Searching for Gravitational Waves in Noisy Detector Data

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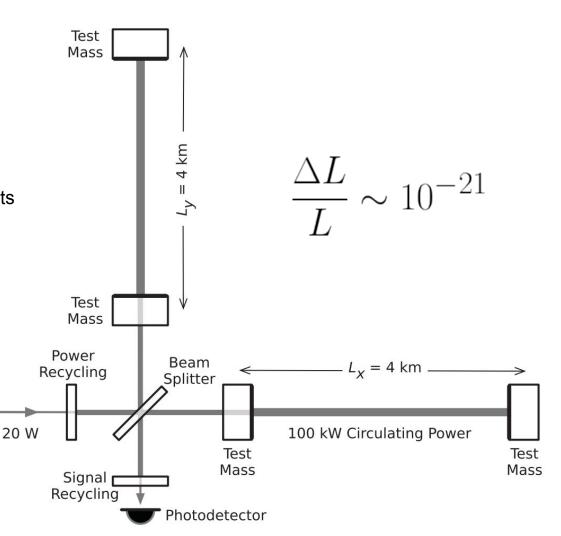
Review

GWs manifest as length changes in our detector

Interferometers make timing measurements to measure changes in length

Laser Source

There are various noise sources within terrestrial detectors



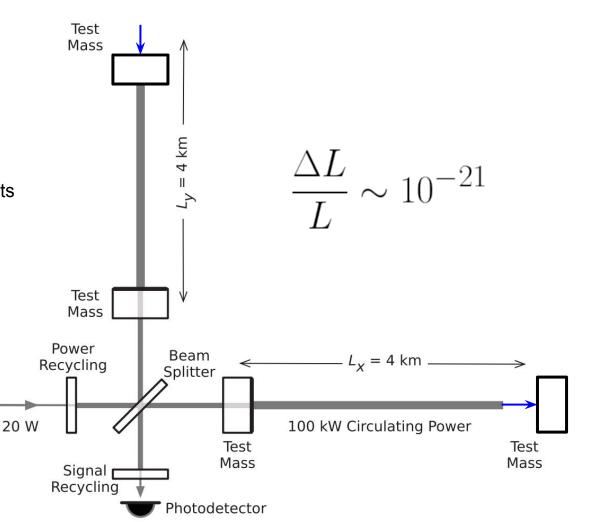
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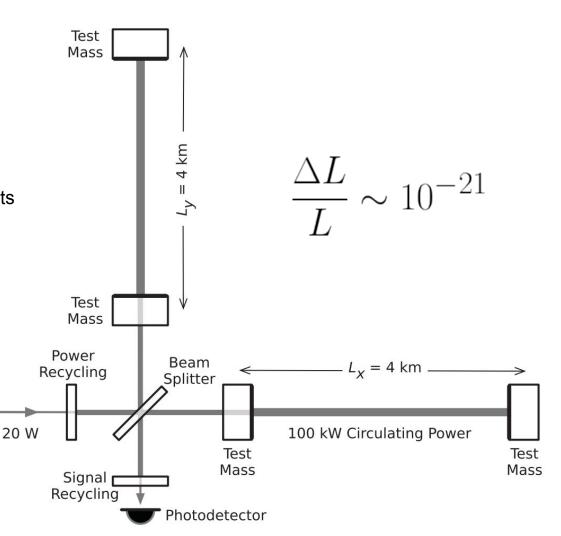
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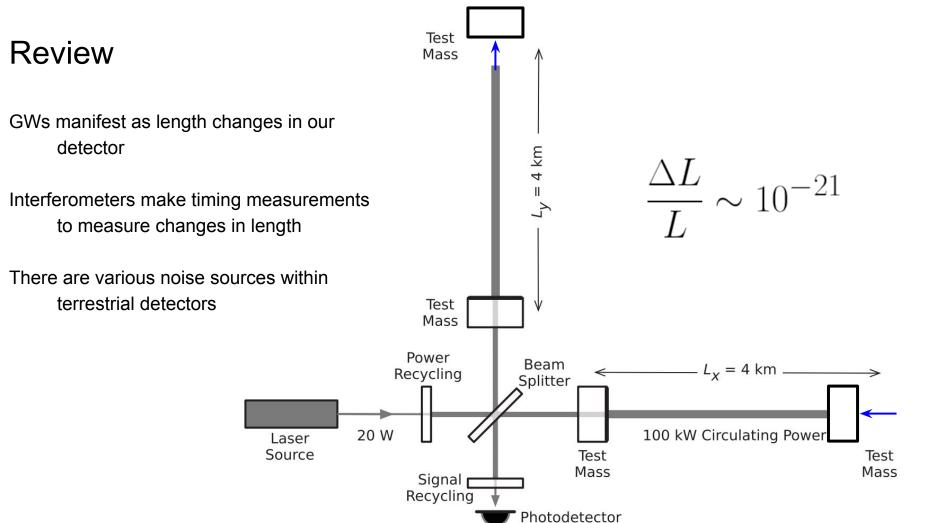
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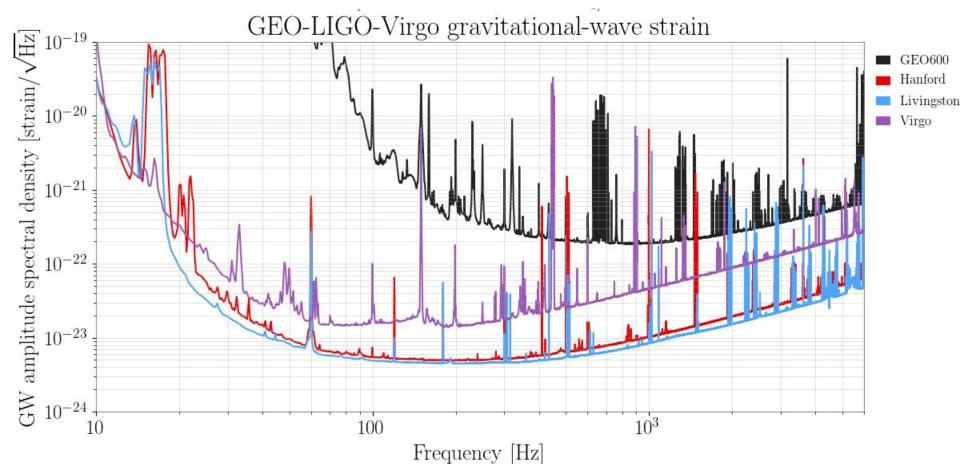
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Interferometer Noise Characteristics



