

# The Physics of Neutron Stars

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KICP

9 November 2019  
Compton Lectures  
University of Chicago

**NOTE!**  
**next week's lecture will be in**  
**KPTC 120**

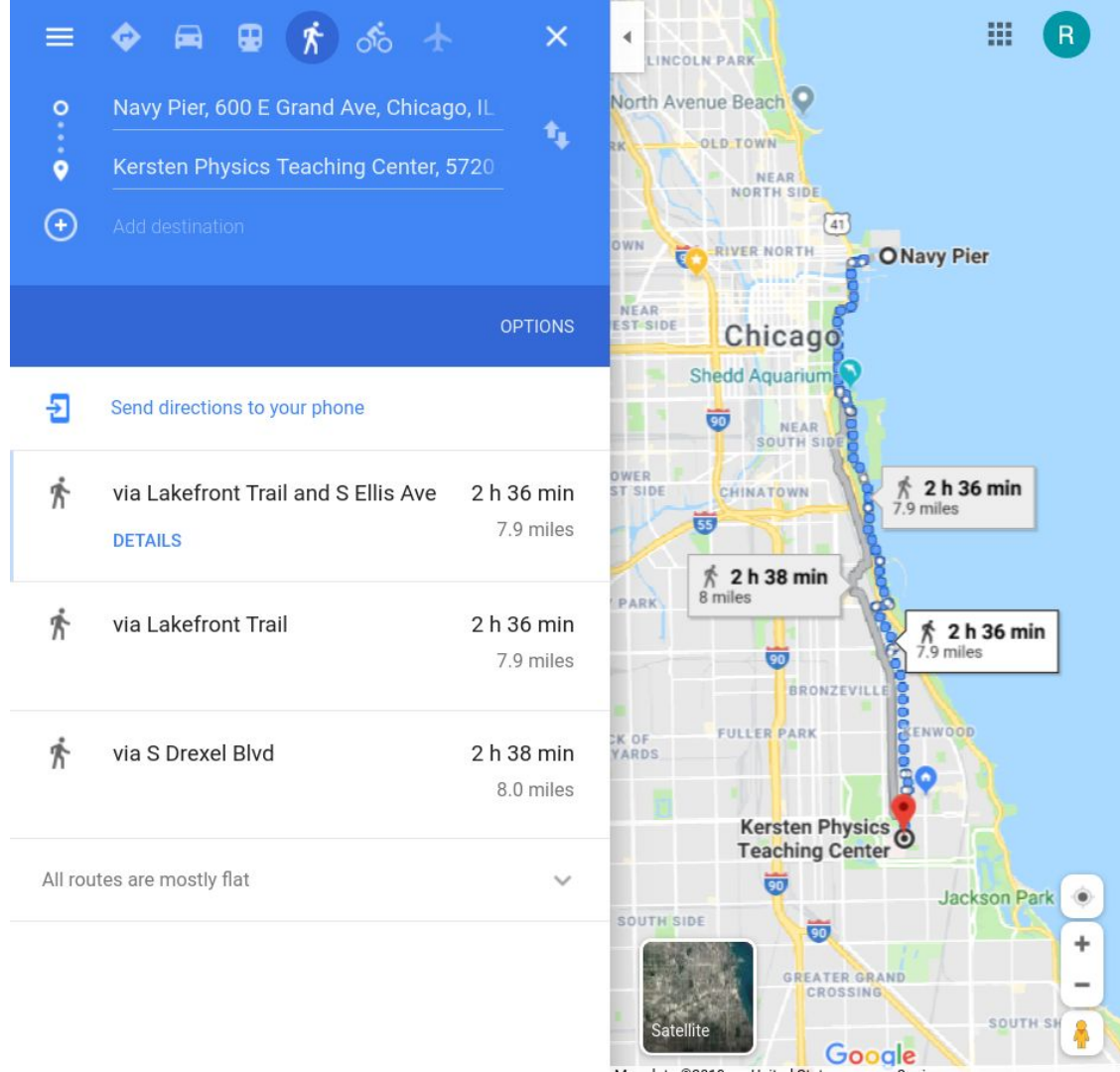
<https://github.com/reedessick/compton-lectures-2019>

# Review

Neutron stars go by many names

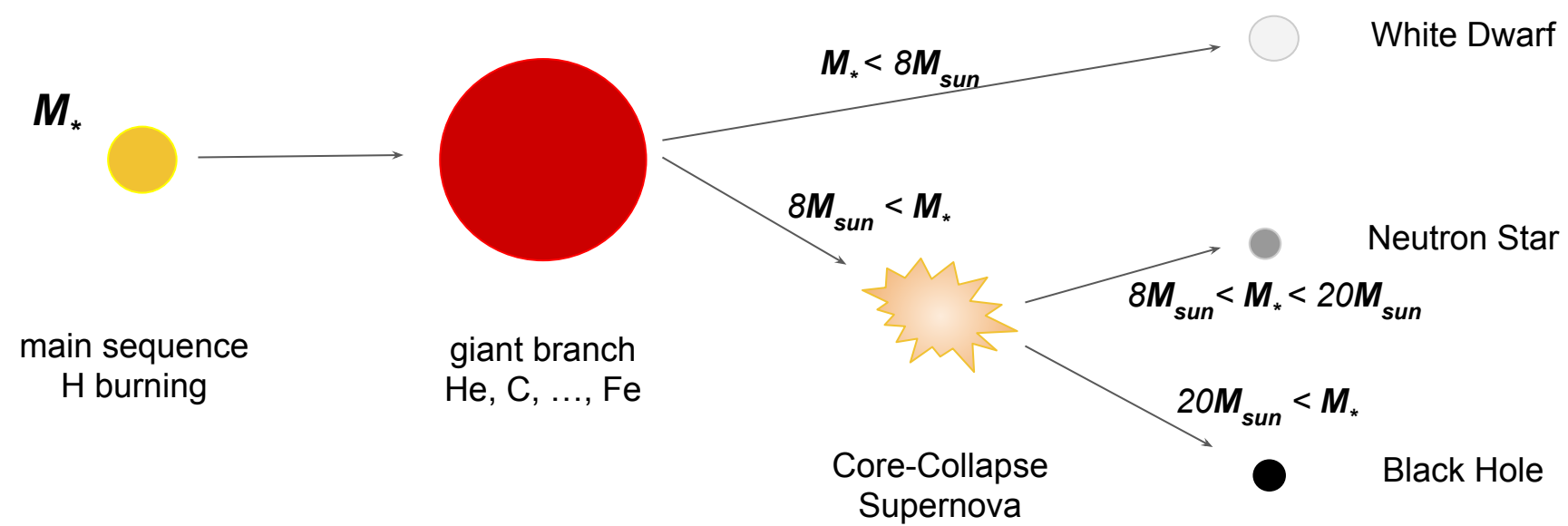
- Pulsar
- Magnetar
- Low-mass X-ray binary
- High-mass X-ray binary

