Solution of Layer Equations

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I. DERIVATION OF LAYER EQUATIONS

A. Outer Region

Let r, θ , φ be right-handed cylindrical coordinates. Throughout the bulk of the plasma, the perturbed magnetic field can be written 1,2

$$\delta B_r = i \frac{m}{r} \delta \psi, \tag{1}$$

$$\delta B_{\theta} = -\frac{\partial \delta \psi}{\partial r},\tag{2}$$

$$\delta B_{\varphi} \simeq 0,$$
 (3)

where

$$\delta\psi(r,\zeta,t) = \psi_1(r,t) e^{i\zeta}, \tag{4}$$

and $\zeta = m \theta - n \varphi$. Here, m and n are poloidal and toroidal mode numbers, respectively.

B. Inner Region

The rational surface lies at radius r_s , where $q(r_s) = m/n$, and q(r) is the safety-factor profile. The inner region is a thin (in r) layer centered on the rational surface. The inner region is assumed to be governed by the four-field model described in Ref. 3. This model is an extension of a model introduced in Ref. 4, and is ultimately based on the four-field model of Hazeltine, et alia.⁵

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C. Useful Definitions

The plasma is assumed to consist of two species. First, electrons of mass m_e , electrical charge -e, number density n_e , and temperature T_e . Second, ions of mass m_i , electrical charge +e, number density n_e , and temperature T_i . Let $p = n_e (T_e + T_i)$ be the total plasma pressure.

It is helpful to define $n_0 = n_e(r_s)$, $p_0 = p(r_s)$,

$$\eta_e = \left. \frac{d \ln T_e}{d \ln n_e} \right|_{r=r_s},\tag{5}$$

$$\eta_i = \left. \frac{d \ln T_i}{d \ln n_e} \right|_{r=r_-},\tag{6}$$

$$\iota = \left(\frac{T_e}{T_i}\right)_{r=r_s} \left(\frac{1+\eta_e}{1+\eta_i}\right),\tag{7}$$

where $n_e(r)$, p(r), $T_e(r)$, and $T_i(r)$ refer to density, pressure, and temperature profiles that are unperturbed by the magnetic perturbation. Note that ι is the ratio of the electron to the ion pressure gradient at the rational surface.

For the sake of simplicity, the perturbed electron and ion temperature profiles are assumed to be functions of the perturbed electron number density profile in the immediate vicinity of the rational surface. In other words, $T_e = T_e(n_e)$ and $T_i = T_i(n_e)$. This implies that $p = p(n_e)$. The MHD velocity, which is the velocity of a fictional MHD fluid, is defined as $\mathbf{V} = \mathbf{V}_E + V_{\parallel i} \mathbf{b}$, where \mathbf{V}_E is the E-cross-B drift velocity, $V_{\parallel i}$ is the parallel component of the ion fluid velocity, and $\mathbf{b} = \mathbf{B}/|\mathbf{B}|$. Here, $\mathbf{B} = [0, B_{\theta}(r), B_{\varphi}]$ is the equilibrium magnetic field. Note that $q(r) = r B_{\varphi}/[R_0 B_{\theta}(r)]$, where R_0 is the plasma minor radius.

D. Dimensionless Fields

The four dimensionless fields in the four-field model—namely, ψ , ϕ , N, and V—have the following definitions:

$$\nabla \psi = \frac{\mathbf{n} \times \mathbf{B}}{r_s B_{\varphi}},\tag{8}$$

$$\nabla \phi = \frac{\mathbf{n} \times \mathbf{V}}{r_s V_A},\tag{9}$$

$$N = -\hat{d}_i \left(\frac{p - p_0}{B_{\varphi}^2 / \mu_0} \right), \tag{10}$$

$$V = \hat{d}_i \left(\frac{\mathbf{n} \cdot \mathbf{V}}{V_A} \right). \tag{11}$$

