

HEMATITE THROUGH THE EYES OF THE EXOMARS 2020 ROVER ROSALIND FRANKLIN: Simulating mineral identification with the PanCam WAC multispectral filters



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1. INTRODUCTION

We present a pipeline for investigating the ability of PanCam to discriminate one particular mineral species against a defined set of background materials. We demonstrate the pipeline for the mineral hematite, an iron oxide indicative of changes in oxidation conditions, with implications for past habitability, and a target of interest for the *Curiosity* rover at Vera Rubin Ridge^{1,2}.

2. PANCAM

PanCam³ is the mast-mounted colour-stereo panoramic camera system for the ExoMars 2020 Rosalind Franklin rover⁴, with an objective of visual geological characterisation, focusing on signatures of ancient habitats. PanCam will measure the VNIR (380nm - 1100nm) spectral reflectance of surfaces with a multispectral suite of 12 narrowband filters⁵.

3. SPECTRAL PARAMETERS

Spectral parameters (SPs) are operations that measure certain features of a mineral reflectance spectrum. Thus, well chosen SPs provide a low-dimensional method of comparing and distinguishing minerals. A set of recommended SPs for the mineralogy of Mars have been reported for application to CRISM hyperspectral orbital data⁶⁻⁷, a number of which span the VNIR, illustrated in Figure 1. Figure 2 illustrates how an SP operation applied across a hyperspectral image can discriminate minerals.