**The mass of the sun in a ball  
16 miles in diameter**

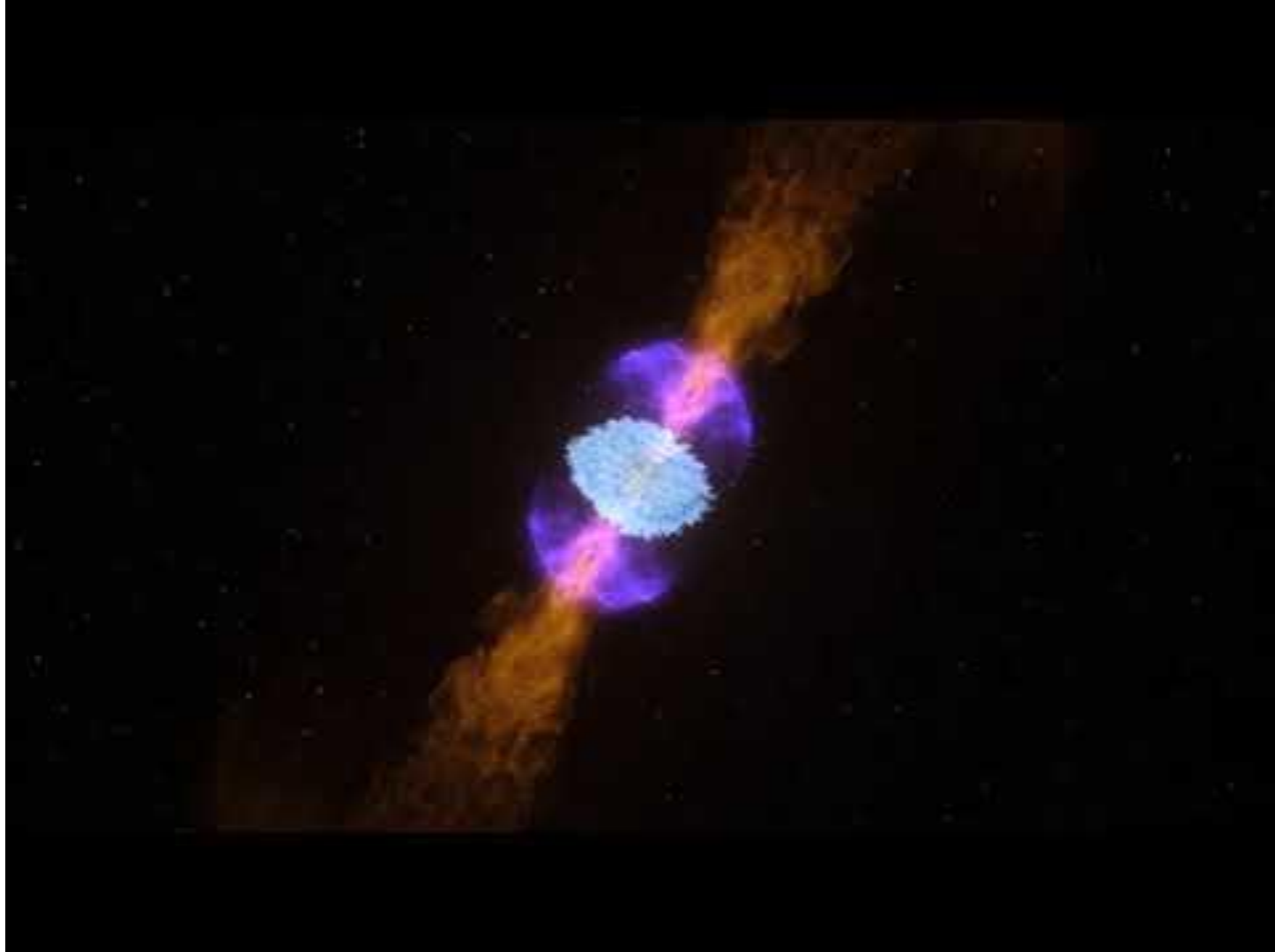


# Review

**Neutron Stars** and **Black Holes** are the end states of massive stars



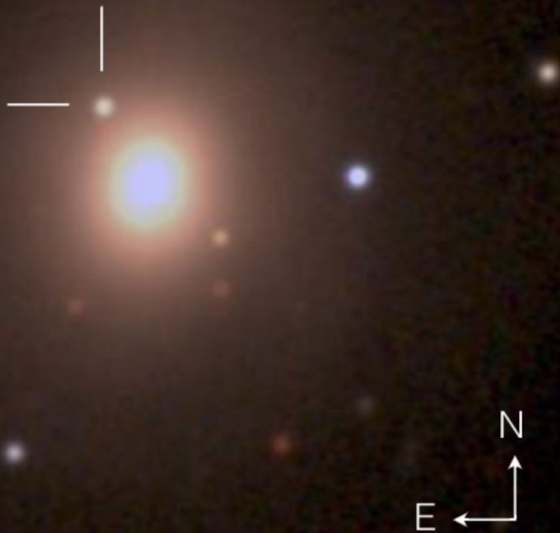
# Review



# Review

GW170817

DECam observation  
(0.5–1.5 days post merger)



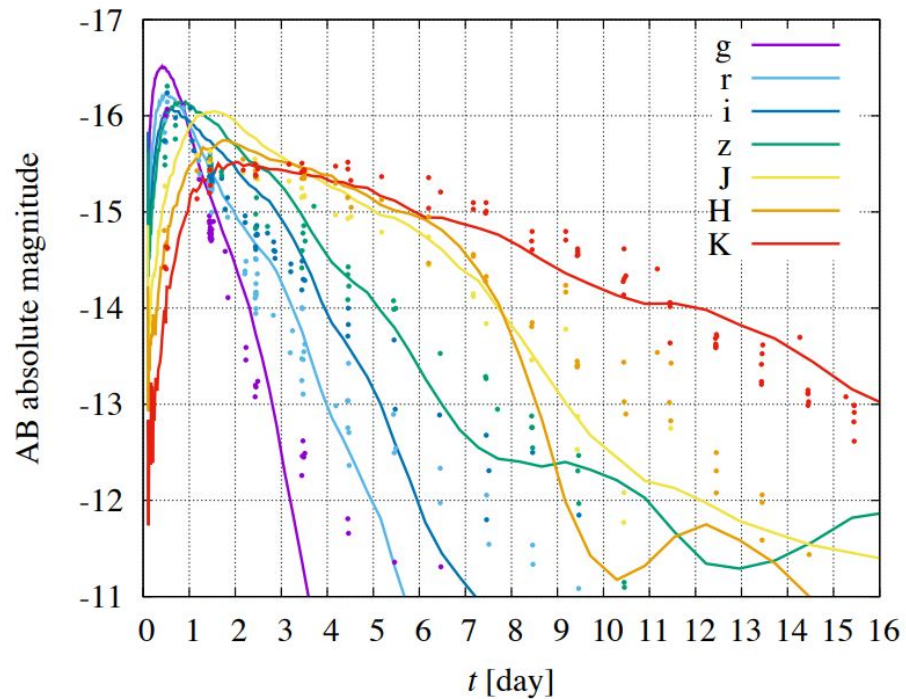
GW170817

DECam observation  
(>14 days post merger)

