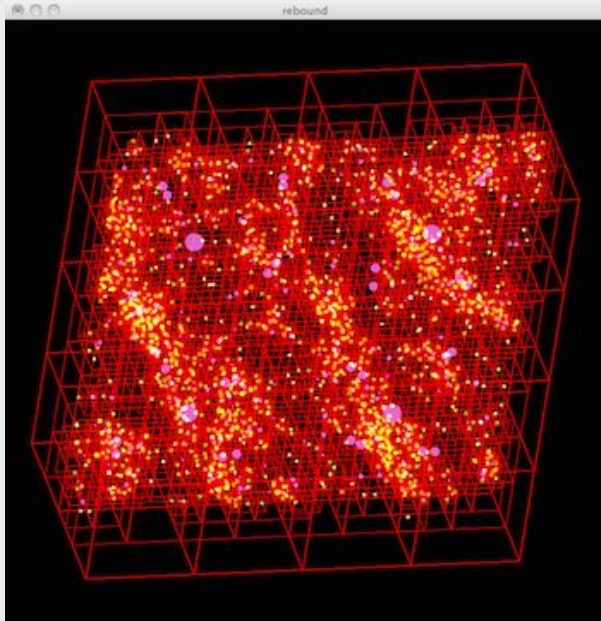


Introducing **Rebound**




An open-source multi-purpose N-body code



Ridlo W. Wibowo

Overview

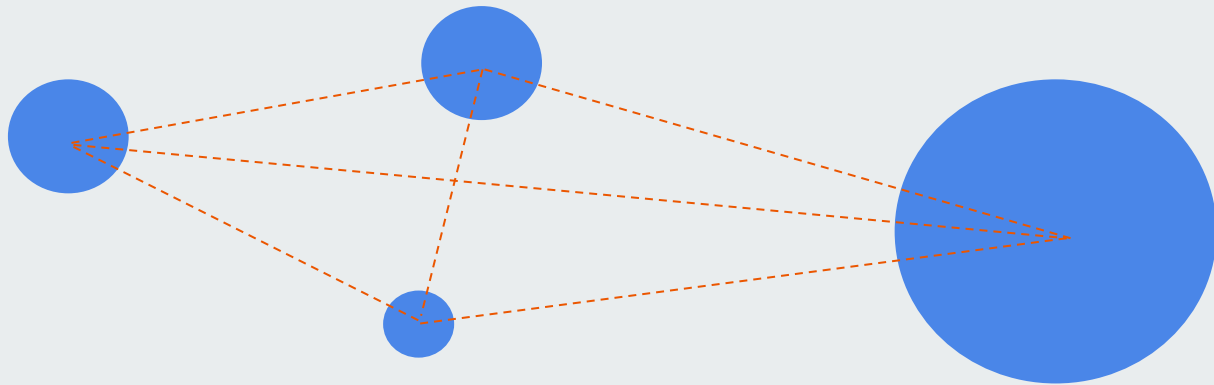


-  N-body simulation
-  What is Rebound?
-  Examples

N-body simulation



$$N > 2$$

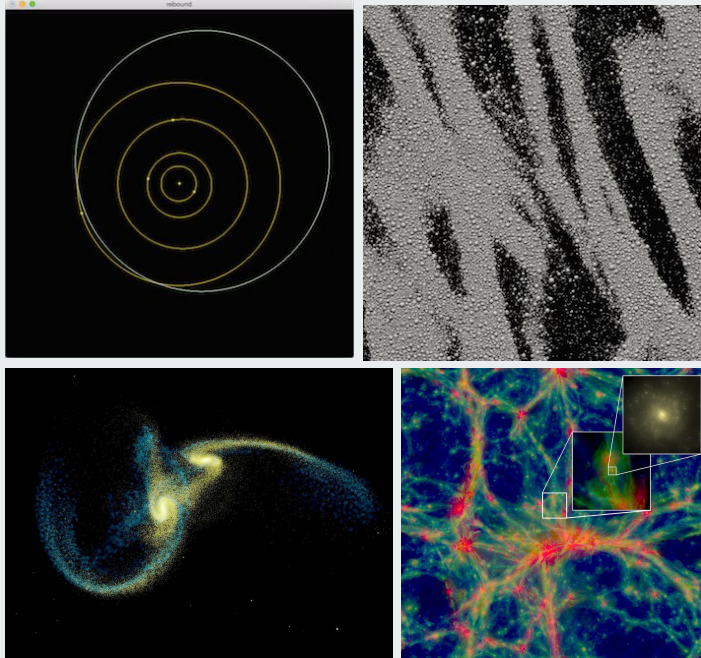


“interaction/force”

