## Physics 305 Demo Notebook 4: Calculating the Displacement Probability Distribution (PDF) from a Time-Series - Exoplanet Light Curve

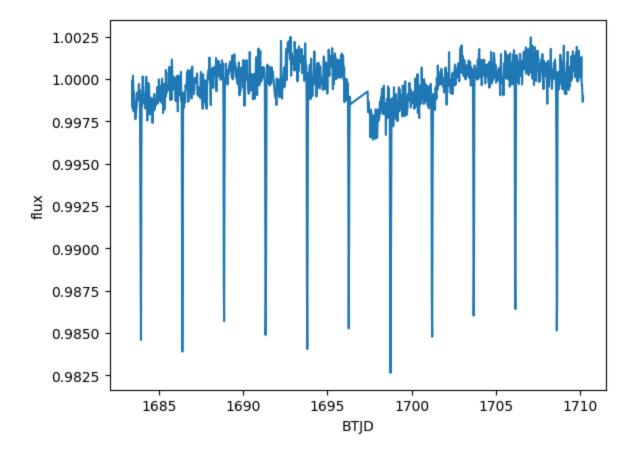
We apply the stochastic process analysis to the light curve dataset of Kepler 1-b observed by TESS.

First, we read the light curve data and plot the timeseries.

```
import numpy as np
In [2]:
         import pandas as pd
         import matplotlib.pyplot as plt
         %matplotlib inline
In [4]: fn = "tess lc kepler1b.txt"
         df = pd.read csv(fn, header=0)
         df.head(2)
Out[4]:
            # Time (BTJD) Normalized SAP_FLUX Normalized SAP_BKG
             1683.405757
                                                     0.008537
          0
                                   0.999869
             1683.426590
                                   0.998307
                                                      0.008400
In [5]: df.columns = ["t", "f", "e"]
         df.head(2)
Out[5]:
                                     е
          0 1683.405757 0.999869 0.008537
          1 1683.426590 0.998307 0.008400
In [6]:
         len(df)
Out[6]: 1171
```

```
In [7]: plt.plot(df['t'], df['f'])
   plt.xlabel("BTJD")
   plt.ylabel("flux")
```

```
Out[7]: Text(0, 0.5, 'flux')
```



```
In [8]: f thresh = 0.9875
         t transit = df.loc[df['f']<=f thresh, 't'].values
        t transit
 In [9]:
Out[9]: array([1683.92659508, 1686.38495208, 1688.86414229, 1691.32249884,
                1691.34333237, 1693.8016885 , 1696.2600432 , 1698.73924098,
                1698.76007459, 1701.19760673, 1701.21844034, 1703.67680558,
                1706.15600399, 1708.61436786, 1708.63520145])
In [10]: deltat transit vals = t transit[1:] - t transit[:-1]
         deltat transit vals
Out[10]: array([2.458357 , 2.47919021, 2.45835655, 0.02083353, 2.45835613,
                2.4583547 , 2.47919778, 0.02083361, 2.43753214, 0.02083361,
                2.45836525, 2.47919841, 2.45836387, 0.02083358])
In [32]: t orbit = np.mean(deltat transit vals[deltat transit vals > 2.0])
         t orbit
Out[32]: 2.462527204276239
```

The time between transits is around 2.46 days-- this is an estimate of the orbital period of Kepler 1-b.

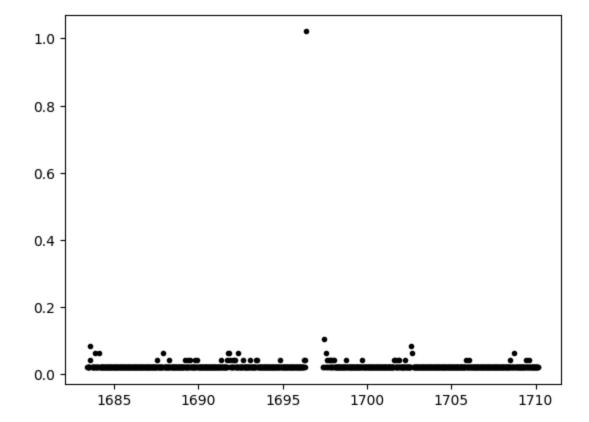
We note that the time interval between data points is not uniform-- and there is a gap in the dataset.

```
In [12]:
         deltat = df["t"].values[1:] - df["t"].values[:-1]
In [13]:
         pd.Series(deltat).describe()
Out[13]: count
                   1170.000000
         mean
                      0.022881
                      0.029862
         std
                      0.020834
         min
         25%
                      0.020834
         50%
                      0.020834
         75%
                      0.020834
         max
                      1.020845
         dtype: float64
         dt = np.mean(deltat)
In [14]:
         dt
Out[14]: 0.022881307263408524
```

