

Take-home Midterm

Due date: March 22, 2024

The objective of this take-home midterm is to write a finite-difference solver for the inviscid flow in a 1D duct. This flow is governed by the 1D Euler equations, which are given in their conservative form as

$$\frac{\partial \rho}{\partial t} + \frac{\partial \rho u}{\partial x} = 0 \quad (1)$$

$$\frac{\partial \rho u}{\partial t} + \frac{\partial}{\partial x}(\rho u^2 + p) = 0 \quad (2)$$

$$\frac{\partial \rho E}{\partial t} + \frac{\partial}{\partial x}[(\rho E + p)u] = 0 \quad (3)$$

where

- ρ is the fluid density,
- u is the fluid velocity,
- p is the thermodynamic pressure,
- E is the total specific energy.

The pressure is related to the other quantities by the equation of state

$$p = (\gamma - 1) \left(\rho E - \frac{1}{2} \rho u^2 \right) \quad (4)$$

for a perfect gas with ratio of specific heats γ .

The Euler equations (1)–(3) constitute a system of coupled nonlinear partial differential equations, which is typically very difficult to solve analytically. Instead, in this midterm, we will solve this system of equations numerically using finite differences and the method of lines on a periodic domain.

Problem 1 · The Euler equations (1)–(3) are written in terms of the conserved variables that are: the mass density ρ (mass per unit volume), the momentum density ρu (momentum per unit volume), and the total energy density ρE (total energy per unit volume). Let us assign these conserved variables to a single vector

$$\mathbf{q} = \begin{bmatrix} q_0 \\ q_1 \\ q_2 \end{bmatrix} = \begin{bmatrix} \rho \\ \rho u \\ \rho E \end{bmatrix} \quad (5)$$

Q1 → Using the equation of state (4), express the pressure as a function of the components of \mathbf{q} , i.e., as a function of q_0 , q_1 , and q_2 .

Hint: For instance, the fluid velocity u can be expressed as

$$u = \frac{q_1}{q_0} \quad (6)$$

Solution: Using the EOS, we can write pressure as a function of the conservative variables as

$$\begin{aligned} p &= (\gamma - 1) \left[\rho E - \frac{1}{2} \rho u^2 \right] \\ &= (\gamma - 1) \left[q_2 - \frac{1}{2} \frac{q_1^2}{q_0} \right] \end{aligned}$$

Problem 2 · The Euler equations (1)–(3) may be written in vector form as

$$\frac{\partial \mathbf{q}}{\partial t} + \frac{\partial \mathbf{f}(\mathbf{q})}{\partial x} = \mathbf{0} \quad (7)$$

where $\mathbf{f}(\mathbf{q})$ is the flux vector given as

$$\mathbf{f}(\mathbf{q}) = \begin{bmatrix} f_0(\mathbf{q}) \\ f_1(\mathbf{q}) \\ f_2(\mathbf{q}) \end{bmatrix} \quad (8)$$

Q2 → Write $\mathbf{f}(\mathbf{q})$ in terms of the components of \mathbf{q} .

Hint: For instance, it is pretty clear from Eq. (1) that

$$f_0(\mathbf{q}) = q_1 \quad (9)$$

Solution: For the flux vector, we know that

$$\begin{aligned} f_0 &= \rho u \\ f_1 &= \rho u^2 + p \\ f_2 &= (\rho E + p) u \end{aligned}$$

which we can rewrite using conservative variables as

$$\begin{aligned} f_0 &= q_1 \\ f_1 &= (\gamma - 1) q_2 + \frac{1}{2} (3 - \gamma) \frac{q_1^2}{q_0} \\ f_2 &= \gamma \frac{q_1 q_2}{q_0} - \frac{1}{2} (\gamma - 1) \frac{q_1^3}{q_0^2} \end{aligned}$$

Problem 3 · Now that we have expressed the Euler equations (1)–(3) in vector form, it is useful to express them in the quasi-linear form

$$\frac{\partial \mathbf{q}}{\partial t} + \mathbf{M} \frac{\partial \mathbf{q}}{\partial x} = \mathbf{0} \quad (10)$$

where \mathbf{M} is a 3×3 flux-Jacobian matrix defined as

$$\mathbf{M}(\mathbf{q}) = \frac{\partial \mathbf{f}(\mathbf{q})}{\partial \mathbf{q}} = \begin{bmatrix} \frac{\partial f_0(\mathbf{q})}{\partial q_0} & \frac{\partial f_0(\mathbf{q})}{\partial q_1} & \frac{\partial f_0(\mathbf{q})}{\partial q_2} \\ \frac{\partial f_1(\mathbf{q})}{\partial q_0} & \frac{\partial f_1(\mathbf{q})}{\partial q_1} & \frac{\partial f_1(\mathbf{q})}{\partial q_2} \\ \frac{\partial f_2(\mathbf{q})}{\partial q_0} & \frac{\partial f_2(\mathbf{q})}{\partial q_1} & \frac{\partial f_2(\mathbf{q})}{\partial q_2} \end{bmatrix} \quad (11)$$

Q3 → Find the expression of the matrix \mathbf{M} .

Remark: This matrix is **not** the same as the flux-Jacobian we have derived in class for the 1D Euler equations in non-conservative form. In class, we have considered the variables $(q_0, q_1, q_2) = (\rho, u, p)$; in this midterm, we consider the conserved variables $(q_0, q_1, q_2) = (\rho, \rho u, \rho E)$.

Solution: Rewriting the Euler equations in the quasi-linear form

$$\frac{\partial \mathbf{q}}{\partial t} + \mathbf{M}(\mathbf{q}) \frac{\partial \mathbf{q}}{\partial x} = 0, \quad (12)$$

we get the sought-after flux Jacobian as a function of conservative variables as

$$\mathbf{M}(\mathbf{q}) = \begin{bmatrix} 0 & 1 & 0 \\ \frac{1}{2}(\gamma - 3) \frac{q_1^2}{q_0^2} & (3 - \gamma) \frac{q_1}{q_0} & (\gamma - 1) \\ -\gamma \frac{q_1 q_2}{q_0^2} + (\gamma - 1) \frac{q_1^3}{q_0^3} & \gamma \frac{q_1}{q_0} - \frac{3}{2}(\gamma - 1) \frac{q_1^2}{q_0^2} & \gamma \frac{q_1}{q_0} \end{bmatrix} \quad (13)$$

Problem 4 · The Jacobian matrix \mathbf{M} admits a complete set of eigenvalues and eigenvectors $(\Omega_n, \mathbf{X}_n), n \in \{0, 1, 2\}$, such that

$$\mathbf{M}\mathbf{X}_n = \Omega_n \mathbf{X}_n, \quad \forall n \in \{0, 1, 2\} \quad (14)$$

which leads to the eigendecomposition

$$\mathbf{M} = \mathbf{X}\mathbf{\Omega}\mathbf{X}^{-1} \quad (15)$$

with

$$\mathbf{\Omega} = \text{diag}(\Omega_0, \Omega_1, \Omega_2) \quad (16)$$

$$\mathbf{X} = [\mathbf{X}_0 \quad \mathbf{X}_1 \quad \mathbf{X}_2] \quad (17)$$

Q4 → Find the eigenvalues and eigenvectors of \mathbf{M} and order them such that $\Omega_0 \leq \Omega_1 \leq \Omega_2$.

