

NOvA Test Beam detector calibration

Technical Note

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Abstract

What is this about and what will I describe in here

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1 Introduction

Why is Test Beam? "The idea, as with any test beam experiment, is to expose a detector to a beam of very well-characterized particles, so that we can improve our understanding of how the detector responds to such particles. We make use of upstream detectors to collect data on the beam particles before they interact in the NOvA detector. For example, we will be able to see what a 1 GeV proton actually looks like in our detector, without having to simulate it, and we can test how well we would have reconstructed the energy using our existing techniques. We may find we are able to make improvements to our tools to better match what we see in the detector with how we reconstruct it. Or we may find we already do a pretty good job. Either way, with a full cross-comparison like this, we can be more confident in our analysis of the data and reduce the level of uncertainty we consider are associated with the relevant measured quantities. Ultimately, the aim will be to reduce the level of uncertainty on the neutrino oscillation analyses and to make even better, more accurate measurements of the Standard Model." Why is Test Beam calibration done:

- To be able to directly compare TB to the standard detectors
- To be able to verify our calibration procedures using TB data
- To study the particle response as a function of energy
- To determine an energy resolution
- To compare currently used energy scales to data and understand if we can use TB data for absolute energy scale in all NOvA detectors

Only considering improvement in the calibration systematic uncertainty, the NOvA Test Beam program can reduce the overall systematic uncertainty for the main NOvA measurements by about 10% [docdb:33012] (talk also contains a list of other talks on impact of TB in different NOvA areas).

Also use information from:

- NOvA Test Beam Technical Statement of Work
- NOvA Test Beam program (paper for DOE) [docdb:25074]
- NOvA Test Beam task force report [docdb:15750]
- Overview presentation of NOvA Test Beam [docdb:20495]
- Test Beam support document [docdb:22172]
- NOvA Test Beam program proceedings [docdb:55808]

Mike's proceedings from ICHEP 2020 [?].

2 Overview of Test Beam detector

What is Test Beam, how does the detector and beamline look like. What are the main differences between the NOvA standard detectors and the Test Beam detector? Placed in MC7b together with a beamline instrumentation (no need to describe beamline). Get a picture of Test Beam.

General parameters

Smaller version of the standard NOvA detectors made up of 2 vertical and 2 horizontal modules, compared to XxY, or XxY for ND and FD respectively. " This detector is identical in every way to the Near and Far Detectors but is scaled down, similar to how the Near Detector is a scaled down version of the Far Detector. It consists of 63 planes, and is constructed as a 2x2 module (c.f. ND is 3x3). The total mass is about 30 tons."63 planes, 32 vertical and 31 horizontal; 64 cells in a plane (2x32 extrusions), readout out in two 32 cell groups; FD: maxPlane=900, maxCell=390. ND: maxPlane=220, maxCell=100. TB: maxPlane=63, maxCell=64 Detector block taken from NDOS, which had only 31 planes [docdb:28943], beginning and ending with a vertical plane. During comissioning of Test Beam, glued in a horizontal plane in between the two block. [docdb:29543]

Describe the orientation of axes. What's right and left, what's positive and negative w, which planes are vertical and which horizontal (vertical are even starting at 0 and ending in 62 and horizontal are odd).

Readout for the horizontal planes on the opposite side compared to the standard NOvA detectors (why?). Causing underfilled cells in the middle and top of the detector, until... [docdb:49827] Describe orientation and readout of the TB detector. Mention the two surveys and the alignment troubles.

Scintillator

Scintillator oil (5600 gallons in total [from Detector Information wiki]) used from ND+NDOS and spare from Texas in the back of the detector (plus some from Ash River afaik). Block 0 filled with ND+NDOS scintillator oil kept in the tanker [docdb:37152]. Found out the rest of the scintillator in the tanks in Fermilab is contaminated and unusable [docdb:38349] and used NDOS oil from barells to top off horizontal modules. Shipped scintillator from Ash River and filled the first 21 planes of Block 1 [docdb:41961]. Last 10 planes filled with the NDOS-drained oil from Texas [docdb:34046 describing Texas oil and docdb:41961 showing final filling state]. Maybe show a plot of response at cell centre and mark the 3 different scintillators.

