

AXION-LIKE PARTICLES SEARCH WITH ULTRACOLD NEUTRONS

3.1 MOTIVATION

3.2 THEORETICAL BACKGROUND

3.3 TIME SERIES CONSTRUCTION

3.3.1 *nEDM @ PSI data taking scheme*

The nEDM experiment at PSI probes the Ramsey resonance curve of ultra-cold neutrons (UCNs). Its operation consists of subsequent *cycles*, wherein an ensemble of polarised UCNs undergoes a Ramsey cycle in a magnetic field. At the end the polarisation of the ensemble is measured. The measurement is performed in a controllable electric field, which takes up three values: parallel to the magnetic field, antiparallel to it and zero. The electric field configuration is altered every couple of tens of cycles.

The Ramsey resonance curve of the neutrons is probed only in four *working points*, as shown in Fig. 4. The points lie on the curve's steep slope. Thereby the resonance position is determined more precisely than it would should the curve be probed homogeneously or close to the resonance.

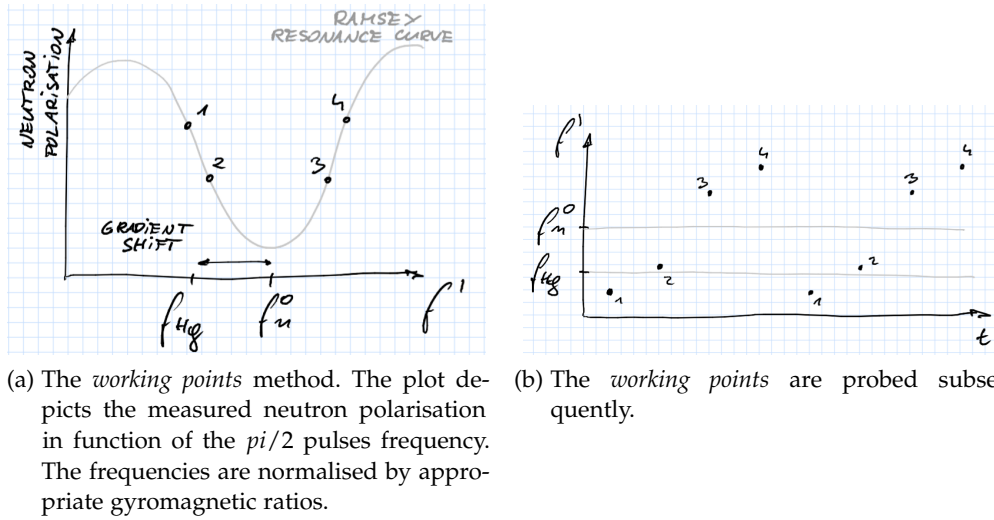


Figure 4: The principle of working points

The resonance frequency f_n^0 is determined with a fit of the resonance curve. Because the points are probed one after another, it is only possible after a set of data points, *cycles*, have been measured. In order to extract the neutron precession frequency for each individual *cycle*, one assumes that the only parameter of the res-

onance curve that varies on a cycle-to-cycle basis is the position of the resonance. With this assumption the shape of the curve, fitted to the whole set of *cycles*, is used to calculate back the resonant frequency in each *cycle* of the set.

Together with the UCNs there is a polarised ^{199}Hg gas precessing. Its precession is monitored with light, allowing for direct determination of the ^{199}Hg Larmor frequency and thus the magnetic field strength. In order to cancel the first-order magnetic field changes one looks at the value:

$$R := f_n / f_{\text{Hg}} . \quad (4)$$

However, the UCNs are cold enough to have their centre of mass shifted downwards a few centimeters by the gravity. The ^{199}Hg gas, being much hotter than the UCNs, fills the precession volume homogeneously. In a presence of a vertical magnetic field gradient this causes the two species to see different magnetic fields.

In the nEDM experiment at PSI data taking is divided into *runs*. A single *run* is carried out automatically with, in most cases, no human intervention. The machine cycles through the working points by itself. Also the charging of the electrodes creating the electric field is automatised. In between *runs* manual interference happens, most importantly the magnetic field vertical gradient is altered.

In order to mitigate the bias due to the human factor, the nEDM experiment implements *data blinding*. The data is artificially altered in a way that mimics a non-zero neutron electric dipole moment, big enough to be visible in the data. The exact value is secret and will be revealed only after the analysis is complete.

