

### Expectations for Exercises

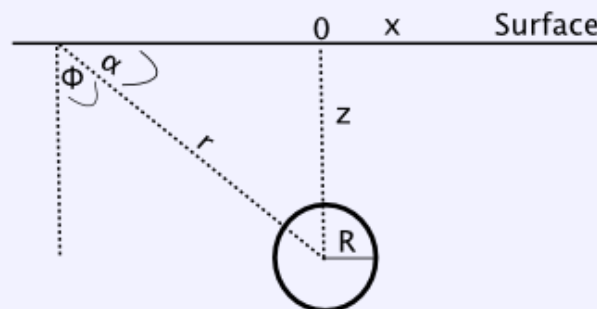
Exercises are an important part of the Geophysics lecture. They will treat some aspects of the lecture in more detail, but also cover new ground. We expect that you work on the exercises at home and we will discuss questions and solutions interactively together. The joint meetings may start with randomly chosen students presenting their approach. It is ok if the solution is not available at this stage and there will be no interrogation. However, please don't show up unprepared, because this will inevitably be awkward.

## 1 Exercises for Gravity Method 1

### 1.1 Detection of a spherical object in the sub-surface

- Consider a spherical object with radius  $R$  located at depth  $z$  below the surface, and a gravimeter that is moved along the surface across the anomaly (Fig. 1). Derive an analytical expression for the expected *vertical* anomaly  $g_z$  as a function of the lateral distance  $x$  and depth  $z$ . Calculate the maximum anomalies that you would expect for some realistic settings (e.g., a limestone cave.)
- Use a piece of paper or a software of your choice (e.g., Excel, SciDAVis, Matlab, Python) and visualize the expected vertical gravity anomaly for a specific setting as a function of lateral distance  $x$ . Show how this profile changes as you vary the depth of the object.
- Sometimes results in the gravity method are expressed in  $mGal$  with  $1 \text{ Gal} = 0.01 \text{ m s}^{-2}$ . Transfer your results to those units and post a picture of your graph in the forum.
- We now investigate how we can estimate the depth of the object from an anomaly as derived in (a,b). For this you have to derive an expression that links the half-width of the anomaly (i.e.  $g_z = \frac{1}{2}g_{z,max}$  for  $x = x_{1/2}$ ) to the depth  $z$  of the object. Derive this relationship which has the form  $z = cx_{1/2}$ . Why is such a relationship be useful?

### Solutions



- (a) Because we have a spherical object, we can use the shell theorem and treat it as a point mass. The gravitational force exhibited by the sphere is:

$$\vec{g} = -G \frac{M}{r^2} \hat{r}$$

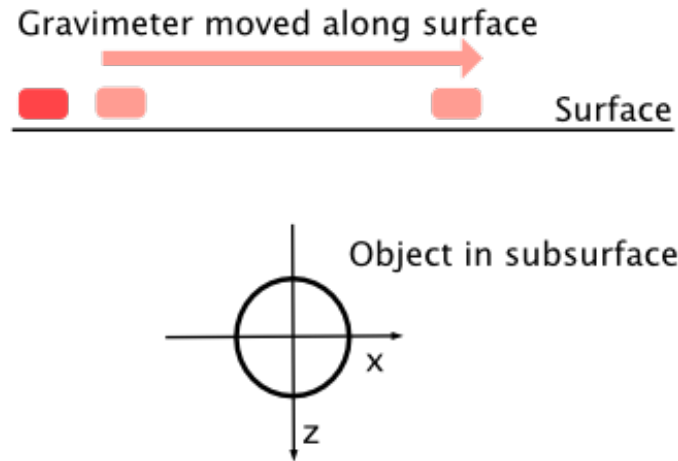


Figure 1: Sketch for problem 1.1. It is easiest to choose a coordinate system with the origin inside the subsurface object.