The mean time interval is 0.022881 days or around 32.9 minutes.

```
In [15]: plt.plot(df.iloc[:-1]["t"].values, deltat, 'k.')
```

Out[15]: [<matplotlib.lines.Line2D at 0x7c7a8b0255a0>]



```
In [16]: deltat_thresh = 0.8
    t_gap = df.iloc[:-1].loc[deltat>deltat_thresh, "t"].values[0]

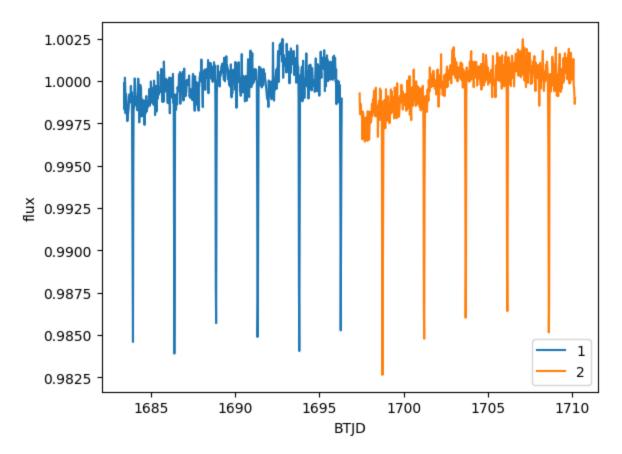
In [17]: t_gap
Out[17]: 1696.364210804781
```

To get around the gap in the dataset, let us split the dataset into two-- before and after the gap.

```
In [18]: df1 = df.loc[df["t"] < t_gap]
    df2 = df.loc[df["t"] > t_gap]
    n1, n2 = len(df1), len(df2)
    n1, n2, n1+n2, len(df)
Out[18]: (586, 584, 1170, 1171)
```

```
In [19]: plt.plot(df1['t'], df1['f'], label="1")
    plt.plot(df2['t'], df2['f'], label="2")
    plt.xlabel("BTJD")
    plt.ylabel("flux")
    plt.legend(loc="best")
```

Out[19]: <matplotlib.legend.Legend at 0x7c7a8ae8f4f0>



We perform interpolation to make the time variable uniform.

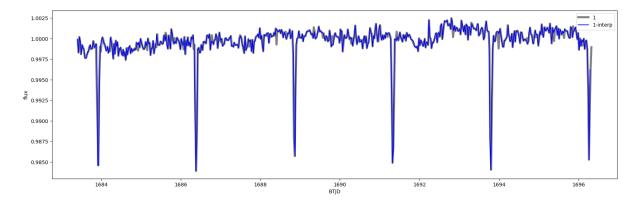
```
In [33]: import scipy.interpolate as interp
interp_method = "nearest"

n1_interp = np.floor((df1["t"].max() - df1["t"].min())/dt)
f1 = interp.interp1d(df1["t"], df1["f"], kind=interp_method)
tt1 = np.arange(n1_interp)*dt + df1["t"].min()
ff1 = f1(tt1)

n2_interp = np.floor((df2["t"].max() - df2["t"].min())/dt)
f2 = interp.interp1d(df2["t"], df2["f"], kind=interp_method)
tt2 = np.arange(n2_interp)*dt + df2["t"].min()
ff2 = f2(tt2)
```

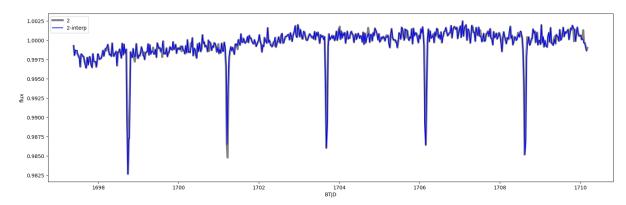
```
In [35]: plt.figure(figsize=(20,6))
  plt.plot(df1['t'], df1['f'], label="1", color='k', alpha=0.5, lw=4)
  plt.plot(tt1, ff1, color='b', label="1-interp")
  plt.xlabel("BTJD")
  plt.ylabel("flux")
  plt.legend(loc="best")
```