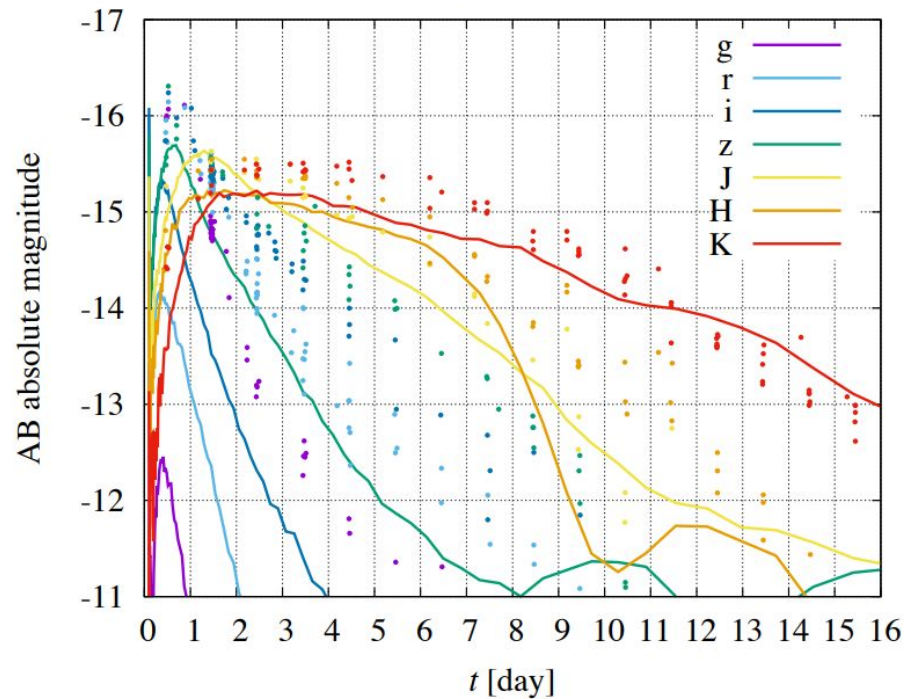


# Review

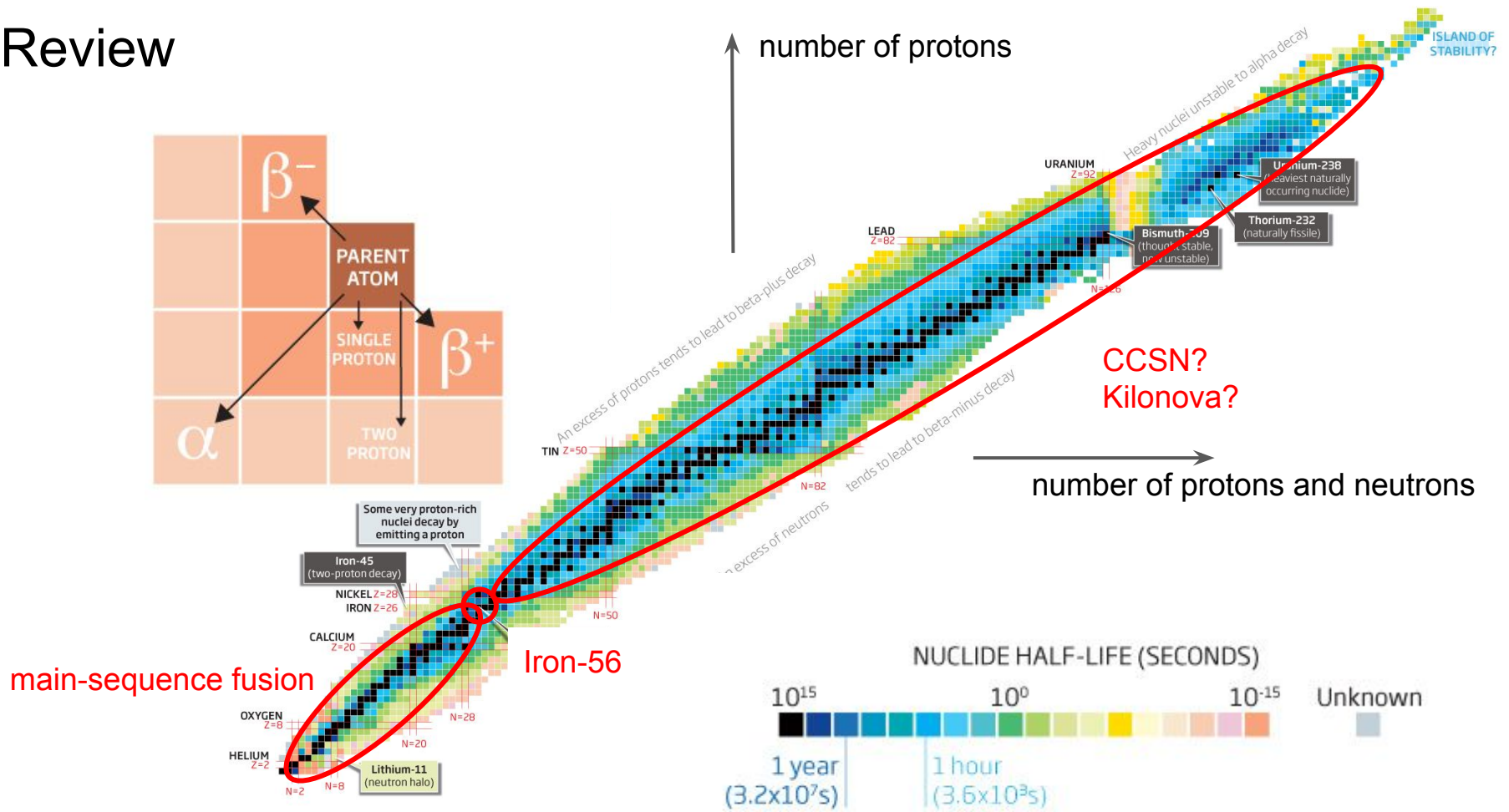
$D=40 \text{ Mpc}, 20^\circ \leq \theta < 28^\circ$