3.3.2 An oscillating nEDM signal

The main purpose of the experiment is to measure the constant neutron electric dipole moment. This would appear as a shift in R dependent on the electric field. Zero electric field would cause no shift, while the parallel and anti-parallel configurations of the magnetic and electric fields would shift R in opposite directions. Due to the *data blinding* the shift is expected even in case of zero nEDM.

Should the neutron electric dipole moment oscillate, R would oscillate as well. Even if the electric field is kept constant. A reversal of the electric field polarity would change phase of the oscillations by π . At zero electric field no oscillations would be visible. The effect is depicted in Fig. 5a.

The vertical magnetic field vertical gradient changes when a new *run* is started. Thereby R is shifted by a big value, changing the DC level of the oscillating nEDM signal, which is illustrated in Fig. 5b. Moreover, even during a single *run* the gradient drifts, as clearly visible in Fig. 6.

The nEDM team spares no effort to measure the gradient. Nevertheless, the achieved precision ($\approx 1 \text{ pT/cm}$) is only comparable to the one of f_n (in the order of 1 pT). The exact way how the gradient should be determined is highly non-trivial and there is ongoing research in this respect. Actually, even the height difference between the neutrons and ^{199}Hg centres of mass (a few millimeters) is still discussed.

One should note, that any, including an oscillating one, nEDM effect affects only the position of the neutrons' resonance. The shape of the resonance curve

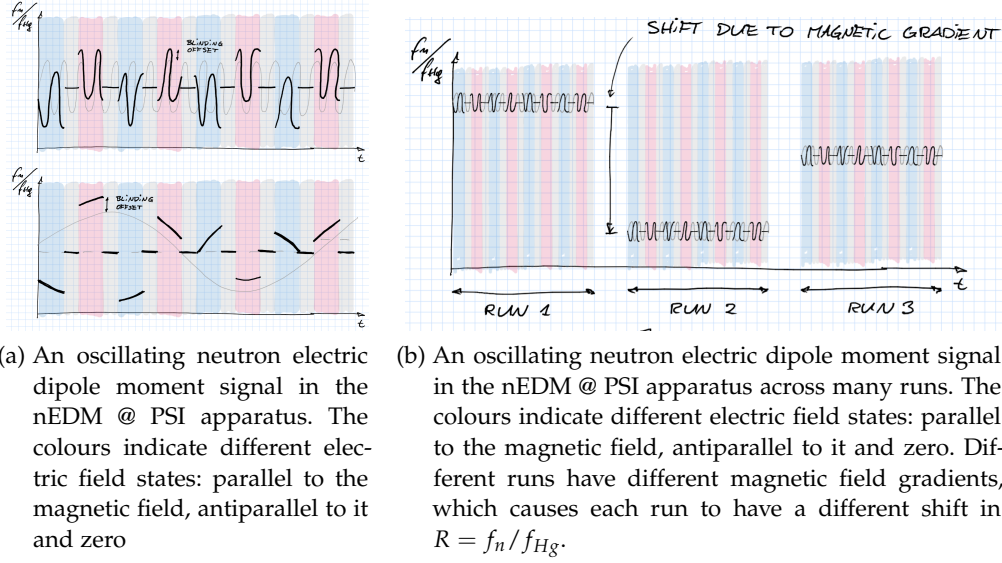


Figure 5: The data taking scheme in the nEDM experiment at PSI.

is unaffected. Therefore, the method to extract neutron Larmor frequency f_n , and thereby R , for each *cycle* is valid also in case of an oscillating nEDM.

3.3.3 R time series demodulation

The time series of R is not eligible for a Fourier-type analysis. The series has to be first demodulated into a coherent signal. To account for the electric field changes, data taken at one configuration need to be *flipped* around the DC level. Unfortunately, determining the DC level is not trivial.

Taking a look at Fig. 6 it becomes clear, that the R value drifts. The main reason is a drift of the vertical gradient of the magnetic field. This effect could be corrected for using the gradient measurement, as shown in Fig. 7. It is not straightforward, though, as there is not yet an established method to determine the gradient. The points taken at zero electric field provide additional information about the DC level. This, however, would have to be interpolated. The gradient drift correction is likely to turn out to be the most delicate part of the analysis.