Hint: To keep the expressions relatively compact, it is useful to introduce the speed of sound c given by

$$c = \sqrt{\frac{\gamma p}{\rho}} \quad (18)$$

Solution: Given the flux Jacobian from Problem 3, we can rewrite it in terms of primitive variables as

$$\mathbf{M}(\mathbf{q}) = \begin{bmatrix} 0 & 1 & 0 \\ \frac{1}{2}(\gamma-3)u^2 & (3-\gamma)u & (\gamma-1) \\ -\gamma uE + (\gamma-1)u^3 & \gamma E - \frac{3}{2}(\gamma-1)u^2 & \gamma u \end{bmatrix} \quad (19)$$

where we can make use of the total specific enthalpy

$$H = E + \frac{p}{\rho} = \frac{c^2}{\gamma-1} + \frac{1}{2}u^2 \quad (20)$$

to simplify it to

$$\mathbf{M}(\mathbf{q}) = \begin{bmatrix} 0 & 1 & 0 \\ \frac{1}{2}(\gamma-3)u^2 & (3-\gamma)u & (\gamma-1) \\ \frac{1}{2}(\gamma-2)u^3 - \frac{c^2u}{\gamma-1} & \frac{c^2}{\gamma-1} - (\gamma-\frac{3}{2})u^2 & \gamma u \end{bmatrix}. \quad (21)$$

Using a symbolic toolbox, we can solve for the eigenvalues and eigenvectors to get

$$\begin{array}{lll} \Omega_0 = u - c & \Omega_1 = u & \Omega_2 = u + c \\ \mathbf{X}_0 = \begin{bmatrix} 1 \\ u - c \\ H - uc \end{bmatrix} & \mathbf{X}_1 = \begin{bmatrix} 1 \\ u \\ \frac{1}{2}u^2 \end{bmatrix} & \mathbf{X}_3 = \begin{bmatrix} 1 \\ u + c \\ H + uc \end{bmatrix} \end{array}$$

Problem 5 · We will use a flux-splitting strategy to solve the system of equations (10). As a result, we need to construct finite-difference operators that are biased in the positive and negative directions. We choose to use the finite-difference operators that approximate a given function $f(x)$ as

$$\left. \frac{df}{dx} \right|_n^+ = \sum_{j=-3}^1 a_j^+ f_{n+j} \quad (22)$$

$$\left. \frac{df}{dx} \right|_n^- = \sum_{j=-1}^3 a_j^- f_{n+j} \quad (23)$$

In other terms, $d\cdot/dx|_n^+$ will be upwind if information propagates from left to right, and $d\cdot/dx|_n^-$ will be upwind if information propagates from right to left.

Q5.1 → Find the coefficients $\{a_j^+\}_{j=-3}^1$ and $\{a_j^-\}_{j=-1}^3$ that lead to the finite-difference approximations (22) and (23).

Q5.2 → Derive the leading truncation error term for each scheme. What is the order of accuracy of these two finite-difference schemes?

Solution: To find the coefficients a_j^- , we use Taylor series approximations about f_i as follows:

$$\begin{aligned}
 f_{i-1} &= f_i - \Delta x \left. \frac{df}{dx} \right|_{x_i} + \frac{\Delta x^2}{2!} \left. \frac{d^2 f}{dx^2} \right|_{x_i} - \frac{\Delta x^3}{3!} \left. \frac{d^3 f}{dx^3} \right|_{x_i} + \frac{\Delta x^4}{4!} \left. \frac{d^4 f}{dx^4} \right|_{x_i} + \text{h.o.t.} \\
 f_i &= f_i \\
 f_{i+1} &= f_i + \Delta x \left. \frac{df}{dx} \right|_{x_i} + \frac{\Delta x^2}{2!} \left. \frac{d^2 f}{dx^2} \right|_{x_i} + \frac{\Delta x^3}{3!} \left. \frac{d^3 f}{dx^3} \right|_{x_i} + \frac{\Delta x^4}{4!} \left. \frac{d^4 f}{dx^4} \right|_{x_i} + \text{h.o.t.} \\
 f_{i+2} &= f_i + (2\Delta x) \left. \frac{df}{dx} \right|_{x_i} + \frac{(2\Delta x)^2}{2!} \left. \frac{d^2 f}{dx^2} \right|_{x_i} + \frac{(2\Delta x)^3}{3!} \left. \frac{d^3 f}{dx^3} \right|_{x_i} + \frac{(2\Delta x)^4}{4!} \left. \frac{d^4 f}{dx^4} \right|_{x_i} + \text{h.o.t.} \\
 f_{i+3} &= f_i + (3\Delta x) \left. \frac{df}{dx} \right|_{x_i} + \frac{(3\Delta x)^2}{2!} \left. \frac{d^2 f}{dx^2} \right|_{x_i} + \frac{(3\Delta x)^3}{3!} \left. \frac{d^3 f}{dx^3} \right|_{x_i} + \frac{(3\Delta x)^4}{4!} \left. \frac{d^4 f}{dx^4} \right|_{x_i} + \text{h.o.t.}
 \end{aligned}$$

Using a Taylor Table approach to solve for the stencils, we get the linear system of equations

$$\begin{bmatrix} 1 & 1 & 1 & 1 & 1 \\ -\Delta x & 0 & \Delta x & 2\Delta x & 3\Delta x \\ \Delta x^2 & 0 & \Delta x^2 & (\Delta x^2)^2 & (2\Delta x^2)^2 \\ -\Delta x^3 & 0 & \Delta x^3 & (\Delta x^2)^3 & (2\Delta x^2)^3 \\ \Delta x^4 & 0 & \Delta x^4 & (\Delta x^2)^4 & (2\Delta x^2)^4 \end{bmatrix} \begin{bmatrix} a_{-1} \\ a_0 \\ a_1 \\ a_2 \\ a_3 \end{bmatrix} = \begin{bmatrix} 0 \\ 1 \\ 0 \\ 0 \\ 0 \end{bmatrix} \quad (24)$$

which we solve using the symbolic toolbox within Matlab and get the stencils

$$a^- = \frac{1}{12\Delta x} \begin{bmatrix} -3 \\ -10 \\ 18 \\ -6 \\ 1 \end{bmatrix}. \quad (25)$$

Similarly, we can solve for the other scheme, which uses the same number of points but is biased in the opposite direction and leads to

$$a^+ = \frac{1}{12\Delta x} \begin{bmatrix} -1 \\ 6 \\ -18 \\ 10 \\ 3 \end{bmatrix} \quad (26)$$

The leading truncation error for the stencils follow

$$\begin{aligned}
 \epsilon^- &= \frac{\Delta x^4}{20} \left. \frac{\partial^5 f}{\partial x^5} \right|_i \\
 \epsilon^+ &= \frac{\Delta x^4}{20} \left. \frac{\partial^5 f}{\partial x^5} \right|_i
 \end{aligned}$$

and the schemes are thus 4th order accurate.

Problem 6 · We want to know whether the schemes constructed in Problem 5 are dispersive and/or dissipative.
 Q6.1 → Using Fourier error analysis, derive the expression of the modified wavenumber κ^* associated with each scheme.

Q6.2 → Plot the real and imaginary parts of $\kappa^* \Delta x$ as a function of $\kappa \Delta x$ between 0 and π , and explain if one should expect the schemes to be dispersive and/or dissipative.

