

The Initial Data Problem of Numerical Relativity

Bowen-York Initial Data

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Overview

1 Introduction

2 Initial Data

3 The Model



Can't even trust black holes nowadays smh...

Introduction

Einstein Field Equations

1915 Relation between Geometry and Matter [Ein15]

$$\underbrace{G_{ab}}_{\text{Geometry}} = 8\pi \underbrace{T_{ab}}_{\text{Matter-Energy}}$$

1952 Can be posed as a Cauchy problem (with *constraints*!) [Fou52]

- Not trivial. In GR space and time are on equal footing!

1969 Extension from local to “global” (MGHD) [CG69]

- Foliation is possible in *globally hyperbolic* spacetimes [strongest causality condition; physically relevant spacetimes must satisfy it]

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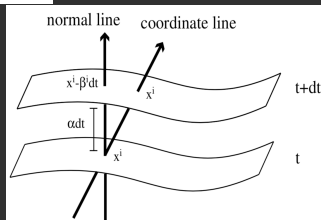
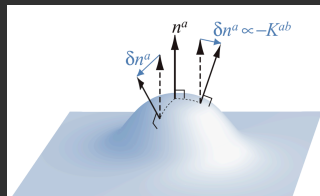
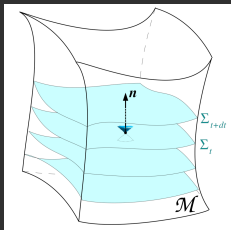
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3+1 Approach to Numerical Relativity



ADM Equations

✧ Evolution Equations:

$$\partial_t \gamma_{ij} = 2D_{(i}\beta_{j)} - 2\alpha K_{ij}$$

$$\begin{aligned} \partial_t K_{ij} = & \alpha(R_{ij} + KK_{ij} - 2K_{ik}K^k_{\ j}) - 8\pi\alpha \left(S_{ij} - \frac{1}{2}\gamma_{ij}(S - \rho) \right) \\ & - D_i D_j \alpha + \beta^k D_k K_{ij} + 2K_{k(j} D_{i)} \beta^k \end{aligned}$$

✧ Constraint Equations:

$$\begin{aligned} R + K^2 - K_{ij}K^{ij} &= 16\pi\rho \\ D_j \left(K^{ij} - \gamma^{ij}K \right) &= 8\pi S^i \end{aligned}$$

Problems with ADM



- Straightforward and relatively easy...but practically useless
 - Numerical simulations violently unstable
 - Equations in this form are “weakly hyperbolic”
- “More hyperbolic” form needed
 - use conformal rescaling
 - messier, but effective

Initial Data

Initial Data Problem

Solving the constraints

$$\begin{aligned} R + K^2 - K_{ij}K^{ij} &= 0 \\ D_j \left(K^{ij} - \gamma^{ij}K \right) &= 0 \end{aligned}$$

is not so trivial.

- there's a total 12 DoF on the system $(\{\gamma_{ij}, K_{ij}\})$;
- the constraints are just 4 equations (thus removing 4 DoF);
- no a priori preference for which eight of the total data to use as *free* parameters and which four as *constrained* quantities.

York-Lichnerowicz Conformal Transformations

Re-scale the 3-metric γ_{ij} and the (traceless part of) the extrinsic curvature A_{ij} :

$$\begin{aligned}\bar{\gamma}_{ij} &= \psi^{-4} \gamma_{ij} \\ \tilde{A}_{ij} &= \psi^2 A_{ij}\end{aligned}$$

where

$$A_{ij} = K_{ij} - \frac{1}{3} \gamma_{ij} K.$$

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These conformal transformations lead to the **constraints**

$$\begin{aligned}\bar{D}^2 \psi + \frac{1}{8} \left(\psi^{-7} \tilde{A}_{ij} \tilde{A}^{ij} - \psi \bar{R} \right) - \frac{1}{12} \psi^5 K^2 &= 0 \\ \bar{D}_j \tilde{A}^{ij} - \frac{2}{3} \psi^6 \bar{D}^i K &= 0\end{aligned}$$

