



SPECTER: Efficient Evaluation of the Spectral EMD

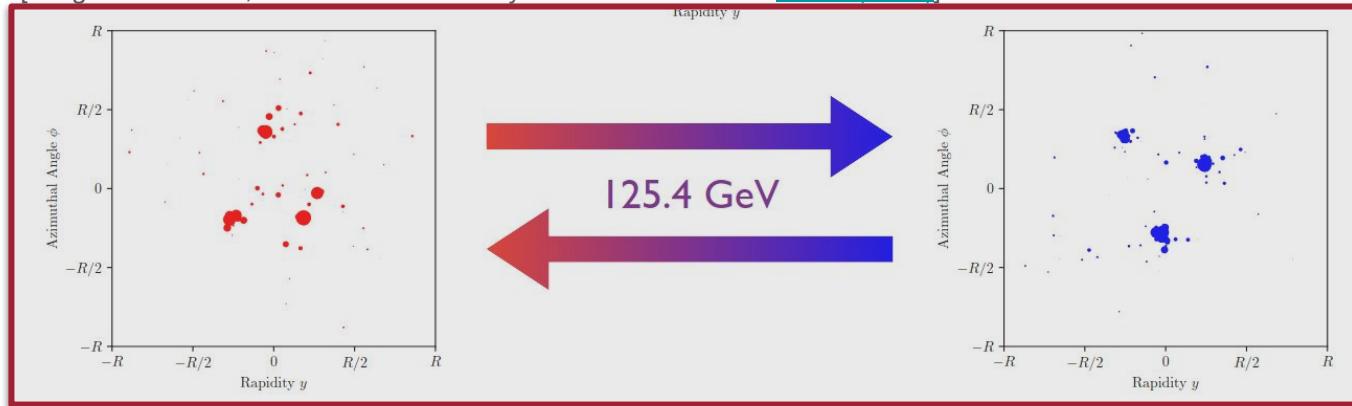
Rikab Gambhir

Email me questions at rikab@mit.edu!
Based on [RG, Larkoski, Thaler, 24XX.XXXX]

Review: Optimal Transport

The **Wasserstein Metric**, a.k.a. **Earth/Energy Mover's Distance (EMD)**^{*}
 allows us to quantitatively ask “How far are two distributions?”

[Image from Thaler; The Hidden Geometry of Particle Collisions [Slides \(2019\)](#)]



$$\text{EMD}_{\beta,R}(\mathcal{E}, \mathcal{E}') = \min_{\{f_{ij} \geq 0\}} \sum_{i=1}^M \sum_{j=1}^{M'} f_{ij} \left(\frac{\theta_{ij}}{R} \right)^\beta + \left| \sum_{i=1}^M E_i - \sum_{j=1}^{M'} E'_j \right|,$$

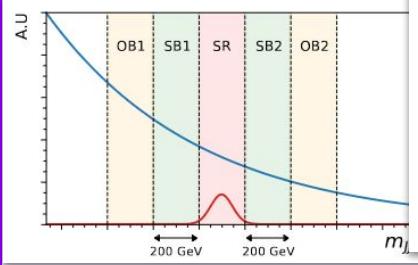
$$\sum_{i=1}^M f_{ij} \leq E'_j, \quad \sum_{j=1}^{M'} f_{ij} \leq E_i, \quad \sum_{i=1}^M \sum_{j=1}^{M'} f_{ij} = \min \left(\sum_{i=1}^M E_i, \sum_{j=1}^{M'} E'_j \right)$$

Very useful for LHC collision data and jets, which are distributions of energy!

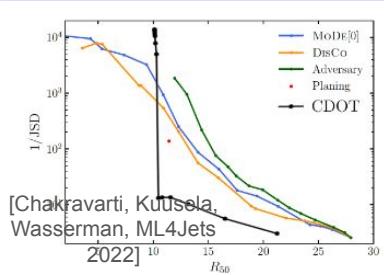
*For this talk, “Optimal Transport”, “Wasserstein Metric”, “Energy Mover’s Distance”, and “Earth Mover’s Distance” are all synonyms.

The **Wasserstein Metric**, a.k.a. **Earth/Mover's Distance (EMD)** has seen increasing interest in jet physics:

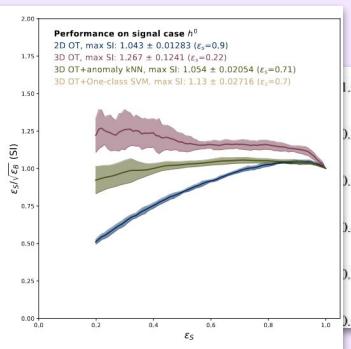
Searches and Anomaly Detection



[Raine, Klein, Sengupta, Golling, [2203.09470](#)]

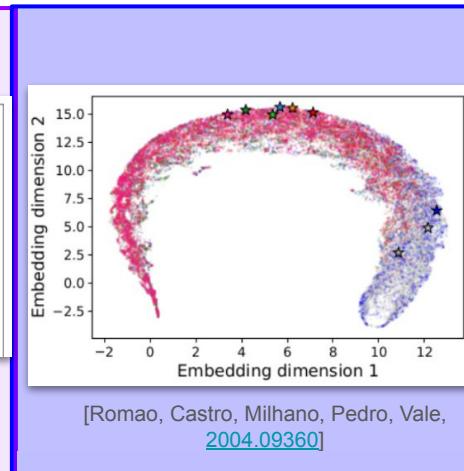


[Chakravarti, Kuusela, Wasserman, ML4Jets, [2022](#)]

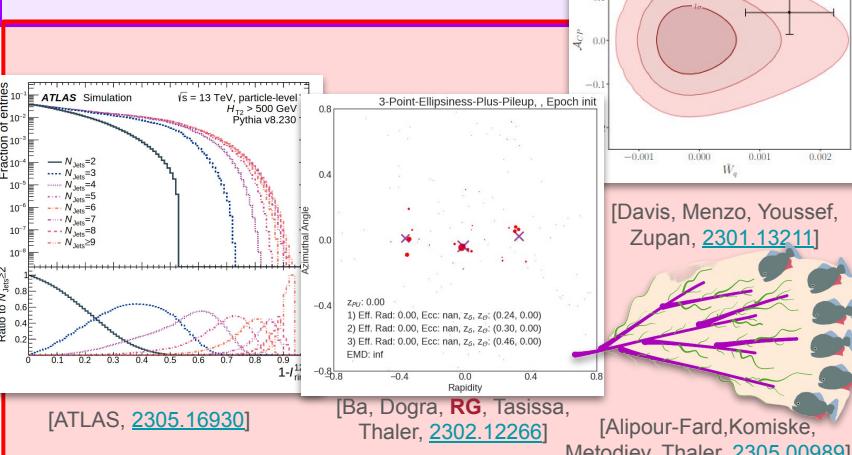


Signal: W
Background: QCD
— KNN (AUC=0.845)
— SVM (AUC=0.869)
— LDA (AUC=0.704)

[Cai, Cheng, Craig, Craig, [2008.08604](#)]



[Romao, Castro, Milhano, Pedro, Vale, [2004.09360](#)]



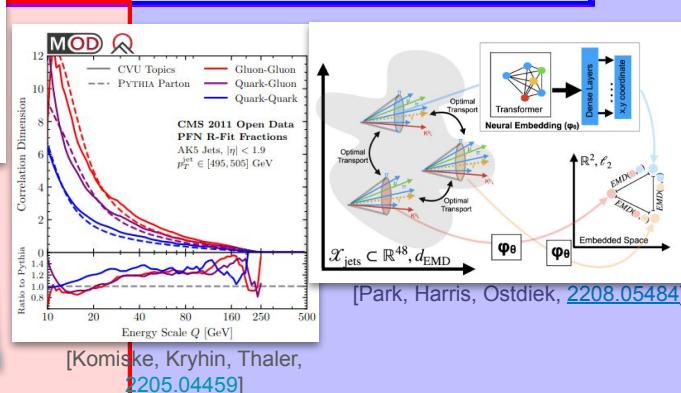
[ATLAS, [2305.16930](#)]

[Ba, Dogra, RG, Tasissa, Thaler, [2302.12266](#)]

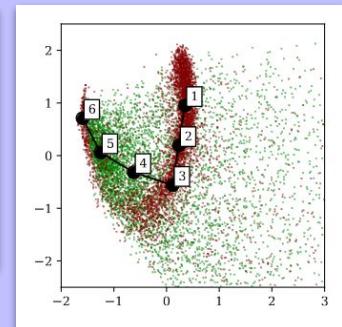
[Alipour-Fard, Komiske, Metodiev, Thaler, [2305.00989](#)]

Not an exhaustive list, let me know if I haven't included your recent EMD application!

Correlations and Embeddings



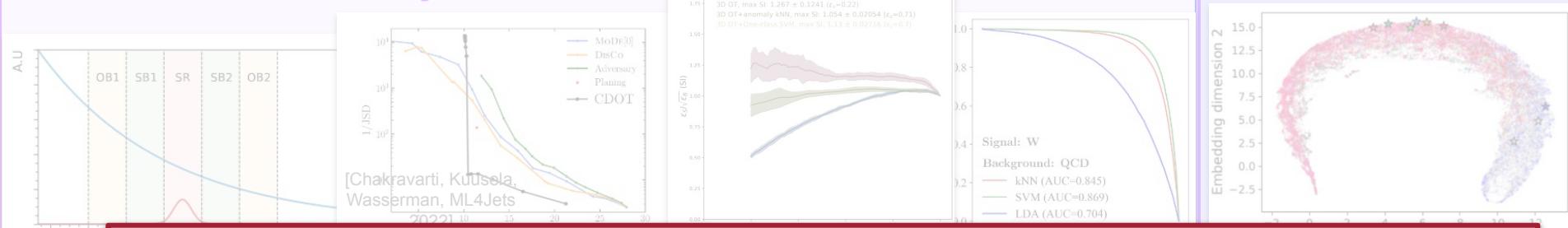
[Komiske, Kryhin, Thaler, [2205.04459](#)]



[Cai, Cheng, Schmitz, Thorpe, DOI:[10.1137/21M1400080](#)]

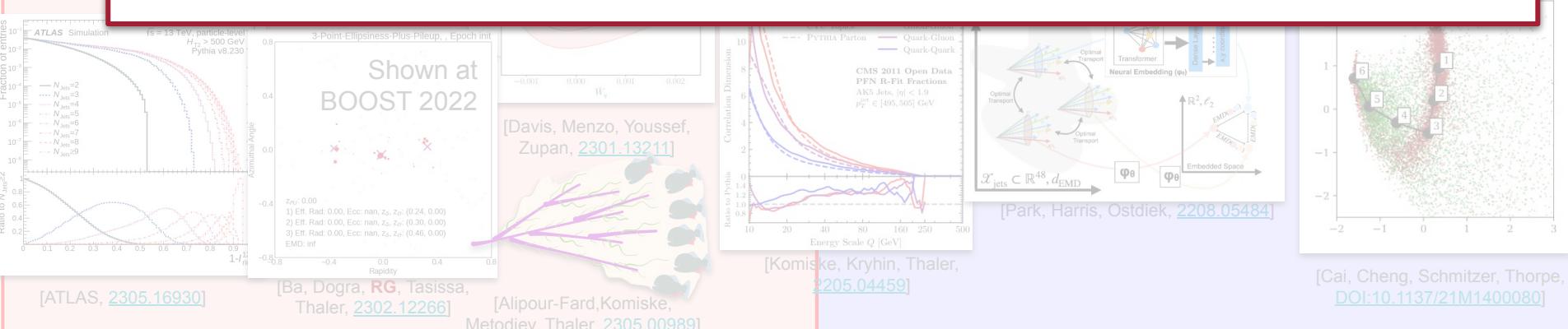
The **Wasserstein Metric**, a.k.a. **Earth/Mover's Distance (EMD)** has seen increasing interest in jet physics:

Searches and Anomaly Detection



But! The EMD is **hard/expensive to calculate**, and even **harder to minimize**...

both computationally, and also in QCD calculations...



Safe Observables for Colliders + Jets

Correlations and Embeddings

Not an exhaustive list, let me know if I haven't included your recent EMD application!

Today ...

An EMD-like metric with associated EMD-like observables that is *easier* and *faster* to calculate using the **Spectral EMD** (SEMD) and **SPECTER**.

With the **Spectral EMD**, we can now (1) evaluate distances between events in closed form, (2) develop EMD-based observables that are fast to numerically evaluate, and (3) often write closed-form expressions for these observables



SPECTER

Logo made with DALL-E. Preliminary.

