

Numerical Relativity

The Initial Data Problem

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Overview

- 1 Introduction
- 2 ADM Formalism
- 3 BSSN Formalism
- 4 Initial Data



Can't even trust black holes nowadays smh...

Introduction

Equations...Equations ...Equations anyone?



$$\begin{aligned}
 \partial_t \bar{A}_{ij} &= \chi [\alpha (R_{ij} + K K_{ij} - 2 K_{ik} K^k_j) \\
 &\quad - 8\pi\alpha (S_{ij} - \underbrace{\frac{1}{2} \gamma_{ij}}_{\text{no TF}} (S - \rho)) - D_i D_j \alpha \\
 &\quad + \mathcal{L}_{\vec{\beta}} K_{ij}]^{\text{TF}} + \chi^{-1} \bar{A}_{ij} [\frac{2}{3} \chi (\alpha K - \partial_i \beta^i) \\
 &\quad + \beta^i \partial_i \chi] \\
 &= [\alpha \chi R_{ij} + \alpha \chi \chi^{-1} K \left(\bar{A}_{ij} + \frac{1}{3} \bar{\gamma}_{ij} K \right) \\
 &\quad - 2\alpha \chi \chi^{-1} \left(\bar{A}_{ik} + \frac{1}{3} \bar{\gamma}_{ik} K \right) \left(\bar{A}^k_j + \delta^k_j K \right) \dots \\
 &\quad \dots - 8\pi\alpha \chi S_{ij} - \chi D_i D_j \alpha]^{\text{TF}} + \dots
 \end{aligned}$$

Einstein Field Equations

1915 Relation between Geometry and Matter [Ein15]

$$\underbrace{G_{ab}}_{\text{Geometry}} = 8\pi \underbrace{T_{ab}}_{\text{Matter-Energy}}$$

1952 Can be posed as a Cauchy problem (with *constraints*!) [Fou52]

- Not trivial. In GR space and time are on equal footing!

1969 Extension from local to “global” (MGHD) [CG69]

- Foliation is possible in *globally hyperbolic* spacetimes [strongest causality condition; physically relevant spacetimes must satisfy it]

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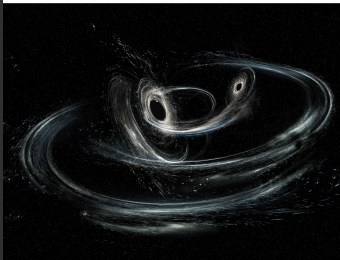
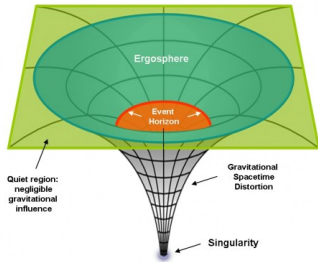
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Why numerical solutions? (Symmetry/Asymmetry)

Black Hole Regions

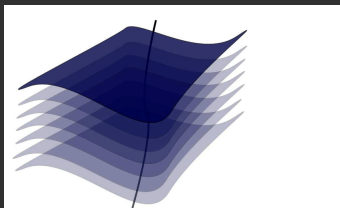


- Symmetries \implies exact solutions to EFE's;
 - Schwarzschild, Kerr, Reissner-Nordström, Kerr-Newman, ...
- No symmetries \implies solutions to EFE's unlikely;
 - Indispensable for study of black hole mergers (GW's!)

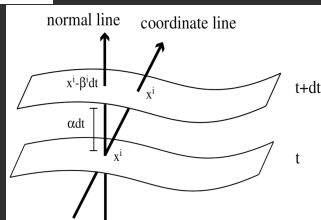
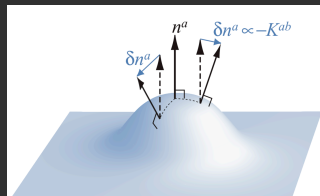
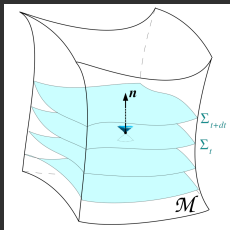
ADM Formalism

3+1 Approach to Numerical Relativity

- Formulate EFE's as a Cauchy problem with constraints.
 - Not trivial! In GR space and time are on equal footing 😊
- Foliation is possible in *globally hyperbolic* spacetimes.
 - Strongest causality condition (physically relevant spacetimes must satisfy it)