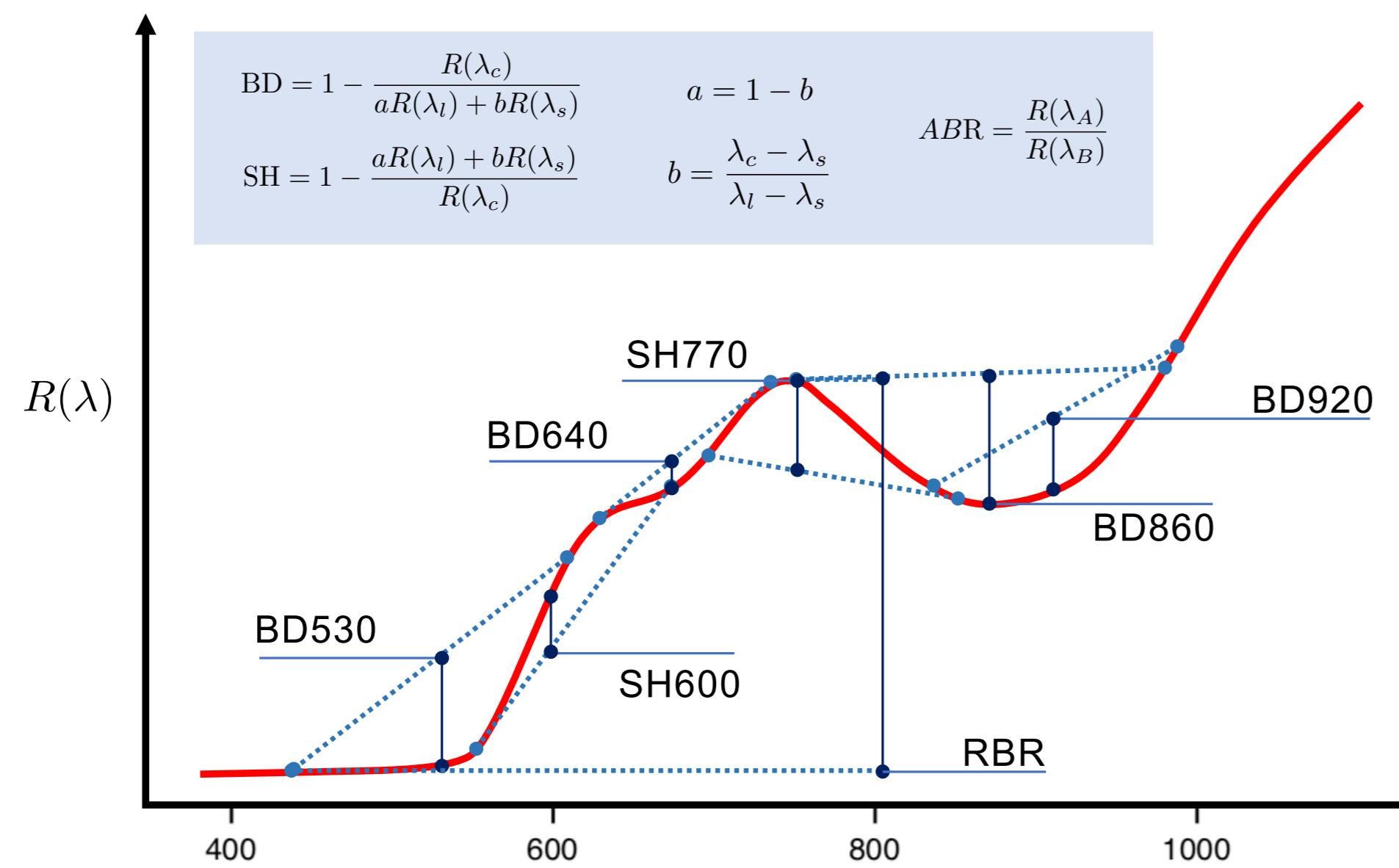


Figure 1. Illustration of the CRISM VNIR spectral parameters on an exemplary hematite spectra (RPN_OOX_04). BD= Band Depth, SH= Shoulder Height, R= Ratio. See 6-7.

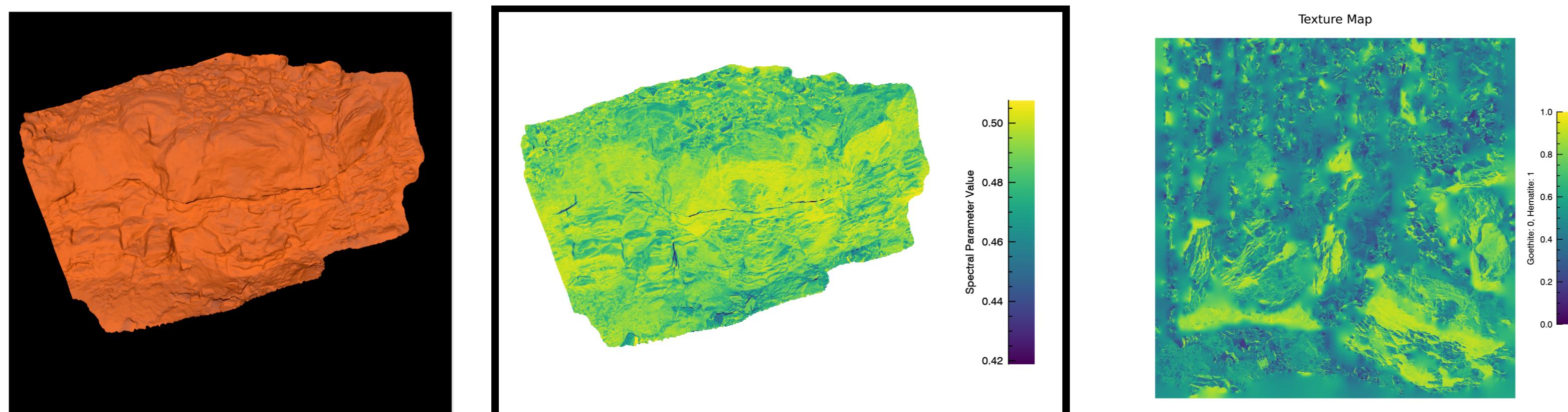


Figure 2. Simulated rock outcrop hosting a mixture of hematite and goethite. Left shows sRGB colour image under uniform illuminant, demonstrating the inability of the colour space to discriminate between the materials. Centre gives the SH600 spectral parameter map of the outcrop, mapped to a viridis colour scale. Right gives the ground-truth texture of mineral mixing. Simulation rendered in an adapted distribution of spectral-PBRT at hyperspectral resolution⁸.

4. PROBLEM STATEMENT

The CRISM SP set was curated for hyperspectral data, but PanCam is multispectral, sampling the morphology at much lower resolution. Here we investigate the robustness of the SPs of Figure 1 at discriminating hematite against a set of expected minerals of the ExoMars landing site at Oxia Planum⁹⁻¹⁰. Laboratory mineral spectra are drawn from a database (see Figure 3) and sampled with Gaussian filters according to the CRISM recommended wavelengths and FWHM, and the nearest corresponding filters of PanCam, as shown in Table 1. Information for the *Curiosity* rover Mastcam multispectral filters are included for comparison.