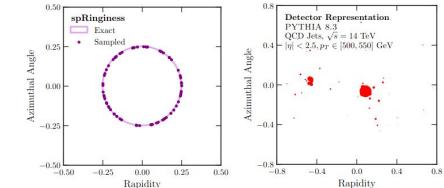
$$\begin{aligned} \text{SEMD}_{\beta,p=2}(s_A, s_B) = & \sum_{i < j \in \mathcal{E}_A} 2E_i E_j \omega_{ij}^2 + \sum_{i < j \in \mathcal{E}_B} 2E_i E_j \omega_{ij}^2 \\ & - 2 \sum_{n \in \mathcal{E}_A^2, l \in \mathcal{E}_B^2} \omega_n \omega_l (\min [S_A(\omega_n^+), S_B(\omega_l^+)] - \max [S_A(\omega_n^-), S_B(\omega_l^-)]) \\ & \times \Theta(S_A(\omega_n^+) - S_B(\omega_l^-)) \Theta(S_B(\omega_l^+) - S_A(\omega_n^-)) , \end{aligned}$$

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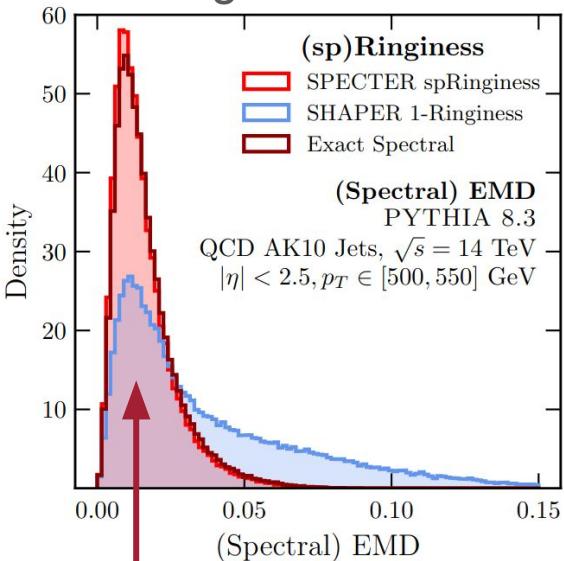
An EMD-like metric with associated EMD-like observables that is easier and faster to calculate using the **Spectral EMD** (SEMD) and **SPECTER**.

With the **Spectral EMD**, we can now (1) evaluate distances between events in closed form, (2) develop EMD-based observables that are fast to numerically evaluate, and (3) often write closed-form expressions for these observables

Old Method: ~ 3 hours
New (Numeric): ~ 5 sec
New (Closed Form): ~ Practically Instant!

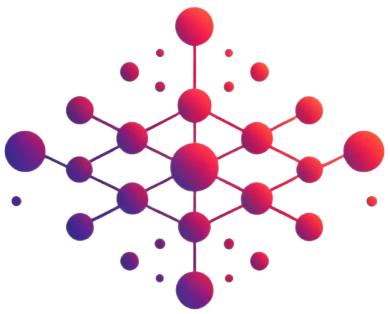


How “ring-like” are Jets?



With these tools, we can evaluate the **red curve** in seconds, equivalent to $\sim 10^7$ OT problems*!

Or, for rings, we can evaluate using an exact **closed form** expression instantaneously!



SPECTER

Logo made with DALL-E. Preliminary.

$$\begin{aligned} \text{SEMD}_{\beta,p=2}(s_A, s_B) = & \sum_{i < j \in \mathcal{E}_A} 2E_i E_j \omega_{ij}^2 + \sum_{i < j \in \mathcal{E}_B} 2E_i E_j \omega_{ij}^2 \\ & - 2 \sum_{n \in \mathcal{E}_A^2, l \in \mathcal{E}_B^2} \omega_n \omega_l (\min [S_A(\omega_n^+), S_B(\omega_l^+)] - \max [S_A(\omega_n^-), S_B(\omega_l^-)]) \\ & \times \Theta(S_A(\omega_n^+) - S_B(\omega_l^-)) \Theta(S_B(\omega_l^+) - S_A(\omega_n^-)), \end{aligned}$$

* $10^7 = 100k$ events $\times \sim 150$ epochs

The Spectral EMD

Similar to the ordinary EMD, the **Spectral EMD (SEMD)**¹ is an IRC-safe metric between events or jets, computed on their **spectral representations**:

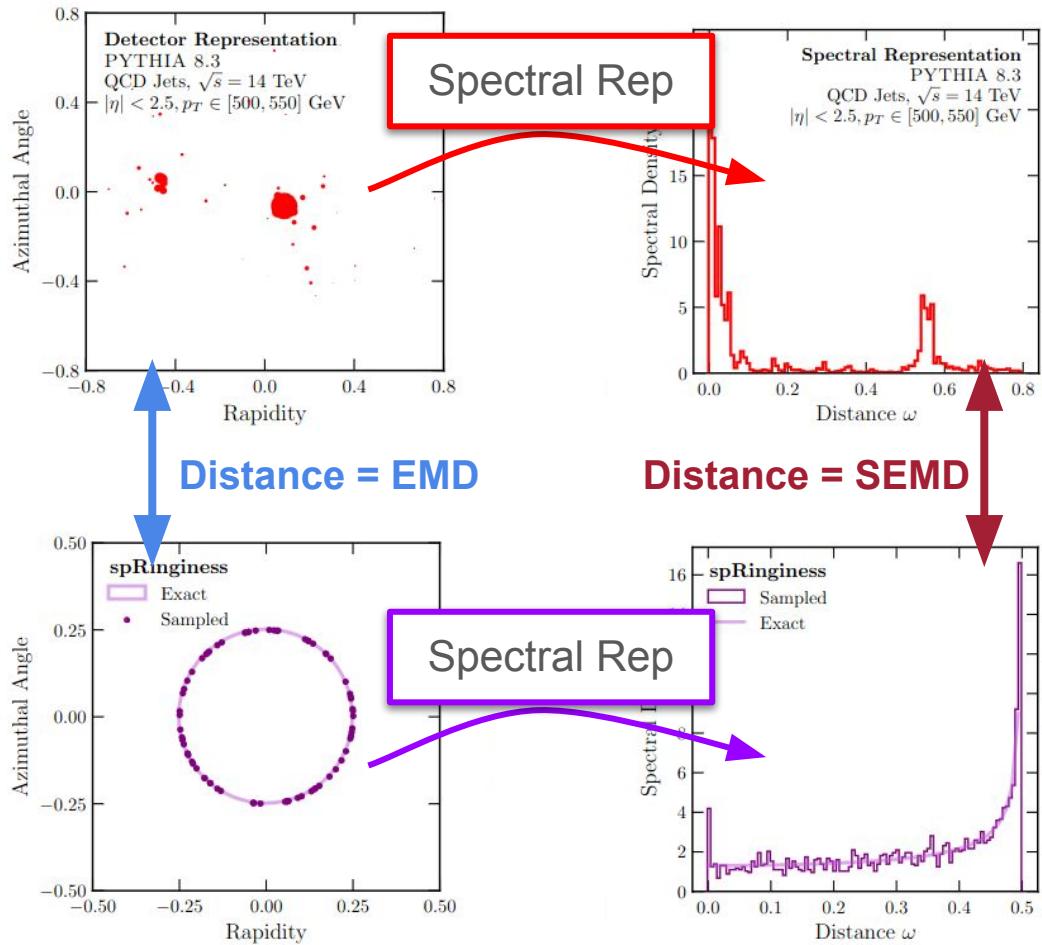
$$s(\omega) \equiv \sum_{i,j \in \mathcal{E}} E_i E_j \delta(\omega - \omega_{ij})$$

= list of energy-weighted pairwise distances

Then the p -SEMD is given by:

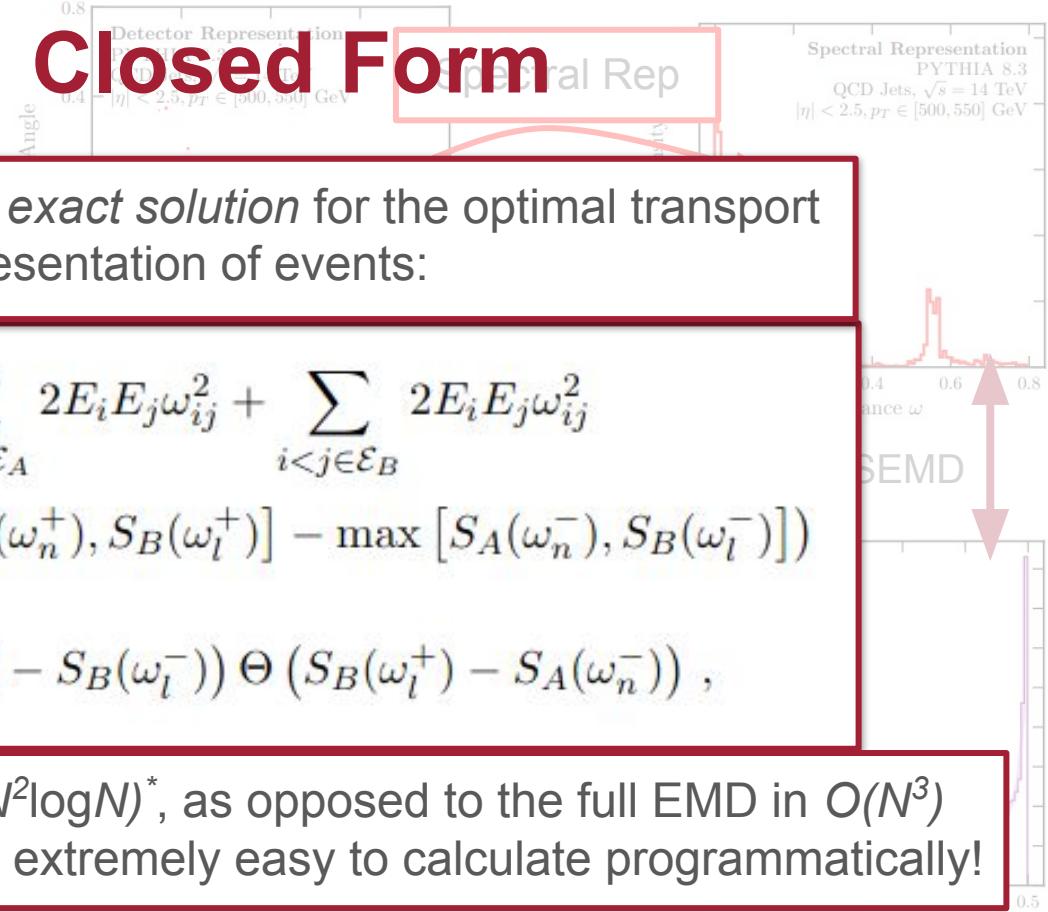
$$\text{SEMD}_p(s_A, s_B) \equiv \int_0^{E_{\text{tot}}^2} dE^2 |S_A^{-1}(E^2) - S_B^{-1}(E^2)|^p$$

Inverse of integral of $s(\omega)$



The SEMD automatically respects all isometries of the metric ω . In this case, the SEMD is invariant under rotations or translations of either event or jet

The Spectral EMD: Closed Form



For $p = 2$, possible to find an *exact solution* for the optimal transport problem on the spectral representation of events:

$$\begin{aligned} \text{SEMD}_{\beta,p=2}(s_A, s_B) = & \sum_{i < j \in \mathcal{E}_A} 2E_i E_j \omega_{ij}^2 + \sum_{i < j \in \mathcal{E}_B} 2E_i E_j \omega_{ij}^2 \\ & - 2 \sum_{n \in \mathcal{E}_A^2, l \in \mathcal{E}_B^2} \omega_n \omega_l (\min [S_A(\omega_n^+), S_B(\omega_l^+)] - \max [S_A(\omega_n^-), S_B(\omega_l^-)]) \\ & \times \Theta(S_A(\omega_n^+) - S_B(\omega_l^-)) \Theta(S_B(\omega_l^+) - S_A(\omega_n^-)), \end{aligned}$$

