

RS-WISP-01: Geo-Stationary Orbit Generation

WorkerInSpace Senior Researcher

Hongseok Kim

1/14/2025

I. Scope

Review of the two-Body Model and generate LEO parking orbit, Geo-Transfer Orbit and GEO Stational Orbit

Starting from two-body model

$$\ddot{\vec{r}} = -\mu \frac{\vec{r}}{|\vec{r}|^3}$$

II. Satellite Initial Orbit Parameter and Simulation Parameter Setting

```
clear; clc;

Tsim = 86400;
delt = 1;
N = floor(Tsim/delt);
tVec = [0:N-1]' * delt;

%% Define Initial Time
initial_epoch = posixtime(datetime('2024-03-21 00:00:00'));

elapsed_days = 0;
%% Define Current time
current_epoch = initial_epoch + elapsed_days * 24 * 3600;
epoch_datetime = datetime(current_epoch +
tVec, 'ConvertFrom', 'posixtime', 'Format', 'MM/dd/yyyy HH:mm:ss');
epoch = [epoch_datetime.Year, epoch_datetime.Month, epoch_datetime.Day,
epoch_datetime.Hour, epoch_datetime.Minute, epoch_datetime.Second];
disp(epoch_datetime(1))
```

03/21/2024 00:00:00

```
% Orbit Constant
```

```
J2 = 0.00108248;
```

```
R_Earth = 6378.1363;
```

```
mu = 398600.4415;
```

```
%% Speed at Apoapsis
```

```
vel_aps = zeros(5,1);
```

```
%% Define Initial Orbit Parameter
```

```
SMA = 24434.864;
```

```
ecc = 0.728693;
```

```
inc = 0;
```

```
RAAN = 0;
```

```
AOP = 0;
```

```
theta = 180;
```

```
orbit_period = 2*pi* sqrt(SMA^3/mu)
```

```
orbit_period = 3.8012e+04
```

```
time_vector = [0:1:orbit_period, orbit_period];
```

```
figure;
```

```
hold on
```

```
% Define Orbit Propagator
```

```
[r0,v0]=kepler2ijk_hs(SMA,ecc,inc,RAAN,AOP,theta);
```

```
[t,y] = orbit_propagation(r0,v0,time_vector);
```

```
r = y(:,1:3);
```

```
v = y(:,4:6);
```

```
plot3(r(:,1),r(:,2),r(:,3),'LineWidth',2)
```

```
vel_aps(1) = norm(v(1,:));
```

```
% r(end,:)
```

```
% v(end,:)
```

```
const = SMA * (1 + ecc);
```

```
% ecc = 0.6
```

```
ecc = 0.6;
```

```
SMA = const/(1+ecc);
```

```
orbit_period = 2*pi* sqrt(SMA^3/mu);
```

```
time_vector = [0:1:orbit_period, orbit_period];
```

```
[r0,v0]=kepler2ijk_hs(SMA,ecc,inc,RAAN,AOP,theta);
```

```

[t,y] = orbit_propagation(r0,v0,time_vector);
r = y(:,1:3);
v = y(:,4:6);
vel_aps(2) = norm(v(1,:));
plot3(r(:,1),r(:,2),r(:,3),'LineWidth',2)

% ecc = 0.4
ecc = 0.4;
SMA = const/(1+ecc);
orbit_period = 2*pi* sqrt(SMA^3/mu);
time_vector = [0:1:orbit_period, orbit_period];
[r0,v0]=kepler2ijk_hs(SMA,ecc,inc,RAAN,AOP,theta);
[t,y] = orbit_propagation(r0,v0,time_vector);
r = y(:,1:3);
v = y(:,4:6);
vel_aps(3) = norm(v(1,:));
plot3(r(:,1),r(:,2),r(:,3),'LineWidth',2)

% ecc = 0.2
ecc = 0.2;
SMA = const/(1+ecc);
orbit_period = 2*pi* sqrt(SMA^3/mu);
time_vector = [0:1:orbit_period, orbit_period];
[r0,v0]=kepler2ijk_hs(SMA,ecc,inc,RAAN,AOP,theta);
[t,y] = orbit_propagation(r0,v0,time_vector);
r = y(:,1:3);
v = y(:,4:6);
vel_aps(4) = norm(v(1,:));
plot3(r(:,1),r(:,2),r(:,3),'LineWidth',2)

% ecc = 0
ecc = 1e-5;
SMA = const/(1+ecc);
orbit_period = 2*pi* sqrt(SMA^3/mu);
time_vector = [0:1:orbit_period, orbit_period];
[r0,v0]=kepler2ijk_hs(SMA,ecc,inc,RAAN,AOP,theta);
[t,y] = orbit_propagation(r0,v0,time_vector);
r = y(:,1:3);
v = y(:,4:6);
vel_aps(5) = norm(v(1,:));
plot3(r(:,1),r(:,2),r(:,3),'LineWidth',2)

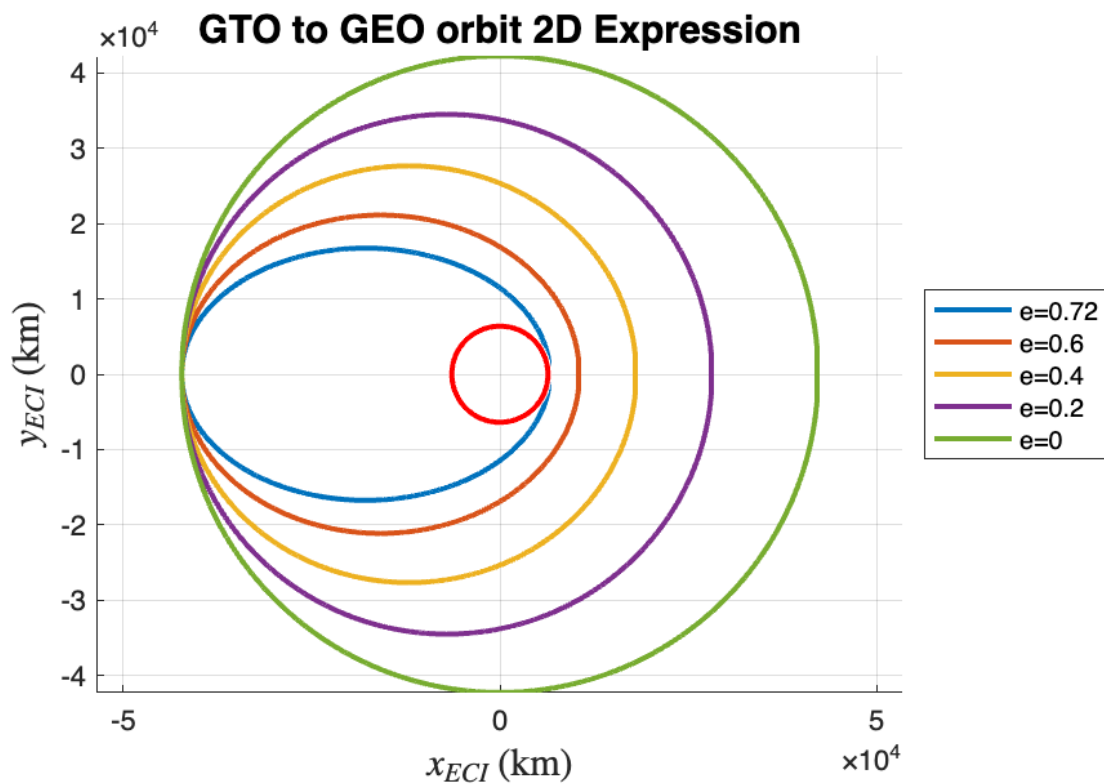
```

```

% Draw Earth
viscircles([0,0],6378);

axis('equal')
view(2)
legend('e=0.72','e=0.6','e=0.4','e=0.2','e=0','Location','eastoutside')
hold off
title('GTO to GEO orbit 2D Expression','FontSize',15,'FontWeight','bold')
xlabel('$x_{ECI}$ (km)', 'Interpreter', 'latex','FontSize',15)
ylabel('$y_{ECI}$ (km)', 'Interpreter', 'latex','FontSize',15)
grid on