Some Key Ingredients



Some Key Ingredients

■ Intrinsic metric & curvature

$$\begin{aligned}\gamma_{ab} &= g_{ab} + n_a n_b \\ K_{ab} &= -\gamma_a^c \gamma_b^d \nabla_c n_d \\ &= -\frac{1}{2} \mathcal{L}_{\vec{n}} \gamma_{ab}\end{aligned}$$

■ Projections

$$\begin{aligned}\gamma_a^e \gamma_b^f \gamma_c^g \gamma_d^h {}^{(4)}R_{efgh} &= R_{abcd} + K_{ac} K_{bd} - K_{ad} K_{cb} \\ \gamma_a^e \gamma_b^f \gamma_c^g n^h {}^{(4)}R_{efgh} &= D_b K_{ac} - D_a K_{bc} \\ \gamma_a^q \gamma_b^r n^c n^d {}^{(4)}R_{qcrd} &= \mathcal{L}_{\vec{n}} K_{ab} + \frac{1}{\alpha} D_a D_b \alpha + K_b^c K_{ac}\end{aligned}$$

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ADM Equations

✧ Evolution Equations:

$$\partial_t \gamma_{ij} = 2D_{(i}\beta_{j)} - 2\alpha K_{ij}$$

$$\begin{aligned} \partial_t K_{ij} = & \alpha(R_{ij} + KK_{ij} - 2K_{ik}K^k_{\ j}) - 8\pi\alpha \left(S_{ij} - \frac{1}{2}\gamma_{ij}(S - \rho) \right) \\ & - D_i D_j \alpha + \beta^k D_k K_{ij} + 2K_{k(j} D_{i)} \beta^k \end{aligned}$$

✧ Constraint Equations:

$$\begin{aligned} R + K^2 - K_{ij}K^{ij} &= 16\pi\rho \\ D_j \left(K^{ij} - \gamma^{ij}K \right) &= 8\pi S^i \end{aligned}$$

Problems with ADM



- Straightforward and relatively easy...but practically useless
 - Numerical simulations violently unstable
 - Equations in this form are “weakly hyperbolic”
- “More hyperbolic” form needed
 - use conformal rescaling
 - messier, but effective

BSSN Formalism

Advantages of BSSN



- Messier, but effective! 😊
 - Stable numerical simulations
 - Equations in this form are “more hyperbolic”
- (One of) *the* standard(s) in NR

Setting up the stage

Theorem

Two **strongly causal** Lorentzian metrics $g_{ab}^{(1)}$ and $g_{ab}^{(2)}$ for some manifold \mathcal{M} determine the same future and past sets at all points (events) if and only if the two metrics are **globally conformal**, i.e., if $g_{ab}^{(1)} = \Psi g_{ab}^{(2)}$, for some smooth function $\Psi \in C^\infty(\mathcal{M})$. In this case, both spacetimes $(\mathcal{M}, g_{ab}^{(1)})$ and $(\mathcal{M}, g_{ab}^{(2)})$ belong to the same conformal class and share the same causal structure.

BSSN Approach

✧ BSSN Variables:

$$\{\chi, \bar{\gamma}_{ij}, \bar{A}_{ij}, K, \bar{\Gamma}^i\}$$

✧ Conformal Rescalings:

$$\begin{aligned}\bar{\gamma}_{ij} &= \chi \gamma_{ij} \\ \bar{\gamma} &= 1 \\ \chi &= \gamma^{-1/3} \\ \bar{A}_{ij} &= \chi K_{ij} - \frac{1}{3} \bar{\gamma}_{ij} K \\ \bar{\Gamma}^i &= \bar{\gamma}^{jk} \bar{\Gamma}_{jk}^i \\ &= -\partial_j \bar{\gamma}^{ij}\end{aligned}$$

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$$\bar{\gamma} = 1$$

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$$\bar{A}_{ij} = \chi K_{ij} - \frac{1}{3} \bar{\gamma}_{ij} K$$

$$\begin{aligned} \bar{\Gamma}^i &= \bar{\gamma}^{jk} \bar{\Gamma}_{jk}^i \\ &= -\partial_j \bar{\gamma}^{ij} \end{aligned}$$