3.4 TIME SERIES ANALYSIS

3.4.1 Definition of the periodogram

The analysis starts with calculating the *periodogram*, an estimator of the power spectral density, of the demodulated R time series. Therefore an array of N frequencies ω_i is chosen. For each a linear least-squares fit is performed with a model:

$$R(t) = A \sin \omega t + B \cos \omega t . \quad (5)$$

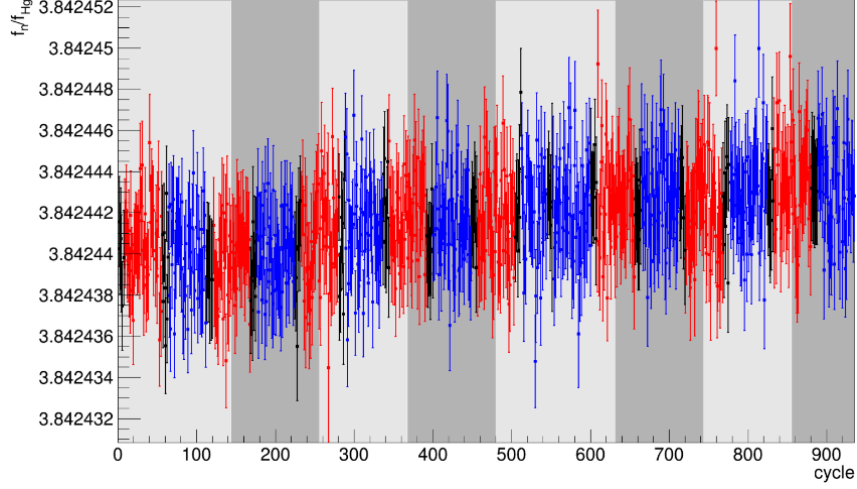


Figure 6: A typical time series of R in the nEDM experiment. The colours depict electric field states, black being no electric field. A drift is clearly visible.

Then the power at the frequency ω is defined:

$$P(\omega) = \frac{N}{4} (A^2 + B^2) . \quad (6)$$

The definition is similar to the one used by [Van Tilburg et al. \[12\]](#), but uses normalisation as defined by [Scargle \[9\]](#). One should remember, that the periodogram is an estimator. Its statistical properties are thoroughly described by [Scargle \[9\]](#).

3.4.2 Null hypothesis test

Once the periodogram of the real data, demodulated R time series, is calculated we look for a signal. A signature of a periodic signal is a peak in the periodogram. The really interesting statement is the answer to the question:

How likely it is, that the highest peak in the periodogram cannot be only a random fluctuation?

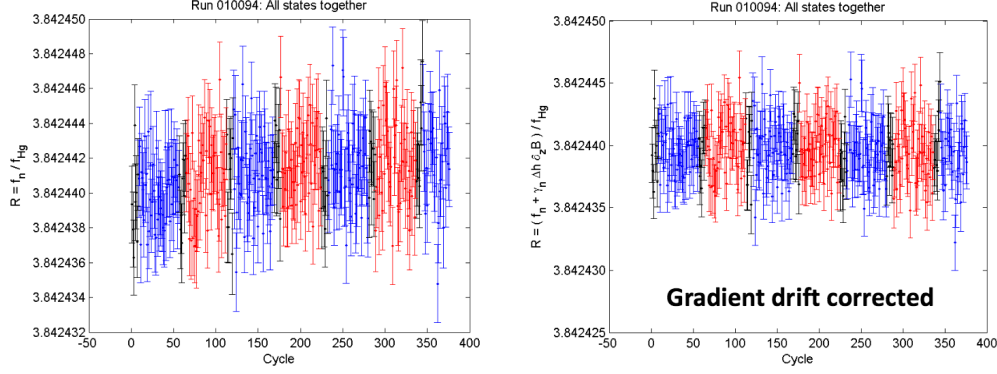
To describe the question mathematically, let us denote the time series from the real data by D . The periodogram is then a set of $P^D(\omega_i)$. We are interested in *the maximum of the periodogram*:

$$P_{max}^D := \max_i P^D(\omega_i) \quad (7)$$

$$\omega_{max}^D := \arg \max_i P^D(\omega_i) \quad (8)$$

The height of the maximum, P_{max}^D , is a random variable. We consider the distribution of P_{max} given the null hypothesis, H_0 , where an array of normally distributed random variables is taken. The probability that a peak at least as high as the one observed arises as a random fluctuation is:

$$\Pr \left(P_{max} > P_{max}^D \mid H_0 \right) . \quad (9)$$



(a) Another time series of R in the nEDM experiment. The colours depict electric field states, black being no electric field. A drift is clearly visible. (b) The data as on Fig. 7a corrected for gradient fluctuations.