Readout

Same readout electronics as in the Far Detector, except for 8 Near Detector Front End Boards, 4 in planes ... (1st quarter of the detector) and another 4 in planes ... (3rd quarter of the detector). "The Near Detector (ND) and Far Detector (FD) use different versions of the front-end electronics (FEBs, front-end boards), designed to be able to collect data at different rates (the ND is a higher-rate environment than the FD since it is much closer to the beam source). The FD uses FEBv4s (i.e. version 4) and the ND uses FEBv5s. The Test Beam detector uses

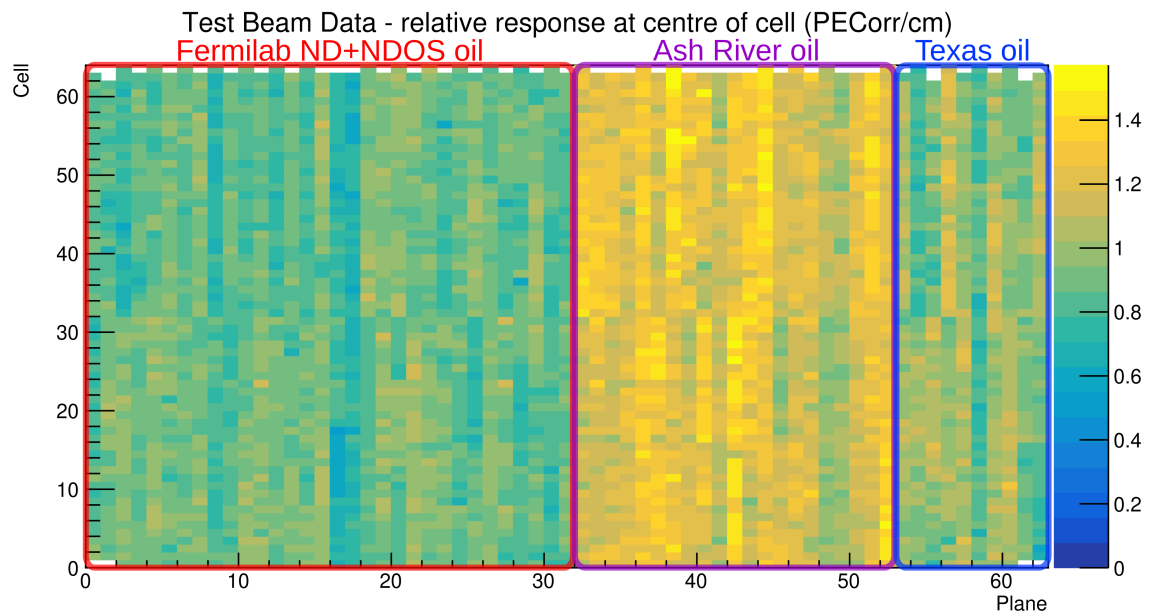


Figure 1: Uncorrected energy response in the centre of cells across the Test Beam detector showing a clear distinction between the different scintillator oils.

both in order to be able to make comparisons and validate both versions of the electronics. Most of the FEBs on the Test Beam detector are v4s; of the 126 total, 118 are v4 and 8 are v5."

Environment

Placed in the Fermilab Test Beam Facility with no overburden. Describe environmental controls, temperature dependence etc. Maybe add plots from environmental control (temperature differences etc.) with descriptions of where were the readings taken.

3 The NOvA calibration

Describe the NOvA calibration procedure in general. Describe the benefit of NOvA Test Beam and differences.

4 Calibration samples

How do we create the calibration samples and what cuts are applied?

