

Today Instead of V constant in time over the period of application,

Today Instead of V constant in time we have $V = V(t)$

we take

$$V(t) = 2v \cos \omega t = v (e^{i\omega t} + e^{-i\omega t}), \quad \omega > 0$$

for $t \geq 0$

General: $P_{n \leftarrow m}(t) = \frac{1}{\hbar^2} \left| \int_0^t \langle \psi_n | V(t_1) | \psi_m \rangle e^{i\omega_{nm}t_1} dt_1 \right|^2$

$$\langle \psi_n^0 | V(t_1) | \psi_m^0 \rangle = v_{nm} e^{i\omega t_1} + v_{nm} e^{-i\omega t_1}, \quad v_{nm} = \langle \psi_n^0 | v | \psi_m^0 \rangle$$

$$\begin{aligned}
 \text{So } P_{n \leftarrow m} &= \frac{1}{\hbar^2} \left| \int_0^t v_{nm} e^{i\omega t_1} e^{i\omega_{nm} t_1} dt_1 + \int_0^t v_{nm} e^{-i\omega t_1} e^{i\omega_{nm} t_1} dt_1 \right|^2 \\
 &= \frac{|v_{nm}|^2}{\hbar^2} \left| \int_0^t e^{i(\omega + \omega_{nm})t_1} dt_1 + \int_0^t e^{i(\omega_{nm} - \omega)t_1} dt_1 \right|^2 \quad \text{for } \omega \ll 1 \\
 &\quad \text{if constant in time} \\
 &\quad (\omega = 0) \\
 &= \frac{|v_{nm}|^2}{\hbar^2} \left| \frac{e^{i(\omega_{nm} + \omega)t} - 1}{\underbrace{\omega_{nm} + \omega}_{\text{1st term}}} + \frac{e^{i(\omega_{nm} - \omega)t} - 1}{\underbrace{\omega_{nm} - \omega}_{\text{2nd term}}} \right|^2
 \end{aligned}$$

We introduce the rotating wave approximation (RWA)

Where the driving frequency ω is quite close to the (ω_{nn})
($\omega > 0$)

If $w_{nm} > 0$, 2nd term dominates

If $w_{nm} < 0$, 1st term dominates

denominators are almost zero.

In the RWA, we ignore/drop the non-dominant term.

Take the case $w_{nm} > 0$ as an example.

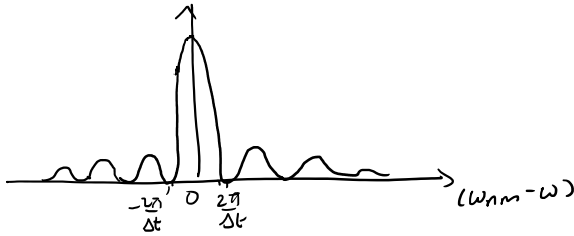
Take the case $\omega_{nm} > 0$ as an example.
Using the RWA, $P_{n \leftarrow m} = \frac{|U_{nm}|^2}{\hbar^2} \left| \frac{e^{i(\omega_{nm}-\omega)t} - 1}{\omega_{nm} - \omega} \right|^2 = \frac{4|U_{nm}|^2}{\hbar^2 (\omega_{nm} - \omega)^2} \sin^2 \left(\frac{(\omega_{nm} - \omega)t}{2} \right)$

$$W1Z1, \omega=0$$

$$P_{ngm} = \frac{|V_{nm}|^2}{\hbar^2} \left| \frac{e^{i\omega_{nm}t} - 1}{\omega_{nm}} \right|^2$$

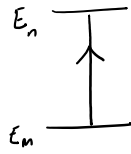
$$\omega_{nm} \rightarrow \omega_{nm} - \omega$$

Fixed Δt , How does $P_{n \leftarrow m}$ depend on ω_{nm} ? (ω is fixed)



$P_{n \leftarrow m}$ is max when $\omega_{nm} = \omega$ (resonance condition)
 driving frequency.
 transition freq $\frac{E_n - E_m}{\hbar}$ (determined by H_0)

Great
 Van matrix elements
 should be non-zero
 ↓
 selection rules



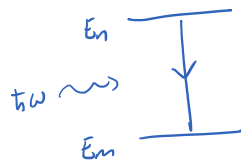
Absorption / Excitation

Intro QM:

$$E_n = E_m + \hbar\omega$$

energy of the photon.

$$\hat{V}(t) = 2\hat{V}\cos\omega t \text{ (EM wave ; photons)}$$



stimulated emission / deexcitation
 (used in lasers)

