#### In [49]:

```
1
   import astropy
   from astropy.io import fits
 2
   from astropy.table import Table
 3
   import matplotlib.pyplot as plt
 5
   import numpy as np
   from astropy import constants as con
 7
   from astropy import units as u
 8
   from astropy.io import ascii
   import pandas as pd
9
10 from scipy.interpolate import interpld
   from mpl toolkits import mplot3d
11
12 from matplotlib import animation
13
   from mpl toolkits.mplot3d import Axes3D
   %matplotlib inline
14
```

#### In [50]:

```
1 yilun = ascii.read('spt_to_teff.txt')
2 np.unique(yilun[0])
```

#### Out[50]:

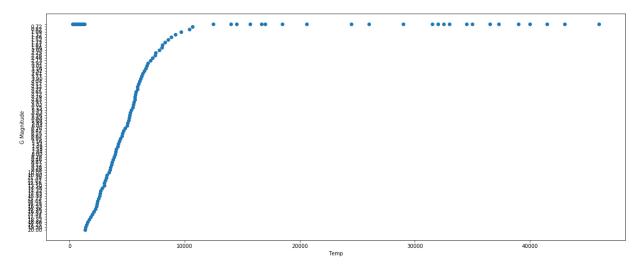
#### In [51]:

```
#gband mag as a funct of temp
plt.figure(figsize=(20,8))
temp = yilun['Teff']

m_gorl = yilun['M_G']
plt.gca().invert_yaxis()
plt.scatter(temp, m_gorl)
plt.xlabel('Temp')
plt.ylabel('G_Magnitude')
#plt.hlines(0,0,40000)
```

#### Out[51]:

Text(0, 0.5, 'G Magnitude')



```
In [52]:
```

```
1 len(temp)
```

#### Out[52]:

124

# In [53]:

```
fixed_temp = np.array([])
fixed_mgorl = np.array([])
for i in np.arange(0, 124):
    if temp[i] > 1300:
        fixed_temp = np.append(fixed_temp, temp[i])
        fixed_mgorl = np.append(fixed_mgorl, m_gorl[i])
        #print(fixed_temp)
        #print(fixed_mgorl)
```

```
In [56]:
```

```
1 '''
2 plt.figure(figsize=(20,8))
3 plt.scatter(fixed_temp, fixed_mgorl)
4 plt.plot(np.linspace(1350, 4600, 1000), f(np.linspace(1350, 4600, 1000)))
5 plt.gca().invert_yaxis()
6 plt.xlabel('Effective Temperature, in K')
7 plt.ylabel('G Magnitude')
8 plt.title('Relation Between Teff and G Mag')
9 '''
```

#### Out[56]:

"\nplt.figure(figsize=(20,8))\nplt.scatter(fixed\_temp, fixed\_mgorl)\np
lt.plot(np.linspace(1350, 4600, 1000), f(np.linspace(1350, 4600, 100
0)))\nplt.gca().invert\_yaxis()\nplt.xlabel('Effective Temperature, in
K')\nplt.ylabel('G Magnitude')\nplt.title('Relation Between Teff and G
Mag')\n"

```
In [57]:
```

1 np.sort(temp)