Search Techniques

modeled signals

If you know what the signal looks like, run a *matched filter*.

Search Techniques

modeled signals

Search Techniques

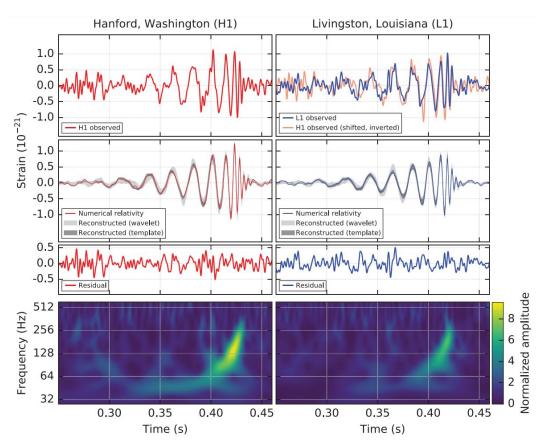
unmodeled signals

If you do not know what the signal looks like, use a *coherent likelihood* or *cross-correlation between detectors*

typically less sensitive than matched filter searches

Search Techniques

unmodeled signals



What is a "background"?

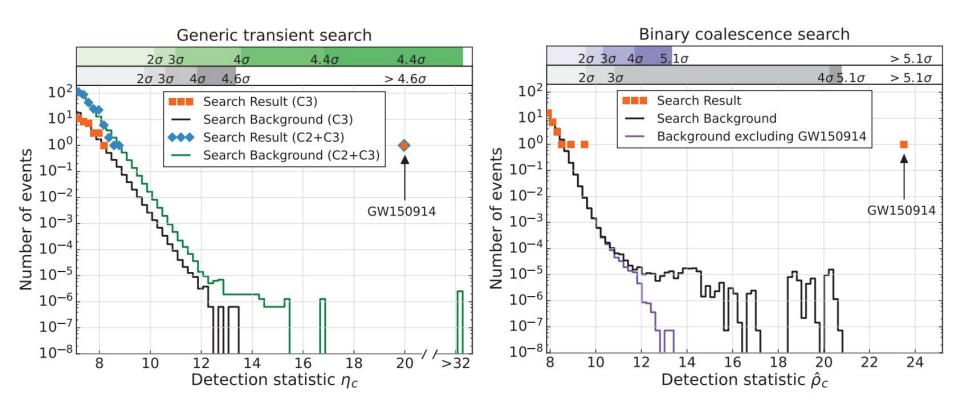
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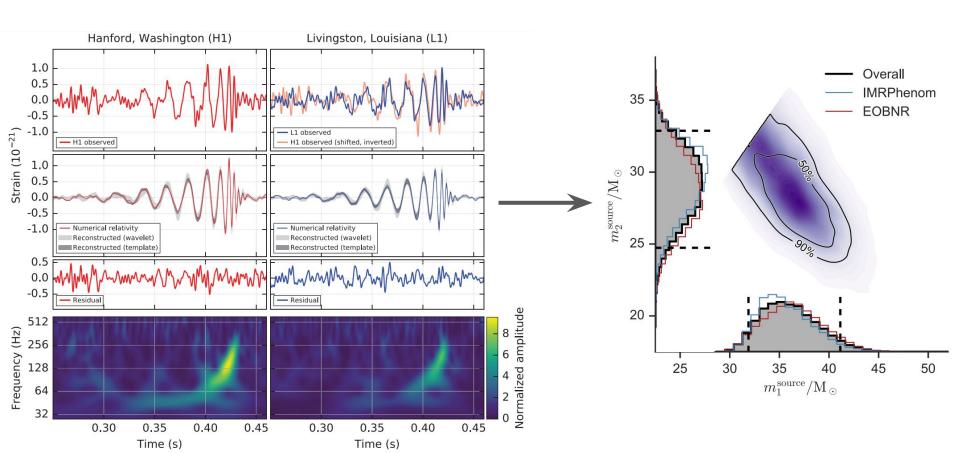
How well can we distinguish "background" (noise) from "foreground" (signals)?

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How well can we distinguish "background" (noise) from "foreground" (signals)?

How do we measure background distributions?





Why can't we know exactly what the signal parameters were?