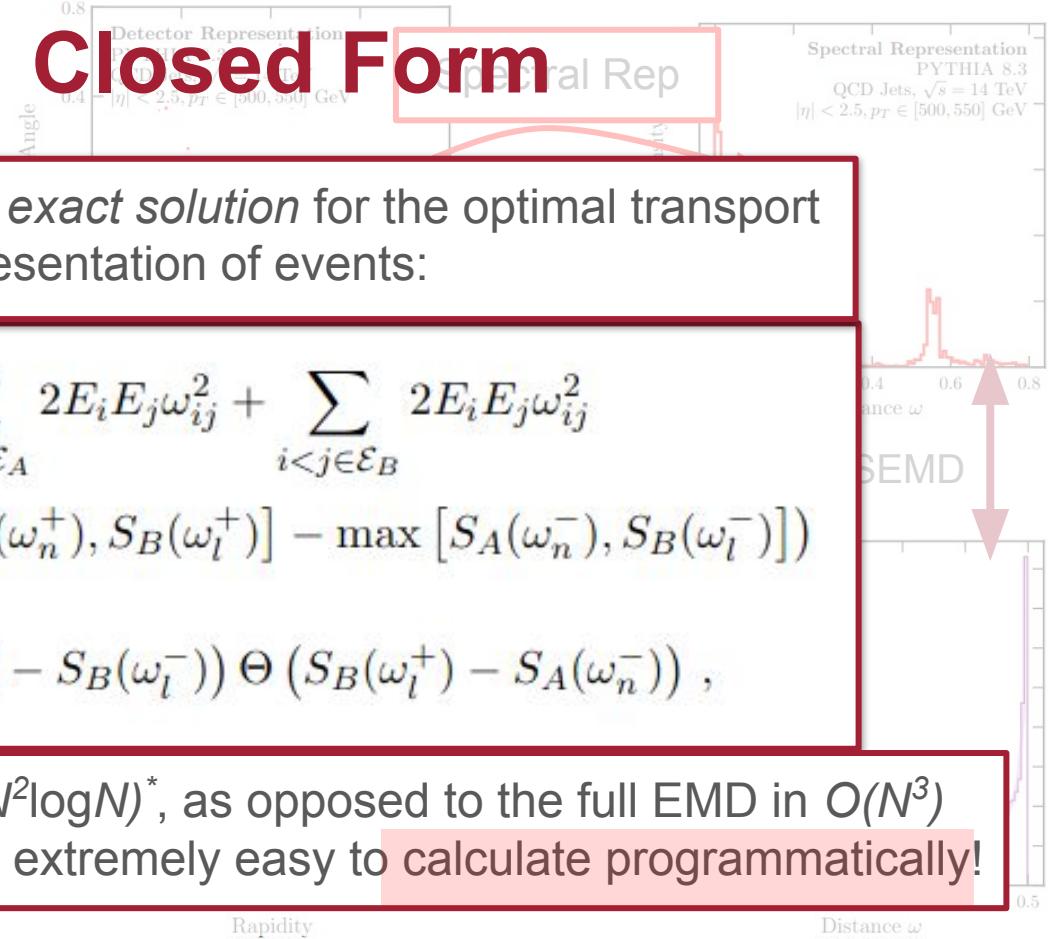
Can be computed *exactly* in $O(N^2 \log N)^*$, as opposed to the full EMD in $O(N^3)$.
Closed form, easy derivatives and extremely easy to calculate programmatically!

$$\text{SEMD}_p(s_A, s_B) \equiv \int_0^{E_{\text{tot}}^2} dE^2 |S_A^{-1}(E^2) - S_B^{-1}(E^2)|^p$$

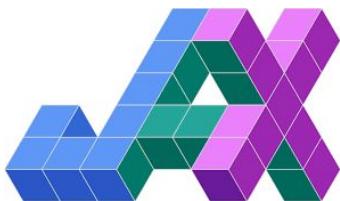
Inverse of integral of $s(\omega)$

*A 1D OT is usually $O(K \log K)$, where K is the number of points. Here, $K \sim N^2$ for particle pairs

The Spectral EMD: Closed Form



Can be computed *exactly* in $O(N^2 \log N)$ ^{*}, as opposed to the full EMD in $O(N^3)$.
Closed form, easy derivatives and extremely easy to calculate programmatically!



Our framework for doing
this, built in Python with JAX

SPECTER

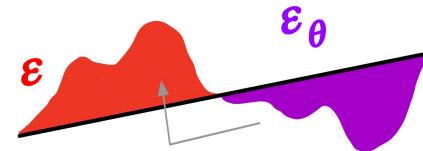
See also: Sinkhorn, Sliced Wasserstein, WGANs, Linearized EMDs, ...

^{*}A 1D OT is usually $O(K \log K)$, where K is the number of points. Here, $K \sim N^2$ for particle pairs

Technical Details ...

For $p = 2$, possible problem on the sp

For events A, B , the p **spectral EMD** is defined as (1D OT!):



EMD = Work done to move "dirt" optimally

$$\text{SEMD}_{\beta,p}(s_A, s_B) \equiv \int_0^{E_{\text{tot}}^2} dE^2 |S_A^{-1}(E^2) - S_B^{-1}(E^2)|^p$$

$$\text{SEMD}_{\beta,p=2}(s_A, s_B) = \sum_{i < j \in \mathcal{E}_A} 2E_i E_j \omega_{ij}^2 + \sum_{i < j \in \mathcal{E}_B} 2E_i E_j \omega_{ij}^2$$

$$- 2 \sum_{n \in \mathcal{E}_A^2, l \in \mathcal{E}_B^2} \omega_n \omega_l (\min [S_A(\omega_n^+), S_B(\omega_l^+)] - \max [S_A(\omega_n^-), S_B(\omega_l^-)]) \\ \times \Theta(S_A(\omega_n^+) - S_B(\omega_l^-)) \Theta(S_B(\omega_l^+) - S_A(\omega_n^-)),$$

S = **cumulative spectral function**

\pm indicates whether or not to include ω in the sum

The trick: Sum over pairs n of particles within each event.

Looks like $O(N^4)$, but with clever sorting & indexing in 1D, reduces to **$O(N^2)$** !

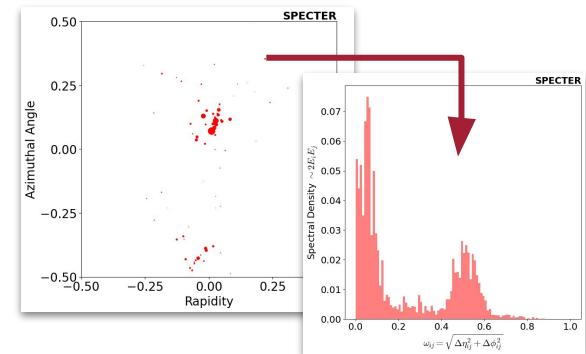
The **spectral density function**

$$s(\omega) = \sum_{i=1}^N \sum_{j=1}^N E_i E_j \delta(\omega - \omega(\hat{n}_i, \hat{n}_j))$$

Pairwise Distances

Reduces events to 1D, while preserving all* information about the event, up to translations and rotations.

*up to measure 0, but important degeneracies, ask me about this later!



EMDs, ...

[If we have time] The Algorithm

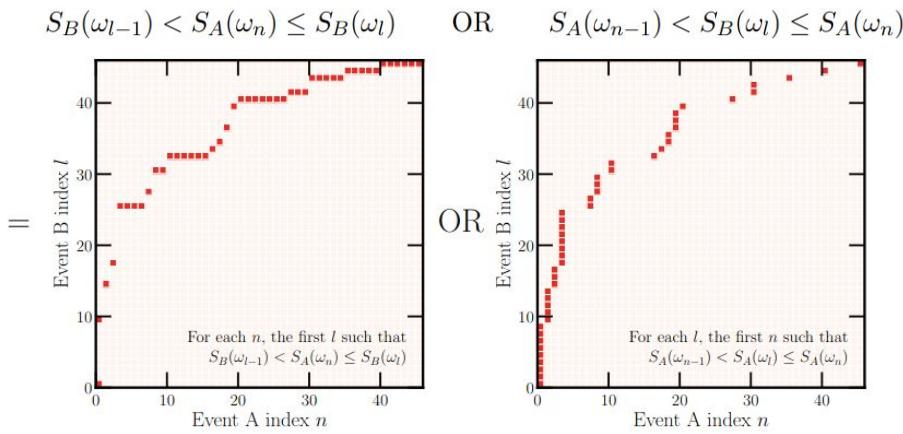
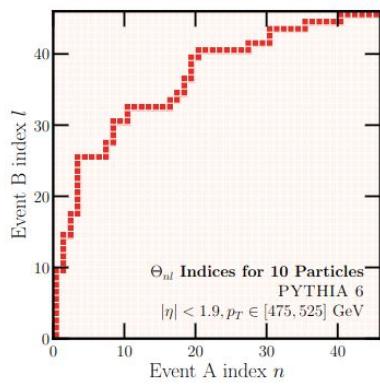
$$\text{SEMD}_{\beta,p=2}(s_A, s_B) = \sum_{i < j \in \mathcal{E}_A} 2E_i E_j \omega_{ij}^2 + \sum_{i < j \in \mathcal{E}_B} 2E_i E_j \omega_{ij}^2$$

$$- 2 \sum_{n \in \mathcal{E}_A^2, l \in \mathcal{E}_B^2} \omega_n \omega_l (\min [S_A(\omega_n^+), S_B(\omega_l^+)] - \max [S_A(\omega_n^-), S_B(\omega_l^-)]) \\ \times \Theta(S_A(\omega_n^+) - S_B(\omega_l^-)) \Theta(S_B(\omega_l^+) - S_A(\omega_n^-)),$$

This sum looks like it goes as $O(N^4)$ as a sum of pairs of pairs, but it turns out only $O(N^2)$ terms survive the Θ -functions!

The Trick: Pre-compute which pairs will activate the Θ -functions by using the fact that in 1D, distances ω can be sorted!

SPECTER



The inequalities can be evaluated efficiently on sorted lists, bringing the total runtime to **$O(N^2 \log N)$** .

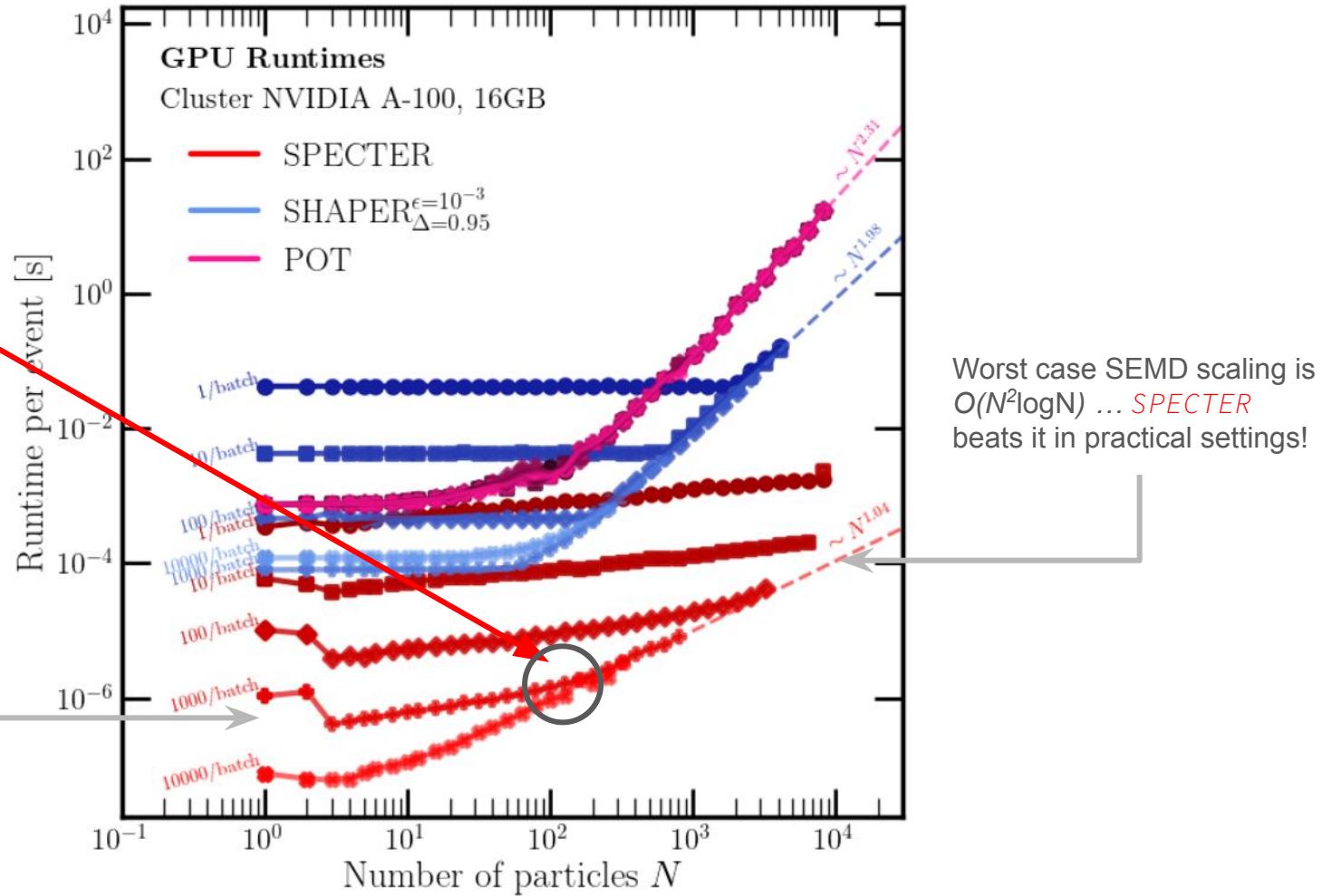
Ask me afterwards if you want more details on the algorithm and how the code works!

SPECTER is FAST (*BOOSTED*)!

Running on a single GPU on my local compute cluster ...

SEMDs between jets with ~ 100 particles can be computed at a rate of **one million SEMDs per second!**

Highly parallelized:
Efficiency gains by processing many events at once!

