain[s=5][p= |
| 59 | gain[s=5][p= |
| 60 | phase[s=1][r= |
| 61 | phase[s=1][r= |
| 62 | phase[s=2][r= |
| 63 | phase[s=2][r= |
| 64 | phase[s=3][r= |
| 65 | phase[s=3][r= |
| 66 | phase[s=4][r= |
| 67 | phase[s=4][r= |
| 68 | phase[s=5][r= |
| 69 | phase[s=5][r= |
| 70 | uvp_PZD |

stepsize: 1.0

reexecute

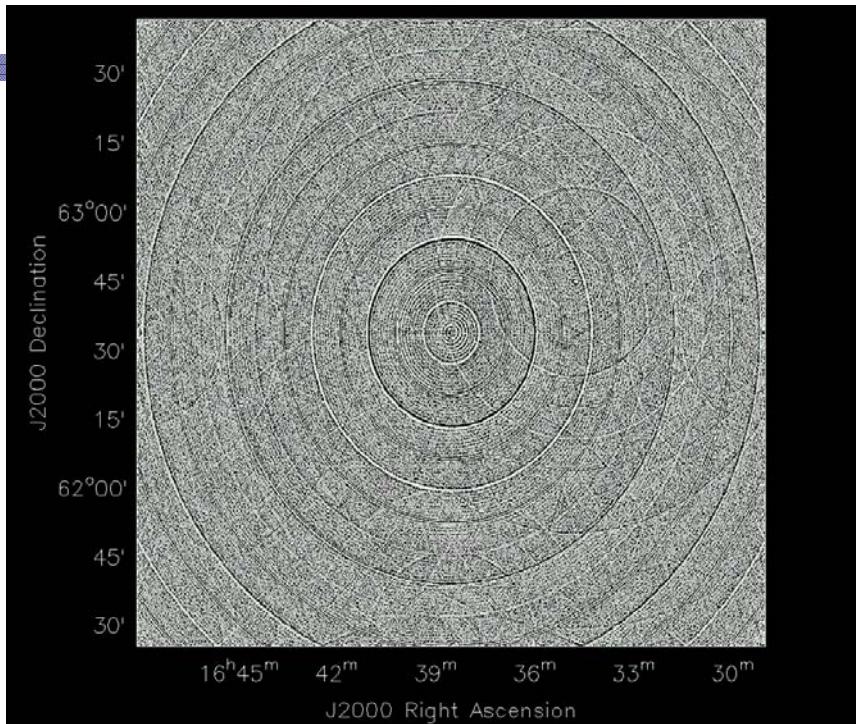
Execute Zero

VSZ:209/211M RSS:109/110M DS:188/189M

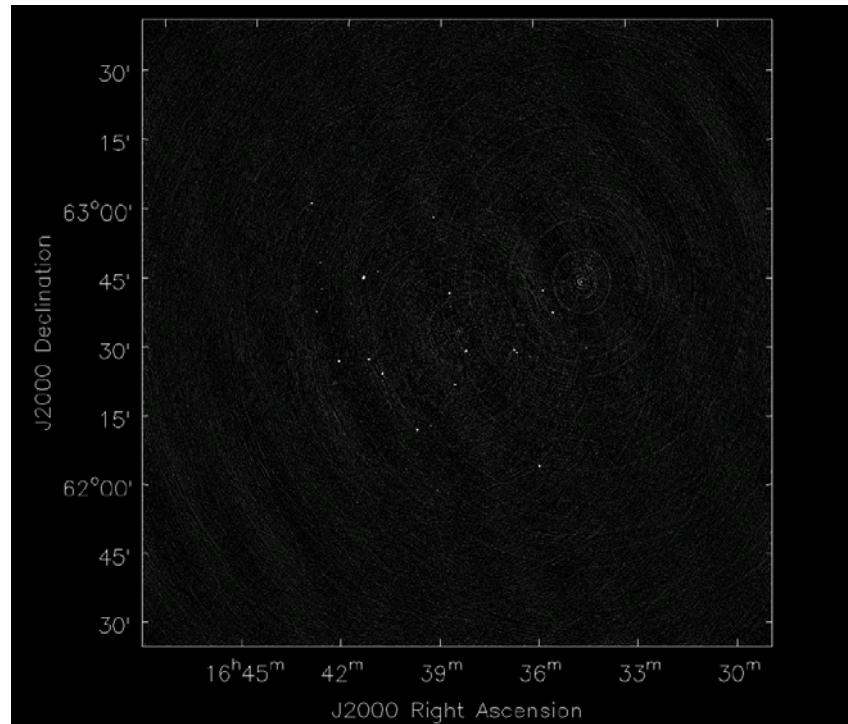
Results so far

- ◆ Stable kernel and browser
