

# Attoseconds and the exascale: on laser plasma surface interactions



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# Abstract

The commissioning of multi-petawatt class laser facilities around the world is gathering pace. One of the primary motivations for these investments is the acceleration of high-quality, low-emittance electron bunches. Here we explore the interaction of a high-intensity femtosecond laser pulse with a mass-limited dense target to produce MeV attosecond electron bunches in transmission and confirm with three-dimensional simulation that such bunches have low emittance and nano-Coulomb charge. We then perform a large parameter scan from non-relativistic laser intensities to the laser-QED regime and from the critical plasma density to beyond solid density to demonstrate that the electron bunch energies and the laser pulse energy absorption into the plasma can be quantitatively described via the Zero Vector Potential mechanism. These results have wide-ranging implications for future particle accelerator science and associated technologies.



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# A List of Symbols and Abbreviations

**Note** . . . . . All quantities are defined in SI units unless otherwise specified throughout the thesis.

$\alpha$	. . . . . Twiss parameter = $-\frac{\langle x_i x'_i \rangle}{\epsilon_{\text{rms}}^i}$
$a_0$	. . . . . Normalised vector potential = $\frac{e \mathbf{E}_L }{m_e c \omega_L}$
$A$	. . . . . Atomic mass number
$\mathbf{A}$	. . . . . Three-vector potential
$\mathbf{A}^\mu$	. . . . . Four-vector potential = $(\phi/c, \mathbf{A})$
$\beta$	. . . . . Normalised speed = $v/c$ or Twiss parameter = $\frac{\langle x_i \rangle}{\epsilon_{\text{rms}}^i}$
$\beta$	. . . . . Normalised velocity = $\mathbf{v}/c$
$\mathbf{B}$	. . . . . Magnetic field
$\mathbf{B}_L$	. . . . . Magnetic field of a laser pulse
$c$	. . . . . Speed of light = $3.00 \times 10^8 \text{ m s}^{-1}$
$\delta$	. . . . . Skin depth
$\delta(\mathbf{x})$	. . . . . Dirac-delta function
$d\mathbf{s}$	. . . . . Vector line element
$\Delta$	. . . . . Change in a variable
$D$	. . . . . Number of dimensions
$\nabla$	. . . . . Partial derivative = $\hat{\mathbf{x}}_\mu \frac{\partial}{\partial x_\mu}$ , $\mu = x, y, z$
$\epsilon_0$	. . . . . Permittivity of free space = $8.854 \times 10^{-12} \text{ F m}^{-1}$
$\epsilon_{\text{rms}}$	. . . . . Transverse geometric emittance
$\epsilon_{n,\text{rms}}$	. . . . . Transverse normalised emittance
$e$	. . . . . Absolute charge of an electron = $1.602 \times 10^{-19} \text{ C}$
$e^\pm$	. . . . . Electron (-) or positron (+)
$\mathbf{E}$	. . . . . Electric field
$E_S$	. . . . . Schwinger electric field = $1.3 \times 10^{18} \text{ V m}^{-1}$

$\mathbf{E}_L$	Electric field of a laser pulse
$f$	Distribution function
$\mathbf{F}_L$	Lorentz force = $q(\mathbf{E} + \mathbf{v} \times \mathbf{B})$
$\gamma$	Lorentz/gamma factor = $\frac{1}{\sqrt{1-\beta^2}}$ or Twiss parameter = $\frac{\langle x_i'^2 \rangle}{\epsilon_{rms}^i}$ or high-energy photon
$\eta$	Efficiency
$I$	Electromagnetic field intensity
$\mathcal{H}$	Hamiltonian
$\theta$	Angle of incidence of a laser pulse
$\mathbf{J}$	Current density
$\mathbf{k}$	Electromagnetic three-wave vector
$\mathbf{K}^\mu$	Electromagnetic four-wave vector = $(\omega/c\mathbf{k})$
$K$	Boltzmann constant = $1.38 \times 10^{-23} \text{ J K}^{-1}$
$\lambda_D$	Debye length $\equiv \sqrt{\frac{\epsilon_0 K T_e}{n_e e^2}}$
$\lambda_L$	Laser pulse wavelength
$\Lambda_\nu^\mu$	Lorentz transformation matrix
$L$	Length
$\mathcal{L}$	Lagrangian
$\mu_0$	Vacuum permeability = $1.257 \times 10^{-6} \text{ N A}^{-2}$
$m_e$	Mass of an electron = $9.11 \times 10^{-31} \text{ kg}$
$n$	Harmonic order
$n, N$	Number
$n_c$	Plasma critical density = $\frac{\omega_L^2 m_e \epsilon_0}{e^2}$
$n_e$	Plasma electron number density
$\bar{n}_e$	Normalised plasma electron number density = $\frac{n_e}{n_c}$
$n_i$	Plasma ion number density
$\mathbf{n}$	Vector normal to a surface
$N_A$	Avogadro's number
$N_D$	Number of particles within the Debye sphere = $\frac{4}{3}\pi\lambda_D^3 n$
$\omega$	Angular frequency of an oscillation or low energy photon
$\omega_L$	Laser pulse angular frequency

$\omega_p$	Plasma frequency = $\sqrt{\frac{e^2 n_e}{m_e \epsilon_0}}$
$\phi$	Angle of laser polarisation out of the plane of interaction or scalar electromagnetic potential
$\mathbf{p}$	Three-momentum
$P$	Pressure
$\mathbf{P}^\mu$	Four-momentum = $(U/c, \mathbf{p})$
$Q$	Charge
$\rho$	Density
$r_L$	Relativistic Larmor radius = $\frac{\gamma m_e v}{e \mathbf{B} }$
$\mathbf{r}$	Radial vector
$\hat{\mathbf{r}}$	Radial unit vector
$R$	Rate
$\sigma$	Surface area
$S$	Relativistic similarity parameter = $\bar{n}_e/a_0$
$S(\mathbf{x})$	Shape function
$\tau$	Average time or standard deviation of time
$t$	Time
$T$	Kinetic energy
$T_e$	Plasma electron temperature
$u, v$	Speed
$\mathbf{u}, \mathbf{v}$	Velocity
$U$	Energy
$v_\phi$	Phase velocity
$V$	Potential or volume
$w_L$	Beam waist
$w_p$	Quasi-particle weight
$\chi_\gamma$	Photon quantum parameter
$x, y, z$	Standard Cartesian coordinates in 3D
$\mathbf{x}, \mathbf{y}, \mathbf{z}$	Vectors along each axis of the Cartesian coordinate system
$\hat{\mathbf{x}}, \hat{\mathbf{y}}, \hat{\mathbf{z}}$	Unit vectors along each axis of the Cartesian coordinate system
$Z$	Ion charge state in units of $e$

<b>Subscripts . . .</b>	The following are defined the subscripts
C . . . . .	Pseudocapacitor
e . . . . .	Electron
$\gamma$ . . . . .	Photon
L . . . . .	Laser pulse or Longitudinal
T . . . . .	Transvere to a laser pulse
pol . . . . .	Along the polarisation vector of a laser pulse
$\perp$ . . . . .	Perpendicular
$\parallel$ . . . . .	Parallel
$i, j$ . . . . .	Indices
s . . . . .	Species
$x, y, z$ . . . . .	Along the specific Cartesian axes
$\infty$ . . . . .	
<b>The hat symbol</b>	To represent a normalised vecotr
<b>primed coordinates</b>	<del>and reference frames</del>
<b>1D, 2D, 3D . .</b>	One-, two- or three-dimension(al)
<b>BW . . . . .</b>	Breit-Wheeler
<b>CLF . . . . .</b>	Central Laser Facility
<b>CPA . . . . .</b>	Chirped Pulse Amplification
<b>CSE . . . . .</b>	Coherent Synchrotron Emission
<b>FDTD . . . . .</b>	Finite Difference Time Domain
<b>HB . . . . .</b>	Hole Boring
<b>HED . . . . .</b>	High-Energy-Density
<b>HHG . . . . .</b>	High Harmonic Generation
<b>HPC . . . . .</b>	High Performance Computing
<b>OHREX . . . .</b>	ORION High Resolution X-ray
<b>PIC . . . . .</b>	Particle-In-Cell
<b>QED . . . . .</b>	Quantum Electro-Dynamics
<b>ROM . . . . .</b>	Relativistic Oscillating Mirror
<b>RPM . . . . .</b>	Relativistic Plasma Mirror
<b>RR . . . . .</b>	Radiation Reaction
<b>SF-QED . . . .</b>	Strong-Field Quantum Electro-Dynamics
<b>ZVP . . . . .</b>	Zero Vector Potential

# 1

## Introduction

### Contents

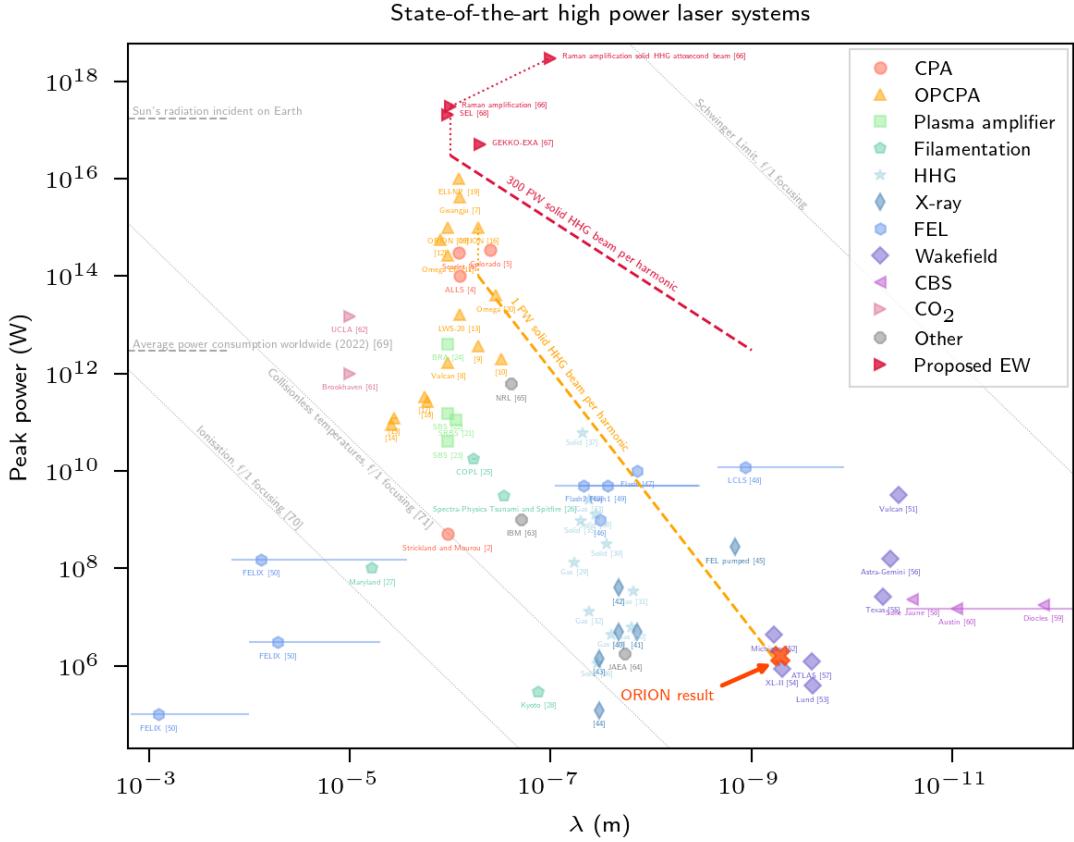
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### 1.1 High-energy-density physics

We sit on the verge of the exa-scale revolution both in terms of computing [1] and laser power. With rapid adoption of the pioneering advancements of



**Figure 1.1:** A by no means exhaustive plot of laser systems across the globe, both commissioned and theorised [2, 4–71] to provide an overview of the parameter space currently and soon to be accessible. The experimental result obtained at the ORION laser facility [16] is presented, and will be discussed in detail in chapter ???. The dashed orange line is the accompanying theoretical prediction. The red dashed line indicates the parameter space accessible via the methods discussed in this thesis when coupled with a sub-etawatt Raman amplified laser beam. Such a beam would provide intensities in the water window many orders of magnitude beyond that currently accessible in state-of-the-art facilities.

Chirped Pulse Amplification (CPA) by Strickland and Mourou [2] and of Optical Parametric Chirped Pulse Amplification [3], multi-petawatt facilities have now been commissioned and built across the globe and plans for sub-exawatt facilities are underway. Figure 1.1 details the parameter space accessible by current and proposed facilities [2, 4–68], clearly, there is no indication of wavering interest or support for the field of High-Energy-Density (HED) physics.

HED physics is the laboratory study of the behaviour of matter with a pressure above 10 GPa, approximately one million atmospheres and containing free electrons

not confined to a solid state [72], typically in the plasma state of matter. HED conditions are found for a vast range of densities and temperatures (from zero to a million million Kelvin) operating in both the quantum and relativistic realms. The applications are equally diverse, including but not limited to inertial confinement fusion, particle acceleration for scientific or medical purposes and light sources for diagnostic tools. Ubiquitous in the natural universe, beyond our solar system, all that can be observed in the sky is radiation emitted from HED plasmas [**chen**].

This thesis concerns itself with a specific interaction, that of a high-power short-pulse laser incident on a flat solid target. Through the process discussed in this thesis, using state-of-the-art 10 PW laser facilities such as [19], electron bunches and light pulse of unprecedented charge and brightness can be produced, both with attosecond duration, thus uncovering new avenues for attosecond resolution diagnostics and to access the Schwinger limit. Seemingly counter-intuitively, as the laser power increases, via relativistic effects and for certain conditions, greater coherency in the electron dynamics can be observed and the signals amplified. The red dashed line of figure 1.1 could be accessed using next generation sub-ettawatt facilities. Before delving into this fascinating phenomenon, the remainder of this chapter will provide some of the relevant background information. Chapter 2 introduces the Zero Vector Potential of attosecond absorption in this laser-solid interaction. The following chapter discusses the associated process of Surface High Harmonic Generation and the results of a recent experiment at the ORION laser facility where the absolute intensity of individual X-ray harmonics was measured and compared to the theory.

## 1.2 Electromagnetism fundamentals

The spatio-temporal propagation of electric  $\mathbf{E}(t, \mathbf{x})$  and magnetic  $\mathbf{B}(t, \mathbf{x})$  fields must satisfy Maxwell's equations [73]

$$\nabla \cdot \mathbf{B} = 0, \quad (1.1a)$$

$$\nabla \cdot \mathbf{E} = \frac{\rho}{\epsilon_0}, \quad (1.1b)$$

$$\nabla \times \mathbf{B} = \mu_0 \mathbf{J} + \mu_0 \epsilon_0 \partial_t \mathbf{E}, \quad (1.1c)$$

$$\nabla \times \mathbf{E} = -\partial_t \mathbf{B}. \quad (1.1d)$$

Here,  $\epsilon_0 = 8.85 \times 10^{-12} \text{ F m}^{-1}$  and  $\mu_0 = 1.26 \times 10^{-6} \text{ N A}^{-2}$  are the vacuum permittivity and permeability respectively and  $\rho(t, \mathbf{x})$  and  $\mathbf{J}(t, \mathbf{x})$  the total charge and current densities of charged particles present.

A particle with charge  $q$  and velocity  $\mathbf{v}$  in the presence of electromagnetic fields experiences the Lorentz force,

$$\mathbf{F}_L = q(\mathbf{E} + \mathbf{v} \times \mathbf{B}). \quad (1.2)$$

The electromagnetic fields can be obtained from the scalar,  $\phi$ , and vector,  $\mathbf{A}$ , potentials as [74]

$$\mathbf{E} = -\nabla \phi - \partial_t \mathbf{A}, \quad (1.3)$$

$$\mathbf{B} = \nabla \times \mathbf{A}. \quad (1.4)$$

### The Vlasov-Maxwell system of equations

A collisionless and fully ionised plasma is fully described in the kinetic description by the Vlasov-Maxwell system of equations [75]. Each plasma species,  $s$ , of particles with mass  $m_s$  and charge  $q_s$  is described by its distribution function  $f_s(t, \mathbf{x}, \mathbf{p})$  at time  $t$ , position  $\mathbf{x}$  and momentum  $\mathbf{p} = m_s \gamma \mathbf{v}$ . The distribution satisfies the Vlasov equation, that is,

$$(\partial_t + \frac{\mathbf{p}}{m_s \gamma} \cdot \nabla + \mathbf{F}_L \cdot \nabla_{\mathbf{p}}) f_s = 0, \quad (1.5)$$

where  $\mathbf{F}_L$  is the Lorentz force given in equation 1.2. The electric  $\mathbf{E}(t, \mathbf{x})$  and magnetic  $\mathbf{B}(t, \mathbf{x})$  fields that generate the force must satisfy Maxwell's equations (equations 1.1).

This self-consistent system of equations describes the dynamics of plasma particles within electromagnetic fields. The particles modify the fields via their charge and current densities,

$$\rho(t, \mathbf{x}) = \sum_s q_s \int d^3 p f_s(t, \mathbf{x}, \mathbf{p}), \quad (1.6)$$

and

$$\mathbf{J}(t, \mathbf{x}) = \sum_s q_s \int d^3 p \mathbf{v} f_s(t, \mathbf{x}, \mathbf{p}), \quad (1.7)$$

respectively.

### 1.3 The definition of a classical plasma

As outlined in F. Chen's definitive textbook 'Introduction to Plasma Physics and Controlled Fusion' [73], a plasma must fulfil three criteria, namely,

1. Ionisation: a plasma must consist of both charged and neutral particles, of course this alone cannot define a plasma, any gas will contain some degree of ionisation, however, with reference to figure 1.1, clearly, modern high power laser systems will instantaneously fully ionise a target upon incidence;
2. Quasineutrality: while locally there can be (often extreme) electromagnetic forces and charge concentrations at work, over the length scales of the plasma, such forces are screened out and the plasma bulk remains net neutral in charge;
3. Collective behaviour: unlike in a gas, where collisions will dominate the dynamics, the particles in a plasma generate electromagnetic fields that interact at a distance and thus a particle's motion depends not only on its immediate vicinity but on the surrounding plasma conditions, indeed often it is the so-called *collisionless* plasmas where collisions can be safely neglected that are of most interest, as is the focus of this thesis.

These conditions can be quantitatively described by the Debye length, the plasma parameter and the plasma frequency as laid out in the following sections.

### 1.3.1 The Debye length

The Debye length describes the extent to which a plasma can shield electromagnetic fields within and so remain quasineutral. Consider an infinitely extending plasma with a test charge placed at some point, then what would be the scalar potential  $\phi(\mathbf{x})$  around it? If the plasma had no kinetic energy, the charged particles would arrange themselves immediately adjacent to the test charge and once this equilibrium state was reached there would be no electromagnetic fields present. More realistically, the plasma will have some temperature, likely a very large temperature and so some particles will be able to escape the potential of the test charge and thus leak electromagnetic fields into the plasma bulk. Poisson's equation (equation 1.1b in the static case) reads

$$\epsilon_0 \nabla^2 \phi = -e(Zn_i - n_e), \quad (1.8)$$

where  $\epsilon_0 = 8.854 \times 10^{-12} \text{ F m}^{-1}$  is the permittivity of free space,  $e = 1.602 \times 10^{-19} \text{ C}$  is the charge of an electron,  $Z$  is the plasma ion charge in units of  $e$  and  $n_i$  and  $n_e$  are the number densities of plasma ions and electrons.

Since the electrons are significantly more mobile than the ions due to their lower mass, it is in general the electrons and not the ions that respond to the test charge and the ions can be assumed to provide a constant background of positive charge density. If the number density of electrons follows a Boltzmann temperature distribution in the presence of a potential energy  $-e\phi$ , then

$$n_e = n_{e,0} e^{e\phi/KT_e}, \quad (1.9)$$

where  $n_{e,0}$  is the electron number density far from the test charge,  $n_i = n_{e,0}/Z$ ,  $K = 1.38 \times 10^{-23} \text{ J K}^{-1}$  is the Boltzmann constant and  $T_e$  is the electron temperature in Kelvin. Note that in plasmas it is very common for different species to have differing temperatures depending on the mechanism for energy absorption and the timescales for collisions compared to the timescale of the study.

Substituting equation 1.9 into equation 1.8 and Taylor expanding the exponential term in the limit that the plasma is weakly coupled ( $e\phi \ll KT_e$ ),

$$\nabla^2\phi = \frac{\phi}{\lambda_D^2}, \quad (1.10)$$

where

$$\lambda_D \equiv \sqrt{\frac{\epsilon_0 K T_e}{n_e e^2}}, \quad (1.11)$$

is the *Debye length* and describes the thickness of the charge sheath surrounding the test charge. For quasineutrality to hold for the plasma bulk, its spatial dimensions,  $L$ , must extend beyond a few Debye lengths, *i.e.*

$$L \gg \lambda_D. \quad (1.12)$$

### 1.3.2 The plasma parameter

In order for the derivation of section 1.3.1 to be statistically valid, there must be a large number of charged particles within the shielding sheath. The number of particles within the *Debye sphere* is

$$N_D = \frac{4}{3}\pi\lambda_D^3 n, \quad (1.13)$$

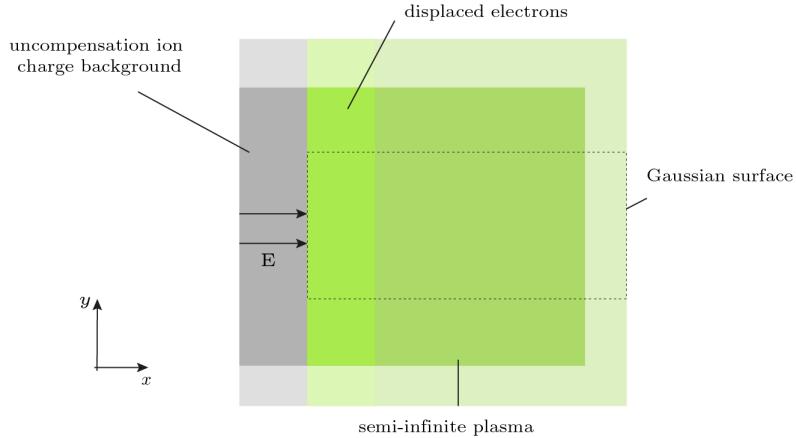
where,  $N_D$  is the *plasma parameter*. Note that, as discussed above, in most cases it is most suitable to choose the number density  $n$  to be the number density of electrons,  $n_e$ . To ensure the plasma is suitably ionised (criterion 1) and that the plasma engages in collective behaviour (criterion 3),

$$N_D \ggg 1. \quad (1.14)$$

### 1.3.3 Collisionality and the plasma frequency

Collective behaviour not only depends on the ability for large numbers of particles to interact via electromagnetic forces but also that these forces dominate over collisions in describing particle trajectories. Taking  $\omega$  as the typical frequency of plasma oscillations and  $\tau$  as the average time between collisions, for a plasma (as opposed to a gas)

$$\omega\tau > 1 \quad (1.15)$$



**Figure 1.2:** The electrons of a semi-infinite plasma are displaced inwards by some external force leaving in their wake an uncompensated space charge of ‘immobile’ positive ions. By constructing a Gaussian surface along the dashed line, using Gauss’ Law, the electric field associated with the positive space charge can be calculated.

is required. It now remains to determine what is the typical frequency of collisions in a given plasma. While the types of plasma waves and their associated frequencies of oscillation are multitudinous, the characteristic frequency, the *plasma frequency*,  $\omega_p$ , is the most straightforward. It describes the response of electrons to charge imbalances within an infinite uniform plasma at rest in the absence of magnetic fields or temperature fluctuations. As noted in section 1.3.1, the ions provide a constant background of positive charge.

Consider an semi-infinite plasma existing for  $x > 0$ , with electron density  $n_e$  and ion density  $n_e/Z$  of charge state  $Z^1$ . Suppose the electron fluid is displaced by some perfectly isotropic force into the plasma bulk a distance  $(\Delta x)\hat{\mathbf{x}}$  as in figure 1.2. The total charge of displaced electrons within a surface area of  $\sigma$  is

$$Q = -en_e\sigma\Delta x. \quad (1.16)$$

Applying the integral form of Gauss’ law (from equation 1.1b) to the surface detailed in figure 1.2, the uncompensated charge leads to

$$-\sigma E\hat{\mathbf{x}} = \frac{Q}{\epsilon_0}\hat{\mathbf{x}} = -\frac{en_e\sigma\Delta x}{\epsilon_0}\hat{\mathbf{x}} \quad (1.17)$$

---

<sup>1</sup>This description has direct relevance to the Zero Vector Potential mechanism which will be made clear in chapter 2.

at the electron surface. By the Lorentz force, equation 1.2, the displaced electrons will experience a restoring force,  $-eE\hat{x}$ , perpendicular to the surface due to the electron-ion charge imbalance. The equation of motion for electrons on that surface is therefore

$$m_e \frac{d^2\Delta x}{dt^2} = -eE = -\frac{e^2 n_e}{\epsilon_0} \Delta x. \quad (1.18)$$

Equation 1.18 clearly describes a simple harmonic oscillator with a characteristic frequency given by the plasma frequency,

$$\omega_p = \sqrt{\frac{e^2 n_e}{m_e \epsilon_0}}. \quad (1.19)$$

## 1.4 The Lawson-Woodward theorem

The Lawson-Woodward theorem states that there can be no net electron energy gain using laser fields [76], quite at odds with one of the primary aims of this thesis, that is, the acceleration of electrons. There are, however, several conditions that must be met, namely,

1. The interaction region is infinite;
2. The interaction occurs in a vacuum;
3. The electron is ultra-relativistic ( $v \approx c$ ) along the acceleration gradient;
4. No electro- or magnetostatic fields are present;
5. Non-linear effects are neglected.

Several of these will be applicable to the various accelerations of electrons considered. It is this final condition that is most damning to the application of the theorem. Throughout this thesis the ultra-relativistic laser pulses under consideration ensure non-linear effects cannot be neglected. It is indeed such non-linearities that are of most interest.

## 1.5 Laser-solid density plasma linear interaction

Consider small transverse electromagnetic waves propagating through a plasma. Linearising equation 1.2 for a single plasma electron by assuming only small field variations and thus small velocity variation,

$$m_e \dot{\mathbf{v}}_e = -e\mathbf{E}. \quad (1.20)$$

Combining the time derivative of equation 1.1c and the curl of equation 1.1d,

$$\nabla \times \nabla \times \mathbf{E} = -\mu_0 \dot{\mathbf{J}} - \mu_0 \epsilon_0 \ddot{\mathbf{E}}. \quad (1.21)$$

Considering only fast oscillations such that ions are effectively immobile,

$$\mathbf{J} = -n_e e \mathbf{v}_e \quad (1.22)$$

and using the identity  $\nabla \times \nabla \times \mathbf{E} = \nabla(\nabla \cdot \mathbf{E}) - \nabla^2 \mathbf{E}$ ,

$$\nabla(\nabla \cdot \mathbf{E}) - \nabla^2 \mathbf{E} = -\frac{\mu_0 n_e e^2}{m_e} \mathbf{E} - \mu_0 \epsilon_0 \ddot{\mathbf{E}}. \quad (1.23)$$

Assuming plane wave solutions of the form

$$\mathbf{E} = \mathbf{E}_0 e^{i(\mathbf{k} \cdot \mathbf{x} - \omega t)}, \quad (1.24)$$

where  $\mathbf{k}$  is the wave-vector and  $\omega$  the frequency of oscillations and noting the waves are transverse  $\mathbf{k} \cdot \mathbf{E} = 0$ ,

$$|\mathbf{k}|^2 \mathbf{E} = -\frac{\mu_0 n_e e^2}{m_e} \mathbf{E} + \mu_0 \epsilon_0 \omega^2 \mathbf{E} \quad (1.25)$$

and hence the dispersion relation for electromagnetic waves propagating in a plasma is

$$\omega^2 = c^2 |\mathbf{k}|^2 + \omega_p^2. \quad (1.26)$$

Equation 1.26 exhibits a *cutoff* dependent on the plasma density via  $\omega_p$ . The *critical density*,  $n_c$ , is defined as the density above which a laser pulse of frequency  $\omega_L$  cannot propagate through a plasma. This occurs for  $\omega_L = \omega_p$ , thus,

$$n_c = \frac{m_e \epsilon_0 \omega_L^2}{e^2}. \quad (1.27)$$

As the plane wave has a spatial dependence  $\sim \exp(\mathbf{k} \cdot \mathbf{x})$ , if  $n_e > n_c$ ,  $\mathbf{k}$  is imaginary and the wave no longer propagates through the plasma and instead exponentially attenuates over a skin depth,

$$\delta = \frac{1}{|\mathbf{k}|} = \frac{c}{\sqrt{\omega_p^2 - \omega_L^2}} \quad (1.28)$$

and is reflected. For typical high power lasers with wavelengths in the visible or near-infrared, fully ionised solids tend to have densities well above the critical density and thus

## 1.6 Relativity

Modern high-power lasers operate in the domain of relativistic mechanics and in general interactions are highly non-linear. It is useful to introduce some of the basic principles of relativity. Note that while there has been growing interest in the curvature of spacetime by relativistic lasers [77], this effect remains undetectable at present. Thus, throughout this thesis the inner product of 4-tensors is defined using the Minkowski Metric [74].