Here, $\mathbf{n} = (0, \epsilon/q_s, 1)$, $\epsilon = r/R_0$, $q_s = m/n$, $V_A = B_{\varphi}/\sqrt{\mu_0 n_0 m_i}$ is the Alfvén speed, $d_i = \sqrt{m_i/(n_0 e^2 \mu_0)}$ is the collisionless ion skin-depth, and $\hat{d}_i = d_i/r_s$. Our model also employs the auxiliary dimensionless field

$$J = -\frac{2\epsilon_s}{q_s} + \hat{\nabla}^2 \psi, \tag{12}$$

where $\epsilon_s = r_s/R_0$, and $\hat{\nabla} = r_s \nabla$. Note that

$$\psi(r,\zeta,t) = \frac{\psi_0(r) + \psi_1(r,t) e^{i\zeta}}{r_s B_{\varphi}},$$
(13)

where

$$\psi_0(r) = \frac{B_{\varphi}}{R_0} \int_{r_*}^r r' \left[\frac{n}{m} - \frac{1}{q(r')} \right] dr', \tag{14}$$

and $\psi_1(r,t)$ is defined in Eqs. (1)–(4).

E. Four-Field Model

Our four-field model takes the form:^{2–4,6}

$$\frac{\partial \psi}{\partial \hat{t}} = [\phi, \psi] - \iota_e [N, \psi] + \hat{\eta}_{\parallel} J + \hat{E}_{\parallel}, \tag{15}$$

$$\frac{\partial \hat{\nabla}^2 \phi}{\partial \hat{t}} = [\phi, \hat{\nabla}^2 \phi] - \frac{\iota_i}{2} \left(\hat{\nabla}^2 [\phi, N] + [\hat{\nabla}^2 \phi, N] + [\hat{\nabla}^2 N, \phi] \right) + [J, \psi]$$

$$+\hat{\chi}_{\varphi}\hat{\nabla}^{4}(\phi + \iota_{i} N), \qquad (16)$$

$$\frac{\partial N}{\partial \hat{t}} = [\phi, N] + c_{\beta}^2 [V, \psi] + \hat{d}_{\beta}^2 [J, \psi] + \hat{\chi}_{\perp} \hat{\nabla}^2 N, \tag{17}$$

$$\frac{\partial V}{\partial \hat{t}} = [\phi, V] + [N, \psi] + \hat{\chi}_{\varphi} \hat{\nabla}^2 V. \tag{18}$$

Here, $[A, B] \equiv \hat{\nabla} A \times \hat{\nabla} B \cdot \mathbf{n}$, $\iota_e = \iota/(1+\iota)$, $\iota_i = 1/(1+\iota)$, $\hat{t} = t/(r_s/V_A)$, $\hat{\eta}_{\parallel} = \eta_{\parallel}/(\mu_0 r_s V_A)$, $\hat{E}_{\parallel} = E_{\parallel}/(B_{\varphi} V_A)$, $\hat{\chi}_{\varphi} = \chi_{\varphi}/(r_s V_A)$, $\hat{\chi}_{\perp} = \chi_{\perp}/(r_s V_A)$, where η_{\parallel} is the parallel plasma electrical resistivity at the rational surface, E_{\parallel} the parallel inductive electric field that maintains the equilibrium toroidal plasma current in the vicinity of the rational surface, χ_{φ} the anomalous perpendicular ion momentum diffusivity at the rational surface, and χ_{\perp} the anomalous perpendicular energy diffusivity at the rational surface. Moreover, $d_{\beta} = c_{\beta} d_i$, and $\hat{d}_{\beta} = d_{\beta}/r_s$, where $c_{\beta} = \sqrt{\beta/(1+\beta)}$, and $\beta = (5/3) \mu_0 p_0/B_{\varphi}^2$. Here, d_{β} is usually referred to as the ion

sound radius. Note that we have neglected parallel energy transport in Eq. (17) because we expect the radial width of the inner region to be less than the critical width above which parallel transport forces the temperature and density profiles to be magnetic flux-surface functions.⁷