# Out[57]:

<Column name='Teff' dtype='int64' length=124>

• • •

# In [58]:

```
data = Table.read('asu.fit')
print(data)
```

_r	_RAJ2000	_DEJ2000	Source	• • •	Teff	Rad
Lum arcmin	deg	deg			K	solRad
solLum						
				• • •		
0.0116	245.89664236128	-26.5255826857852	6631424	• • •	0.0	0.0
0.0	245.8971323101844	-26.5257618225833	6536960		0.0	0.0
0.0	213.0371323101011	20.3237010223033	0330700	•••	0.0	0.0
0.0265	245.896769646258	-26.5260997864899	6539264	• • •	0.0	0.0
0.0 0.0301	245 896858746915	-26.525258143415	6631808		0.0	0.0
0.0301	243.030030740313	-20:323230143413	0031000	• • •	0.0	0.0
0.0415	245.8973597661302	-26.5261756755971	6539392		5095.63	0.0
0.0	0.45 0.071060572011	06 5064160504000	6520640		0 0	0 0
0.0443	245.8971060573211	-26.5264169504993	6539648	• • •	0.0	0.0
	245.8959742841654	-26.5255185055872	6631936		5143.45	0.0
0.0						
	245.8969773452325	-26.5264051264131	6538368	• • •	0.0	0.0
0.0	245.8975538856744	-26.5250953593613	6632192		4508.75	0.0
0.0	243.03/3330030/44	20.3230733373013	0032172	•••	4500.75	0.0
	245.8961584044747	-26.5263737945874	6539008		5143.45	0.0
0.0						
	• • •	• • •	• • •	• • •	• • •	• • •
0.9945	245.8878103617531	-26.5401557475941	16593536		5076.0	0.0
0.0						
0.9945	245.8882688926921	-26.5403794827975	6361472	• • •	0.0	0.0
	245.8827857078983	-26.5148357156183	7046656		0.0	0.0
0.0			,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,			
	245.9072999693412	-26.5393939638479	6175360	• • •	0.0	0.0
0.0	245 070401074076	-26.52827501348	7761702		5000.0	0.99
0.9933	245.0/04910/40/0	-20.5262/501546	7701792	• • •	3000.0	0.99
	245.8840173611604	-26.5136539133512	7079424		0.0	0.0
0.0						
	245.8805703431035	-26.5337641229179	16642944	• • •	5143.45	0.0
0.0	245 9146830999414	-26.530264331309	6549376		4650 03	1.62
1.108	243.7140030777414	-20.330204331307	0343370	• • •	4030.03	1.02
0.9994	245.9126454428589	-26.516903343388	6606976		0.0	0.0
0.0	045 045406400555	0.6	628022		4000	2 5
0.9996	245.9154261835508	-26.5255252911882	6378880	• • •	4920.0	0.0
	= 1913 rows					
-						

```
final project python pt 2 - Jupyter Notebook
In [59]:
   np.unique(data[:0])
Out[59]:
array([],
      dtype=[(' r', '<f8'), (' RAJ2000', '<f8'), (' DEJ2000', '<f8'),</pre>
('Source', '<i4'), ('FG', '<f8'), ('e_FG', '<f8'), ('Gmag', '<f8'),
('e_Gmag', '<f8'), ('BPmag', '<f8'), ('e_BPmag', '<f8'), ('RPmag', '<f
8'), ('e_RPmag', '<f8'), ('BP-RP', '<f8'), ('Teff', '<f8'), ('Rad', '<
f8'), ('Lum', '<f8')])
In [60]:
   bprp arr1 = np.array(data['BP-RP'])
In [61]:
    lum arr = np.array(data['Lum'])
In [62]:
 1 teff arr = np.array(data['Teff'])
   print(len(teff_arr))
1913
In [63]:
   dat arr = np.array(data['Gmag'])
In [64]:
    bprp_array = np.array([])
    gmag_array = np.array([])
 3
    lum array = np.array([])
 4
    teff_array = np.array([])
    for i in data['BP-RP']:
 5
 6
        if i != 0.0:
 7
            bprp array = np.append(bprp array, i)
 8
 9
    for k in np.arange(0, 1913):
```

gmag\_array = np.append(gmag\_array, dat\_arr[k])

teff\_array = np.append(teff\_array, teff\_arr[k])

lum array = np.append(lum array, lum arr[k])