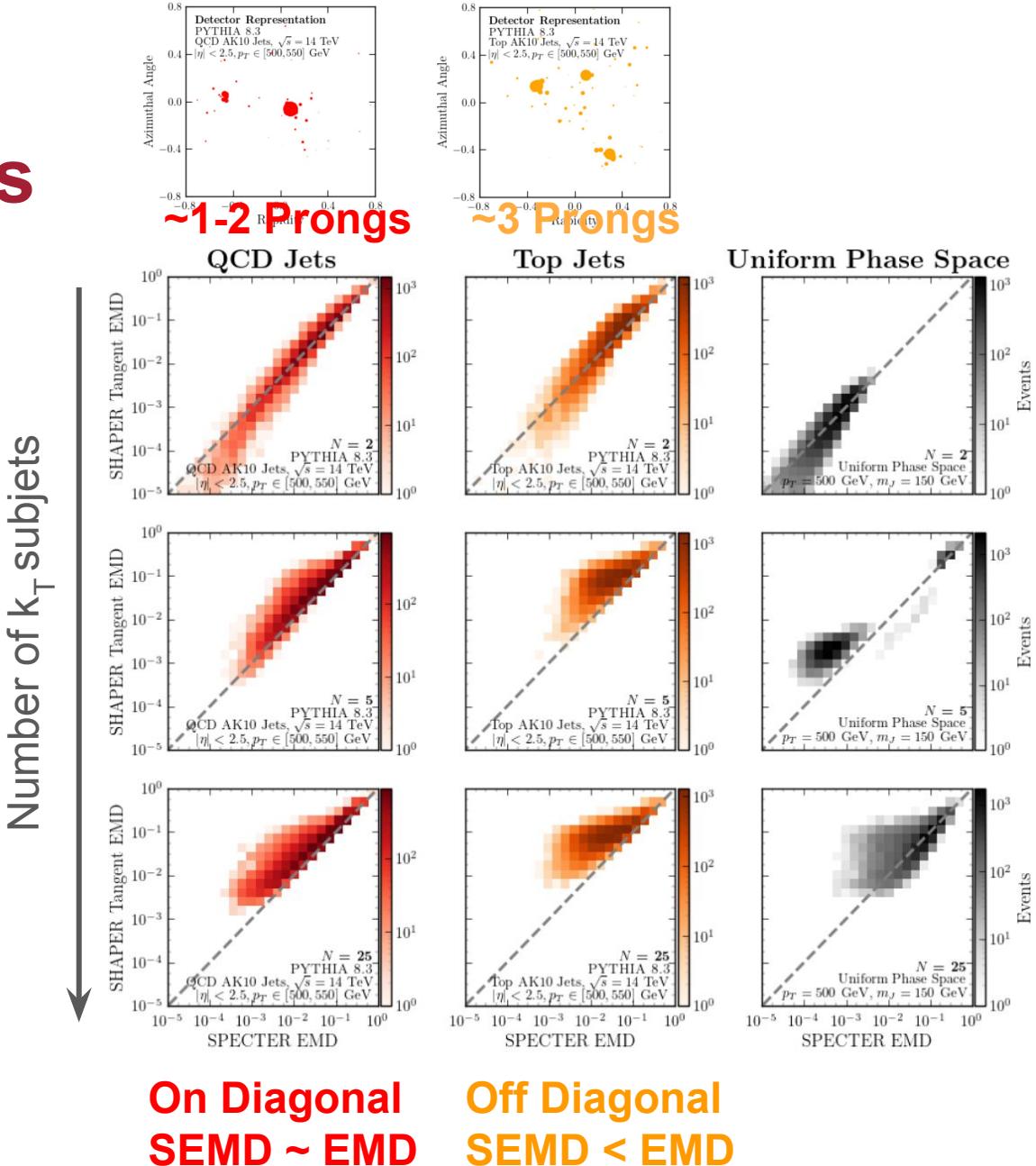


Pairwise (S)EMDs

We can now easily evaluate SEMDs between pairs of events!

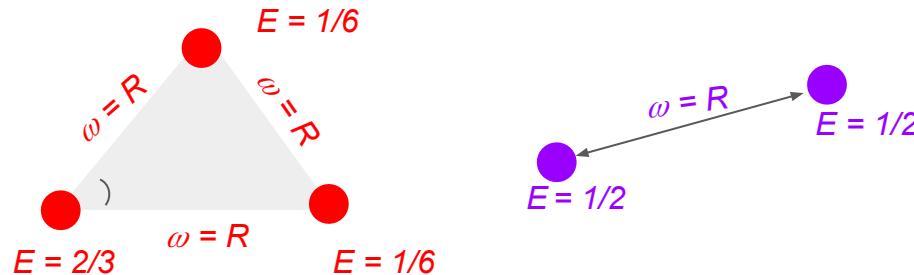
The SEMD and EMD are *not* the same metric, but they are correlated, and this correlation can be different for different types of physics!

The SEMD is invariant to translations and rotations of the jets, but the EMD does not and this needs to be minimized over.



The EMD vs. The SEMD

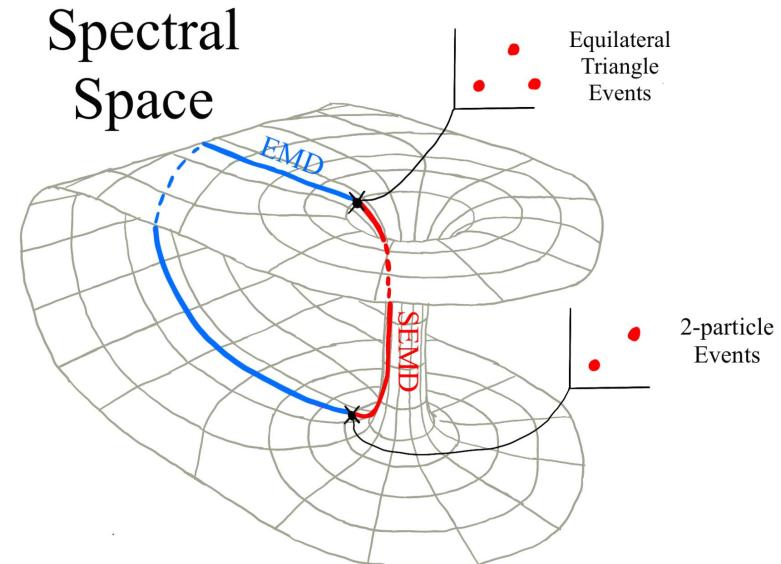
The SEMD is *topologically different* from the ordinary EMD!



This three particle event looks very different from this two particle event, and thus they have a large EMD.

But their SEMD is zero! The spectral function only cares about pairwise distances and degenerate configurations can occur.

This can come into play when events have 3 hard prongs, e.g. top jets!



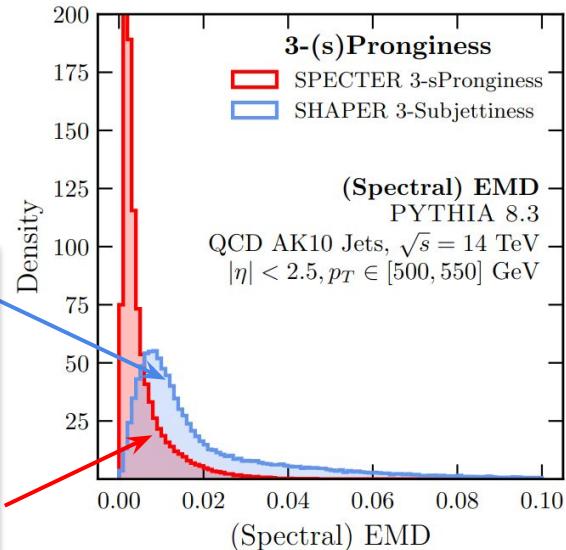
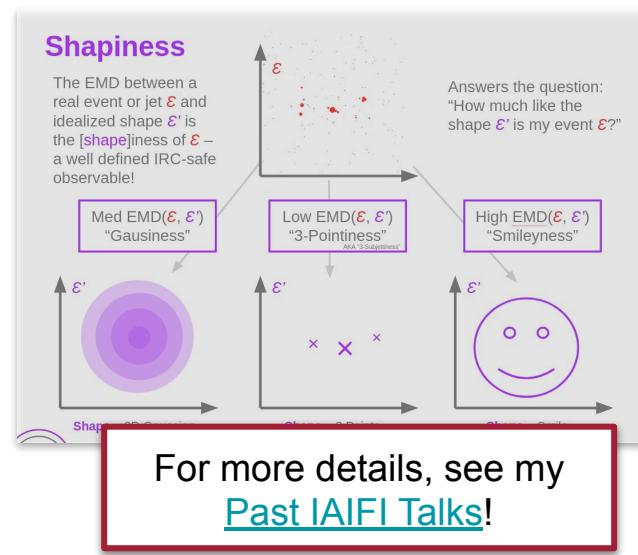
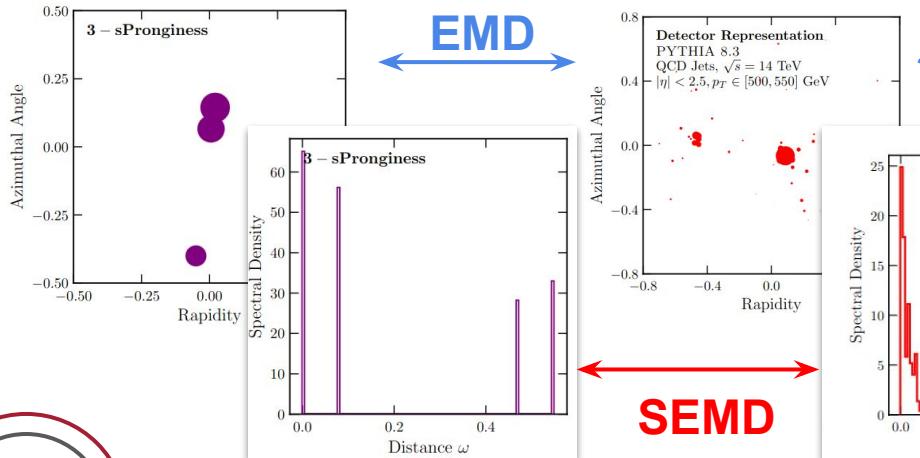
The space of events gets “pinched” at degenerate configurations when looking at only their spectral representations

SEMD Observables

With a geometry based metric, we can now define IRC-safe **shape observables** by finding events that minimize the metric:

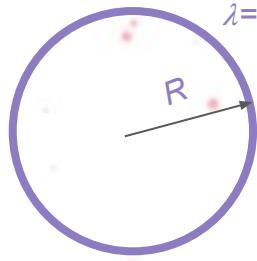
$$\mathcal{O}(\mathcal{E}) = \min_{\mathcal{E}' \in \mathcal{M}} [\text{SEMD}(s(\mathcal{E}), s(\mathcal{E}'))]$$

e.g. How 3-pointy are jets? (3-subjettiness)
Minimize the metric over 3-particle events



Full Example: How “ring-like” are jets?

Step 1: Define the shape with parameters



$$\lambda = E_{tot} / 2\pi R$$

```
def sample_circle(params, N, seed):  
    thetas = jax.random.uniform(seed, shape)  
    x = params["Radius"] * jnp.cos(thetas)  
    y = params["Radius"] * jnp.sin(thetas)
```

Unlike ordinary EMD, not necessary to specify center / orientation!

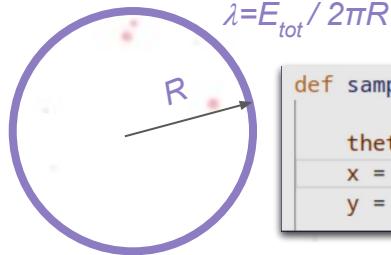
Shapes are parameterized distributions of energy on the detector space.

Many of your favorite observables, like N -(sub)jettiness, thrust, and angularities take the form of finding the shape that best fits an event’s energy distribution.

Custom shapes define custom IRC-safe observables – to define a shape, all you need is to define a parameterized energy distribution and how to sample points from it!

Full Example: How “ring-like” are jets?