Many useful quantities can be arranged into contravariant four-vectors that undergo a Lorentz transformation for a change of frame of reference [74], specifically,

$$\mathbf{A}'^\mu = \Lambda_\nu^\mu \mathbf{A}^\mu, \quad (1.29)$$

where  $\Lambda_\nu^\mu$  is the appropriate Lorentz transformation, and primed symbols typically denote boosted frames of reference. Without loss of generality, the coordinate system can be defined such that the boosted frame travels along the  $\mathbf{x}$ -axis with respect to the initial frame. Thus, the Lorentz transformation is defined

$$\Lambda_\mu^\nu = \begin{pmatrix} \gamma & -\beta\gamma & 0 & 0 \\ -\beta\gamma & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}. \quad (1.30)$$

Generally, a *beta factor* is a normalised speed or velocity of an object,

$$\beta = \frac{v}{c}, \quad (1.31)$$

Symbol	Name	Components	Invariant
$\mathbf{X}^\mu$	4-displacement	$(ct, \mathbf{x})$	$-c^2\tau^2$
$\mathbf{A}^\mu$	4-potential	$(\phi/c, \mathbf{A})$	
$\mathbf{J}^\mu$	4-current density	$(c\rho, \mathbf{J})$	$-c^2\rho_0^2$
$\mathbf{K}^\mu$	4-wave vector	$(\omega/c, \mathbf{K})$	
$\mathbf{P}^\mu$	4-momentum	$(U/c, \mathbf{p})$	$-m^2c^2$

**Table 1.1:** Four-vectors of relevance to this thesis. New parameters are the proper time,  $\tau$ , the proper charge density,  $\rho_0$ , energy,  $U = \gamma mc^2$ , three-momentum,  $\mathbf{p} = \gamma m\mathbf{v}$ .

here it refers to the frame velocity and its associated Lorentz or *gamma factor* is

$$\gamma = \frac{1}{\sqrt{1 - \beta^2}}. \quad (1.32)$$

Four-vectors relevant this thesis are listed in table 1.1. Transformations of electromagnetic fields under reference frame boosts are given in appendix A.1 and Maxwell's equations are Lorentz covariant.

Focusing now on the 4-potential  $\mathbf{A}^\mu$ , choosing the Lorenz gauge,

$$\partial_\mu \mathbf{A}^\mu = \nabla \cdot \mathbf{A} + \frac{1}{c^2} \partial_t \phi = 0, \quad (1.33)$$

then Maxwell's equations can be written

$$\partial_\nu \partial^\nu \mathbf{A}^\mu = -\frac{1}{c^2 \epsilon_0} \mathbf{J}^\mu. \quad (1.34)$$

Equation 1.34 can then be solved to yield

$$\mathbf{A}(\mathbf{x}, t) = \frac{\mu_0}{4\pi} \int \frac{\mathbf{J}(\mathbf{x}', t_a)}{|\mathbf{x} - \mathbf{x}'|} d^3 \mathbf{x}', \quad (1.35)$$

where  $t_a = t - |\mathbf{x} - \mathbf{x}'|/c$  is the advanced time.

### 1.6.1 Ultra-relativistic similarity theory

Consider a relativistically intense laser pulse normally incident on a collisionless plasma as in figure 1.2 and neglect ion motion. The electron distribution is fully described by the Vlasov equation (equation 1.5) with the self-consistent electric

and magnetic fields satisfying Maxwell's equations (equations 1.1). Suppose the incident laser pulse has an initial vector potential

$$\mathbf{A}(t=0) = \mathbf{a}((y^2 + z^2)/R^2, x/c\tau) \cos k_L x. \quad (1.36)$$

This envelope form for the potential,  $\mathbf{a}((y^2 + z^2)/R^2, x/c\tau)$ , is sensible provided  $k_L R \gg 1$  and  $\omega_L \tau \gg 1$ , where  $R$  is the focal spot radius and  $\tau$  the pulse duration. For fixed laser envelope, the laser-plasma dynamics depends on just four dimensionless variables: the normalised focal spot size,  $k_L R$ , the normalised pulse duration,  $\omega_L \tau$ , the normalised laser vector potential amplitude

$$a_0 = \max \left| \frac{e\mathbf{a}}{m_e c^2} \right|, \quad (1.37)$$

in terms of the peak laser electric field amplitude  $\mathbf{E}_L$ ,

$$a_0 = \frac{e|\mathbf{E}_L|}{m_e c \omega_L}, \quad (1.38)$$

and the normalised plasma density

$$\bar{n}_e = \frac{n_e}{n_c}. \quad (1.39)$$

By normalising the system of equations, details given in appendix, and combining these last two expressions into the *relativistic similarity parameter*,

$$S = \frac{\bar{n}_e}{a_0}, \quad (1.40)$$

it is possible to show that in the ultra-relativistic limit ( $a_0 \gg 1$ ), the dynamics of the system is similar for constant  $S$  [78] with plasma electrons following similar trajectories where

$$\mathbf{p} \sim a_0. \quad (1.41)$$

There is also a more physical meaning to the  $S$  parameter. Consider again section 1.5 on the propagation of linear electromagnetic waves through a plasma but now for the case of an ultra-relativistic laser pulse. For an electron rotating in an electromagnetic field,

$$\mathbf{F}_\perp = \gamma m_e \mathbf{a}_\perp, \quad (1.42)$$

where  $\mathbf{a}_\perp$  is the acceleration perpendicular to the motion and thus the response of the electrons is reduced by a factor of  $\gamma$ . While some find the *relativistic mass* correction to be somewhat unhelpful nomenclature for the phenomenon [74], it has nevertheless become commonplace within the literature of relativistic plasma physics [70]. Turning the handle, one finds that the relativistic plasma frequency is

$$\omega_p^{\text{rel}} = \sqrt{\frac{e^2 n_e}{\gamma m_e \epsilon_0}}. \quad (1.43)$$

Using equation 1.41, and taking  $v \approx c$ , then  $\gamma \approx a_0$  and the normalised relativistic cutoff density is simply  $S$ . Thus, the ultra-relativistic similarity parameter is just a measure of the overdensity of a plasma once relativistic corrections have been applied, *i.e.* for  $S > 1$ , a laser pulse will be reflection, however for  $\bar{n}_e > 1$  and  $S < 1$ , one enters the regime of relativistically self-induced transparency [79]. It is now possible to define the parameter space of interest in this thesis: relativistic laser plasma surface interactions occur for  $a_0 \gg 1$  and  $S > 1$ .

### 1.6.2 Relativistic lasers and plasmas

The descriptor *relativistic* is applied liberally in this thesis. When applied to electromagnetic fields or laser pulses it refers to

$$a_0 \geq 1. \quad (1.44)$$

When applied to particles, their Lorentz factors are

$$\gamma = \frac{1}{\sqrt{1 - \beta^2}} \geq 2, \quad (1.45)$$

corresponding to a speed,  $u \geq 0.87c$ . *Ultra-relativistic* implies these quantities are much larger than the conditions provided. A relativistic laser pulse will accelerate electrons to relativistic velocities in a fraction of a laser pulse cycle. Consider an electron in the presence of a uniform electric field of magnitude  $a_0 = 100$ , an intensity accessible by current state of the art laser facilities. The work done on that particle by the field is

$$T = (\gamma - 1)m_e c^2 = \int \mathbf{E} \cdot d\mathbf{x}, \quad (1.46)$$

The field will accelerate an electron to relativistic velocities in a distance less than 1 % of a corresponding laser pulse wavelength.

### 1.6.3 Conservation of generalised transverse momentum

Consider a holonomic system of  $N$  relativistic particles under the influence of electromagnetic forces. A particle  $j$  with charge  $e_j$  and mass  $m_j$  experiences a scalar potential,

$$V_j = e_j(\phi - \mathbf{A} \cdot \mathbf{v}_j) \quad (1.47)$$

and hence the system is described by the Lagrangian [80]

$$\mathcal{L} = \sum_{j=1}^N \left( -m_j c^2 \sqrt{1 - \beta_j^2} - e_j(\phi - \mathbf{A} \cdot \mathbf{v}_j) \right), \quad (1.48)$$

The generalised momentum corresponding to coordinate  $x_j$  is

$$p_{j,x} = \frac{\partial L}{\partial \dot{x}_j} = \frac{m_j \dot{x}_j}{\gamma_j} + e_j A_x, \quad (1.49)$$

thus, the generalised momentum describes both the linear mechanical momentum and the momentum of the electromagnetic field. Via Noether's theorem, if  $L$  is independent of  $x_j$ , *i.e.* spatially homogeneous along  $x$  for particle  $j$ , then

$$\dot{p}_{j,x} = 0 \quad (1.50)$$

since

$$\frac{d}{dt} \left( \frac{\partial L}{\partial \dot{x}_j} \right) = \frac{\partial L}{\partial x_j}. \quad (1.51)$$

Taking the Lorenz gauge, consider a linearly polarised Gaussian laser pulse, with axis of polarisation along  $x$  incident on a solid target at rest. Then  $A_x$  is approximately constant along  $x$  near the beam centre<sup>2</sup>. Integrating equation 1.50, fully constraining the gauge by setting the potential to zero at initially, and noting that there is no linear momentum at the target, the generalised transverse momentum conservation equation for an electron in the laser field is

$$p_T = eA, \quad (1.52)$$

---

<sup>2</sup>Constant relative to the scale of typical electron trajectories in such an interaction.

where  $p_T$  is the electron momentum along the polarisation axis of the laser pulse and  $A$  is the laser pulse 3-vector potential amplitude. As a sanity check, this expression complies with the ultra-relativistic similarity result of equation 1.41.

Note that this is only valid provided the electron does not radiate along the direction of polarisation as discussed by Sokolov *et al* [81]. The implications of *Radiation Reaction* are discussed in the following section.

## 1.7 QED effects

Next generation laser facilities will enable the testing of decades old theoretical predictions of Strong-Field Quantum Electro-Dynamics (SF-QED). Already Fedeli *et al* have shown in simulations that current PW-class laser facilities can access the regime using an all-optical set up based on laser-solid surface interactions [82]. The Schwinger Limit  $E_S = 1.32 \times 10^{18} \text{ V m}^{-1}$  is field intensity at which the vacuum becomes non-linear. If by some means an electron can be directed towards a plane electromagnetic wave, by consideration of the Lorentz transformations of electromagnetic fields (equations A.1), it is possible that provided the electron is sufficiently relativistic, in its own frame of rest it will ‘see’ electromagnetic fields intense enough to access vacuum non-linearities. The first two frontiers of SF-QED that will be accessible are Radiation Reaction and multi-photon Breit-Wheeler electron pair production. Brief introductions to these phenomena are now presented.

### 1.7.1 High-energy photon emission and radiation reaction

When a charged particle undergoes an acceleration, it emits electromagnetic radiation. If the electromagnetic field is sufficiently strong, *i.e.* approaching the Schwinger Limit in the rest frame of the particle, then a non-negligible fraction of the particle momentum can be transferred to the emitted high energy photon, substantially impacting the dynamics of the accelerated particle. This back reaction is known as Radiation Reaction (RR).

Smilei (detailed in the following section) implements the process of high-energy photon emission as Inverse Compton Scattering on the basis of several assumptions [83], namely,

1. radiating particles are ultra-relativistic and therefore radiation is emitted in the direction of travel of the particle;
2. the field varies slowly over the timescale of photon emission, this is the *locally-constant field approximation* and requires ultra-relativistic field strengths;
3. but they are small with respect to the Schwinger Limit, specifically requiring the invariants  $\sqrt{c^2\mathbf{B}^2 + \mathbf{E}^2}$  and  $\sqrt{c\mathbf{B} \cdot \mathbf{E}} < E_S$ ;
4. real particles radiated independently of their neighbours, this requires the emitted wavelength being shorter than the typical inter-particle spacing.

Provided such conditions hold, the rate of photon emission depends on two invariants [84], the electron quantum parameter

$$\chi = \frac{\gamma}{E_S} \sqrt{(\mathbf{E} + \mathbf{v} \times \mathbf{B})^2 - (\mathbf{v} \cdot \mathbf{E})^2/c^2}, \quad (1.53)$$

where  $\mathbf{v}$  is the electron velocity and the emitted photon quantum parameter

$$\chi_\gamma = \frac{\gamma_\gamma}{E_S} \sqrt{(\mathbf{E}_\perp + \mathbf{c} \times \mathbf{B})^2 - (\mathbf{c} \cdot \mathbf{E})^2/c^2}, \quad (1.54)$$

where  $\gamma_\gamma$  is the normalised photon energy  $= \hbar\omega_\gamma/m_e c^2$ . The exact relationship is complex and in the the fully quantum domain ( $\chi > 1$ ), is it not practical to solve the integrations required for all particles and instead values are extracted from precalculated tables and into a Monte-Carlo algorithm.

### 1.7.2 Multi-photon Breit-Wheeler pair production

Multi-photon Breit-Wheeler pair production, also known as non-linear Breit-Wheeler is the decay of a high energy photon, typically produced via RR, into an electron-positron pair in the presence of a strong electromagnetic field, explicitly,

$$\gamma + n\omega \rightarrow e^- + e^+. \quad (1.55)$$

The strength of the effect is dependent on the Lorentz invariant photon quantum parameter, equation 1.54 Cite Smilei and say, in a constant E-field, the rate of pair production increases rapidly up to  $\chi_\gamma \approx 10$  at which point it saturates and slowly reduces.

## 1.8 Simulating the interaction

Modelling laser-plasma interactions is a notoriously challenging endeavour; due to the complexity of the many-bodied systems involved (a fully ionised centimetre cubed block of plastic contains on the order of  $10^{23}$  particles) and the stochasticity of particle motion, it is frequently impossible to construct models *ab initio*. Instead, hydrodynamic simulation codes such as HYADES [85] and FLASH [86] and Particle-In-Cell simulation codes such as Smilei [75], Osiris [87] and EPOCH [88] are used to construct phenomenological models and to direct experimentation.

### 1.8.1 Supercomputing resources

Modern High Performance Computing (HPC) systems are poised to enter the exascale regime ( $> 10^{18}$  Floating Point Operations Per Second). With limited improvements in microprocessor technologies, such power is achieved through massive parallelisation across processing units. To enable the study of the dynamics of up to billions of macroparticles, PIC codes test the limits of current supercomputing architectures. ARCHER2, the UK’s national supercomputer came online in November 2021, with it delivering over ten times the resources of its predecessor (ARCHER) [89]. An HPE Cray EX supercomputing system with a peak performance estimated at  $28 \text{ Pflops s}^{-1}$  across 5860 nodes each with dual AMD EPYCTM 7742 64-core processor for a total of 750,080 cores, ARCHER2 was able to supply the resources required to run the costly PIC simulations for this research. The substantially cheaper HYADES simulations were performed using the Rutherford Appleton Laboratory’s SCARF HPC cluster [90].

### 1.8.2 Particle-In-Cell codes

#### Discretisation of the Vlasov-Maxwell equations

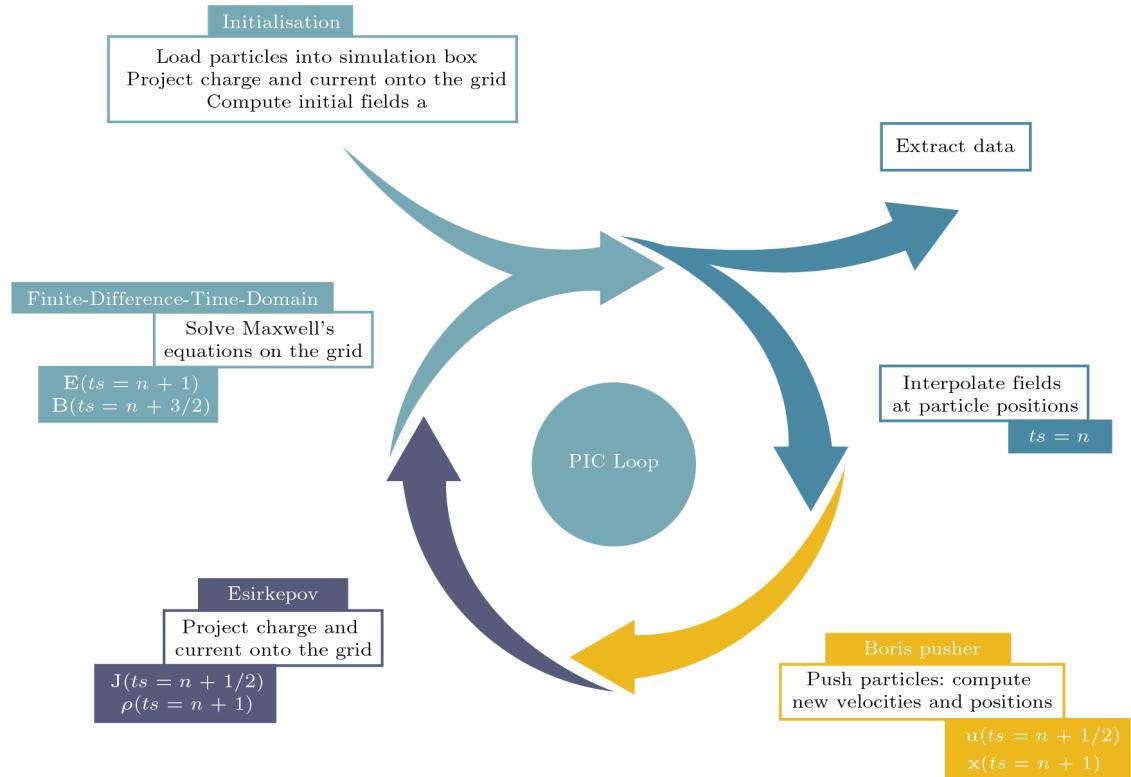
Finding numerical solutions to the Vlasov-Maxwell equations is no straightforward task, while codes exist that are capable, such as Valis [91], the requirement of high resolution in both position and momentum is exceedingly costly and the use of such codes is limited with respect to their size, duration and spatial dimensions. A more tractable approach is to discretise the distribution function into  $N_s$  *quasi-particles*<sup>3</sup>. These are often referred to as *macro-particles* in practice and typically represent a large number of real particles, such that

$$f_s(t, \mathbf{x}, \mathbf{p}) = \sum_{p=1}^{N_s} w_p S(\mathbf{x} - \mathbf{x}_p(t)) \delta(\mathbf{p} - \mathbf{p}_p(t)), \quad (1.56)$$

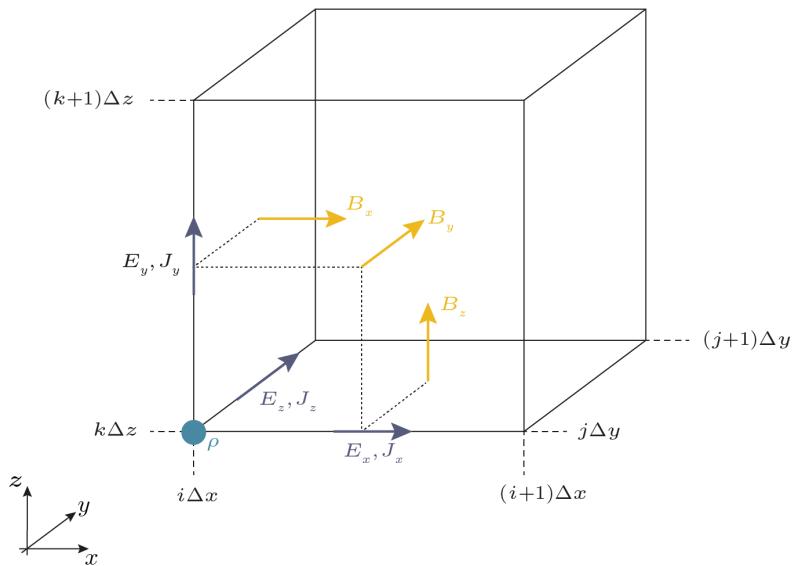
where  $w_p$  is the quasi-particle's weight,  $\mathbf{x}_p$  and  $\mathbf{p}_p$  are its position and momentum respectively,  $\delta$  is the Dirac-delta distribution and  $S(\mathbf{x})$  the shape-function chosen to represent the quasi-particle. The Vlasov equation is then integrated along the continuous trajectories of the quasi-particles while Maxwell's equations are solved on a discrete spatial grid of *cells*. Such a code is aptly named a *Particle-In-Cell* (PIC) code. A schematic of the standard PIC code algorithm is presented in figure 1.3. After particle and field initialisation, fields are interpolated at particle positions. The well-established momentum-conserving *Boris pusher* algorithm computes the new macro-particle velocities and positions [93]. Particles are advanced in time using a *leap-frog* scheme, where positions are defined at integer,  $n$ , time steps and momenta at half integer,  $n + 1/2$ . The charge conserving Esirkepov algorithm [94] projects the new charge and current densities onto the grid to then solve Maxwell's equations on the grid using the Finite-Difference-Time-Domain approach [95]. To ensure space and time centering of the electromagnetic field derivatives in Maxwell's equations, electric and magnetic fields are discretised on the staggered *Yee grid* as represented in figure 1.4. with electric fields defined at integer time steps and magnetic fields at half-integer time steps.

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<sup>3</sup>Originally introduced by Langdon and Birdsall as *clouds* [92].



**Figure 1.3:** A schematic of the PIC code loop and the algorithms performed from time step,  $ts$ , from  $n$  to  $n + 1$ .



**Figure 1.4:** A representation of the staggered 3D Yee grid for the cell at  $(i\Delta x, j\Delta y, k\Delta z)$  for spatial centering of the curl operations, including the locations where all system properties are defined.

## Smilei

Smilei (for Simulating Matter Irradiated by Light at Extreme Intensities) is a modern collaborative, massively-parallel, fully relativistic and open source plasma physics PIC code and the major workhorse for this thesis<sup>4</sup>. Produced by M. Grech's team(CHECK THIS) at École Polytechnique [75], its recent development was motivated by the rapid advancements of multi-petawatt facilities both globally and locally with the recent completion of the 10 PW Apollon laser facility and by the availability of supercomputing power which has ‘skyrocketed’ in recent years [75]. Indeed, the 3D PIC simulations discussed in this thesis required the parallelisation of almost 20 % of the computing resources available on ARCHER2.

## Reference units

Given the broad range of magnitudes linked to multi-petawatt and femtosecond laser pulses, solid density plasmas, micrometer wavelengths, and attosecond electron bunches, it becomes significantly more convenient to transform them into dimensionless normalised values. Smilei operates in such units. This normalisation is not chosen *a priori*, instead results can be scaled by an arbitrary reference angular frequency. This is extremely useful when working with boosted frames of reference. As this thesis focuses on the interaction of a laser pulse with plasma, the laser pulse angular frequency,  $\omega_L$  is set as the frequency of reference. A list of the most common normalisations are given in table 1.2.

## Simulation parameters

*Silver-Müller* boundary conditions for the simulation box edges [96] are able to absorb and inject electromagnetic waves and particles. Note that there can be non-physical reflection of electromagnetic waves at such boundaries leading to some error.

The quasi-particle shape function  $S(\mathbf{x})$  determines the projection of particle charge onto the grid. It is symmetric in all dimensions with respect to  $\mathbf{x}$  and

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<sup>4</sup>At points benchmarks against the EPOCH and Osiris PIC codes were performed.

Units of	SI units	Normalisation
velocity	$\text{m s}^{-1}$	$c$
charge	C	$e$
mass	kg	$m_e$
momentum	$\text{kg m s}^{-1}$	$m_e c$
energy/temperature	J	$m_e c^2$
time	s	$\omega_L^{-1}$
length	m	$c/\omega_L$
number density	$\text{m}^{-3}$	$n_c$
electric field	$\text{V m}^{-1}$	$m_e c \omega_L / e$

**Table 1.2:** Smilei normalisations for common quantities with the laser angular frequency  $\omega_L$  set at the reference angular frequency.

extends over  $n$  cells of width  $\Delta x$  in each direction where  $n$  is the interpolation order. It can be written as a product across  $D$  dimensions,

$$S(\mathbf{x}) = \prod_{\mu=1}^D s^{(n)}(x^\mu). \quad (1.57)$$

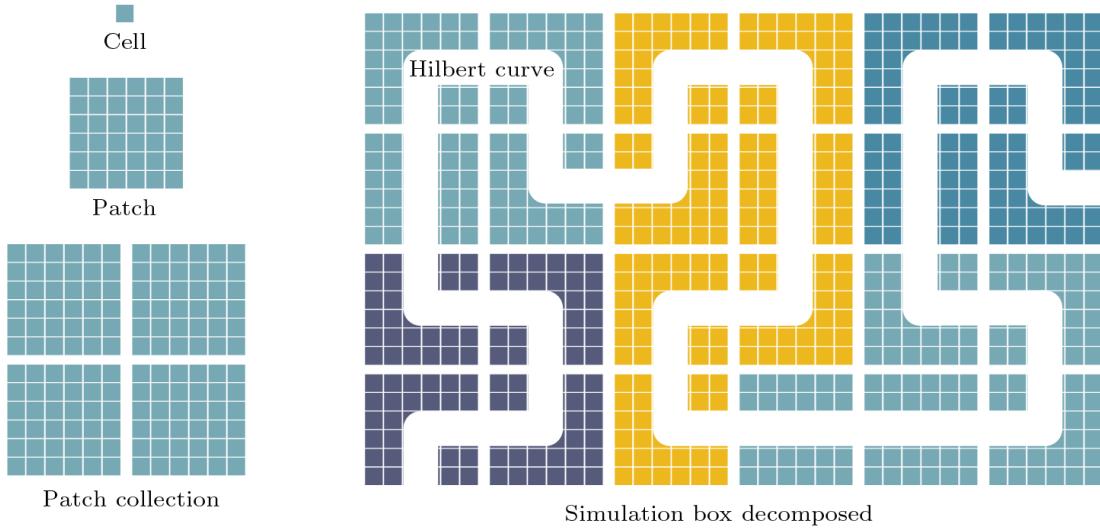
Smilei implements orders 2, 3 and 4, the explicit shape functions are

$$s^2(n) = \begin{cases} \frac{1}{\Delta x} \left(1 - \left|\frac{x}{\Delta x}\right|\right) & \text{if } |x| \leq \Delta x, \\ 0 & \text{otherwise,} \end{cases} \quad (1.58a)$$

$$s^3(n) = \begin{cases} \frac{3}{4\Delta x} \left(1 - \frac{4}{3} \left(\frac{x}{\Delta x}\right)^2\right) & \text{if } |x| \leq \frac{1}{2}\Delta x, \\ \frac{9}{8\Delta x} \left(1 - \frac{2}{3} \left|\frac{x}{\Delta x}\right|\right)^2 & \text{if } \frac{1}{2}\Delta x < |x| \leq \frac{3}{2}\Delta x, \\ 0 & \text{otherwise,} \end{cases} \quad (1.58b)$$

$$s^4(n) = \begin{cases} \frac{2}{3\Delta x} \left(1 - \frac{3}{2} \left(\frac{x}{\Delta x}\right)^2 + \frac{3}{4} \left|\frac{x}{\Delta x}\right|^3\right) & \text{if } |x| \leq \Delta x, \\ \frac{4}{3\Delta x} \left(1 - \frac{1}{2} \left|\frac{x}{\Delta x}\right|^3\right) & \text{if } \Delta x < |x| \leq 2\Delta x, \\ 0 & \text{otherwise.} \end{cases} \quad (1.58c)$$

While the correct implementation of collisions in PIC codes remains an open problem [97], Smilei has implemented relativistic binary collisions between macroparticles using a Monte-Carlo based scheme [98]. The aforementioned QED processes of Inverse Compton scattering and non-linear Breit-Wheeler pair production are included using the built-in Smilei packages [75]. These processes can lead to cascades



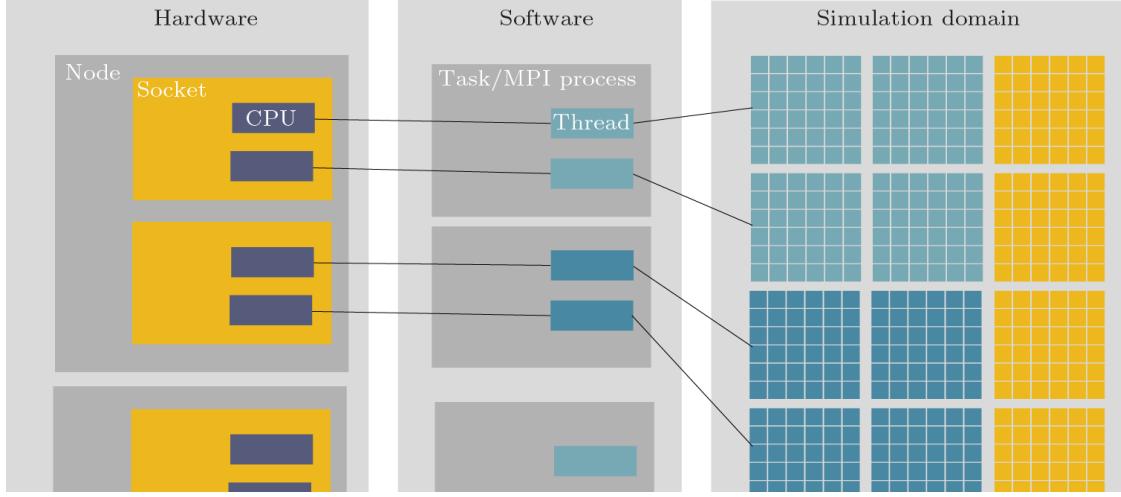
**Figure 1.5:** Smilei simulation box decomposition into cells, patches and MPI patch collections. Cells are grouped into patches, patches are grouped into MPI patch collections. MPI patches are assigned contiguously along the Hilbert curve.

of many particles being added to the simulations. Macroparticle merging can increase simulation efficiency and reduce the memory footprint. Smilei implements such a scheme, inspired by that designed by Vranic *et al* [99], that is computationally efficient and conserves energy, momentum and charge within a cell. While Smilei contains modules to handle ionisation, with reference to figure 1.1 these are deemed unnecessary for the laser intensities considered in this thesis.