Solution: In order to quantify the dispersive and dissipative behavior of the two schemes at hand, we utilize the Fourier error analysis, as follows.

$$\begin{aligned} \left. \frac{\partial f}{\partial x} \right|_n^- &= \frac{1}{12\Delta x} [-3f_{n-1} - 10f_n + 18f_{n+1} - 6f_{n+2} + f_{n+3}] \\ &= \frac{e^{i\kappa x}}{12\Delta x} [-3e^{-i\kappa\Delta x} - 10 + 18e^{i\kappa\Delta x} - 6e^{2i\kappa\Delta x} + e^{3i\kappa\Delta x}] \\ &= \frac{e^{i\kappa x}}{12\Delta x} \left[15\cos(\kappa\Delta x) - 6\cos(2\kappa\Delta x) + \cos(3\kappa\Delta x) - 10 \right] \\ &\quad + i \left[21\sin(\kappa\Delta x) - 6\sin(2\kappa\Delta x) + \sin(3\kappa\Delta x) \right] \end{aligned}$$

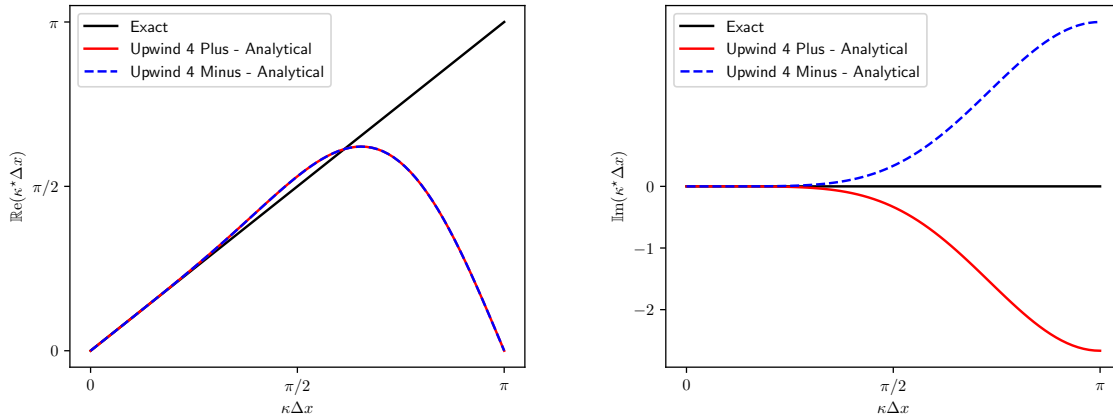
Consequently, the modified wave number follows the expression

$$\kappa^* \Delta x|^- = \frac{1}{12} \left[21\sin(\kappa\Delta x) - 6\sin(2\kappa\Delta x) + \sin(3\kappa\Delta x) \right] - i \left[15\cos(\kappa\Delta x) - 6\cos(2\kappa\Delta x) + \cos(3\kappa\Delta x) - 10 \right]$$

Equally, we can derive the modified wave number for the second scheme as

$$\kappa^* \Delta x|^+ = \frac{1}{12} \left[21\sin(\kappa\Delta x) - 6\sin(2\kappa\Delta x) + \sin(3\kappa\Delta x) \right] - i \left[-15\cos(\kappa\Delta x) + 6\cos(2\kappa\Delta x) - \cos(3\kappa\Delta x) + 10 \right]$$

As the plot below shows, the two schemes have little dispersion for a wide range of frequencies. The imaginary part of the the modified wavenumber shows dispersive behavior, where the D^+ is stable but dissipative and D^- is unstable for the linear advection equation (with for $a > 0$). For the case of $a < 0$, the behavior is switch and D^- is stable but dissipative and D^+ is unstable.



Problem 7 · We want to solve the system of equations (10) in a stable manner by splitting the flux-Jacobian matrix \mathbf{M} as

$$\mathbf{M} = \underbrace{\mathbf{X}\mathbf{\Omega}^+\mathbf{X}^{-1}}_{\mathbf{M}^+} + \underbrace{\mathbf{X}\mathbf{\Omega}^-\mathbf{X}^{-1}}_{\mathbf{M}^-} \quad (27)$$

with

$$\Omega^+ = \frac{1}{2} (\Omega + |\Omega|) \quad (28)$$

$$\Omega^- = \Omega - \Omega^+ \quad (29)$$

This leads to the system of equations

$$\frac{\partial \mathbf{q}}{\partial t} + \mathbf{M}^+ \frac{\partial \mathbf{q}}{\partial x} + \mathbf{M}^- \frac{\partial \mathbf{q}}{\partial x} = \mathbf{0} \quad (30)$$

Following the method of lines and making use of our knowledge of the sign of Ω^+ and of Ω^- , we approximate the spatial derivative in the term $\mathbf{M}^+ \partial \mathbf{q} / \partial x$ with the biased scheme given in Eq. (22), and the spatial derivative in the term $\mathbf{M}^- \partial \mathbf{q} / \partial x$ with the biased scheme given in Eq. (23), so as to transform Eq. (30) into a system of ordinary differential equations (i.e., the semi-discrete 1D Euler problem).

Q7 → Develop a code for solving the semi-discrete 1D Euler problem on the periodic domain $[0, L]$ discretized into N_x grid points.

Hints: You may reuse parts of the code that was developed in HW3 for the semi-discrete 1D linear advection problem. Remember that in this midterm, we consider the three conserved variables $(\rho, \rho u, \rho E)$, therefore there will be three discrete variables at each grid point (i.e., $3N_x$ discrete variables in total). You may order the vector of discrete variables in whichever way you see fit; this should not affect your end results.

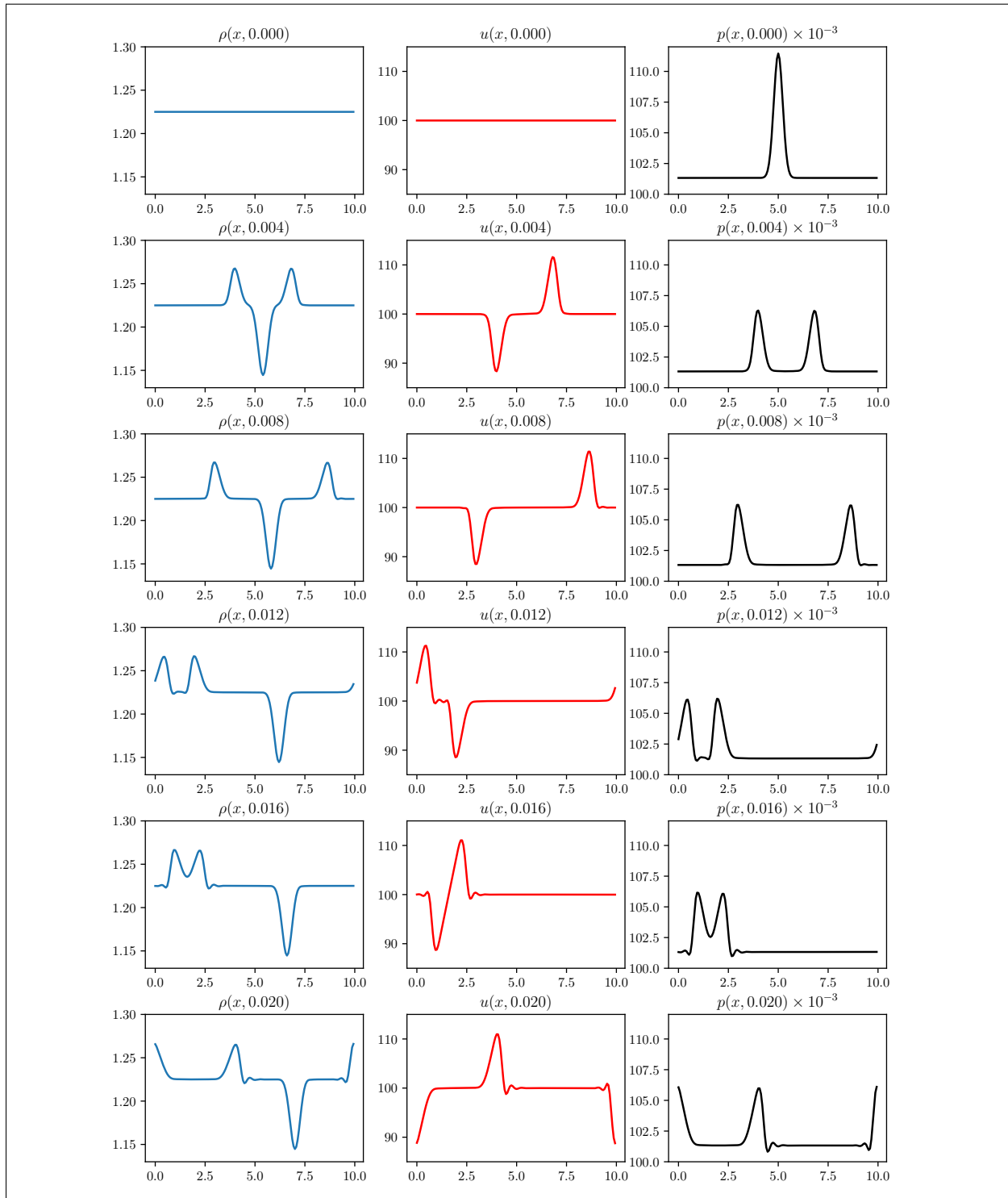
Solution: The code can be found in the Appendix.