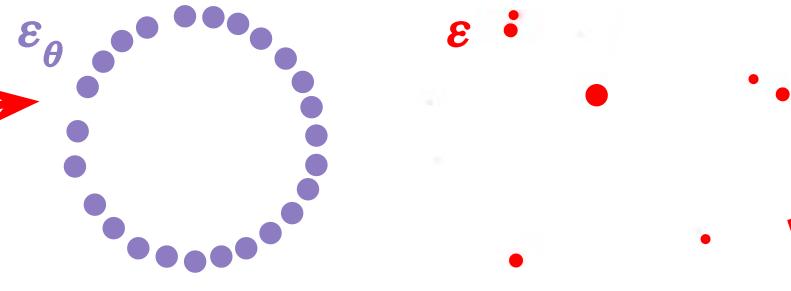
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Step 2: Sample from Parameterized Shapes



Step 3: Calculate the spectral metric between events and shapes

$$\begin{aligned} \text{SEMD}_{\beta,p=2}(s_A, s_B) = & \sum_{i < j \in \mathcal{E}_A} 2E_i E_j \omega_{ij}^2 + \sum_{i < j \in \mathcal{E}_B} 2E_i E_j \omega_{ij}^2 \\ - 2 & \sum_{n \in \mathcal{E}_A^2, l \in \mathcal{E}_B^2} \omega_n \omega_l (\min [S_A(\omega_n^+), S_B(\omega_l^+)] - \max [S_A(\omega_n^-), S_B(\omega_l^-)]) \\ & \times \Theta(S_A(\omega_n^+) - S_B(\omega_l^-)) \Theta(S_B(\omega_l^+) - S_A(\omega_n^-)), \end{aligned}$$

Key difference from previous work: We use the SEMD, *not* the EMD!

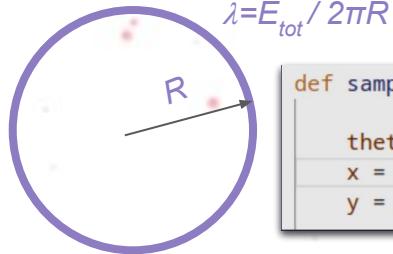
The $p = 2$ spectral EMD between two sets of discrete points has a closed-form solution with only binary discrete minimizations.

We discretize our shape by randomly sampling points from it.

If the spectral functions are sorted, can compute the SEMD in $\sim O(N^2 \log N)$ time!

Full Example: How “ring-like” are jets?

Step 1: Define the shape with parameters



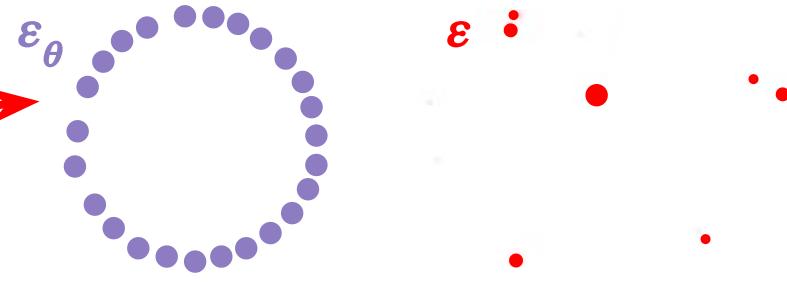
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    x = params["Radius"] * jnp.cos(thetas)  
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```

Unlike ordinary EMD, not necessary to specify center / orientation!

We have an explicit formula for the spectral EMD, so we can automatically differentiate through it

Standard ML procedure: Sample, calculate gradients, gradient descent, repeat! Analogous to WGANs.

Step 2: Sample from Parameterized Shapes

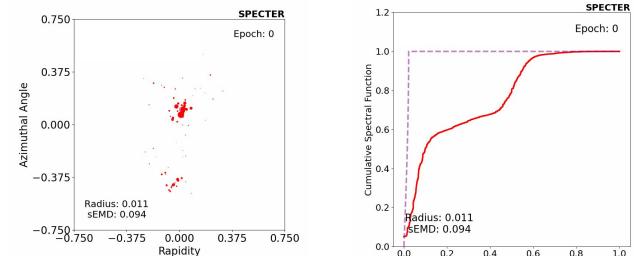


Step 3: Calculate the spectral metric between events and shapes

$$\text{SEMD}_{\beta,p=2}(s_A, s_B) = \sum_{i < j \in \mathcal{E}_A} 2E_i E_j \omega_{ij}^2 + \sum_{i < j \in \mathcal{E}_B} 2E_i E_j \omega_{ij}^2 - 2 \sum_{n \in \mathcal{E}_A^2, l \in \mathcal{E}_B^2} \omega_n \omega_l (\min [S_A(\omega_n^+), S_B(\omega_l^+)] - \max [S_A(\omega_n^-), S_B(\omega_l^-)]) \times \Theta(S_A(\omega_n^+) - S_B(\omega_l^-)) \Theta(S_B(\omega_l^+) - S_A(\omega_n^-)),$$

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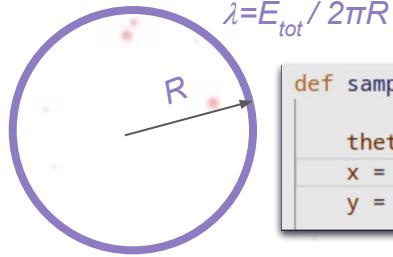
Step 4: Minimize w.r.t. parameters using grads



Pictured: Animation of optimizing for the radius R

Full Example: How “ring-like” are jets?

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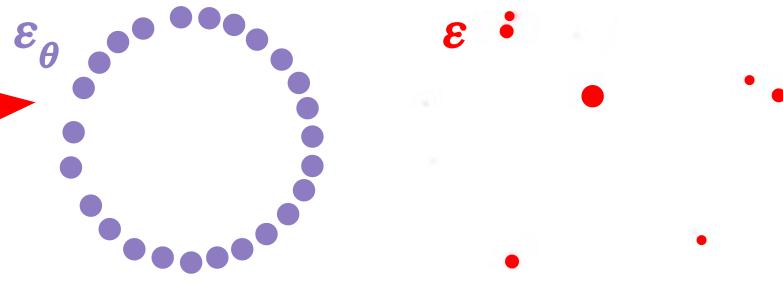
SPECTER is our code interface for performing these steps: sampling from user-defined shapes, calculating spectral functions and differentiable EMDS, and optimizing over parameters.

Built in highly parallelized and compiled JAX

SPECTER

Our code framework for these calculations

Step 2: Sample from Parameterized Shapes

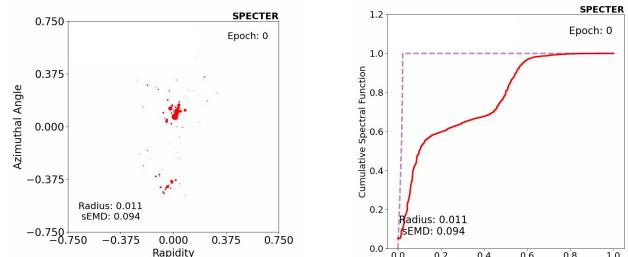


Step 3: Calculate the spectral metric between events and shapes

$$\begin{aligned} \text{SEMD}_{\beta,p=2}(s_A, s_B) = & \sum_{i < j \in \mathcal{E}_A} 2E_i E_j \omega_{ij}^2 + \sum_{i < j \in \mathcal{E}_B} 2E_i E_j \omega_{ij}^2 \\ - 2 & \sum_{n \in \mathcal{E}_A^2, l \in \mathcal{E}_B^2} \omega_n \omega_l (\min [S_A(\omega_n^+), S_B(\omega_l^+)] - \max [S_A(\omega_n^-), S_B(\omega_l^-)]) \\ & \times \Theta(S_A(\omega_n^+) - S_B(\omega_l^-)) \Theta(S_B(\omega_l^+) - S_A(\omega_n^-)), \end{aligned}$$

Key difference from previous work: We use the SEMD, *not* the EMD!

Step 4: Minimize w.r.t. parameters using grads



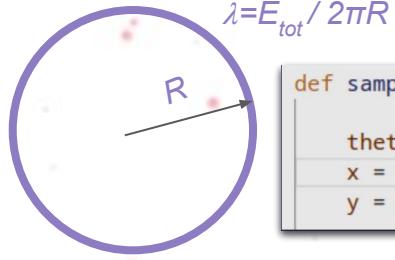
Pictured: Animation of optimizing for the radius R

SPECTER is a “sequel” to SHAPER, introduced last ML4Jets. SPECTER is *not* an acronym, don’t ask me what it stands for.

Full Example: How “ring-like” are jets?

SPECTER Our code framework for these calculations

Step 1: Define the shape with parameters

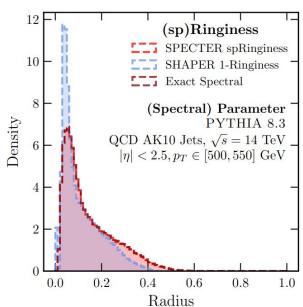
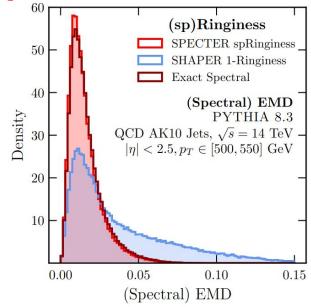


Unlike ordinary EMD, not necessary to specify center / orientation!

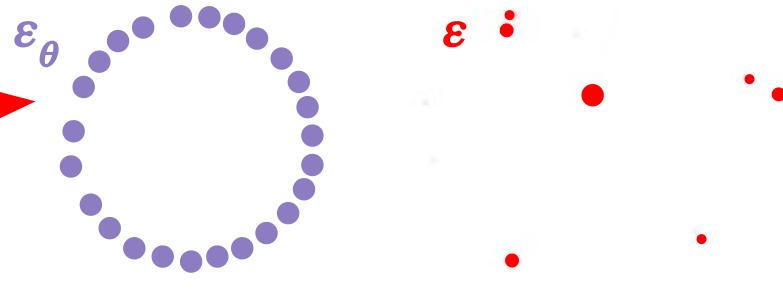
```
def sample_circle(params, N, seed):
    thetas = jax.random.uniform(seed, shape)
    x = params["Radius"] * jnp.cos(thetas)
    y = params["Radius"] * jnp.sin(thetas)
```

Pictured: 100k Jets, PYTHIA 8 QCD Jets

Step 5: Plots!



Step 2: Sample from Parameterized Shapes

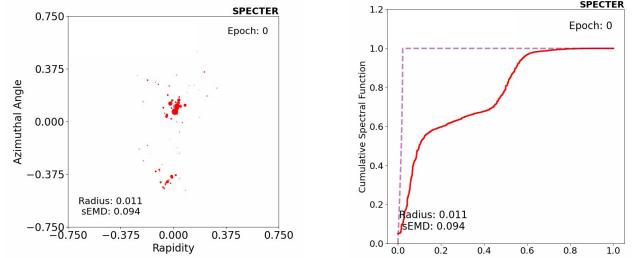


Step 3: Calculate the spectral metric between events and shapes

$$\text{SEMD}_{\beta,p=2}(s_A, s_B) = \sum_{i < j \in \mathcal{E}_A} 2E_i E_j \omega_{ij}^2 + \sum_{i < j \in \mathcal{E}_B} 2E_i E_j \omega_{ij}^2 - 2 \sum_{n \in \mathcal{E}_A^2, l \in \mathcal{E}_B^2} \omega_n \omega_l (\min [S_A(\omega_n^+), S_B(\omega_l^+)] - \max [S_A(\omega_n^-), S_B(\omega_l^-)]) \times \Theta(S_A(\omega_n^+) - S_B(\omega_l^-)) \Theta(S_B(\omega_l^+) - S_A(\omega_n^-)),$$

Key difference from previous work: We use the SEMD, *not* the EMD!

Step 4: Minimize w.r.t. parameters using grads



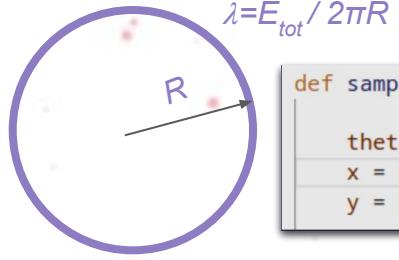
Pictured: Animation of optimizing for the radius R

Full Example: How “ring-like” are jets?

SPECTER

Our code framework
for these calculations

Step 1: Define the shape with parameters



```
def sample():
    theta = np.random.uniform(0, 2 * np.pi)
    x = p[0] + R * np.cos(theta)
    y = p[1] + R * np.sin(theta)
```

Unlike ordinary EMD, not necessary to sample points

Alternatively...

The spectral EMD, and its optimization, are often partially or **completely solvable** in closed form!

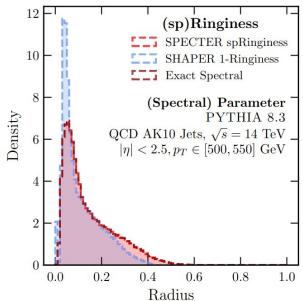
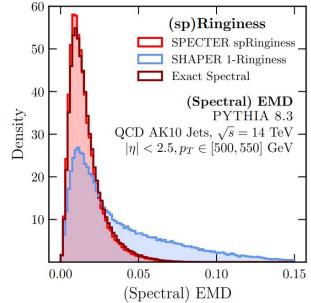
$$R_{\text{opt}} = \frac{2}{\pi} \sum_{\substack{n \in \mathcal{E}^2 \\ \omega_n < \omega_{n+1}}} \omega_n \left[\sin \left(\frac{\pi}{2E_{\text{tot}}^2} \sum_{\substack{m \in \mathcal{E}^2 \\ \omega_m < \omega_{m+1}}} (2EE)_m \right) - \sin \left(\frac{\pi}{2E_{\text{tot}}^2} \sum_{\substack{m \in \mathcal{E}^2 \\ \omega_m < \omega_{m+1}}} (2EE)_m \right) \right]$$

$$\text{SEMD}_{\beta,p=2}(s, s_{\text{jet ring}}) = \sum_{i < j \in \mathcal{E}} 2E_i E_j \omega_{ij}^2 - 2E_{\text{tot}}^2 R_{\text{opt}}^2$$

For many shapes, we can completely short circuit having to perform expensive optimization over an optimal transport problem entirely!