$D=40 \text{ Mpc}, 86^\circ \leq \theta < 90^\circ$

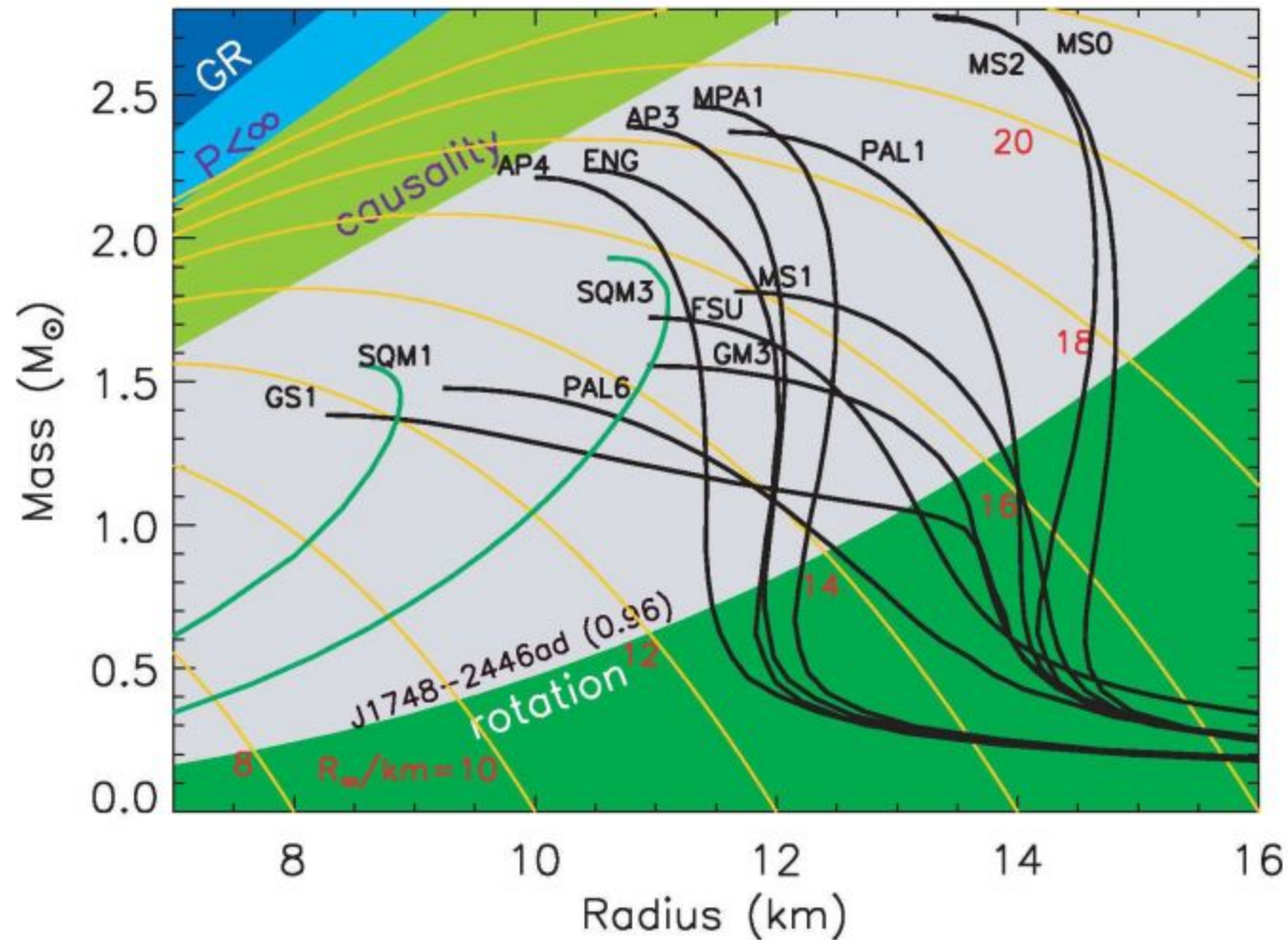


# Review



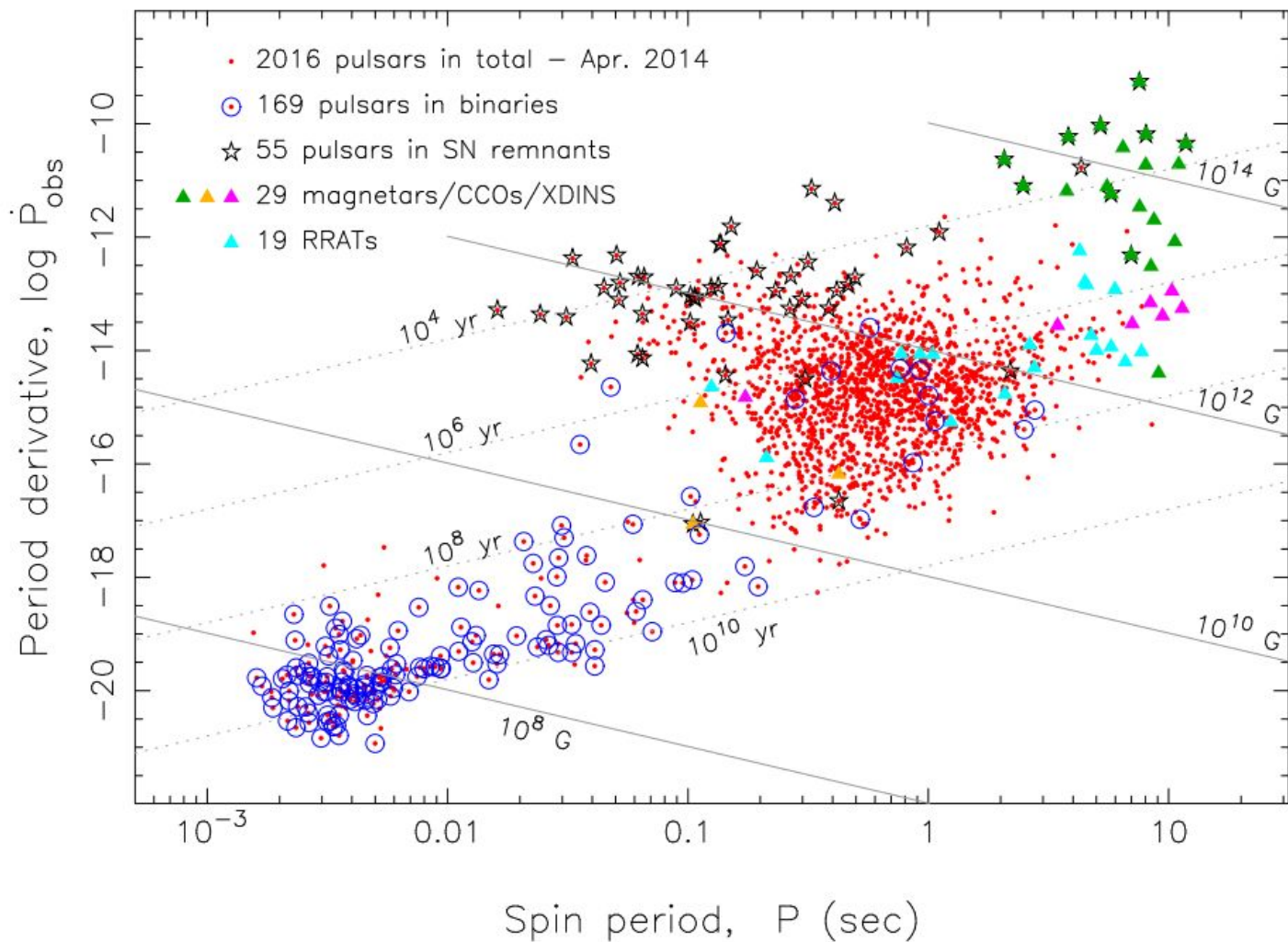


What are  
Neutron Stars?



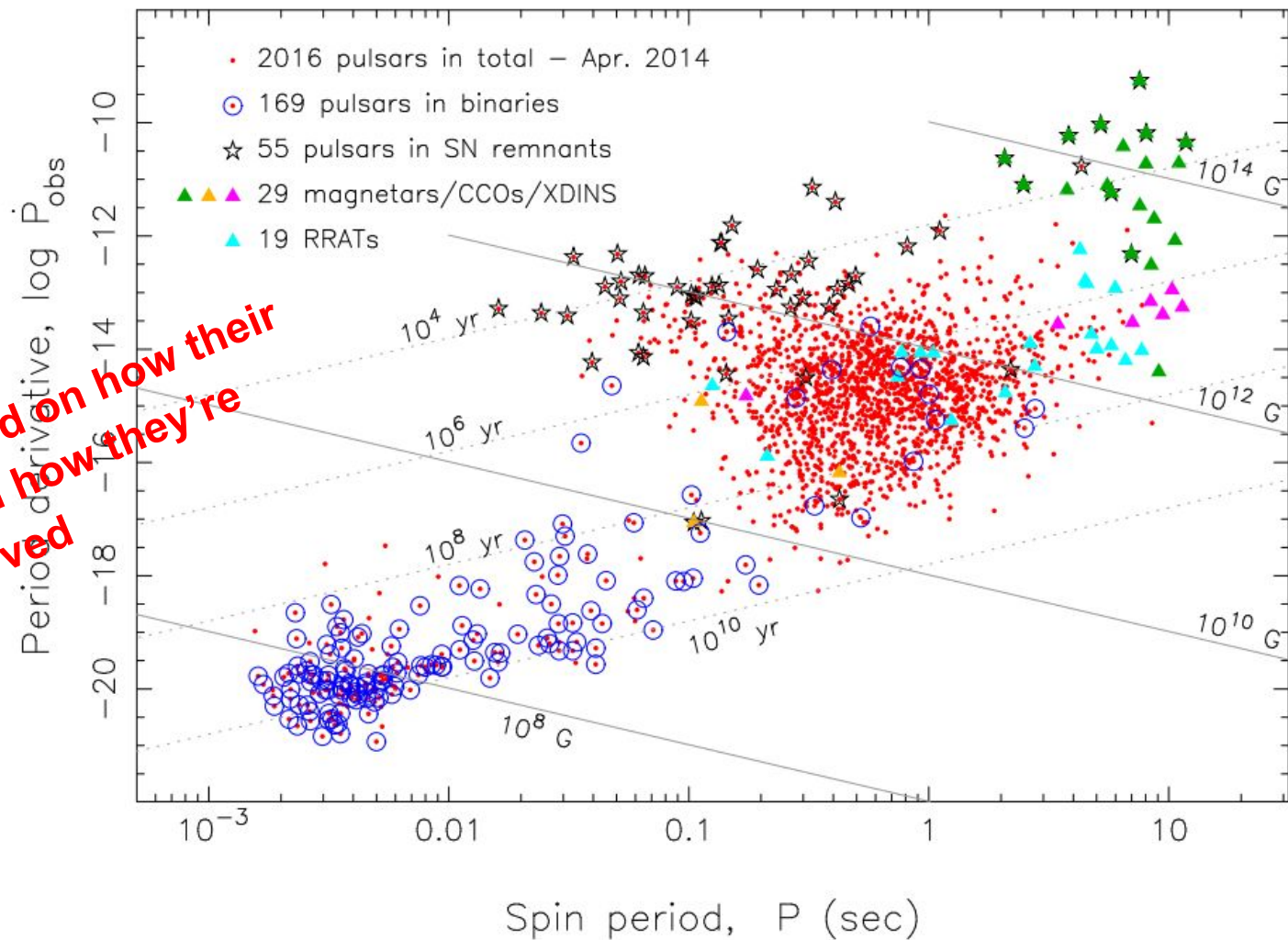


# What are Neutron Stars?



# What are Neutron Stars?

different names based on how their  
environment and how they're  
observed



# What are Neutron Stars?

- How do stars support themselves against gravity?
- What is an Equation of State?