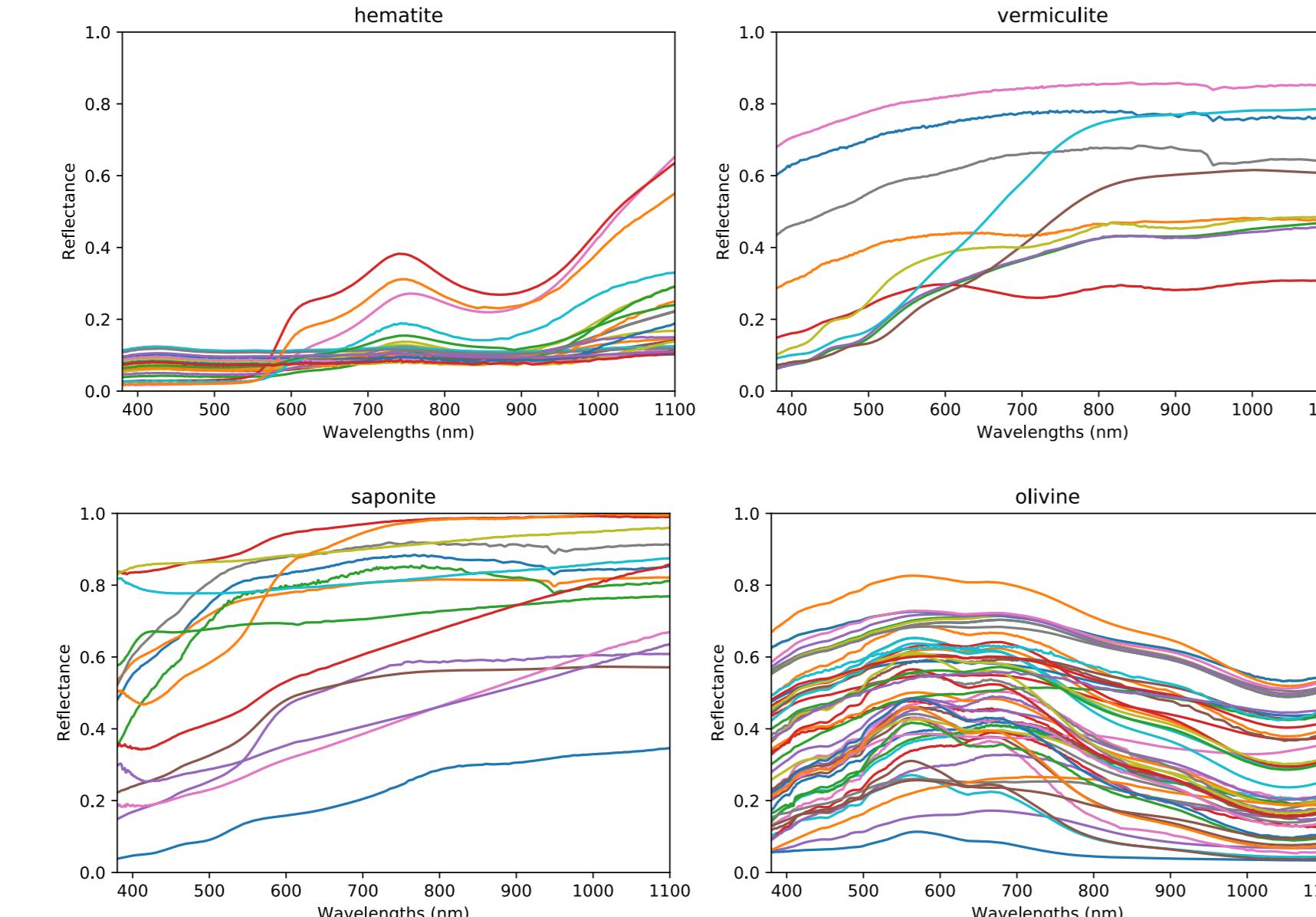


Figure 3. Input mineral reflectance data for target (hematite) and background (vermiculite, saponite and olivine). Data sourced via the Western Washington Spectral Database¹¹, a compilation of several widely used spectral libraries, including USGS speclib06¹², and HOSERLab¹³.