Problem 8 · Solve the semi-discrete 1D Euler periodic problem with:

- $L = 10$ m, $N_x = 150$
- $\gamma = 1.4$
- $\rho(x, t = 0) = 1.225$ kg/m³
- $u(x, t = 0) = 100$ m/s
- $p(x, t = 0) = p_\infty \left(1 + \frac{1}{10} \exp(-10(x - L/2)^2)\right)$, with $p_\infty = 101325$ Pa.

Q8 → Plot the spatial variations of (ρ, u, p) at 4 millisecond intervals between $t = 0$ and $t = 0.02$ s, and explain what you observe.

Solution: Lastly, we solve the Euler equations with the given initial condition of a Gaussian pressure pulse which triggers all three characteristics. The unsteady solution for all three primitive variables ρ , u , and p is shown in the figures below. We can clearly see three flow phenomena that consist of two downstream propagating and one upstream propagating characteristic, each with its individual advection speed corresponding to its eigenvalue. As expected, treating each characteristic individually with an upwind scheme results in a stable solution for the entire system and the incorporated numerical dissipation helps reduce any non-physical, numerical oscillations around sharp gradients. At later time steps, smaller such oscillations can be observed which arise from the dispersive properties of the numerical scheme.



Problem 9: This problem is required for students taking AE 410/CSE 461 for four credit hours.
It is not required for students taking AE 410/CSE 461 for three credit hours.

It is customary in CFD development to verify ones code before using it, so as to guarantee the correctness of the simulation results. There exist several ways to do so, including:

Option 1: Develop an exact solution to the governing equations using the method of characteristics, and compare your numerical solution against this analytical solution.

Option 2: Research and implement the method of manufactured solutions.

Option 3: Compare your numerical solution against the numerical solution of another, already established code (this is called cross-code verification).

Q9 → Verify the code you have developed for Problems 7/8 using one of the listed approaches.

Solution: In order to make sure our previously developed code works properly, we verify it using the method of manufactured solutions. For this, we prescribe a manufactured solution that is L -periodic and smooth and analytically solve for the required source terms that we add to the original ODE to acquire the modified equations

$$\frac{\partial \mathbf{q}}{\partial t} = -\frac{\partial \mathbf{f}(\mathbf{q})}{\partial x} + \begin{bmatrix} S_0 \\ S_1 \\ S_2 \end{bmatrix}. \quad (31)$$

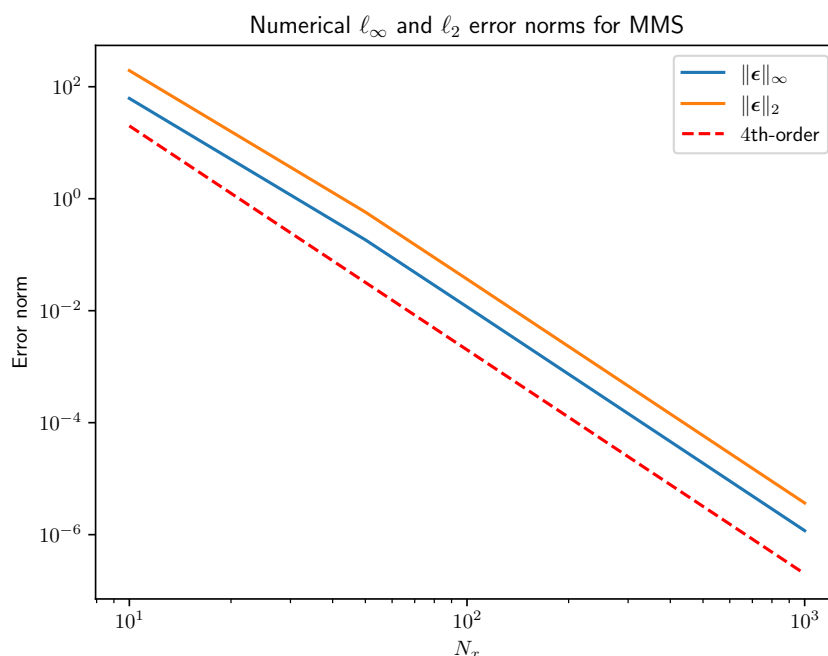
The manufactured solution is then a solution to the modified equations, which we can evaluate numerically to check if the order of accuracy implemented into the RHS evaluation holds true. For the manufactured solutions, we chose

$$\rho = a_1 + a_2 \sin(kx)$$

$$u = a_3 + a_4 \sin(kx)$$

$$p = a_5 + a_6 \sin(kx)$$

with $k = 2\pi/L$. We solve for the resulting source terms $S_i(x)$ symbolically and implement into the existing function that contains and evaluates the RHS of the ODE. Since the manufactured solution is steady, the RHS for given solution should be equal to zero, aside from numerical errors (i.e. finite difference approximation), which we know is of 4^{th} order accuracy. Conducting a grid convergence study, we see that the resulting error follows the expected order of accuracy for the L_2 - and L_∞ -norm.



Solution: Full codes:

Python (0-based indexing)

```

1  # %% [markdown]
2  # Problem 4: Eigendecomposition of the flux Jacobian
3
4  # %%
5  import sympy as sym
6  from IPython.display import display, Latex
7
8  # Declare symbolic variables
9  rho, u, c, g, H = sym.symbols('rho u c gamma H')
10
11 # Construct Jacobian matrix with symbolic variables
12 M = sym.Matrix([[0, 1, 0], [(g-3)*u*sym.Rational(1,2), (3-g)*u, g-1], [(g-2)*u*u*sym.
    Rational(1,2)-c*c*u/(g-1), c*c/(g-1)-(g-sym.Rational(3,2))*u*u, g*u]])
13
14 # Compute and print eigenvalues and eigenvectors
15 EV = M.eigenvecs()
16 for i in range(3):
17     # Extract eigenvalue and eigenvector
18     Eval = EV[i][0]
19     Evec = EV[i][2][0]
20     # Normalize eigenvector by first component
21     Evec /= Evec[0]
22     # Substitute c^2 by (gamma-1)*(H-u^2/2)
23     Evec = Evec.subs(c**2, (g-1)*(H-u**2*sym.Rational(1,2)))
24     # Simplify
25     Evec = sym.simplify(Evec)
26     # Print eigenvalue and eigenvector
27     display(Latex('\Omega_{} = {}'.format(i, sym.latex(Eval))))
28     display(Latex('\mathbf{X}{}_{} = {}'.format('{', '}', i, sym.latex(Evec))))
29
30 # %% [markdown]
31 # Problem 6: Modified wavenumber for the 4th-order upwind scheme
32
33 # %%
34 import numpy as np
35 import matplotlib as mpl
36 import matplotlib.pyplot as plt
37 import cmath
38 plt.rcParams['text.usetex'] = True
39 plt.rcParams['figure.dpi'] = 300
40 plt.rcParams['savefig.dpi'] = 300
41 plt.rc('text.latex', preamble=r'\usepackage{amsmath} \usepackage{amssymb}')
42
43 # We only plot the first N/2+1 eigenvalues, corresponding to k*dx in [0,pi]
44 Nx = 150
45 halfNx = int(0.5*Nx+1)
46 kndx = np.zeros(halfNx)
47
48 # Initialize abscissae
49 for n in range(halfNx):
50     kndx[n] = 2*np.pi*n/Nx
51
52 # Exact modified wavenumbers
53 knmoddx_exact_upwind4_plus = (21*np.sin(kndx)-6*np.sin(2*kndx)+np.sin(3*kndx))/12 - (-15*
    np.cos(kndx)+6*np.cos(2*kndx)-np.cos(3*kndx)+10) * 1j/12
54 knmoddx_exact_upwind4_minus = (21*np.sin(kndx)-6*np.sin(2*kndx)+np.sin(3*kndx))/12 - (15*
    np.cos(kndx)-6*np.cos(2*kndx)+np.cos(3*kndx)-10) * 1j/12
55
56 # Plot real part
57 fig, ax = plt.subplots(1, 1, figsize=(5, 4))
58 ax.plot(kndx, kndx, 'k', label=r'Exact')
59 ax.plot(kndx, knmoddx_exact_upwind4_plus.real, color='red', label=r'Upwind 4 Plus -
    Analytical')

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60 ax.plot(kndx, knmoddx_exact_upwind4_minus.real, color='blue', linestyle='dashed', label=r'
    Upwind 4 Minus - Analytical')
61 plt.ylabel(r'\mathbb{R}\mathrm{e}(\kappa^{\star} \Delta x)'); plt.xlabel(r'\kappa \Delta x$
    ')
62 plt.legend()
63 tick_labels = [r'$0$', r'$\pi/2$', r'$\pi$']
64 ticks = [0, np.pi/2, np.pi]
65 ax.set_xticks(ticks); ax.set_yticks(ticks)
66 ax.set_xticklabels(tick_labels); ax.set_yticklabels(tick_labels)
67 plt.savefig('modwavenumber_real.pdf')
68 plt.show()
69
70 # Plot imaginary part
71 fig, ax = plt.subplots(1, 1, figsize=(5, 4))
72 ax.plot(kndx, 0*kndx, 'k', label=r'Exact')
73 ax.plot(kndx, knmoddx_exact_upwind4_plus.imag, color='red', label=r'Upwind 4 Plus -
    Analytical')
74 ax.plot(kndx, knmoddx_exact_upwind4_minus.imag, color='blue', linestyle='dashed', label=r'
    Upwind 4 Minus - Analytical')
75 plt.ylabel(r'\mathbb{I}\mathrm{m}(\kappa^{\star} \Delta x)'); plt.xlabel(r'\kappa \Delta x$
    ')
76 plt.legend()
77 ytick_labels = [r'$0$', r'$-1$', r'$-2$']
78 yticks = [0, -1, -2]
79 ax.set_xticks(ticks); ax.set_yticks(yticks)
80 ax.set_xticklabels(tick_labels); ax.set_yticklabels(ytick_labels)
81 plt.savefig('modwavenumber_imag.pdf')
82 plt.show()
83
84 # %% [markdown]
85 # Problem 7: Integrator for the 1D Euler equation using flux-splitted finite differences:
86
87 # %%
88 from scipy.linalg import circulant
89 from scipy.integrate import solve_ivp
90
91 # Operator from HW1
92 def D_operator_periodic(N, L, R, a):
93     first_row = np.zeros(N); first_row[0:L+R+1] = a; first_row = np.roll(first_row, -L)
94     return np.array(circulant(first_row)).transpose()
95
96 # Functor for the 1D Euler equation
97 def Euler1D(t, Q, Dp, Dm):
98     # Ratio of specific heats
99     gamma = 1.4
100     # Number of computational nodes (obtained from the size of D)
101     Nx = int(len(Dp))
102     # Extract conserved variables from the vector Q
103     rho = Q[:Nx]
104     rhoU = Q[Nx:2*Nx]
105     rhoE = Q[2*Nx:]
106     # Reconstruct velocity from conserved variables
107     U = rhoU/rho
108     U_sq = U*U
109     # Reconstruct pressure from conserved variables using the EOS
110     p = (gamma-1)*(rhoE-np.multiply(rhoU,U)/2)
111     # Compute speed of sound
112     c_sq = gamma*p/rho
113     c = np.sqrt(c_sq)
114     # Compute total specific enthalpy
115     H = c_sq/(gamma-1)+U_sq/2
116     # Compute spatial derivatives of conserved variables with +/- FD biased schemes
117     drhodxp = Dp@rho; drhodxm = Dm@rho
118     drhoUdpx = Dp@rhoU; drhoUdpxm = Dm@rhoU
119     drhoEdxp = Dp@rhoE; drhoEdxm = Dm@rhoE
120     # Initialize Qdot = F(t,Q) to zero

```

```

121     F = np.zeros(len(Q))
122     # Fill in Qdot = F(t,Q) at each discrete grid point
123     for i in range(Nx):
124         # Construct local diagonal matrix of eigenvalues
125         Omega = np.zeros((3,3)); Omega[0][0]=U[i]; Omega[1][1]=U[i]-c[i]; Omega[2][2]=U[i]
126             +c[i]
127         # Split this matrix into positive and negative eigenvalues
128         Omegap = (Omega + np.abs(Omega))/2; Omegam = Omega-Omegap
129         # Compute the corresponding matrix of eigenvectors and its inverse
130         X = np.array([[1,U[i],U_sq[i]/2],[1,U[i]-c[i],H[i]-U[i]*c[i]],[1,U[i]+c[i],H[i]+U[i]
131             i*c[i]]]).transpose()
132         Xinv = np.linalg.inv(X)
133         # Construct flux-splitting matrices M+/- = X@Omega+/-@X^-1
134         Mp = X@Omegap@Xinv
135         Mm = X@Omegam@Xinv
136         # Compute local derivatives and F(t,Q) for the ith cell
137         dQdxi = np.array([drhodxi[i],drhoUdxi[i],drhoEdxi[i]])
138         dQdxm = np.array([drhodxm[i],drhoUdxm[i],drhoEdxm[i]])
139         F_local = -Mp@dQdxi - Mm@dQdxm
140         # Update the global vector that contains all F(t,Q)
141         for j in range(3): F[j*Nx+i] = F_local[j]
142     return F
143
144 # General integrator function
145 def Integrator(periodic,operator,problem,L,T,Nx,Nt,U0):
146     # Initialize spatial domain
147     x = np.linspace(0, L, Nx, endpoint=(not periodic))
148     dx = x[1] - x[0]
149
150     # Initialize temporal domain
151     t = np.linspace(0, T, Nt, endpoint=True)
152
153     # Construct spatial matrix operator
154     match (operator,periodic):
155         case ('UpwindOrder4FirstDeriv',True): # Periodic 3rd-order upwind
156             Dp = D_operator_periodic(Nx,3,1,[-1,6,-18,10,3]/(12*dx))
157             Dm = D_operator_periodic(Nx,1,3,[-3,-10,18,-6,1]/(12*dx))
158         case _:
159             raise Exception("The %s operator '%s' is not yet implement!" % ('periodic' if
160                 periodic else 'non-periodic', operator))
161
162     # Solve and return solutions!
163     match problem:
164         case 'Euler1D':
165             # Solve initial value problem; see documentation at:
166             # https://docs.scipy.org/doc/scipy/reference/generated/scipy.integrate.
167             # solve_ivp.html
168             sol = solve_ivp(Euler1D, [0, T], U0, args=(Dp,Dm), t_eval=t, rtol=1.0e-9, atol
169                 =1.0e-9, first_step=1e-6)
170             # Transpose solution vector so U has the format (Nt x Nx)
171             U = sol.y.transpose()
172             # Return outputs
173             return t, U
174         case _:
175             raise Exception("The case '%s' is not yet implement!" % problem)
176
177 # %% [markdown]
178 # Problem 8: Solution of an example of 1D Euler periodic problem
179
180 # %%
181 # Initialize space and temporal domains
182 L = 10; T = 0.02
183 Nx = 150; Nt = 6
184 x = np.linspace(0, L, Nx, endpoint=False)
185 dx = x[1] - x[0]
186

