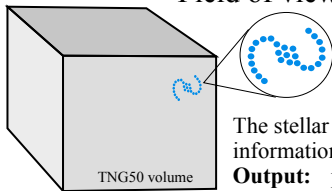


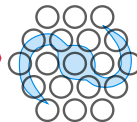
### Field of view definition (Sect. 3.1)



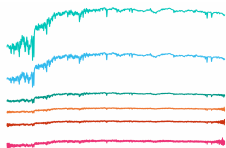
The stellar particles' and gas cells' information is saved for a given galaxy.

**Output:** particle file.

### From particles to light (Sect. 3.2)



The stellar particles and gas cells are projected on the sky and randomly shifted to mimic the atmospheric seeing (FWHM=1.43").



Each stellar particle is associated to a synthetic spectrum from the MaStar\_sLOG template (see Sánchez+ 2022) and linearly combined to form the fiber spectrum. A simple screen model accounts for dust attenuation (Ibarra-Medel+ 2019).

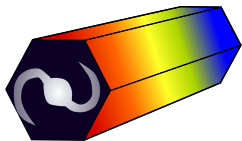
**Output:** RSS file.

### Control



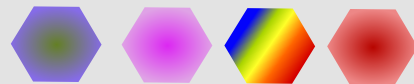
For control, the intrinsic values of the stellar particles in the FOV of the fibers are saved as an extension of the RSS file.

### Building MaNGA-like datacubes (Sect. 3.3)



Gaussian noise is added per fiber spectrum. The fiber spectra are recombined to form MaNGA's spatial grid.

**Output:** 3D datacube.



The values per fiber are recombined to form the maps with the intrinsic stellar particles' information.

**Output:** intrinsic and assigned 2D maps.

### From spectra to stellar population maps (Sect. 3.4)



The cubes are analyzed with the pyPIPE3D (Sánchez+ 2016 a, Sánchez+ 2016 b, Lacerda+ 2022), using the MaStar\_sLOG template.

**Output:** maps of age, metallicity and kinematics comparable to those derived from the MaNGA observed galaxies (Sánchez+ 2022).

× 10,000 simulated galaxies