where  $\hat{r}$  is the unit vector pointing from the center of mass to our gravimeter at the surface. The minus sign makes sure that the force is attractive (and not repulsive). The mass anomaly  $M$  formulated in a relative sense of the sphere is given by:

$$M = \frac{4}{3}\pi R^3 \Delta\rho$$

At distance  $x$  the distance to the sphere is  $r = \sqrt{x^2 + z^2}$ . The gravimeter only measures the vertical component of  $\vec{g}$ :

$$g_z = -|\vec{g}| \cos \phi = -|\vec{g}| \frac{z}{r}$$

which brings us to:

$$\begin{aligned} g_z &= -\frac{4}{3}\pi R^3 \Delta\rho \frac{G}{r^2} \frac{z}{r} \\ &= -\frac{4}{3}\pi R^3 \Delta\rho \frac{G}{r^3} z \\ &= -\frac{4}{3}\pi R^3 \Delta\rho G \frac{z}{(x^2 + z^2)^{\frac{3}{2}}} \end{aligned}$$

We will encounter the maximum value of the anomaly at  $x = 0$  (directly above the target):

$$g_{z,max} = \frac{4}{3}\pi R^3 \Delta\rho G \frac{1}{z^2}$$

which for the limestone cave example is approximately -0.35 mGal (see below).

(b, c)