```



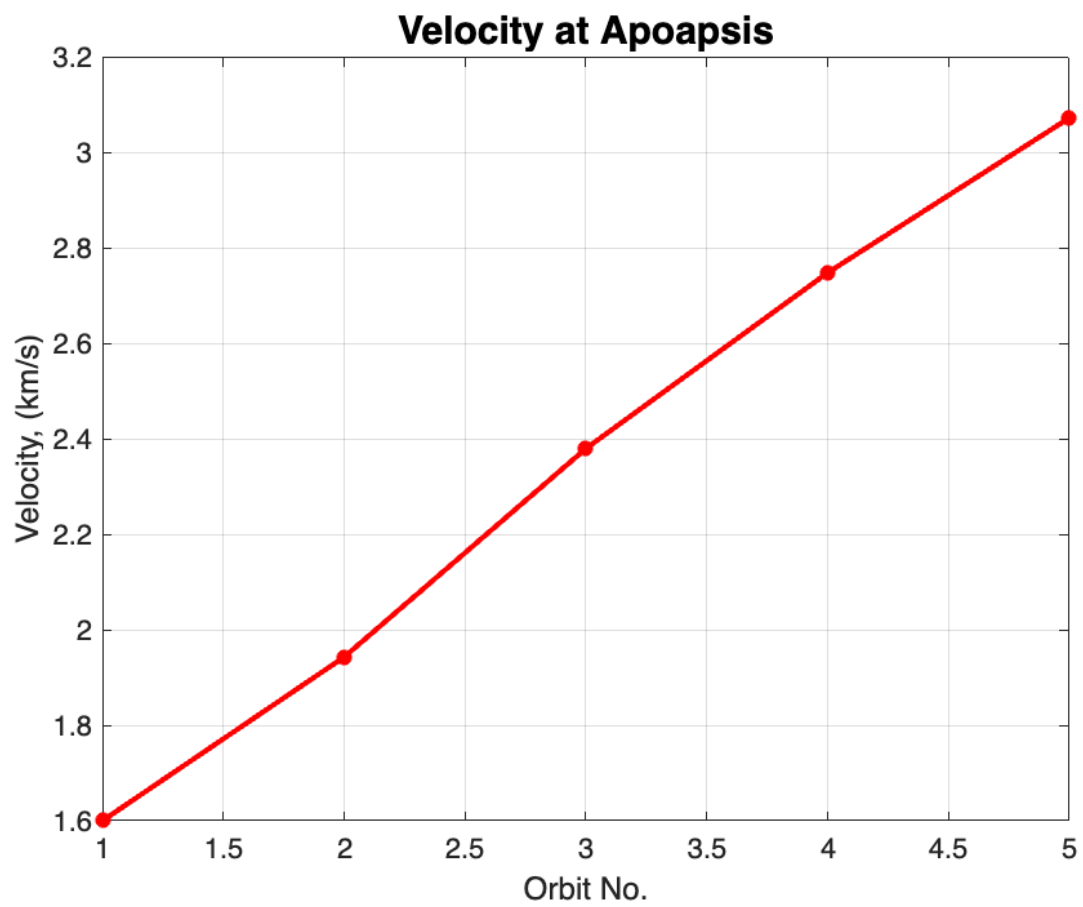
```

figure;

plot(1:5,vel_aps,'*-','LineWidth',2,'Color','r')
grid on
title('Velocity at Apoapsis','FontSize',15,'FontWeight','bold')

