# PHY 905 Project 4: Molecular Dynamics

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Molecular dynamics simulations provide a means of studying many-body systems through, in this case, numerically solving Newton's equations of motion. We ran a simulation of 500 argon atoms interacting via a Lennard-Jones two-body potential and sampled the system's energy, temperature and diffusion constant. By comparing the behavior of the diffusion constant over time for various equilibrium temperatures, we found the melting point for this system to be roughly 287.5 K.

### INTRODUCTION

Molecular dynamics simulations attempt to capture the dynamical evolution of a many body system. These simulations are applied in a wide range of disciplines from studying the internal dynamics of polymer chains [1] to understanding the structure and function of biological macromolecules [3]. The rise of high-performance computing has benefited molecular dynamics simulations through allowing for enhanced methods of calculating the interatomic forces [4]. The increasing scope and complexity of molecular dynamics simulations can provide further insights into the many problems to which this method can be applied.

In this project, we simulate a simple system of argon atoms interacting via a pairwise Lennard-Jones potential [5, 7]. By calculating the diffusion constant, we can estimate the melting point for this system. In the next section, we discuss the statistical thermodynamics background of our project. Thereafter, we describe the simulation code we used and the results we obtained. Finally, we will briefly discuss avenues for further development of the tools developed for this project.

### THEORY

Statistical thermodynamics provides the theoretical basis for the type of classical molecular dynamics simulation we employed in this project. The code we developed solves Newton's equations of motion numerically to evolve the system along a deterministic trajectory in its phase space. We can then compute various macroscopic quantites from the system's microscopic properties.

## Lennard-Jones Potential

We model the interaction between argon atoms with a Lennard-Jones potential - plotted in FIG. 1. This potential is often used to model interactions between noble gas atoms. It is parameterized by the strength  $\epsilon$  and range of the interaction  $\sigma$ . Values of  $\epsilon/k_B=119.8$  K and  $\sigma=5.26$  Å are optimal values for argon.

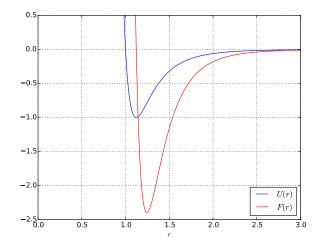


FIG. 1: The Lennard-Jones potential (blue) and derived force (red).

$$U(r_{ij}) = 4\epsilon \left[ \left( \frac{\sigma}{r_{ij}} \right)^{12} - \left( \frac{\sigma}{r_{ij}} \right)^{6} \right], \tag{1}$$

# Thermodynamic quantities

Since our system contains a fixed number of particles, it is contained in a fixed volume and the total energy remains constant, we identify it as a microcanonical ensemble. We can calculate the average energy of the system using the equipartition theorem. Since we are dealing with monoatomic argon, there are only three translational degrees of freedom, so the average energy is given by

$$\langle E_k \rangle = \frac{3}{2} N k_B T$$

Re-writing this relation allows us to compute an estimate of the system's temperature

$$T = \frac{2}{3} \frac{E_k}{N_{\text{atoms}} k_B}.$$
 (2)

The diffusion constant is another observable that we can extract from our simulation. The Einstein relation between the mean square displacement and the diffusion constant (D) allows us to calculate D from the atoms' average displacements at each time step:

$$D = \frac{\langle r^2(t) \rangle}{6t} \tag{3}$$

where

$$\langle r^2(t) \rangle = \frac{1}{N} \sum_{i}^{N} r_i^2(t)$$
$$= \frac{1}{N} \sum_{i}^{N} |\mathbf{r}_i(t) - \mathbf{r}_i(0)|^2$$

The average of all the atomic displacements gives a measure of how mobile the atoms are. For example, in a solid the atoms won't move far from their lattice sites so  $\langle r^2(t) \rangle$  will be small. This value will be larger for liquids and larger still for gases. Therefore, we can interpret the diffusion constant as a comparative measure of the state of the system.

#### Units

The small scale of the system we are simulating lends itself to working in a specialized set of units. We used the following four relations to define our system of units

1 unit of mass = 1 a.m.u = 
$$1.661 \times 10^{-27}$$
kg,  
1 unit of length =  $1.0 \text{Å} = 1.0 \times 10^{-10}$ m,  
1 unit of energy =  $1.651 \times 10^{-21}$ J,  
1 unit of temperature =  $119.735$ K.