F. Matching to Plasma Equilibrium

The unperturbed plasma equilibrium is such that $\mathbf{B} = [0, B_{\theta}(r), B_{\varphi}], p = p(r), \mathbf{V} = [0, V_E(r), V_{\varphi}(r)],$ where $V_E(r) \simeq E_r/B_{\varphi}$ is the (dominant θ -component of the) E-cross-B drift velocity. Now, the inner region is assumed to have a radial thickness that is much smaller than r_s . Hence, we only need to evaluate plasma equilibrium quantities in the immediate vicinity of the rational surface. Equations (8)–(11) and (13)–(14) suggest that

$$\psi(\hat{x}) = \frac{\hat{x}^2}{2\hat{L}_s},\tag{19}$$

$$\phi(\hat{x}) = -\hat{V}_E \,\hat{x},\tag{20}$$

$$N(\hat{x}) = -\hat{V}_* \,\hat{x},\tag{21}$$

$$V = \hat{V}_{\parallel}, \tag{22}$$

where $\hat{L}_s = L_s/r_s$, $L_s = R_0 \, q_s/s_s$, $\hat{V}_E = V_E(r_s)/V_A$, $\hat{V}_* = V_*(r_s)/V_A$, $V_*(r) = (dp/dr)/(e \, n_0 \, B_\varphi)$ is the (dominant θ -component of the) diamagnetic velocity, and $\hat{V}_{\parallel} = \hat{d}_i \, V_\varphi(r_s)/V_A$. Here, $s_s = s(r_s)$, and $s(r) = d \ln q / \ln r$ is the magnetic shear. We also deduce that

$$J = -\left(\frac{2}{s_s} - 1\right) \frac{1}{\hat{L}_s},\tag{23}$$

and $\hat{E}_{\parallel} = (2/s_s - 1)(\hat{\eta}_{\parallel}/\hat{L}_s)$.

G. Linearized Layer Equations

In accordance with Eqs. (12), (13), and (19)–(23), let us write

$$\psi(\hat{x},\zeta,\hat{t}) = \frac{\hat{x}^2}{2\hat{L}_s} + \tilde{\psi}(\hat{x},\hat{t}) e^{i\zeta}, \qquad (24)$$

$$\phi(\hat{x},\zeta,\hat{t}) = -\hat{V}_E \hat{x} + \tilde{\phi}(\hat{x},\hat{t}) e^{i\zeta}, \qquad (25)$$

$$N(\hat{x}, \zeta, \hat{t}) = -\hat{V}_* \,\hat{x} + \iota_e \,\tilde{N}(\hat{x}, \hat{t}) \,\mathrm{e}^{\mathrm{i}\,\zeta},\tag{26}$$

$$V(\hat{x},\zeta,\hat{t}) = \hat{V}_{\parallel} + \iota_e \tilde{V}(\hat{x},\hat{t}) e^{i\zeta}, \tag{27}$$

$$J(\hat{x},\zeta,\hat{t}) = -\left(\frac{2}{s_s} - 1\right)\frac{1}{\hat{L}_s} + \hat{\nabla}^2 \tilde{\psi}(\hat{x},\hat{t}) e^{i\zeta}.$$
 (28)

Substituting Eqs. (24)–(28) into Eqs. (15)–(18), and only retaining terms that are first order in perturbed quantities, we obtain the following set of linearized layer equations:

$$\tau_H \frac{\partial \tilde{\psi}}{\partial t} + i \left(\omega_E + \omega_{*e}\right) \tau_H \tilde{\psi} = -i \hat{x} \left(\tilde{\phi} - \tilde{N}\right) + S^{-1} \hat{\nabla}^2 \tilde{\psi}, \tag{29}$$

$$\tau_H \frac{\partial \hat{\nabla}^2 \tilde{\phi}}{\partial t} + i \left(\omega_E + \omega_{*i} \right) \tau_H \hat{\nabla}^2 \tilde{\phi} = -i \hat{x} \hat{\nabla}^2 \tilde{\psi} + S^{-1} P_{\varphi} \hat{\nabla}^4 \left(\tilde{\phi} + \frac{\tilde{N}}{\iota} \right), \tag{30}$$

$$\tau_H \frac{\partial \tilde{N}}{\partial t} + i \omega_E \tau_H \, \tilde{N} = -i \omega_{*e} \, \tau_H \, \tilde{\phi} - i \, c_\beta^2 \, \hat{x} \, \tilde{V} - i \, \iota_e \, \hat{d}_\beta^2 \, \hat{x} \, \hat{\nabla}^2 \tilde{\psi}$$

$$+ S^{-1} P_{\perp} \hat{\nabla}^2 \tilde{N}, \tag{31}$$

$$\tau_H \frac{\partial \tilde{V}}{\partial t} + i \omega_E \tau_H \tilde{V} = i \omega_{*e} \tau_H \tilde{\psi} - i \hat{x} \tilde{N} + S^{-1} P_{\varphi} \hat{\nabla}^2 \tilde{V}.$$
 (32)