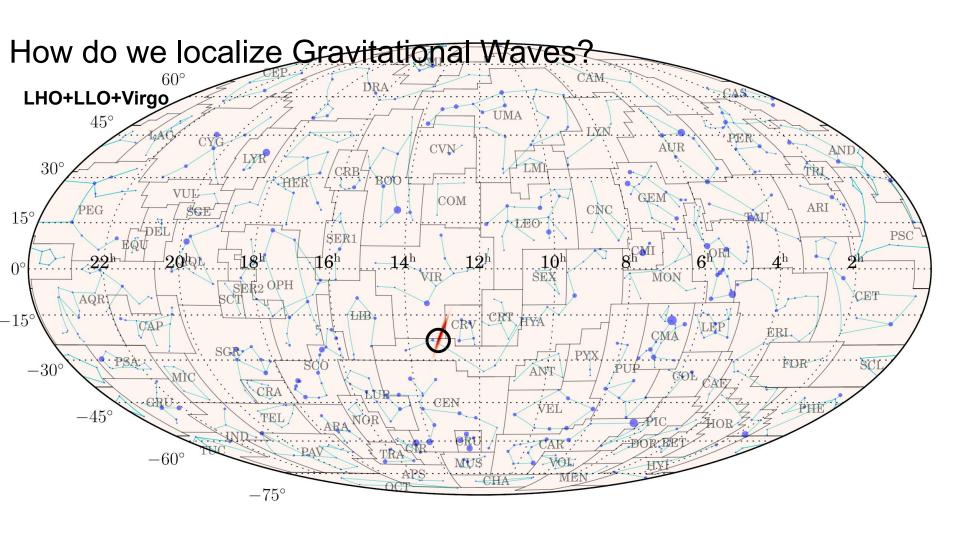
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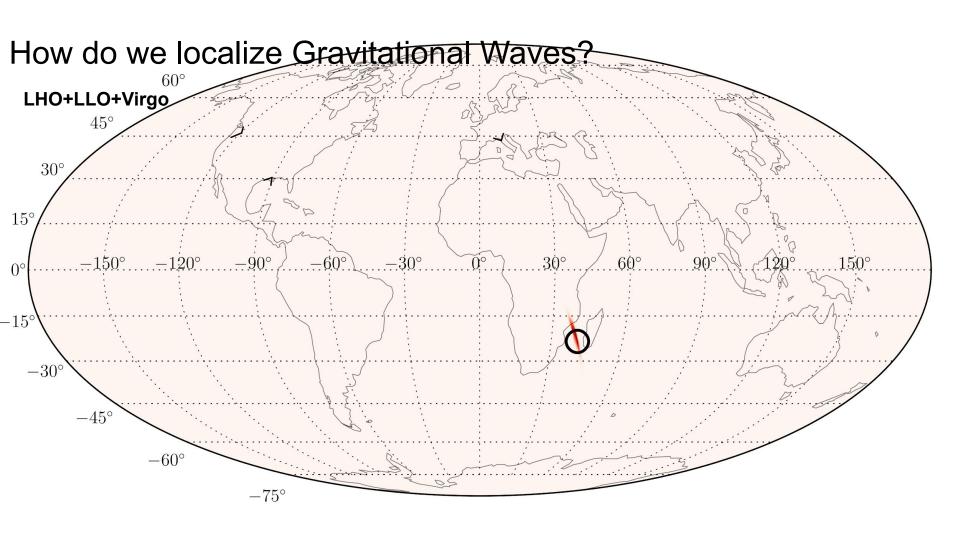
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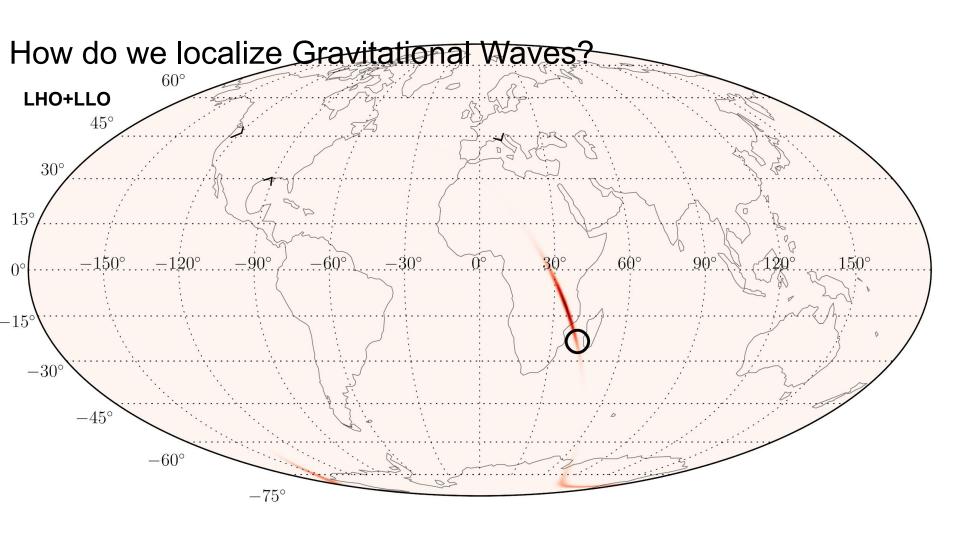