### Parallelisation in practice

In the following discussion, where differences in language occur, objects are given as their ARCHER2 (Smilei) names. The Smilei simulation box consists of a grid of cells as in figure 1.5. The box is decomposed into *patches* consisting of many cells. Patches are arranged into *MPI patch collections* assigned contiguously along a Hilbert curve.

Archer consists of many *CPUs (cores)* that can each perform computational tasks. CPUs are grouped together into *nodes*. Memory is shared within a node such that all CPUs (cores) in a node can operate on the same data. When optimised, ARCHER2's memory in each node is split into 8 *sockets*. These 8 sockets each perform a *task (MPI process)*. Each task (MPI process) has 16 CPUs (cores) assigned that each perform a *thread*. A thread is a sequence of instructions from the program.



**Figure 1.6:** Representation of the interaction of the ARCHER2 hardware and software components when running Smilei. Each CPU is assigned a thread, each task is carried out by one socket. A task tackles one MPI patch collection with threads working through patches.

Each task (MPI process) handles one *MPI patch collection*. Threads work through patches. Figure 1.6 represents this division of labour. Threads do not need to wait for other threads in their process to tackle new patches in their MPI patch collection. This is a form of local *dynamic load balancing*. If an MPI patch collection is overloaded, a patch is offloaded to a contiguous MPI patch collection.

There are several considerations for simulation optimisation. Tasks (MPI processes) should always be assigned more patches than threads. To apply the Hilbert curve, it is necessary that the number of patches in each dimension is a power of 2. The less cells in a patch, the more efficient the load balancing, however, the cost of synchronisation between patches increases, although generally the gain in efficiency from load balancing by increasing patch number will win out at later times with increasing efficiency for increased frequency of load balancing [75]. Note that smaller patches are preferable when there are small regions with large numbers of particles, as in laser-solid surface interactions, however, the minimum patch size is dependent on the shape function of the macro-particles.

### Sources of error inherent to PIC codes

Despite their relatively intuitive interpretation, PIC codes are famously finickety and prone to errors, most notably that of numerical self heating<sup>5</sup>. There are three conditions required for stability.

To ensure stability, or at least to minimise instability, there are several conditions which must be met. Naturally, the time step,  $\Delta t$ , and cell size ( $\Delta x \Delta y \Delta z$ ) must adequately capture all interesting features of a given simulation. Typically such features are plasma wave oscillations,

$$\Delta t \omega_p \ll 1, \quad (1.59)$$

and laser pulse electromagnetic field oscillations or higher order harmonics of the laser pulse if that is of interest, for the  $n$ th harmonic

$$\Delta t \omega_L n \ll 1. \quad (1.60)$$

Note this also ensures that macroparticles are smaller than the wavelengths of the system, a requirement to ensure they mimic real particles [100]. Relativistic PIC codes must satisfy the much-acclaimed [101] Courant-Friedrichs-Lowy condition,

$$\frac{1}{c\Delta t} > \sqrt{\frac{1}{(\Delta x)^2} + \frac{1}{(\Delta y)^2} + \frac{1}{(\Delta z)^2}}, \quad (1.61)$$

thus preventing light and relativistic particles from crossing a cell in one timestep and generating numerical Cerenkov radiation [102]. As with real plasmas and real particles, to avoid numerical charge fluctuation and ensure collective behaviour of macroparticles,

$$\frac{4}{3}\pi\lambda_D^3 n_{\text{macro}} = N_{D,\text{macro}} \gg 1, \quad (1.62)$$

where  $n_{\text{macro}}$  is the macroparticle density [102]. To avoid numerical heating, the cell size must resolve the Debye length,

$$\frac{\lambda_D}{\Delta x} \geq 1, \quad (1.63)$$

---

<sup>5</sup>Standard PIC code algorithms are charge and momentum conserving but not energy conserving [75].

failure to do so may cause plasma self-heating until the Debye length matches the cell size [102]. Interestingly, Brackbill *et al* [103] observed in their simulations that setting  $\lambda_D/\Delta x = 1$  was most effective at reducing spurious heating. Arber *et al* [104] performed extensive simulations exploring this instability. If the Debye length is not resolved, after an initial period of rapid self-heating, the temperature increases linearly and can be modelled as

$$\frac{dT_{\text{eV}}}{dt_{\text{ps}}} = \alpha_H \frac{n_{23}^{3/2} \Delta x_{\text{nm}}^2}{N_{\text{ppc}}}, \quad (1.64)$$

where  $T_{\text{eV}}$  is the temperature in electron volts,  $t_{\text{ps}}$  is the time in picoseconds,  $n$  is the plasma electron number density in units of  $10^{23} \text{ cm}^{-3}$ ,  $\Delta x_{\text{nm}}$  is the cell size in nanometres and  $N_{\text{ppc}}$  is the number of macroparticles per cell with  $\alpha_H$ , the heating coefficient, determined from their simulations. For a top-hat macroparticle shape function they found  $\alpha_H = 3000$  with an order of magnitude reduction for every increase in order of the shape function and further improvements from using current smoothing techniques. Note that the heating curves are roughly self similar at all points and thus while equation 1.64 was established in the linear regime only, its scalings remain useful to compare simulations at all times.

The final instability that shall be discussed is the *finite grid instability*. This is the aliasing error associated with particles properties being deposited at grid points. Is most catastrophic for cold drifting plasmas and depends on the *beam Debye length*,

$$B = \frac{u}{\omega_p \Delta x}, \quad (1.65)$$

where  $u$  is the beam speed. While their theory predicts stability for  $B > 0.25$ , Brackbill *et al* [103] observed instability growth for all beam temperatures in their simulations, although they found the percentage error is a small fraction for  $B > 10$ .

# 2

## The Zero Vector Potential Absorption Mechanism

### Contents

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### 2.1 Motivations and an overview

Throughout the history of experimental science, light has always been the primary tool of investigation and discovery. Through the creation of synchrotron radiation sources and more recently XFELs, electron bunches have been employed to create increasingly specialised light sources for the study of matter of all kinds. At SLAC, the United States' forefront electron accelerator, applications hail from many

disciplines: science, medicine, industry and homeland security. At Diamond Light Source, the UK's national synchrotron science facility, studies range from novel drugs to ancient paintings. Unsurprisingly, extensive research efforts have been enlisted to produce electron bunches of ever greater charge, energy and coherency. Multi-petawatt laser facilities are now available across the globe for the study of laser-plasma interactions in the ultra-relativistic regime  $a_0 \gg 1$ . Here, there is a novel method for high-charge electron bunch creation. Via relativistic effects, a laser pulse organises the electrons at the surface of a solid density plasma into coherent bunches that can be ejected at high speeds. Those discussed in this thesis have properties comparable to those of forefront accelerators but on ultra-short timescales. The quality, charge and duration of the attosecond electron bunches described here would enable the study of the most fundamental interactions of our universe. Electron bunch formation from solid targets has received much interest in recent years [105–112] with some experimental evidence for attosecond electron bunches from intense laser-solid interactions [113–115]. This interest stems partially from their ability to produce higher charge bunches at lower intensities compared to gas targets [113]. Here is proposed a new mass-limited target setup to generate electron bunches of extreme charge density. These electron bunches are fully characterised in 3D PIC simulations to compare their quality to those of existing electron bunch production methods. Their energies are described via an extended version of the ZVP Mechanism and the corresponding implications for laser to plasma energy absorption are also considered including the intersection of ZVP with SF-QED effects of next generation laser facilities. ZVP theory has direct relevance to High Harmonic Generation (HHG) due to the known intrinsic link between electron bunches and the reflection mechanism [116, 117].

These results have excited the community to perform experiments to realise these electron bunches for the creation of ultra-bright X-ray pulses: after the recent successful campaign at the ORION laser facility, studying HHG from such electron bunches, details are presented here for the upcoming experiment to observe the electron bunches directly at the GEMINI PW laser at the Central Laser Facility, UK.

This Chapter is organised as follows. Section 2.2 outlines the ZVP mechanism including quantitative calculations of system properties. In Section 2.3.1 the first 3D PIC simulations to observe the ZVP mechanism are presented. Further PIC simulations are presented in Section 2.4 focusing now on the hot electron bunches produced via the mechanism their properties, energy scalings and implications for SF-QED. Section ?? details the planned experiment to observe the ZVP mechanism at the GEMINI-PW laser facility including. Finally, concluding remarks are given in Section 2.5.

## 2.2 Introduction

Of primary interest in this thesis is the interaction of a relativistically intense short-pulse laser interacting with a solid density plasma target with a sharp density gradient. Now is presented one model for the interaction: the ZVP mechanism as proposed by *Baeva et al* [118] and later developed by *Savin et al* [117, 119]. Alongside the theories of the Relativistic Electron Spring [120] and Coherent Synchrotron Emission [121], ZVP is a post-ponderomotive model of attosecond absorption and reflection. The model considers a quasi-static surface equilibrium state analogously to ion acceleration in the hole boring [122] and light sail regimes [123].

An explanation for the absorption of laser energy into dense plasmas was first suggested by Wilks and Kruer [124]. A ponderomotive mechanism where plasma electrons are heated directly by the laser pulse via the  $\mathbf{J} \times \mathbf{B}$  force. This thesis is interested in the so-called ‘post-ponderomotive’ regime where the frequency of relativistic plasma oscillations ( $\omega_p \sim \sqrt{S}$ ) are greater than the  $\mathbf{J} \times \mathbf{B}$  induced plasma electron oscillations at  $2\omega_L$ . The plasma electrons’ response is then fast enough to compensate the ponderomotive pressure of the laser pulse with the formation of electrostatic fields between electrons and ions and so respond adiabatically to the applied  $\mathbf{J} \times \mathbf{B}$  force. Hence, plasma electrons cannot be heated directly by the laser pulse. Interestingly, this condition suggests a criterion for the ZVP regime,  $S > 4$ , slightly more constraining than  $S > 1$  as is typically stated for the ZVP regime [119]. The bulk plasma must have  $S > 1$  to prevent relativistic transparency and so

ensure a laser-surface interaction. Then the ponderomotive pressure of the laser will typically compress the front surface to densities such that the overdensity condition is satisfied, provided the target is of sufficient thickness. Note that to neglect the pre-adiabatic formation phase requires a sufficiently steep density gradient around the relativistic critical density surface (where  $S = 1$ ) to shift the main interaction to a region where this condition on the overdensity is satisfied. Pre-plasma formation and scale length will be discussed in great detail in the following chapters. For now, assume the technology exists to control this phenomenon.

Provided all conditions are met, the ponderomotive pressure of the laser pulse compresses electrons at the front surface of the plasma and so shifts the laser-plasma surface interaction to plasma densities well beyond the relativistic critical density, leaving in its wake a positive space charge of ions. This electron-ion charge separation leads to the formation of a *pseudo-capacitor* electrostatic field. Having entered a regime of adiabaticity, the plasma skin layer is confined within a potential well consisting of the ponderomotive pressure of the laser pulse and the Coulomb potential of the pseudocapacitor field. Thus is formed a high density and longitudinally thin electron bunch (sometimes referred to as an electron sheath in the literature [120]) at the plasma surface.

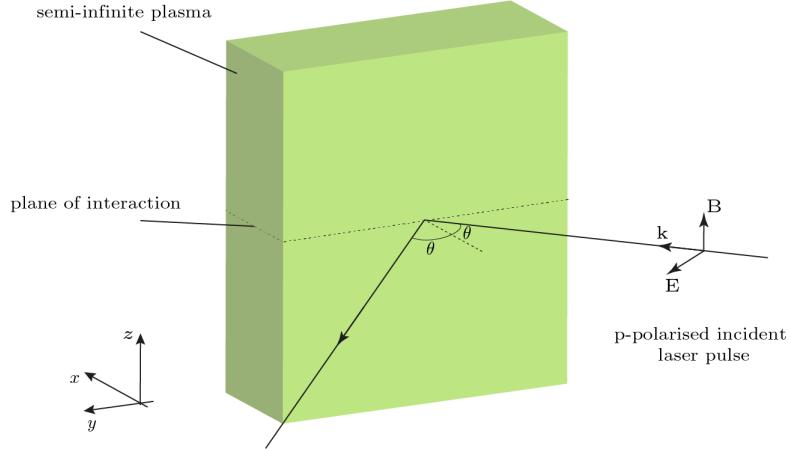
To understand this system, consider now a relativistic linearly polarised laser pulse obliquely incident, at an angle of incidence  $\theta$ , on a semi-infinite plasma, existing for  $x \geq 0$  as in Figure 2.1.

The Hamiltonian of a single electron confined within the potential well [80] is

$$\mathcal{H} = c\sqrt{m_e^2c^2 + |\mathbf{p}|^2} - e\Phi. \quad (2.1)$$

Here, the second term of Equation 2.1 describes the contribution to the electron's energy from the electrostatic potential,  $\Phi$ , of the pseudo-capacitor. The first term is the electron energy,  $U$ , extracted from the invariant of the relativistic 4-momentum of the electron,  $\mathbf{P}^\mu = (U/c, \mathbf{p})$ ,

$$\mathbf{P}_\mu \mathbf{P}^\mu = \frac{U^2}{c^2} - |\mathbf{p}|^2 = m_e^2c^2. \quad (2.2)$$



**Figure 2.1:** Diagram of a p-polarised laser pulse incident at angle  $\theta$  specularly reflected from a solid density plasma. By considering the Lorentz force equation, it is clear that all forces and therefore all plasma particle dynamics are confined to a plane.

Decomposing the electron's 3-momentum into orthogonal components:  $p_{\text{prop}}$ , along the laser propagation direction,  $p_{\text{pol}}$ , along the polarisation axis of the laser pulse and  $p_{\perp}$ , perpendicular to both, there are two simplifications to be made. Firstly, by canonical conservation of transverse momentum,  $p_{\text{pol}} = eA$ , where  $A$  is the laser vector potential amplitude. Secondly, in the case of a p-polarised laser pulse (the known optimum for ZVP electron bunch generation [117] and HHG [125]), with reference to Figure 2.1 and the Lorentz force law, the forces at play confine the electron trajectory to the  $p_{\text{prop}}-p_{\text{pol}}$  plane and the essential interaction geometry is two-dimensional. This is provided one considers length scales smaller than the focal spot of the laser pulse on the target, such that variation of the ponderomotive pressure with the third dimension can be neglected.

Explicitly, the Hamiltonian can be written as

$$\mathcal{H} = c\sqrt{m_e^2c^2 + p_{\text{prop}}^2 + e^2A^2} - e\Phi. \quad (2.3)$$

Equation 2.3 describes a potential well that constrains the electrons, on one side longitudinal space charge— $e\Phi$  term prevents further propagation into the target while on the other term associated with the vector potential of the laser pulse,  $e^2A^2$ . It is clear that should the vector potential pass through zero<sup>1</sup>, one of potential

---

<sup>1</sup>Note that in writing  $p_{\text{pol}} = eA$  one has determined the gauge in which the zero becomes a defined quantity

wall walls is totally suppressed, enabling electrons in the skin layer to escape the plasma, breaking adiabaticity. The necessity of vector potential zeros for this violent reconstruction of the plasma surface led Baeva *et al* [118] to coin the term ‘Zero Vector Potential’ mechanism to describe this process. Indeed, while elementary electromagnetism tells us a laser pulse will exponentially decay within a skin layer of a plasma without passing through zero, Baeva *et al* [118] demonstrated in PIC simulations that in this non-linear regime, zeros do exist and do propagate through the skin layer. The explanation relies on a Doppler shift in the laser field due to the relativistic motion of the ablating plasma surface, and a mathematical formalism of this process proceeds as follows.

As the ZVP mechanism is a relativistic phenomenon, it is absolutely essential to perform a relativistic analysis. Since all accelerated electrons travel at approximately speed  $c$ , surface electrons undergo similar trajectories. Acting collectively they oscillate in the laser pulse field. Consider first a transformation to the frame of reference where the laser pulse is normally incident to the plasma surface, this frame travels at velocity  $\mathbf{v} = (c \sin \theta) \hat{\mathbf{y}}$  with electrons streaming at  $-\mathbf{v}$ . Using Equation 2.2,  $U = \gamma m_e c^2$  and integrating Equation 1.50 in the boosted frame noting  $p_T(t = 0) = \gamma_{\mathbf{v}} m_e v = m_e c \sin \theta \cos \theta$ ,

$$\gamma^2 = 1 + (a_0 + \sin \theta \cos \theta)^2 + \left( \frac{p_{\text{prop}}}{m_e c} \right)^2, \quad (2.4)$$

where all parameters are in the boosted frame. Using  $\mathbf{p} = \gamma m_e \mathbf{v}$ , the longitudinal velocity is

$$v_{\text{prop}} = \frac{\tilde{p}_{\text{prop}} c}{\sqrt{1 + (a_0 + \sin \theta \cos \theta)^2 + \tilde{p}_{\text{prop}}^2}}, \quad (2.5)$$

where  $\tilde{p}_{\text{prop}} = p_{\text{prop}}/m_e c$ . Thus, should the transverse vector potential pass through  $-\sin \theta \cos \theta$ , zero for normal incidence, the surface can propagate towards the laser pulse at very close to speed  $c$ . Transforming back to the laboratory frame, at the peak of ablation ( $\mathbf{u} \approx -c \hat{\mathbf{x}}$ ) and using the equations for relativistic velocity addition,

$$\mathbf{u}'_{\parallel} = \frac{\mathbf{u} - \mathbf{v}}{1 - \mathbf{u} \cdot \mathbf{v}/c^2}, \quad (2.6)$$

$$\mathbf{u}'_{\perp} = \frac{\mathbf{u}_{\perp}}{\gamma_v(1 - \mathbf{u} \cdot \mathbf{v}/c^2)}, \quad (2.7)$$

where  $\gamma_v = 1/\sqrt{1 - |\mathbf{v}|^2/c^2}$  [74], one finds that this peak ablation at speed  $\approx c$  occurs now in the specular reflection direction. Simultaneity is broken and ripples co-move along the surface with the incident laser pulse wavefronts.

Transform now to the rest frame of the ablating front. Beyond the relativistic critical density surface, the vector potential of the laser pulse decays evanescently. At the spatial centre of the laser pulse, it can be described simply by

$$\mathbf{A}'_{\text{L}}(t', r') = A'_0 \cos(\omega'_{\text{L}} t') \exp(-r'/\delta') \hat{\mathbf{r}}'_{\text{pol}} = A'_0 \hat{\mathbf{r}}'_{\text{pol}}, \quad (2.8)$$

where the primed symbols indicate that these quantities are measured in the rest frame of the expanding front.  $A'_0$  is the vector potential amplitude and  $\omega'_{\text{L}}$  is the frequency of the laser pulse,  $r'$  is the propagation distance of the laser into the plasma,  $\delta'$  is the skin depth and  $\hat{\mathbf{r}}'_{\text{pol}}$  a unit vector defining the polarisation direction of the laser pulse. Un-primed coordinates will indicate the lab frame measurements.

While previous demonstrations of the existence of vector potential zeros assumed that the ablation occurs normal to plasma surface, it is necessary to confirm that zeros are still predicted for specular ablation. Consider a p-polarised laser pulse confined to the  $x$ - $y$  plane incident with an angle of incidence  $\theta$  on an ablating overdense plasma expanding with velocity  $-v_f \hat{\mathbf{x}}$  in the lab frame, as in Figure 2.2.

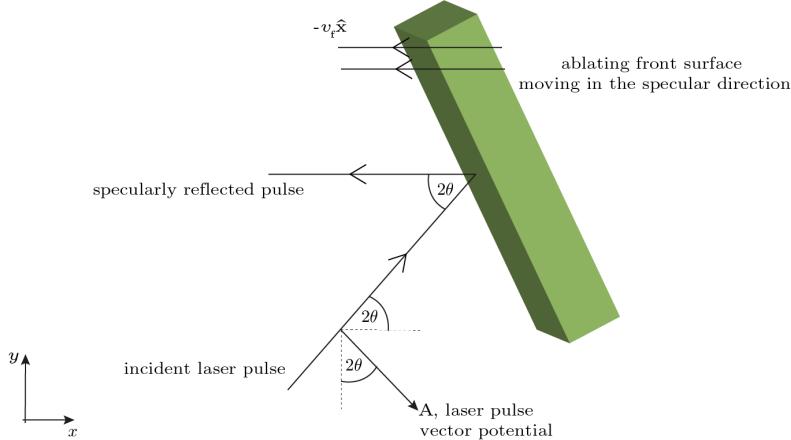
The direction of polarisation is

$$\hat{\mathbf{r}}_{\text{pol}} = \hat{\mathbf{x}} \sin 2\theta - \hat{\mathbf{y}} \cos 2\theta \quad (2.9)$$

and the velocity of the rest frame of the ablating front relative to the lab frame is  $-v_f \hat{\mathbf{x}}$ .

Applying the Lorentz transformation to the electromagnetic 4-potential,  $\mathbf{A}^{\mu}$ , where  $\Lambda_{\mu}^{\nu}$  is given by Equation 1.30, immediately from the  $y$ -coordinate transformation,

$$A'_{\text{L}} \cos 2\theta' = A_{\text{L}} \cos 2\theta. \quad (2.10)$$



**Figure 2.2:** Diagram of a  $p$ -polarised laser pulse incident on an ablating overdense plasma. The laser is incident obliquely at an angle of  $\theta$  and is reflected specularly. The plasma ablates specularly also. Note that this process occurs on sub-laser pulse cycle timescales and it is therefore only the electrons that are ablating in this process. The interaction geometry is confined to a 2D plane.

Applying the headlight effect for a source moving at an angle  $2\theta$  to the boosted frame (a full derivation is given in Appendix A.2),

$$\cos(2\theta') = \frac{\cos(2\theta) - \beta}{1 - \beta \cos(2\theta)} \quad (2.11)$$

and rearranging Equation 2.10, the vector potential in the lab frame is

$$A_L = \frac{1 - \beta \sec(2\theta)}{1 - \beta \cos(2\theta)} A'_0 \cos(\omega'_L t') \exp(-r'/\delta'). \quad (2.12)$$

Writing the boosted frame space-time coordinates in terms of the lab frame coordinates,

$$ct' = \gamma(ct - \beta x), \quad (2.13)$$

$$x' = \gamma(x - \beta ct), \quad (2.14)$$

yields

$$A_L = A_0 \cos(\omega_L t - kx) \exp\left(-\frac{\sqrt{(x - \beta ct)^2 + (y/\gamma)^2}}{\delta}\right), \quad (2.15)$$

where

$$A_0 = \frac{1 - \beta \sec(2\theta)}{1 - \beta \cos(2\theta)} A'_0, \quad (2.16)$$

$$\omega_L = \gamma \omega'_L, \quad (2.17)$$

$$k = \frac{\beta \gamma \omega'_L}{c}, \quad (2.18)$$

$$\delta = \frac{\delta'}{\gamma}. \quad (2.19)$$

The oscillatory term in Equation 2.15 demonstrates the propagation of vector potential zeros within the plasma target. From the structure of this term it would appear that these zeros are expelled from the plasma along the specular direction at a speed

$$v_\phi = \frac{\omega_L}{k} = \frac{c}{\beta} = -\frac{c^2}{v_f}. \quad (2.20)$$

While in their original ZVP paper Baeva laid some doubt on their version of this calculation, instead suggesting similarity theory predicts zeroes propagate at speed  $c$  [118], the theory of High Harmonic Generation rejects this alternative. It is well known that the emission of reflected radiation occurs primarily at the point where the transverse vector potential passes through zero [116] but equally that the width of the radiated pulse depends on the advance time emission point [126]. If the zeroes moved at speed  $c$ , then the observed emitted pulse would be infinitely thin, producing radiation with perfect coherence in all cases. This is unphysical and simulations have suggested finite advanced time bunch widths decreasing rapidly with increasing laser intensity [126]. This is precisely what is predicted by Equation 2.20. As the laser intensity increases,  $v_f$  naturally increases. Thus, the zeroes propagate closer to the speed of light and reducing the advanced time bunch width.

To summarise, for a sufficiently intense laser pulse, electrons at the surface of an irradiated solid target are accelerated by the laser to relativistic velocities at a fraction of a laser pulse cycle and therefore electrons both follow similar trajectories and respond adiabatically to the  $\mathbf{J} \times \mathbf{B}$  force of the laser pulse. They form into a high charge density spatially thin coherent electron bunch on the front surface of the plasma but displaced inwards from the approximately immobile ions via the ponderomotive pressure of the laser. This charge separation generates a longitudinal electrostatic pseudocapacitor field that confines electrons to a potential well on the front surface of the plasma, preventing further propagation of the electron bunch into the plasma bulk. When the zero of the vector potential passes through the electron bunch, the ponderomotive pressure instantaneously

vanishes and electrons are ejected specularly from the target, co-propagating with the zeroes of the vector potential.

In the work by Cousens *et al* on electron trajectories in this regime, it is noted that a Coherent Synchrotron Emission (CSE) radiation burst occurs when the transverse momentum goes to zero [cousens2020], corresponding to the passage of the zero of the transverse vector potential. CSE and ZVP are therefore intrinsically linked, however, while CSE focuses on reflection and the HHG spectrum, ZVP is concerned with laser pulse energy absorption.

After expulsion from the plasma, the pseudocapacitor is discharged as the electron bunch accelerates across it. Upon encountering the subsequent laser pulse peak amplitude, the bunch is then rotated back towards the plasma and launched into the bulk at high energy along the laser propagation axis (by conservation of transverse momentum in the plasma bulk), as it does so emitting coherent synchrotron radiation in transmission.

### 2.2.1 ZVP electron bunch energies

In [118], Baeva *et al* propose energy scalings in one-dimension (1D) for an electron bunch produced in the ZVP regime as a function of the incident laser pulse intensity and plasma density. This was extended to three-dimensions (3D) by Savin *et al* [117]. What follows is that discussion with closer consideration of both the constants of proportionality and their consequences. Note that throughout the electron bunch is treated as infinitessimally thin, as proved to be a reasonable assumption in previous work on ZVP [117, 125, 127]. Sub-bunch dynamics have been explored in more detail by Gonoskov *et al* [128].

The expressions for energies in equations 2.26 and 2.29 require the electron bunch to fully discharge the pseudo-capacitor before interaction with the subsequent laser pulse peak. Since the electron bunch travels at speed  $\approx c$ , the peak displacement (and thus the pseudo-capacitor width) must satisfy

$$\Delta x \leq \frac{\lambda}{8}. \quad (2.21)$$

Using Equation 2.24, it is clear Equation 2.21 is satisfied for  $S \geq 1.3$  at normal incidence.