Fermi's Golden Rule for constant transition rates

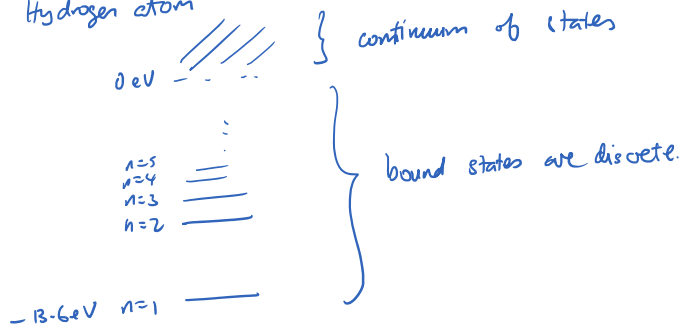
Transition rate = Transition probability per unit time. (→ allows to define lifetimes)

Constant transition rate \Rightarrow transition probability \propto time
 (in time)

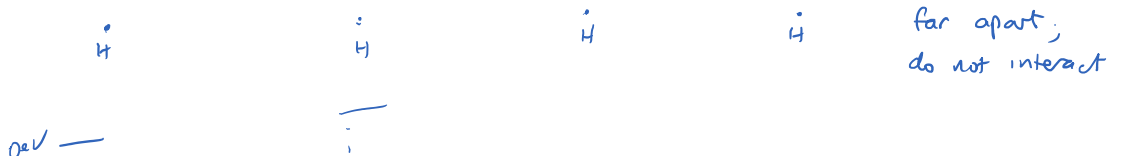
Problem Formulation

- Time-dependent perturbation theory
- V is either constant in time or harmonic (within the RWA) ie. close to resonance.
- ★ Transition from an initial state to a continuum of final states

Eg. Hydrogen atom



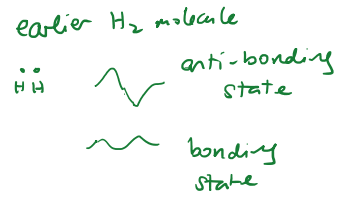
Eg. Solid state physics (condensed phase) - bands.



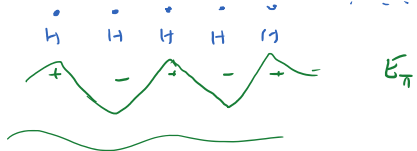


discrete bound states

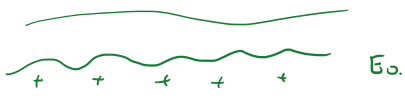
Now bring the H atoms together closer



adjacent all s-orbitals are exactly out of phase



all s-orbitals are in-phase



In between 0 and π , we have other relative phases. - this results in a continuum of energies between E_0 and E_π .

Eg. where we have a continuum of final states

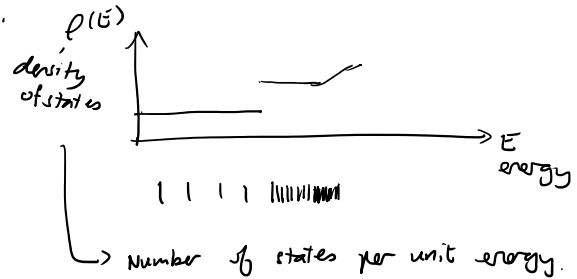
- final state is unbound; or
- final state is a molecular or atomic state coupled to a surface.

(Q) How does having a continuum of final states lead to $P_{trans} \propto (\Delta E)$ at resonance?

- How do we describe a continuum of states?
We use the "density of states".

$$P_{f \in i} = \sum_{j=1}^{large N} P_{f_j \in i}$$

continuum
energy spacing is small



$$= \int \underbrace{\rho(E_f)}_{\text{density of states}} \underbrace{P_{f \in i}(E_f)}_{\text{"weights" for the sum}} dE_f$$

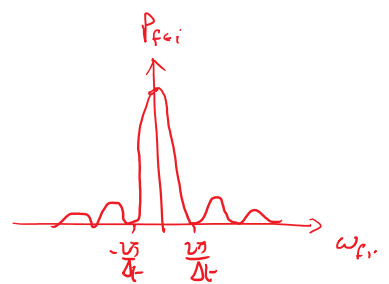
Now we work out the transition probability from perturbation theory.
consider $\omega = 0$. ← P_{trans} from W12L1

$$\begin{aligned} \text{Probability} &= \int P_{f \in i} \rho(E_f) dE_f \\ &= \int \frac{4|V_{fi}|^2}{\hbar^2} \frac{\sin^2 \frac{\omega_{fi} t}{2}}{\omega_{fi}^2} \rho(E_f) dE_f \end{aligned}$$

no additional approx. so far.