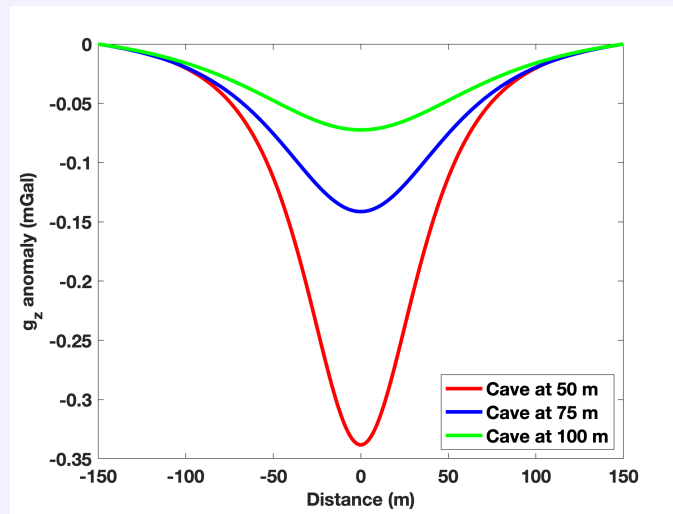
<sup>1</sup> clear all

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2  close all
3
4  %This is vector x for the lateral distance
5  x = -150:0.01:150;
6  %Depth of the sphere
7  z = 50;
8  %Radius of the sphere
9  R = 25;
10 %Constants for the limestone cave scenario
11 % (SI stands for System International units)
12 DensityLimestone_InGramsPerCentimeterCubed = 2.0;
13 DensityLimestone_SI = DensityLimestone_InGramsPerCentimeterCubed *1/1000 * 100^3;
14 DenistyAir_SI = 1.2;
15 %Density Contrast in kg/m^3
16 DeltaRho_SI = DenistyAir_SI-DensityLimestone_SI;
17 %Gravitational constants
18 G_SI = 6.674e-11;
19
20
21 %This is the formula we want to plot
22 gz_SI = 4/3*pi*R^3*DeltaRho_SI*G_SI*z./((z.^2+x.^2).^(3/2));
23 gz_SI_max = 4/3*pi*R^3*DeltaRho_SI*G_SI*1/(z^2);
24 gz_mGal = gz_SI*100*1000;
25
26 fig=figure(1)
27 plot(x,gz_mGal-gz_mGal(1),'r-','LineWidth',3);hold on
28
29 %And for a different depths
30 z=75;
31 gz_SI = 4/3*pi*R^3*DeltaRho_SI*G_SI*z./((z.^2+x.^2).^(3/2));
32 gz_mGal = gz_SI*100*1000;
33 plot(x,gz_mGal-gz_mGal(1),'b-','LineWidth',3);hold on
34
35 z=100;
36 gz_SI = 4/3*pi*R^3*DeltaRho_SI*G_SI*z./((z.^2+x.^2).^(3/2));
37 gz_mGal = gz_SI*100*1000;
38 plot(x,gz_mGal-gz_mGal(1),'g-','LineWidth',3);hold on
39
40 legend('Cave at 50 m', 'Cave at 75 m','Cave at 100 m','Location','SouthEast')
41 xlabel('Distance (m)')
42 ylabel('g_z anomaly (mGal)')
43 %Export to a png. (This can be done much better.)
44 set(findall(fig, '-property', 'FontSize'), 'FontSize', 15, 'fontWeight', 'bold')
45 print('-dpng', '-r300', 'GravityAnomaly.png')

```

../Src/Exercises/Gravimetry/Gravimetry01\_SphereVisualization.m



(d) This requires some arithmetic manipulation to isolate  $z$ .

$$\begin{aligned}
 g_z &= \frac{1}{2} g_{z,max} \\
 \rightarrow \frac{z}{(x_{1/2}^2 + z^2)^{\frac{3}{2}}} &= \frac{1}{2z^2} \\
 \rightarrow \frac{z^3}{(x_{1/2}^2 + z^2)^{\frac{3}{2}}} &= \frac{1}{2} \\
 \rightarrow \frac{1}{(\frac{x_{1/2}^2}{z^2} + 1)^{\frac{3}{2}}} &= \frac{1}{2} \\
 \rightarrow z = \frac{x_{1/2}}{\sqrt{2^{\frac{2}{3}} - 1}} &\approx 1.305 x_{1/2}
 \end{aligned}$$

Those expressions can be useful to estimate the depth of an object from an observed anomaly without using a full forward model. However, this only works if the object is close to spherical (which we usually don't know necessarily ahead of time.) Similar estimates exist for other shapes (e.g., horizontal and vertical cylinders).

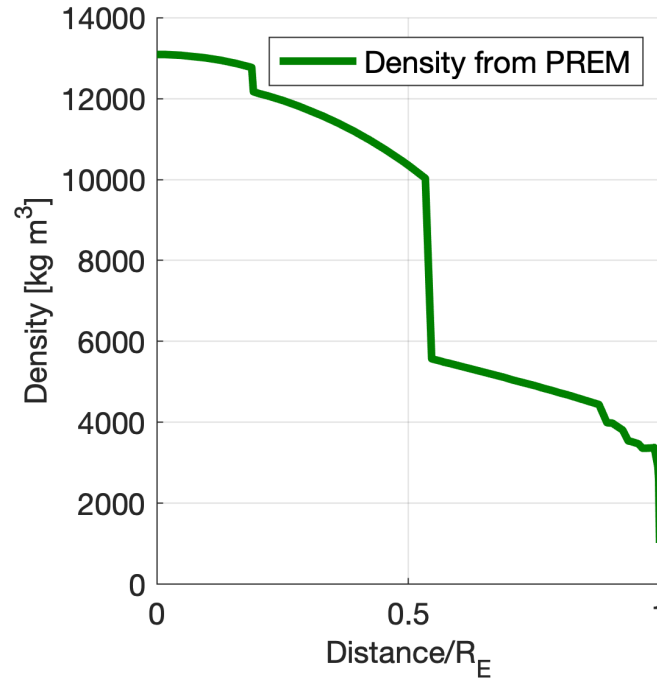
## 1.2 Gravitational acceleration inside the Earth

Knowing the gravitational acceleration *inside* the Earth is important, e.g., for understanding processes related to mantle convection.

- Use the consequences of the shell theorem to predict the gravitational acceleration  $g(r)$  inside a spherical Earth. First assume that the Earth's density is constant, then that it decreases linearly with distance from the center. Assume that the density in the core is approximately  $\rho_{core} = 13 \frac{g}{cm^3}$  and in the crust approximately  $\rho_{crust} = 2.7 \frac{g}{cm^3}$ .
- Draw the results on a x-y graph on paper or using a software of your choice.
- The PREM model provides an observationally constrained estimate of the density distribution inside the Earth. Knowing the shape of this profile, what is your guess of the gravitational acceleration