BSSN Evolution Eqs

✧ Evolution Equations:

$$\partial_t \chi = \frac{2}{3} \chi (\alpha K - \partial_i \beta^i) + \beta^i \bar{D}_i \chi$$

$$\partial_t \bar{\gamma}_{ij} = -2\alpha \bar{A}_{ij} + \beta^k \partial_k \bar{\gamma}_{ij} + \bar{\gamma}_{ik} \partial_j \beta^k + \bar{\gamma}_{kj} \partial_i \beta^k - \frac{2}{3} \bar{\gamma}_{ij} \partial_k \beta^k$$

$$\partial_t K = \alpha \left(\bar{A}_{ij} \bar{A}^{ij} + \frac{1}{3} K^2 \right) + 4\pi\alpha(\rho + S) - D^2 \alpha + \beta^i \bar{D}_i K$$

$$\begin{aligned} \partial_t \bar{A}_{ij} = & [\chi (\alpha R_{ij} - 8\pi\alpha S_{ij} - D_i D_j \alpha)]^{\text{TF}} - \alpha (2\bar{A}_{ik} \bar{A}^k_j + \bar{A}_{ij} K) \\ & + \beta^k \partial_k \bar{A}_{ij} + \bar{A}_{ik} \partial_j \beta^k + \bar{A}_{kj} \partial_i \beta^k - \frac{2}{3} \bar{A}_{ij} \partial_k \beta^k \end{aligned}$$

$$\begin{aligned} \partial_t \bar{\Gamma}^i = & -2\alpha \left(\frac{3}{2\chi} \bar{A}^{ij} \bar{D}_j \chi + \frac{2}{3} \bar{D}^i K + 8\pi \bar{S}^i - \bar{\Gamma}^i_{jk} \bar{A}^{jk} \right) - 2\bar{A}^{ij} \bar{D}_j \alpha \\ & + \beta^j \partial_j \bar{\Gamma}^i + \bar{\gamma}^{jk} \partial_j \partial_k \beta^i - \bar{\Gamma}^j \partial_j \beta^i + \frac{2}{3} \bar{\Gamma}^i \partial_j \beta^j + \frac{1}{3} \bar{\gamma}^{ij} \partial_j \partial_k \beta^k \end{aligned}$$

BSSN Approach (cont)

✧ Constraint Equations:

$$\begin{aligned}
 16\pi\bar{\rho} &= \bar{R} + 2\bar{D}^2(\log\chi) - \frac{1}{2}\bar{D}_k(\log\chi)\bar{D}^k(\log\chi) \\
 &\quad + \frac{4}{3\chi}K^2 - \frac{1}{\chi}\bar{A}_{ij}\bar{A}^{ij} \\
 8\pi\bar{S}^i &= \bar{D}_j\bar{A}^{ij} - \frac{3}{2\chi}\bar{A}^{ij}\bar{D}_j\chi - \frac{2}{3}\bar{D}^iK
 \end{aligned}$$

Initial Data

Initial Data Problem

Solving the constraints

$$\begin{aligned} R + K^2 - K_{ij}K^{ij} &= 16\pi\rho \\ D_j \left(K^{ij} - \gamma^{ij}K \right) &= 8\pi S^i \end{aligned}$$

is not so trivial.

- there's a total 12 DoF on the system $(\{\gamma_{ij}, K_{ij}\})$;
- the constraints are just 4 equations (thus removing 4 DoF);
- no a priori preference for which eight of the total data to use as *free* parameters and which four as *constrained* quantities.

York-Lichnerowicz Conformal Transformations

Re-scale the 3-metric γ_{ij} and the (traceless part of) the extrinsic curvature A_{ij} :

$$\begin{aligned}\bar{\gamma}_{ij} &= \psi^{-4} \gamma_{ij} \\ \tilde{A}_{ij} &= \psi^2 A_{ij}\end{aligned}$$

where

$$\begin{aligned}\psi &= \chi^{-1/4} = \gamma^{1/12} \\ A_{ij} &= K_{ij} - \frac{1}{3} \gamma_{ij} K.\end{aligned}$$