Consider again the semi-infinite block of plasma presented in Figure 1.2, normally irradiated by a laser pulse with wavelength  $\lambda_L$  and peak electric field,  $E_L$ . It is now the ponderomotive pressure of the laser that displaces the electron fluid in this picture. The electron surface moves inwards until the pressure exerted by the peak instantaneous ponderomotive pressure of the laser pulse cycle,

$$\mathbf{P}_L = \epsilon_0 E_L^2 \hat{\mathbf{x}} = \epsilon_0 \left( \frac{a_0 \omega_L m_e c}{e} \right)^2 \hat{\mathbf{x}} \quad (2.22)$$

is equal and opposite to the pressure exerted by the pseudo-capacitor field,

$$\mathbf{P}_C = \frac{QE_C}{\sigma} \hat{\mathbf{x}} = -\frac{(en_e \Delta x)^2}{\epsilon_0} \hat{\mathbf{x}} \quad (2.23)$$

from equations 1.16 and 1.17. Equating the magnitudes of  $\mathbf{P}_L$  and  $\mathbf{P}_C$ , the maximum displacement inwards of electrons is

$$\Delta x \hat{\mathbf{x}} = \frac{c}{\omega_L} \frac{a_0}{\bar{n}_e} \hat{\mathbf{x}} = \frac{1}{kS} \hat{\mathbf{x}}, \quad (2.24)$$

where  $k$  is the wave-vector of the laser pulse. Correspondingly,

$$E_C = \frac{en_e}{\epsilon_0} \Delta x = \frac{\omega_L c m_e a_0}{e} = E_L. \quad (2.25)$$

Applying the results of equations 2.24 and 2.25, when the ponderomotive pressure vanishes and the electron bunch is launched across the pseudo-capacitor, the relativistic kinetic energy gained by a single electron is

$$T = \int \mathbf{F} \cdot d\mathbf{s} = \int_{\Delta x}^0 -eE_C dx = \int_{\Delta x}^0 -\frac{en_e x}{\epsilon_0} dx = \frac{1}{2} m_e c^2 \frac{a_0^2}{\bar{n}_e} \quad (2.26)$$

or an electron gamma factor,

$$\gamma = \frac{1}{\sqrt{1 - \beta^2}} = 1 + \frac{a_0^2}{2\bar{n}_e}. \quad (2.27)$$

Assuming all displaced electrons are captured by the potential well and launched as a coherent bunch, the total number of electrons in the bunch is

$$N_e = n_e \sigma \Delta x = \frac{\sigma a_0 n_c}{k} = \sigma \epsilon_0 E_L \quad (2.28)$$

and hence, the total kinetic energy of the electron bunch is

$$U_{\text{ZVP}} = N_e T = \frac{\sigma n_c}{k} \times \frac{1}{2} m_e c^2 \frac{a_0^3}{\bar{n}_e}. \quad (2.29)$$

It is now interesting to compare Equation 2.29 to the laser energy deposited upon the plasma surface and therefore consider what fraction of the laser energy can be absorbed via the ZVP mechanism. Using  $E_C = E_L$ , Equation 2.29 can be rewritten as

$$U_{\text{ZVP}} = \frac{1}{2\omega_L S} \sigma c \epsilon_0 E_L^2. \quad (2.30)$$

For the case of normal incidence, bunches are produced at a frequency of  $2\omega_L$ , naturally following the frequency of the  $\mathbf{J} \times \mathbf{B}$  force. Assuming a sinusoidal plane wave incident with surface area  $\sigma$ , the energy available during the pushing phase (a quarter-cycle) is

$$U_{L,1/4} = \sigma \frac{T}{4} \langle I_L \rangle = \frac{2\pi}{8\omega_L} \sigma c \epsilon_0 E_L^2. \quad (2.31)$$

Hence,

$$\eta_{\text{ZVP}} = \frac{U_{\text{ZVP}}}{U_{L,1/4}} = \frac{2}{\pi S}. \quad (2.32)$$

Interestingly, this analytical result predicts the trend observed by A. Savin [119] in PIC simulations both in magnitude and in scaling. Indeed, A. Savin demonstrated in numerical simulation

$$\eta_{\text{ZVP}} \sim S^{-1.000(3)}, \quad (2.33)$$

however, this result led A. Savin to conclude that increasing  $S$  reduces absorption, increasing the energy in the reflected HHG beam thus increasing high harmonic efficiency, seemingly in tension with the results of other works [120, 126]. The resolution arises from awareness of two distinct conversion efficiencies that describe the reflected harmonic spectrum: the total conversion efficiency into the full reflected beam and the conversion efficiency for individual harmonics. While the overall conversion into the reflected beam decreases for decreasing  $S$ , the slope of the harmonic spectrum becomes shallower and HHG efficiency increases. Indeed, high

harmonic efficiency necessitates high reflection inefficiencies due to the production of ZVP electron bunches as higher energy bunches produce more coherent reflected radiation [126]. Harmonic efficiencies will be discussed in more detail in the following Chapters, however, at currently accessible laser intensities, it would appear that X-ray harmonics are produced with greater efficiency for high  $S$  [129], a regime generally neglected in parameter scans given the computational and experimental challenges of accessing it.

Unfortunately, it is not possible to link the energy scaling derived in this Section to CSE since the harmonic emission occurs before the ZVP acceleration phase. In Dromey *et al*, the CSE regime was demonstrated for the first time in transmission through thin foils, however laser pulses are now of sufficient intensity to access CSE in reflection from solid targets. This has been identified as a more efficient approach if it can be reached due to the production of higher density and shorter duration electron bunches and therefore brighter and more coherent CSE [130].

### 2.2.2 ZVP bunches oblique incidence scaling

This section is inspired by ideas from the work of Gonoskov *et al* [120] and Vincenti *et al* [122] to extend the theory of the ZVP mechanism for energy absorption to the more practical<sup>2</sup> case of oblique incidence.

Provided the plasma-vacuum boundary is sufficiently steep, the plasma electrons will respond adiabatically to the laser pulse and arrange themselves to form a pseudocapacitor longitudinal electric field  $E_C$  at the plasma surface. At all points in this adiabatic ‘pushing’ phase, the surface electrons will be in a quasi-static equilibrium *i.e.* there will be a balance between the electromagnetic forces on them. Consider again the laser pulse incident on a solid density plasma existing for  $x > 0$  at angle  $\theta$ . Transforming to the frame of reference in which the laser is normally incident (quantities in this frame are indicated by the primed symbol), the electron and ion bulk plasma species stream at velocity  $\mathbf{v}_d = -c \sin \theta \hat{\mathbf{y}}$ . Applying the Lorentz force law along the longitudinal direction ( $\hat{\mathbf{x}}$ ), for a displacement of the electron fluid

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<sup>2</sup>Not only is this more feasible in experiment but has been shown to optimise HHG [120].

$x'_e$  (one assumes that the expression for a single electron at the surface describes the surface since all electrons follow similar trajectories), travelling at speed  $\mathbf{v}'$ ,

$$-e(\mathbf{v}'(x'_e) \times (\mathbf{B}'_L(x'_e) + \mathbf{B}'_i(x'_e)) \cdot \hat{\mathbf{x}} + E'_C(x'_e)) = 0, \quad (2.34)$$

where the laser magnetic field,

$$B'_L = \frac{m_e \omega'_L a_0 \sin(\omega'_L t' - k' x'_e)}{e} \hat{\mathbf{z}} \quad (2.35)$$

and  $B_i$  is the magnetic field generated by the uncompensated ion current,  $\mathbf{J}_i = Z n'_i(x'_e) \mathbf{v}_d$ , where the electron fluid has been displaced. As before, from Equation 1.17,

$$E'_C = \frac{en'_e x'_e}{\epsilon_0}. \quad (2.36)$$

Applying Maxwell-Ampère's Law, Equation 1.1c, and noting that by symmetry there can be no variation in the magnetic field with  $y'$  or  $z'$ , it becomes clear that

$$-\frac{d(\mathbf{B}'_i)_{z'}}{dx'} = \mu_0 (\mathbf{J}_i)_{y'}. \quad (2.37)$$

Integrating Equation 2.37 from  $-\infty$  to  $x'_e$ , noting that  $\mathbf{B}_i = 0$  at infinity and assuming a constant density profile  $n'_i$  for  $x > 0$  with  $Z n_i = n_e$ ,

$$\mathbf{B}'_i(x'_e) = \mu_0 en'_e x'_e c \sin(\theta) \hat{\mathbf{z}}. \quad (2.38)$$

Using equations 2.35, 2.36 and 2.38 and making the reasonable approximation that the relativistic electrons on the surface move at speed  $v'_y \approx \pm c$  at peak displacement ( $x'_e = x'_p$ ), 2.34 can be written as

$$-e \left( \pm c \left( \pm \frac{m_e \omega'_L a_0}{e} + \mu_0 en'_e x'_p c \sin \theta \right) + \frac{en'_e x'_p}{\epsilon_0} \right) = 0. \quad (2.39)$$

For the laser to be in the pushing phase, the first term must be negative, corresponding to  $\mathbf{v}'$  and  $\mathbf{B}'_L$  having the opposite sign, hence,

$$c \left( -\frac{m_e \omega'_L a_0}{e} \pm \mu_0 en'_e x'_p c \sin \theta \right) + \frac{en'_e x'_p}{\epsilon_0} = 0, \quad (2.40)$$

where here the  $\pm$  tracks the sign of  $\mathbf{v}'$ . After some manipulation, one arrives at

$$x'_p = \frac{1}{k' S' (1 \pm \sin \theta)}. \quad (2.41)$$

Transforming back to the lab frame,

$$x_p = \frac{\cos^2 \theta}{kS(1 \pm \sin \theta)}. \quad (2.42)$$

Convincingly, this reduces to Equation 2.24 for  $\theta = 0$  and predicts the suppression and enhancement of the two surface oscillations per laser pulse cycle. Explicitly, for a laser pulse propagating at  $y = x \tan \theta$ , the peak displacement of the electron surface is enhanced for  $\mathbf{B}_L$  in the  $+\hat{\mathbf{z}}$ -direction and suppressed for  $\mathbf{B}_L$  in the  $-\hat{\mathbf{z}}$ -direction.

Again, integrating to find the work done as the electron bunch accelerates across the pseudocapacitor,

$$T(\theta) = \int \mathbf{F} \cdot d\mathbf{s} = \frac{1}{2} m_e c^2 \frac{a_0^2}{\bar{n}_e} \frac{\cos^4 \theta}{(1 \pm \sin \theta)^2}. \quad (2.43)$$

While it was to be expected that a component of the electric field acting into or out of the plasma would change the magnitude of the displacement, this result suggests that increasing the angle of incidence can increase the electron energy gain in the enhanced cycle more than the decrease in the suppressed cycle. Note that this is to be expected as it is known that HHG efficiency is improved for non-zero angle of incidence [120, 126]. It would be interesting to explore whether the presence of an external magnetic field could be applied to mimic the effect of oblique incidence by replacing the magnetic field due to the uncompensated ion current.

While this model would suggest an optimal angle for electron energy and therefore of HHG of  $\pi/2$ , if  $\theta > \pi/4$ , then, if the relativistic electron bunch is travelling at  $c$  along the specular reflection direction, the subsequent laser peak amplitude will never ‘catch up’ with the electron bunch, and electrons can escape, generating high charge electron bunches in reflection as observed in experiment by Lin *et al* [113], but reducing the efficiency of HHG.

Finally, the total bunch energy as a function of  $\theta$ ,

$$U_{ZVP}(\theta) = \frac{\sigma n_c}{k} \times \frac{1}{2} m_e c^2 \frac{a_0^3}{\bar{n}_e} \frac{\cos^6 \theta}{(1 \pm \sin \theta)^3}. \quad (2.44)$$

As anticipated, oblique incidence can increase laser absorption efficiency for higher  $S$  plasmas which are currently more easily accessed compared to the more optimal low  $S$  plasmas.

### 2.2.3 Defining characteristics of the ZVP mechanism

In her original paper on the ZVP mechanism, T. Baeva *et al* [118] outlined 6 defining characteristics of the ZVP mechanism, namely,

1. The existence of vector potential zeros moving through the skin layer in the laboratory frame;
2. The existence of zeroes in the incident laser pulse vector potential required for the formation of fast electron bunches;
3. The generation of fast electron bunches with ultra-short temporal duration;
4. That such fast electron bunches follow the energy scalings of equations 2.26 and total energy 2.29;
5. Injection of the fast electron bunches is along the propagation axis of the laser pulse;
6. There must be an intrinsic link between the fast electron bunches and coherent HHG;

with the moving zeros within the skin layer being the key delineator between this post-ponderomotive regime of laser pulse energy absorption and all other proposed mechanisms. While such observational requirements are far beyond the reaches of current experimental know-how, numerical simulations in both 1- [118] and 2-dimensions [117] have confirmed the above points. Now is presented the first 3D simulations attempting to demonstrate these criteria.

## 2.3 Numerical simulations of the ZVP mechanism

This thesis relies on the analysis of 1,2 and 3D PIC simulations, primarily using the massively-parallel and open-source simulation code Smilei [75]. Simulation parameters are provided throughout.

### 2.3.1 The ZVP mechanism in 3D3V

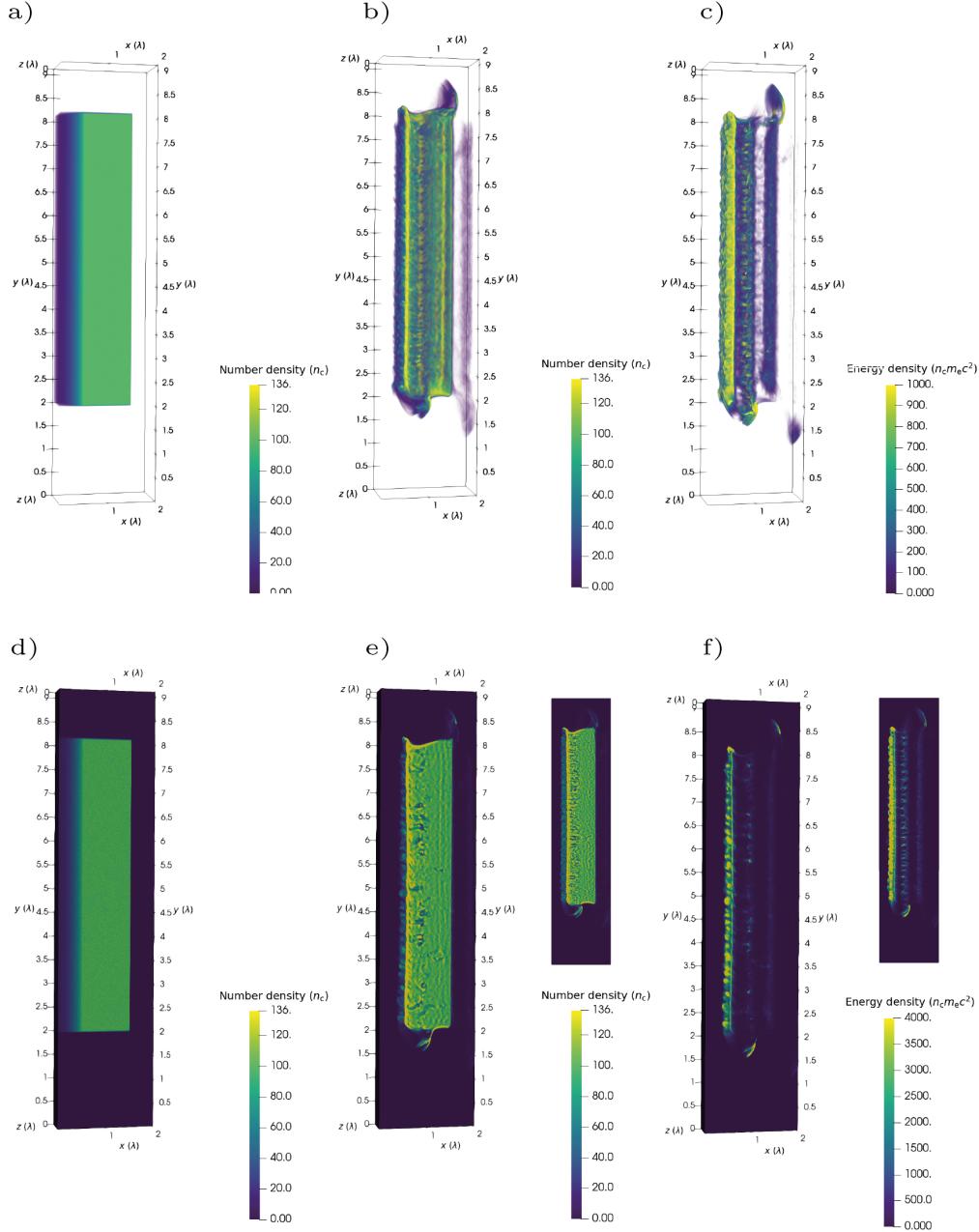
3D simulation results are presented in Figure 2.3 alongside comparison to an equivalent 2D simulation. Simulation parameters are given in Table 2.1, such parameters are compatible with the 10 PW ELI-NP state-of-the-art short pulse laser facility [131] incident on foam targets. Figure 2.3c) clearly demonstrates the existence of high energy density electron bunches propagating through the plasma bulk in the direction of the laser pulse. Note that this ZVP criterion is a direct consequence of conservation of transverse momentum inside the plasma bulk where the laser fields cannot propagate. Figure 2.3b) shows these bunches escape to the rear of the bulk but lose energy as they do so. Looking now at Figure 2.3e) and the internal structure of the plasma bulk. These bunches drive two-stream and filamentation instabilities [132]. The bulk propagating electron bunches are accompanied by higher density electron bunches to either side of the plasma block with the side of emission alternating every half laser pulse cycle.

#### Considerations and convergence of 3D simulations

The plasma specifications were chosen to minimise computational load while ensuring numerical convergence, requiring over 100 billion macroparticles. The electron temperature is raised significantly higher than that which would be expected in such a laser-plasma system so as to resolve the Debye length. Anticipated plasma temperatures are calculated using 1D HYADES simulations in the following Chapter. While this temperature is unphysical and will lead to some small plasma expansion over the course of the simulation (negligible on these timescales), the temperature remains small compared to that imparted to the electron bunches by the laser pulse. The striking similarity between the 2 and 3D simulation results is a natural consequence of the 2D nature of the interaction geometry. It is, however, still reassuring to show that previous work withstands the test of reality's geometry and the simple elegance of the ZVP mechanism is not lost in the chaos.

For the 3D simulations only, particles were initialised randomly not regularly to avoid numerical error. Regularly initialised plasma blocks in 3D simulations

### 2.3. Numerical simulations of the ZVP mechanism



**Figure 2.3:** Simulation results from a 3D PIC simulation of the ZVP mechanism. a) The initialised electron number density. b) The electron number density several cycles later, the plasma bulk is intact, however there is evidence of instabilities and electron bunches propagating through and around the plasma. c) The electron kinetic energy density at the same timestep. Note that the scale has been clipped to enable observation of both electron bunches propagating through and around the plasma bulk. Significantly higher energy density, corresponding to a higher charge density and attosecond duration is observed for the electron bunches propagating around the bulk. d-f) Plots clipped through  $z = \lambda_L/2$  for a-c) respectively for to access the internal structure of the plasma bulk. The accompanying plots for figures e) and f) are equivalent 2D PIC simulations demonstrating excellent agreement.

Laser (3D, normal incidence)		
Parameters	Real	Smilei
Wavelength, $\lambda$ (nm)	1060	$2\pi$
Angular frequency, $\omega_L$ (fs $^{-1}$ )	1.8	1
Beam waist, $w_L$ (nm)	$6\lambda$	$12\pi$
Focal point, $(f_x, f_y, f_z)$ (nm)	$(0.5\lambda, 5\lambda, 0.5\lambda)$	$(\pi, 10\pi, \pi)$
Spatial envelope, $E_i$ , $i = y, z$	$E_i \sim e^{-(i-i_f)^2/w_L^2}$	
Temporal envelope, $E_t$	$E_t \sim e^{-(t-4\lambda/c)^2/((4\lambda/3c)^2 \ln 2)}$	
Simulation box		
Size, $x \times y \times z$ (nm)	$2\lambda \times 9\lambda \times \lambda$	$4\pi \times 18\pi \times 2\pi$
Sim length (fs)	35.22	$20\pi$
Spatial resolution, $\Delta x$ (nm)	$\lambda/128 = 8.28$	0.0491
Temporal resolution, $\Delta t$ (as)	$\Delta x/11c = 2.51$	0.00446
Collisionless, pre-ionised randomly-initialised aluminium plasma		
Electron $x$ profile, $n(x)$	$\begin{cases} n_e & \text{for } 2\lambda \leq x \leq 3\lambda, \\ n_e e^{(x-2\lambda)/0.2\lambda} & \text{for } x \leq 2\lambda. \end{cases}$	
Electron $y$ profile, $n(y)$	$\begin{cases} 1 & \text{for } 2\lambda \leq y \leq 8\lambda, \\ 0 & \text{otherwise.} \end{cases}$	
Electron $z$ profile, $n(z)$	$\begin{cases} 1 & \text{for } 0.125\lambda \leq z \leq 0.875\lambda, \\ 0 & \text{otherwise.} \end{cases}$	
Ion profile, $n_i(x, y, z)$	$n_i = n(x)n(y)n(z)/13$	
Macro-electrons per cell	729	
Macro-ions per cell	8	
Ion temperature, $T_i$ (keV)	0	0
Electron temperature, $T_e$ (keV)	10	0.02
Stability criteria		
$\lambda_D/\Delta x$	0.288	
$1/\Delta t \omega_p$	24.4	
$\Delta x/c\Delta t$	11	
Macro-particles in the Debye sphere	210	

**Table 2.1:** Simulation parameters in both real and normalised Smilei simulation units for the 3D3V simulations.

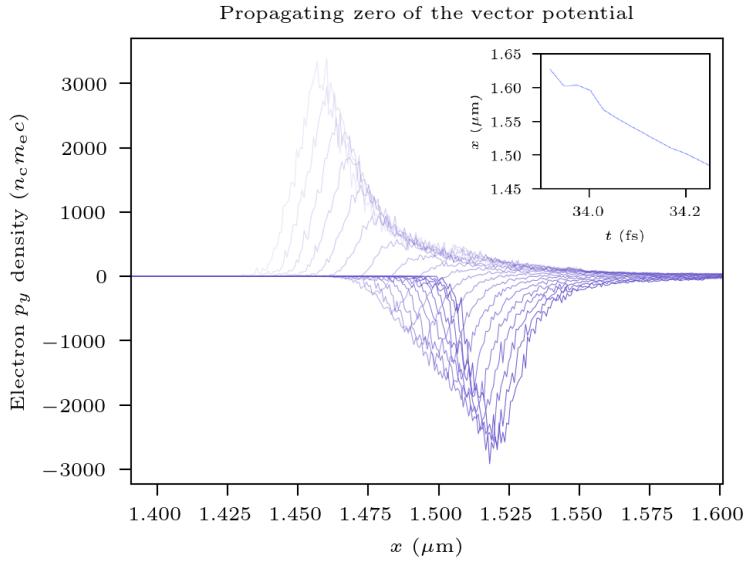
rapidly blow apart due to spurious large amplitude fields generated on the large plasma surfaces due to macroscopic electron-ion charge separation at initialisation. 2D simulations with randomly initialised particles tend to produce nano-structures reminiscent of the cosmic web. The absence of these structures in 3D is telling. Unsurprisingly this error can be reduced by increasing the plasma temperature.

The longitudinal thickness of the target does not impact the interaction and is chosen for computational efficiency, indeed it is standard to consider such targets of thickness  $\geq \lambda_L$  as bulk targets [71], however, for sufficiently long pulse durations, the effect of hole boring necessitates thicker targets to make this approximation as is necessary in the following Chapter.

The initial 3D simulation parameters were chosen to be consistent with previous work on the ZVP mechanism, however, such simulations are cumbersome. In order to query the defining characteristics outlined by Baeva, a lower resolution simulation was performed with similar parameters to the initial simulation. Comparisons between the simulation outputs are made in Figure B.1 in Appendix B.1. Good convergence is qualitatively demonstrated by the presence of characteristic features of the ZVP mechanism. While the instabilities are similar in structure, the change in seeding changes their exact positions. As instabilities are not the focus of this thesis this variation is of no cause for concern.

## Confirmation of ZVP in 3D

Figure 2.4 tracks the transverse momentum distribution along the polarisation axis of the laser pulse of an electron bunch during its ablative journey, clearly demonstrating the existence of a singular zero of the vector potential propagating through the electron bunch and away from the plasma bulk. The zero propagates at a speed of  $\approx 1.4c$ , corresponding to a surface velocity of  $\approx 0.71c$ . In this simulation, the zero propagates through the bunch over a few nanometres and before crossing the pseudocapacitor. Note that consideration of the moving zeroes reframes the problem of the advanced time bunch width [126] (the determiner of attosecond

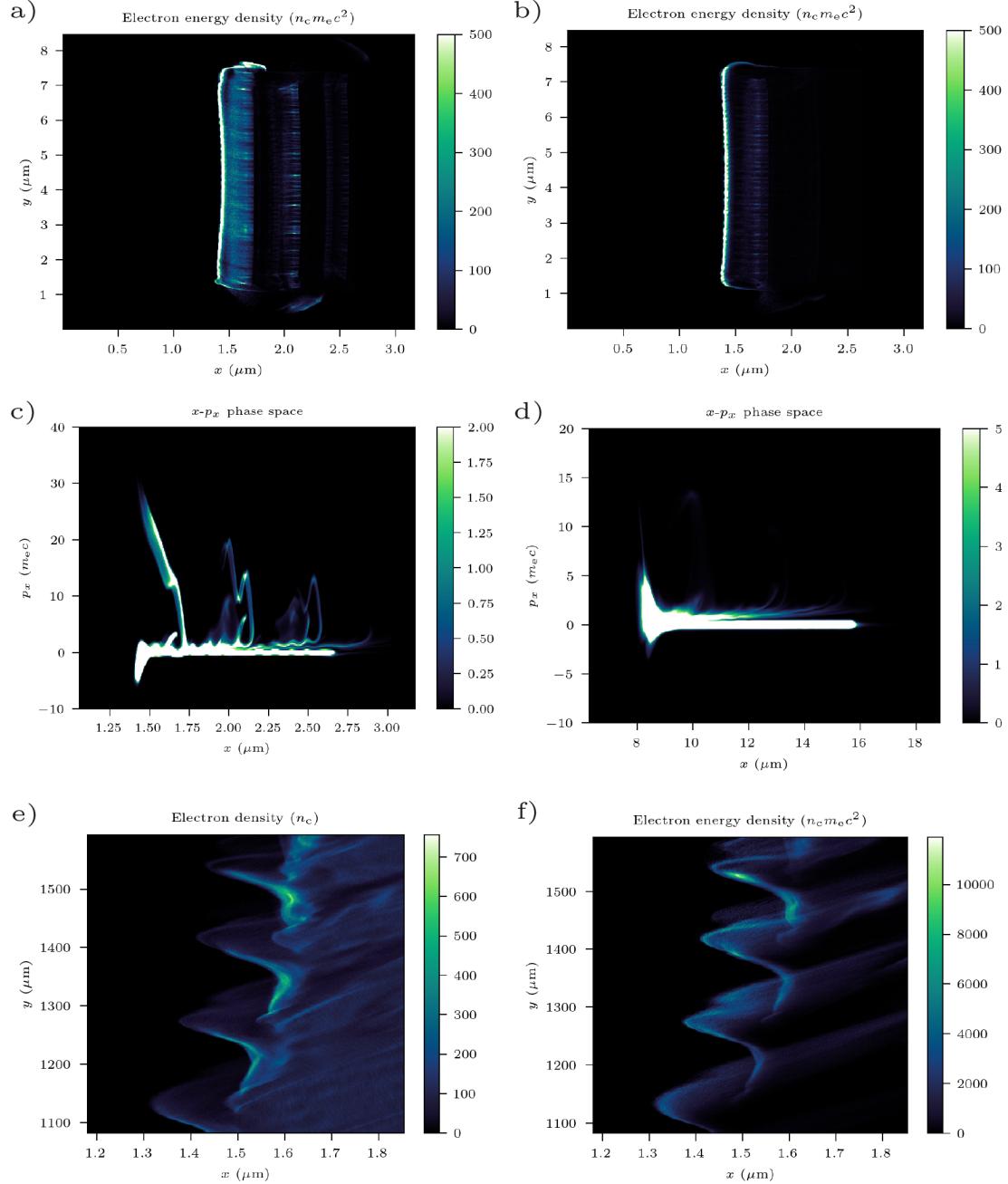


**Figure 2.4:** Propagation of a vector potential zero through the ablating ZVP electron bunch via the proxy of transverse momentum conservation. The zero of the vector potential exists where the transverse momentum is macroscopically zero. Each line represents a timestep with time increasing with decreasing colour. With increasing time the bunch and the zero move in the  $-\hat{\mathbf{x}}$ -direction with the zero overtaking the electron bunch. The inset tracks the position of the zero of the vector potential with time.

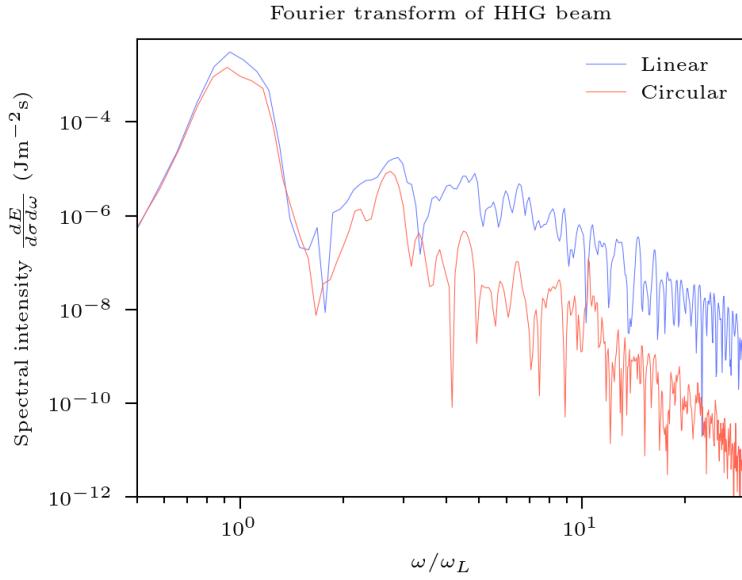
CSE pulse duration). Gonoskov *et al* have determined a scaling for electron bunch width [120] thus the only unknown is the surface velocity.