```

```
xlabel('Orbit No.')  
ylabel('Velocity, (km/s)')
```



Appendix 1: 2-body orbit propagator in ECI

```
function dydt = two_body_dynamics(t,y)

% Define the standard gravitational parameter for Earth (km^3/s^2)
mu = 398600.4415;

% Extract position and velocity from the input state vector
r_vec = y(1:3); % Position vector
v_vec = y(4:6); % Velocity vector

% Calculate the norm of the position vector
r_norm = norm(r_vec);

% Compute the acceleration based on the gravitational force
a1 = -mu*r_vec(1)/(r_norm^3);
a2 = -mu*r_vec(2)/(r_norm^3);
a3 = -mu*r_vec(3)/(r_norm^3);

% Combine velocity and acceleration to form the derivative of the state vector
dydt = [v_vec; a1;a2;a3];
end

function [t,y] = orbit_propagation(r0, v0, tVec)

odeoptions = odeset('RelTol',1e-10,'AbsTol',1e-20);
% Calculate the magnitude of the position and velocity vectors
y0 = [r0;v0];

% Solve the ODE using ode45
[t, y] = ode45(@two_body_dynamics, tVec, y0, odeoptions);
end
```

Appendix 2: ECI to Keplerian Converter (kepler2ijk, ijk2kepler)

```
function [r,v] = kepler2ijk_hs(a,e,inc,RAAN,AOP,theta)

mu = 398600.4415;

RAAN = RAAN/180*pi;
inc = inc/180*pi;
AOP = AOP/180*pi;
```

```

theta = theta/180*pi;

R_3_AOP = [cos(AOP), sin(AOP),0;
           -sin(AOP), cos(AOP),0;
           0,          0,1];

R_1_inc = [1,      0,      0;
           0, cos(inc), sin(inc);
           0,-sin(inc), cos(inc)];

R_3_RAAN = [cos(RAAN), sin(RAAN),0;
            -sin(RAAN), cos(RAAN),0;
            0,          0,1];

Q_rotation = R_3_AOP * R_1_inc * R_3_RAAN;

p = Q_rotation(1,:);
q = Q_rotation(2,:);
w = Q_rotation(3,:);

p_value = a*(1-e^2);
r = p_value* (cos(theta)*p + sin(theta) *q)/(1+e*cos(theta));
r = r';
v = sqrt(mu/p_value)*(-sin(theta)*p+(e+cos(theta))*q)';
end

function [a, e, inc, RAAN, AOP, theta] = ijk2kepler_hs(r,v)

mu = 398600.4415;

h = cross(r,v); % Angular Velocity Vector
z = [0,0,1]';
n = (cross(z,h))/(norm(cross(z,h))); % Normal Vector
e_vector = ((cross(v,h))/mu) - r/(norm(r)); % Eccentricity vector

a = (-mu)/((norm(v))^2 - 2*mu/norm(r)); % 1. Semi-major axis
e = norm(e_vector); % 2. Eccentricity
inc = acos(dot(z,h)/norm(h)) / pi *180; % 3. Inclination

RAAN = atan2(n(2),n(1)); % 4. RAAN
RAAN = RAAN / pi * 180;

```

```

AOP = acos((dot(n,e_vector))/norm(e_vector));           % 5. Argument
of periapse
if(e_vector(3)<0)
    AOP = 2*pi - AOP;
end
AOP = AOP/pi*180;

theta = acos((dot(r,e_vector))/(norm(r)*norm(e_vector))); % 6. True
anomaly
if(dot(r,v) < 0)
    theta = 2*pi-theta;
end
theta = theta / pi * 180;
end

```