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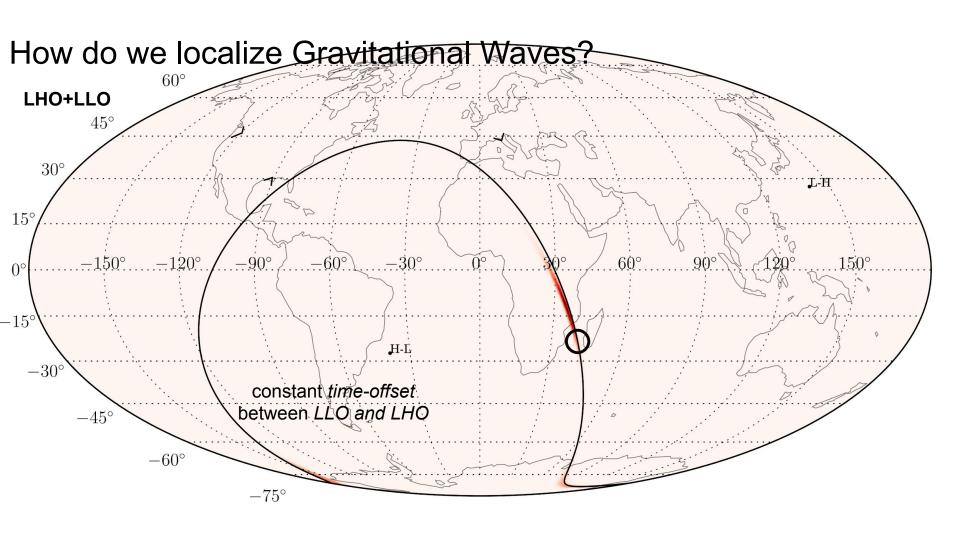
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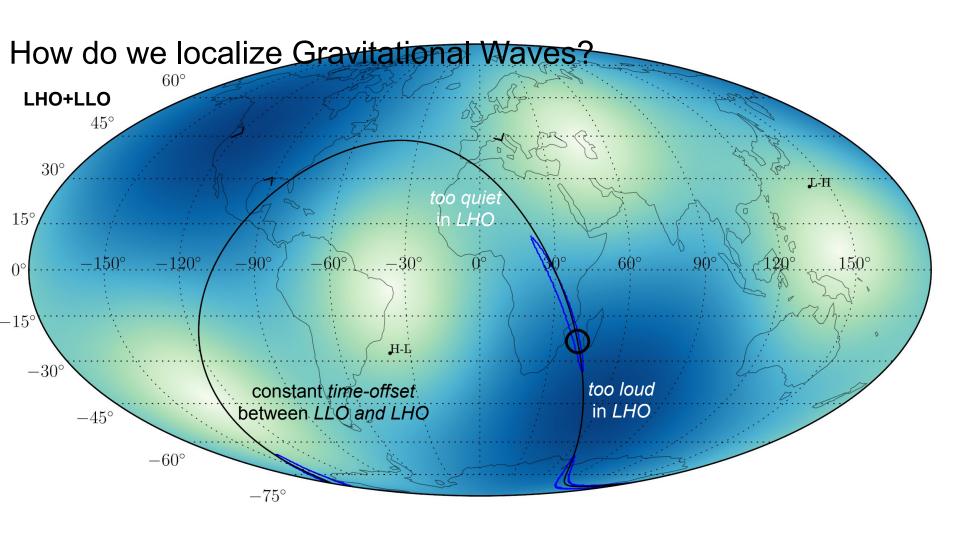
How do LIGO+Virgo come up with their probability distributions for signal parameters?