- ◆ 3c343 (Michiel Brentjens)
 - ❖ Also: Miriad, NEWSTAR, AIPS++
- ◆ M81/82 (Tom Oosterloo): Peeling in Miriad
- ◆ Minimim Ionospheric Model (MIM)
 - ❖ URSI poster by Jim Anderson
- ◆ Model of CLAR beam (Tony Willis)
- ◆ Model of LOFAR WHAT station beam

MeqTree: 3C343.1 & 3C343



Uncalibrated Image



Calibrated Image
2 sources subtracted

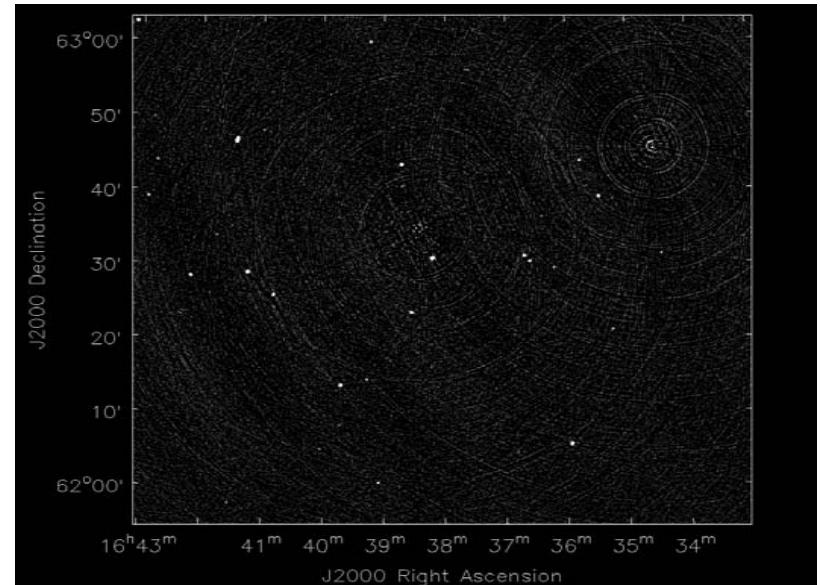
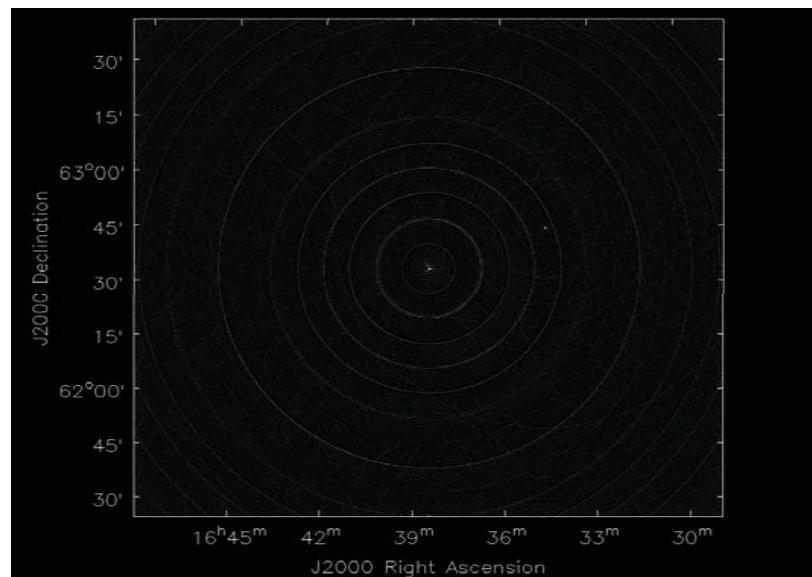
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PSS_5 JEN/2004/25/11