inside the Earth? Why is it more difficult for you to calculate this quantitatively compared to the previous cases of constant and linearly varying density?

- (d) Extra: Calculate the gravitational acceleration inside the Earth based on the PREM model using Matlab/Python/Excel. The data can be found on Ilias.



### Solutions

(a) At location  $l < R$  inside the Earth, we do not need to worry about the mass distribution for distances  $> l > R$  as these will cancel out due to the spherical symmetry. Also the acceleration will only have a radial component. Hence,

$$|\vec{g}(l)| = G \frac{M}{l^2} \quad (1)$$

where  $M$  is mass contained in the sphere with radius  $l$ . How does  $M$  change as we change  $l$ ? In this spherical symmetry it is best to adopt spherical coordinates so that we can write:

$$M = \int \rho dV \quad (2)$$

$$= \int_0^{\pi/2} d\theta \sin(\theta) \int_0^{2\pi} d\phi \int_0^l dl \rho l^2 dV \quad (3)$$

this is a fairly complicated way of writing something that you know by heart (i.e. the density times the volume of a sphere), but this is a good opportunity to make use of spherical coordinates. Most importantly, it is worthwhile to remember that the volume element  $dV = l^2 \sin(\theta) dr d\theta d\phi$  is quite different from the cartesian coordinates that you are used to otherwise. Also take a moment and understand why the integration limits are the way they are. Solving the integral step by step for

the case of a constant density:

$$M = \rho \int_0^{\frac{\pi}{2}} d\theta \sin(\theta) \int_0^{2\pi} d\phi \int_0^l l^2 dl \quad (4)$$

$$= \rho 2\pi \int_0^{\frac{\pi}{2}} d\theta \sin(\theta) \int_0^l l^2 dl \quad (5)$$

$$= \rho 4\pi \int_0^l l^2 dl \quad (6)$$

$$= \rho \frac{4}{3} \pi l^3 \quad (7)$$

Plugging (7) into (1) then results for  $l < R$ :

$$|\vec{g}(l)| = G \rho \frac{4}{3} \pi \frac{l^3}{l^2} = G \rho \frac{4}{3} \pi l \quad (8)$$

so the gravitational acceleration increases linearly with distance from the Earth's center as long as  $l < R$  and  $\rho$  is constant. A linear decrease of density with increasing distance can be parameterized as:

$$\rho(l) = \frac{\rho_{crust} - \rho_{core}}{R} (l - R) + \rho_{crust} = \frac{\rho_{crust} - \rho_{core}}{R} l + \rho_{core} \quad (9)$$

As the density now changes with  $l$  we cannot take it out of the integration. In particular for eq. (6) this means:

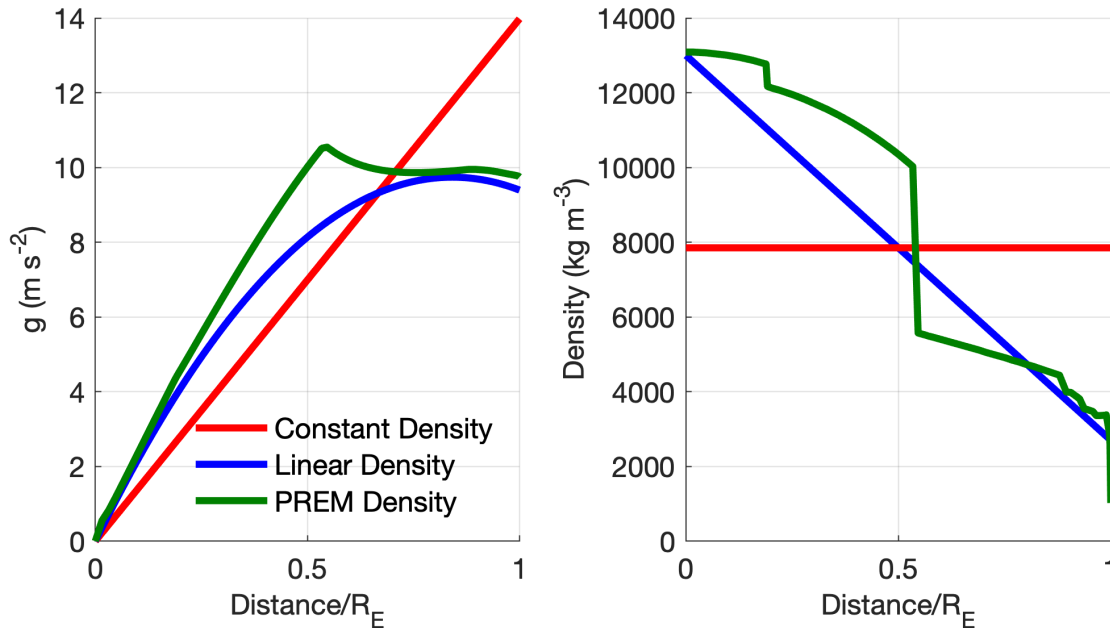
$$M = 4\pi \int_0^l dl \left( \frac{\rho_{crust} - \rho_{core}}{R} l^3 + \rho_{core} l^2 \right) \quad (10)$$

$$= 4\pi \left( \frac{\rho_{crust} - \rho_{core}}{4R} l^4 + \frac{1}{3} \rho_{core} l^3 \right) \quad (11)$$

so that:

$$|\vec{g}(l)| = 4G\pi \frac{1}{l^2} \left( \frac{\rho_{crust} - \rho_{core}}{4R} l^4 + \frac{1}{3} \rho_{core} l^3 \right) \quad (12)$$

$$= 4G\pi \left( \frac{\rho_{crust} - \rho_{core}}{4R} l^2 + \frac{1}{3} \rho_{core} l \right) \quad (13)$$



so that the gravitational acceleration now changes quadratically as shown in the Figure above. Is this realistic? We will have to compare it with a realistic dataset such as the PREM model.

(c) Because we don't have an analytical expression for the PREM density model, we cannot pursue an analytical integration as done in the previous two cases. Hence we do it numerically:

```

1 clear all
2 close all
3
4 %Gravitational constants in SI units
5 G_SI = 6.674e-11;
6 %Density of core and crust in kg m-3
7 rho_crust = 2.7*1003/1000;
8 rho_core = 13.0 * 1003/1000;
9 %Mean density in this model (ouch, too high)
10 rho_mean = (rho_crust+rho_core)/2;
11 %Radius of a spherical Earth in m
12 R = 6371*1000;l=0:1:R;
13
14
15 %% Made up linear density
16 LinearDensity = (rho_crust-rho_core)/R*(1-R)+rho_crust;
17 %% PREM Density
18 PREM = dlmread('PREM.txt',' ',4,0);
19 PREM_Depth = PREM(:,2)*1000;PREM_Density = PREM(:,5)*1003/1000;PREM_dz = (PREM(2,2)
    -PREM(3,2))*1000;
20 [PREM_Depth,ia,ic] = unique(PREM_Depth);PREM_Density=PREM_Density(ia);PREM_Depth(1)
    =1e-3;out = [PREM_Depth PREM_Density];
21 save('PREM_DepthDensity.txt','out','-ASCII');
22 %% Get the corresponding gs
23 gConstantDensity = G_SI*rho_mean*4.0/3.0*pi*l;
24 gLinearDensity = 4*pi*G_SI*((rho_crust-rho_core)/(4*R)*l.^2+1.0/3.0*rho_core*l);
25

```