```

```

182 # Intialize conserved variables
183 Q0 = np.zeros(3*Nx)
184 gamma = 1.4
185 for i in range(Nx):
186     rho = 1.225
187     U = 100
188     p = 101325*(1+0.1*np.exp(-10*(x[i]-5.0)**2))
189     Q0[i] = rho
190     Q0[Nx+i] = rho*U
191     Q0[2*Nx+i] = p/(gamma-1) + rho*U**2/2
192
193 # Solve problem!
194 t, Q = Integrator(True, 'UpwindOrder4FirstDeriv', 'Euler1D', L, T, Nx, Nt, Q0)
195
196 # Extract conserved variables from the vector Q, and reconstruct velocity and pressure
197 rho = Q[:,0:Nx]
198 rhoU = Q[:,Nx:2*Nx]
199 rhoE = Q[:,2*Nx:3*Nx]
200 U = np.divide(rhoU, rho)
201 p = (gamma-1)*(rhoE-rhoU*rhoU/rho/2)
202
203 # %%
204 # Plot solution
205 for j in range(Nt):
206     fig, (ax_rho, ax_u, ax_p) = plt.subplots(1, 3, sharex=True, figsize=(10, 2))
207     ax_rho.plot(x, rho[j])
208     ax_u.plot(x, U[j], 'r')
209     ax_p.plot(x, p[j]*1e-3, 'k')
210     ax_rho.set_title(r'$\rho(x, %.3f)$' % t[j]); ax_rho.set_ylim(1.13, 1.3); ax_rho.
        set_xlabel(r'$x$')
211     ax_u.set_title(r'$u(x, %.3f)$' % t[j]); ax_u.set_ylim(85, 115); ax_u.set_xlabel(r'$x$')
212     ax_p.set_title(r'$p(x, %.3f)$ \times 10^{-3}$' % t[j]); ax_p.set_ylim(100, 112); ax_p.
        set_xlabel(r'$x$')
213     plt.savefig('euler1d_%i.pdf' % j)
214     plt.show()
215
216 # %% [markdown]
217 # Problem 9: Verification with the method of manufactured solutions (MMS)
218
219 # %%
220 x, rho, u, p, g, k = sym.symbols('x rho u p g k')
221 a1, a2, a3, a4, a5, a6 = sym.symbols('a1 a2 a3 a4 a5 a6')
222
223 # Initialize prescribed variables symbolically
224 rho = a1+a2*sym.sin(k*x)
225 u = a3+a4*sym.sin(k*x)
226 p = a5+a6*sym.sin(k*x)
227
228 # Declare flux functions
229 f0 = rho*u
230 f1 = rho*u*u+p
231 f2 = (p/(g-1)+rho*u*u/2+p)*u
232
233 # Calculate source terms for the MMS (symbolically)
234 S0 = sym.diff(f0, x)
235 S1 = sym.diff(f1, x)
236 S2 = sym.diff(f2, x)
237 print('S0 = ', S0)
238 print('S1 = ', S1)
239 print('S2 = ', S2)
240
241 # Create numerical function that returns source terms from the coefficients and positions
242 MMS = sym.lambdify([a1, a2, a3, a4, a5, a6, g, k, x], [S0, S1, S2], 'numpy')
243
244 # %%
245 # Initialize mesh resolutions

```

```

246 L = 1; Nxs = [10,50,100,500,1000]
247 k = 2*np.pi/L
248
249 # Initialize MMS coefficients
250 a1 = 1
251 a2 = 0.5
252 a3 = -1
253 a4 = 4
254 a5 = 42
255 a6 = 0.7
256
257 L2Error = []
258 LinfError = []
259 # Loop over all considered mesh resolutions
260 for n in range(len(Nxs)):
261     # Create grid
262     Nx = Nxs[n]
263     x = np.linspace(0, L, Nx, endpoint=False)
264     dx = x[1] - x[0]
265
266     # Intialize conserved variables according to MMS function
267     Q0 = np.zeros(3*Nx)
268     gamma = 1.4
269     rho = a1+a2*np.sin(k*x)
270     u = a3+a4*np.sin(k*x)
271     p = a5+a6*np.sin(k*x)
272     Q0[:Nx] = rho
273     Q0[Nx:2*Nx] = rho*u
274     Q0[2*Nx:] = p/(gamma-1)+rho*u*u/2
275
276     # Compute dQ/dt = F(t,Q) using the 1D Euler functor
277     Dp = D_operator_periodic(Nx,3,1,[-1,6,-18,10,3]/(12*dx))
278     Dm = D_operator_periodic(Nx,1,3,[-3,-10,18,-6,1]/(12*dx))
279     F = Euler1D(0,Q0,Dp,Dm)
280
281     # Compute error (i.e., the difference between -F and the MMS source terms)
282     Error = F + np.reshape(MMS(a1,a2,a3,a4,a5,a6,gamma,k,x),3*Nx)
283     L2Error.append(np.max(np.abs(Error)))
284     LinfError.append(np.sqrt(np.sum(Error**2)/(3*Nx)))
285
286 # Plot numerical error norms
287 plt.title(r'Numerical  $\ell_1$  and  $\ell_2$  error norms for MMS')
288 plt.xlabel(r'$N_x$'); plt.ylabel(r'Error norm')
289 plt.loglog(Nxs,LinfError,label=r'$\epsilon_1$')
290 plt.loglog(Nxs,L2Error,label=r'$\epsilon_2$')
291 plt.loglog([1e1,1e3],[2e1,2e-7], 'r--',label=r'$4^{th}$-order')
292 plt.legend()
293 plt.savefig('MMS.pdf')
294 plt.show()

```

Matlab (1-based indexing)

```

1 function [ D ] = D1_operator( n,dx,R,L,a )
2 %computes a finite difference operator
3 % n = number of grid points
4 % dx = step size
5 % [R,L] = right/left bound of stencil
6 % a = stencils
7
8 % initialize
9 D = zeros(n+1);
10 % fill interior domain
11 for i=1:L:n+1-R
12     D(i,i-L:i+R) = a;

```

```

13 end
14 % fill left boundary
15 for i=1:L
16     D(i,1:i+R) = a(L-i+2:end);
17     D(i,end-L+i:end) = a(1:L-i+1);
18 end
19 % fill right boundary
20 for i=n+2-R:n+1
21     D(i,end-L-1+(i-n):end) = a(1:L+(n+2-i));
22     D(i,1:(i-n+R-1)) = a(L+(n+3-i):end);
23 end
24 D = 1/dx*D;
25 end

1 function qdot = Euler1D(t,q,dx,Dp,Dm)
2 clc; t
3
4 % ideal gas
5 gam = 1.4;
6 %
7 Nx = length(Dp);
8
9 % extract current state in primitive variables
10 rho = q(1:Nx);
11 rhou = q(Nx+1:2*Nx);
12 rhoE = q(2*Nx+1:3*Nx);
13 u = rhou./rho;
14 p = (gam-1)*(rhoE-0.5*rho.*u.^2);
15
16 % compute Lambda
17 a_sq = gam*p./rho;
18 for i=1:Nx
19     if (a_sq(i) < 0)
20         %error('contains negative element');
21     end
22 end
23 a = sqrt(gam*p./rho);
24 Lambda1 = (u-a);
25 Lambda2 = u;
26 Lambda3 = u+a;
27 Os = zeros(Nx,Nx);
28 Lambda = [diag(Lambda1),Os,Os;Os,diag(Lambda2),Os;Os,Os,diag(Lambda3)];
29 Lambdap = 0.5*( Lambda + abs(Lambda) );
30 Lambdam = Lambda-Lambdap;
31
32 % compute matrix containing eigenvectors
33 H = 0.5*u.^2+a.^2/(gam-1);
34
35 % eigenvectors
36 V11 = eye(Nx);
37 V21 = diag((u-a));
38 V31 = diag(H-u.*a);
39 V12 = eye(Nx);
40 V22 = diag(u);
41 V32 = diag(0.5*u.^2);
42 V13 = eye(Nx);
43 V23 = diag(u+a);
44 V33 = diag(H+u.*a);
45 V = [V11,V12,V13;V21,V22,V23;V31,V32,V33];
46
47 % compute A+ and A- matrices
48 Vinv = inv(V);
49 Ap = V*Lambdap*Vinv;
50 Am = V*Lambdam*Vinv;
51
52 % compute global differentiation operator