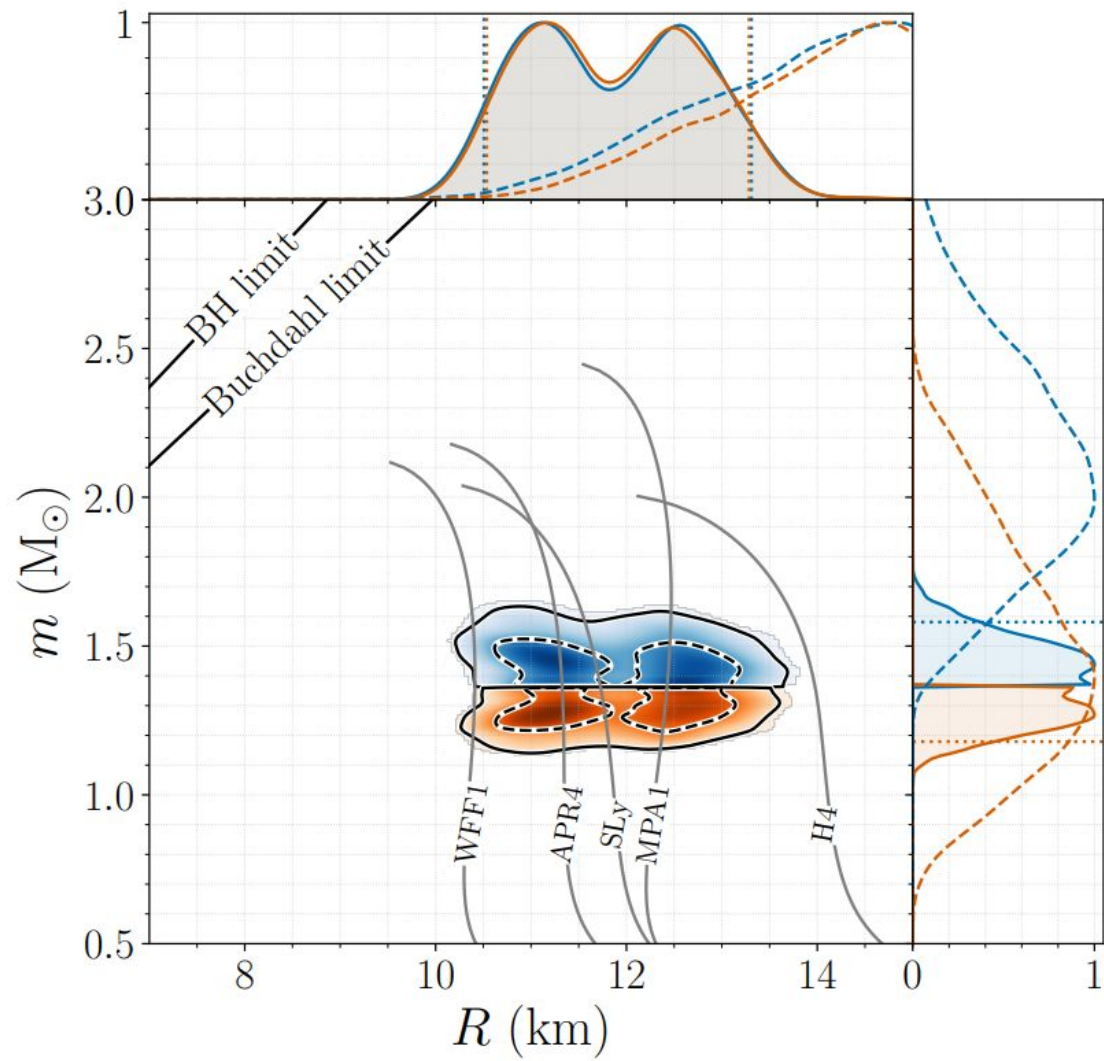
# What are Neutron Stars?

- How do stars support themselves against gravity?
- What is an Equation of State?

# What do NS do in CBCs?

- Linear tidal deformability
- Dynamical tides, nonlinear tides
- Tidal disruption

# EOS Inference with GWs

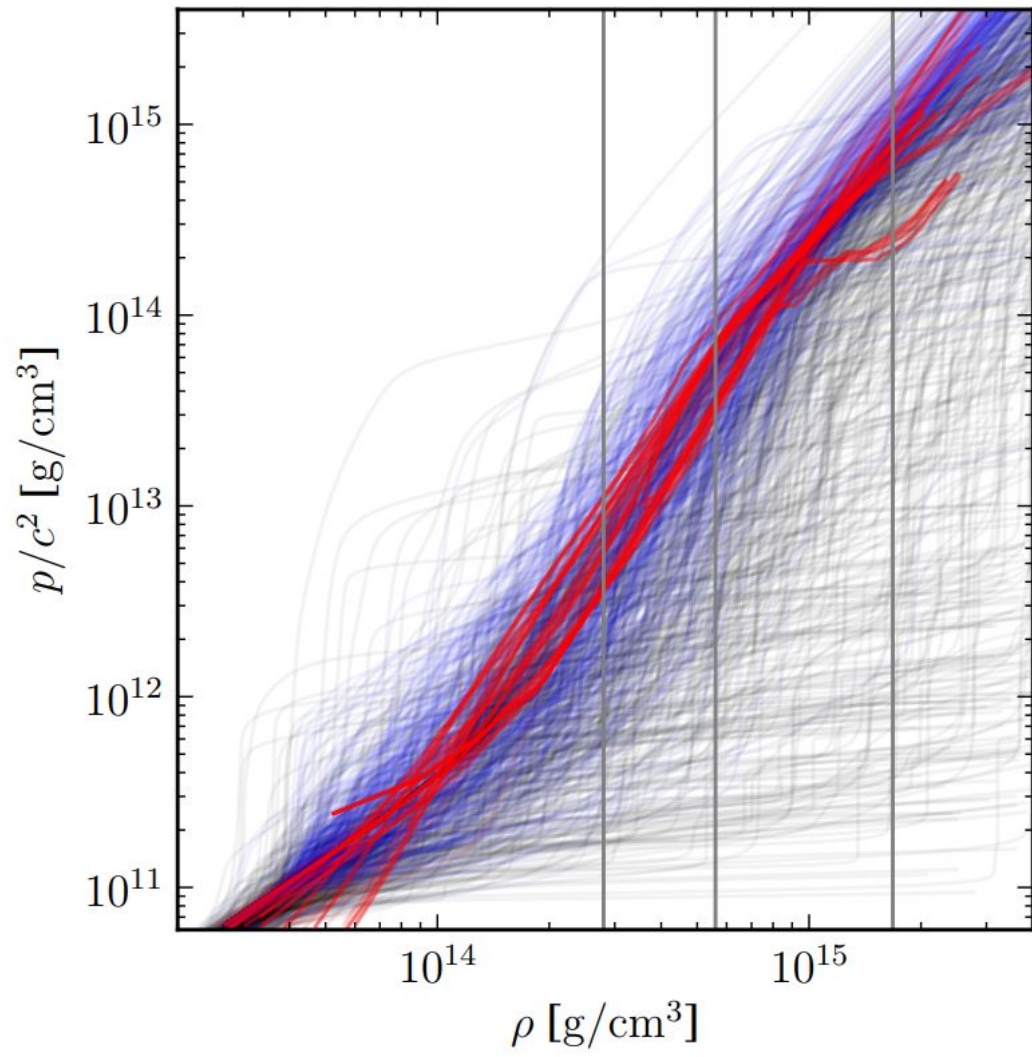


# EOS Inference with GWs

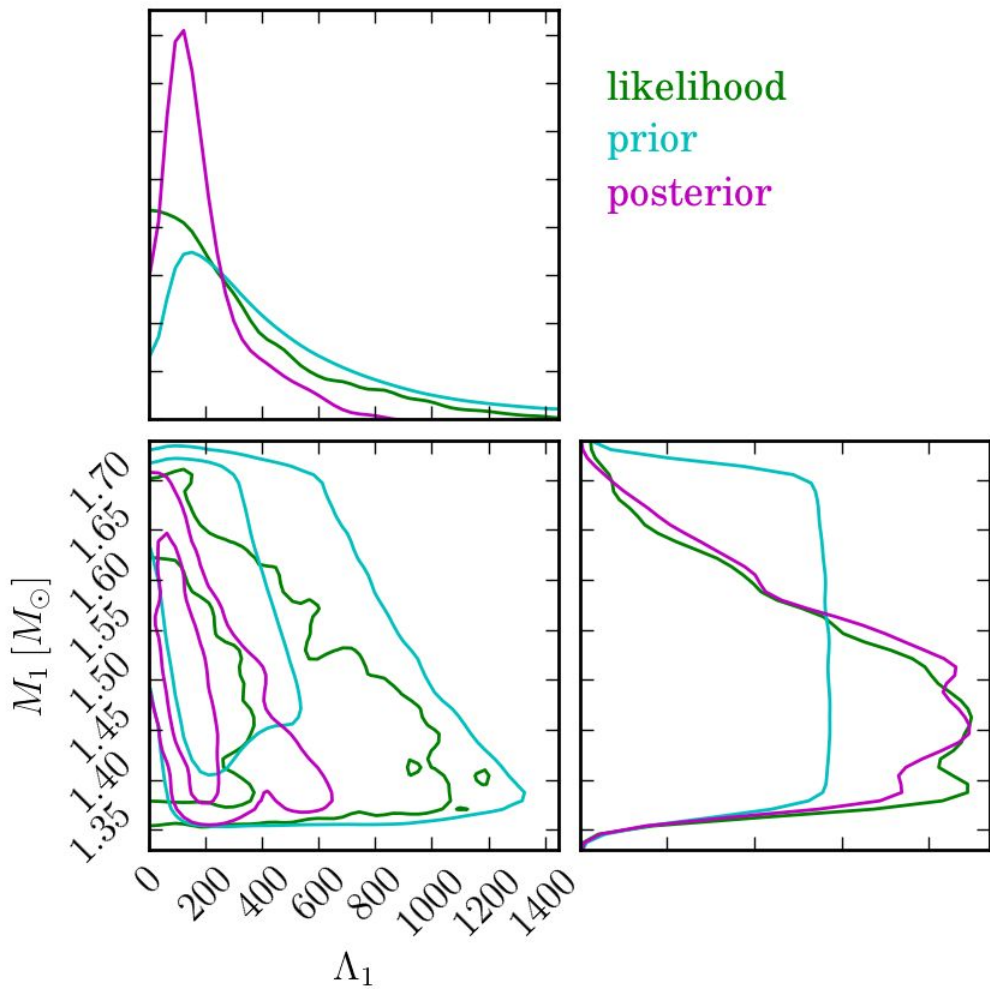
- Self-consistently incorporate information from arbitrary tabulated EOS models
- Automatically incorporate causality constraints and thermodynamic stability
- Allow for large amounts of model freedom
- Incorporate transparent priors