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These conformal transformations lead to the **constraints**

$$\begin{aligned}\bar{D}^2 \psi + \frac{1}{8} \left(\psi^{-7} \tilde{A}_{ij} \tilde{A}^{ij} - \psi \bar{R} \right) - \frac{1}{12} \psi^5 K^2 &= -2\pi \psi^{-3} \tilde{\rho} \\ \bar{D}_j \tilde{A}^{ij} - \frac{2}{3} \psi^6 \bar{D}^i K &= 8\pi \tilde{S}^i\end{aligned}$$

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Conformal Transverse-Traceless (CTT) method

$$\tilde{A}^{ij} = \tilde{A}_L^{ij} + \tilde{A}_{\text{TT}}^{ij}$$

with the **transverse-traceless** (TT) part of \tilde{A}^{ij} satisfying

$$\bar{D}_j \tilde{A}_{\text{TT}}^{ij} = 0 \quad (\text{Transverse})$$

$$\tilde{\gamma}_{ij} \tilde{A}_{\text{TT}}^{ij} = 0. \quad (\text{Traceless})$$

The **longitudinal** (L) part of \tilde{A}^{ij} is expressed in terms of the **conformal Killing operator** $(\mathbb{L}X)^{ij}$:

$$\tilde{A}_L^{ij} \equiv (\mathbb{L}X)^{ij} := 2\bar{D}^{(i}X^{j)} - \frac{2}{3}\tilde{\gamma}^{ij}\bar{D}_kX^k.$$

The vector field X is determined from

$$\bar{D}_j(\mathbb{L}X)^{ij} = \bar{D}_j\tilde{A}^{ij}.$$

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The vector field X is determined from the **conformal vector Laplacian**

$$\bar{\Delta}_{\mathbb{L}}X^i \equiv \bar{D}_j(\mathbb{L}X)^{ij} = \bar{D}_j\tilde{A}^{ij}.$$

Conformal Transverse-Traceless (CTT) method

Expanding $\bar{\Delta}_{\mathbb{L}} X^i$:

$$\begin{aligned}
 \bar{\Delta}_{\mathbb{L}} X^i &= \bar{D}_j (\bar{\mathbb{L}} X)^{ij} = \bar{D}_j \left(\bar{D}^i X^j + \bar{D}^j X^i - \frac{2}{3} \bar{\gamma}^{ij} \bar{D}_k X^k \right) \\
 &= \underbrace{\bar{D}_j \bar{D}^i X^j}_{=\bar{R}^i_j X^j + \bar{D}^i \bar{D}_j X^j} + \bar{D}^2 X^i - \frac{2}{3} \bar{D}^i \bar{D}_j X^j \\
 &= \bar{D}^2 X^i + \bar{R}^i_j X^j + \frac{1}{3} \bar{D}^i \bar{D}_j X^j. \qquad \text{(by Ricci identity)}
 \end{aligned}$$

Replacing $\bar{\Delta}_{\mathbb{L}} X^i \leftrightarrow \bar{D}_j \tilde{A}^{ij}$ in the constraint

$$\bar{D}_j \tilde{A}^{ij} - \frac{2}{3} \psi^6 \bar{D}^i K = 8\pi \tilde{S}^i,$$

we rewrite the **momentum constraints** as

$$\bar{D}^2 X^i + \bar{R}^i_j X^j + \frac{1}{3} \bar{D}^i \bar{D}_j X^j - \frac{2}{3} \psi^6 \bar{D}^i K = 8\pi \tilde{S}^i$$

Conformal Transverse-Traceless (CTT) method

Expanding $\bar{\Delta}_{\mathbb{L}} X^i$:

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Conformal Transverse-Traceless (CTT) method

To sum up, in CTT,

$$\begin{aligned}\bar{D}^2\psi + \frac{1}{8} \left(\psi^{-7} \tilde{A}_{ij} \tilde{A}^{ij} - \psi \bar{R} \right) - \frac{1}{12} \psi^5 K^2 &= -2\pi \psi^{-3} \tilde{\rho} \\ \bar{D}_j \tilde{A}^{ij} - \frac{2}{3} \psi^6 \bar{D}^i K &= 8\pi \tilde{S}^i\end{aligned}$$

becomes

$$\begin{aligned}\bar{D}^2\psi + \frac{1}{8} \left(\psi^{-7} \tilde{A}_{ij} \tilde{A}^{ij} - \psi \bar{R} \right) - \frac{1}{12} \psi^5 K^2 &= -2\pi \psi^{-3} \tilde{\rho} \\ \bar{D}^2 X^i + \bar{R}^i_j X^j + \frac{1}{3} \bar{D}^i \bar{D}_j X^j - \frac{2}{3} \psi^6 \bar{D}^i K &= 8\pi \tilde{S}^i\end{aligned}$$

