

Rapid detector of dark matter transit with GNSS clock data

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Introduction: Dark matter

- In the standard Lambda-CDM cosmology of the universe, only $\sim 5\%$ of the universe consists of ordinary baryonic matter.
 - 26% consists of dark matter
 - 69% consists of dark energy
- Models of galaxies require additional “dark” matter binding them together gravitationally against the centrifugal force of rotation in order to account for observed angular speeds.
- One theory of dark matter posits a mechanism known as *dilatonic* interactions.
 - The dilaton field would couple to the standard model fields, with the side-effect of locally altering the fine structure constant, which determines atomic transition rates.
- A 2015 paper suggested that this theory could be tested with atomic clocks, whose accuracy precisely depends on counting atomic transitions.

Introduction: GNSS system

- Absent the ground-based segment, the GNSS system represents a massive distributed network of atomic clocks with sub-nanosecond precision.
- Proposed dark matter interactions could generate clock biases on the order of nanoseconds.
- Since dark matter is presumed to be weakly interacting (hence the apparent lack of bodies), a dark matter transit would take the form of a moving wavefront.
- If such a wavefront were propagate through the GNSS system, it would leave a signature on the time biases of the satellites.

Approach and related work

- In the last several years, a multi-affiliate research group has unsuccessfully searched for dilaton interactions in GPS data.
- A bulk-processing approach is computationally extensive and simultaneously limited in scope.
- A real-time detection tool could search for a much wider variety of signatures, and quickly alert radio telescope operators of the transit.
- This represents a “responsible data science” approach to physics — one whose primary goal is to inform and facilitate more principled research.

Problem statement

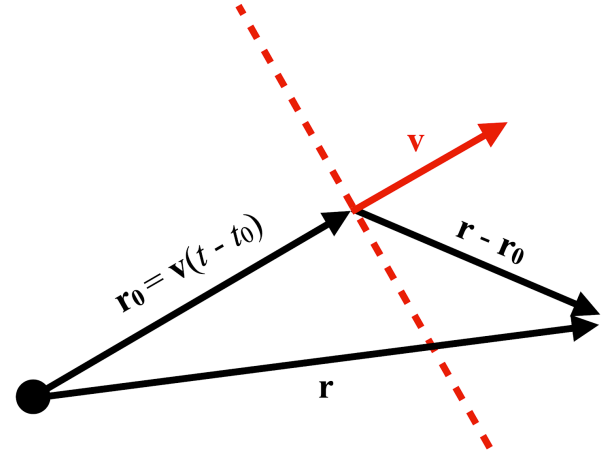
- An infinite flat wavefront propagating through space is described by the expression

$$(\mathbf{r} - \mathbf{v}(t - t_0)) \cdot \frac{\mathbf{v}}{|\mathbf{v}|} = 0$$

- This can be inverted to yield

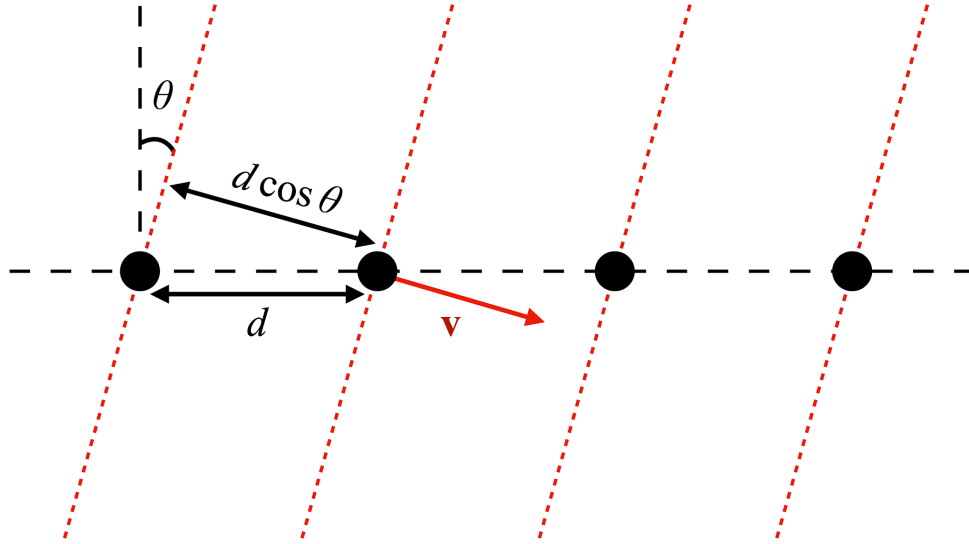
$$t = t_0 + \frac{\mathbf{r} \cdot \mathbf{v}}{|\mathbf{v}|^2}$$

- We can now apply a non-linear least squares fit.



Mathematical formulation: degeneracies

- Near collinearity and coplanarity can cause multiple valid fits.



Approach

- Goals for real-time tool:
 - Rapidity
 - Concurrency
 - Stream processing
- Tool is validated with hardware-in-the-loop testing
- Simulator feeds detector “real-time” data
 - Memory efficiency allows long-term simulation without burdening processor
- Detector assembles list of probable events and passes it to analyzer
 - Events determined by time differencing satellite clock biases
 - Filtering
- Analyzer computes wavefront estimate, attempting to remove bad data / noise measurements
 - Outlier removal
 - Jackknife resampling

Results

Table 1.

Sampling time	v [km/s]	azimuth [deg]	elevation [deg]	t₀ [s]
0.01	399.9958 ± 0.0054	120.0003 ± 0.0011	-45.0000 ± 0.0007	150.0060 ± 0.0005
0.05	400.0185 ± 0.0253	120.0030 ± 0.0053	-44.9989 ± 0.0034	150.0290 ± 0.0024
0.1	400.0597 ± 0.0564	120.0002 ± 0.0116	-44.9997 ± 0.0079	150.0561 ± 0.0054
0.5	400.2843 ± 0.3233	119.9977 ± 0.0669	-44.9957 ± 0.0443	150.2375 ± 0.0311
1.0	400.0609 ± 0.6343	120.2005 ± 0.1314	-45.0166 ± 0.0869	150.5160 ± 0.0610
10.0	405.8272 ± 5.1814	118.4279 ± 1.0875	-45.8324 ± 0.7879	155.4136 ± 0.5054
100.0	438.3547 ± 81.7563	119.4463 ± 13.8810	-41.2214 ± 9.3802	199.5537 ± 6.2997

Conclusion and future work

- Estimate uncertainty scales linearly with sample time.
- Realistic data transmission is far less regular or comprehensive as the simulated transmission.
 - GNSS ephemeris transmissions are on the order of minutes
- Further work would involve modifying the tool structure and algorithms to better suit them to sparse data.
- Expanding data stream paradigm to encompass entire processing pipeline
 - This would allow partial estimates to be formed and refined during the transit itself, giving radio telescopes additional time to observe.

Acknowledgments

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