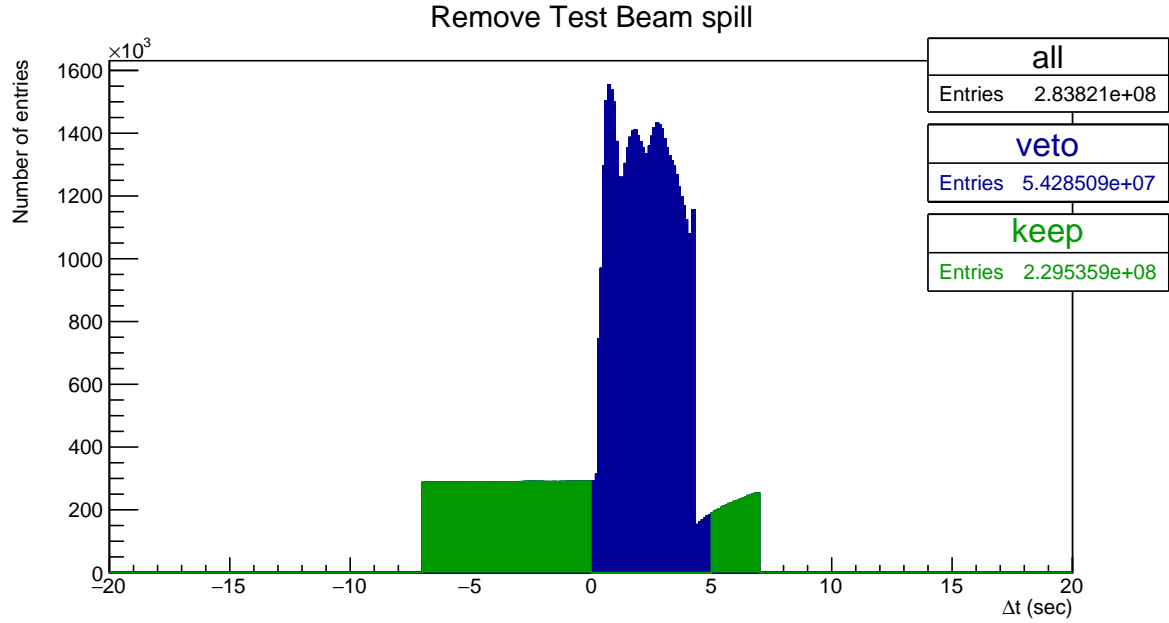


Figure 2: Test Beam beam spill events removed from the calibration samples. Test Beam beam spill is 4.2 seconds long and we remove events (in blue) within a 5 seconds window from the start of the beam spill. The remaining events (green) should mostly consist of cosmic particles. This example and the numbers of entries are for the full period 4 Test Beam sample.

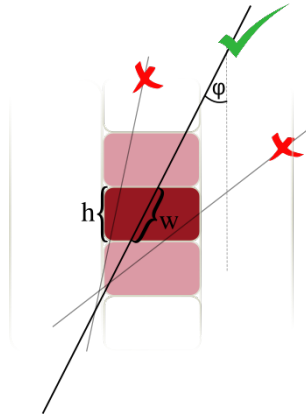


Figure 3: Illustration of the tricell condition. We only use hits that have two surrounding hits in the same plane to be used in the NOvA calibration. This is to ensure a good quality of the pathlength (w) reconstruction, which is calculated from the known cell height (h) and the reconstructed track angle (φ).