Conformal Transverse-Traceless (CTT) method

✧ CTT approach (12 DoF)

- Constrained Data (4 DoF): ψ, X^i
- Free Data (8 DoF): $\tilde{A}_{\text{TT}}^{ij}, \tilde{\gamma}_{ij}, K$, (matter terms (ρ, S^i) , if present)

The physical solution is then constructed from

$$\gamma_{ij} = \psi^4 \tilde{\gamma}_{ij}$$

$$K_{ij} = \psi^{-2} \tilde{A}_{ij} + \frac{1}{3} \gamma_{ij} K.$$

Conformal Thin-Sandwich (CTS) method

- Consider $\bar{\gamma}_{ij}$ on Σ and its vicinity; i.e., $\bar{\gamma}_{ij}$ and $\bar{v}_{ij} \equiv \partial_t \bar{\gamma}_{ij}$
- We'll use (from BSSN)

$$\bar{v}_{ij} \equiv \partial_t \bar{\gamma}_{ij} = 2\bar{D}_{(i}\beta_{j)} - \frac{2}{3}\bar{\gamma}_{ij}\bar{D}_k\beta^k - 2\alpha\bar{A}_{ij},$$

with

$$\bar{A}_{ij} = \psi^{-4}A_{ij} = \psi^{-6}\tilde{A}_{ij}.$$

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with

$$\bar{A}_{ij} = \psi^{-4}A_{ij} = \psi^{-6}\tilde{A}_{ij}.$$

- Thus,

$$\tilde{A}_{ij} = \frac{1}{2\bar{\alpha}} \left((\bar{\mathbb{L}}\beta)_{ij} - \bar{v}_{ij} \right)$$

with the *densitized lapse*

$$\bar{\alpha} \equiv \psi^{-6}\alpha.$$

Conformal Thin-Sandwich (CTS) method

Expand the divergence $\bar{D}_j \tilde{A}^{ij}$,

$$\begin{aligned}\bar{D}_j \tilde{A}^{ij} &= \bar{D}_j \left[\frac{1}{2\bar{\alpha}} \left((\bar{\mathbb{L}}\beta)^{ij} - \bar{v}^{ij} \right) \right] \\ &= \frac{1}{2\bar{\alpha}} \bar{\Delta}_{\bar{\mathbb{L}}} \beta^i - \frac{1}{2\bar{\alpha}^2} (\bar{\mathbb{L}}\beta)^{ij} \bar{D}_j \bar{\alpha} - \frac{1}{2\bar{\alpha}} \bar{D}_j \bar{v}^{ij} + \frac{1}{2\bar{\alpha}^2} \bar{v}^{ij} \bar{D}_j \bar{\alpha} \\ &= \frac{1}{2\bar{\alpha}} \left(\bar{\Delta}_{\bar{\mathbb{L}}} \beta^i - (\bar{\mathbb{L}}\beta)^{ij} \bar{D}_j (\log \bar{\alpha}) - \bar{D}_j \bar{v}^{ij} + \bar{v}^{ij} \bar{D}_j (\log \bar{\alpha}) \right).\end{aligned}$$