## Out[35]: <matplotlib.legend.Legend at 0x7c7a8443aa40>



```
In [36]: plt.figure(figsize=(20,6))
   plt.plot(df2['t'], df2['f'], label="2", color='k', alpha=0.5, lw=4)
   plt.plot(tt2, ff2, color='b', label="2-interp")
   plt.xlabel("BTJD")
   plt.ylabel("flux")
   plt.legend(loc="best")
```

## Out[36]: <matplotlib.legend.Legend at 0x7c7a81a5dfc0>

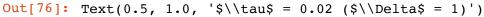


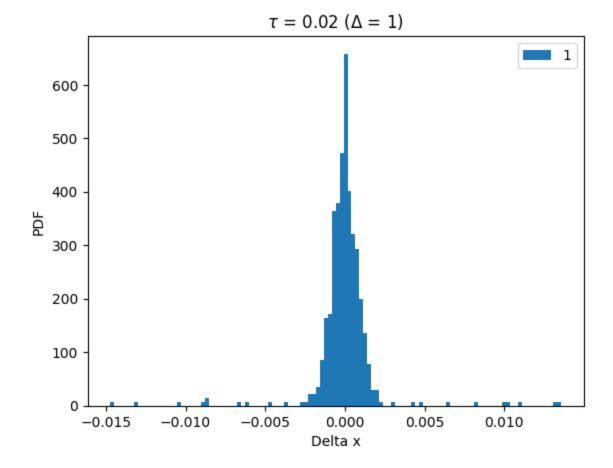
```
In [58]: f_mean, f_min = df['f'].mean(), df['f'].min()
  f_mean - f_min, f_mean, f_min
```

Out[58]: (0.01697816355152182, 0.9996260365540742, 0.9826478730025524)

Let us calculate and plot the PDF for time lag  $\tau = 1$  for the first time series.

```
In [76]: # Plot PDF of displacements
plt.hist(dx, density=True, bins="auto", label="1")
plt.xlabel("Delta x")
plt.ylabel("PDF")
plt.legend(loc="best")
plt.title(r"$\tau$ = %.2f ($\Delta$ = %d)" % (tau, delta))
```

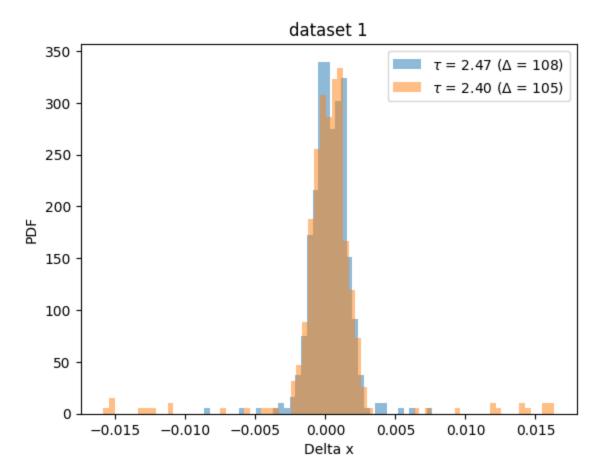




Let us calculate and plot the PDF for time lag corresponding to the orbital period  $\tau$  = 2.47 ( $\Delta$  = 108) and a smaller lag time  $\tau$  = 2.40 ( $\Delta$  = 105), for the first time series.

```
In [111]: | #tau = dt
          #delta = int(np.round(tau/dt))
          delta = 108 # corresponding to lag time equal to the orbital period
          tau = dt*delta
          x = ff1
          deltaA, tauA = delta, tau
          # get truncated copy of x, ending in initial data point of the last pair
          x trunc = x[:-1*delta]
          # get shifted copy of x, starting from end data point of the first pair
          x shift = x[delta:]
          # get displacements
          dx = x_shift - x_trunc
          dxA = dx
          delta = 105
          tau = dt*delta
          deltaB, tauB = delta, tau
          # get truncated copy of x, ending in initial data point of the last pair
          x trunc = x[:-1*delta]
          # get shifted copy of x, starting from end data point of the first pair
          x shift = x[delta:]
          # get displacements
          dx = x \text{ shift } - x \text{ trunc}
          dxB = dx
          # Plot PDF of displacements
          labelA = r"$\tau$ = %.2f ($\Delta$ = %d)" % (tauA, deltaA)
          labelB = r"\alpha = %.2f (\Delta = %d)" % (tauB, deltaB)
          plt.hist(dxA, density=True, bins="auto", alpha=0.5, label=labelA)
          plt.hist(dxB, density=True, bins="auto", alpha=0.5, label=labelB)
          plt.xlabel("Delta x")
          plt.ylabel("PDF")
          plt.legend(loc="best")
          plt.title("dataset 1")
```