Further simulation results are presented in Figure 2.5. In Figures 2.5a)-d), comparisons are made in the electron energy distributions for simulations with zeros present in the vector potential (linearly polarised) and without (circularly polarised). Clearly zeroes are required for the formation of high energy, short duration attosecond bunches. From Figure 2.5a) it is clear that electron bunches are injected and propagate through the plasma bulk in the laser propagation direction. Figure 2.5c) demonstrates the quasi-monoenergeticity of the high energy electron bunches as initially identified in 1D by Baeva. Although the shape in the phase space is more complex in 3D, the attosecond duration at a given energy is retained. Figures 2.5e) and f) describe the surface dynamics in the linearly polarised case. One can observe the high density bunches on the front surface with the peak in energy density occurring after acceleration across the pseudocapacitor.

### 2.3. Numerical simulations of the ZVP mechanism



**Figure 2.5:** Electron dynamics in a 3D PIC simulation. a) and b) Electron energy density for linearly and circularly polarised laser pulses respectively. c) and d) Electron longitudinal momentum for linearly and circularly polarised laser pulses respectively. e) Electron density at the plasma surface streaked in time. f) Electron energy density at the plasma surface streaked in time.



**Figure 2.6:** The Fourier transform of the reflected laser pulse in 3D PIC simulations both with and without zeroes in the vector potential.

Finally, Figure 2.6 compares the spectra of the reflected light in the presence and absence of zeros of the vector potential. Individual harmonics are not resolved due to blueshifts between successive pulse cycles from considerable hole boring, a phenomenon discussed in great detail in the following Chapter. The spectrum from the circularly polarised light is typically over two orders of magnitude below that for linearly polarised light. Thus, all defining characteristics of the ZVP mechanism have been identified in 3D simulations with the notable exception of the energy scalings. Next generation supercomputers will be required to perform such parameter scans.

## 2.4 The ZVP electron bunch

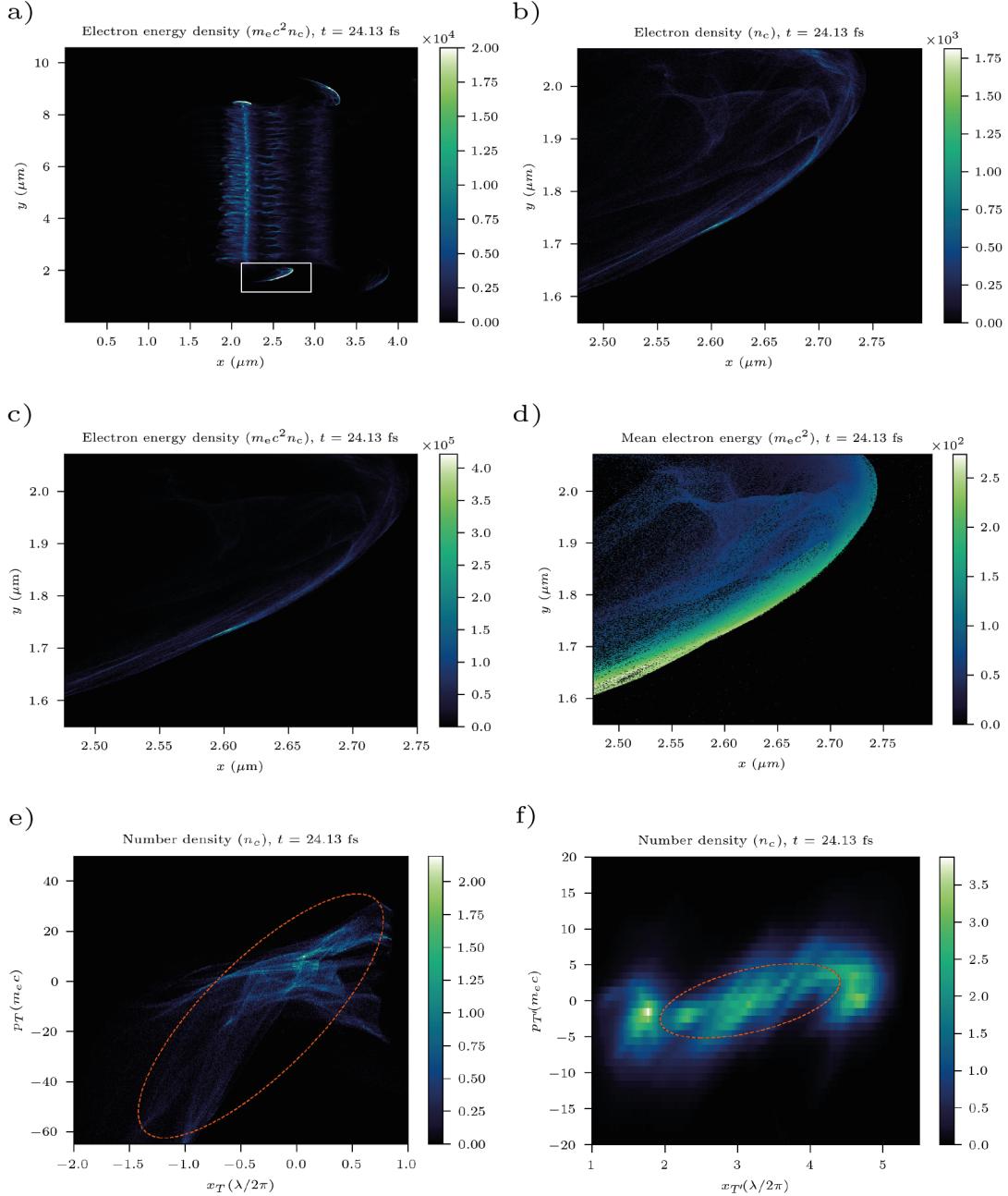
Following the excellent agreement demonstrated between 2 and 3D PIC simulations, the remainder of this Chapter explores ZVP primarily in 2D. Previous work on the ZVP mechanism has highlighted the high energy and short duration of electron bunches, what follows is a full characterisation of their properties. A ZVP electron bunch is an electron bunch produced via the ZVP mechanism. Once produced and accelerated across the pseudocapacitor field, it is launched back in the laser propagation direction. While the bunch has no spatial separation over energies

when propagating with the zero of the vector potential, the turning point of the electrons is longitudinal momentum dependent due to the Coulomb attraction of the ions after overshooting the pseudocapacitor field. Baeva *et al* showed that the electron bunch has a quasi-monoenergetic spectrum: the electron bunch is now of attosecond duration in the spectro-temporal domain and there is a one-to-one relationship between energy and position with the higher energies trailing the lower energies. The full bunch is confined to 130 as while a single energy confined to 5 as. If, however, the plasma bulk is transversely mass-limited relative to the laser spot size, when rotated back towards the plasma block, some of the electron bunch will overshoot and escape the potential well without significant chirping of the bunch as can be seen in Figure 2.3. Such electron bunches retain their high charge density and ultra-short duration with the tradeoff of increased divergence. Note that electron bunch properties are imprinted in the transmitted harmonics: high divergence but attosecond duration.

ZVP electron bunches can therefore be placed into two categories: ultra-high charge, ultra-short duration electron bunches from mass-limited targets, hereafter labelled mass-limited electron bunches, of interest due to their unique properties, and bulk propagating bunches, hereafter labelled bulk bunches, which have lower charge densities, are imprinted with instabilities and are instead of interest due to their connection to energy absorption and reflection in this post-ponderomotive regime. To investigate these two bunch types further, 2D PIC simulations were performed, see the Appendix for parameters.

### Attosecond nano-Coulomb mass-limited electron bunches

The plots of Figure 2.7 describe a typical mass-limited ZVP electron bunch qualitatively. The electron bunch under interrogation is ultra-relativistic with a mean energy of 51(11) MeV and a duration of 35 as. It propagates at an angle of -0.393 rad relative to the laser propagation direction, *i.e.* the  $x$ -axis, and has a transverse geometric emittance in the simulation plane (the  $x$ - $y$  plane) of 35(7) nm rad. The calculation of the transverse geometric emittance, a measure



**Figure 2.7:** 2D PIC simulation results qualitatively describing typical mass-limited ZVP electron bunch structure. a) Electron energy density for the full simulation window, corresponds to Figure 2.3f). The box highlights the bunch presented in the following plots. b) Electron number density of the electron bunch. c) Electron energy density of the electron bunch, the colourbar scale has been increased compared to Figure a) to observe the internal structure. d) The mean electron energy across the electron bunch, suggesting a position dependent energy or quasi-monoenergetic nature to the electron bunch [118]. Cells with no macroparticles are black. e) The transverse phase space in the 2D simulation plane. The ellipse describes the calculated emittance. The skew of the ellipse is a consequence of a low density tail on the phase space beyond the bottom left corner. f) This plot was extracted from the equivalent 3D simulation and describes the transverse emittance in the  $z$ -direction. Again the ellipse marks the emittance. The relatively well-defined border to the phase space and the mild tilt (indicating only mild divergence) are direct consequences of the 2D nature of the interaction.

of the quality of the electron beam, is given in Appendix A.3. Note that while the bunch does not propagate in the laser propagation direction, this does not mean it must be rejected under consideration of the ZVP bunch conditions. We must expand the definition of a ZVP electron bunch. In the mass-limited case, a bunch must propagate at some angle to the laser due to conservation of canonical momentum while it remains in the thrall of the laser pulse. For an equivalent bunch in a corresponding 3D simulation, the transverse geometric emittance in the  $z$  plane is  $7.4(10)$  nm rad. This electron bunch has a total charge of  $0.35$  nC for a slab of plasma of thickness  $0.75\lambda$  in the  $z$ -direction. Note again the two-dimensional nature of the interaction geometry. Electrons a distance less than twice the relativistic Larmor radius,

$$r_L = \frac{\gamma m_e v}{e|\mathbf{B}|} \quad (2.45)$$

will escape to the side of the target when rotated back towards the plasma. Here,  $\gamma$  and  $v$  correspond to the electron velocity. The total number of electrons in the mass-limited bunch is

$$N = 2n_e r_L L_z \Delta x, \quad (2.46)$$

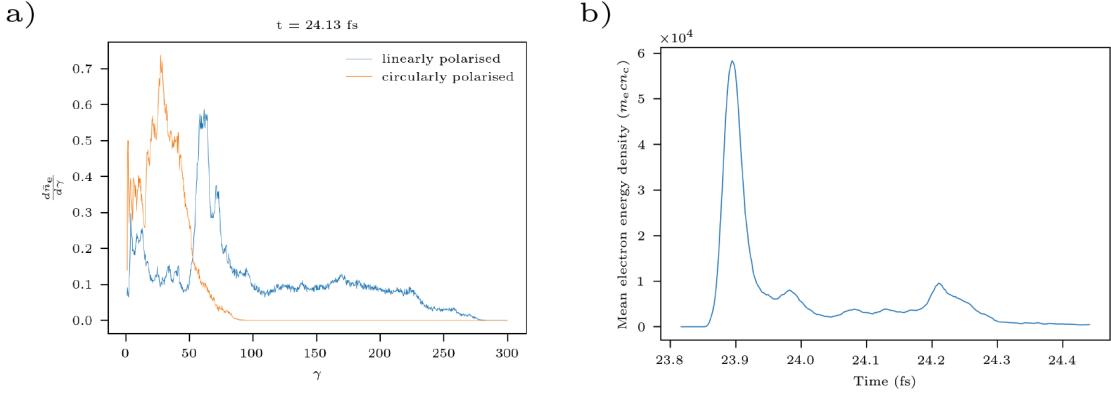
where  $L_z$  is the width of the plasma in the  $z$ -direction. Combining Equations 2.45 and 2.46 with 2.24 and 2.27 and taking  $v \approx c$  for the ultra-relativistic electron bunch,

$$N \approx 2\gamma n_c \frac{L_z}{k^2}. \quad (2.47)$$

Equation 2.47 can be rewritten in terms of fundamental constants as

$$N \approx 2(1 + 0.5 \frac{a_0^2}{n_e}) L_z \frac{m_e \epsilon_0 c^2}{e^2}. \quad (2.48)$$

For these simulation parameters, this corresponds to a total bunch charge,  $Q = eN$ , of  $0.37$  nC, a remarkably successful prediction of the ZVP model. Perhaps counter-intuitively, it would appear the total charge scales inversely with the plasma density. Instead, charge can be increased either by increasing the laser pulse intensity or  $L_z$ . Indeed, provided the laser pulse intensity remains relativistic, the focal spot can



**Figure 2.8:** a) Energy spectra for mass-limited electron bunches formed via linearly and circularly polarised laser pulses. b) Mean energy density streaked in time through the centre of a mass-limited ZVP electron bunch demonstrating the attosecond duration.

be increased indefinitely and there is no limit to the mass-limited electron bunch total charge. For a realistic laser pulse with beam width  $10\lambda$  incident on a larger plasma block, Equation 2.48 predicts a charge of 9.3 nC.

Caution is necessary when the standard definition of emittance is applied to a non-Gaussian beam profile risking over- or underestimating the emittance. When applied to an ideal Gaussian distribution, the elliptical contour that defines the emittance contains 39.3 % of the beam particles [133], however, Figures 2.7 e) and f) contain 75.5 % and 38.6 % of the beam respectively. The large overestimate of the former is a consequence of the long low density tail on the distribution to the bottom left of the plot.

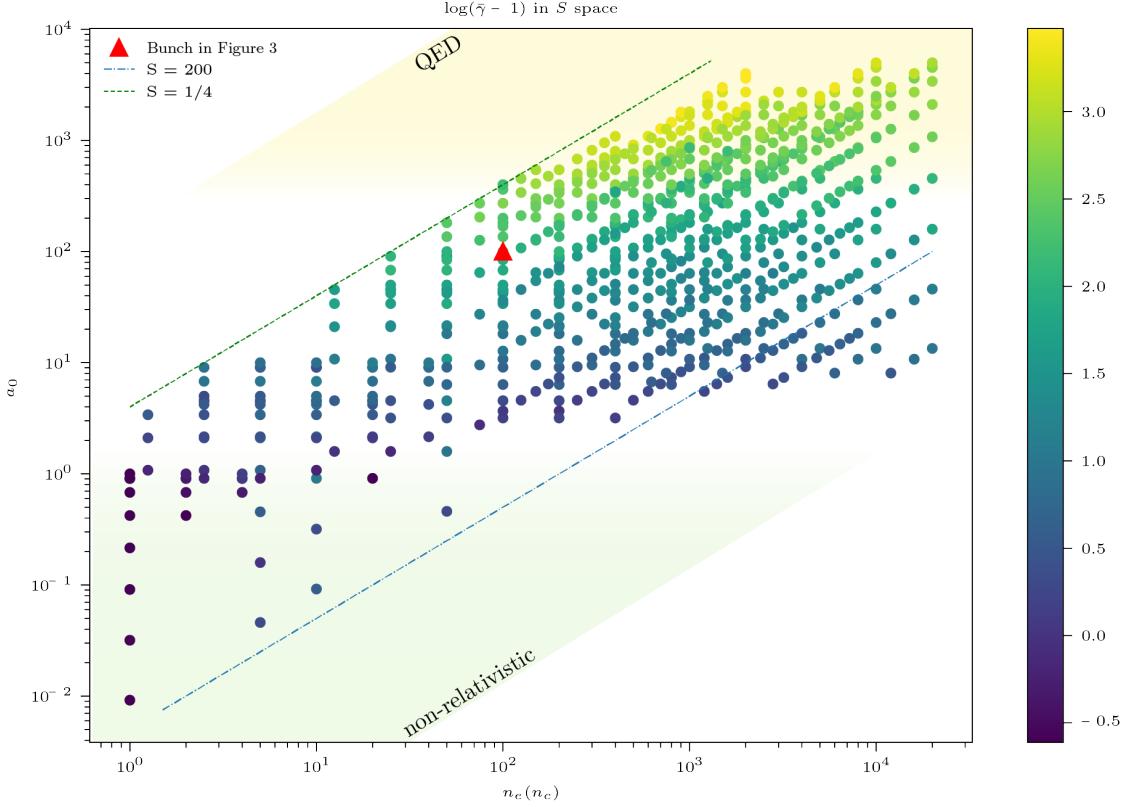
Figure 2.8a compares the electron bunch energy to an equivalent bunch produced by a circularly polarised laser pulse<sup>3</sup>. The mean electron bunch energy is over three times lower as there is no ZVP acceleration phase and there is no quasi-monochromatic structure [118]. Figure 2.8 b) plots the mean electron bunch energy in time at a point as the bunch passes, clearly demonstrating the attosecond duration of the electron bunch.

<sup>3</sup>A circularly polarised laser pulse will expel electrons from a mass-limited target in a corkscrew shape, the bunch is therefore only loosely equivalent.

### Applications: from electron bunches to attosecond light

There are a plethora of applications for high charge, attosecond electron bunches, primarily for the purpose of resolving attosecond scale phenomena [134]. Already femtosecond pump, attosecond probe experiments are underway [135, 136] but the higher intensities and charge densities accessible in the laser-solid regime compared to in laser-gas interactions [113, 126] would represent a step-change in the field by enabling atto-pump atto-probe experiments [109]. Potential applications include: electron microscopy and atomic diffraction to temporally resolve photoelectric processes such as Bragg diffraction [137], ultra-fast electron radiography to probe the evolution of the formation of magnetic fields in dynamical systems [138] or for XFEL seeding [114]: the record XFEL X-ray pulse duration stands at 280 attoseconds [139], substantially longer than the durations accessible using this technique. Electron bunches are also a promising alternative for radiotherapy due to their superior penetration depth in biotissues compared to X-rays [140].

Analogously to the HHG process, rapid acceleration of an electron bunch generates a burst of radiation whose properties (brightness, coherency, duration, spectrum) are determined by the corresponding properties of the electron bunch (charge, emittance, duration and energy). Thus globally electron bunches are used as a diagnostic tool in synchrotrons and XFELs. Other acceleration mechanisms for X-ray generation include: bremmstrahlung radiation from firing the electron bunch at a secondary high- $Z$  target [141], interaction with a counter-propagating laser pulse [142, 143] or injection into a laser or plasma wakefield accelerator, including accessing the solid density plasma wakefield regime [113]. The mass-limited electron bunches produced via the ZVP mechanism have transverse emittances comparable in all planes to those conditioned in state-of-the-art nano-Coulomb electron bunch accelerators [144, 145]. Such facilities typically produce electron bunches with geometric emittances of  $\sim$  mm rad prior to damping ring injection [146] and  $\sim$  nm rad post-injection [144]. Thus, the mass-limited ZVP electron bunches are ideal candidates for the production of bright X-rays of unprecedentedly short



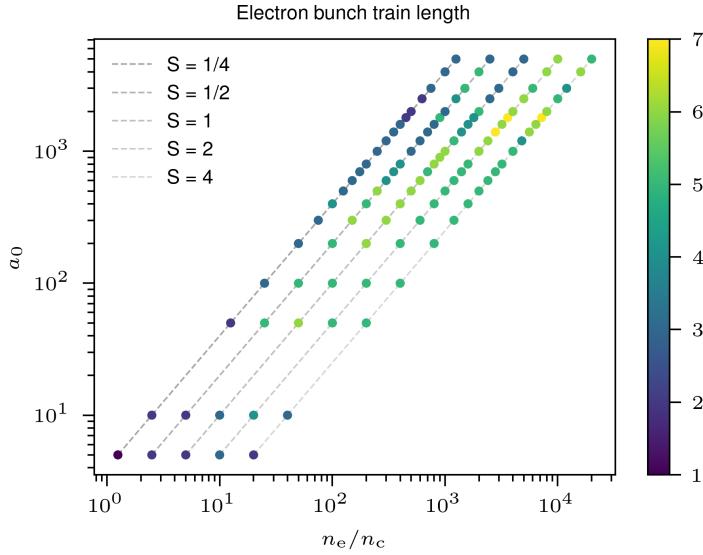
**Figure 2.9:** Mean mass-limited ZVP electron bunch normalised kinetic energies extracted from 2D PIC simulations. The bunch detailed in Figure 2.7 is highlighted.

duration and are suitable for study in the new attosecond regime with applications to physical, chemical and biological systems.

#### Parameter scan of electron bunch mean energy

It is important to confirm that mass-limited electron bunches follow the same scaling relations as has previously been confirmed for bulk ZVP electrons in both 1D [118] and 2D [117] PIC simulations. The mean mass-limited electron bunch kinetic energies were extracted from 120 2D PIC simulations and are plotted in Figure 2.9. The dependence of the energy on both  $a_0$  and  $\bar{n}_e$  is immediately apparent and therefore these electron bunches are accelerated by a non-ponderomotive mechanism.

The parameter scan took plasma block densities ranging from the critical plasma density to well-beyond solid density for the aluminium target and with laser pulse peak intensities ranging from non-relativistic ( $a_0 < 1$ ) through to the Quantum Electro-Dynamics (QED) plasma regime ( $a_0 > 300$ ) up to a peak  $a_0 = 5000$  to



**Figure 2.10:** Mass-limited electron bunch train length as a function of laser intensity and plasma density. Lines of constant peak  $S$  parameter are plotted for reference. Clearly, decreasing  $S$  reduces the bunch length.

investigate the change in scaling observed by Savin *et al* [127] at the onset of QED effects. This study is also the first to extract specific bunch energies as opposed to total simulation box energy gain, representing a far more precise test of ZVP theory. Particle merging was turned on for macro-photons and macro-electrons for  $a_0 \geq 1800$  to prevent overloading of the available supercomputer memory due to the vast number of particles produced in this SF-QED regime.

A total of 856 data points for mass-limited electron bunch mean energies were extracted from the simulations. Bunches are arranged in trains with an average length of 4, however isolated bunches are more useful for attosecond diagnostics. The number of bunches in a given train as a function of simulation peak intensity and plasma density is plotted in Figure 2.10. It is clear that reducing the peak  $S$  reduces the train length. Where  $S < 1$ , the transition to relativistic self induced transparency occurs in the rising edge of the laser pulse corresponding to a breakdown of ZVP and an early cutoff to the bunch train. Other routes to isolated attosecond electron bunches can be likely borrowed from CSE theory given the close relation between the two. The interest in isolated attosecond pulse of radiation has led to the proliferation of ideas, suggested techniques include using the attosecond lighthouse

technique using laser pulse wavefront rotation [147], few cycle laser pulses[148], non-collinear laser pulse gating [149] and circular polarisation gating [150].

Return now to Figure 2.9. The ZVP energy relations must be modified before comparison to the data since these electron bunches experience a further direct laser acceleration phase before reaching the measurement point. Thévenet *et al* [115] suggested that attosecond electron bunches produced in reflection exhibit precisely the phase and energy properties required to ‘surf’ the reflected laser pulse and experience vast acceleration gradients over the Rayleigh length of the laser pulse. This process is known as Vacuum Laser Acceleration. Potentially providing a route to a fully optical scheme to create GeV nano-Coulomb electron bunches. Certainly some of the bunch will have the necessary injection close to the laser propagation axis, however much of the bunch propagates at an angle to the laser and this must be accounted for. Interestingly in Figure 4.1 modelling oblique incidence, more of the bunch appears to propagate along the axis. A parameter scan is necessary to find the optimal parameters.

Electrons are ejected from the plasma in phase with the the subsequent laser pulse cycle peak and diverge outwards from that point. Near to the laser focus, the electrons experience a field

$$\mathbf{E}(x, y, t) = E_0 e^{-(y-f_y)^2/w_L^2} \cos(k_L(x - f_x) - \omega_L t) \hat{\mathbf{y}} = E_0 f(x, y, t) \hat{\mathbf{y}}. \quad (2.49)$$

The work done by this field is then

$$\Delta T = \int e \mathbf{E} \cdot d\mathbf{x}. \quad (2.50)$$

Note that for this process, the laser pulse electric field and electron bunch direction of travel will always be aligned no matter to which side of the plasma bulk the electrons are accelerated to and therefore  $\Delta T$  will always increase the energy of the electron bunch. The electron trajectories can be approximated at linear from the ejection point  $(y_e, x_e)$  to the observation point  $(y', x')$ , correspondingly,

$$x = y \frac{(x' - x_e)}{(y' - y_e)}, \quad t = \sqrt{x^2 + y^2}/c. \quad (2.51)$$

Equation 2.50 can now be integrated along a given electrons path and therefore the gamma factor increases by

$$\Delta\gamma = \int \frac{eE_y(y)dy}{m_e c^2} = a'_0(y)F, \quad (2.52)$$

where  $a'_0$  is the vector potential of the subsequent laser pulse cycle peak and

$$F = \frac{2\pi}{\lambda} \int f(y)dy. \quad (2.53)$$

The gamma factor gain from the ZVP acceleration phase is determined by the electric field at the corner of the plasma block,

$$\Delta\gamma_{ZVP} = \frac{(a_0 e^{-(2\lambda - y_f)^2/L_0^2})^2}{2\bar{n}_e} = 0.31 \frac{a_0^2}{\bar{n}_e}. \quad (2.54)$$

Thus,

$$\gamma = 1 + (0.31) \times \frac{a_0^2}{\bar{n}_e} + F \times a'_0. \quad (2.55)$$

This final term could be neglected or at least reduced somewhat once super-Gaussian spatial laser pulses become standard in this intensity regime or by using a suitable plasma separator [151], as applied in [152]. Both acceleration phases fail to meet the criteria of the Lawson-Woodward theorem. The ZVP phase is dependent on the existence of electrostatic forces, while the secondary phase occurs for a finite interaction region.

The data of Figure 2.9 is taken from two measurement points, enabling two  $F$  values to be investigated. Fits to Equation 2.55 were made for each data set using the Ordinary Least Squares regression model of the statsmodels Python module [seabold2010statsmodels], allowing the constants to vary. The non-linear ZVP energy can be applied to the linear model via construction of the necessary composite parameter. Both fits were successful with  $r^2$ -values of 0.81 and 0.84 and found the same ZVP pre-factor  $F_{ZVP} = 0.47 \pm 0.02$ , slightly above that anticipated from the model. The model is sensitive to the laser pulse intensity, likely much of the bunch originated from closer to the laser pulse focus than the target corner explaining the higher value.

The first data set measurement point was  $(0.41 \mu\text{m}, 0.25 \mu\text{m})$  from the target edge. This gives  $F = 0.31$  compared to  $0.22 \pm 0.02$  predicted by the fit. The second, at  $(0.41 \mu\text{m}, 0.31 \mu\text{m})$  from the target edge giving  $F = 0.39$  compared to  $0.34 \pm 0.02$  from the fit.

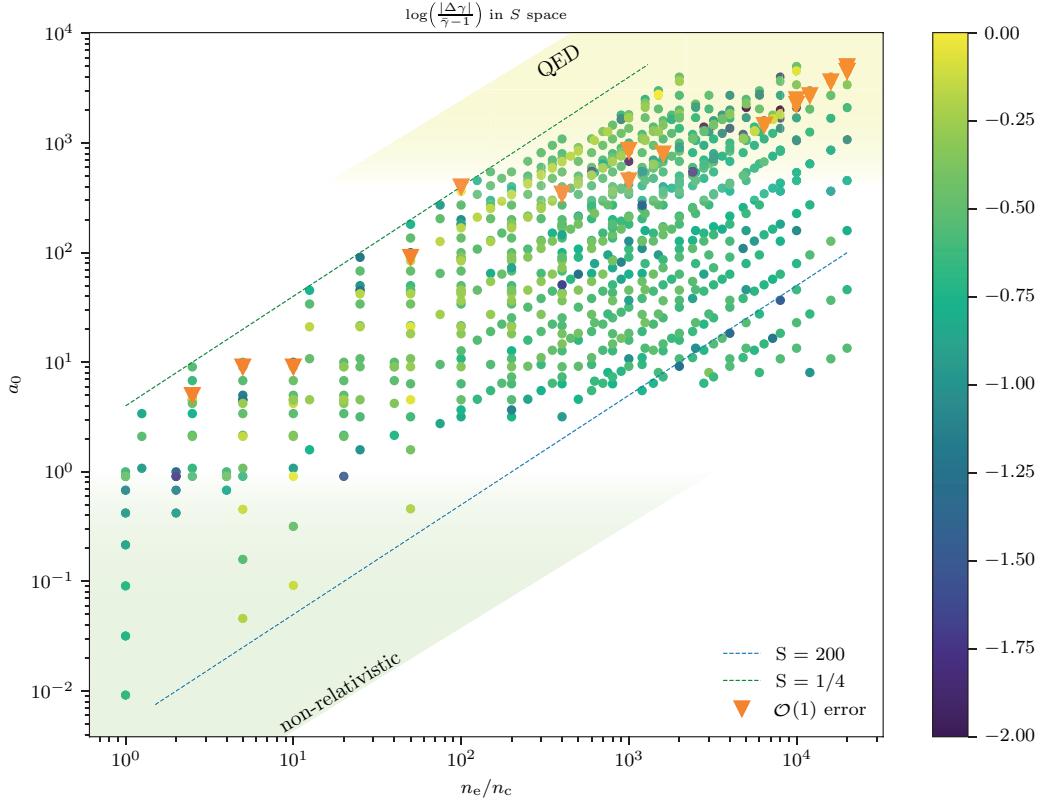
This is the first demonstration of ZVP theory to calculate absolute values and not only the scaling relationship. Such order of magnitude calculation is essential for the comparison to other absorption models and the determination of dominant modes. It is certainly remarkable that such a simple theory for energy absorption has such predictive success in this highly non-linear and seemingly chaotic many particle system. It is interesting that increasing laser intensity to such extremes will, at least for a short time, cause relativistic effects that simplify the dynamics before the total annihilation of a target.