Converting to this system of units, Boltzmann's constant becomes  $k=(1.38\times 10^{-23})(119.7)(1.65\times 10^{-21})=1.$ 

### ALGORITHMS AND METHODS

The C++ code developed for this project utilizes the velocity Verlet method to solve Newton's equations of motion. In this section we briefly summarize the derivation of this method discussed in Hjorth-Jensen [2]. We then provide a description of the codes which can be found at https://github.com/redpath11/phy905\_thr in the projects/project4/src directory.

#### Verlet Method

We numerically solve the equations of motion using the velocity Verlet method. Since we've assumed a Lennard-Jones type pairwise interaction between atoms, the form of the acceleration is known from the gradient of the interaction potential. The Verlet method derives from a summation of two Taylor expansions. Consider the position in one dimension

$$x(t+h) = x(t) + hx'(t) + \frac{1}{2}h^2x''(t) + O(h^3)$$
  
$$x(t-h) = x(t) - hx'(t) + \frac{1}{2}h^2x''(t) + O(h^3).$$

Summing these two expansions and switching to the discretized notation gives

$$x_{i+1} = 2x_i - x_{i-1} + h^2 x_i'' + O(h^4).$$

In general, this algorithm is not self-starting since it requires the value of x at two previous points.

Now, consider a Taylor expansion of the velocity and note the relationshipe between velocity and acceleration v'=a.

$$v_{i+t} = v_i + hv_i' + \frac{1}{2}v_i'' + O(h^3)$$
  

$$a_{i+1} = v_i' + hv_i'' + O(h^2).$$

We will assume that the velocity is a linear function of time  $hv_i'' \approx v_{i+1}' - v_i'$ . We can now write the final algorithmic form for the position and velocities.

$$x_{i+1} = x_i + hv_i + \frac{h^2}{2}v_i' + O(h^3)$$
(4)

$$v_{i+1} = v_i + \frac{h}{2}(v'_{i+1} + v'_i) + O(h^3)$$
 (5)

### Code

## Overview

We supplied were code skeleton (from https://github.com/andeplane/ molecular-dynamics-fys3150) outlining the major classes needed to run a molecular dynamics simulation. From this starting point, we filled in details to apply periodic boundary conditions, initialize the system of argon atoms in face-centered cubic (FCC) lattice, calculate the pairwise forces from the Lennard-Jones potential and compute physical properites of the system. Figure 2 gives a general overview of the code structure by identifying key elements and where they reside in the within the class structure.

### Initialization

Initialization of our system takes place in two general steps: (1) setting the size, initial temperature and

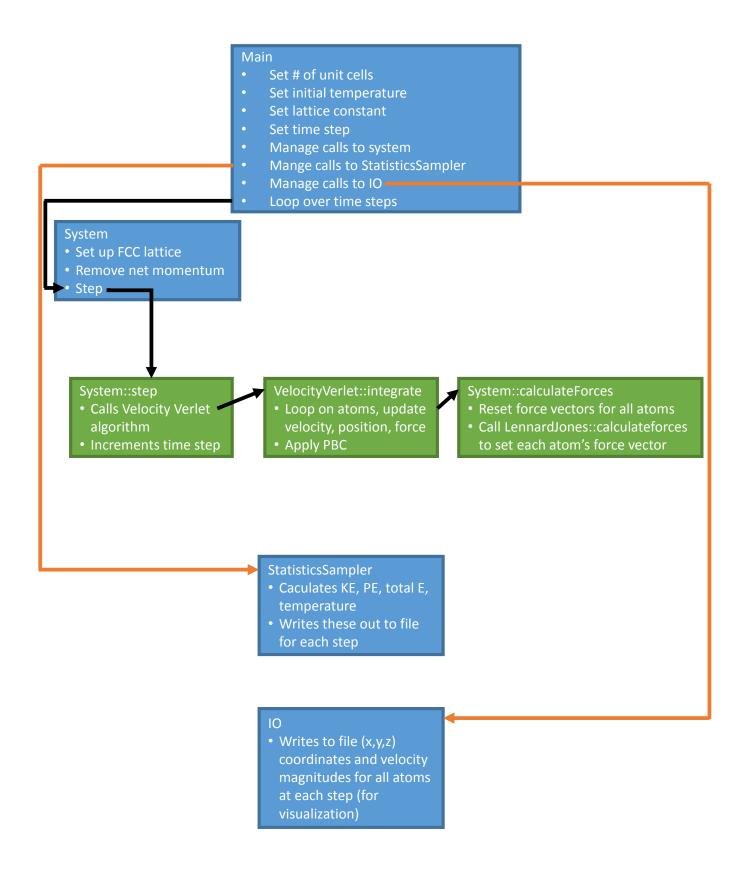


FIG. 2: The flowchart gives a crude overview of how the code operates. Blue boxes represent classes and indicate processes carried out by one or more member functions. Green boxes represent functions crucial to the simulation. Black arrows indicate the hierarchy of function calls for one timestep. Orange arrows depict function calls that result in writing data to a file.

