

# Practical session I

## *From ABINIT to EXC*

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May 17, 2013 ■ 14h00m-18h00m

### Introduction

It is your duty now to prepare the ground-state and screening files in order to solve the Bethe-Salpeter equation using EXC.

The EXC code firstly creates the excitonic Hamiltonian

$$H_{exc}^{e-h} = [(\epsilon_c + \Delta_c - \epsilon_v + \Delta_v) \delta_{cc'vv'} + 2 \langle\langle v \rangle\rangle - \langle\langle W \rangle\rangle]$$

with  $\Delta_n$  being the GW correction (or a scissor operator shift)

$\langle\langle v \rangle\rangle = \int \phi_v(r) \phi_c(r) v(r, r') \phi_{v'}(r') \phi_{c'}(r')$  and  $\langle\langle W \rangle\rangle = \int \phi_v(r) \phi_{v'}(r) W(r, r') \phi_c(r') \phi_{c'}(r')$  where  $W(r, r') = \frac{\varepsilon^{-1}(r, r')}{|r - r'|}$ . So EXC needs a KSS file (for  $\epsilon_n$  and  $\phi_n$ ), a GW file (for  $\Delta_n$ ) or simply a scissor shift, and a screening file SCR (for the  $\varepsilon^{-1}$ ).

### Tasks

1. Create the “\*\_KSS” file containing complete information on the Kohn-Sham bandstructure (use the usual Monkhorst-Pack 4x444 k-grid).
2. Create the “\*\_SCR” file containing the RPA screening function (this should contain 19 q-points).
3. Create the “\*\_GW” file containing the GW corrections to the DFT eigenvalues.
4. Create now another “\*\_KSS” file, using these shifts in the abinit input file:

```
nshiftk 4
shiftk
0.6 0.7 0.8
0.6 0.2 0.3
0.1 0.7 0.3
0.1 0.2 0.8
```

Q1. What is the difference between the two KSS files ? (hint: look at the symmetries)

Q2. How many k-points are contained in the two KSS files ?

### Comments

## Practical session II

### *EXC*

LECTURERS: FRANCESCO SOTTILE, MATTEO GATTI

May 17, 2013 ■ 14h00m-18h00m

### Introduction

We are ready to use the Bethe-Salpeter equation to calculate the absorption spectra of solids.

### Objectives

- Calculate the different kind of spectra (RPA, with and without local fields, GW-RPA, BSE)
- Study the convergence of the spectrum with respect to the number of bands, the number of plane waves in the wavefunctions, the dimension of the dielectric matrix, the set of **k**-points.

In order to run an exciton calculation, use the same input file as for DP, with the additional variable **exciton**. Again, in the transition framework, you can do **rpa**, **alda** or now **exc**.

### Tasks

- Run a RPA calculation, with the EXC code, with and without local fields, and answer the following question.
  - Q1. Is there a difference between a RPA calculation with EXC and DP ? Why?
  - Q2. What is the difference among the files **outexc.mdf** **outrpanlf.mdf** **outgwnlf.mdf**?
- Perform now a GW-RPA calculation (using a scissor for example). Plot the spectra and consider again the three files **outexc.mdf** **outrpanlf.mdf** **outgwnlf.mdf**.
- Finally perform an exciton calculation (still using a scissor if you wish). Plot the spectra and consider again the three files **outexc.mdf** **outrpanlf.mdf** **outgwnlf.mdf**.
- Perform a calculation including coupling (i.e. using the variable **coupling**).
  - Q1. What is the effect of the coupling on the absorption spectrum ?
  - Q2. And what the effect on the eels (hint: generate smaller kss and scr files for the evaluation of the eels spectra, for the convergence parameters might imply long calculations)? [If you have no time to generate new kss files, use the ones included in the **/nfs\_home/tutoadmin/spectroscopy\_lectures/work/Si** directory, but prior to calculation export the variable **F\_UFMTENDIAN** to big].

In the directories **~/spectroscopy\_lectures/work/** “other material” input files for other simple semiconductors and insulators are available. Create the kss and scr with Abinit and try to obtain the absorption spectra (for example, try LiF).

### Comments