The relative error between data and theory is plotted in Figure 2.11. Those points marked by an orange triangle have associated errors of over an order of magnitude. Reassuringly, such points occur only after the onset of QED effects, known to impact the ZVP mechanism [127] and for  $S < 1$ , that is, where the plasma becomes relativistically transparent to the laser pulse, a fundamentally different regime. The model is inconsistent in the non-relativistic domain as can be expected. It is particularly interesting that there is no indication that large  $S$  causes a breakdown of the model, extending the applicability of the model further than previously considered and opening up the field to a wider range of conditions, such as that of shock compressed plasmas. To summarise, it would appear the ZVP model is valid for  $10 \leq a_0 \leq 300$  and  $S \geq 1$ .

### 2.4.1 Energy absorption in the ZVP regime

As stated previously the laser-plasma coupling exists in a state of adiabaticity with the exception of the ZVP acceleration phase and hence Equation 2.29 describes the absorption of laser pulse energy. As two bunches are produced per laser pulse cycle, the rate of energy transfer is therefore

$$R = \frac{U\omega_L}{\pi}, \quad (2.56)$$



**Figure 2.11:** The relative errors for each mean energy data point compared to Figure 2.9. The orange triangles indicate data points for which the model fails to predict the mean energy.

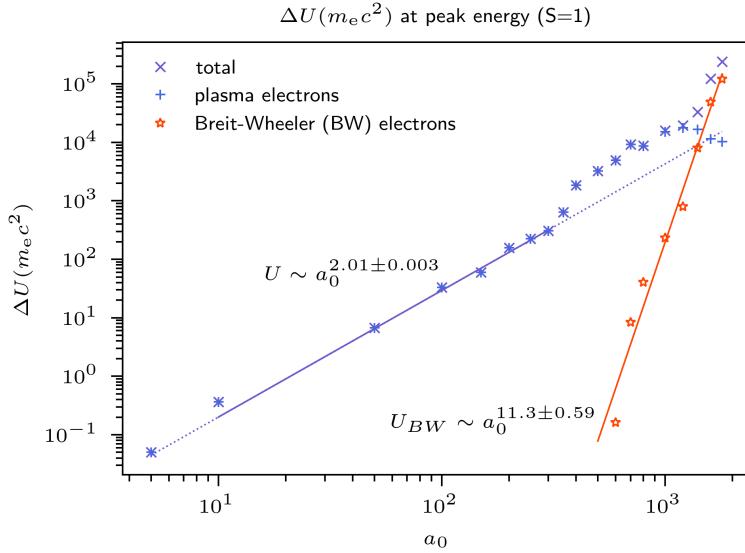
for normal incidence. To observe the scaling for  $U$  in 2D PIC simulations, peak instantaneous electron bunch energies escaping to the rear of the bulk were extracted from those PIC simulations with  $S = 1$ . For constant  $S$ ,

$$U \sim a_0^2. \quad (2.57)$$

Energies are plotted in Figure 2.12. Fitting the total energy within the range of validity established for the ZVP model finds

$$U \sim a_0^{2.01 \pm 0.003}, \quad (2.58)$$

reproducing with great success the anticipated scaling within the ZVP regime. It is satisfying that the number of electrons and bunch mean energies both follow their anticipated ZVP scalings. It was not possible to reproduce the constants of Equation 2.29 as the neutralising return current in the plasma bulk generates an electrostatic



**Figure 2.12:** Peak instantaneous bulk electron bunch total energies after escaping to the plasma bulk rear. Energies begin to deviate from the anticipated scaling at approximately  $a_0 = 300$ . Above  $a_0 = 1000$ , BW pair produced electrons begin to dominate and the peak energies rise rapidly.

field on the rear side of the plasma block, decelerating bulk electron bunches as they escape the plasma. It should be possible to calculate the deceleration by considering the number of electrons expelled by the plasma. It is, however, clear from the simulations that at least some electrons in the escaping bunch are trapped by this rear-side potential well reduces its ability to slow electrons.

While Equation 2.29 describes energy absorption into hot electron bunches, the coupling of such hot collisionless electrons to the bulk plasma given the lack of collisionality is naturally indirect. There are two key mechanisms [153]. Firstly, via a cooler resistive return current of electrons that neutralises the current of the injected hot electrons that escape the potential well of the front surface. Since all hot electrons travel at approximately speed  $c$ , the magnitude of the return current depends not on the total energy absorbed but instead on the total number of electrons injected, as given by Equation 2.28, depending linearly on laser spot area and the electric field magnitude and not on the plasma density<sup>4</sup>. Secondly, via the formation

<sup>4</sup>Note that for a sufficiently thin target, the return current induces an electrostatic field on the back surface of the target which can then reflect hot electron bunches and decelerate them to the point of a return to collisionality. This is a reality for the PIC simulations explored in this thesis, however, since realistic targets are much thicker this shall be neglected.

of large amplitude bulk plasma waves induced in the wake of the hot electron bunches. Sherlock *et al* [153] calculate the magnitude of the induced wakefield to be

$$E_W = \frac{eN_e c}{\omega_p \epsilon_0} = \sigma \sqrt{\frac{m_e \epsilon_0}{n_e}} E_L, \quad (2.59)$$

where here the bunch velocity has been set to  $c$ , bulk electrons will be accelerated by  $E_W$  and their kinetic energy converted to heat via collisions. Interestingly, this reproduces the mid temperature electron scaling with density that was observed by Chrisman *et al* [154] in their study of hot electron energy coupling in cone-guided fast ignition of inertial fusion targets. This is a different possible explanation to their self-declared ‘hand waving argument’. Excluding this study, such formulations for heat transfer to the plasma bulk within the ZVP regime remain untested in simulations.

Note also that as the laser pulse intensity rises, the fraction of energy absorbed by the ion species increases. Savin [119] determined for  $S = 1/2$ ,  $a_0 = 100$ , that this would be almost 20%. Energy is absorbed by ions via the hole boring mechanism as described later in this Thesis.

#### 2.4.2 Unpacking the QED effects of Figure 2.12

In Savin’s acclaimed paper [127], they determined theoretically and demonstrated in simulation that at  $a_0 = 300$ ,  $n_e = 50n_c$ , there is a transition from standard ZVP scalings to an enhanced QED scaling associated with Breit-Wheeler (BW) electrons increasing the pseudocapacitor plate charge. Explicitely,

$$T \sim \frac{a_0^5}{\bar{n}_e}. \quad (2.60)$$

At first glance of Figure 2.12, Savin’s results are inconsistent with this parameter scan. Perhaps the measurement method can explain this via to the well known effect of radiation trapping from to RR [155], also observed in these PIC simulations. After acceleration across the pseudocapacitor, the electron bunch encounters the subsequent laser peak. If the electron bunch gamma factor and laser intensity are both large enough, electrons radiate a significant fraction of their energy and are thus stopped in their tracks. Unable now escape the potential well at the plasma

surface they are trapped and are not observed to escape the plasma until the laser pulse intensity reduces. Such an effect would not impact Savin's scalings but would of course inhibit the observation of the scaling for electrons rear of the plasma block. Note that should this be the case and the collisionless electron bunches remain within the plasma bulk, Savin's ZVP QED model truly applies directly to energy absorption by the plasma bulk.

There is another interpretation. Returning now to Figure 2.12, there are two interesting aspects. Firstly, the sudden jump in total energy at  $a_0 \approx 300$ . This cannot be explained by ZVP theory nor QED theory since the jump is observed with QED effects switched off. Secondly, the even sharper jump in total electron bunch energy above  $a_0 = 1000$ . Decomposing the total energy into bulk electrons and those produced via the BW process, this is clearly a QED effect. Energy in BW produced electrons scales at a staggering

$$U_{\text{BW}} \sim a_0^{11.3 \pm 0.8}, \quad (2.61)$$

while the energy of bulk plasma electrons decreases. Perhaps this is a signal of Savin's QED ZVP electron bunches only at a higher energy due to the substantially greater plasma density of these simulations. The reduction in bulk electron energy can be attributed to an oversaturation of the front surface with BW electrons.

Combining equations 1.54 and 2.29 and assuming the electron radiates all its energy to the photon, for the ZVP mechanism, at the point of emission one finds

$$\chi_\gamma = \frac{\sqrt{2}|\mathbf{E}|}{E_S} \frac{a_0^2}{\bar{n}_e}. \quad (2.62)$$

The probability of BW pair production begins to rapidly increase around  $\chi = 1$ , therefore, the transition to QED will occur at

$$a_0 \approx \left( \frac{E_S}{\sqrt{2}} \frac{e\bar{n}_e}{m_e c \omega_L} \right)^{1/3} = \left( \frac{a_s \bar{n}_e}{\sqrt{2}} \right)^{1/3}, \quad (2.63)$$

where  $a_s = 7.73 \times 10^5$  is the normalised vector potential associated with the Schwinger Field. Repeating the calculation instead as a function of  $S$ , one finds instead

$$a_0 \approx \left( \frac{a_s S}{\sqrt{2}} \right)^{1/2}. \quad (2.64)$$

This corresponds to a transition to rapid pair production with Savin's parameters of  $a_0 \approx 301$  versus  $a_0 \approx 739$ , consistent with both studies. This is interesting, ZVP with Savin parameters predicts a significantly lower transition to QED compared to the Wilks scaling for  $\mathbf{J} \times \mathbf{B}$  heating. Following Savin's QED ZVP theory, one can predict

$$U_{\text{QED}} \sim \frac{a_0^7}{S}, \quad (2.65)$$

therefore much work remains understand the scaling of Figure 2.12 and unify these results, starting with the application of the methods of this analysis to the parameter space explored by Savin. Undoubtedly, the advent of next generation exa-watt scale lasers and access to this regime will be exceedingly interesting if such scalings in bunch energy can be maintained.

Linear Breit-Wheeler can safely be neglected in these simulations. There is simply not enough energy in the system. For two interacting photons of energy  $E_1$ ,  $E_2$ , by consideration of four-momenta, the threshold condition for pair production is

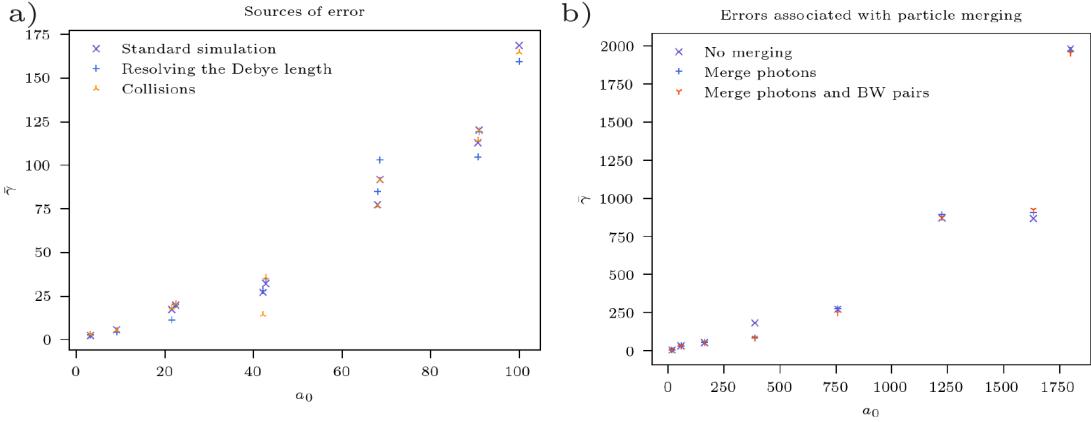
$$E_1 E_2 \geq (m_e c^2)^2. \quad (2.66)$$

For a near-infrared laser photon of energy  $\hbar\omega_L$ , the interacting photon must have an energy in excess of 200 GeV. Despite the extreme acceleration gradients considered in this thesis, the ZVP mechanism predicts photons of energies no greater than 10 GeV and thus linear Breit-Wheeler is suppressed.

He *et al* [156] identified an alternative mechanism for linear BW using solid density targets, where forwards- and back-scattered high energy radiation reaction produced photons interact within a hollowed out plasma channel. In the geometry of interest in this thesis, while photons are produced in both directions, their production is localised to the plasma surface and at no points do their paths cross and thus this mechanism does not occur.

### 2.4.3 Errors

A set of simulations were undertaken to explore the stability of some assumptions of the PIC simulations. The effective of including collisions, increasing the temperature



**Figure 2.13:** An exploration of ZVP 2D PIC simulation stability and assumptions. a) Electron bunch mean energies extracted from simulations with  $a_0$ ,  $\bar{n}_e = 100$ . Here, the  $x$ -axis is the normalised vector potential of the laser pulse cycle that made them, this is why there appears to be results stacked. b) Electron bunch mean energies extracted from simulations with  $a_0$ ,  $\bar{n}_e = 1800$  both with and without particle merging.

to resolve the Debye length and particle merging were explored. The results are presented in Figure 2.13. In general including collisions had little to no effect on the results and changing the Debye length had a small but not systematic impact on the results. Note that while numerical heating is not directly dependent on laser intensity, increases to  $a_0$  would increase the electron bunch density. Thus, resolving the Debye length of the ultra-high density (orders of magnitude greater than solid density) electron bunches produced by the higher laser pulse intensities considered here is no straightforward task. Particle merging was essential for simulations with  $a_0 > 1800$  due to the proliferation of high energy photons and electron-positron pairs. At least at  $a_0 = 1800$  it would appear particle merging had negligible impact, however it is likely the error would increase with  $a_0$ .

## 2.5 Conclusions

The Zero Vector Potential mechanism describes the post-ponderomotive rapid absorption of ultra-relativistic laser energy by a solid density overdense, collisionless and fully ionised plasma on the timescales of ion immobility. The defining characteristics of the mechanism have been identified in 3D PIC simulations including the observation of a zero of the vector potential propagating at speed  $\approx 1.4c$  through

a high density ZVP electron bunch at the front surface of the plasma early in the ablative journey of the electron bunch. Simulations have suggested that from currently operational 10 PW short pulse laser facilities and foam targets, the ZVP mechanism can produce a train of attosecond duration, nano-Coulomb electron bunches, each with a transverse emittance of a few nm rad. Such charge and quality is comparable to state-of-the-art electron bunch accelerators but on paradigm shifting timescales. Such timescales being those on which atomic processes occur, these electron bunches could be manipulated to literally ‘shed light’ onto fundamental biological and chemical processes. Via a massive 2D PIC parameter scan the energies of such mass-limited electron bunches have been compared to those predicted by the ZVP model, identifying a range of validity for the model, specifically  $a_0 > 10$ ,  $S > 1$ . These simulations were also used to confirm the energy absorption scaling in 2D up to and into the QED regime. Chapter 4 details the upcoming GEMINI PW experiment to observe ZVP electron bunches for the first time. On the theme of experimentation, the following Chapter will now switch gears from absorption to reflection for the discussion recent results from the ORION laser facility.

# 3

## Attosecond X-ray harmonics on the ORION laser facility

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## 3.1 A plan

While simulations are relatively cost effective and experiments cumbersome and technically challenging, the proliferation of errors in simulation codes lead to their deviation from reality while reality itself deviates from the idealistic input conditions to the codes. At least for these first few decades of the 21st Century, experimentation remains an essential component of the scientific method. Thus, our expectations must also be scaled back to reality. This chapter reports on our March 2023 experiment on the ORION laser facility, AWE, Aldermaston. The UK's most powerful sub-picosecond laser. Yet, one cannot access the ZVP parameter space at these intensities and pulse durations: the  $S$  parameter is too large and the preplasma scale length too steep. Instead, this chapter will focus on the ROM model of HHG.

Have a look at the intro in baeva's thesis, good outline, include first solid HHG observation and oscillating mirror model.

The early developments in the field are outlined by Teubner *et al* [157].

HHG from solids has demonstrated significantly higher conversion efficiencies than for HHG from gases nor are there any limitations to the applied laser intensity [157]. SHHG has thus been a field of huge interest over the past three decades for the production of bright coherent attosecond harmonics.

SOme applications Applications Bright high harmonics of the laser pulse Novel bright attosecond duration X-ray diagnostics Accessing the Schwinger Limit in vacuum Photon-photon scattering

For CWE: <https://iopscience.iop.org/article/10.1088/0953-4075/43/21/213001/pdf>  
 Note that somewhere in this paper they say CWE optimises for 0.02 lambda (use this to explain why in ORION short sim I used a lambda slightly larger than this)

## 3.2 The experiment

The two experimental goals were to resolve and measure the absolute intensity of X-ray harmonics produced with the ORION SP2 beamline. After debris from the plasma mirror optic damaged the parabola we were required to switch to the

Parameter	ORION	GEMINI	
Beamlines	SP1	SP2	N & S
Power (PW)	0.5	1	0.5
Energy (J)	200	500	12
Wavelength (nm)	527	1053	800
Parabola, $f/\#$	$f/3$	$f/3$	$f/2$
Focal spot FWHM ( $\mu\text{m}$ )	< 20	< 10	2
Duration (fs)	500	500	40
Shot rate	5/day	5/day	3/min
Peak $a_0$ (approx)	10	30	24

**Table 3.1:** The ORION and GEMINI petawatt class short pulse beamlines for comparison. The GEMINI North (N) and South (S) beamlines are equivalent.

higher contrast but lower energy SP1 beamline and only after the experiment did we established it was not possible to resolve X-ray harmonics with the ORION target chamber geometry. However, we were able to measure the absolute intensity for both SP1 and SP2 beamlines and for CVD and PMMA targets.

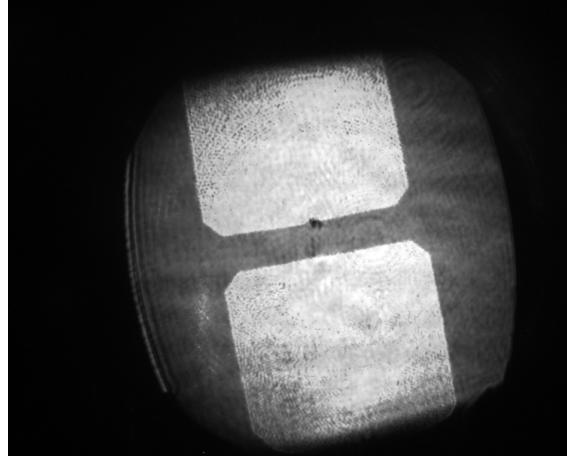
Typical ORION and GEMINI beam parameters are presented in Table 3.1. The GEMINI-PW laser facility of the Central Laser Facility (CLF) at Harwell Campus, Oxfordshire is a petawatt class facility consisting of two 30 fs beams, **N!** (**N!**) and **S!** (**S!**), each delivering a maximum focused intensity of  $2 \times 10^{21} \text{ W cm}^{-2}$  at a repetition rate of 0.05 Hz. Such high frequency of operation has led to a paradigm shift in high power laser physics experimentation with the arrival of statistically significant results.

The ORION SP1 is produced by passing the SP2 through two frequency doubling crystals creating two equivalent beamlets that are then overlaid at the target. The double beamlet structure at the near-field is imaged in Figure 3.1.

The ORION SP2 beamline intensity temporal profile can be modelled as

$$I_{\text{SP2}} \sim (I_{0 \text{ p}} \operatorname{sech}(t/t_p)^2 + I_{0 \text{ pp}} e^{-(t/t_{pp})^2} + I_{0 \text{ ps}} e^{-\text{abs}(t)/t_{ps}} + I_{0 \text{ ns}} e^{-(t/t_{ns})^8}), \quad (3.1)$$

where the constants are detailed in Table 3.2 [158]. The picosecond pump residual arises from parametric fluorescence in the ps OPA, the picosecond pedestal from



**Figure 3.1:** Image of the ORION SP1 double beamlet structure in the near-field. The two beamlets are then superimposed at the target.

Pedestal, $i$	$I_0$	$t_i$
Main pulse, p	1	0.2 ps
Picosecond pump residual, pp	$10^{-3}$	3 ps
Picosecond pedestal, ps	$5 \times 10^{-5}$	8 ps
Nanosecond pedestal, ns	$10^{-11}$	3 ns

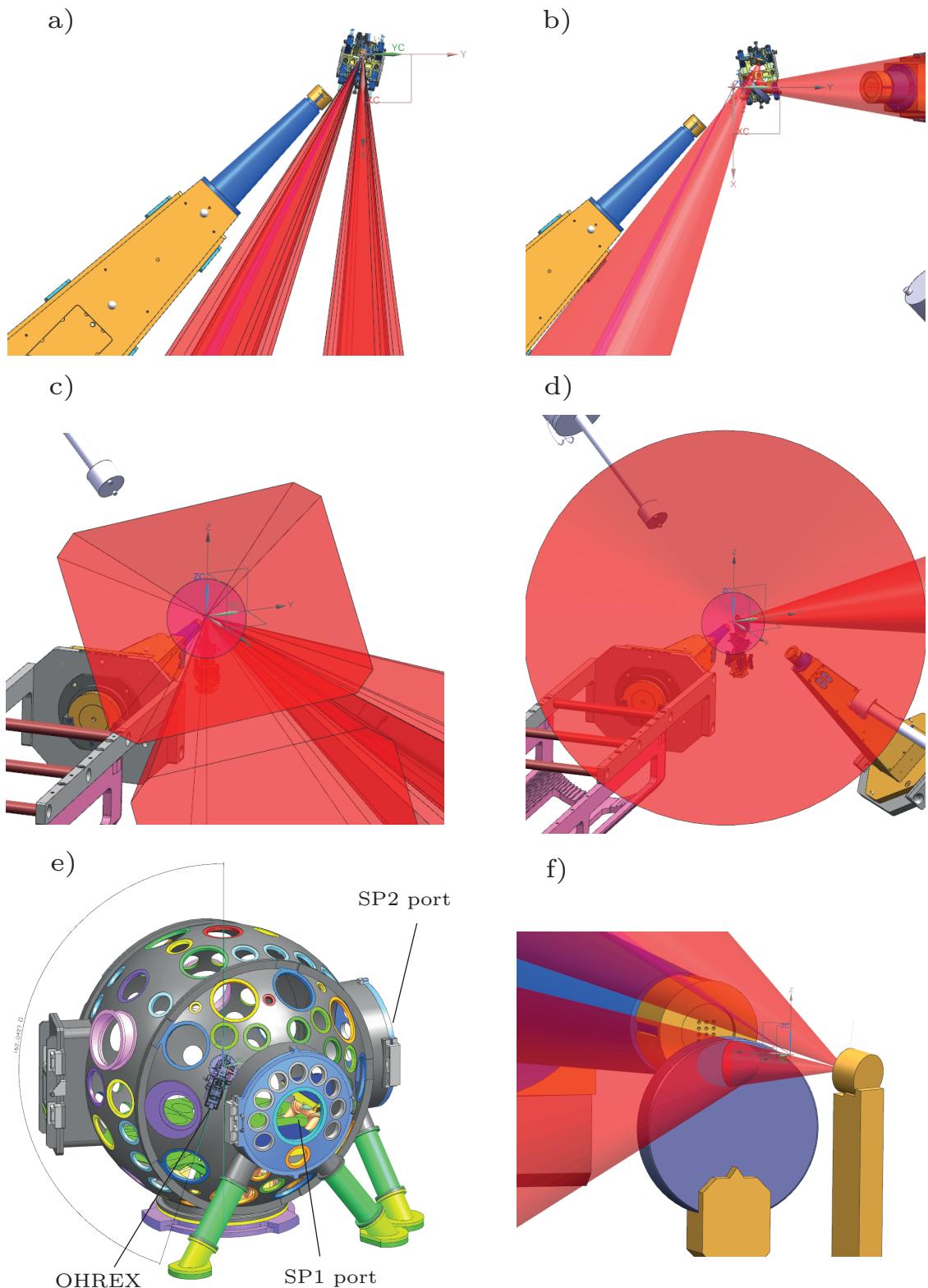
**Table 3.2:** The native ORION SP2 pulse and prepulse pedestal constants as defined in Equation 3.1.

scatter in the stretchers and noise on the OPA pump laser and the nanosecond pedestal from parametric fluorescence from the nanosecond OPAs. The SP1 temporal intensity profile is

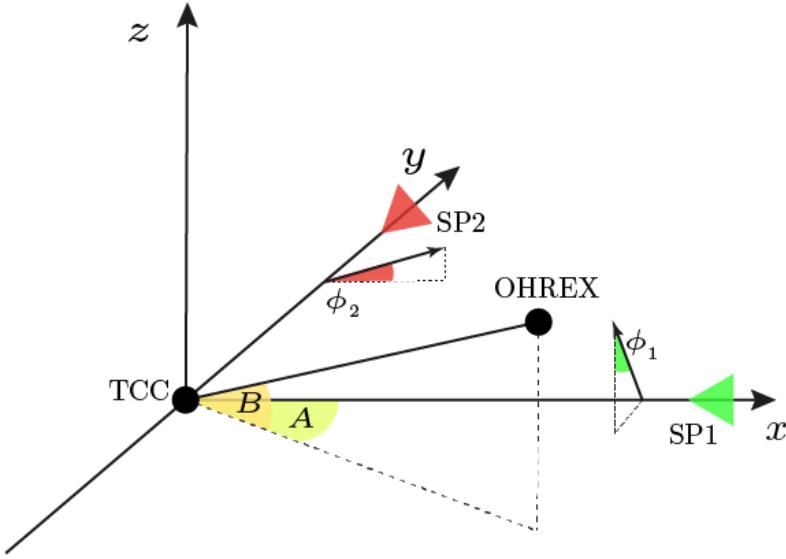
$$I_{\text{SP1}} \sim (I_{\text{SP2}}^2 + I_{\text{SP2}} \times 10^{-8}). \quad (3.2)$$

The second term is due to limitations of the harmonic separation system. The frequency doubling mechanism is not 100 % efficient, *i.e.* some of SP2 beamline remains and must be filtered out. The main pulse of the GEMINI beamlines is more appropriately modelled as a Gaussian.

The experimental design is relatively straightforward. CAD drawings of the experimental setup are presented in Figure 3.2. The laser is focused onto the target at **TCC!** (**TCC!**) with the ORION High Resolution X-ray (OHREX) spectrometer positioned along the specular reflection direction. Additional X-



**Figure 3.2: CAD images of the ORION target chamber set up for this experiment.** a) and b) SP1 and SP2 beamlines respectively from parabola to TCC to OHREX spectrometer. c) and d) SP1 and SP2 beamlines respectively looking from OHREX to TCC. e) Annotated ORION target chamber model. f) SP2 beamline reflecting off plasma mirror before main target incidence.



**Figure 3.3:** ORION target chamber geometry with the location of the target, OHREX spectrometer and the SP1 (green) and SP2 (infra-red) beamlines and their corresponding polarisations,  $\phi_1 = 11.8^\circ$  and  $\phi_2 = 16.4^\circ$ .

ray pinhole diagnostics monitor the spot size. The SP2 beamline reflects off an additional **PM!** (**PM!**) optic at  $45^\circ$  **AOI!** (**AOI!**) to improve the laser prepulse contrast. Then both the SP1 and SP2 beamlines arrive at **TCC!** along the same axis.

Both the non-linear HHG interaction and the linear OHREX crystal spectrometer are sensitive to the polarisation of incident light. It is therefore important to track the polarisations of the beamlines as they propagate through the system.

The ORION target chamber has its own defined geometry with the target located at the origin (**TCC!**), described in Figure 3.3. The polarisation angles are  $\phi_1 = 11.8^\circ$  and  $\phi_2 = 16.4^\circ$ . Following reflection of the SP2 beam off the plasma mirror, both the SP1 and SP2 beamlines propagate in the  $-\hat{x}$ -direction towards the origin. The OHREX crystal is located at

$$\mathbf{r}_{\text{OHREX}} = r_0(\cos B \cos A, -\cos B \sin A, \sin B), \quad (3.3)$$

where  $r_0 = 2.4$  m,  $A = 26.82^\circ$  and  $B = 18.15^\circ$ , setting the rotation angle of the target. This was achieved using the ORION Multi-Target-Mounts. Alignment was performed offline by Dr Ed Gumbrell and no further details will be provided here on that process. As in Figure 3.2e, the OHREX spectrometer sits on the outside of the target chamber tilted an angle  $18^\circ$  to the vertical.

The interaction plane is defined by the vector

$$\mathbf{n} = \frac{\mathbf{r}_{\text{OHREX}}}{r_0} \times \hat{\mathbf{x}} = (0, \sin B, \cos B \sin A). \quad (3.4)$$

The cosine rule can be applied to determine the polarisation of the laser pulses in the interaction plane, for polarisation vector  $\hat{\mathbf{E}}$ ,

$$\frac{\mathbf{n}}{|\mathbf{n}|} \cdot \hat{\mathbf{E}} = \cos \theta, \quad (3.5)$$

where  $\theta$  defines the angle between the polarisation vector and the vector normal to the interaction plane. This corresponds to angles out of the interaction plane of  $42.2^\circ$  for the SP1 beam (rotating anticlockwise out of the interaction plane when looking from TCC to parabola) and  $19.6^\circ$  for the SP2 beam (rotating clockwise out of the interaction plane when looking from TCC to parabola). Again applying the cosine rule, the angle of incidence is  $16^\circ$ .