N-body simulation



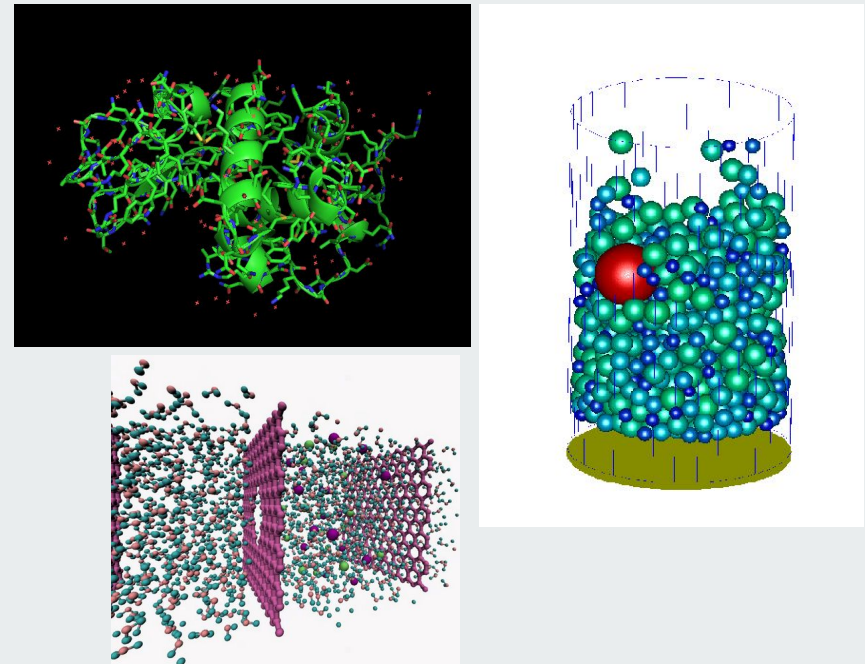
↖
In astronomy/astrophysics



interaction: (Newtonian) **gravitational force**
++ **relativity corr, radiation pressure, drag force, thermal effect, etc**

#principle: how particle interact? boundary of system?
Mostly care about position & velocity only

In other fields: (similar **principle**)
Molecular dynamics (MD), atomic simulation,
granular simulation, etc.



interaction: Coulomb force, normal force,
Lennard-Jones potential, etc

-- sometimes we care about the **potential** (not the force)



rocket and interplanetary flight

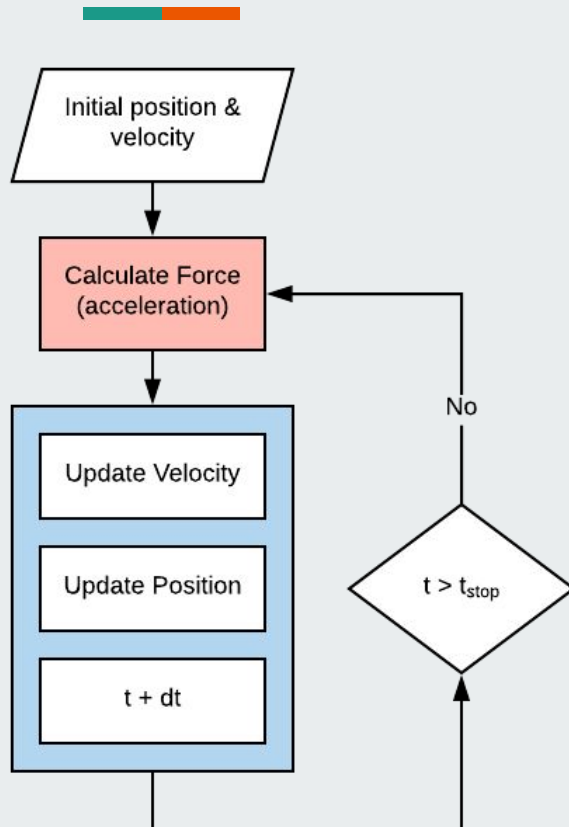
Initial Value Problem

Solving ODE

self-driving car



N-body simulation



Boundary

e.g. open, closed, periodic, ...

How to calculate force? e.g. direct summation, BH-Tree, ...

$$\ddot{\vec{r}}_i = \sum_{j=1, j \neq i}^N \frac{Gm_j}{r^2} \hat{r} + \dots$$

Integration method e.g. euler, **leapfrog**/verlet, RK, ...

$$\begin{aligned} x_{i+\frac{1}{2}} &= x_i + v_i \frac{\Delta t}{2}, & \text{Drift} \\ v_{i+1} &= v_i + a_{i+\frac{1}{2}} \Delta t, & \text{Kick} \\ x_{i+1} &= x_{i+\frac{1}{2}} + v_{i+1} \frac{\Delta t}{2}, & \text{Drift} \end{aligned}$$

Need to be considered:

- ☐ order (precision) of integrator **vs** number of particles **vs** timestep
- ☐ collision **vs** collisionless system
- ☐ symplectic **vs** non-symplectic integrator
- ☐ coordinate used, e.g. cartesian, jacobian, ...
- ☐ type of particles
- ☐ etc.. etc..