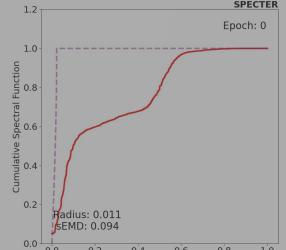
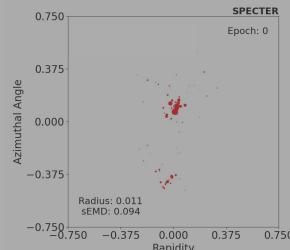
Pictured: 100k Jets, PYTHIA 8 QCD Jets

Step 5: Plots!



Key difference from previous work: We use the SEMD, not the EMD!

Step 4: Minimize w.r.t. parameters using grads



Pictured: Animation of optimizing for the radius R

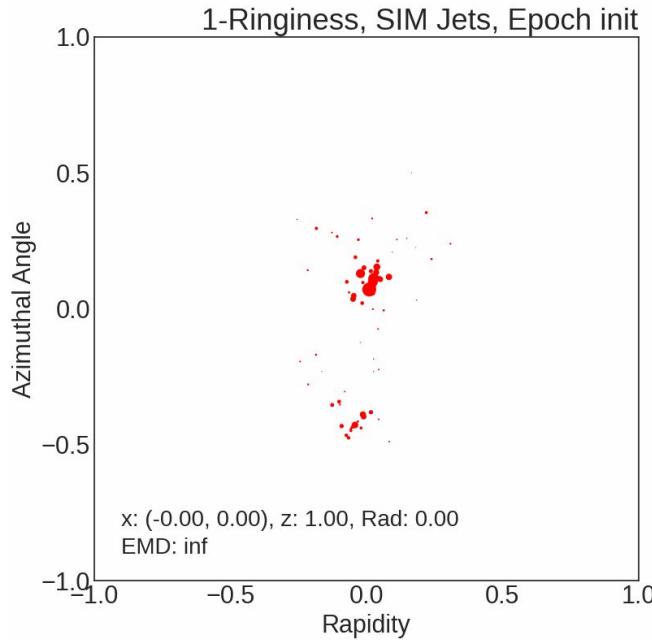
To distinguish SEMD observables from EMD observables, I will add “s” or “sp”

Hearing Jets (sp)Ring

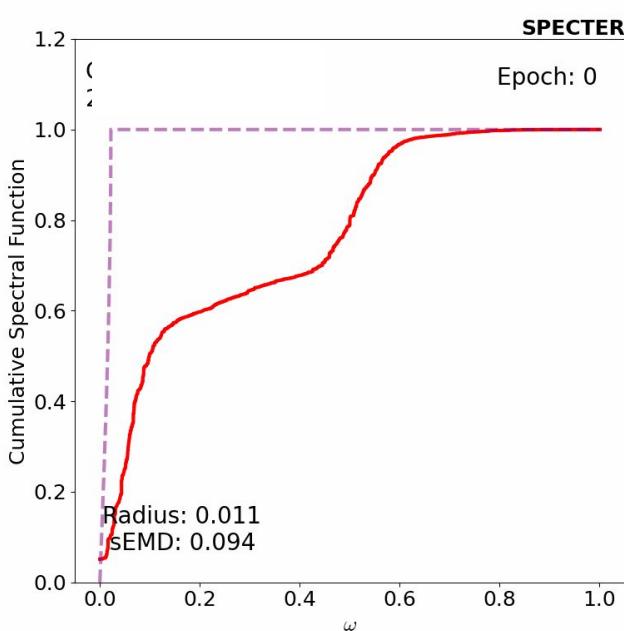
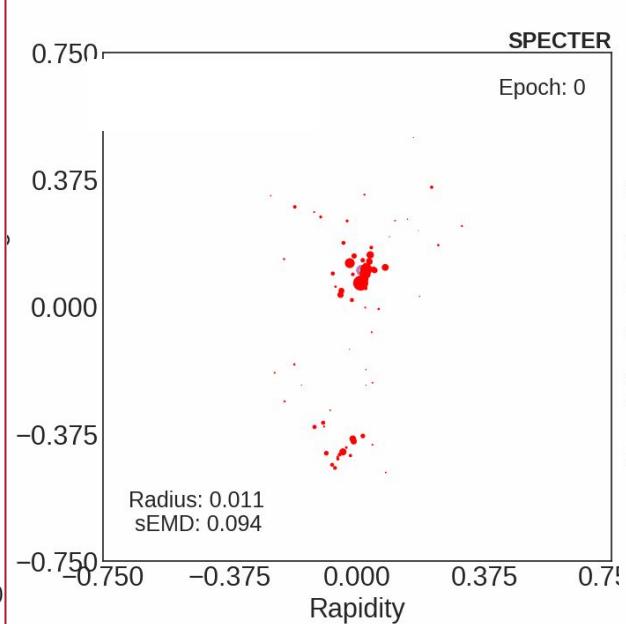
Small R Jet

Large R Jet

EMD



Spectral EMD



Calculated using SHAPER¹
Position of ring must be optimized – can use as jet algorithm

Calculated using **SPECTER**
Translationally invariant – no need to optimize over position
Secretly a 1D optimal transport problem over the spectral function

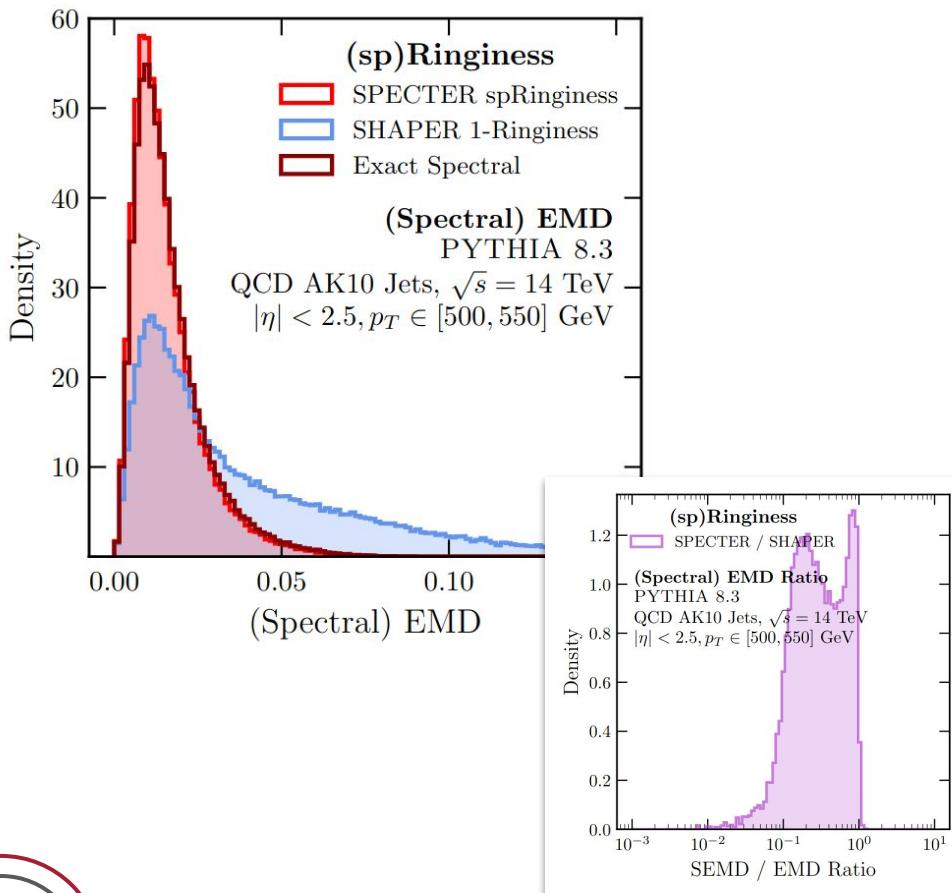
Hearing Jets (sp)Ring

Runtimes (NVIDIA A100 GPU):

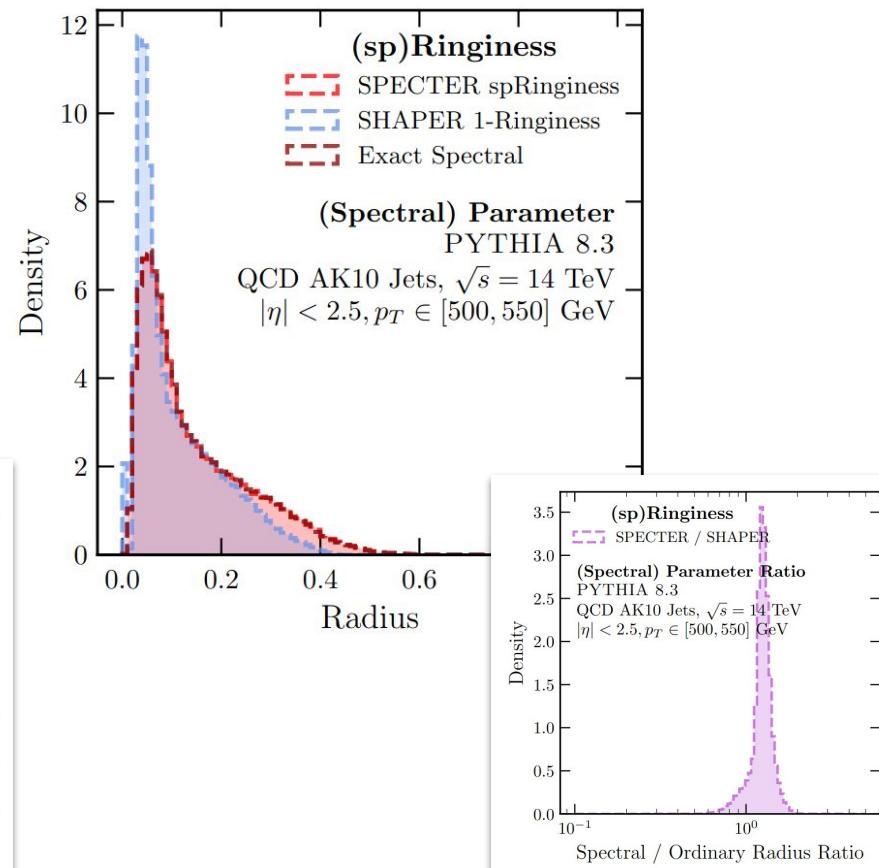
SHAPER (EMD): ~ 36 hours / 100k events

Generalized SPECTER: ~55 seconds / 100k events

Closed Form SPECTER: ~ < 0.1 seconds / 100k events



The SEMD and EMD are qualitatively different, but give similar radii!
They probe the same event length scale

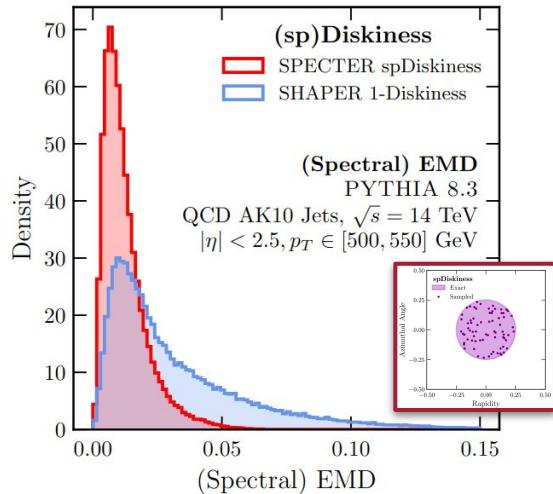


Lots of Observables!

Event and jet shape observables can be defined as the (S)EMD between events and any parameterized set of ideal events!

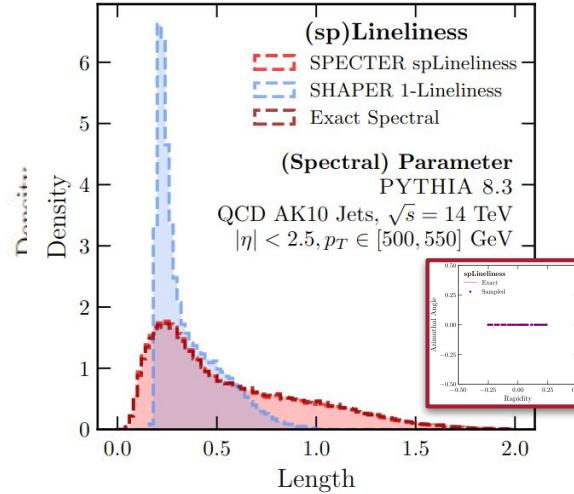
Some examples ...