The polarisation vector can then be split into s- and p-polarised components. The s-polarised component remains unchanged by the interaction but the p-polarised component must be rotated, its direction along  $\mathbf{n} \times \mathbf{r}_{\text{OHREX}}$ . In Section 3.4, I show the s-polarisation component is suppressed at the X-ray photon energies of interest and therefore can be neglected.

The OHREX spectrometer sits flat on the target chamber wall but is rotated an angle  $C = 18^\circ$  to the vertical. The OHREX interation plane is correspondingly rotated. Again, applying the cross-product, the OHREX interaction plane is

$$\mathbf{n}_O = [-\cos(A) \sin(B) \sin(C) - \sin(A) \cos(B) \cos(C), -\cos(A) \cos(B) \cos(C) + \sin(A) \sin(B) \sin(C)] \quad (3.6)$$

Applying the cosine rule between the p-polarised vector and  $\mathbf{n}_O$ , the angle of the polarisation vector of the harmonic beam out of the plane of the OHREX interaction is  $77.23^\circ$  for both SP1 and SP2.

### 3.2.1 Targets

Specialised target holders were designed and built by the AWE and CLF target fabrication teams. These hold the targets at the required angle out of the horizontal plane with an additional fiducial wire for alignment. Targets were mounted on the ORION Multi-Target-Mount and alignment performed offline by Dr Ed Gumbrell.

One might assume that target surface variation, root-mean-square roughness,  $\Delta s$ , need be less than the highest harmonic order that one wished to generate, however, Dromey *et al* demonstrated categorically in both experiment and simulation that surface roughness need only be less than the excursion amplitude of the oscillating electrons at the target front surface thus ensuring specular reflection and preventing scattering into the wings of the harmonic beam [159]. This is generally satisfied by  $\Delta s \ll \lambda_L$ . In their experiment, harmonics were visible provided  $\Delta s < \lambda_L/16$ . Furthermore, preplasma expansion leads an increase in target smoothness prior to the main pulse arrival.

There are multiple considerations when it comes to target density. Simulations have repeatedly emphasised the improvement of harmonic efficiency for reduced similarity parameter,  $S$ , [120, 126] corresponding to low electron density solid targets. In practice accessing the  $S \approx 1$  regime is not possible using a petawatt class laser system with even the lowest density solid plastic targets. However, the 2007 Vulcan experiment successfully reproduced the 8/3 scaling up to the X-ray regime for  $S \approx 50$  [160]. At the same time, as discussed in Section 3.3, target hole boring leads to an increased harmonic beam divergence. High electron density targets are perhaps therefore more suited to the petawatt class laser system thereby maintaining beaming of harmonics.

The surface roughness of a selection of low density plastic targets were analysed by the CLF target fabrication team leading to the choice of PMMA plastic targets. CVD targets were also selected, having the highest ionisation potential (and therefore lowest susceptibility to laser prepulse) of solids that can be polished optically flat albeit at a higher  $S$  parameter. Scans performed by AWE and CLF of the target surfaces are given in Figure 3.4.

**Figure 3.4: ORION experiment 3D scans of target roughness.** a) CVD with roughness  $\Delta s = 4 \text{ nm}$  and  $S \approx 26, 35$  for SP1 and SP2 beamlines respectively. b) PMMA with roughness  $\Delta s = 11.2 \text{ nm}$  and  $S \approx 10, 13$  for SP1 and SP2 beamlines respectively.

### 3.2.2 Contrast and plasma mirrors

As illustrated in Figure 3.2, a plasma mirror optic is included before the main target to improve the contrast of the SP2 beamline [161]. HYADES simulations of the effect of this optic are detailed in Section ???. The distance between the PM and the main target is 15 mm, thus the beam diameter at the PM is  $\approx 15/(f/3) = 5 \text{ mm}$  and the intensity on PM can be calculated. Provided the harmonic beam divergence is less than or equal to the  $f/3$  cone, as is anticipated, the harmonic beam should not be clipped by the PM in this setup.

The intention was to use AR! (AR!) -coated fused silica optimised for  $45^\circ$  AOI! and the SP2 pulse wavelength (BCP45R) [162]. This has a reflectance of 0.398 % at  $\lambda_L = 1053 \text{ nm}$  for unpolarised light. Unfortunately the debris from the PM caused damage to the parabola. For the second SP2 shot, the high reflectivity fused silica PM was replaced with a thin foil of silicon nitride. The reduction in available material of the thin foil reducing the risk of damaging debris. At  $45^\circ$  AOI!,  $\lambda = 1.053 \mu\text{m}$  and  $\phi_2 = 16.4^\circ$ , the silicon nitride PM pre-switch-on has a reflectivity of 5.55 % [163]. One can assume its behaviour after switch-on is consistent with the other PM since it is simply a plasma interaction now.

### 3.2.3 KBRXM

The ORION SP1 and SP2 spot size was routinely monitored on shot by the AWE laser team via two X-ray pinhole cameras and the time-integrating Kirkpatrick-Baez X-ray microscope (KBXRM) diagnostic, sat on the target chamber Port 21 and pointed at the notional TCC!. This reflective X-ray microscope optic has a ten times magnification and spatial resolution of  $15 \mu\text{m}$ , comfortably resolving the laser spot size. Images were again recorded with TR-type IP. While direct mapping between the recorded images and real laser irradiance profiles is a highly non-trivial problem with any proposed model depending strongly on the assumptions made.

However, by making the very reasonable assumption of spatial and temporal blurring of the spot size in the X-ray image due to electron transport away from the laser spot, clearly the measured spot size will be an upper bound on the real spot size. It follows that the upper bound on optical spot widths corresponds to a lower limit of irradiance (since the on shot energy is well-defined). These diagnostics determine sensible upper bounds for the laser spot size are  $20\text{ }\mu\text{m}$  and  $10\text{ }\mu\text{m}$  for the SP1 and SP2 beamlines respectively.

### 3.3 Theory

Note: in the below all we are simply saying is the reflected field is simply the negative of the incoming field with a phase modulation [121]. See also this paper for a simle explanation of OMM

For the moderately relativistic sub-picosecond ORION SP1 and SP2 beamlines the most appropriate description of their non-linear interaction with a solid density target is the Relativistic Oscillating Mirror (ROM) model. The full mathematical formation of which forms the main body of Baeva's 2008 thesis [164], a notably more rigorous derivation than her original paper on the same topic [125], applying the highly general Apparent Reflection Point formulism to the case of  $S > 1$  in relativistic similarity theory, showing that it coincides with the motion of the relativistic critical density surface. The following covers the more qualitative, physically motivated understanding of why this theory works, which is fundamentally that while the surface motion is generally complex and parameter space dependent, at the important point the surface motion is always parabolic. By application of the Bourdier method, Baeva demonstrated that the envelope of the spectral content of the reflected beam does not deviate from the normal incidence linearly polarised case for either s- or p-polarised incidence. Thus the following discussion will be restricted to normal incidence. One starts with relativistic similarity theory and Equation ??,  $\mathbf{p} \sim a_0$ . As discussed in Section 2.2, microscopically electrons follow similar relativistic trajectories yet from Equation 2.5,  $v_{\text{prop}}$ , the velocity of electrons along the propagation axis of the laser pulse, not necessarily relativistic except at

the passing of the zero of the vector potential where  $v_{\text{prop}} \rightarrow c$ . Macroscopically, the well-defined critical density surface within the skin depth (located at  $S = 1$ ), has a gamma factor that tracks the electrons' velocity along the propagation direction ( $x$ -direction) at that point,

$$\gamma_s = \frac{1}{\sqrt{1 - v_{\text{prop}}^2/c^2}} = \sqrt{\frac{1 + a(t, x)^2 + p_{\text{prop}}^2/(m_e c)^2}{1 + a(t, x)^2}}. \quad (3.7)$$

Equation 3.7 indicates surface gamma factors of order unity except at zeros in the vector potential at which point  $\gamma_s \sim a_0$ . In the vicinity of the zero,  $a \sim a_0 \sin(\omega_L t) \approx a_0 \omega_L t$ . Thus, while the velocity of the surface is a smoothly varying function around the maximum,  $v_{s, \text{max}} \sim 1 - \mathcal{O}(a_0^{-2})$ , the *gamma factor spikes* as the vector potential passes through zero and it is at this point that the surface emits high frequency photons. One can then approximate the duration of this high frequency component of the pulse. The surface will begin radiating such photons at some point near the gamma spike and continue radiating until the gamma factor drops down again after the gamma spike. In ZVP we observe the zero overtaking the emitted photons and thus radiation from the extended electron bunch that is emitted first arrives at the observer last. Here, we consider one point in the sub-light speed surface emitting radiation, thus the radiation overtakes the surface and points emitted first are observed first. The radiation emitted therefore has a duration of

$$\Delta l \approx (c - |v_{s, \text{max}}|)\Delta t \approx \frac{\Delta t}{\gamma_{s, \text{max}}^2}, \quad (3.8)$$

where  $\Delta t$  is the temporal duration of the gamma spike.

Writing the surface velocity as a smooth function around its maximum,

$$\mathbf{v}_s(t) = -(v_{s, \text{max}} - c\alpha(\omega_L t)^2)\hat{\mathbf{x}}. \quad (3.9)$$

Without loss of generality the peak of the gamma spike is set to  $t = 0$  for cleaner notation. Around the gamma spike, the surface gamma factor is thus

$$\gamma_s(t) \approx \frac{1}{\sqrt{1 - (\mathbf{v}_{s, \text{max}}/c)^2 - 2\alpha(\omega_L t)^2}}. \quad (3.10)$$

Since  $1 - (\mathbf{v}_{s,\max}/c)^2 = 1/\gamma_{s,\max}^2$ , one can calculate the FWHM of the gamma spike is

$$\Delta t \sim \frac{1}{\sqrt{\alpha}\omega_L \gamma_{s,\max}}. \quad (3.11)$$

Hence,  $\Delta t_{\text{FWHM}} \sim a_0^{-1}$  and  $\Delta l \sim 1/a_0^{-3}$  and the cut off frequency corresponding to the highest frequency accessible for a pulse of that duration is

$$\frac{\omega_c}{\omega_L} \sim \sqrt{\alpha} \gamma_{s,\max}^3 \sim a_0^3. \quad (3.12)$$

Within the ROM model these high frequency pulses are coherently produced with the low order harmonics and at intensities that require filtration to access the possible attosecond duration of these pulses.

One can access the harmonic spectrum envelope for lower energies than the harmonic cut off. From the Leontovich boundary conditions corresponding to zero energy flux at the reflection point. This corresponds to a balance of the electric fields at the plasma surface,  $x_s$ , hence for the incoming and reflected waves,

$$E_i(x_s(t) - ct) = -E_r(x_s(t) + ct) \quad (3.13)$$

for all  $t$ . Since we are making the approximation of a plane wave, the phase structure at the reflection point will be maintained as the reflected wave propagates away from the interaction region to our observation point ( $E_r(x_s(t) - ct) = E_r(x' - ct')$  for an obesrvation point  $(x', t')$  near an electric field peak.). Thus, determining the modulation of the reflected beam is akin to solving Equation 3.13. Writing  $\Phi = x_s(t) - ct$ ,  $\Psi = x_s(t) + ct$  and taking the time derivative of Equation ??,

$$\frac{dE_r(\Psi)}{d\Psi} = \frac{c - \mathbf{v}_s(t)}{c + \mathbf{v}_s(t)} \frac{dE_i(\Phi)}{d\Phi} \quad (3.14)$$

where  $v_s(t)$  is given in Equation 3.9. Integrating Equation 3.9 the reflected wave phase is

$$\Psi(t) = \Psi(0) + \frac{ct}{\gamma_{s,\max}^2} + \frac{1}{3} c \alpha \omega_L t^3. \quad (3.15)$$

Very close to the gamma spike peak the linear term will dominate the interaction, corresponding to the highest order harmonics as discussed previously. The non-linear term dominates for

$$\delta\Psi = |\Psi(t) - \Psi(0)| \gg \left(\frac{3}{\alpha}\right)^{1/2} \frac{c}{\omega_L \gamma_{s,\max}^3}. \quad (3.16)$$

Note the correspondence between Equation 3.12 and the right hand side of Equation 3.16. In this regime we can write

$$t = \left(\frac{3\delta\Psi}{c\alpha\omega_L}\right)^{1/3} \quad (3.17)$$

For these intermediate times around the gamma spike the linear term can be neglected, this is equivalent to writing  $v_{s,\max} = c$ . Thus,

$$\frac{dE_r(\Psi)}{d\Psi} = \left(\frac{2c - c\alpha(\omega_L t)^2}{c\alpha(\omega_L t)^2}\right) \frac{dE_i(\Phi)}{d\Phi} \quad (3.18)$$

We assume the incoming radiation is slowly varying around the short duration of the gamma spike,  $dE_i(\Phi)/d\Phi \approx dE_i(\Phi)/d\Phi|_{\Phi=\Phi(0)}$  and so we arrive at an equation we can integrate,

$$\frac{dE_r(\Psi)}{d(\delta\Psi)} = \frac{2}{\alpha^{1/3}} \left(\frac{c}{3\omega_L^2 \delta\Psi}\right)^{2/3} \frac{dE_i(\Phi)}{d\Phi}|_{\Phi=\Phi(0)} \quad (3.19)$$

obtaining

$$E_r(\Psi) = -E_i(\Phi(0)) - \left(\frac{3c^2}{\alpha\omega_L^2}\right)^{1/3} \frac{dE_i(\Phi)}{d\Phi}|_{\Phi=\Phi(0)} \times (\delta\Psi)^{1/3}. \quad (3.20)$$

One can see therefore that the reflected radiation gets the quasi-singularity

$$E_r(x, t) = \text{const}_1 - \text{const}_2 \times (ct - x)^{1/3}. \quad (3.21)$$

Consequently, the harmonic content follows the envelope obtained by the Fourier transform,

$$|E(\omega)|^2 \sim \left|\int (ct - x)^{1/3} e^{-i\omega t} dt\right|^2 \sim \frac{1}{\omega^{8/3}}. \quad (3.22)$$

Explain that since we are in the 1D model,  $E_r$  does not evolve from where it starts at the surface.

### 3.3.1 The ROM model

Apply Bourdier method and conservation of generalised momentum, we see that and thus non-linearity

and thus the thing we need to solve is blah.

The laser pulse leads to macroscopic oscillation of the plasma surface and therefore a corresponding oscillation of the ARP. One can understand this leading to a sub cycle Doppler shift in the reflected pulse. These pulses of radiation separated by the laser pulse frequency must therefore in the spectral domain consist of HH.

The Relativistic Oscillating Mirror (ROM) model describes an extension.

The following outlines the key ideas

Ok: assume I have the expression from Baeva's paper (do this tomorrow)

### 3.3.2 The normalisation factor

Baeva's theory provides us with the relative intensity of harmonics. However, for comparison with the absolute spectral intensity of harmonics in experiment, the normalisation factor is required. This can be calculated from conservation of energy. For arbitrary harmonic order scaling,  $n^{-p}$ , the spectral intensity of the harmonic beam is

$$I_\omega(\omega) = \frac{dE_\omega(J)}{dAd\omega} = I_0 \sum_{n=1, \text{ odd}}^{n_C} n^{-p} S_n \left( \frac{\omega}{\omega_L} - n \right) \quad (3.23)$$

up to the cutoff,  $n_C$ . Here  $S_n(\omega/\omega_L - n)$  is the spectral shape function of the  $n^{\text{th}}$  harmonic in reciprocal space and  $I_0$  is the normalisation factor of interest.

From conservation of energy,

$$\int I_\omega(\omega) d\omega dA = ER, \quad (3.24)$$

where  $E$  the total energy of the laser pulse and  $R$  is the reflectivity of the Relativistic Plasma Mirror (RPM). Intuition suggests the spectral shape function for the  $n^{\text{th}}$  harmonic retains the spectral shape of the incident laser pulse, *i.e.* for a laser pulse with a Gaussian temporal profile, this corresponds to a Gaussian centred at  $n\omega_L$

of width  $\sigma_L = 1/t_L$ , where  $t_L$  is the laser pulse width. Simulations show this is a reasonable approximation for ORION parameters <sup>1</sup>.

Simulations also show that for the SP1 laser, the incident laser pulse significantly suppresses the even harmonics, hence,

$$ER = \int_0^\infty I_0 \sum_{n=1, \text{ odd}}^{n_C} n^{-p} e^{-(\omega/\omega_L - n)^2/\sigma_L^2} d\omega. \quad (3.25)$$

The integral and summation order can be reversed. Since  $\sigma_L \ll \omega_L$ , all integrals in the summation are  $\approx \int_{-\infty}^\infty$  and therefore,

$$ER \approx I_0 \sum_{m=0}^{n_C/2-1} (1 + 2m)^{-p} \sqrt{\pi} \sigma_L \omega_L, \quad (3.26)$$

thus,

$$ER \approx I_0 \sqrt{\pi} \sigma_L \omega_L ((1 - 2^{-p}) \zeta(p) - 2^{-p} \zeta(p, \frac{n_C + 1}{2})), \quad (3.27)$$

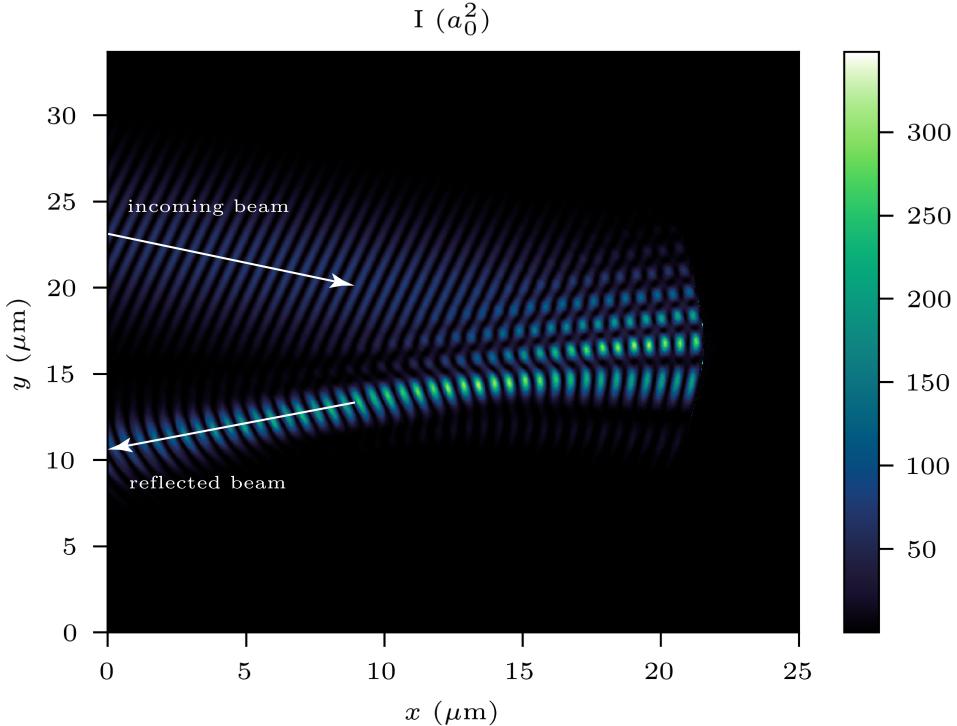
where  $\zeta(p)$  and  $\zeta(p, (n_C + 1)/2)$  are the Riemann Zeta and Hurwitz Zeta functions respectively. The final term can be neglected for a petawatt class laser pulse where  $n_C \gg 1$ . In the case of an ideal  $p$ -polarised laser pulse,

$$ER \approx I_0 \sqrt{\pi} \sigma_L \omega_L \zeta(p). \quad (3.28)$$

For the ROM regime,  $p = 8/3$  and the RPM is extremely efficient,  $R \approx 1$ . Thus,  $I_0$  can be estimated from the system parameters.

### 3.3.3 Hole boring

As observed in the ZVP plots, on long timescales relative to the laser pulse cycles, via the ponderomotive pressure of the laser, the plasma front moves inwards. This is hole boring [124]. One can derive a hole boring velocity by considering conservation of momentum in this quasi-static state. Since the hole boring velocity is laser pulse intensity dependent, the spatial profile of the laser will be imprinted on the surface. Typically Gaussian in shape, for high power laser systems, this can to first order generate a focusing RPM and a beaming of the specularly reflected signal, as in



**Figure 3.5:** Electromagnetic field intensity in a 2D PIC simulation of a relativistic ( $a_0 = 30$ ) laser pulse incident on a solid density plasma. The incoming beam is specularly reflected off the target which is curved by the radiation pressure leading to beaming in the reflected harmonic beam.

Figure 3.5. To access the highest possible electromagnetic field intensities, the laser pulse is focused on target to the diffraction limit. However, since the diffraction limit scales linearly with the wavelength, higher-order harmonics can be refocused to a smaller spot via this mechanism allowing access to unprecedented peak intensities. Vincenti *et al* demonstrated intensity gains of over 1000 with currently accessible parameters in 3D PIC simulations [165], suggesting a realistic route to the Schwinger Limit using next-generation laser facilities. Regardless of any blue-skies purposes, it is clear any prediction of HHG beam intensity must account for hole boring and is therefore an essential component of the ORION experiment analysis.

Applying momentum balance between the laser pulse and particles in the rest

---

<sup>1</sup>Hole boring discussed in the following section can lead to a Doppler shift in harmonic energy across the laser pulse corresponding to a Doppler broadening of the spectral shape function in reciprocal space, however, unlike the ZVP simulations of the previous chapter, this is negligible for ORION parameters.

frame of the RPM surface, the hole boring velocity is

$$\frac{v_{\text{HB}}}{c} = \sqrt{\frac{R \cos \theta}{2} \frac{Z m_e}{A m_p} \frac{n_c}{n_e(x_i(t, y))} a_L(t, y)} = \Pi a_L(t, y), \quad (3.29)$$

where  $R$  is the RPM reflectivity,  $\theta$  is the angle of incidence,  $16^\circ$  in the ORION experiment,  $Z$  and  $A$  are the atomic and atomic mass numbers respectively for the plasma ions,  $n_c$  is the plasma critical density,  $n_e(x_i(t, y))$  the electron number density and  $x_i(t, y)$  the depth of hole boring, from the Supplementary Material [122]. For the ORION laser pulse parameter space  $\Pi \ll a_L \forall t$ , *i.e.* we are in the relativistic electron and non-relativistic ion regime. Hence, the relativistic correction derived by Robinson *et al* [166] to Equation 3.29 can be neglected. Due to the high contrast and long duration of the ORION beamlines, there is minimal pre-plasma formation on the front surface and therefore the number density is simply the number density of the material in solid form and  $n_e$  is independent of  $x_i(t, y)$ . Robinson *et al* [167] generalised momentum conservation to multiple species, simply replace the mass density with the composite mass density  $\rho = \sum_j m_{ij} n_{ij}$ , then

$$\frac{An_e}{Z} \rightarrow \sum_j \frac{A_j n_{ej}}{Z_j}, \quad (3.30)$$

where  $n_{ej}$  is the number density of electrons that originated from the  $j$  ion.

The spatio-temporal envelope of the normalised vector potential of the laser pulse incident on the target surface is modelled as

$$a_L(t, y) = a_0 e^{-\frac{y^2}{2w_L^2}} g(t - t_0) \quad (3.31)$$

where  $w_L$  is the beam waist on target and  $g(t)$  the temporal envelope, a Gaussian or sech profile and  $t_0$  the main pulse peak time.

Integrating Equation 3.29,

$$x_i(y) = \int v_{\text{HB}} dt = \Pi \int_{-\infty}^t a_L(t, y) c dt. \quad (3.32)$$

At the peak of the main pulse,

$$x_i(y) = \Pi a_0 c e^{-\frac{y^2}{2w_L^2}} G, \quad (3.33)$$

where  $G = \int_{\infty}^{t_0} g(t - t_0) dt \sim t_L$  and  $t_L$  is the laser pulse temporal width.

The total denting is a combination of the peak electron-ion charge separation,  $x_e$  (which leads to the intrinsic phase of the HHG beam [168]) and Equation 3.32. Note that for the long pulse duration of the ORION laser,  $x_i \gg x_e$  and therefore  $x_e$  can be neglected.

Applying a Taylor expansion to the spatial profile of Equation 3.32 around the laser spot centre,

$$x_i = \text{constant} - \frac{y^2}{4f_p} + \mathcal{O}(y^4), \quad (3.34)$$

where, to first order, this is the equation of a parabolic mirror with focal length

$$f_p(t) = \frac{w_L^2}{4\Pi a_0 c G}. \quad (3.35)$$

Following the Vincenti *et al* derivation [122], the denting parameter is defined as,

$$\delta_T = x_i|_{(y=0)} - x_i|_{(y=\sqrt{2}\omega_L)}. \quad (3.36)$$

Hence,

$$\delta_T = \frac{w_L^2}{2f_p} = 2\Pi a_0 c G, \quad (3.37)$$

and is independent of laser focal spot size.

If the spatial profile of the  $n^{\text{th}}$  harmonic beam can be adequately described by a Gaussian at the plasma mirror plane, with a beam width described by the harmonic source size,  $w_n$ ,

$$h_n \sim e^{-r^2/w_n^2}, \quad (3.38)$$

then the beam profile is known at all distances,  $z$  from the target. Its divergence, defined as

$$\theta_n = \lim_{z \rightarrow \infty} \frac{w_n(z)}{z}, \quad (3.39)$$

is therefore

$$\theta_n = \theta_n^0 \sqrt{1 + \Psi_n^2}, \quad (3.40)$$

where  $\theta_n^0 = \lambda_n/\pi w_n$  is the harmonic divergence in the absence of RPM denting and

$$\Psi_n = \frac{2\pi}{\cos \theta} \left( \frac{w_n(0)}{w_L} \right)^2 \frac{\delta_T}{\lambda_n} \quad (3.41)$$

is the dimensionless focusing parameter. If  $\Psi_n \gg 1$ , as is true for the short wavelength X-ray harmonics of interest,

$$\theta_n \approx \frac{w_n(0)}{f_p \cos \theta} \quad (3.42)$$

and the divergence is dominated by RPM curvature.

Far from focus, at the detector plane  $z = 2.4$  m from the target,

$$w_n \approx z \tan \theta_n. \quad (3.43)$$

The corresponding magnification factor at detection is thus

$$\gamma_n(z) = \frac{w_n(z)}{w_n(0)}, \quad (3.44)$$

thus the laser intensity at detection is reduced by a factor  $\gamma_n(z)^{-2}$ .

At large distances,

$$\gamma_n \approx \frac{z \tan(w_n(0)/(f_p \cos \theta))}{w_n(0)} \quad (3.45)$$

Taking the Taylor expansion of the tangent, one sees that the magnification factor is only weakly dependent on the harmonic source size ( $f_p$  is independent of the source size), whereas the magnification is strongly dependent on the the laser spot size ( $\sim w_L^4$ ).

At the new RPM focal point,  $z = z_f$ , the demagnification factor is [122]

$$\gamma_n(z_f) = \frac{1}{\sqrt{1 + \Psi_n^2}}, \quad (3.46)$$

this determines the new peak intensity accessible via hole boring.

### 3.3.4 Harmonic source size

Include plots and fits for near roll off.

## 3.4 Simulations

### 3.4.1 1D PIC simulations of HHG

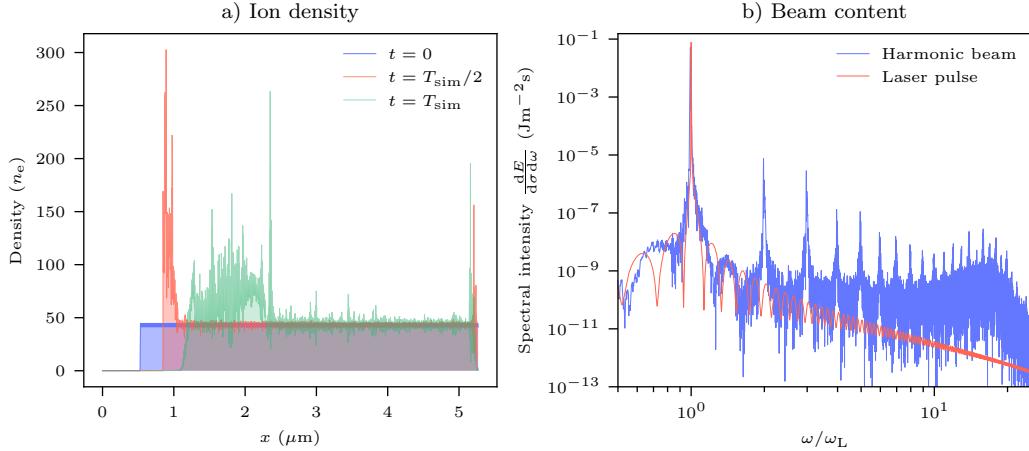
Numerical simulations of the SP1 and SP2 interactions with CVD and PMMA were performed using the Smilei PIC code to confirm the results of the previous section for ORION beamline parameters. Application of the Bourdier method (plasma initialised streaming at  $\mathbf{v} = c \sin \theta \hat{\mathbf{y}}$ ) enabled the accurate and computationally efficient reconstruction of the interactions in 1D, a necessity for the sub-ps ORION laser pulse duration. Despite the 1D geometry and optimised supercomputer architecture, resolution numerical heating and noise remained a serious issue. Edwards and Mikhailova noted for their single laser pulse cycle simulations that the highest resolved harmonic required 10 cells per wavelength and up to 5000 particles per cell was necessary for their noise requirements [126] while numerical heating scales linearly with time. The fourth interpolation order for the particle shape function was found to be an efficient choice for the minimisation of numerical heating when the Debye length is not resolved.