Then plugging in the momentum constraints,

$$\begin{aligned}\bar{D}_j \tilde{A}^{ij} - \frac{2}{3} \psi^6 \bar{D}^i K &= 8\pi \tilde{S}^i \\ \frac{1}{2\bar{\alpha}} [\bar{\Delta}_{\bar{\mathbb{L}}} \beta^i - (\bar{\mathbb{L}}\beta)^{ij} \bar{D}_j (\log \bar{\alpha}) - \underbrace{\bar{D}_j \bar{v}^{ij} + \bar{v}^{ij} \bar{D}_j (\log \bar{\alpha})}_{= -\bar{\alpha} \bar{D}_j (\bar{\alpha}^{-1} \bar{v}^{ij})}] - \frac{2}{3} \psi^6 \bar{D}^i K &= 8\pi \tilde{S}^i\end{aligned}$$

$$\bar{\Delta}_{\bar{\mathbb{L}}} \beta^i - (\bar{\mathbb{L}}\beta)^{ij} \bar{D}_j (\log \bar{\alpha}) - \bar{\alpha} \bar{D}_j (\bar{\alpha}^{-1} \bar{v}^{ij}) - \frac{4}{3} \bar{\alpha} \psi^6 \bar{D}^i K = 16\pi \bar{\alpha} \tilde{S}^i.$$

Conformal Thin-Sandwich (CTS) method

Expand the divergence $\bar{D}_j \tilde{A}^{ij}$,

$$\begin{aligned}
 \bar{D}_j \tilde{A}^{ij} &= \bar{D}_j \left[\frac{1}{2\bar{\alpha}} \left((\bar{\mathbb{L}}\beta)^{ij} - \bar{v}^{ij} \right) \right] \\
 &= \frac{1}{2\bar{\alpha}} \bar{\Delta} \bar{\mathbb{L}}\beta^i - \frac{1}{2\bar{\alpha}^2} (\bar{\mathbb{L}}\beta)^{ij} \bar{D}_j \bar{\alpha} - \frac{1}{2\bar{\alpha}} \bar{D}_j \bar{v}^{ij} + \frac{1}{2\bar{\alpha}^2} \bar{v}^{ij} \bar{D}_j \bar{\alpha} \\
 &= \frac{1}{2\bar{\alpha}} \left(\bar{\Delta} \bar{\mathbb{L}}\beta^i - (\bar{\mathbb{L}}\beta)^{ij} \bar{D}_j (\log \bar{\alpha}) - \bar{D}_j \bar{v}^{ij} + \bar{v}^{ij} \bar{D}_j (\log \bar{\alpha}) \right).
 \end{aligned}$$

Then plugging in the momentum constraints,

$$\begin{aligned}
 \bar{D}_j \tilde{A}^{ij} - \frac{2}{3} \psi^6 \bar{D}^i K &= 8\pi \tilde{S}^i \\
 \frac{1}{2\bar{\alpha}} [\bar{\Delta} \bar{\mathbb{L}}\beta^i - (\bar{\mathbb{L}}\beta)^{ij} \bar{D}_j (\log \bar{\alpha}) - \underbrace{\bar{D}_j \bar{v}^{ij} + \bar{v}^{ij} \bar{D}_j (\log \bar{\alpha})}_{= -\bar{\alpha} \bar{D}_j (\bar{\alpha}^{-1} \bar{v}^{ij})}] - \frac{2}{3} \psi^6 \bar{D}^i K &= 8\pi \tilde{S}^i
 \end{aligned}$$

$$\bar{\Delta} \bar{\mathbb{L}}\beta^i - (\bar{\mathbb{L}}\beta)^{ij} \bar{D}_j (\log \bar{\alpha}) - \bar{\alpha} \bar{D}_j (\bar{\alpha}^{-1} \bar{v}^{ij}) - \frac{4}{3} \bar{\alpha} \psi^6 \bar{D}^i K = 16\pi \bar{\alpha} \tilde{S}^i.$$

Conformal Thin-Sandwich (CTS) method

Hence, in CTS,

$$\begin{aligned}\bar{D}^2\psi + \frac{1}{8} \left(\psi^{-7} \tilde{A}_{ij} \tilde{A}^{ij} - \psi \bar{R} \right) - \frac{1}{12} \psi^5 K^2 &= -2\pi \psi^{-3} \tilde{\rho} \\ \bar{D}_j \tilde{A}^{ij} - \frac{2}{3} \psi^6 \bar{D}^i K &= 8\pi \tilde{S}^i\end{aligned}$$

becomes

$$\begin{aligned}\bar{D}^2\psi + \frac{1}{8} \left(\psi^{-7} \tilde{A}_{ij} \tilde{A}^{ij} - \psi \bar{R} \right) - \frac{1}{12} \psi^5 K^2 &= -2\pi \psi^{-3} \tilde{\rho} \\ \bar{\Delta}_{\bar{\mathbb{L}}} \beta^i - (\bar{\mathbb{L}} \beta)^{ij} \bar{D}_j (\log \bar{\alpha}) - \bar{\alpha} \bar{D}_j (\bar{\alpha}^{-1} \bar{v}^{ij}) - \frac{4}{3} \bar{\alpha} \psi^6 \bar{D}^i K &= 16\pi \bar{\alpha} \tilde{S}^i\end{aligned}$$