How disk-like are jets?



[No closed form for sDiskiness – not all observables have closed forms]

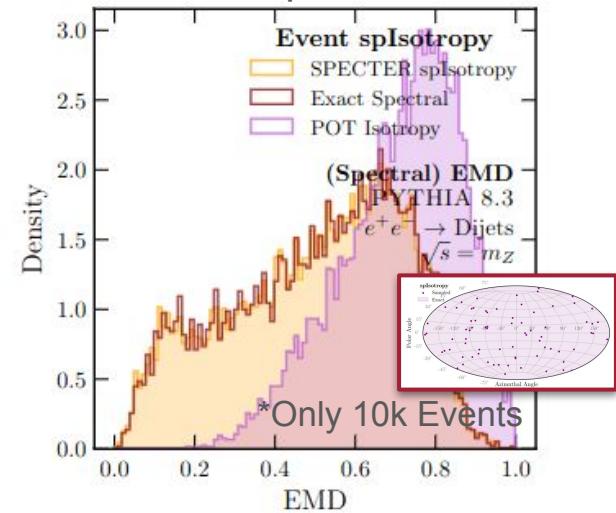
If QCD jets were lines, how long would they be?



Closed-form lines!

Everything I said today applies to full events on the celestial sphere as well as localized jets!
 Different topologies are possible!

How isotropic¹ are events?



$$\mathcal{O}_{\text{Isotropy}}(s_{\mathcal{E}}) = \sum_{i < j \in \mathcal{E}} 2E_i E_j \omega_{ij}^2 + \frac{\pi^2 - 4}{2} E_{\text{tot}}^2 - 2 \sum_{\substack{n \in \mathcal{E}^2 \\ \omega_n < \omega_{n+1}}} \omega_n [f^+(n) - f^-(n)]$$

$$- S^-(\omega_n) - \frac{2}{3E_{\text{tot}}} (E_{\text{tot}}^2 - S^-(\omega_n))^{3/2}]$$

$$f^{\pm}(n) = \sqrt{S^{\pm}(\omega_n)} \sqrt{E_{\text{tot}}^2 - S^{\pm}(\omega_n)} + S^{\pm}(\omega_n) \cos^{-1} \left(1 - 2 \frac{S^{\pm}(\omega_n)}{E_{\text{tot}}^2} \right)$$

$$- E_{\text{tot}}^2 \sin^{-1} \left(\frac{\sqrt{S^{\pm}(\omega_n)}}{E_{\text{tot}}} \right)$$

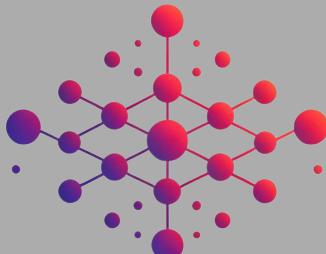
Closed-form isotropy!

Things to think about:

- **Speed**: I am not a great computer scientist; *SPECTER* could probably be made even faster with more clever and better programming.
- **Degeneracies and Topology**: The EMD and SEMD are different, especially for equilateral triangle configurations – how often do these configurations occur in different theories?
- **Closed form Observables**: Not every shape has a completely closed-form solution, but it is usually possible to partially simplify and reduce the problem to 1D minimization, 1D root finding, or simple 1D numeric integrals. Can we understand this better?
- **Perturbative Calculations**: Closed-form and simple expressions means perturbative calculations may be possible – can we predict the radius of a jet to LO, NLO, LL, NLL, ...?
- **Theory Space**: There have been proposals to use the (S)EMD between events as a ground metric for an OT distance between theories. With *SPECTER*, this could now be numerically viable!

Happy to talk with you about any and all of these afterwards!

Conclusion



SPECTER

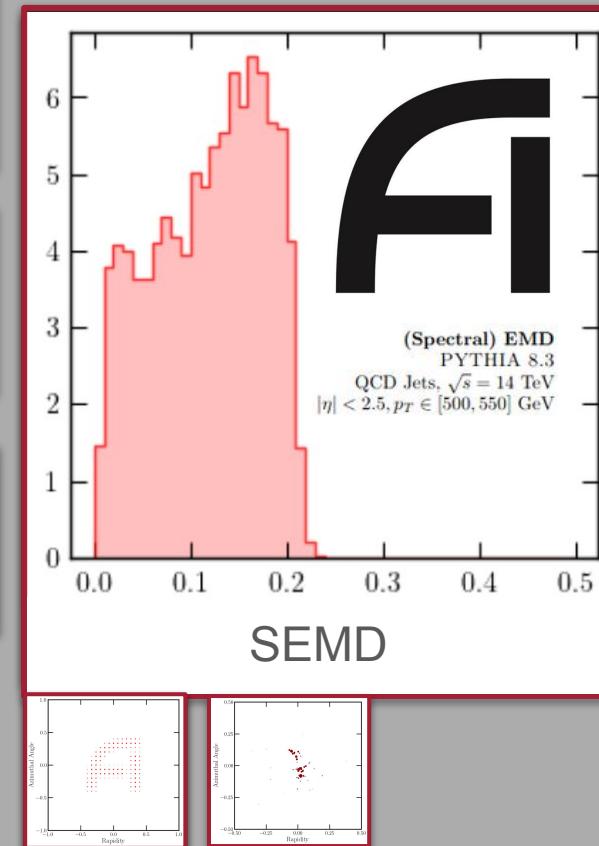
Pictured: The **spectral-IAIFI-ness** of QCD Jets!

The **spectral EMD** can be used as an alternative to the EMD. It is **fast** and **easy to minimize**.

SPECTER is a code package for efficiently evaluating the spectral EMD and calculating shape observables.

With the spectral EMD, many jet observables can be understood in **closed form**.

SPECTER will be *pip*-installable!
Coming soon!



More questions? Email me at rikab@mit.edu

Appendices

The EMD

Definition:

$$\text{EMD}_{\beta,R}(\mathcal{E}_A, \mathcal{E}_B) = \min_{\{f_{ab}\}} \sum_{a \in J_A} \sum_{b \in J_B} f_{ab} \frac{\Omega(\hat{n}_a, \hat{n}_b)^\beta}{R^\beta} + \left| \sum_{a \in J_A} E_a - \sum_{b \in J_B} E_b \right|$$

such that

$$f_{ab} \geq 0, \quad \sum_{b \in J_B} f_{ab} \leq E_a, \quad \sum_{a \in J_A} f_{ab} \leq E_b, \quad \sum_{a \in J_A} \sum_{b \in J_B} f_{ab} = \min \left(\sum_{a \in J_A} E_a, \sum_{b \in J_B} E_b \right)$$

Ground Metrics

For local jets on the rapidity-azimuth plane:

$$\omega_{ij} = \sqrt{(\phi_i - \phi_j)^2 + (y_i - y_j)^2}$$

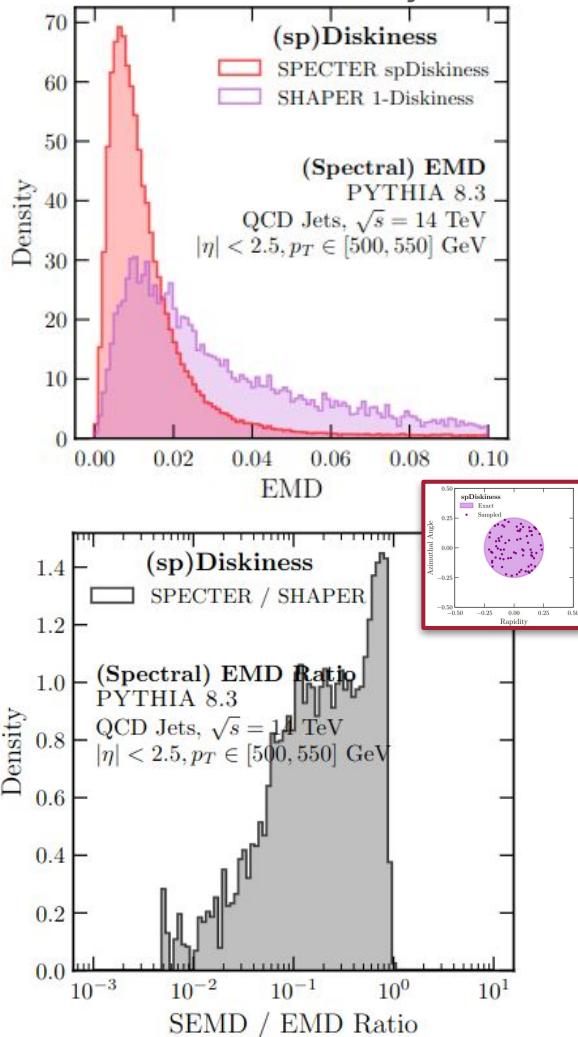
For global events on the sphere:

$$\begin{aligned}\omega_{ij} &= |\theta_{ij}| \\ &= \left| \cos^{-1} \left(1 - \frac{\mathbf{p}_i \cdot \mathbf{p}_j}{E_i E_j} \right) \right|\end{aligned}$$

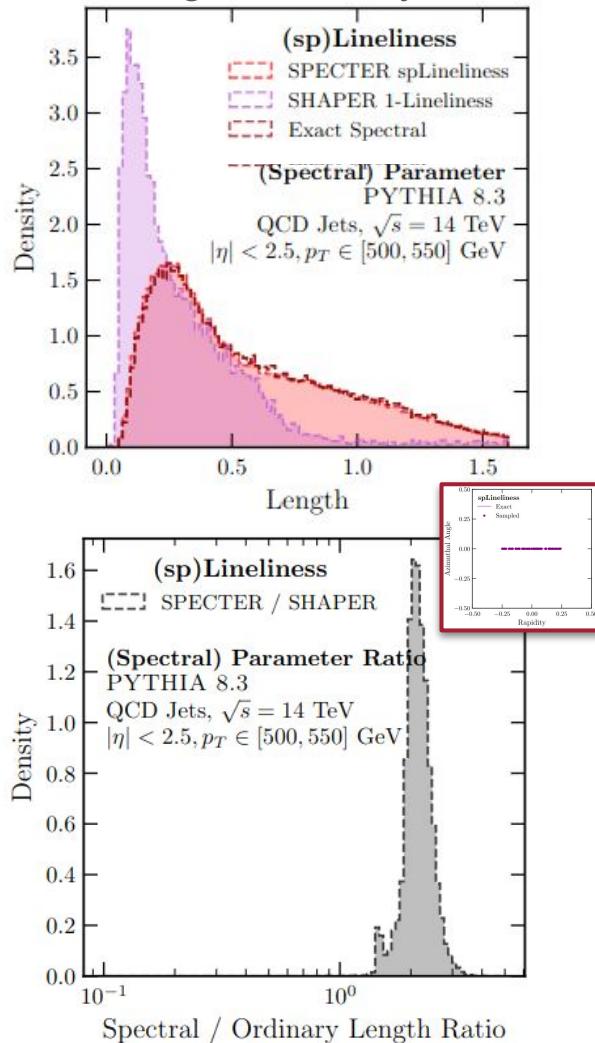
In principle, could have picked chord length rather than arc length

SEMD to EMD Ratios

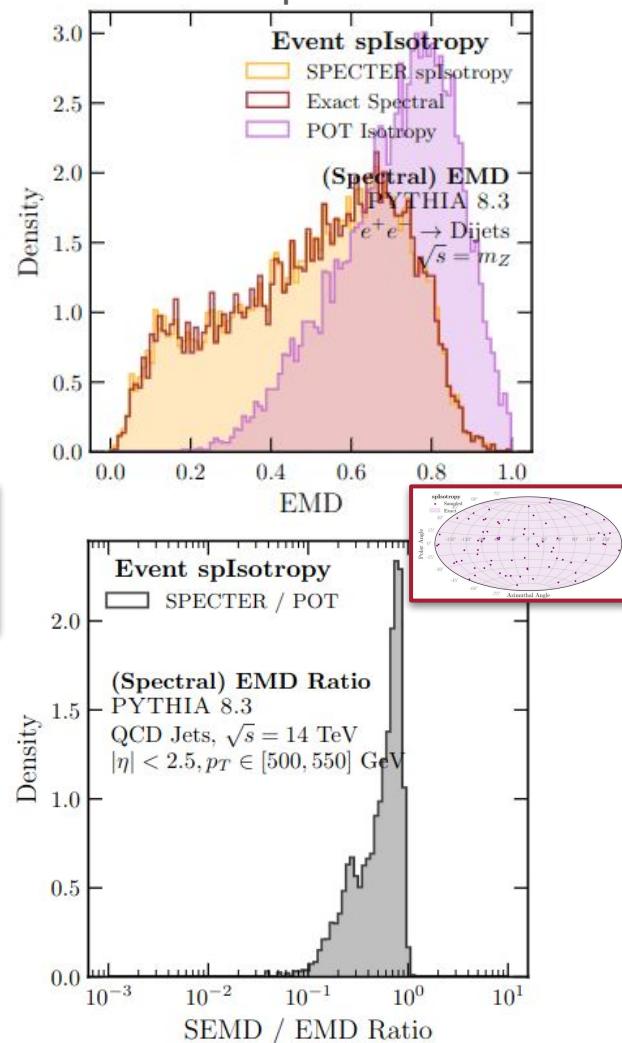
How disk-like are jets?



