

Problem 1 (Exercise 7.8 (Schutz)). Suppose that in some coordinate system the components of the metric $g_{\alpha\beta}$ are independent of some coordinate x^μ .

(a) Show that the conservation law $T^\nu_{\mu;\nu} = 0$ for any stress-energy tensor becomes

$$\frac{1}{\sqrt{-g}} \left(\sqrt{-g} T^\nu_\mu \right)_{,\nu} = 0. \quad (1)$$

(b) Suppose that in these coordinates $T^{\alpha\beta} \neq 0$ only in some bounded region of each spacelike hypersurface $x^0 = \text{const}$. Show that Eq. (1) implies

$$\int_{x^0=\text{const}} \sqrt{-g} T^\nu_\mu n_\nu d^3x \quad (2)$$

is independent of x^0 , if n_ν is the unit normal to the hypersurface.

(c) Consider flat Minkowski space in a global inertial frame with spherical polar coordinates (t, r, θ, ϕ) . Show from (b) that

$$J = \int_{t=\text{const}} T^0_\phi r^2 \sin \theta dr d\theta d\phi \quad (3)$$

is independent of t . This is the total angular momentum of the system.

(d) Express the integral in (c) in terms of the components of $T^{\alpha\beta}$ on the Cartesian basis (t, x, y, z) , showing that

$$J = \int (x T^{y0} - y T^{x0}) dx dy dz. \quad (4)$$

This is the continuum version of the nonrelativistic expression $(\mathbf{r} \times \mathbf{p})_z$ for a particle's angular momentum about the z axis.

Solution to (a). Let's first expand Eq. (1):

$$\begin{aligned} \frac{1}{\sqrt{-g}} \partial_\nu \left(\sqrt{-g} T^\nu_\mu \right) &= \partial_\nu T^\nu_\mu + \frac{1}{\sqrt{-g}} T^\nu_\mu \partial_\nu \sqrt{-g} \\ &= \partial_\nu T^\nu_\mu + \frac{1}{2g} T^\nu_\mu \partial_\nu g. \end{aligned} \quad (5)$$

Now recall that on a previous exercise (past homework) we showed that

$$\Gamma^\mu_{\mu\nu} = \frac{1}{2} \partial_\nu (\log |g|) = \frac{1}{2g} \partial_\nu g. \quad (6)$$

Moreover, note that

$$\begin{aligned} \Gamma^\sigma_{\mu\nu} T^\nu_\sigma &= \frac{1}{2} T^\nu_\sigma g^{\sigma\alpha} (\partial_\nu g_{\alpha\mu} + \partial_\mu g_{\alpha\nu} - \partial_\alpha g_{\mu\nu}) \\ &= \frac{1}{2} T^\nu_\sigma g^{\sigma\alpha} \partial_\mu g_{\alpha\nu} \\ &= 0. \end{aligned} \quad (7)$$

On the first equality the terms

$$\frac{1}{2} T^\nu_\sigma g^{\sigma\alpha} (\partial_\nu g_{\alpha\mu} - \partial_\alpha g_{\mu\nu})$$

vanished by relabeling. On the last equality we get zero because, by assumption on this exercise, we are using a coordinate system where the components of the metric tensor are independent of some coordinate x^μ ; therefore $\partial_\mu g_{\alpha\nu} = 0$.

Now, expanding $T^\nu_{\mu;\nu}$,

$$\begin{aligned} \nabla_\nu T^\nu_\mu &= \partial_\nu T^\nu_\mu + \Gamma^\nu_{\sigma\nu} T^\sigma_\mu - \overbrace{\Gamma^\sigma_{\mu\nu} T^\nu_\sigma}^{= 0 \text{ by Eq. (7)}} \\ &= \partial_\nu T^\nu_\mu + \frac{1}{2g} T^\sigma_\mu \partial_\sigma g. \end{aligned} \quad (\text{By Eq. (6)}) \quad (8)$$

Relabeling $\sigma \mapsto \nu$ we get the same result as (5). Thus we have proved the validity of (1). \square

Solution to (b). Here we use Gauss's Law. However, let me clarify something that is quite confusing from the notation of Schutz: Eq. (2) should actually read something like

$$\int_{x^0=\text{const}} \sqrt{\gamma} T^\nu_\mu n_\nu d^3x, \quad (9)$$

where γ is the determinant of the 3-metric γ_{ij} of the hypersurface $x^0 = \text{constant}$. Schutz uses g for the determinant of both the 4-metric and the 3-metric, which can lead to some confusion ...

Anyhow, proceeding from Eq. (1) and expanding,

$$\begin{aligned} 0 &= \frac{1}{\sqrt{-g}} \left(\sqrt{-g} T^\nu_\mu \right)_{,\nu} = T^\nu_{\mu;\nu} \\ \Rightarrow 0 &= \int_{\mathcal{M}} \sqrt{-g} T^\nu_{\mu;\nu} d^4x = \int_{\mathcal{M}} \left(\sqrt{-g} T^\nu_\mu \right)_{,\nu} d^4x \\ &= \int_{\partial\mathcal{M}} \sqrt{\gamma} T^\nu_\mu n_\nu d^3x. \end{aligned} \quad (\text{By Gauss's Law}) \quad (10)$$

Hence, in some coordinate system where the components of the metric $g_{\alpha\beta}$ are independent of some coordinate x^μ , the integral (10) vanishes regardless of whether $\partial\mathcal{M} = x^0$ constant or it's some other spacelike hypersurface. \square

Solution to (c). We have previously shown that such change of coordinates transforms the volume form as

$$d^3x = r^2 \sin \theta dr d\theta d\phi. \quad (11)$$

Moreover, on the hypersurface $x^0 = t = \text{constant}$, the unit normal has components $n_\nu = (1, 0, 0, 0)$; thus

$$T^\nu_\mu n_\nu = T^0_\mu n_0 = T^0_\mu.$$

Putting all this into Eq. (2), with $\mu = \phi$, we recover Eq. (3). \square

Solution to (d). Recall that the change of coordinates from Euclidean to spherical polar in \mathbb{R}^3 is given by

$$(x, y, z) \mapsto (r \cos \phi \sin \theta, r \sin \phi \sin \theta, r \cos \theta).$$

Now we transform T^0_ϕ to Cartesian:

$$\begin{aligned} T^0_\phi &= \Lambda^\alpha_\phi T^0_\alpha = \frac{\partial x}{\partial \phi} T^0_x + \frac{\partial y}{\partial \phi} T^0_y + \frac{\partial z}{\partial \phi} T^0_z \\ &= -r \sin \phi \sin \theta T^0_x + r \sin \phi \sin \theta T^0_y + 0 \cdot T^0_z \\ &= -y T^0_x + x T^0_y \\ &= -y T^{0x} + x T^{0y}, \end{aligned}$$

where on the last equality we used the fact that the Euclidean metric in Cartesian coordinates is just δ_{ij} . Combining this with Eq. (11) the result follows. \square