Figure 7: Correcting the R time series for fluctuations of the vertical magnetic field gradient.

This value is called the *false alarm probability* (see eg. Pandola [8]). It can be numerically calculated with a Monte–Carlo method by generating random data according to the null hypothesis and counting the relative number of cases when $P_{max} > P_{max}^D$. To claim a discovery the *false alarm probability* has to be at most in the range of 10^{-7} (so-called 5-sigma).

It should be stressed that the distribution of P_{max} is very different from the distribution of $P(\omega_{max})$. The highest peak will always occur in a far tail of the $P(\omega_{max})$ distribution, which is natural. By looking for the highest peak we check a big number of random variables $P(\omega_i)$ and pick the one that does lie the furthest in the tail of the distribution. The distribution of P_{max} is centered around much higher values, as the highest peak is bound to occur *somewhere*. It is explained in Fig. 8.

3.4.3 Signal hypotheses tests

Should no claim for a discovery be possible, the next question to ask is:

Which oscillations would produce a visible peak, but did not, and can be thus excluded?

In order to answer this question, the data need to be tested against being compatible with a number of model signal hypotheses. As an oscillation is characterised by its amplitude and frequency, the space of the hypotheses to test is two-dimensional.

The probability that a hypothetical oscillation of amplitude A and frequency ω would produce more power at ω than observed is:

$$\Pr \left(P(\omega) > P^D(\omega) \mid H(\omega, A) \right) . \quad (10)$$

This probability is the *confidence level* (C.L.) at which the hypothesis $H(\omega, A)$ can be rejected. Figure 9 presents comparing a not-yet-real data power spectrum with

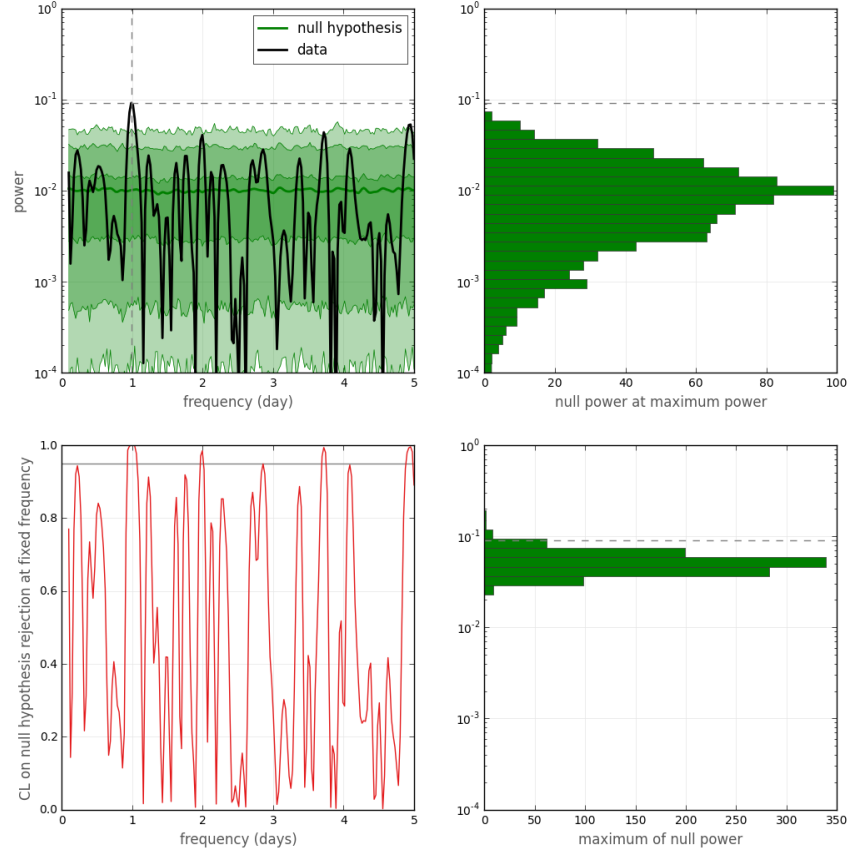


Figure 8: Testing not-yet-real data against the null hypothesis. UPPER-LEFT: The periodogram of the data (black line) on top of the periodogram of the null hypothesis (green line with Monte-Carlo generated distribution bands). UPPER-RIGHT: The distribution of $P(\omega_{max})$, given the null hypothesis (vertical slice through the green part of the plot in the upper-left corner). LOWER-LEFT: CDF of $P(\omega)$ evaluated at $P^D(\omega)$ in function of ω . LOWER-RIGHT: The distribution of P_{max} .

a hypothetical signal. This test may be performed many times, each covering a *pixel* of the space of possible hypotheses, forming an image as in Fig. 10a. The set of hypotheses excluded at a certain C.L. (often 95%) forms an *exclusion region*.