```

26 %% Numerical integration using the trapezoidal rule
27 for kk=2:length(PREM_Depth)
28     gPREM(kk) = 4*pi*G_SI.*1./ (PREM_Depth(kk).^2) .*trapz(PREM_Depth(1:kk),
29     PREM_Density(1:kk) .*PREM_Depth(1:kk).^2);
29 end
30
31 fig=figure(1)
32 subplot(1,2,1)
33 plot(1/R,gConstantDensity,'r-','LineWidth',3);hold on
34 plot(1/R,gLinearDensity,'b-','LineWidth',3);hold on
35 plot(PREM_Depth/max(PREM_Depth),gPREM,'-','LineWidth',3,'color',[0 0.5 0])
36 ylabel('g (m s^{-2})')
37 xlabel('Distance/R.E')
38 legend('Constant Density','Linear Density','PREM Density','Location','SouthEast')
39 box off;grid on;
40 legend('box','off')
41 subplot(1,2,2)
42 plot(1/R,LinearDensity,'b-','LineWidth',3);hold on
43 plot(1/R,LinearDensity*0+rho_mean,'r-','LineWidth',3);
44 plot(PREM_Depth/max(PREM_Depth),PREM_Density,'-','LineWidth',3,'color',[0 0.5 0])
45 ylabel('Density (kg m^{-3})')
46 xlabel('Distance/R.E')
47 box off;grid on;
48
49 %Export to a png. (This can be done much better.)
50 set(gcf,'PaperUnits','centimeters','PaperPosition',[0 0 20 10])
51 set(findall(fig,'-property','FontSize'),'FontSize',12)
52 print('-dpng','-r300','../..../Exercises/Figures/Gravimetry/
53     Gravimetry01-GravityInsideEarth.png')
54
55 fig=figure(2)
56 plot(PREM_Depth/max(PREM_Depth),PREM_Density,'-','LineWidth',3,'color',[0 0.5 0])
57 box off;
58 grid on;
59 legend('Density from PREM');ylabel('Density [kg m^3]');xlabel('Distance/R.E')
60 set(gcf,'PaperUnits','centimeters','PaperPosition',[0 0 10 10])
61 set(findall(fig,'-property','FontSize'),'FontSize',12)
62 print('-dpng','-r300','../..../Exercises/Figures/Gravimetry/Gravimetry01-PREM.png')

```

../Src/Exercises/Gravimetry/Gravimetry03-GravityInsideSphere.m

### 1.3 Determining the mean density of the Earth with your own gravimeter

#### Group Work

Self-organize in groups with 5-6 team members. Choose a group name and a team captain that communicates with the instructors. It is ideal if this group stays together throughout the term for the applied exercises. Make sure that you are inclusive during the group formation.

The vertical gravitational acceleration can be determined by measuring the traveltime of a freely falling object for a known distance. Assuming that the radius of the Earth and the gravitational constant are



known ( $G \approx 6.674 \cdot 10^{-11} \text{ m}^3\text{kg}^{-1}\text{s}^{-2}$ ,  $R_E \approx 6370$ ) the mass of the Earth can be determined with the basic principles derived in class.

- (a) Your task is to design a home-built free-fall gravimeter that does the job. You will quickly realize that the time measurement (i.e. the time from the start to the end of the fall) is one critical aspect in the system design. A helper tool that we suggest is the *acoustic stopwatch* that can be accessed via a smartphone and the *phyphox* app. Figure out an efficient way how the traveltime of a freely falling object can be determined with this type of acoustic trigger. (Tip: You may need a hammer, a metal bar and a bar clamp for your acoustic trigger. Be creative.)
- (b) Collect a dataset of traveltimes and derive the mass of the earth. Provide an error estimate. Can you identify a measurement bias? How does your result compare to literature values?
- (c) Given your result, what is the mean density of the Earth? How does that compare to rock densities found at the Earth's surface? What is a main conclusion that you can draw from this?
- (d) Post your result, your dataset and a picture of your system setup in the ILIAS forum.
- (e) Extra: Are you a tinkerer/Bastler? If so, feel free to improve the system design, e.g., by using light barriers in combination with a raspberry pi nano. We are more than happy to buy material for you, the only constraint is that it works reliably and that it remains cheapish (let's say <100 Euro). **If you succeed, you will win a gift certificate of that can be used in a Tübingen pub of your choice to celebrate your victory with your peers.** Moreover, your system will then be used for eternity for the following classes giving you much honor within the geo- and environmental student community.

### Solutions

I made a video with a possible setup here:

[Link to Video](https://www.youtube.com/watch?v=q5SdryMp8iw&feature=youtu.be)

<https://www.youtube.com/watch?v=q5SdryMp8iw&feature=youtu.be>