Here, $\tau_H = L_s/(m\,V_A)$ is the hydromagnetic time, $\omega_E = (m/r_s)\,V_E(r_s)$ the E-cross-B frequency, $\omega_{*\,e} = -\iota_e\,(m/r_s)\,V_*(r_s)$ the electron diamagnetic frequency, $\omega_{*\,i} = \iota_i\,(m/r_s)\,V_*(r_s)$ the ion diamagnetic frequency, $S = \tau_R/\tau_H$ the Lundquist number, $\tau_R = \mu_0\,r_s^2/\eta_{\parallel}$ the resistive diffusion time, $\tau_{\varphi} = r_s^2/\chi_{\varphi}$ the toroidal momentum confinement time, and $\tau_{\perp} = r_s^2/\chi_{\perp}$ the energy confinement time. Furthermore, $P_{\varphi} = \tau_R/\tau_{\varphi}$ and $P_{\perp} = \tau_R/\tau_{\perp}$ are magnetic Prandtl numbers. Note that $\iota = -\omega_{*\,e}/\omega_{*\,i}$ and $\iota_e = \omega_{*\,e}/(\omega_{*\,e} - \omega_{*\,i})$.

In the following, in accordance with the analysis of Refs. 4 and 3, we shall neglect the term $-i c_{\beta}^2 \hat{x} \tilde{V}$ in Eq. (31). This approximation can be justified a posteriori. It is equivalent to the neglect of ion parallel dynamics in the inner region, and effectively converts our four-field model into the following three-field model:

$$\tau_H \frac{\partial \tilde{\psi}}{\partial t} + i \left(\omega_E + \omega_{*e}\right) \tau_H \tilde{\psi} = -i \hat{x} \left(\tilde{\phi} - \tilde{N}\right) + S^{-1} \hat{\nabla}^2 \tilde{\psi}, \tag{33}$$

$$\tau_H \frac{\partial \hat{\nabla}^2 \tilde{\phi}}{\partial t} + i \left(\omega_E + \omega_{*i} \right) \tau_H \hat{\nabla}^2 \tilde{\phi} = -i \hat{x} \hat{\nabla}^2 \tilde{\psi} + S^{-1} P_{\varphi} \hat{\nabla}^4 \left(\tilde{\phi} + \frac{\tilde{N}}{\iota} \right), \tag{34}$$

$$\tau_H \frac{\partial \tilde{N}}{\partial t} + i \omega_E \tau_H \tilde{N} = -i \omega_{*e} \tau_H \tilde{\phi} - i \iota_e \hat{d}_\beta^2 \hat{x} \hat{\nabla}^2 \tilde{\psi} + S^{-1} P_\perp \hat{\nabla}^2 \tilde{N}.$$
 (35)

Finally, if we write

$$\frac{\partial}{\partial t} = g - i\,\omega_E,\tag{36}$$

where g is the complex growth-rate in a frame of reference that co-rotates with the MHD fluid at the rational surface, then we get

$$(g + i \omega_{*e}) \tau_H \tilde{\psi} = -i \hat{x} (\tilde{\phi} - \tilde{N}) + S^{-1} \hat{\nabla}^2 \tilde{\psi}, \tag{37}$$

$$(g + i \omega_{*i}) \tau_H \hat{\nabla}^2 \tilde{\phi} = -i \hat{x} \hat{\nabla}^2 \tilde{\psi} + S^{-1} P_{\varphi} \hat{\nabla}^4 \left(\tilde{\phi} + \frac{\tilde{N}}{\iota} \right), \tag{38}$$

$$g \tau_H \tilde{N} = -i \omega_{*e} \tau_H \tilde{\phi} - i \iota_e \hat{d}_\beta^2 \hat{x} \hat{\nabla}^2 \tilde{\psi} + S^{-1} P_\perp \hat{\nabla}^2 \tilde{N}.$$
 (39)