The exclusion region, depicted by a white line in Fig. 10a, exhibits a number of thin peaks going down in very low amplitudes. Seemingly for some frequencies signals with amplitude far below the sensitivity of the experiment are excluded. This is disturbing. Consider, however, that as the power was evaluated for many frequencies, inevitably at some of them, roughly 5%, the power is low enough to be rejected at 95% C.L. even when tested against a white noise. However completely fine from the statistical point of view, physicists do not accept a situation, where a hypothesis is rejected based on an experiment which was not sensitive to it. One possible solution is called the *CLs method*. The method is defined, as well the problem itself discussed, in the booklet of Particle Data Group [6]. With use of the *CLs method* the exclusion final region is obtained, as shown in Fig. 10b.

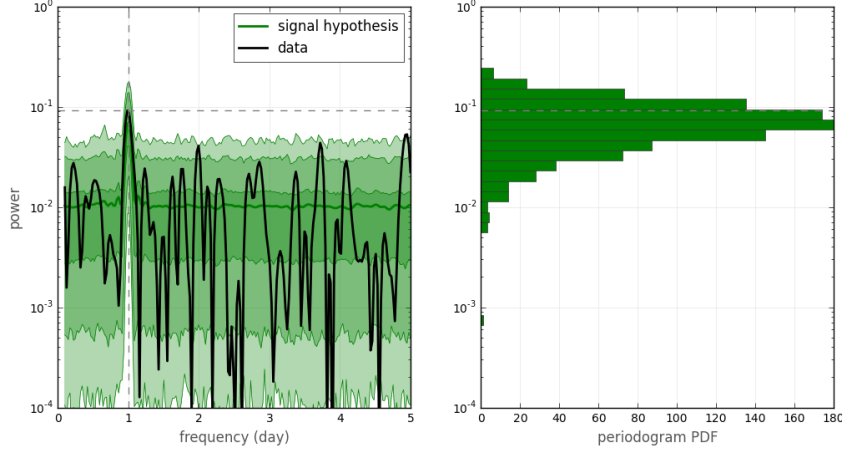


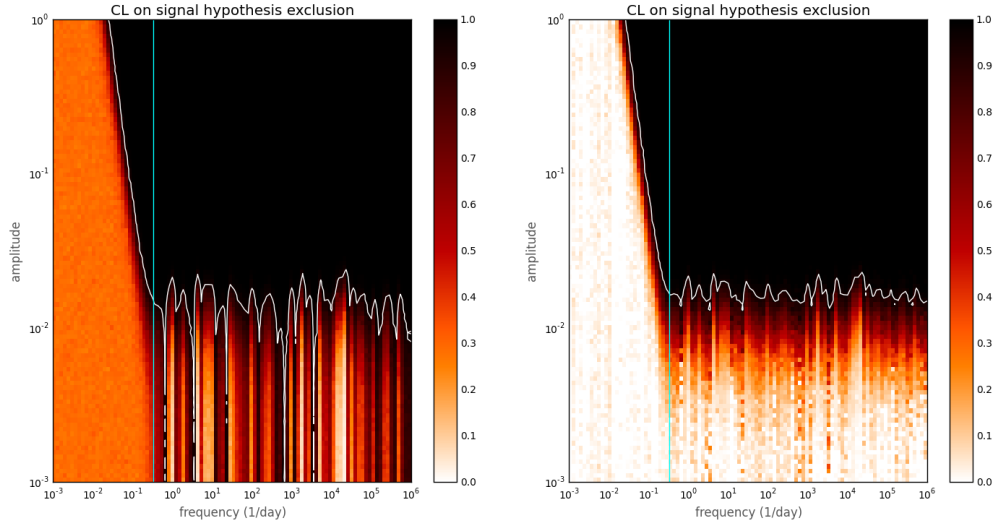
Figure 9: LEFT: A periodogram of not-yet-real data on top of distribution of a periodogram of a hypothetical signal (green). RIGHT: The distribution of power of the hypothetical signal at its model frequency.