-> hard to make a “multi-purpose” N-body code!

Rebound

Rebound is an N-body integrator, i.e. a software that can integrate the motion of particles under the influence of gravity. The particles can represent stars, planets, moons, ring or dust particles.

- Multi-purpose N-body code → *now mainly used for Solar System dynamics and planetary science.
- Written in C with easy-to-use Python 'interface'.

Installation:

```
pip install rebound
```

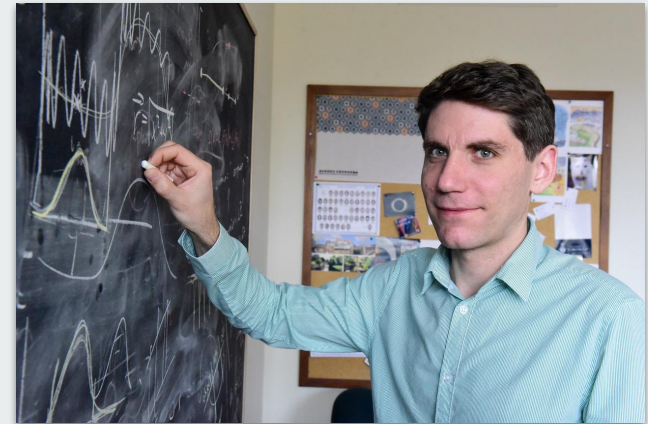
Source code:

<http://github.com/hannorein/rebound>

Doc:

<https://rebound.readthedocs.io/en/latest/>

current version: **Rebound 3.6.6**



[Prof. Hanno Rein](#)

Rebound

Main modules

Boundary/geometry

- None (**default**)
- Open boundary condition
- Periodic boundary condition
- Shearing-sheet (Hill's approx.)

Integrator

- Euler
- Leapfrog
- IAS15 (**default**)
- Wisdom-Holman Fast (WHFast)
- Symplectic Epicycle integrator (SEI)
- Janus (**experimental**)
- Mercurius (**experimental**)
- Hermes (**experimental**)

Gravity solver

- None (no self gravity)
- Direct sum (**default**)
- BH-Tree
- OpenCL (**in progress**)
- FFT (**in progress**)

Collision Detection

- None (**default**)
- Direct nearest-neighbor search
- BH-Tree
- Plane sweep algorithm

Rebound

Additional features

See complete list in [API doc](#) (for [C](#) and [Python](#))

- ❑ **SimulationArchive** → data analysis + enables fully reproducible simulations!
- ❑ Additional Force (native and also “widget” code: [reboundx](#))
- ❑ Tools and Misc. functions:
 - ❑ Simple collision resolve function (merge or elastic)
 - ❑ Easy plot orbit (in python)
 - ❑ Coordinate conversion, e.g. convert Cartesian ↔ Keplerian
 - ❑ OpenGL integration (you need to install it first), WebGL,..
 - ❑ JPL Ephemeris download
 - ❑ Particle ID (& hash)
 - ❑ ...

Examples



Provide a lot of examples in C and python: <https://rebound.readthedocs.io/en/latest/examples.html>

C: <https://github.com/hannorein/rebound/tree/master/examples>

Jupyter Notebook: https://github.com/hannorein/rebound/tree/master/ipython_examples

For this coffee talk:

1. [Quick start](#)
2. [Small bodies & Solar System](#) (how to set parameters)
3. [Planetary Migration](#) & [Debris disk](#)
(additional effect/force, introducing [ReboundX](#))
4. ...

Examples



Some papers:

- **Dynamical Stability of Imaged Planetary Systems in Formation: Application to HL Tau**, Tamayo, D., Triaud, A. H. M. J., Menou, K., & Rein, H. 2015, ApJ, 805, 100
- **No circumbinary planets transiting the tightest Kepler binaries - a possible fingerprint of a third star**, Martin, D. V., Mazeh, T., & Fabrycky, D. C. 2015, Mon Not R Astron Soc, 453, 3554
- **Gap Clearing by Planets in a Collisional Debris Disk**, Nesvold, E. R., & Kuchner, M. J. 2015, ApJ, 798, 83
- **Numerical simulation of tidal evolution of a viscoelastic body modelled with a mass-spring network**, Frouard, J., Quillen, A. C., Efroimsky, M., & Giannella, D. 2016, Mon Not R Astron Soc, 458, 2890
- **Stability of multiplanetary systems in star clusters**, Cai (蔡栩), M. X., Kouwenhoven, M. B. N., Portegies Zwart, S. F., & Spurzem, R. 2017, Mon Not R Astron Soc, 470, 4337

When you **should/(not)** try to use Rebound?

Please, you may want ask me about your N-body problem! :)



Thank you ❤️