Result by Michiel Bremer



The MeqTree System is a reduction module that implements
Generalized Self-Cal and allows automated Picingeling



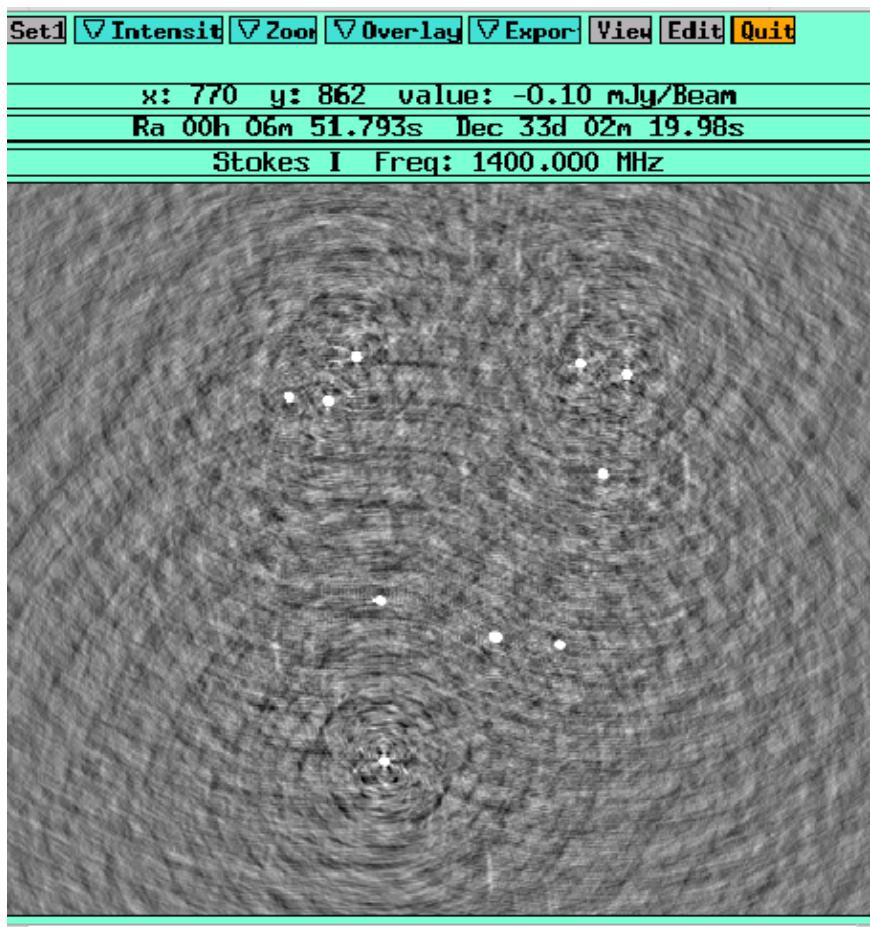
Result by Michiel Brentjes

LOFAR Self-Cal will be based on experience with the MeqTree System