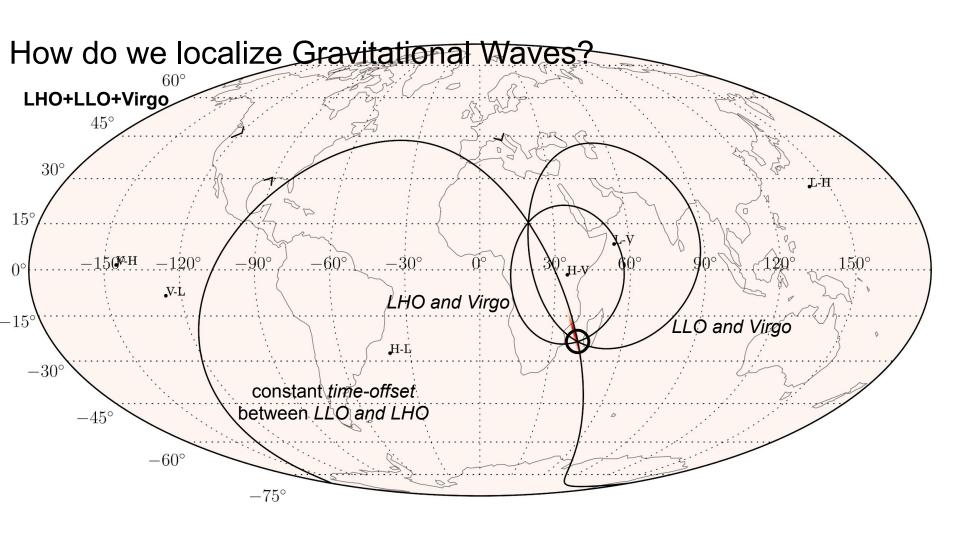




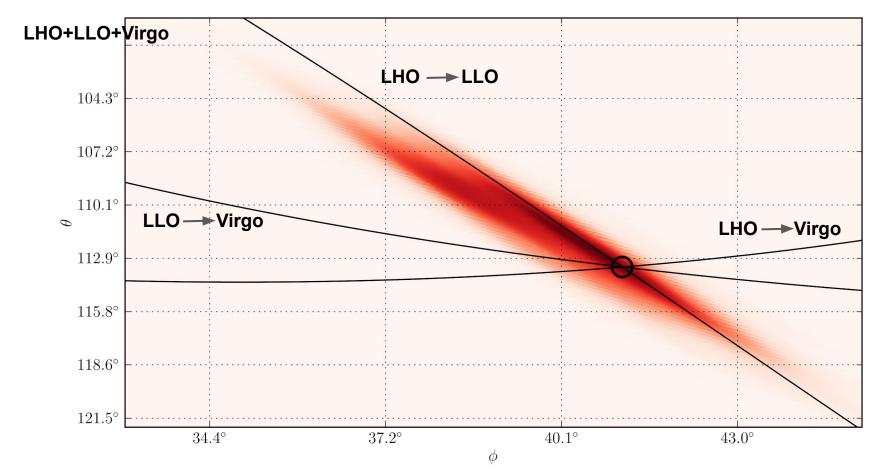


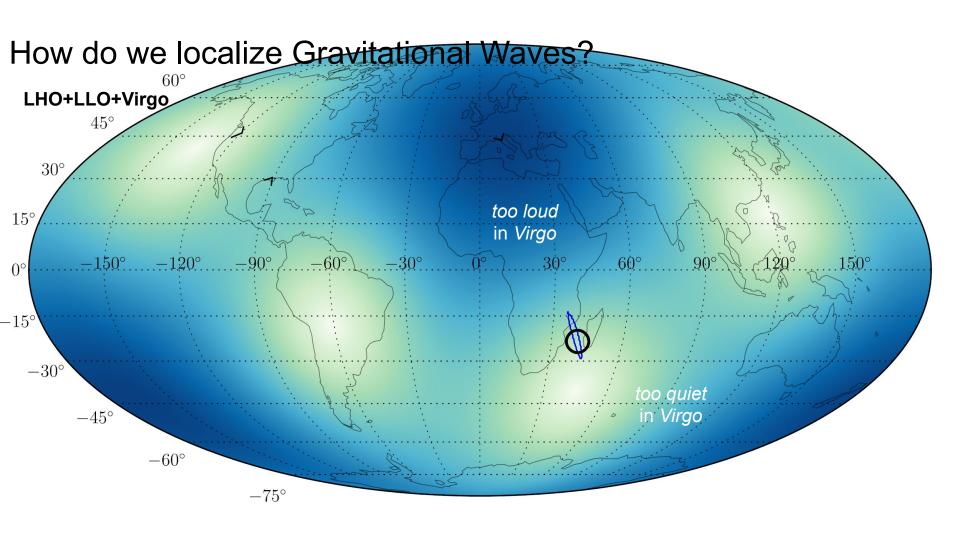


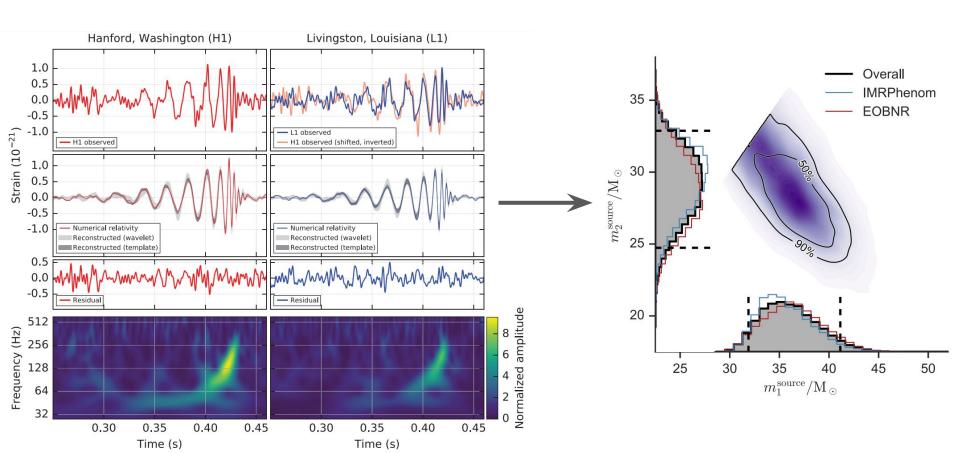




How do we localize Gravitational Waves?







Next time

Gravitational-Wave Populations and Cosmology

- What do we mean by a population?
- Catalogs of detected events
- Cosmological effects on Gravitational Waves

Suggested Reading

- Observation of Gravitational Waves from a Binary Black Hole Merger. Phys. Rev. Lett. 116, 061102 (2016).
- Properties of the Binary Black Hole Merger GW150914. Phys. Rev. Lett. 116, 241102 (2016).