Out[111]: Text(0.5, 1.0, 'dataset 1')



We notice that the outliers with large  $|\Delta x|$  are not in the PDF for  $\tau=2.47$  because the data points corresponding to the transit times (dips in the light curve) are paired with one another and result in a smaller dx-which is not the case for other values of lag times (that are not multiples of the orbital period).

Now, let us calculate and plot the PDF for different lag times  $\Delta = 1, 10, 100, 108, 216, 432$ , where we included values coinciding with the orbital period and multiples of the orbital period (2x and 3x).

```
In [122]: delta_orbit = np.floor(t_orbit/dt)
    delta_orbit, t_orbit

    delta_vals = np.array([1, 10, 1e2, 108, 216, 432]).astype(int)
    tau_vals = delta_vals*dt
    n_tau = len(tau_vals)

    print(delta_vals)
    print(tau_vals)
```

[ 1 10 100 108 216 432] [0.02288131 0.22881307 2.28813073 2.47118118 4.94236237 9.88472474]

```
In [124]: # initialize array to store displacements
           # note that the size n is larger than the actual number of values to be
           stored
           # values are initialized to NaN (which are not included in the PDF calcu
           lation)
          dx tau = np.empty((n, n tau, n samp))*np.nan
           for i samp in np.arange(n samp):
             for i, tau in enumerate(tau vals):
              delta = delta vals[i]
              # get truncated copy of x, ending in initial data point of the last
          pair
              x trunc = x samp[:-1*delta, i samp]
              # get shifted copy of x, starting from end data point of the first p
          air
              x shift = x samp[delta:, i samp]
              # get displacements
              dx = x \text{ shift } - x \text{ trunc}
              # store in output array
              dx tau[:len(dx), i, i samp] = dx
               #print(i, tau, delta, len(dx))
```

```
In [125]: # define plotting function
    def plot_pdf(dx, tau, label=""):

    # Plot PDF of displacements
    plt.hist(dx, density=True, bins="auto", alpha=0.4, label=label)
    plt.xlabel("Delta x")
    plt.ylabel("PDF")

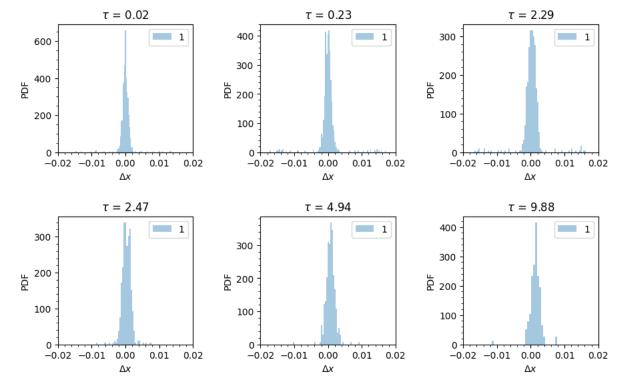
    plt.ylabel("PDF")

    plt.xlabel(r"$\Delta x$")
    #plt.legend(loc="best")
    plt.title(r"$\tau$ = %.2f" % tau)
```

```
In [126]: # generate PDF plots for different values of tau
i_samp = 0
nx = 2
ny = 3

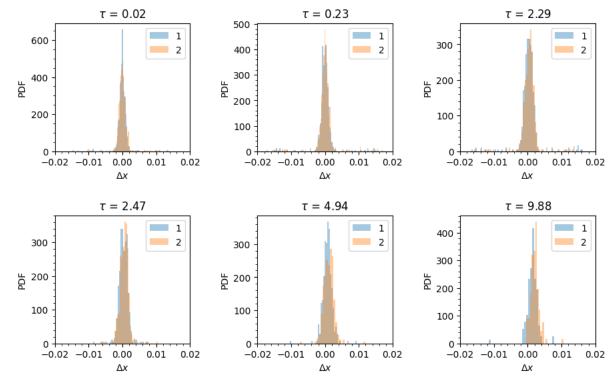
plt.figure(figsize=(12,6))
for i in np.arange(n_tau):
    plt.subplot(nx,ny,i+1)
    plot_pdf(dx_tau[:,i,i_samp], tau_vals[i], label="%d" % (i_samp+1))
    plt.xlim((-0.02, 0.02))
    plt.minorticks_on()
    plt.legend(loc="upper right")

plt.subplots_adjust(bottom=0.1, right=0.8, top=0.9, hspace=0.5, wspace=0.5)
```



```
In [128]: # generate plot showing PDF of both timeseries

plt.figure(figsize=(12,6))
for i in np.arange(n_tau):
   plt.subplot(nx,ny,i+1)
   for i_samp in np.arange(n_samp):
        plot_pdf(dx_tau[:,i,i_samp], tau_vals[i], label="%d" % (i_samp+1))
        plt.xlim((-0.02, 0.02))
        plt.minorticks_on()
        plt.legend(loc="upper right")
   plt.subplots_adjust(bottom=0.1, right=0.8, top=0.9, hspace=0.5, wspace=0.5)
```