3.5 DEDICATED MEASUREMENT

One could consider performing a dedicated measurement for the ALP search. With a short, more densely sampled measurement one could explore the high mass (thus high frequency) region of possible axions, as was done eg. by [Van Tilburg et al. \[12\]](#).

The measurement would not need to last longer than a day. More dense sampling is possible by shortening the length of a *cycle*. One should note that this worsens at the same time the sensitivity of the measurement. The lowest possible period of measurement in the nEDM apparatus is around 100 s.

3.6 RESULTS



- (a) The not-yet-real data tested against hypothetical signals. Each pixel is one signal hypothesis. The white line connects points of 95% C.L., surrounding an exclusion region. Note how deep into low amplitudes the line goes for couple of frequencies. See the text for the explanation.
- (b) The not-yet-real data tested against hypothetical signals using the *CLs method*, in which hypotheses to which the experiment is not sensitive to get a statistical penalty.

Figure 10: Exclusion region — signals that can be excluded at 95% confidence level.

BIBLIOGRAPHY

- [1] S. Afach, G. Bison, K. Bodek, F. Burri, Z. Chowdhuri, M. Daum, M. Fertl, B. Franke, Z. Grujic, V. Hélaine, R. Henneck, M. Kasprzak, K. Kirch, H.-C. Koch, A. Kozela, J. Krempel, B. Lauss, T. Lefort, Y. Lemièrre, M. Meier, O. Naviliat-Cuncic, F. M. Piegsa, G. Pignol, C. Plonka-Spehr, P. N. Prashanth, G. Quémener, D. Rebreyend, S. Roccia, P. Schmidt-Wellenburg, A. Schnabel, N. Severijns, J. Voigt, A. Weis, G. Wyszynski, J. Zejma, J. Zenner, and G. Zsigmond. Dynamic stabilization of the magnetic field surrounding the neutron electric dipole moment spectrometer at the Paul Scherrer Institute. *Journal of Applied Physics*, 116(8):084510, aug 2014. ISSN 0021-8979. doi: 10.1063/1.4894158. URL <http://scitation.aip.org/content/aip/journal/jap/116/8/10.1063/1.4894158>.
- [2] Jon Bentley. *Programming Pearls*. Addison-Wesley, Boston, MA, USA, 2nd edition, 1999.
- [3] Robert Bringhurst. *The Elements of Typographic Style*. Version 3.2. Hartley & Marks Publishers, Point Roberts, WA, USA, 2008.
- [4] R.C. Compton. Gradient coil apparatus for a magnetic resonance system, 1982.
- [5] Thomas H. Cormen, Charles E. Leiserson, Ronald L. Rivest, and Clifford Stein. *Introduction to Algorithms*. The MIT Press, Cambridge, MA, USA, 2nd edition, 2001.
- [6] Particle Data Group. Statistics. 090001(September 2013):1–36, 2014.
- [7] Donald E. Knuth. Big Omicron and Big Omega and Big Theta. *SIGACT News*, 8(2):18–24, April/June 1976.
- [8] L. Pandola. Search for time modulations in the Gallex/GNO solar neutrino data. *Astroparticle Physics*, 22(2):219–226, nov 2004. ISSN 09276505. doi: 10.1016/j.astropartphys.2004.07.007. URL <http://www.sciencedirect.com/science/article/pii/S092765050400129X>.
- [9] J. D. Scargle. Studies in astronomical time series analysis. II - Statistical aspects of spectral analysis of unevenly spaced data. *The Astrophysical Journal*, 263:835, dec 1982. ISSN 0004-637X. doi: 10.1086/160554. URL <http://adsabs.harvard.edu/abs/1982ApJ...263..835S>.
- [10] Ian Sommerville. *Software Engineering*. Addison-Wesley, Boston, MA, USA, 4th edition, 1992.
- [11] R Turner. Gradient coil design: a review of methods. *Magn Reson Imaging*, 11(7):903–920, 1993. ISSN 0730-725X (Print). doi: 10.1016/0730-725X(93)90209-V. URL <http://www.ncbi.nlm.nih.gov/entrez/>

[query.fcgi?cmd=Retrieve{&}amp;db=PubMed{&}amp;dopt=Citation{&}amp;list{_-}uids=8231676](#).

- [12] Ken Van Tilburg, Nathan Leefer, Lykourgos Bougas, and Dmitry Budker. Search for Ultralight Scalar Dark Matter with Atomic Spectroscopy. *Physical Review Letters*, 115(1), 2015. doi: 10.1103/PhysRevLett.115.011802.