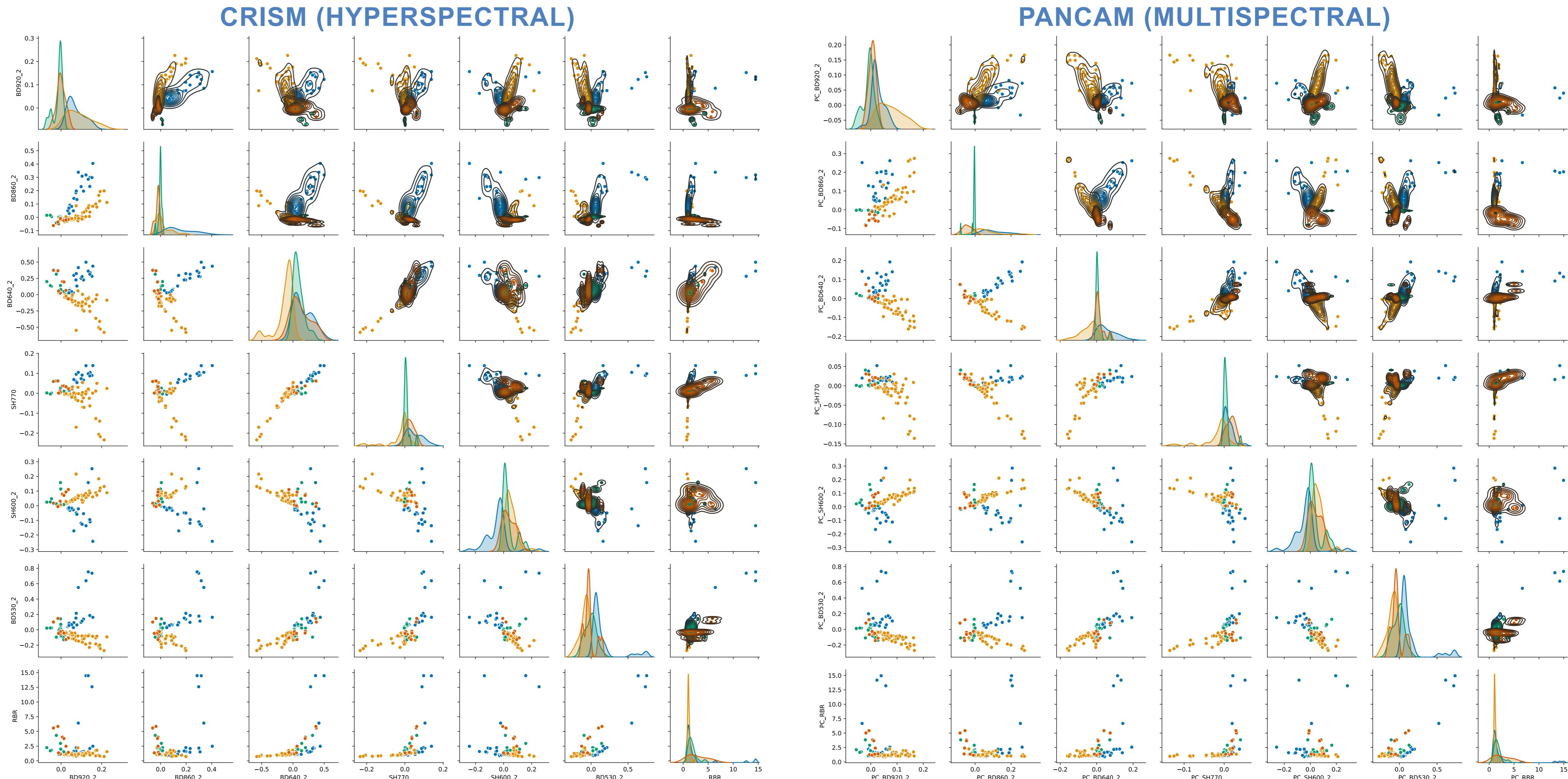


Figure 4. Paired scatterplots of the SPs computed with CRISM and PanCam sampling, indicating data distribution of hematite (blue) against background minerals (orange, green, red) across the SP vector space. Univariate plots (diagonal) indicate single SP discrimination, bivariate indicate advantage of combining SPs. Upper gives density contours.

5. VISUALISATION

A Python software pipeline is in development, utilizing the Pandas library for multivariate data analysis. Figure 4 gives a visual comparison of the computed SPs for CRISM and PanCam sampling, demonstrating the ability of each system to discriminate hematite against the background minerals, via clustering in the SP vector space. Single SPs (diagonal) indicate poor discrimination. Combined SPs, such as BD530/BD860, indicate robust discrimination. Figure 5 visualizes the consistency of SP measurements between CRISM, PanCam and Mastcam, indicated by deviation from linearity.