**if** bprp arr1[k] != 0.0:

10

11

12

13

```
In [65]:
```

```
1 len(teff_array)
```

## Out[65]:

841

#### In [66]:

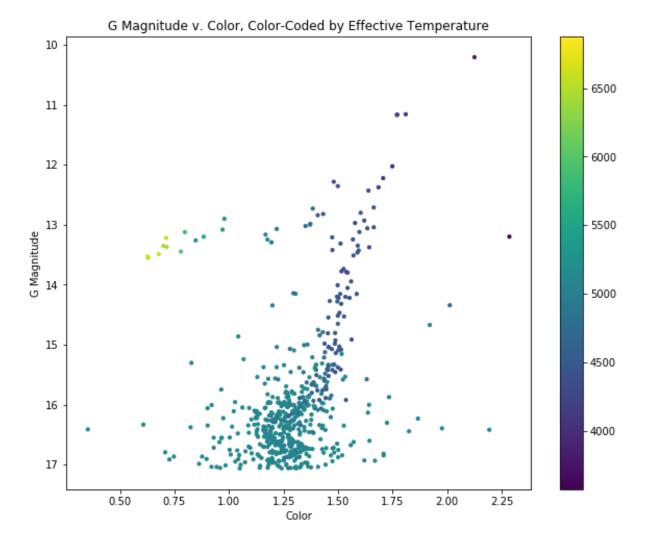
```
finalteff_array = np.array([])
   finalbprp array = np.array([])
   finalgmag_array = np.array([])
   finallum_array = np.array([])
 5
   for k in np.arange(0, 841):
 6
       if teff array[k] != 0.0:
 7
            finalgmag array = np.append(finalgmag array, gmag array[k])
 8
           finallum array = np.append(finallum array, lum array[k])
 9
           finalteff array = np.append(finalteff array, teff array[k])
10
            finalbprp_array = np.append(finalbprp_array, bprp_array[k])
   print(len(finalgmag array))
```

## In [67]:

```
plt.figure(figsize=(10,8))
plt.scatter(finalbprp_array, finalgmag_array, c = finalteff_array, s = 10)
plt.gca().invert_yaxis()
plt.xlabel('Color')
plt.ylabel('G Magnitude')
plt.title('G Magnitude v. Color, Color-Coded by Effective Temperature')
plt.colorbar()
```

## Out[67]:

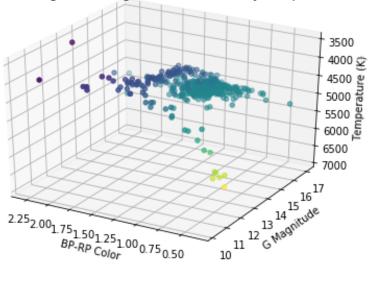
<matplotlib.colorbar.Colorbar at 0x7f25fc7b1d60>



## In [92]:

```
1  ax = plt.axes(projection ="3d")
2  ax.invert_yaxis()
3  ax.invert_zaxis()
4  ax.scatter3D(finalbprp_array, finalgmag_array, finalteff_array, c = finalteff_a
5  ax.set_xlabel('BP-RP Color')
6  ax.set_ylabel('G Magnitude')
7  ax.set_zlabel('Temperature (K)')
8  ax.set_title('3D Color-Magnitude Diagram, Color-Coded by Temperature')
9  plt.tight_layout()
```

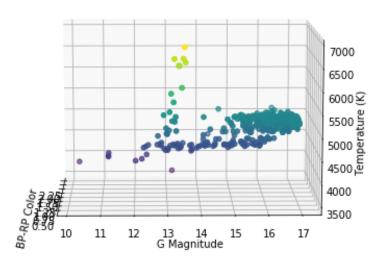
## 3D Color-Magnitude Diagram, Color-Coded by Temperature



#### In [95]:

```
1
   # Create a figure and a 3D Axes
 2
 3
   fig = plt.figure()
 4
   ax = Axes3D(fig)
 5
   # Create an init function and the animate functions.
 6
 7
   # Both are explained in the tutorial. Since we are changing
   # the the elevation and azimuth and no objects are really
   # changed on the plot we don't have to return anything from
 9
   # the init and animate function. (return value is explained
10
   # in the tutorial.
11
12
   def init():
       ax.invert yaxis()
13
       ax.invert zaxis()
14
       ax.scatter(finalbprp array, finalgmag array, finalteff array, c = finalteff
15
16
       ax.set xlabel('BP-RP Color')
17
18
       ax.set ylabel('G Magnitude')
19
       ax.set_zlabel('Temperature (K)')
       ax.set title('3D Color-Magnitude Diagram, Color-Coded by Temperature')
20
21
       return fig,
22
23
   def animate(i):
24
       ax.view init(elev=10., azim=i)
25
       return fig,
26
   # Animate
27
   anim = animation.FuncAnimation(fig, animate, init func=init,
28
29
                                   frames=360, interval=20, blit=True)
   # Save
30
31
   anim.save('finalproj anim1.mp4', fps=30)
```