H. Asymptotic Matching

The layer equations, (37)–(39) possess a trivial twisting parity solution, and a non-trivial tearing parity solution.⁸ The latter solution is such that $\tilde{\psi}(-\hat{x}) = \tilde{\psi}(\hat{x})$, $\tilde{\phi}(-\hat{x}) = -\tilde{\phi}(\hat{x})$, and $\tilde{N}(-\hat{x}) = -\tilde{N}(\hat{x})$. When this solution is asymptotically matched to the outer solution, we find that

$$\tilde{\psi}(\hat{x}) \to \left(\frac{\Psi_s}{r_s B_{\varphi}}\right) \left(1 + \Delta_s \frac{|\hat{x}|}{2}\right)$$
 (40)

at the interface between the inner and the outer regions. Here, Ψ_s is the (complex) asymptotic reconnected helical magnetic flux at the rational surface, whereas Δ_s is the (real, dimensionless) tearing stability index.

II. FOURIER TRANSFORMED LAYER EQUATIONS

A. Stretched Radial Coordinate

Let us define the stretched radial coordinate $X = S^{1/3} \hat{x}$. Assuming that $X \sim \mathcal{O}(1)$ in the resonant layer that constitutes the inner region (i.e., assuming that the thickness of the layer is roughly of order $S^{-1/3} r_s$), and making use of the fact that $S \gg 1$ in a conventional tokamak plasma, Eqs. (37)–(39) reduce to the following set of linear layer equations:

$$(\hat{g} + i Q_e) \tilde{\psi} = -i X \left(\tilde{\phi} - \tilde{N} \right) + \frac{d^2 \tilde{\psi}}{dX^2}, \tag{41}$$

$$(\hat{g} + iQ_i)\frac{d^2\tilde{\phi}}{dX^2} = -iX\frac{d^2\tilde{\psi}}{dX^2} + P_{\varphi}\frac{d^4}{dX^4}\left(\tilde{\phi} + \frac{\tilde{N}}{\iota}\right),\tag{42}$$

$$\hat{g}\,\tilde{N} = -\mathrm{i}\,Q_e\,\tilde{\phi} - \mathrm{i}\,D^2\,X\,\frac{d^2\tilde{\psi}}{dX^2} + P_\perp\,\frac{d^2\tilde{N}}{dX^2},\tag{43}$$

where

$$\hat{g} = S^{1/3} g \tau_H, \tag{44}$$

$$Q_{e,i} = S^{1/3} \,\omega_{*e,i} \,\tau_H,\tag{45}$$

$$D = S^{1/3} \iota_e^{1/2} \hat{d}_{\beta}. \tag{46}$$

According to Eq. (40), the asymptotic behavior of the tearing-parity solutions of Eqs. (41)–(43) is such that

$$\tilde{\psi}(X) \to \left(\frac{\Psi_s}{r_s B_{\varphi}}\right) \left[1 + \frac{\hat{\Delta}_s}{2} |X| + \mathcal{O}(X^2)\right]$$
 (47)

as $|X| \to \infty$, where

$$\hat{\Delta}_s = S^{-1/3} \, \Delta_s. \tag{48}$$

It follows from Eqs. (41)–(43) that

$$\bar{\phi}(X) \to i \,\hat{g}\left(\frac{\Psi_s}{r_s \, B_{\varphi}}\right) \left[\frac{1}{X} + \frac{\hat{\Delta}_s}{2} \operatorname{sgn}(X) + \mathcal{O}(X)\right]$$
 (49)

as $|X| \to \infty$.