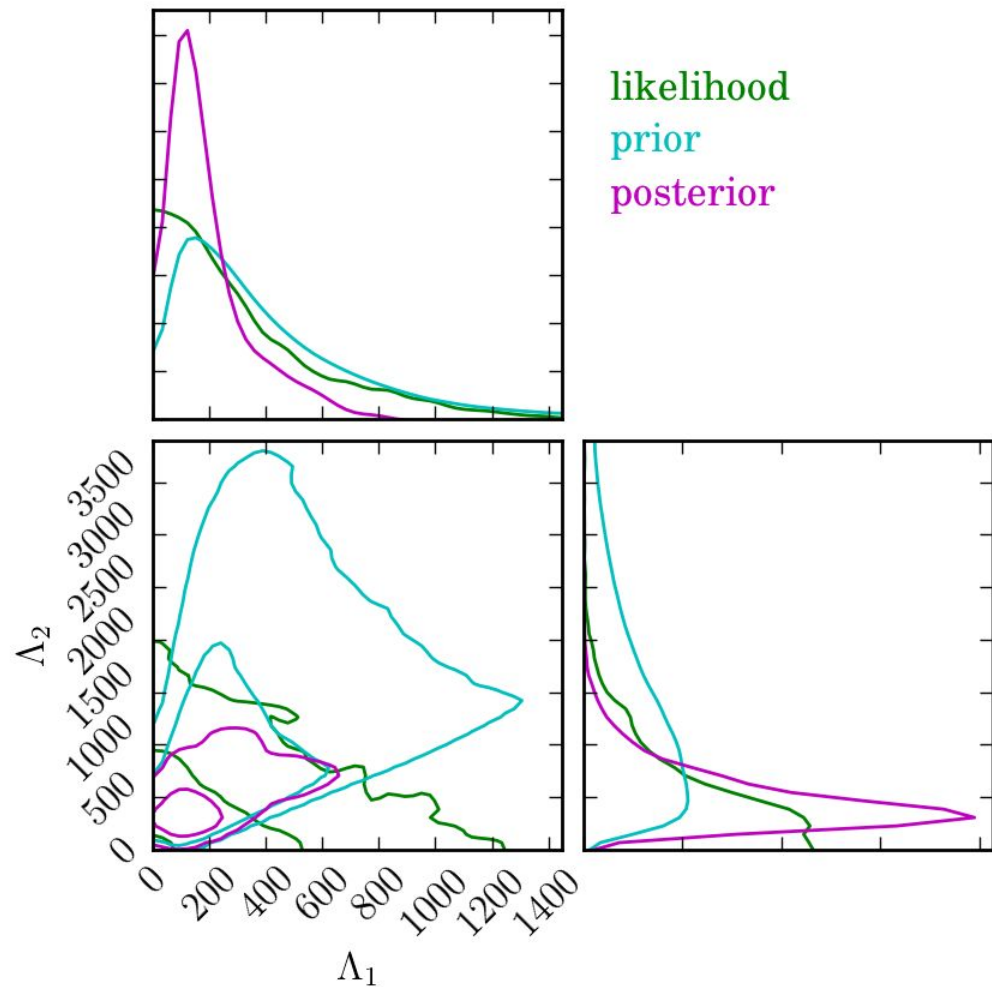
# EOS Inference with GWs



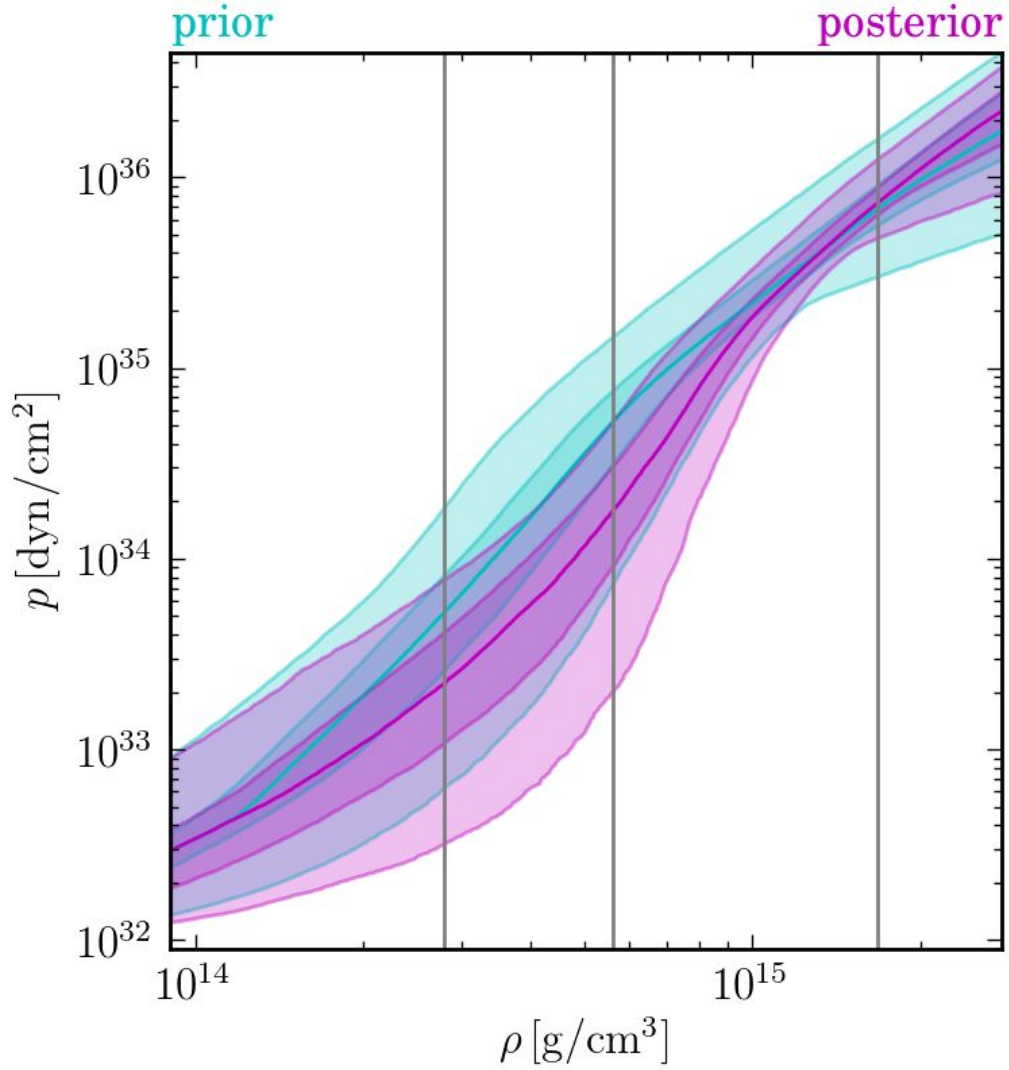
# EOS Inference with GWs



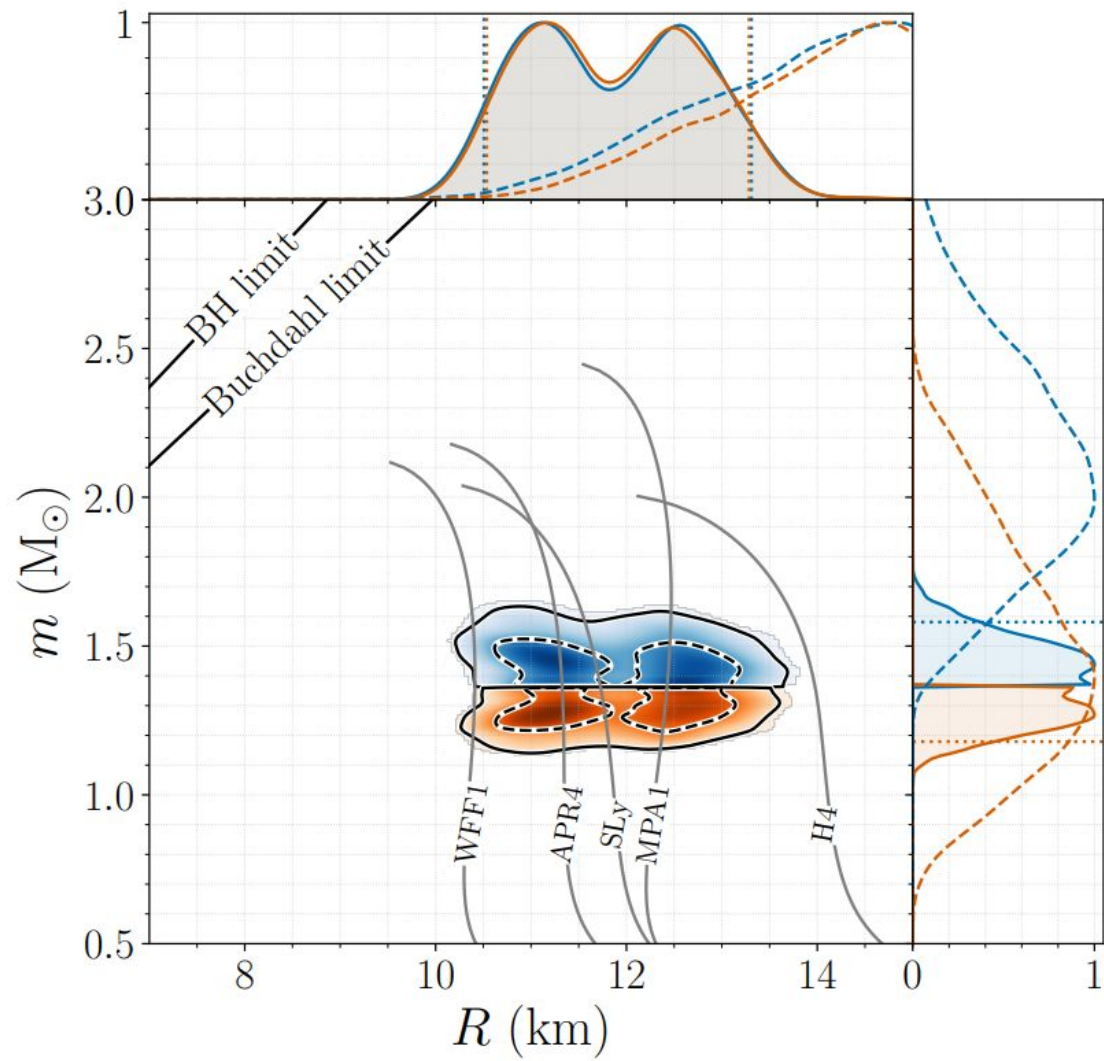