3D Color-Magnitude Diagram, Color-Coded by Temperature

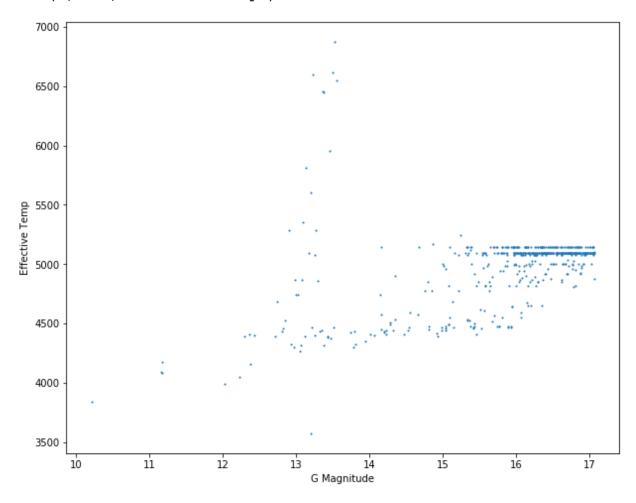


# In [87]:

```
plt.figure(figsize=(10,8))
plt.scatter(finalgmag_array, finalteff_array, s=1)
plt.xlabel('G Magnitude')
plt.ylabel('Effective Temp')
```

# Out[87]:

Text(0, 0.5, 'Effective Temp')



# In [35]:

```
1 len(lum_array)
```

# Out[35]:

## In [36]:

```
plt.figure(figsize=(10,8))
plt.gca().invert_yaxis()

plt.scatter(bprp_array, gmag_array, s=1)

plt.xlabel('BP-RP Color')

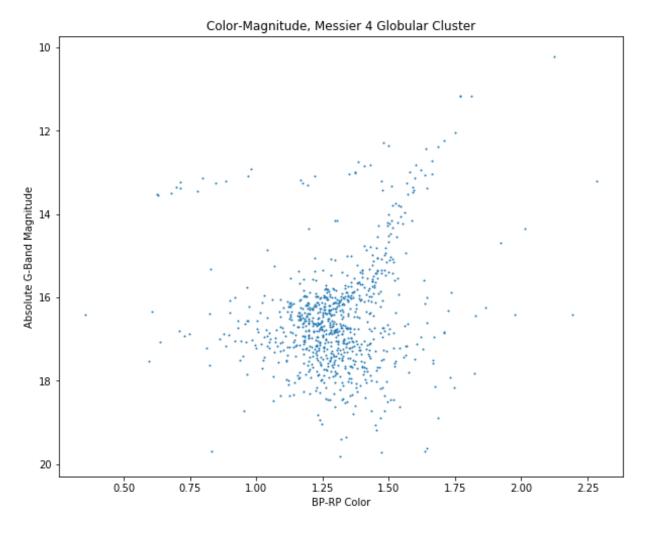
plt.ylabel('Absolute G-Band Magnitude')

plt.title('Color-Magnitude, Messier 4 Globular Cluster')

#try color coding by temp
#try color coding by temp
```

## Out[36]:

Text(0.5, 1.0, 'Color-Magnitude, Messier 4 Globular Cluster')