```

```

53 Dpmat = [Dp,Os,Os;Os,Dp,Os;Os,Os,Dp];
54 Dmmat = [Dm,Os,Os;Os,Dm,Os;Os,Os,Dm];
55
56 % compute RHS
57 qdot = -( Ap*(Dpmat*q) + Am*(Dmmat*q));
58 return

1 function [t,q] = integrator(x,T,Nt,q0)
2 %
3 % Inputs
4 %
5 % x :: grid point vector
6 % T :: length of temporal domain, t runs from 0 to T
7 % Nt :: number of points to use in t (for reporting solution, ode45 chooses dt internally)
8 % t :: time vector
9 % q :: solution vector
10
11 % spatial stepsize
12 dx = x(2)-x(1);
13
14 % temporal domain
15 dt = T/(Nt-1);
16 t = dt*linspace(0,Nt-1,Nt);
17
18 % spatial operators
19 ap = [-1,6,-18,10,3]/12;
20 am = [-3,-10,18,-6,1]/12;
21 [ Dp ] = D1_operator( length(x)-1,dx,1,3,ap );
22 [ Dm ] = D1_operator( length(x)-1,dx,3,1,am );
23
24 % call ode45
25 [t,q] = ode45(@(t,y) Euler1D(t,y,dx,Dp,Dm), t, q0);
26 end

1 clear all; clc; close all;
2
3 syms gam u a
4
5 A = [0,1,0;...
6      -1/2*(3-gam)*u^2,(3-gam)*u,(gam-1);...
7      u*(1/2*(gam-2)*u^2-a^2/(gam-1)),a^2/(gam-1)-(gam-3/2)*u^2,gam*u];
8
9 [vecs,vals] = eig(A);
10
11 res1 = A*vecs(:,1)-vals(1,1)*vecs(:,1)
12 res2 = A*vecs(:,2)-vals(2,2)*vecs(:,2)
13 res3 = A*vecs(:,3)-vals(3,3)*vecs(:,3)

1 clear all; close all; clc
2
3 syms dx a1 a2 a3 a4 a5
4
5 %% upwind
6 A = [1,1,1,1,1;...
7      -dx,0,dx,2*dx,3*dx;...
8      dx^2,0,dx^2,(2*dx)^2,(3*dx)^2;...
9      -dx^3,0,dx^3,(2*dx)^3,(3*dx)^3;...
10     dx^4,0,dx^4,(2*dx)^4,(3*dx)^4];
11 b = [0;1;0;0;0];
12
13 x = inv(A)*b
14
15
16 %% downwind
17 A = [1,1,1,1,1;...
18     dx,0,-dx,-2*dx,-3*dx;...

```



```

19     dx^2,0,dx^2,(2*dx)^2,(3*dx)^2;...
20     dx^3,0,-dx^3,-(2*dx)^3,-(3*dx)^3;...
21     dx^4,0,dx^4,(2*dx)^4,(3*dx)^4;
22     b = [0;1;0;0;0];
23
24     x = inv(A)*b

1     clear all; close all; clc
2     n = 150;
3
4     for k=1:n
5         dxkm(k)      = 2*pi*(k-1)/(n-1);
6     end
7
8     modWaveNumAn1 = 1/(sqrt(-1)*12)*( ( 15*cos(dxkm)-6*cos(2*dxkm)+cos(3*dxkm)-10 ) ) ...
9         +1/(sqrt(-1)*12)*( sqrt(-1)*( 21*sin(dxkm)-6*sin(2*dxkm)+sin(3*dxkm) ) );
10
11    modWaveNumAn2 = 1/(sqrt(-1)*12)*( -( 15*cos(dxkm)-6*cos(2*dxkm)+cos(3*dxkm)-10 ) ) ...
12        +1/(sqrt(-1)*12)*( sqrt(-1)*( 21*sin(dxkm)-6*sin(2*dxkm)+sin(3*dxkm) ) );
13
14    % plot modified wavenumber
15    line      = 3;
16    markersize = 8;
17    gcafont    = 20;
18    labelszise = 24;
19    legendsize = 24;
20
21    figure(3), clf; hold on
22    subplot(1,2,1); hold on
23    plot(dxkm,real(modWaveNumAn1),'k-','LineWidth',3,'LineWidth',line)
24    plot(dxkm,real(modWaveNumAn2),'--','LineWidth',3,'LineWidth',line)
25    plot([0,pi],[0,pi],'r-','LineWidth',2)
26    xlim([0,pi]); ylim([0,pi])
27    set(gca,'FontSize',gcafont)
28    xlabel('$k_m \Delta x$','Interpreter','Latex','FontSize',labelsize)
29    ylabel('$\Re(\tilde{k}_m \Delta x)$','Interpreter','Latex','FontSize',labelsize)
30    leg=legend('$D^+$','$D^-$','exact');
31    set(leg,'Interpreter','Latex','FontSize',legendsize,'Location','NorthWest')
32
33    subplot(1,2,2); hold on
34    plot(dxkm,-imag(modWaveNumAn1),'k-','LineWidth',3,'LineWidth',line)
35    plot(dxkm,-imag(modWaveNumAn2),'--','LineWidth',3,'LineWidth',line)
36    plot([0,pi],[0,0],'r-','LineWidth',2)
37    xlim([0,pi]);% ylim([0,pi])
38    set(gca,'FontSize',gcafont)
39    xlabel('$k_m \Delta x$','Interpreter','Latex','FontSize',labelsize)
40    ylabel('$\Im(\tilde{k}_m \Delta x)$','Interpreter','Latex','FontSize',labelsize)
41    leg=legend('$D^+$','$D^-$','exact');
42    set(leg,'Interpreter','Latex','FontSize',legendsize,'Location','NorthWest')
43
44    set(gcf,'PaperPositionMode','Auto')
45    set(gcf,'Position',[100 100 1000 300])
46    print(gcf,'midterm_P6','-djpeg90')

1     clear all; clc; close all;
2
3     Nxs      = [10,50,100,500,1000];
4     L        = 1;
5     gam      = 1.4;
6     pi2L     = 2*pi/L;
7     t        = 1;
8     ap       = [-1,6,-18,10,3]/12;
9     am       = [-3,-10,18,-6,1]/12;
10
11    %% source terms
12    a(1) = 1;