Conformal Thin-Sandwich (CTS) method

✧ CTS approach (16 DoF)

- Constrained Data (4 DoF) : ψ, β^i
- Free Data (12 DoF) : $\bar{\gamma}_{ij}, \bar{v}_{ij}, K, \bar{\alpha}$, (matter terms (ρ, S^i) , if present)

The physical solution is then constructed from

$$\alpha = \psi^6 \bar{\alpha}$$

$$\gamma_{ij} = \psi^4 \bar{\gamma}_{ij}$$

$$K_{ij} = \psi^{-2} \tilde{A}_{ij} + \frac{1}{3} \gamma_{ij} K.$$

Extended Conformal Thin-Sandwich (XCTS) method

✱ Replaces CTS's free datum $\bar{\alpha}$ with $\partial_t K$

- **Constrained Data:** ψ, β^i, α (or, equivalently, $\bar{\alpha}$);
- **Free Data:** $\bar{\gamma}_{ij}, \bar{v}_{ij}, K, \partial_t K$, (matter terms (ρ, S^i) , if present).

✱ Additional constraint equation

$$\bar{D}^2(\alpha\psi) = \alpha\psi \left(\frac{7}{8}\psi^{-8}\tilde{A}_{ij}\tilde{A}^{ij} + \frac{5}{12}\psi^4 K^2 + \frac{1}{8}\bar{R} + 2\pi\psi^{-4}(2\tilde{S} + \tilde{\rho}) \right) + \psi^5 \left(\beta^i \bar{D}_i K - \partial_t K \right)$$

Extended Conformal Thin-Sandwich (XCTS) method

- ✧ Replaces CTS's free datum $\bar{\alpha}$ with $\partial_t K$
 - Constrained Data: ψ, β^i, α (or, equivalently, $\bar{\alpha}$);
 - Free Data: $\bar{\gamma}_{ij}, \bar{v}_{ij}, K, \partial_t K$, (matter terms (ρ, S^i) , if present).
- ✧ Additional constraint equation

$$\bar{D}^2(\alpha\psi) = \alpha\psi \left(\frac{7}{8}\psi^{-8}\tilde{A}_{ij}\tilde{A}^{ij} + \frac{5}{12}\psi^4 K^2 + \frac{1}{8}\bar{R} + 2\pi\psi^{-4}(2\tilde{S} + \tilde{\rho}) \right) + \psi^5 \left(\beta^i \bar{D}_i K - \partial_t K \right)$$

The slippery slope of having $\bar{\gamma}_{ij}$ as free data

- Conformally-related metric $\bar{\gamma}_{ij}$ as free data
 - flat? ($\bar{\gamma}_{ij} = f_{ij}$) \longrightarrow spurious GW bursts at start of sim
 - Kerr-Schild-like? ($\bar{\gamma}_{ij} = f_{ij} + 2H\ell_i\ell_j$)
 - more control over $\bar{\gamma}_{ij}$ in CTS than in CTT, as $\partial_t \bar{\gamma}_{ij}$ is also free
 - even more control in XCTS ($\partial_t \bar{\gamma}_{ij}$ and $\partial_t K$ are free)
- Waveless Approximation ([SUFO4])

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 - setting $\partial_t\bar{\gamma}_{ij} = \partial_t K = 0$ is great for equilibrium/quasi-equilibrium settings.
- Waveless Approximation ([SUF04])
 - replaces $\bar{\gamma}_{ij} \leftrightarrow \partial_t\tilde{A}_{ij}$ as free data
 - impose helical symmetry of $\partial_t\bar{\gamma}_{ij}$ & $\partial_t\tilde{A}_{ij}$ in the near zone, and set them = 0 in the far zone. (Very accurate for binaries!)

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THANK YOU!