$$= \int_{-\infty}^{\infty} \frac{4|V_{fi}|^2}{\hbar^2} \frac{\sin^2 \frac{\omega_{fi} t}{2}}{\omega_{fi}^2} \rho(E_f) dE_f$$

new



new

Approx #1:

Secondary peaks contribute very little to the integral.

$$= \frac{4 |V_{fi}|^2 \rho(E_{f_0})}{\hbar^2} \int_{-\infty}^{\infty} \frac{\sin^2 \frac{\omega_{fi} t}{2}}{\omega_{fi}^2} dE_{fi} \quad \text{where } E_{f_0} = E_i.$$

new

Approx #2:

middle of primary peak contributes the most

and picks up $E_{fi} = 0$, and $|V_{fi}| \rho(E_{fi})$ is approximately constant over the energy interval

$(-\frac{2\pi}{\Delta t}, \frac{2\pi}{\Delta t})$

$$= \frac{4}{\hbar^2} |V_{f_0 i}|^2 \rho(E_{f_0}) \int_{-\infty}^{\infty} \frac{\sin^2 \left(\frac{E_{fi} t}{2\hbar} \right)}{\left(\frac{E_{fi}}{\hbar} \right)^2} dE_{fi}, \quad \omega_{fi} = \frac{E_{fi}}{\hbar}$$

$$\int_{-\infty}^{\infty} \frac{\sin^2 u}{u^2} du = \pi$$

$$\text{let } u = \frac{E_{fi} t}{2\hbar}$$

$$du = \frac{t}{2\hbar} dE_{fi}$$

$$= 4 |V_{f_0 i}|^2 \rho(E_{f_0}) \int_{-\infty}^{\infty} \frac{\sin^2 u}{\left(\frac{2\hbar}{t} \right)^2 u^2} \left(\frac{2\hbar}{t} \right) du$$

$$= \frac{2}{\hbar} |V_{f_0 i}|^2 \rho(E_{f_0}) t \int_{-\infty}^{\infty} \frac{\sin^2 u}{u^2} du$$

$$= \frac{2\pi}{\hbar} |V_{f_0 i}|^2 \rho(E_{f_0}) t, \quad E_{f_0} \approx E_i$$

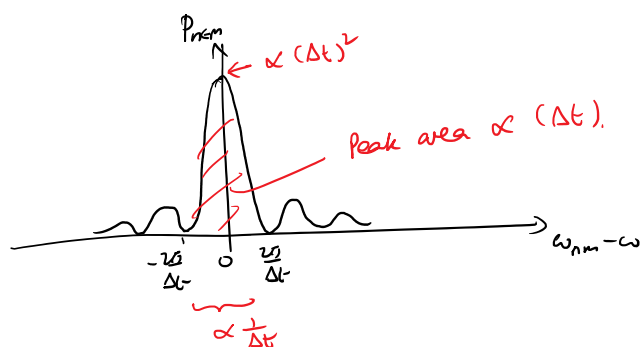
We see that the probability is proportional to t .

$$\text{Constant transition rate} = \frac{2\pi}{\hbar} |V_{f_0 i}|^2 \rho(E_{f_0}), \quad E_{f_0} \approx E_i, \quad \omega = 0.$$

$$\langle f_0 | \hat{V} | i \rangle$$

for $\omega \neq 0$

$$\text{Constant transition rate} = \frac{2\pi}{\hbar} |V_{f_0 i}|^2 \rho(E_{f_0}), \quad E_{f_0} \approx E_i \pm \hbar \omega, \quad (\pm \text{ depends on signs of } \omega_{fi}, \omega)$$



(essence of Fermi Golden rule)

Now we focus on the matrix element $|V_{nm}|$ which leads to 'selection rules'.

Example: Light-matter interaction.

Recall: Zeeman effect EM wave

$$H = H_0 + \frac{e}{m} \vec{A} \cdot \vec{p}$$

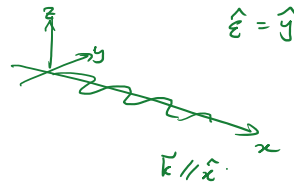
$(q = -e)$ (ignored A^+ term)

Light is an EM wave.