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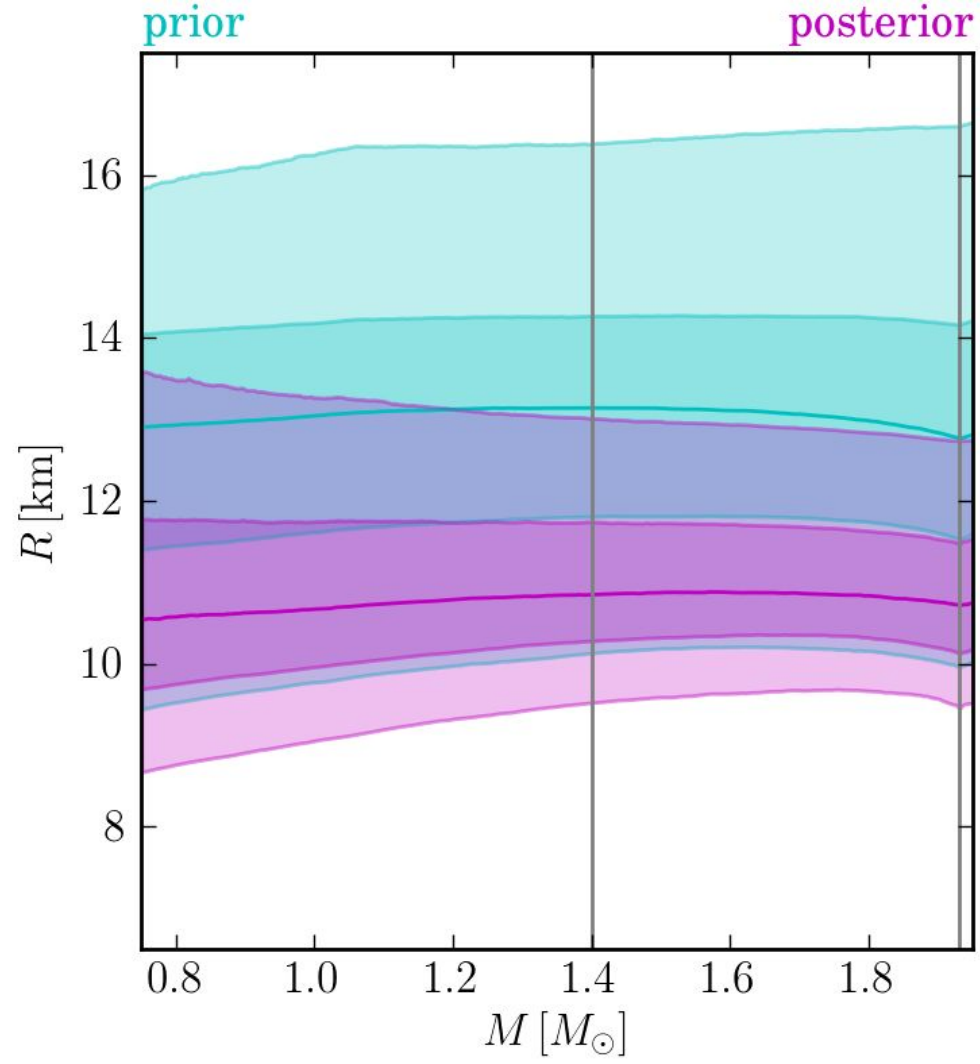
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# EOS Inference with GWs