B. Fourier Transformation

Equations (41)–(43) are most conveniently solved in Fourier transform space.⁴ Let

$$\hat{\phi}(p) = \int_{-\infty}^{\infty} \tilde{\phi}(X) e^{-i p X} dX, \qquad (50)$$

et cetera. The Fourier transformed linear layer equations become

$$(\hat{g} + i Q_e) \hat{\psi} = \frac{d}{dp} (\hat{\phi} - \hat{N}) - p^2 \hat{\psi}, \tag{51}$$

$$(\hat{g} + i Q_i) p^2 \hat{\phi} = \frac{d(p^2 \hat{\psi})}{dp} - P_{\varphi} p^4 \left(\hat{\phi} + \frac{\hat{N}}{\iota}\right), \tag{52}$$

$$\hat{g}\,\hat{N} = -\mathrm{i}\,Q_e\,\hat{\phi} - D^2\,\frac{d(p^2\,\hat{\psi})}{dp} - P_\perp\,p^2\hat{N},\tag{53}$$

where, for a tearing parity solution, Eq. (49) yields 9

$$\hat{\phi}(p) \to = \pi \,\hat{g} \left(\frac{\Psi_s}{r_s \, B_{\varphi}}\right) \left[\frac{\hat{\Delta}_s}{\pi \, p} + 1 + \mathcal{O}(p)\right].$$
 (54)

as $p \to 0$.

Let

$$Y_e(p) \equiv \hat{\phi}(p) - \hat{N}(p) = \pi \left(\hat{g} + i Q_e\right) \left(\frac{\Psi_s}{r_s B_{\varphi}}\right) \hat{Y}_e(p). \tag{55}$$

Equations (51)–(53) can be combined to give

$$\frac{d}{dp}\left[A(p)\frac{d\hat{Y}_e}{dp}\right] - \frac{B(p)}{C(p)}p^2\hat{Y}_e = 0,$$
(56)

where

$$A(p) = \frac{p^2}{\hat{g} + i Q_e + p^2},\tag{57}$$

$$B(p) = \hat{g} (\hat{g} + i Q_i) + (\hat{g} + i Q_i) (P_{\varphi} + P_{\perp}) p^2 + P_{\varphi} P_{\perp} p^4,$$
 (58)

$$C(p) = \hat{g} + i Q_e + [P_{\perp} + (\hat{g} + i Q_i) D^2] p^2 + \iota_e^{-1} P_{\varphi} D^2 p^4.$$
 (59)

Note that $\iota_e = Q_e/(Q_e - Q_i)$. Furthermore, because

$$\hat{Y}_e(p) = \left(\frac{\hat{g} + i Q_e}{\hat{g}}\right) \hat{\phi}(p) \tag{60}$$

as $p \to 0$, Eqs. (54) and (55) yield the following small-p boundary condition that Eq. (56) must satisfy: ⁹

$$\hat{Y}_e(p) \to \frac{\hat{\Delta}_s}{\pi p} + 1 + \mathcal{O}(p)$$
 (61)

as $p \to 0$. Equation (56) must also satisfy the physical boundary condition

$$\hat{Y}_e(p) \to 0 \tag{62}$$

as $p \to \infty$.

III. TECHNICAL DETAILS

A. Large-p Limit

In the large-p limit, if we write

$$A = 1 + \frac{\alpha}{p^2},\tag{63}$$

$$\frac{B}{C} = \beta + \frac{\gamma}{p^2},\tag{64}$$

and look for a solution of Eq. (56) of the form

$$\hat{Y}_e(p) \propto p^x \exp\left(\frac{-\sqrt{\beta} p^2}{2}\right)$$
 (65)

then we find that

$$x = \frac{\gamma - \sqrt{\beta} \left(1 - \sqrt{\beta} \,\alpha\right)}{2\sqrt{\beta}}.\tag{66}$$

It is easily seen that

$$\alpha = -(\hat{g} + i Q_e), \tag{67}$$

$$\beta = \frac{P_{\varphi} P_{\perp}}{\iota_e^{-1} P_{\varphi} D^2},\tag{68}$$

$$\gamma = \frac{P_{\varphi} P_{\perp}}{\iota_e^{-1} P_{\varphi} D^2} \left[1 + (\hat{g} + i Q_i) \frac{P_{\varphi} + P_{\perp}}{P_{\varphi} P_{\perp}} - (P_{\perp} + [\hat{g} + i Q_i] D^2) \frac{1}{\iota_e^{-1} P_{\varphi} D^2} \right].$$
(69)