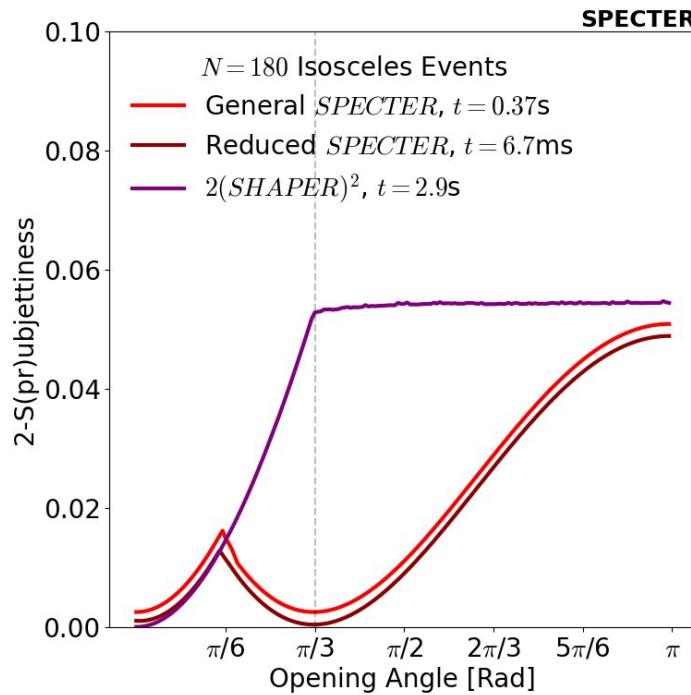
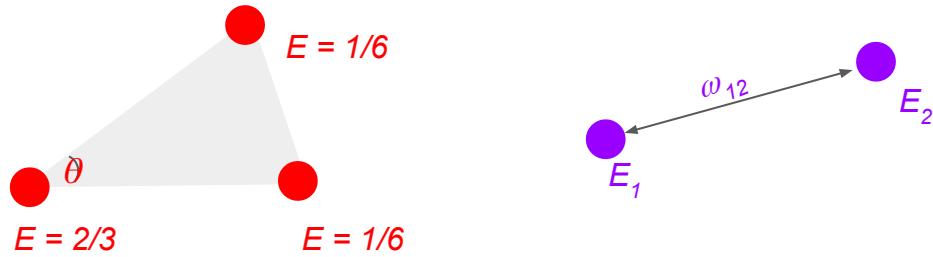
If QCD jets were lines, how long would they be?



How isotropic¹ are events?



Degeneracies (Continued)



For this precise energy configuration, equilateral triangles are *exactly* degenerate with 2 particle events – so the spectral EMD only sees 2 particles!

Only measure 0 configuration of events – but events *near* this give spectral EMDs *near* zero against 2 particle events.

*with the right energy weights.

Shapiness

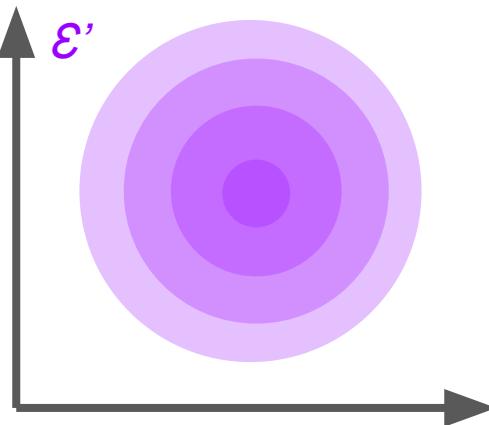
The EMD between a real event or jet \mathcal{E} and idealized shape \mathcal{E}' is the [shape]iness of \mathcal{E} – a well defined IRC-safe observable!

Answers the question:
“How much like the
shape \mathcal{E}' is my event \mathcal{E} ? ”

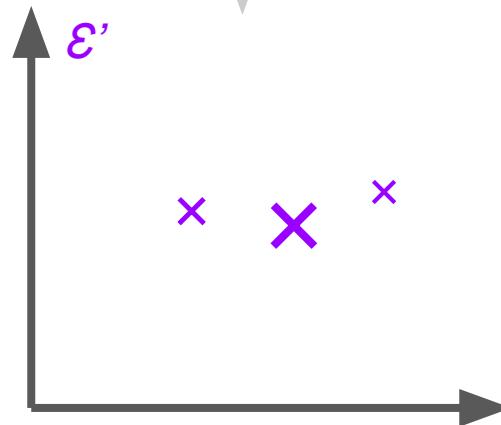
Med EMD($\mathcal{E}, \mathcal{E}'$)
“Gausiness”

Low EMD($\mathcal{E}, \mathcal{E}'$)
“3-Pointiness”
AKA “3-Subjettiness”

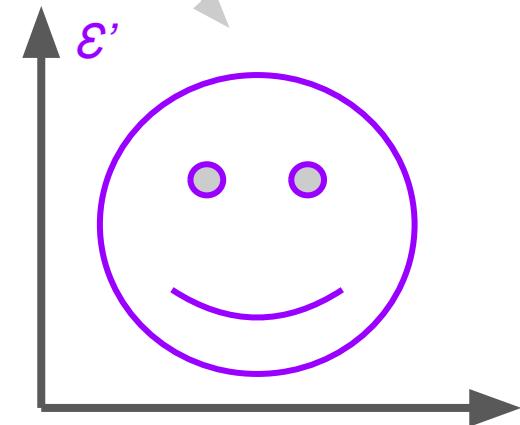
High EMD($\mathcal{E}, \mathcal{E}'$)
“Smileyness”



Shape = 2D Gaussian



Shape = 3 Points



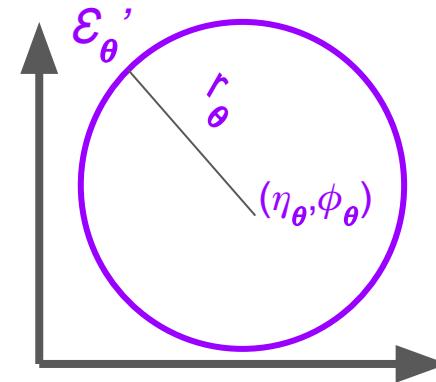
Shape = Smile

Mathematical Details - Shapiness

Rather than a single shape, consider a **parameterized manifold \mathcal{M} of energy flows**.

e.g. The manifold of uniform circle energy flows:

$$\mathcal{E}_\theta'(y) = \begin{cases} \frac{1}{2\pi r_\theta} & |\vec{y} - \vec{y}_\theta| = r_\theta \\ 0 & |\vec{y} - \vec{y}_\theta| \neq r_\theta \end{cases}$$



Then, for an event \mathcal{E} , define the **shapiness $\mathcal{O}_{\mathcal{M}}$** and **shape parameters $\theta_{\mathcal{M}}$** , given by:

$$\mathcal{O}_{\mathcal{M}}(\mathcal{E}) \equiv \min_{\mathcal{E}_\theta \in \mathcal{M}} \text{EMD}^{(\beta, R)}(\mathcal{E}, \mathcal{E}_\theta)$$

$$\theta_{\mathcal{M}}(\mathcal{E}) \equiv \operatorname{argmin}_{\mathcal{E}_\theta \in \mathcal{M}} \text{EMD}^{(\beta, R)}(\mathcal{E}, \mathcal{E}_\theta)$$

Observables \leftrightarrow Manifolds of Shapes

Observables can be specified solely by defining a **manifold of shapes**:

$$\mathcal{O}_{\mathcal{M}}(\mathcal{E}) \equiv \min_{\mathcal{E}_\theta \in \mathcal{M}} \text{EMD}^{(\beta, R)}(\mathcal{E}, \mathcal{E}_\theta),$$

$$\theta_{\mathcal{M}}(\mathcal{E}) \equiv \operatorname{argmin}_{\mathcal{E}_\theta \in \mathcal{M}} \text{EMD}^{(\beta, R)}(\mathcal{E}, \mathcal{E}_\theta),$$

Many well-known observables^{*} already have this form!

Observable	Manifold of Shapes
N -Subjettiness	Manifold of N -point events
N -Jettiness	Manifold of N -point events with floating total energy
Thrust	Manifold of back-to-back point events
Event / Jet Isotropy	Manifold of the single uniform event
	... and more!

All of the form “How much like **shape** does my **event** look like?”

Generalize to *any* shape.

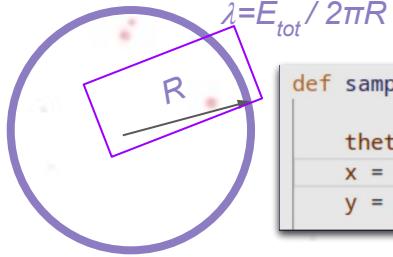
^{*}These observables are usually called event shapes or jet shapes in the literature – we are making this literal!

Full Example: How “ring-like” are jets?

SPECTER

Our code framework
for these calculations

Step 1: Define the shape with parameters

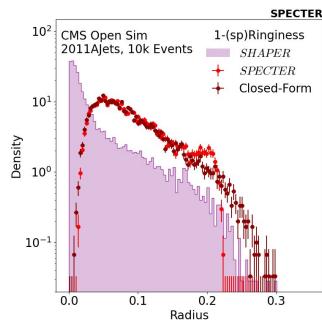
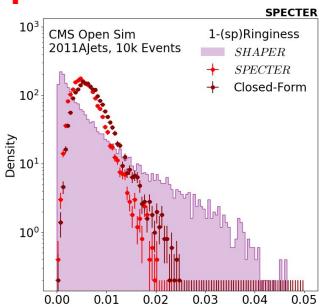


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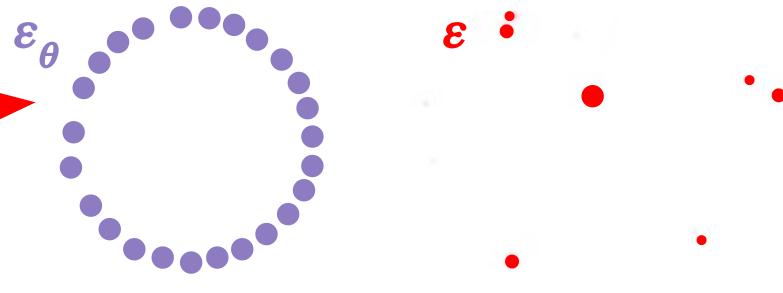
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    x = params["Radius"] * jnp.cos(thetas)
    y = params["Radius"] * jnp.sin(thetas)
```

Pictured: 10k Jets, CMS 2011AJets Open Sim

Step 5: Plots!



Step 2: Sample from Parameterized Shapes

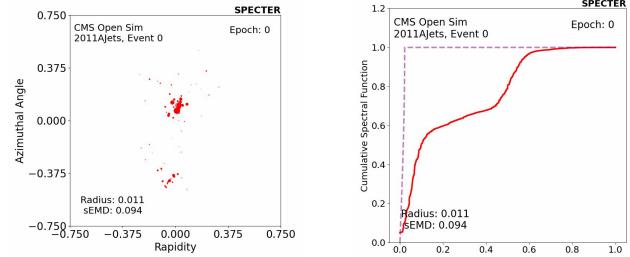


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$$\begin{aligned} \text{SEMD}_{\beta,p=2}(s_A, s_B) = & \sum_{i < j \in \mathcal{E}_A} 2E_i E_j \omega_{ij}^2 + \sum_{i < j \in \mathcal{E}_B} 2E_i E_j \omega_{ij}^2 \\ & - 2 \sum_{n \in \mathcal{E}_A^2, l \in \mathcal{E}_B^2} \omega_n \omega_l (\min [S_A(\omega_n^+), S_B(\omega_l^+)] - \max [S_A(\omega_n^-), S_B(\omega_l^-)]) \\ & \times \Theta(S_A(\omega_n^+) - S_B(\omega_l^-)) \Theta(S_B(\omega_l^+) - S_A(\omega_n^-)), \end{aligned}$$

Key difference from previous work: We use the SEMD, *not* the EMD!

Step 4: Minimize w.r.t. parameters using grads



Pictured: Animation of optimizing for the radius R

CPU Runtimes