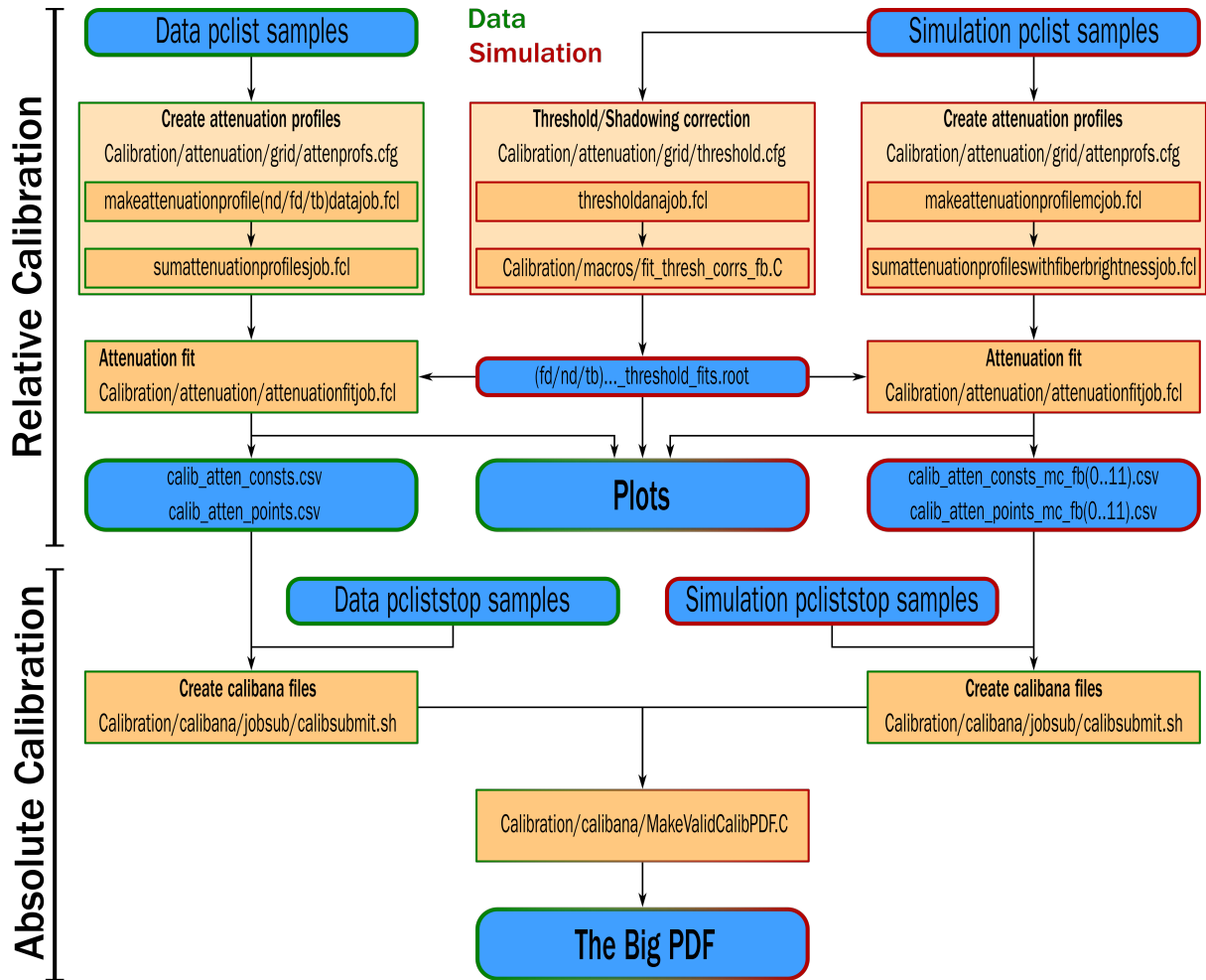


Figure 4: Flow chart showing the jobs (orange background) and files (blue background) needed and produced during the full NOvA calibration process. The left chain is showing the data calibration process (with green border) and is applied to every data calibration sample separately (periods or epochs). The center and right chains are showing the simulation calibration process, which is redone only if there's a change to the detector simulation. The absolute calibration at the bottom combines data and simulation. The entire process is done separately for each NOvA detector.

Tricell condition

4.1 Calibration uncertainties

5 NOvA Test Beam detector calibration

5.1 Overview

History of TB calibration. What led to the final version of TB calibration. What can be done next.

5.2 Test Beam data and simulation

Dates and times when the data taking occurred.

Period naming, possibly epochs (for P3). List of data samples, plus MC samples that were used and pointer to the data-based simulation technote.

Specific running conditions: Underfilled cells Faulty FEBs (Period 2 and Period 3) Why do we do the calibration generally and why do we need to do in for Test Beam specifically

Temperature study (small overview)

5.2.1 Definitions

List all final data and simulation definitions used.