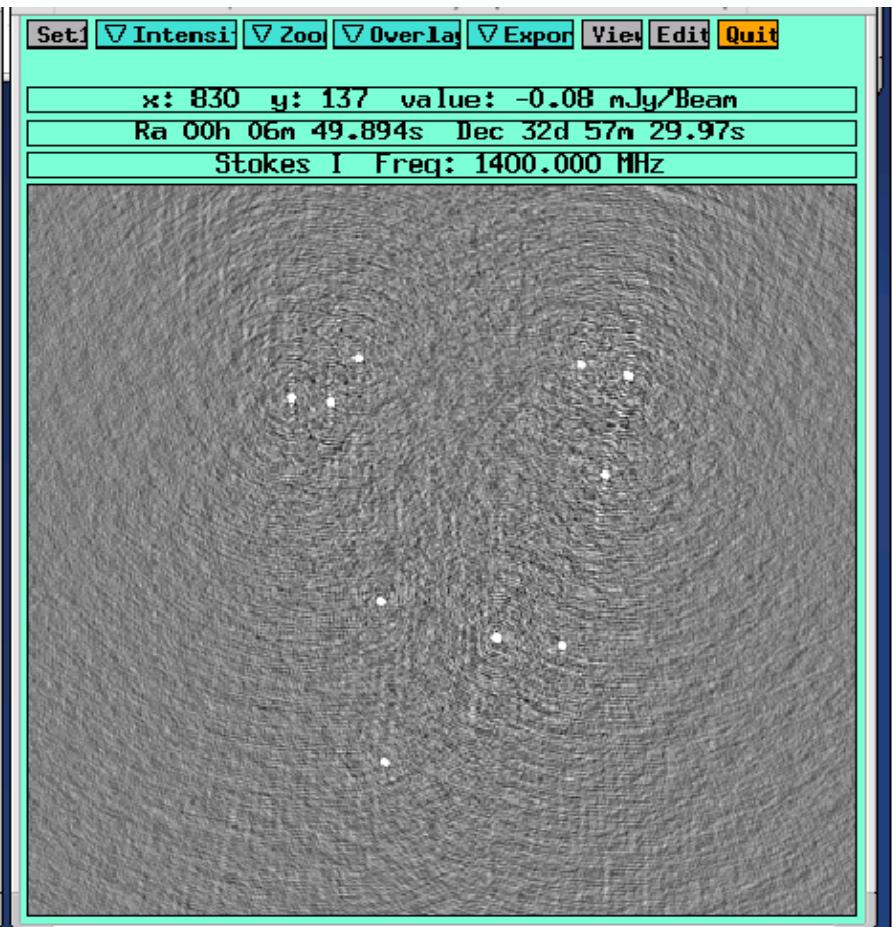
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First external user (Tony Willis, Penticton, Canada): CLAR simulations

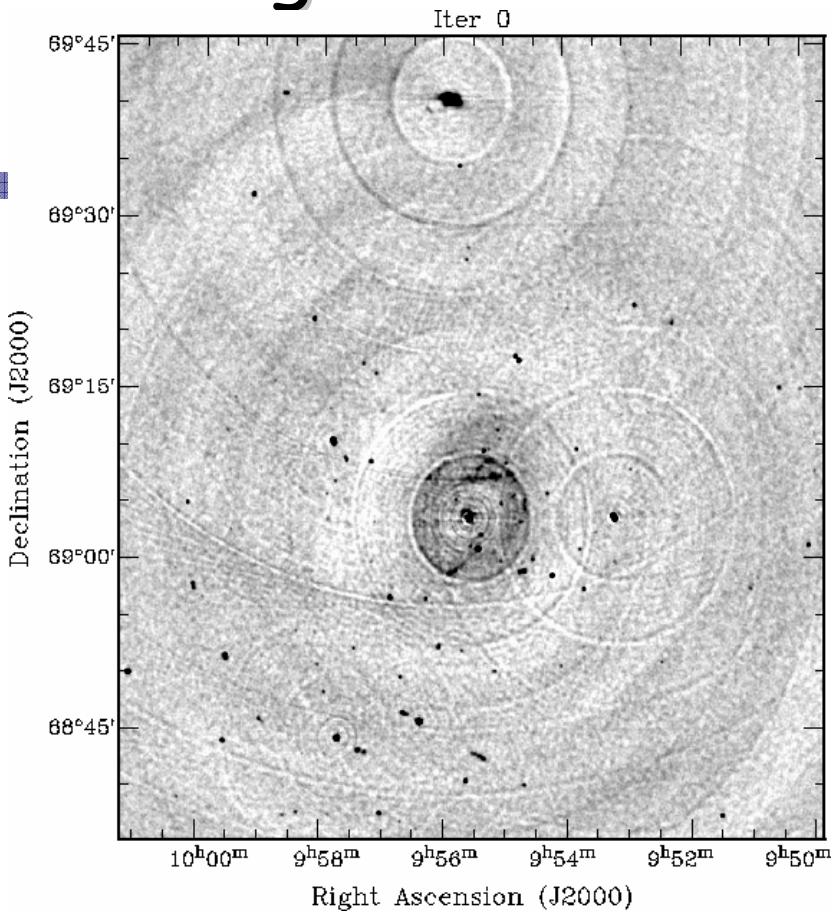
AIPS++ (20 min chunks, image-plane)