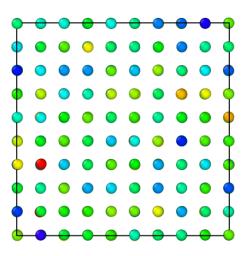
Lennard-Jones potential characteristics and (2) setting the initial positions and velocities for the atoms. The first step is implemented in the main function. The size of the system is determined by the number of unit cells in the FCC lattice and the lattice constant (the size of one unit cell). We studied a system of argon atoms for which the lattice constant is 5.26 Å. The  $\epsilon$ ,  $\sigma$  parameters specify the L-J potential, for argon these are  $\epsilon/k_B=119.8\mathrm{K}$ ,  $\sigma=3.405 Å$ .

The second initialization step is carried out in the System class. This class contains a function System::createFCCLattice to position  $4 \times N_x \times N_y \times N_z$  atoms in a cubic lattice and give them random velocities drawn from a Maxwell-Boltzmann distribution.

$$P(v_i) = \sqrt{\frac{m}{2\pi kT}} \exp\left[\frac{-mv_i^2}{2kT}\right]$$

Doing so gives the system some initial kinetic energy (i.e. temperature). It should be noted that this does not generate a system in equilibrium, but rather it populates some random microstate with a random initial kinetic energy. The System::removeTotalVelocity function subtracts the net velocity from each atom so that the center of mass of the system is stationary.

```
double ssize = ((double)
      numberOfUnitCellsEachDimension) *
      latticeConstant;
  setSystemSize(vec3(ssize, ssize, ssize));
  const double b = latticeConstant / 2.;
  const double x[4] = \{0., b, 0., b\};
  const double y[4] = \{0.,b,b,0.\};
const double z[4] = \{0.,0.,b,b,\};
  for (int i=0; i < number Of Unit Cells Each Dimension; i
  { } \{// \text{ place unit cells along x }
    double bx = ((double) i)*b*2. + offset;
    for (int j=0; j<numberOfUnitCellsEachDimension; j
     {// place unit cells along y
       double by = ((double) j)*b*2. + offset;
       for (int k=0;k<numberOfUnitCellsEachDimension
       : k++)
       {// place unit cells along z
         double bz = ((double) k)*b*2. + offset;
         for (int a=0; a<4; a++)
         {// place 4 atoms
            Atom *atom;
            atom = new Atom(UnitConverter::
      massFromSI(6.63352088e-26));
            atom\rightarrowposition.set(x[a]+bx,y[a]+by,z[a
21
            atom->initialPosition.set(x[a]+bx,y[a]+
      by, z[a]+bz);
            atom->resetVelocityMaxwellian(
      temperature);
            m_atoms.push_back(atom);
         }// END place 4 atoms
       }// END place unit cells along z
    }// END place unit cells along y
  }// END place unit cells along x
```



(a) Top view

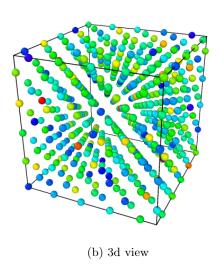


FIG. 3: Positions and velocity magnitudes (indicated by the color) visualized in Ovito for the system initial state.

An initialized system consisting of 500 atoms (5 unit cells) is shown in FIG. 3. We used Ovito [6] to visualize the system. The density of our system is then given by  $(Nm)/(N_xN_yN_zb^3)$  where N is the number of atoms, m is the mass of one atom,  $N_i$  is the number of unit cells in each dimension.

### Simulation

A general overview of the code structure is depicted in FIG. 2. We used the velocity Verlet algorithm to update the positions and velocities of the atoms in our simulation at every time step. Furthermore, we implemented periodic boundary conditions to confine the atoms in our

system to a finite volume. Each time the positions are updated, we run the following routine to ensure that all atoms are contained inside the specified volume. We also need to adjust the initial position vector accordingly in order to properly compute the diffusion constant. Finally, to improve the efficiency of the code, we exploit the  $1/r^6$  behavior of the potential and set the contribution to the net force to zero when atoms are more than 2.5 Å apart. To compensate for this, we subtract the constant  $V_{LJ}(r=2.5)$  from the total potential.

```
for (Atom *atom : m_atoms)
{
    for (int j=0;j<3;j++)
    {
        if (atom->position[j] += m_systemSize[j];
            atom->initialPosition[j] += m_systemSize[j];
        }
        if (atom->position[j]>=m_systemSize[j]) {
            atom->position[j]-= m_systemSize[j];
            atom->initialPosition[j] -= m_systemSize[j
```