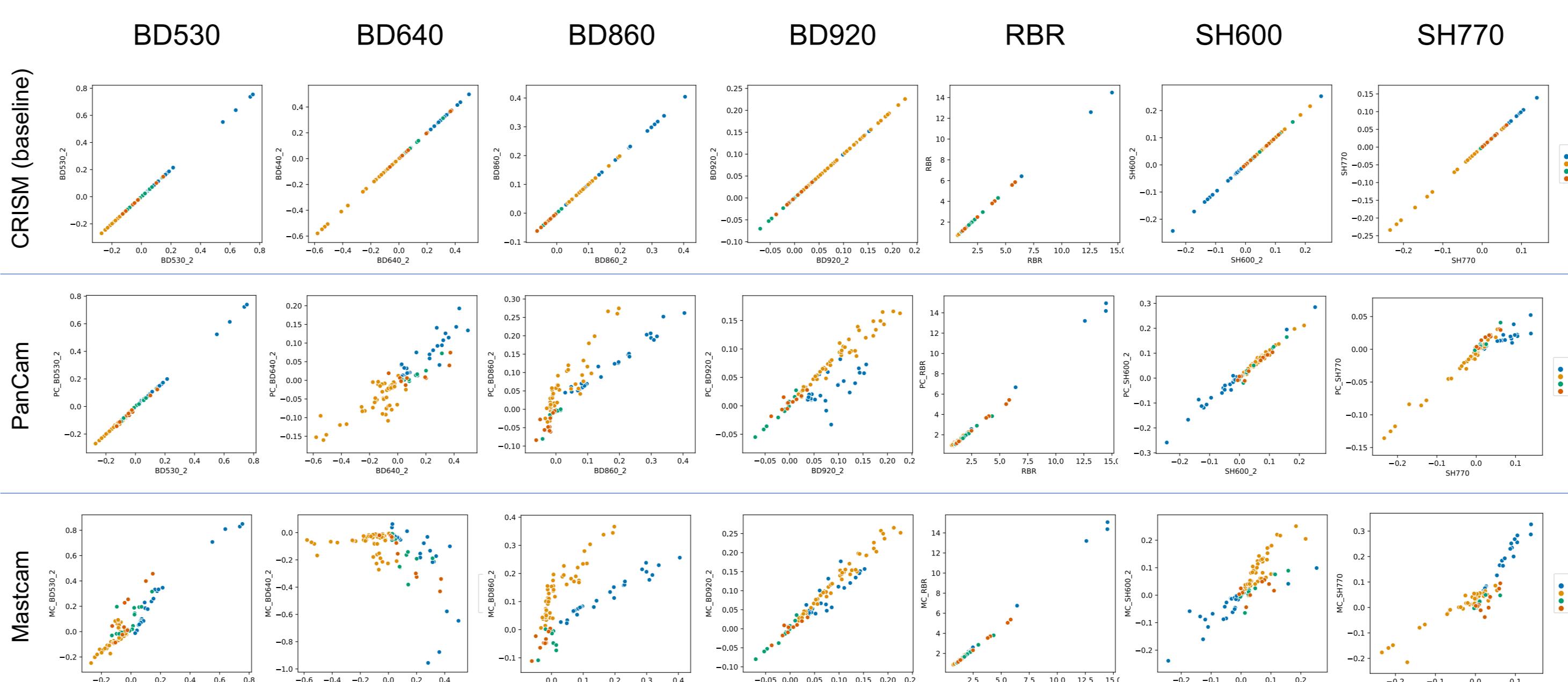


Figure 5. Plotting SPs of each instrument against the CRISM SPs provides a visualization of the consistency of SP evaluation between instruments. CRISM vs CRISM explicitly shows the expected linear relationship, with axes indicating the expected range. Notably, Mastcam poorly captures BD640, and both instruments poorly capture BD860. Despite this, Figure 4 demonstrated the utility of BD860 in hematite discrimination, which is ultimately the more important metric of success.

6. CONCLUSIONS & FUTURE WORK

The method presented in this work, for the specific case of hematite in contrast to minerals expected at Oxia Planum, is extendable to arbitrary targets and background sets: software is in development for scaling the method to large sets (i.e. more comprehensive background mineralogy), and for exploring the wider SP phase-space by varying the filters. Quantitative measures of clustering will enable scoring of SPs and combinations, enabling a recommendation of PanCam filters required to discriminate a target against an expected background.

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REFERENCES

- Fraeman et al., 2013, Geol. 2, Fraeman et al., 2018, LPSC49, 1557, 3, Coates et al., 2017, Astrobiology, 4, Vago et al., 2017, Astrobiology, 5, Cousins et al., 2012, Plan. & Space. Sci., 6, Pelkey et al., 2007, JGR-Planets, 7, Viviano-Beck et al., 2014, JGR-Planets, 8, Stabbins et al., 2018, LPSC49, 2099, 9, Quantin-Nataf et al., 2019, Mars-9, 6317, 10, Mandon et al., 2019, Mars-9, 6173, 11, Rice et al., 2016, <http://spectro.geol.wvu.edu/>, 12, Clark et al., 2007, USGS Digital Spectral Library, 13, Cloutis et al., 2006, LPSC37, 2121