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Conformal Transverse-Traceless (CTT) method

$$\tilde{A}^{ij} = \tilde{A}_L^{ij} + \tilde{A}_{\text{TT}}^{ij}$$

with the **transverse-traceless** (TT) part of \tilde{A}^{ij} satisfying

$$\bar{D}_j \tilde{A}_{\text{TT}}^{ij} = 0 \quad (\text{Transverse})$$

$$\tilde{\gamma}_{ij} \tilde{A}_{\text{TT}}^{ij} = 0. \quad (\text{Traceless})$$

The **longitudinal** (L) part of \tilde{A}^{ij} is expressed in terms of the **conformal Killing operator** $(\bar{\mathbb{L}}X)^{ij}$:

$$\tilde{A}_L^{ij} \equiv (\bar{\mathbb{L}}X)^{ij} := 2\bar{D}^{(i}X^{j)} - \frac{2}{3}\tilde{\gamma}^{ij}\bar{D}_kX^k.$$

Relation between X^i and \tilde{A}^{ij} is given by the **conformal vector Laplacian**

$$\bar{\Delta}_{\bar{\mathbb{L}}}X^i \equiv \bar{D}_j(\bar{\mathbb{L}}X)^{ij} = \bar{D}_j\tilde{A}^{ij}.$$

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Conformal Transverse-Traceless (CTT) method

✧ CTT approach (12 DoF)

- Constrained Data (4 DoF): ψ, X^i
- Free Data (8 DoF): $\tilde{A}_{\text{TT}}^{ij}, \tilde{\gamma}_{ij}, K$

The physical solution is then constructed from

$$\gamma_{ij} = \psi^4 \tilde{\gamma}_{ij}$$

$$K_{ij} = \psi^{-2} \tilde{A}_{ij} + \frac{1}{3} \gamma_{ij} K$$

The Model

Bowen-York Model

$$\tilde{A}_{\text{TT}}^{ij} = 0$$

$$K = 0$$

$$\tilde{\gamma}_{ij} = f_{ij} \quad (= \delta_{ij}).$$

(Maximal Slicing)

(Conformal Flatness)

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With these assumptions, the momentum constraints become

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Two crucial solutions

$$X_J^i = \frac{1}{r^2} \left(\bar{\epsilon}^{ijk} \ell_j J_k \right)$$

$$X_P^i = -\frac{1}{4r} \left(7P^i + \ell^i \ell_j P^j \right).$$

Bowen-York Model

With $X_{\vec{f}}^i, X_{\vec{p}}^i$ as above, plugging into

$$\tilde{A}^{ij} = \tilde{A}_{\vec{L}}^{ij} = (\mathbb{L}X)^{ij} = 2\bar{D}^{(i}X^{j)} - \frac{2}{3}\bar{\gamma}^{ij}\bar{D}_kX^k,$$

we end up with

$$\begin{aligned}\tilde{A}_{\vec{f}}^{ij} &= \frac{6}{r^3}\ell^{(i}\bar{\epsilon}^{j)kl}J_k\ell_l \\ \tilde{A}_{\vec{p}}^{ij} &= \frac{3}{2r^2}\left[2P^{(i}\ell^{j)} + \ell_kP^k\left(\ell^i\ell^j - \delta^{ij}\right)\right]\end{aligned}$$

Puncture Data

Using

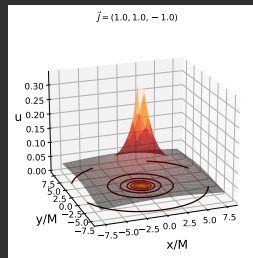
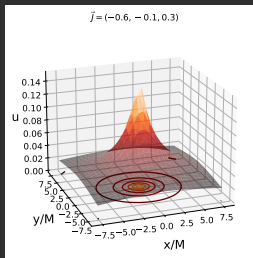
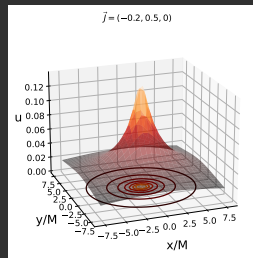
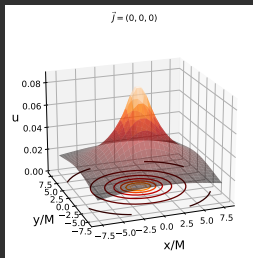
$$\psi = 1 + \frac{1}{\vartheta} + u, \quad \text{where} \quad \frac{1}{\vartheta} \equiv \frac{M}{2r}$$

$$\varrho \equiv \frac{\vartheta^7}{8} \tilde{A}_{ij} \tilde{A}^{ij}$$

we end up with the final form of the Hamiltonian constraint:

$$\partial^2 u = -\varrho (\vartheta(1 + u) + 1)^{-7}$$

Results



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THANK YOU!