Once the positions are updated at each time step, we sample several macroscopic properites of the system. This is handled by the StatisticsSampler class. These functions loop over all atoms to calculate the total kinetic energy, the temperature and the diffusion constant for the system. This class also retrieves the potential energy that is calculated when the forces are updated. These quantities are written out to a file. Finally, the positions and velocity magnitudes at each time step for all atoms are also written to a .xyz file that can be read into Ovito to produce a movie of the simulation.

### RESULTS AND DISCUSSION

## **Energy Conservation**

Our first test of the simulation was to ensure that energy conservation is obeyed. This test revealed a major bug in the way we were applying the periodic boundary conditions. When this is not done correctly, the potential energy and forces are incorrectly calculated because the distances between atoms is wrong. Once this issue was resolved, we obtained the results shown in Table I for a simulation of 1000 time steps and an initial temperature of 300 K.

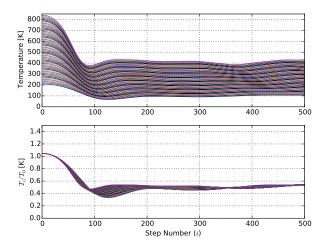


FIG. 4: The top panel plots calculated system temperature (eq.2) vs. time step number for starting temperatures  $T_0$  ranging from 200 to 800 K (the different lines). The bottom panel presents the same data as the top panel but with the temperature at each step normalized by the  $T_0$  for that simulation. We varied the initial temperature between 200 K and 800 K and used a time step value  $dt = 10^{-15}$  s. For all the simulations, the equilibrium temperature is roughly half the initial temperature.

### Equillibrium Temperature

In a preceding section about initializing the simulation, we mentioned that the "initial" temperature we set is not the actual temperature of the system but rather a way to set the total energy of the system. When we randomly assign velocities to each atom, we are starting the system in a random one of its virtually uncountable microstates. Through simulating the interactions between atoms, we evolve the system through other microstates at each time step where, according to the ergodic hypothesis, the probability of reaching a given microstate is given by the probability distribution governing the ensemble. Now, the initial state that we populate when we construct the system isn't likely to be an equilibrium state for the system, but as we run the simulation, the system evolves to reach a region of the phase space corresponding to equilibrium microstates. This process is represented by the peak at 0 and subsequent fall off in FIG. 4. Once the system reaches equilibrium (around step 175), the temperature fluctuates randomly around its equilibrium value  $(T_{eq})$ . We found that the equilibrium temperature is roughly half the initial temperature.

TABLE I: Results from a simulation with the initial temperature $T_0 = 300$ K and a time step size $dt = 10^{-14}$ s.
Values in the time column are in MD units $(1.00224 \times 10^{-13} \text{ s})$ .

Timestep	Time	Temperature [K]	KineticEnergy [eV]	PotentialEnergy [eV]	TotalEnergy [eV]
1	0.099777	312.026	20.1662	-38.3532	-18.187
101	10.0774	158.627	10.252	-28.4274	-18.1754
201	20.0551	163.31	10.5547	-28.7268	-18.1721
301	30.0327	171.405	11.0779	-29.252	-18.1741
401	40.0104	160.382	10.3655	-28.5349	-18.1694
501	49.988	169.434	10.9505	-29.1232	-18.1726
601	59.9657	159.62	10.3162	-28.4912	-18.175
701	69.9434	164.443	10.6279	-28.8028	-18.1748
801	79.921	167.812	10.8457	-29.0173	-18.1716
901	89.8987	172.008	11.1168	-29.2901	-18.1732

#### **Diffusion Constant**

Next, we investigated the melting point of the system by comparing the diffusion constant from simulations with different starting temperatures  $T_0$  (or equivalently  $T_{eq} \approx T_0/2$ ). In FIG. 5, we show the sampled temperature and diffusion constant vs. step number from simulations with  $T_i$  ranging from 100 K to 900 K. These simulations used a step size  $dt = 10^{-14}$  s and ran for 1000 steps. We looked at the diffusion coefficent as a function of time to gain insight into the state of the system. Recall that a system of atoms in a solid state will be constrained to the vicinity of their lattice sites and therefore the system will have a smaller average displacement when compared to a system in a liquid state. In this case, the atoms move more freely so the system will exhibit a larger average displacement. Since the diffusion constant (D) is proportional to the average displacement (eq. 3), it provides an indication of the state of the system.