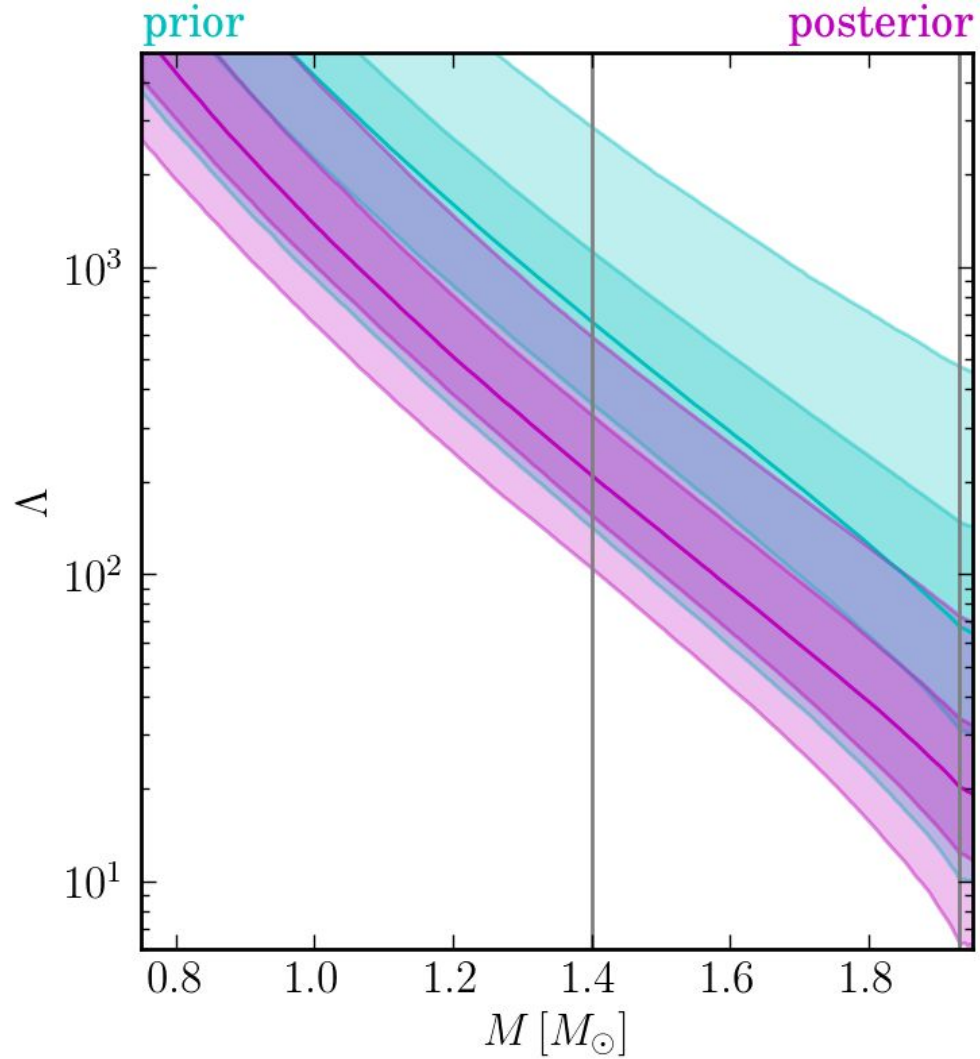


# EOS Inference with GWs

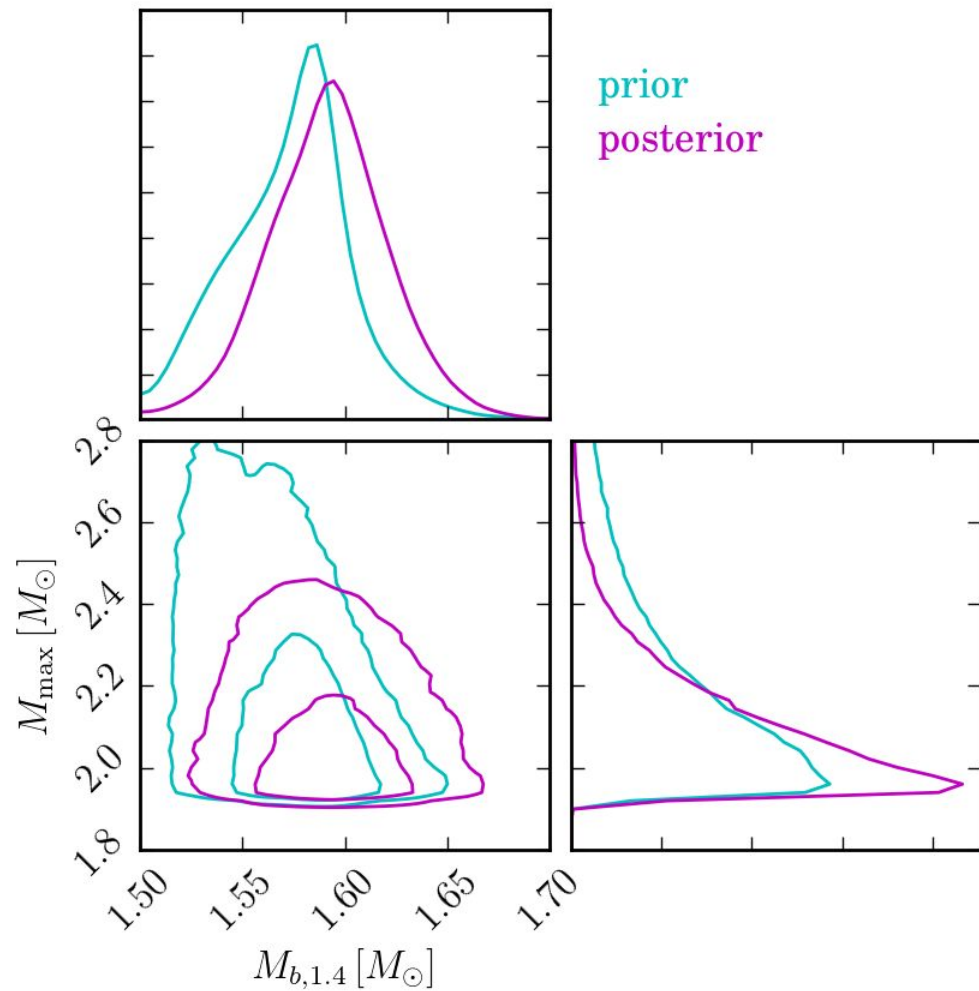




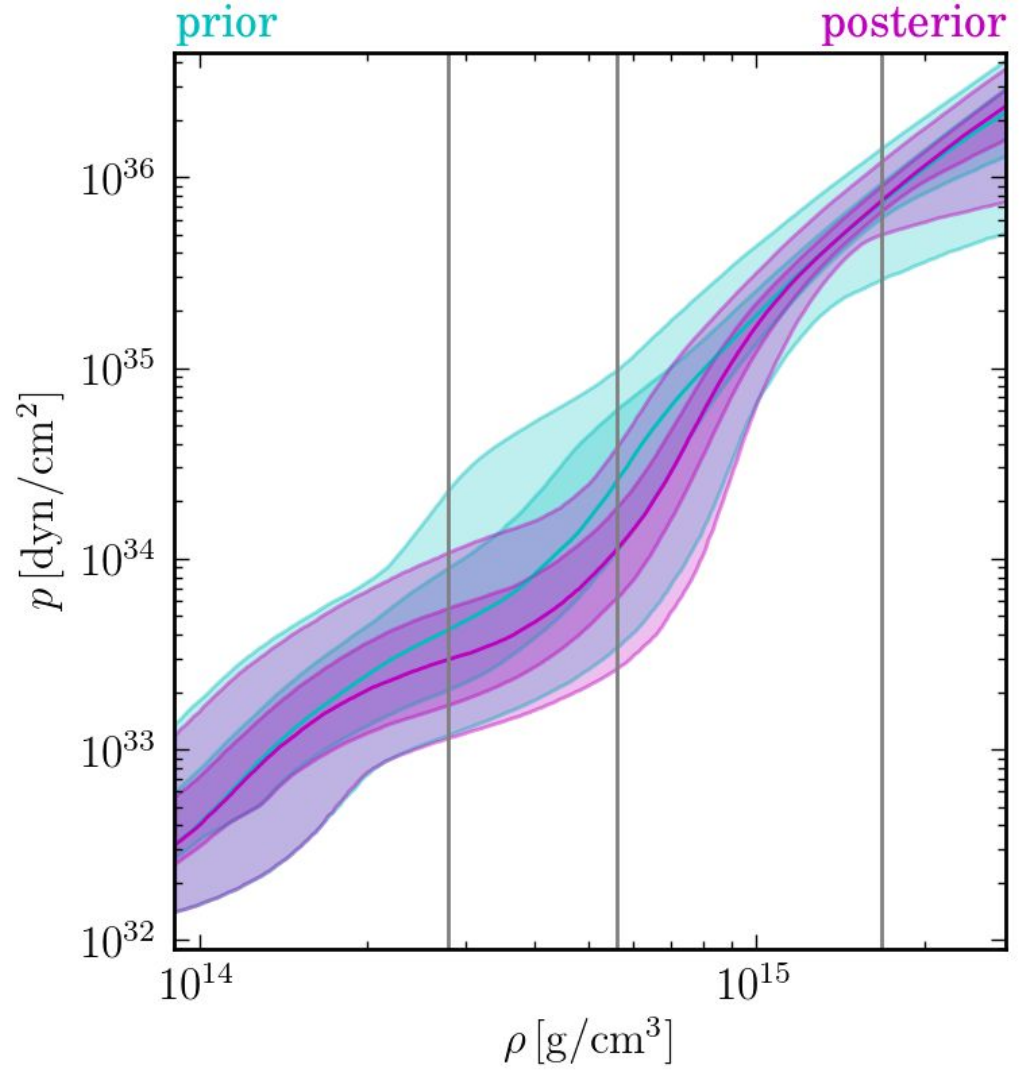
# EOS Inference with GWs



# EOS Inference with GWs



# EOS Inference with GWs



# Next time

## Testing General Relativity with Gravitational Waves

- What types of tests have been performed?
- What have we learned about gravity from these tests?
- How can we use GR's accuracy to improve our detectors?

**NOTE!**  
Next week's lecture will be in  
**KPTC 120**

# Suggested Reading

- [P. Landy and R. Essick. \*Non-parametric inference of the neutron star equation of state from gravitational wave observations\*. arXiv:1811.12529 \(2019\).](#)
- [R. Essick, P. Landry, and D. Holz. \*Nonparametric inference of neutron star composition, equation of state, and maximum mass with GW170817\*. arXiv:1910.09740 \(2019\).](#)
- [\*How Big are Neutron Stars?\* <https://www.ligo.org/science/Publication-GW170817EoS/index.php>](#)

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