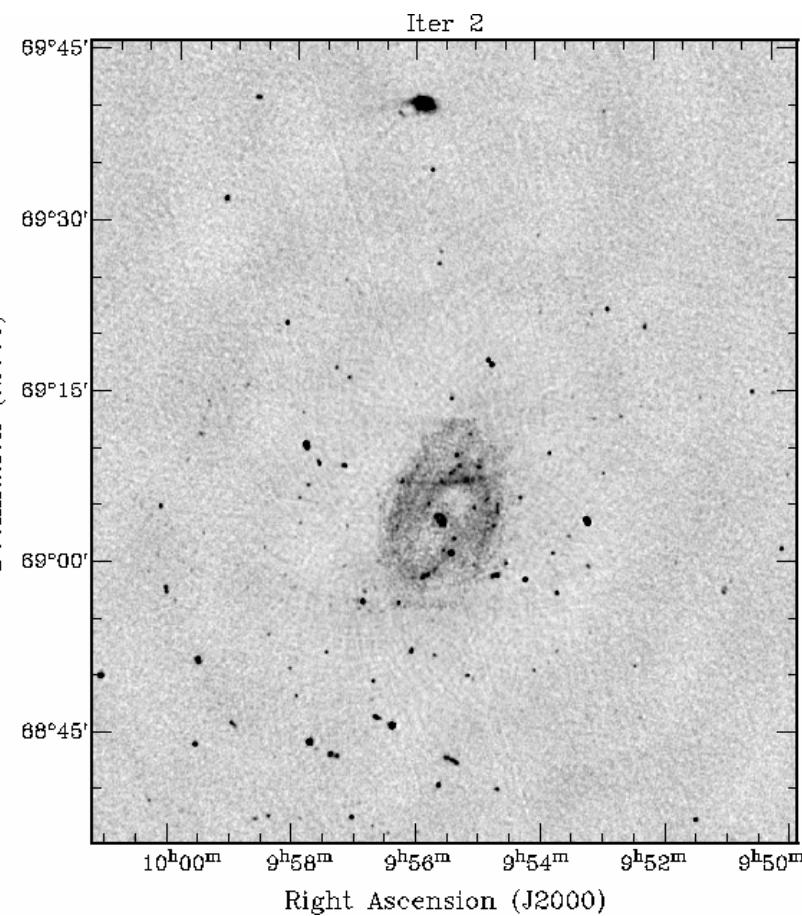
MeqTree (continuous, station-specific)



Peeling on M81 & M82



standard selfcal entire field



2-patch peeling (miriad)

(Result by Tom Oosterloo)

URSI (Oct 2005)



Conclusions

- ◆ LOFAR calibration will not be easy
- ◆ We have a lot of new ideas
- ◆ But more are needed
- ◆ The stage is set for M.E. development