5.3 Detector Brightness

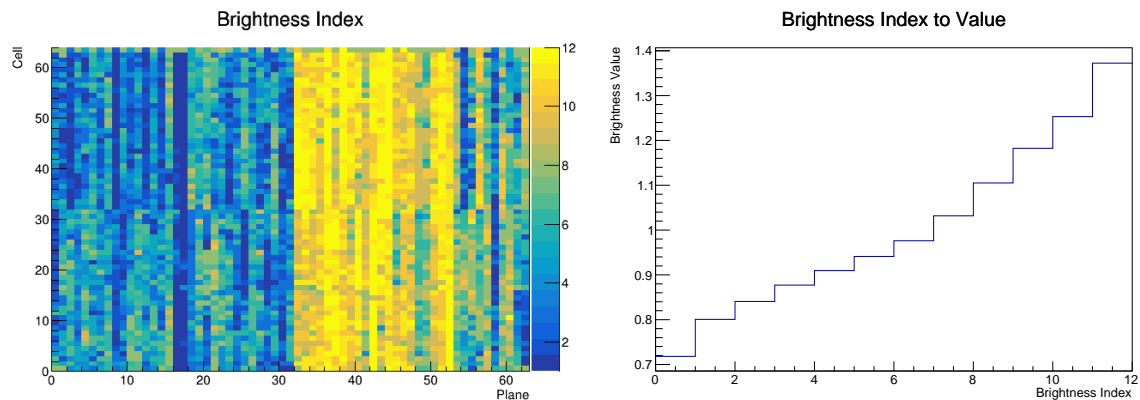


Figure 5: The Test Beam detector is (like the standard NOvA detectors) divided into 12 brightness bins (left plot), each representing a relative difference in energy response (right plot) due to different brightnesses of the fibers, scintillators, or readout.

5.4 Period 1

Only a month of data, only first half of detector filled, primary/secondary beam halo, or over-saturation leading to FEB shutoffs [docdb:38349 and 41331]. Only used for comissioning, not used for any data analysis or calibration.

5.5 Period 2

What was done for the period 2 tb calibration, short overview of what has been done: test beam data were calibrated all at the same time without splitting them to separate epochs. We originally used Teresa's calibration MC sample, but after we saw disagreement, we developed a new MC based off of the period 3 data, which we ended up using for both period 2 and period 3. For fibre brightness we are also using the same MC from period 3 data as it represents the detector in its best condition.

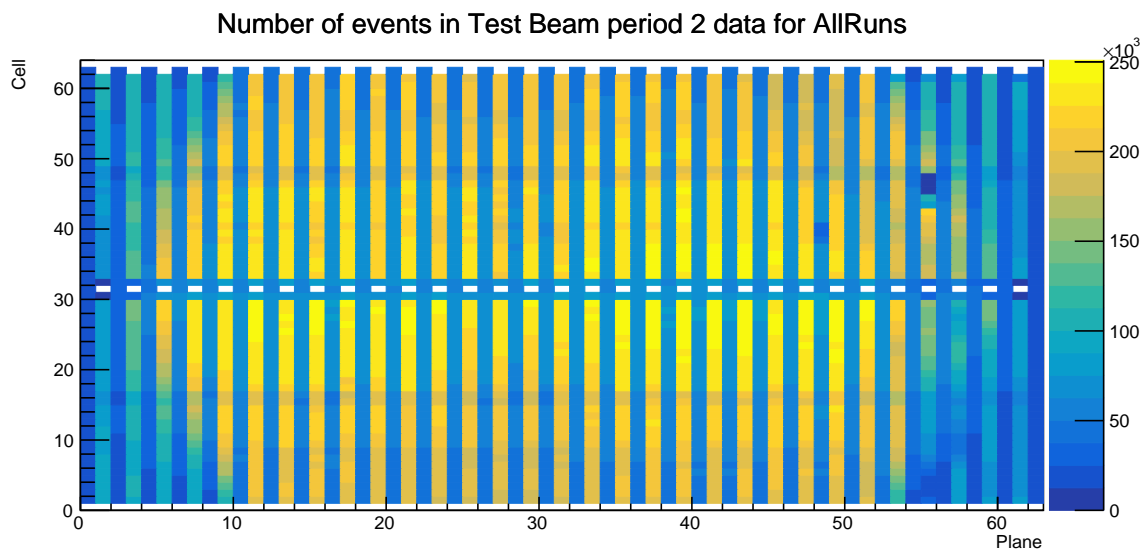


Figure 6: Distribution of events in the period 2 Test Beam data calibration sample.

5.6 Period 3

Separation of Period 3 data into different epochs based on the running conditions (include plot of the running conditions). We are separating data into pre- and post- filling states. We're using only the fully-refilled post-FEB swap data from period 3 as a basis for the simulation creation.

5.6.1 Period 4

5.7 Simulation

We used a data-based simulation of cosmic muons for the Test Beam detector calibration. The details are described in the technote XXX. We used this and this data as a basis and this and this data for the fiber brightness file.