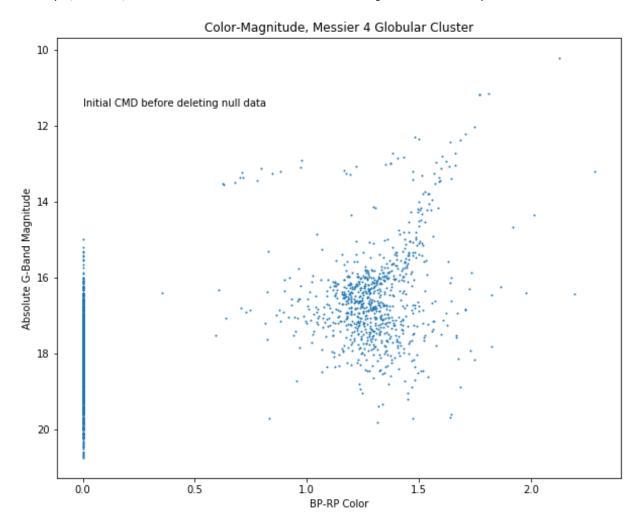


#### In [37]:

```
plt.figure(figsize=(10,8))
plt.gca().invert_yaxis()
plt.scatter(data['BP-RP'], data['Gmag'], s=1)
plt.xlabel('BP-RP Color')
plt.ylabel('Absolute G-Band Magnitude')
plt.title('Color-Magnitude, Messier 4 Globular Cluster')
plt.text(0, 11.5, 'Initial CMD before deleting null data')
```

## Out[37]:

Text(0, 11.5, 'Initial CMD before deleting null data')



```
In [38]:
```

```
1 Teff_array = np.array(data['Teff'])
2 print(len(Teff_array))
```

```
In [39]:
    print(con.sigma_sb)

Name = Stefan-Boltzmann constant
Value = 5.6703744191844314e-08
Uncertainty = 0.0
Unit = W / (K4 m2)
Reference = CODATA 2018

In [40]:
    sigmasb = 5.6703744191844314 * 10**(-8)

In [41]:
    len(teff_array)
Out[41]:
```

841

## In [42]:

```
1 teff = teff_array * u.K
2 oldlum = lum_array * con.L_sun
3 oldlum.to(u.W)
```

## Out[42]:

```
[0, 0, 0, \dots, 4.241424 \times 10^{26}, 0, 0] \text{ W}
```

```
In [43]:
   lum = oldlum / u.W
 2
   print(lum)
 3
   for n in lum:
 4
       if n != 0.00000000e+00:
 5
           print(n)
[0.00000000e+00 0.0000000e+00 0.0000000e+00 0.0000000e+00
0.0000000e+00 1.36621320e+27 0.0000000e+00 0.0000000e+00
 3.85862400e+26 0.00000000e+00 1.24367892e+28 0.00000000e+00
0.0000000e+00 0.00000000e+00 2.05257360e+27 0.00000000e+00
0.0000000e+00 0.0000000e+00 0.0000000e+00 0.0000000e+00
 4.15184880e+27 0.00000000e+00 4.34784240e+27 5.56744320e+27
0.0000000e+00 0.0000000e+00 0.0000000e+00 0.0000000e+00
 3.35876376e+28 0.00000000e+00 0.0000000e+00 1.21730400e+26
0.0000000e+00 0.00000000e+00 5.49700800e+26 0.00000000e+00
 0.0000000e+00 0.0000000e+00 0.0000000e+00 0.0000000e+00
0.0000000e+00 0.0000000e+00 0.0000000e+00 0.0000000e+00
0.0000000e+00 0.0000000e+00 0.0000000e+00 5.25201600e+26
0.0000000e+00 0.00000000e+00 1.90584636e+28 0.00000000e+00
 0.0000000e+00 0.0000000e+00 0.0000000e+00 0.0000000e+00
 0.0000000e+00 0.0000000e+00 0.0000000e+00 0.0000000e+00
In [44]:
   rad_array = np.array([])
 2
   for j in np.arange(0, 841):
 3
       r = (lum_array[j] / (4 * np.pi * sigmasb * (teff_array[j]**4)))**0.5
 4
       rad array = np.append(rad array, r)
   #r = (lum array / (4 * np.pi * con.sigma sb * teff array**4))**0.5
<ipython-input-44-650e08bd932c>:3: RuntimeWarning: invalid value encou
ntered in double scalars
 r = (lum array[j] / (4 * np.pi * sigmasb * (teff array[j]**4)))**0.5
In [ ]:
 1
```