Light $V(t) = \frac{e}{m} \vec{A}(t) \cdot \vec{p}$

$$\vec{A} = A_0 \hat{\epsilon} \cos(\vec{k} \cdot \vec{r} - \omega t)$$

$\hat{\epsilon}$ denotes the polarization of light (unit vector)
 \vec{k} : wave vector.



$$\vec{E} = -\frac{\partial \vec{A}}{\partial t}$$

$$V(t) = \frac{e}{m} A_0 \cos(\vec{k} \cdot \vec{r} - \omega t) \hat{\epsilon} \cdot \vec{p}$$

$$= \frac{e A_0}{m} \frac{1}{2} (e^{i \vec{k} \cdot \vec{r}} e^{-i \omega t} + e^{-i \vec{k} \cdot \vec{r}} e^{i \omega t}) \hat{\epsilon} \cdot \vec{p}$$

So far, no approx:

Approx:

$$e^{\pm i \vec{k} \cdot \vec{r}} = 1 \quad \text{--- called the "dipole approx".}$$

$$\text{So } V(t) = V(e^{i \omega t} + e^{-i \omega t}) \quad , \quad V = \frac{e A_0}{2m} \hat{\epsilon} \cdot \vec{p}$$

Why is this mostly valid?
 λ of light \sim few hundred nm

$$\vec{r} \sim \vec{A}$$

$$|\vec{k}| = \frac{2\pi}{\lambda} \quad |\vec{k} \cdot \vec{r}| \approx 0, \quad r \sim a_0$$

$\lambda \gg a_0$ Bohr radius

Approx #2 — RWA.

$$P_{e \rightarrow g}(t) = \frac{4 |V_{eg}|^2 \sin^2((\omega - \omega_{eg})t/2)}{\hbar^2 (\omega - \omega_{eg})^2}$$

$$\text{where } |V_{eg}|^2 = \frac{1}{4} \left| \frac{e A_0}{m} \right|^2 \underbrace{|\langle e | \hat{\epsilon} \cdot \vec{p} | g \rangle|^2}_{\text{momentum matrix elements}}$$

Can also be written in terms of dipole matrix elements:

$$\langle e | \hat{\epsilon} \cdot \underbrace{q \vec{r}}_{\text{dipole}} | g \rangle$$

To express V_{eg} in terms of \vec{r} matrix elements, we use:

$$[\vec{r}, H]$$

$$\vec{p} = \frac{m}{i\hbar} [\vec{r}, H_0]$$

to express \vec{p} in terms of \vec{r}

$$\begin{aligned}
 & [\vec{r}, H_0] \\
 &= [\vec{r}, \frac{\vec{p}^2}{2m} + \tilde{V}(r)] \\
 &= [\vec{r}, \frac{\vec{p}^2}{2m}] \\
 &= \frac{1}{2m} [\vec{r}, p^2] \\
 &= \frac{1}{2m} (2i\hbar \vec{p}) \\
 &= i\hbar \frac{\vec{p}}{m}
 \end{aligned}$$

$$\boxed{\vec{p} = \frac{m}{i\hbar} [\vec{r}, H_0]}$$

valid when
potential \tilde{V} for H_0
depends only on r .

$$\begin{aligned}
 & [\vec{r}, \vec{p} \cdot \vec{p}]_i \\
 &= [r_i, \vec{p} \cdot \vec{p}] \\
 &= [r_i, p_j p_j] \quad (\text{summation notation}) \\
 &= [r_i, p_j] p_j + p_j [r_i, p_j] \\
 &= 2i\hbar \delta_{ij} p_j \\
 &= 2i\hbar p_i
 \end{aligned}$$

momentum matrix element

$$\begin{aligned}
 \langle e | \hat{E} \cdot \vec{p} | g \rangle &= \langle e | \hat{E} \cdot \frac{m}{i\hbar} [\vec{r}, H_0] | g \rangle \\
 &= \frac{m}{i\hbar} \langle e | \hat{E} \cdot (\vec{r} H_0 - H_0 \vec{r}) | g \rangle \\
 &= \frac{m}{i\hbar} \langle e | \hat{E} \cdot (\vec{r} E_g - E_e \vec{r}) | g \rangle \\
 &= \frac{m}{i\hbar} \underbrace{(E_g - E_e)}_{\text{energy term}} \underbrace{\langle e | \hat{E} \cdot \vec{r} | g \rangle}_{\text{position matrix element}}
 \end{aligned}$$

Next lecture — work out selection rules for the specific case where the states are spherical harmonics.

— adiabatic approx.

Canvas — Past two years' final exam papers.