```

```

13 a(2) = 0.5;
14 a(3) = -1;
15 a(4) = 4;
16 a(5) = 0.7;
17 a(6) = 42;
18
19 for ii=1:length(Nxs)
20     Nx      = Nxs(ii);
21
22     %% compute x-domain
23     x        = linspace(0,L,Nx+1);
24     x        = x(1:end-1);
25     dx       = x(2)-x(1);
26
27     %% compute initial condition
28     rho0     = (a(1) + a(2)*sin(pi2L*x))';
29     u0       = (a(3) + a(4)*sin(pi2L*x))';
30     rho0     = u0.*rho0;
31     p0       = (a(5) + a(6)*sin(pi2L*x))';
32     rhoE0    = p0/(gam-1)+0.5*rhou0.^2./rho0;
33     q0       = [rho0;rhou0;rhoE0];
34
35     %% compute RHS with source terms
36     [ Dp ]    = D1_operator( length(x)-1,dx,1,3,ap );
37     [ Dm ]    = D1_operator( length(x)-1,dx,3,1,am );
38     qdot      = Euler1D(t,q0,dx,Dp,Dm);
39     [S1,S2,S3] = sourceterms(a,t,x,L,gam);
40
41     %% compute error and norms
42     err       = qdot + [S1;S2;S3];
43     L2norm(ii) = norm(err);
44     Linfnorm(ii) = max(abs(err));
45 end
46
47 %% plot err norms
48 figure(1), clf;
49 loglog(Nxs,L2norm,'LineWidth',2); hold on
50 loglog(Nxs,Linfnorm,'k-','LineWidth',2);
51 loglog([1e1,1e3],[1e3,1e-4],'r--','LineWidth',2)
52 loglog([1e1,1e3],[1e2,1e-6],'r:','LineWidth',2)
53 xlabel('$N_x$', 'Interpreter', 'Latex', 'FontSize', 30)
54 set(gca, 'FontSize', 20)
55 leg=legend('L2-norm', 'L_{\infty}-norm', '$3.5^{th}$-order', '$4^{th}$-order');
56 set(leg, 'Interpreter', 'Latex', 'FontSize', 24)
57 set(gcf, 'PaperPositionMode', 'Auto')
58 set(gcf, 'Position', [100 100 800 300])
59 print(gcf, 'midterm_P8_errnorms', '-djpeg90')

1 clear all; clc; close all;
2
3 legfont      = 25;
4 axisfont     = 18;
5 labelfont    = 30;
6
7 %% input parameter
8 T            = 0.02;
9 Nt           = 6;
10 L            = 10;
11 mu           = 5;
12 sigma        = 0.1;
13 Nx           = 150;
14
15 %% compute x-domain
16 x            = linspace(0,L,Nx+1);
17 x            = x(1:end-1);
18

```

```

19 %% compute initial condition
20 gam      = 1.4;
21 rho0      = 1.225*ones(Nx,1);
22 rhou0     = 1.225*100*ones(Nx,1);
23 pinf      = 101325;
24 p0        = pinf*(1+0.1*exp(-(x-mu).^2/sigma));
25 rhoE0     = p0/(gam-1)+0.5*rhou0.^2./rho0;
26 q0        = [rho0;rhou0;rhoE0];
27
28 %% compute unsteady solution and eigenvalues of differentiation operators
29 [t,q] = integrator(x,T,Nt,q0);
30
31 %% plot
32 figure(1), clf; set(gcf,'Position',[100 100 1000 400])
33 for tt=1:length(t)
34     clf
35     rho      = q(tt,1:Nx);
36     rhou     = q(tt,Nx+1:2*Nx);
37     rhoE     = q(tt,2*Nx+1:3*Nx);
38
39     u        = rhou./rho;
40     p        = (gam-1)*(rhoE-0.5*rho.*u.^2);
41
42     a(1)=subplot(1,3,1);
43     plot(x,rho0,'r--','LineWidth',2); hold on
44     plot(x,rho,'k-','LineWidth',2);
45     leg1=legend('IC','$\rho$');
46     ylim([1.13,1.3])
47     a(2)=subplot(1,3,2);
48     plot(x,rhou0./rho0,'r--','LineWidth',2); hold on
49     plot(x,u,'k-','LineWidth',2);
50     leg2=legend('IC','$u$');
51     ylim([85,115])
52     title(['t=',num2str(t(tt),'%0.3f')],'FontSize',16)
53     a(3)=subplot(1,3,3);
54     plot(x,p0,'r--','LineWidth',2); hold on
55     plot(x,p,'k-','LineWidth',2);
56     leg3=legend('IC','$p$');
57     ylim([100,112]*1e3)
58
59     set(a,'FontSize',axisfont)
60     xlabel(a(1),'$x$','Interpreter','Latex','FontSize',labelfont)
61     xlabel(a(2),'$x$','Interpreter','Latex','FontSize',labelfont)
62     xlabel(a(3),'$x$','Interpreter','Latex','FontSize',labelfont)
63     set(leg1,'Interpreter','Latex','FontSize',legfont)
64     set(leg2,'Interpreter','Latex','FontSize',legfont)
65     set(leg3,'Interpreter','Latex','FontSize',legfont)
66
67     set(gcf,'PaperPositionMode','Auto')
68     set(gcf,'Position',[100 100 1000 300])
69     fname = ['midterm_P9_soln_tt0',num2str(tt),'.jpg'];
70     print(gcf,fname,'-djpeg90')
71     pause(0.01)
72 end

1 clear all; close all; clc
2
3 syms x pi2L t a1 a2 a3 a4 a5 a6 a7 a8 a9 a10 gam
4
5 % prescribed variables
6 rho = a1 + a2*sin(pi2L*x);
7 u   = a3 + a4*sin(pi2L*x);
8 p   = a5 + a6*sin(pi2L*x);
9
10 % derived variables
11 q1 = rho;

```

```

12 q2 = rho*u;
13 q3 = p/(gam-1)+0.5*rho*u^2;
14 rhoE= q3;
15
16 % fluxes
17 f1 = rho*u;
18 f2 = (gam-1)*rhoE + 0.5*(3-gam)*rho*u^2;
19 f3 = gam*u*rhoE-0.5*(gam-1)*rho*u^3;
20
21 % source terms
22 S1 = diff(f1,x)
23 S2 = diff(f2,x)
24 S3 = diff(f3,x)

1 function [ S1,S2,S3 ] = sourceterms(a,t,x,L,gam)
2
3 pi2L = 2*pi/L;
4 a1 = a(1);
5 a2 = a(2);
6 a3 = a(3);
7 a4 = a(4);
8 a5 = a(5);
9 a6 = a(6);
10
11 % new source terms
12 S1 = a4.*pi2L.*cos(pi2L*x).*(a1 + a2*sin(pi2L*x)) + a2*pi2L*cos(pi2L*x).*(a3 + a4*sin(pi2L
    *x));
13 S2 = (gam - 1).*((a6.*pi2L.*cos(pi2L*x))/(gam - 1) + (a2.*pi2L.*cos(pi2L*x).*(a3 + a4.*
    sin(pi2L*x)).^2)/2 + 2.*a4.*pi2L.*cos(pi2L*x).*(a3 + a4.*sin(pi2L*x)).*(a1/2 + (a2
    .*sin(pi2L*x))/2)) - a2.*pi2L.*cos(pi2L*x).*(gam/2 - 3/2).*(a3 + a4.*sin(pi2L*x))
    .^2 - 2.*a4.*pi2L.*cos(pi2L*x).*(gam/2 - 3/2).*(a1 + a2.*sin(pi2L*x)).*(a3 + a4.*sin
    (pi2L*x));
14 S3 = gam.*(a3 + a4.*sin(pi2L*x)).*((a6.*pi2L.*cos(pi2L*x))/(gam - 1) + (a2.*pi2L.*cos(
    pi2L*x).*(a3 + a4.*sin(pi2L*x)).^2)/2 + 2.*a4.*pi2L.*cos(pi2L*x).*(a3 + a4.*sin(
    pi2L*x)).*(a1/2 + (a2.*sin(pi2L*x))/2)) - a2.*pi2L.*cos(pi2L*x).*(gam/2 - 1/2).*(a3
    + a4.*sin(pi2L*x)).^3 + a4.*gam.*pi2L.*cos(pi2L*x).*((a3 + a4.*sin(pi2L*x)).^2.*(
    a1/2 + (a2.*sin(pi2L*x))/2) + (a5 + a6.*sin(pi2L*x))/(gam - 1)) - 3.*a4.*pi2L.*cos(
    pi2L*x).*(gam/2 - 1/2).*(a1 + a2.*sin(pi2L*x)).*(a3 + a4.*sin(pi2L*x)).^2;
15
16 S1 = S1';
17 S2 = S2';
18 S3 = S3';
19 end

```

Submission guidelines · Instructions on how to prepare and submit your report are available on the course's Canvas page at <https://canvas.illinois.edu/courses/43781/assignments/syllabus>