Next, let us get the MSD curves.

```
In [105]: n_tau = 500
    delta_vals = np.arange(n_tau).astype(int)+int(1)
    #delta_vals = np.round(tau_vals/dt).astype(int)
    tau_vals = delta_vals*dt
    #tau_vals, delta_vals
```

```
In [101]: # initialize array to store displacements
           # note that the size n is larger than the actual number of values to be
          stored
           # values are initialized to NaN (which are not included in the MSD calcu
           lation)
          dx_tau = np.empty((n, n_tau, n_samp))*np.nan
           for i samp in np.arange(n samp):
             for i, tau in enumerate(tau vals):
               delta = delta vals[i]
               # get truncated copy of x, ending in initial data point of the last
          pair
               x_trunc = x_samp[:-1*delta, i_samp]
               # get shifted copy of x, starting from end data point of the first p
           air
               x_shift = x_samp[delta:, i_samp]
               # get displacements
               dx = x \text{ shift } - x \text{ trunc}
               # store in output array
               dx tau[:len(dx), i, i samp] = dx
               #print(i, tau, delta, len(dx))
```

```
In [102]: # initialize array to store MSD values
    msd_tau = np.empty((n_tau, n_samp))*np.nan

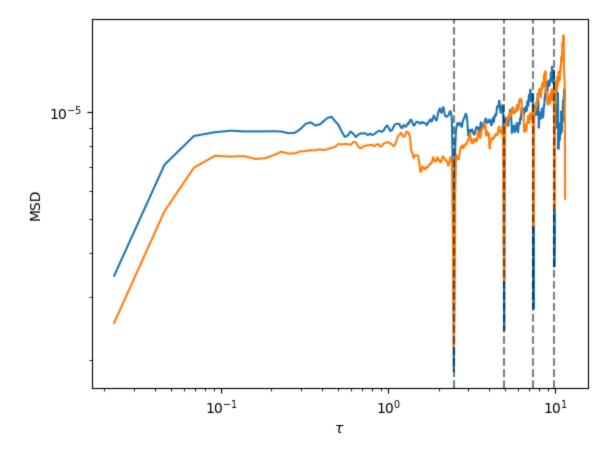
for i_samp in np.arange(n_samp):
    for i, tau in enumerate(tau_vals):
        dx2_sum = np.nansum(dx_tau[:, i, i_samp]**2) # returns sum treating
    NaNs as zero
        denom = n-delta_vals[i]
        msd_tau[i, i_samp] = dx2_sum/denom
```

```
In [103]: # plot MSD curve
    for i_samp in np.arange(n_samp):
        plt.plot(tau_vals, msd_tau[:, i_samp], label="%d" % (i_samp+1))

plt.axvline(t_orbit, color='k', alpha=0.5, ls="--")
    plt.axvline(t_orbit*2, color='k', alpha=0.5, ls="--")
    plt.axvline(t_orbit*3, color='k', alpha=0.5, ls="--")
    plt.axvline(t_orbit*4, color='k', alpha=0.5, ls="--")

plt.xscale("log")
    plt.yscale("log")
    plt.xlabel(r"$\tau$")
    plt.ylabel("MSD")
```

## Out[103]: Text(0, 0.5, 'MSD')



We find features in the MSD-- sharp drops-- occurring at lag times equal to the orbital period and its multiples (2x, 3x, and 4x).