In the bottom panel of FIG. 5, we see that D, towards the end of the simulation, separates the simulations into two groups. The lower band, indicates  $T_0$  values for which D remains constant over the course of the simulation. In the upper group, D gradually rises over the course of the simulation. For the largest  $T_0$  simulations, D reaches a steady state before the end of the simulation. We interpret this increase in D as a phase transition where the system changes from a solid to a liquid (it melts). Furthermore, the slope indicates how quickly the process happens - for systems initialized with a large amount of kinetic energy (higher  $T_0$ ), the transition occurs faster than for systems initialized with a lower  $T_0$ . Closer inspection of FIG. 5 reveals a small group of systems with a higher final D than the non melting cases. These systems are likely very close to the melting point so the phase transition takes a long time or they only partially melt.

The information in FIG. 5 can be represented more succinctly by noting that the temperature (T) and D are relatively constant over the last 200 time steps of all the simulations. Therefore, we plot D vs. T both

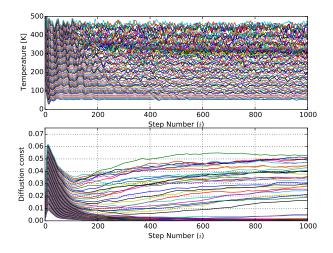


FIG. 5: In the top panel, we show the temperature vs. time step number. In the bottom panel, we plot, as a function of step number, the diffusion constant in units of  ${\rm \AA}^2/s_{\rm MD}$  where  $s_{\rm MD}$  is the unit of time used in the simulations. The different lines correspond to simulations with different initial temperatures ranging from 100 K to 900 K in steps of 10 K. Each simulation used 1000 steps and a step size of  $10^{-14}~{\rm s}$ .

averaged over the last 200 time steps in FIG. 6. The top panel of FIG. 6 shows that for all the simulations, D remains roughly constant over the last 200 time steps. The bottom panel shows a discontinuity in the diffusion constant as a function of temperature suggesting that the melting point occurs just above 300 K.

To ensure that the simulation time wasn't too short to capture the melting process for systems initialized close to the melting point, we ran two batches of simulations with 10000 time steps each. We simulated systems with  $T_eq$  ranging from 280 K to 292.5 K and 195 K to 205 K. The D vs. time step index from these simulations are shown in FIG. 7 and FIG. 8, respectively and show the melting point to lie between 285 and 287.5 K.

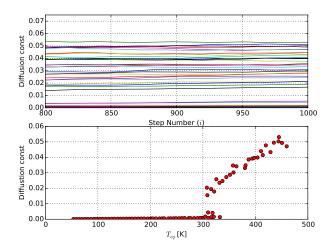


FIG. 6: In the top panel, we plot the diffusion constant in units of  $\mathring{A}^2/s_{\rm MD}$  vs. the time step index. The bottom panel shows the diffusion constant vs. the temperature where both observables have been averaged over the last 200 time steps.

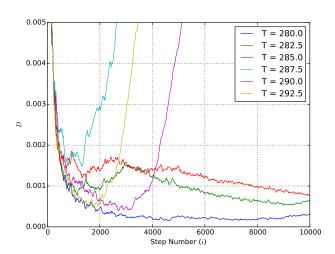


FIG. 7: Diffusion constant vs. time step for systems with 5 different  $T_e p$  raging from 280 K to 292.5 K in steps of 2.5 K. Each simulation was run with 10000 time steps.

### CONCLUSIONS

We have developed a simple molecular dynamics simulation for argon atoms interacting via a simple two-body force. We extracted the system energy, temperature and diffusion constant and used these to determine the melting point as roughly 287.5 K. In order to compare this result to the actual argon melting point, we need to calculate the pressure of our simulated system. This is a

logical next step that will involve additions to the sim-

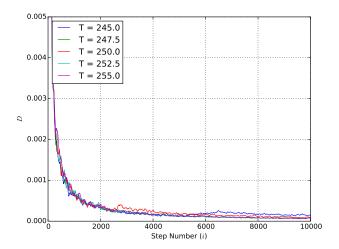


FIG. 8: Diffusion constant vs. time step for systems with 5 different  $T_e p$  raging from 245 K to 255 K in steps of 2.5 K. Each simulation was run with 10000 time steps.

ulation package. Another avenue for further study, is to vary the size of the system and explore how this impacts the random fluctuations of the macroscopic observables.

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