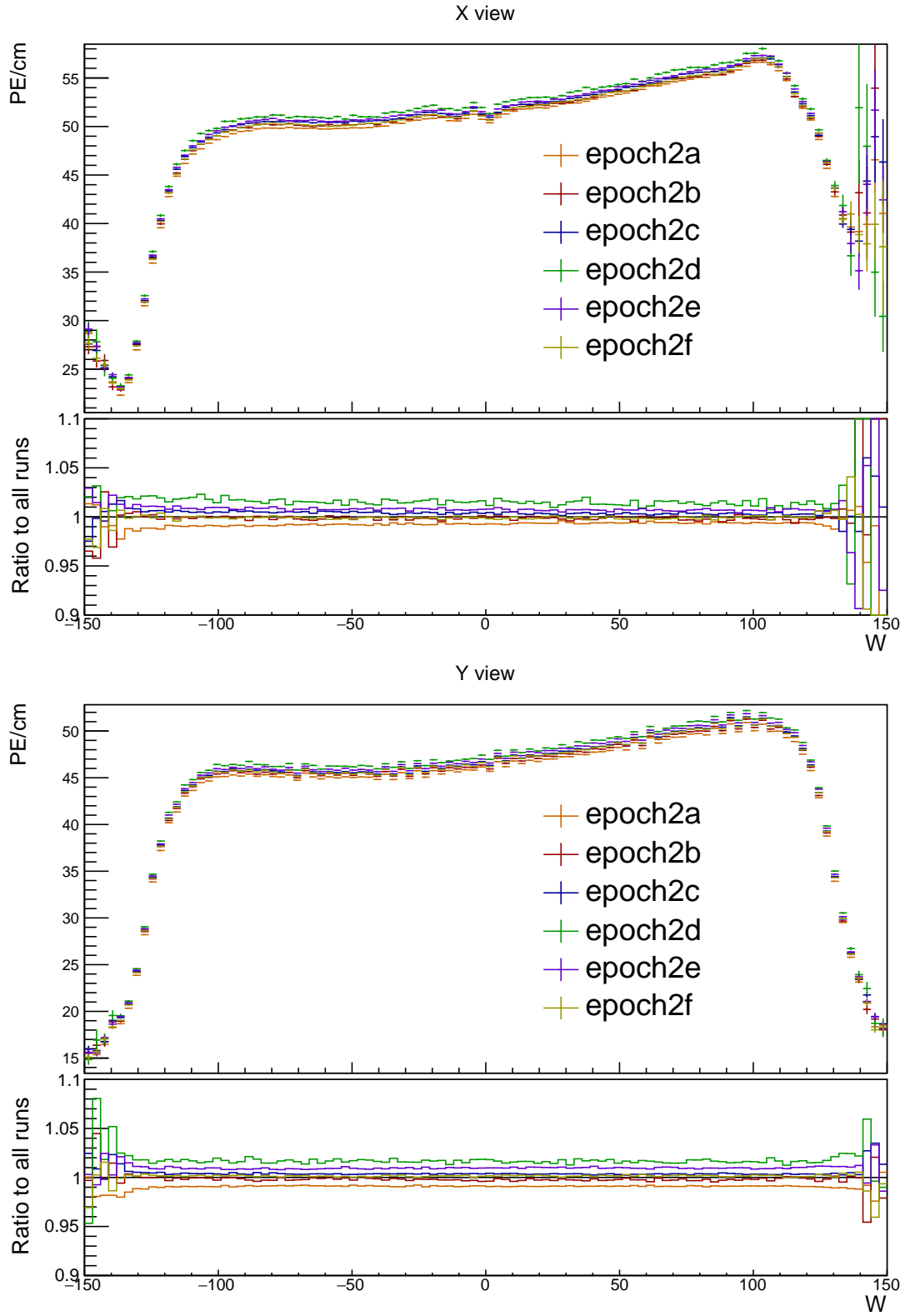


Figure 7: Brightness map

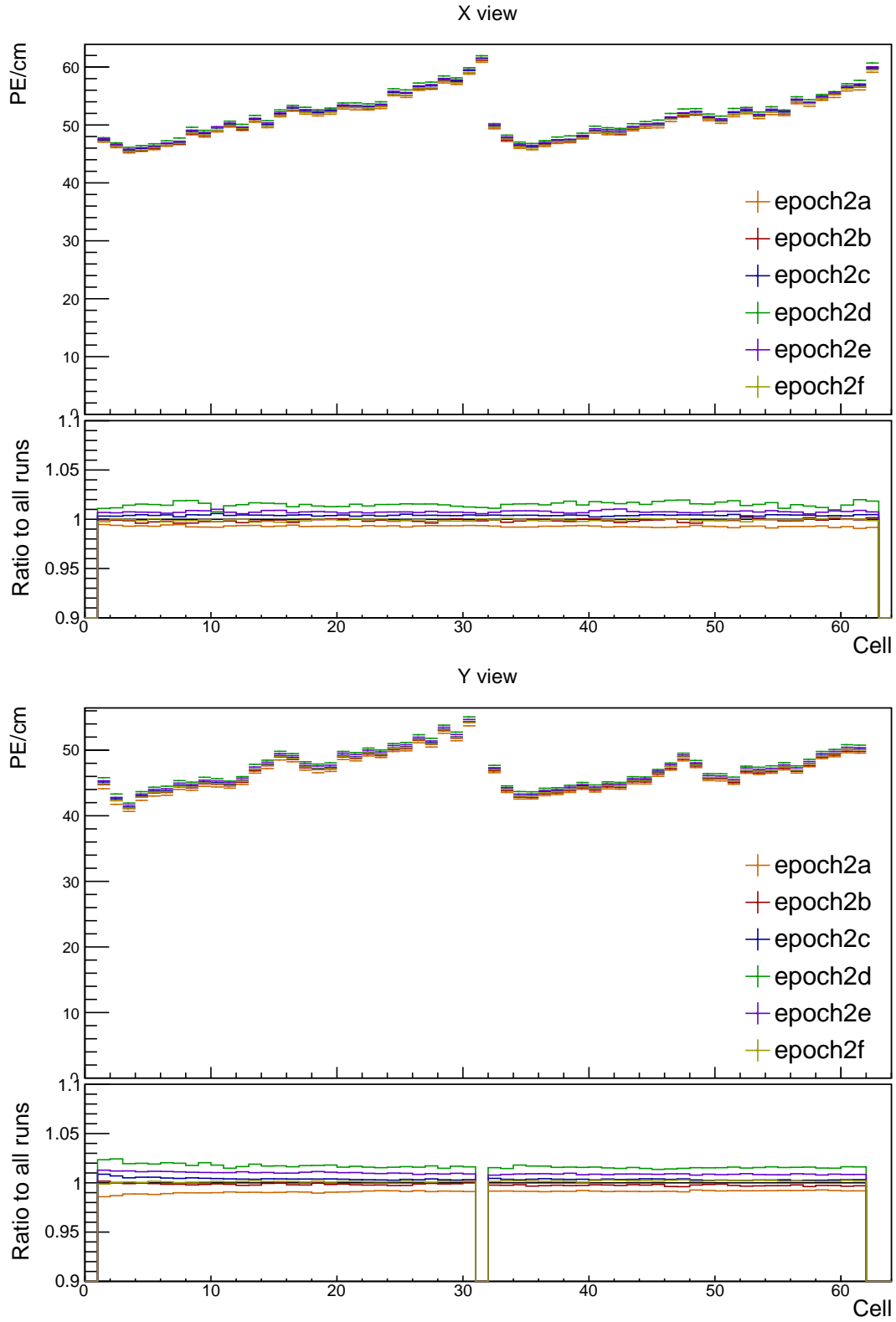


Figure 8: Brightness map

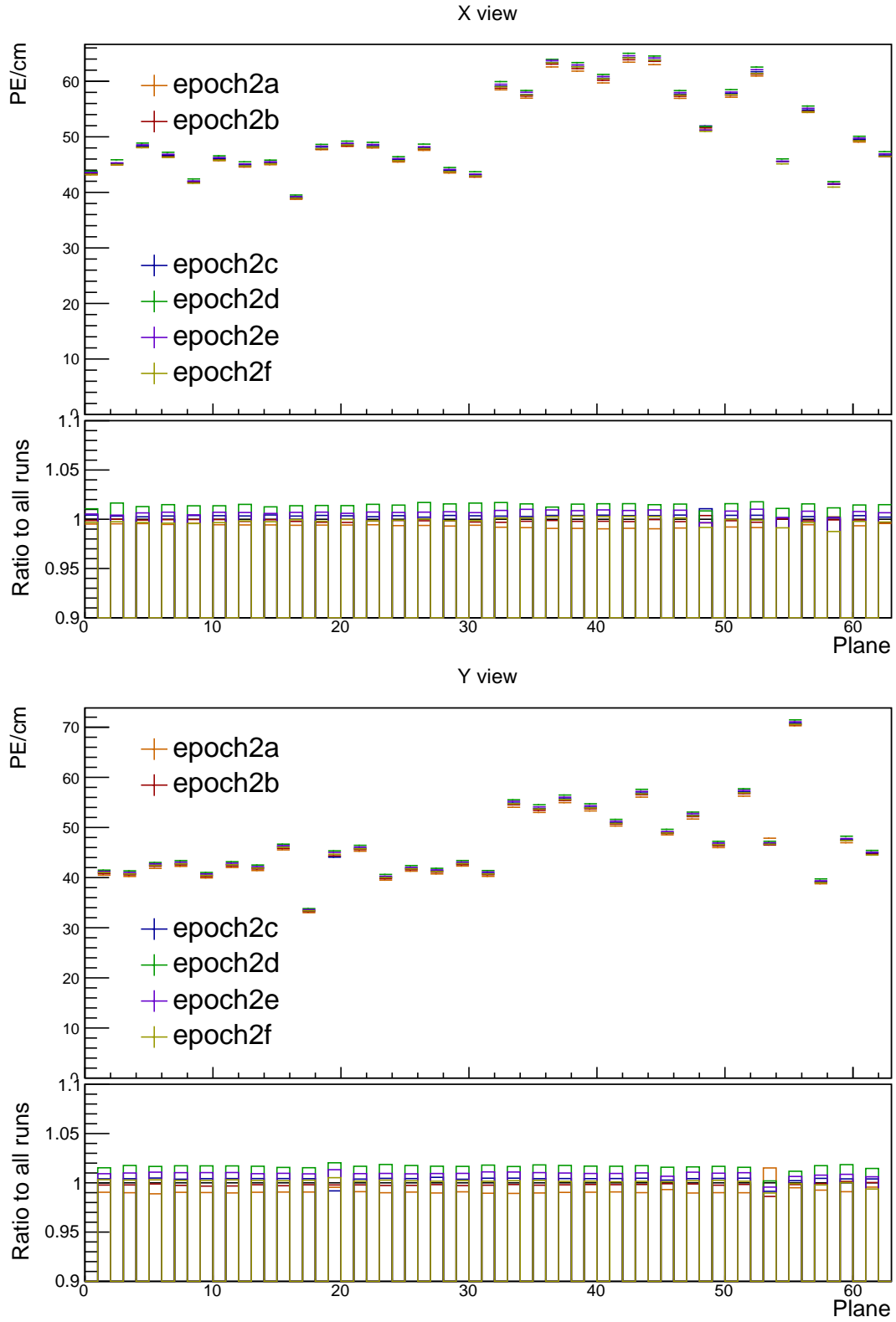


Figure 9: Brightness map

Should I put this into a separate Appendix (or just a stand alone technote) as well? First we needed to select the right data that could serve as a basis for the simulation. Comparing this with the previous version and how did the totlength cut change our cuts. Why do we need to remove the beam events and how do we remove them ($\text{CosZ} < 0.98$). Include plots showing some distributions and how they are affected by the cuts. How many events do we have after the cuts?

5.8 Relative calibration

Attenuation profiles have a constant binnin fNBins=100 (in w), same for ND, FD and TB. This results in an effectively finer binning for TB compared to ND and FD. For FD $w = (-900, +900)$, ND: $(-250, +250)$, TB: $(-150, +150)$. TB: 3cm/bin, ND: 5cm/bin, FD: 18cm/bin. What effect could this have on the relative calibration results? Particularly on the calibration shape?

5.9 Results

Table of final results. Final CSVs are locate in the /nova/ana/testbeam/calibration and they have been included in the vXX.XX calibration tag.

Plots of absolute calibration results

5.10 Validation

Comparisons with older version of calibration and maybe with the FD and ND