In order to be in the large-p limit, we require

$$p \gg |\hat{g} + i Q_e|^{1/2},\tag{70}$$

$$p \gg \left| \frac{(\hat{g} + i Q_i) (P_{\varphi} + P_{\perp})}{P_{\varphi} P_{\perp}} \right|^{1/2}, \tag{71}$$

$$p \gg \left| \frac{\hat{g} \left(\hat{g} + i Q_i \right)}{P_{\omega} P_{\perp}} \right|^{1/4}, \tag{72}$$

$$p \gg \left| \frac{P_{\perp} + (\hat{g} + i Q_i) D^2}{\iota_e^{-1} P_{\varphi} D^2} \right|^{1/2}, \tag{73}$$

$$p \gg \left| \frac{\hat{g} + i Q_e}{\iota_e^{-1} P_{\varphi} D^2} \right|^{1/4}, \tag{74}$$

$$p \gg \left(\frac{\iota_e^{-1} P_{\varphi} D^2}{P_{\varphi} P_{\perp}}\right)^{1/4}. \tag{75}$$

B. Low-D Limit

If

$$D^2 \ll \left| \frac{P_\perp}{\hat{g} + i Q_e} \right|, \quad \frac{\iota_e P_\perp}{P_\wp^{2/3}} \tag{76}$$

then the terms in the layer equations involving D^2 are negligible. In this case, in the large-p limit, we can write

$$A = 1 + \frac{\alpha}{p^2},\tag{77}$$

$$\frac{B}{C} = \beta \, p^2 + \gamma,\tag{78}$$

where

$$\alpha = -(\hat{g} + i Q_e), \tag{79}$$

$$\beta = P_{\varphi},\tag{80}$$

$$\gamma = -\mathrm{i} \left(Q_e - Q_i \right) \frac{P_{\varphi}}{P_{\perp}} + \hat{g} + \mathrm{i} \, Q_i. \tag{81}$$

The solution of Eq. (56) becomes

$$\hat{Y}_e \propto p^{-1} \exp\left(x \, p - \frac{\sqrt{\beta} \, p^3}{3}\right),\tag{82}$$

where

$$x = \frac{\alpha \beta - \gamma}{2\sqrt{\beta}}. (83)$$

In order to be in the large-p limit, we require

$$p \gg |\hat{g} + i Q_e|^{1/2}, \tag{84}$$

$$p \gg \left| \frac{(\hat{g} + i Q_i) (P_{\varphi} + P_{\perp})}{P_{\varphi} P_{\perp}} \right|^{1/2}, \tag{85}$$

$$p \gg \left| \frac{\hat{g} \left(\hat{g} + i Q_i \right)}{P_{\varphi} P_{\perp}} \right|^{1/4}, \tag{86}$$

$$p \gg \left| \frac{\hat{g} + i Q_e}{P_\perp} \right|^{1/2}, \tag{87}$$

$$p \gg P_{\varphi}^{-1/6}. \tag{88}$$

C. Ricatti Transformation

Let

$$W = \frac{p}{\hat{Y}_e} \frac{d\hat{Y}_e}{dp}.$$
 (89)

Equation (56) transforms to give

$$\frac{dW}{dp} = -\frac{A'}{p}W - \frac{W^2}{p} + \frac{B}{AC}p^3,$$
(90)

where

$$A' = \frac{\hat{g} + i Q_e - p^2}{\hat{g} + i Q_e + p^2}.$$
 (91)

According to Sects. IIB and IIIA, this equation must be solved subject to the boundary condition that

$$W(p) = x - \sqrt{\beta} \, p^2 \tag{92}$$

at large-p, and

$$W(p) = -1 + \frac{\pi p}{\hat{\Delta}_s} \tag{93}$$

at small-p. However, according to Sect. IIIB, in the low-D limit,

$$W(p) = -1 + x p - \sqrt{\beta} p^3 \tag{94}$$

at large-p.

D. Method of Solution

We launch a solution of Eqs (90) from large p, subject to the boundary condition (92) or (94), as appropriate, and integrate it to small p. We can deduce the value of $\hat{\Delta}_s$ from Eq. (93).

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DATA AVAILABILITY STATEMENT

The digital data used in the figures in this paper can be obtained from the author upon reasonable request.

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