Fortunately, to make comparisons to the derivations of Section 3.3 only a low resolution is necessary. Simulation parameters are detailed in Table 3.3. Parameters are given in real units in the laboratory frame and normalised Smilei units in the boosted Bourdier frame. Note that polarisation angles and normalised vector potentials are unchanged by the frame transformation.

Figure 3.6 presents typical simulation results for the SP1 laser geometry, SP2 simulations produced similar results. Figure 3.6a illustrates how the ion density profile evolves via hole boring. While the density increases to several times its initial level near the plasma surface, this need not be accounted for in the hole boring calculation since these additional ions have already acquired the necessary momentum and are simply propagating inwards with the surface. Figure 3.6b clearly shows the harmonic content of the specularly reflected beam with harmonics retaining the spectral shape of the incident beam. Even harmonics are significantly suppressed relative to odd harmonics. The bump in the distribution around the 18<sup>th</sup>

SP1 Laser (1D, 16° angle of incidence)		
Parameters	Real	Smilei
Wavelength, $\lambda$ (nm)	527	$2\pi$
Angular frequency, $\omega_L$ (fs $^{-1}$ )	3.58	1
Polarisation, $\phi$ (°)	-42.2	-42.2
Normalised vector potential, $a_0$	7 - 10	7 - 10
Temporal FWHM, $t_{FW}$ (fs)	470	1690
Temporal envelope, $E_t$	$E_t \sim e^{-(t-t_{FW})^2 \ln 2/t_{FW}^2}$	
SP2 Laser (1D, 16° angle of incidence)		
Parameters	Real	Smilei
Wavelength, $\lambda$ (nm)	1053	$2\pi$
Angular frequency, $\omega_L$ (fs $^{-1}$ )	1.79	1
Polarisation, $\phi$ (°)	19.6	19.6
Normalised vector potential, $a_0$	20 - 30	20 - 30
Temporal FWHM, $t_{FW}$ (fs)	470	843
Temporal envelope, $E_t$	$E_t \sim e^{-(t-t_{FW})^2 \ln 2/t_{FW}^2}$	
Simulation box		
Parameters	Real	Smilei
Size, $x$ (nm)	$10\lambda$	$20\pi$
Sim length, $T_{sim}$ (fs)	$2t_{FW} + 5\lambda/c$	
Spatial resolution, $\Delta x$ (nm)	$\lambda/128$	
Temporal resolution, $\Delta t$ (as)	$\Delta x/2c$	
Macro-electrons per cell	1000	
Macro-ions per cell	1000	
Ion temperature, $T_i$ (keV)	0	0
Electron temperature, $T_e$ (keV)	0.115	$2.25 \times 10^{-4}$
Collisionless, pre-ionised, randomly-initialised plasma		
Parameters	Real	Smilei
CVD density, $n_e$ (m $^{-3}$ )	$1.05 \times 10^{24}$	$\begin{cases} 294 \text{ for SP1} \\ 1175 \text{ for SP2} \end{cases}$
PMMA density, $n_e$ (m $^{-3}$ )	$3.90 \times 10^{23}$	$\begin{cases} 109 \text{ for SP1} \\ 435 \text{ for SP2} \end{cases}$
Electron $x$ profile, $n(x)$	$n_e$ for $\lambda \leq x \leq 9\lambda$	

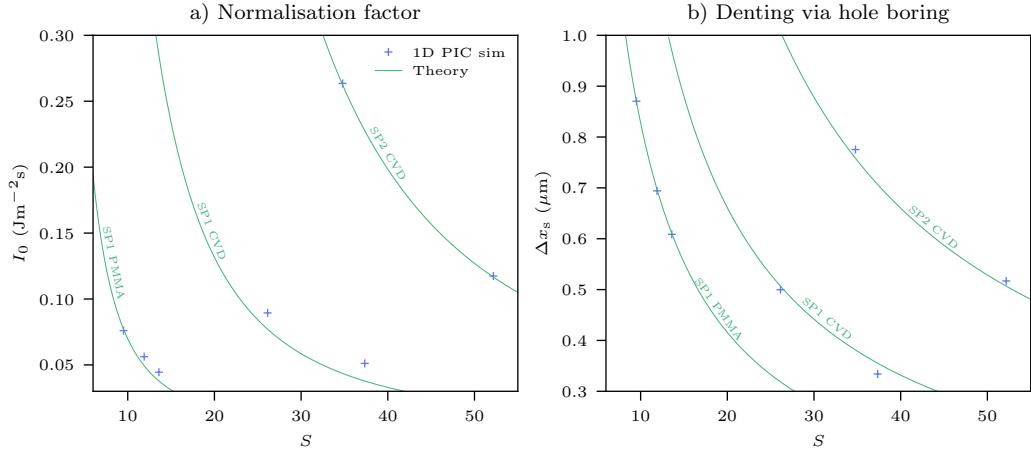
**Table 3.3: Smilei parameters to compare Section 3.3 analytical results to numerical simulation for ORION parameter space.** Real parameters are taken in the laboratory frame with the given units, while Smilei parameters are in the boosted frame with normalised units. The CVD ion distribution consists of carbon ions with number density  $n/6$  while the PMMA ion distribution consists of carbon, oxygen and hydrogen ions with number densities  $DR14FTiB5pt/54.02nA54i8n/2024$  respectively, each with 1000 macroparticles per cell.



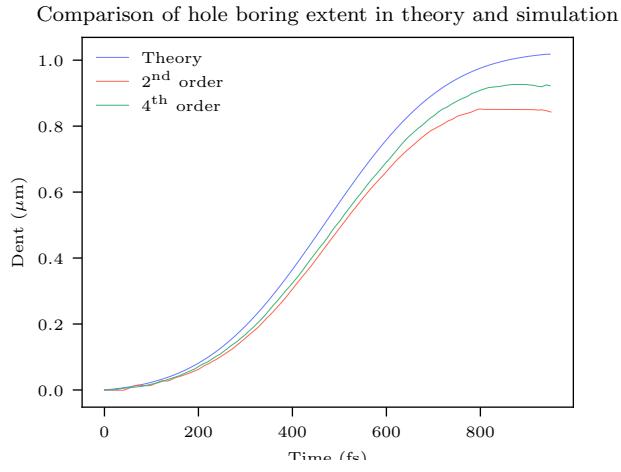
**Figure 3.6: 1D PIC simulation of the SP1 ORION beamline geometry at  $a_0 = 7$  interacting with a pre-ionised CVD target.** a) Ion density at initialisation ( $t = 0$ ), halfway ( $t = T_{\text{sim}}/2$ ) and at the simulation end ( $t = T_{\text{sim}}$ ) demonstrating hole boring. b) Absolute spectral intensity of the laser pulse and reflected harmonic beam obtained via Fourier transforms. The harmonic beam retains the laser pulse spectral width. Even harmonics are significantly suppressed at this polarisation. The large bump in the distribution around the 18<sup>th</sup> harmonic is spuriously generated via numerical heating.

harmonic is an entirely numerical effect, increasing the simulation resolution shifts the bump to the right. These simulations had a resolution of 128 cells per laser wavelength, following Edwards and Mikhailova one can only trust this distribution up to the 12<sup>th</sup> harmonic, at which point the distribution begins to change its shape. Clearly energy conservation does not hold for this simulation, however it is still possible to extract  $I_0$  if we make the reasonable assumption that Debye heating does not affect the HHG process and that the spectrum is simply the sum of the two contributions (justified by the observation of the continuation of harmonics over the bump in the distribution). Thus,  $I_0$  can simply be extracted as the peak intensity of the distribution corresponding to the intensity at  $\omega/\omega_L = 1$ ).

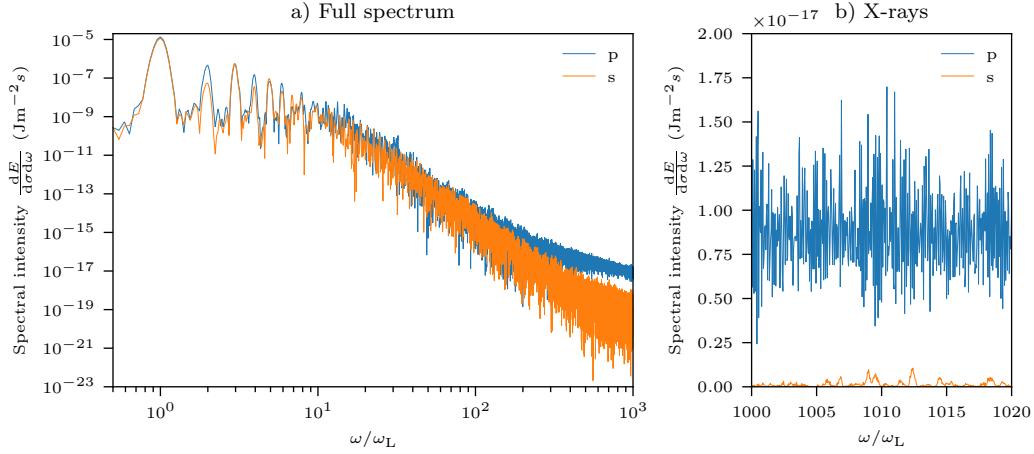
Figure 3.7 compares the analytically calculated parameters with those extracted from the simulations demonstrating excellent agreement for the parameter space accessible with the ORION laser facility and for the assumptions made in the analysis. The effect of numerical heating on the hole boring extent is evident in Figure X where the change in the surface position is plotted over time for the analytical calculation and PIC simulations with 2<sup>nd</sup> and 4<sup>th</sup> order shape function. The deviation from the



**Figure 3.7: Comparison between simulation and analytical predictions of hole boring and HHG normalisation factor in the ORION laser facility parameter space for CVD and PMMA targets.** a) The normalisation factor,  $I_0$ , where the spectral intensity,  $I = I_0 n^{-p}$  for low order harmonics. b) The depth of hole boring compared to the initialised plasma surface. Note that the steep density gradient as visible in Figure 3.6a ensures the surface position is well defined. Both parameters are expressed in the laboratory frame.



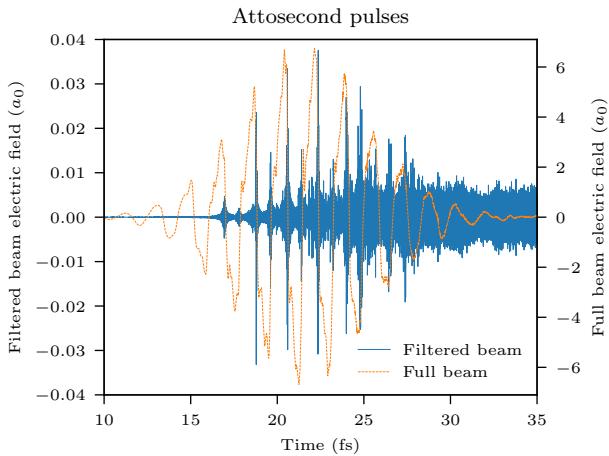
**Figure 3.8: The effect of the PIC code particle shape function on numerical heating demonstrated via the extent of hole boring.** The deviation from the theoretical prediction increases with time, the error decreases for increasing order of the shape function.



**Figure 3.9: Reflected beam harmonic content up to the keV range in a high resolution 1D PIC simulation of a test pulse with SP1 ORION beam geometry incident on a CVD target.** The absolute spectral intensity is calculated from the Fourier transform of the reflected beam electric field. The s- and p-polarised components of the beam are analysed separately for both a) the full spectrum and b) the X-ray harmonics accessible with the OHREX quartz ( $10\bar{1}1$ ) crystal.

theoretical prediction increases with time. Numerical heating leads to higher electron temperatures at the plasma surface and correspondingly a spurious preplasma expansion analogous to the prepulse preplasma expansion. This is why artificially increasing the plasma temperature to increase the Debye length and reduce numerical heating, while improving HHG accuracy, reduces hole boring accuracy since on sub-ps timescales, there is the potential for significant preplasma expansion.

The X-ray harmonic spectrum produced from a CVD target irradiated by an SP1 type laser pulse is resolved in Figure 3.9. The simulation has 16384 cells per laser wavelength, the minimum multiple of 2 required to resolve the harmonics accessed with the OHREX quartz crystals. It was necessary to use a short test laser pulse with a  $5\lambda_L$  FWHM. The simulation time before the peak of laser pulse was increased to reduce the error in the spectrum from unphysical hard cutoffs to the laser pulse profile. A steep preplasma scale length was applied to the plasma distribution of  $0.04\lambda_L$  to model the peak of the ORION pulse and suppress **CWE!** (**CWE!**). At the X-ray intensities of Figure 3.9b, no harmonic structure is visible, indeed for the non-optimal ORION target chamber geometry, harmonics begin to merge from around the 20<sup>th</sup>. Note the s-polarised part of the



**Figure 3.10: Attosecond pulse train from the filtered reflected laser pulse.** The filtered beam contains only harmonics above  $n = 50$ .

reflected beam is over 50 times weaker than the p-polarised beam at the 1000<sup>th</sup> harmonic. It would be interesting to extract the reflected beam temporal profile after filtering of sub X-ray harmonics. This can be achieved by applying the inverse Fourier transform to the Fourier transformed temporal profile with the lower order harmonics removed. Unfortunately, beyond  $n = 50$ , the contribution in the signal from noise prevented the structure from being observed. Figure 3.10 compares the reflected pulse structure to the reflected pulse structure with harmonics below  $n = 50$  removed. The filtered beam consists of a series of attosecond pulses of radiation, the brightest pulse is of duration 30.3 as. Note that the intensity of the first few laser cycles is insufficient for the generation of harmonics above  $n = 50$  and there is therefore no signal here, reducing the pulse train length.

### 3.4.2 Hydrodynamic simulations of preplasma formation

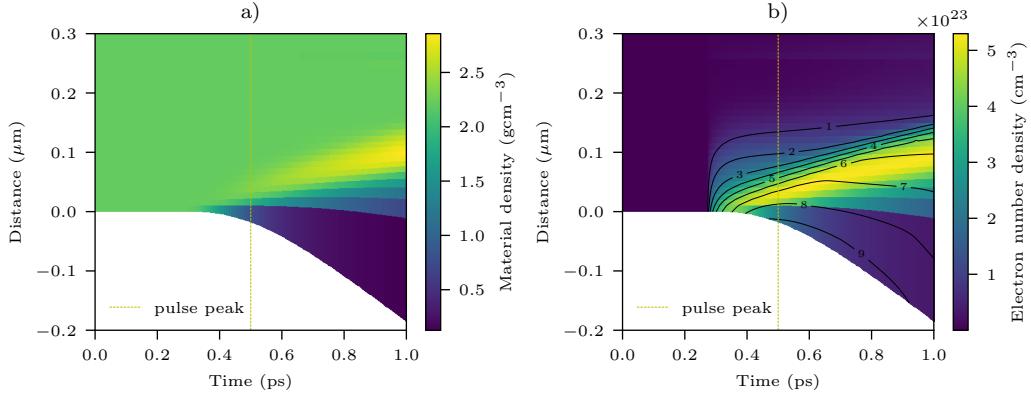
?? Following Dollar *et al* [169], hydrodynamic simulations of the ORION laser systems were performed using HYADES to confirm the suitability of the contrast conditions.

To do: calculate intensity on target and PM. Discuss how there is potentially a range of laser intensities thus had ‘high’ intensity and ‘low’ intensity sims (state

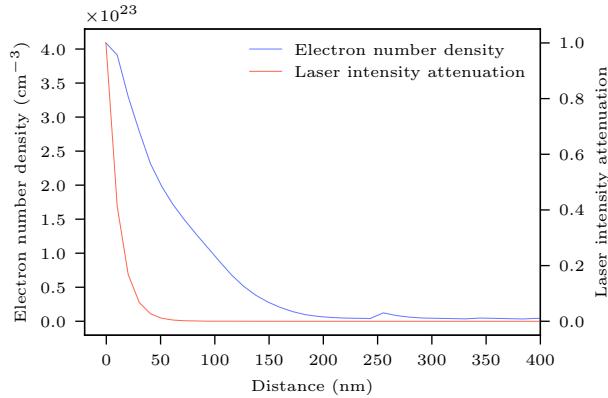
what these mean for the two beamlines.) All simulations were performed with s- and p-polarisation.

Parameters: 1D planar geometry Mesh of 511 evenly space points covering  $20\mu\text{m}$  of the target (SiO<sub>2</sub>, PMMA, CVD, etc) Equation of state data taken from the in built HYADES data tables. For low temperatures and pressures, the data for the minimum value in the tables is used, for temperatures and pressures above the highest value in the tables, the values are linearly extrapolated from the edge. The local thermal equilibrium average atom model is used for ionisation. At low temperatures ( $T < 0.15\text{eV}$ ) the absorption and refractive indices are input. The threshold for laser energy absorption is set to  $1 \times 10^6\text{ergs cm}^{-2} = 0.1\text{W cm}^{-2}$ . Electrons and ions are initialised at a temperature  $1.551 \times 10^{-2}\text{eV}$ .

Results: No pre-ionisation or preplasma formation was observed in the simulations of the high contrast SP1 laser pulse with any target materials. This is in stark contrast with the SP2 laser, in the absence of a PM, the preplasma expansion of the targets is huge. Necessitating the use of a PM and further HYADES simulations. First, the interaction between the SP2 and SiO<sub>2</sub> plasma mirror was simulated. Note that it is not possible to model the AR coating in hydrodynamic codes, instead it was assumed that the thin coating has minimal impact on the interaction. No ionisation was observed from the high intensity laser prepulse starting from 10 ns out from the main pulse. Figure 3.11 plots the interaction between a low intensity main pulse and the plasma mirror. Similar results were obtained for s- and p-polarisation. These simulations suggest ionisation begins at the start of the main pulse with some preplasma expansion but more significantly is the hole boring. Note that hole boring on the PM need not be accounted for since the spot size is significantly larger and therefore this change in surface curvature has minimal effect. The collisionless skin depth and plasma frequency can be calculated at each point in the plasma via the electron density and the corresponding laser attenuation at each point as in Figure 3.12 at the peak of the main pulse. The laser is rapidly attenuated and therefore ‘switched-on’ despite not attaining total ionisation at the front surface (reducing its reflectivity).

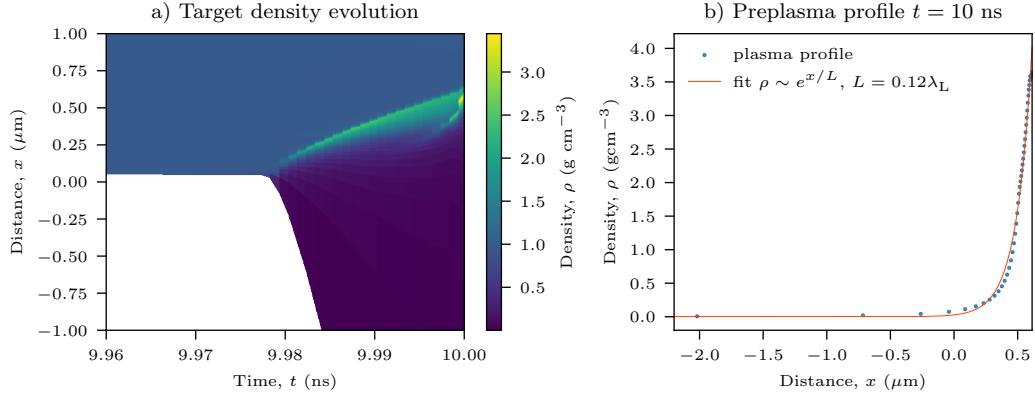


**Figure 3.11: Silicon dioxide plasma mirror switch on from irradiation by the SP2 main pulse at  $45^\circ$  angle of incidence and p-polarisation.** The laser pulse travels in the  $+\hat{x}$ -direction. The main pulse peak intensity is marked at 500 ps into the simulation. a) Material density as a function of time, a short preplasma is formed alongside hole boring. b) Electron number density as a function of time. The contours map the ionisation level.



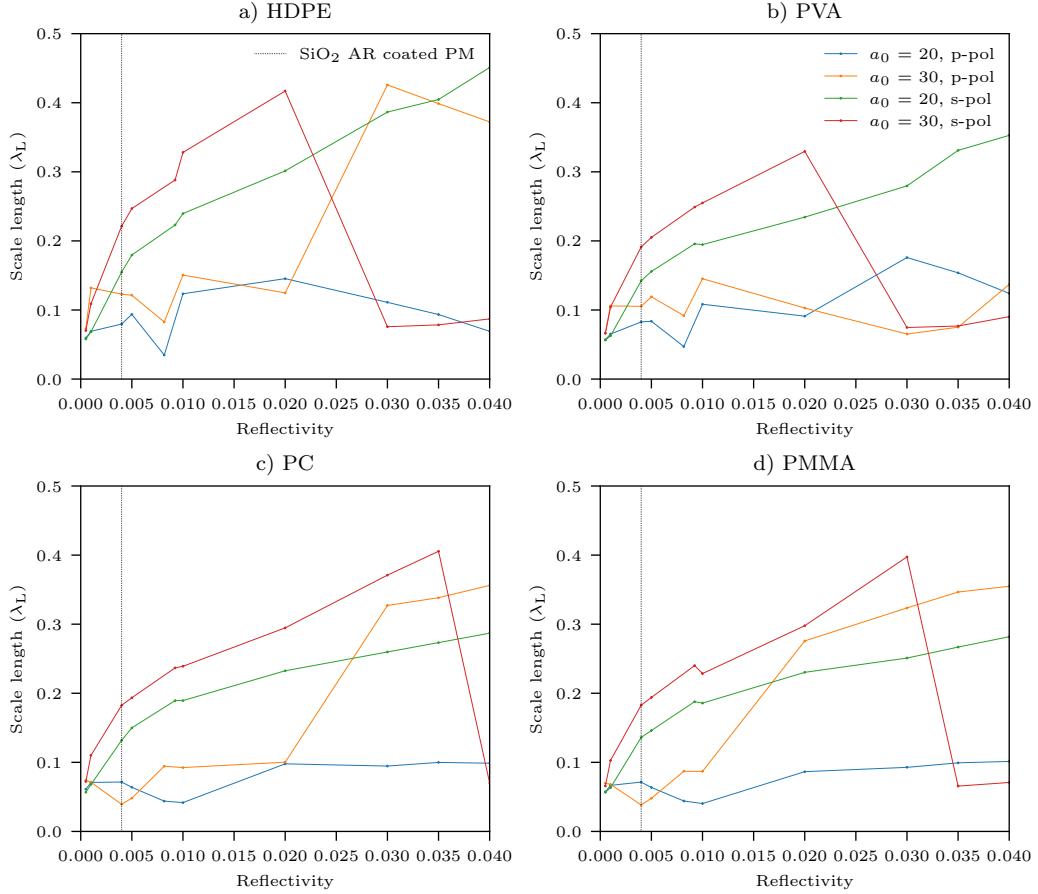
**Figure 3.12: The attenuation of the SP2 laser pulse calculated as it propagates through a switched-on PM.**

Satisfied that the PM operates as anticipated, a parameter scan of the targets was then conducted. Figure 3.13a describes the typical evolution of target density with the application of the SP2 prepulse, in this case after that prepulse has been attenuated by a plasma mirror with a reflectivity of 0.004. By the time of the main pulse arrival in this simulation, hole boring has overtaken preplasma expansion. Figure 3.13b demonstrates the exponential fit to obtain the corresponding close to optimal preplasma scale length of  $L = 0.12\lambda_L$  at the time of arrival of the SP2 main pulse.



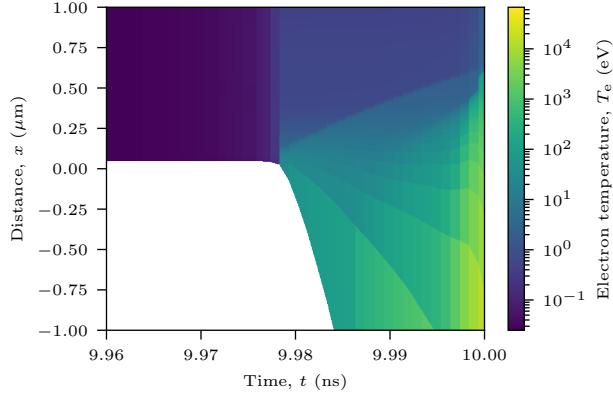
**Figure 3.13:** Typical preplasma formation from the p-polarised  $16^\circ$  angle of incidence of the ORION SP2 laser prepulse on an HDPE target. In this simulation the laser is modelled from 10 ns before the main pulse with  $a_0 = 30$  and attenuated by a PM of reflectivity 0.004. a) Typical preplasma expansion elucidated via target density. By the time of the main pulse, hole boring is dominant over plasma expansion. b) An exponential fit ( $\rho \sim e^{x/L}$ ) for the preplasma scale length,  $L$ , at the time of arrival of the main pulse. For this simulation the optimum fit is  $L = 0.12\lambda_L$ .

Figure 3.14 repeats the analysis of 3.13 for a variety of target materials, reflectivities, polarisations and laser intensities. The s-polarised light simulations follow a clear trend: increasing incident laser energy leads to an increased scale length up to a transition point that will be determined by the point at which hole boring overcomes plasma expansion. This will depend on the interplay of target ion density, ionisation potential and heating. Interestingly, the trend is much less consistent for p-polarised light. It is known that heating increases for p-polarised light as there is now a component of the laser pulse electric field acting into the plasma, however, the crossing of the high and low intensities suggests a dynamic balance between hole boring and plasma heating. Regardless, the change in scale length is not significant at these PM reflectivities. Unfortunately, it was not possible to conduct a similar parameter scan for the CVD targets, it would appear that the EOS tables for such a target material are not suited for this interaction type, the two tables offered provided inconsistent and unphysical results with specific simulations crashing at random. However, one can assume that the preplasma scale lengths of irradiated CVD targets would be lower than those established in Figure 3.14 due to the high damage threshold of the material.



**Figure 3.14: Preplasma scale length parameter scan for a variety of plastic targets.** Target materials are a) HDPE, b) PVA, c) PC and d) PMMA. The dashed line highlights the reflectivity of the  $\text{SiO}_2$  AR coated PM for the SP2 wavelength. Preplasma scale length remains small for the range of parameters explored.

It is interesting to return to the question of temperature raised with regards to the initialisation of PIC code main pulse simulations as presented in Chapter ???. Figure 3.15 plots the plasma electron temperature from 40 ps before the main pulse arrives. Once plasma expansion initiates the extent of heating rapidly diverges across the preplasma and bulk. Arguably it does not make sense to define a temperature for a system clearly not in equilibrium. Indeed, there is no standard practice and yet it can have significant effect such as has been seen in simulations on hole boring and **paper on prepulses and cooler being better**. Clearly, plasma temperature will be highly sensitive to the precise experimental conditions. At the time of the main pulse arrival at plasma surface (as defined by the peak of the electron density



**Figure 3.15:** The typical electron temperature of a plastic target after irradiation by a petawatt class laser prepulse. This is the same simulation as presented in Figure 3.13. Namely, an  $a_0 = 30$  p-polarised laser pulse attenuated by a plasma mirror of reflectivity 0.004 incident on an HDPE target at  $16^\circ$ . The bulk and preplasma temperatures differ by multiple orders of magnitude.

distribution), the temperature is 112 eV, consistent with the value chosen for PIC simulations. Note that hydrodynamic codes would be unsuitable for the prediction of the plasma temperature increase from the main pulse interaction given the highly non-linear dynamics in this regime could not be accurately captured by such a code.

To do tomorrow: final results of this section, presentation, paper? then aim to finish theory by end of the conference?? then its just the experimental layout which should be quick enough

## 3.5 Experimental data processing

### 3.5.1 Image plate calibration

Image plates (IPs) are reusable recording media that detect ionising radiation and are particularly suitable for the detection of X-rays produced in laser-plasma interactions. Their response is well understood and their sensitivities to a wide spectrum photon energies have been absolutely calibrated on the ORION facility [170]. Albeit for the FLA3000 scanner not the FLA7000 used in this experiment. However, the deviation in response is negligible for the photon energies measured. In this experiment the Fuji Biological Analysis System (BAS) TR-type IPs were used. They have a phosphor layer composed of  $\text{BaFBr}_{0.085}\text{I}_{0.15}$  with density  $2.61 \text{ g cm}^{-3}$

and thickness  $60\ \mu\text{m}$  but no mylar layer. This makes them suitable for low energy X-ray detection. When scanned, the IP releases blue photons via photostimulated luminescence (PSL), which is then collected by a photomultiplier tube. The PSL value is generalised across scanner types from the measured ‘Grey’ ( $G$ ) value by

$$\text{PSL} = (0.23284G^2 \times 10^{-9}) \left( \frac{\Delta x}{100} \right)^2 W \times 10^{-L/2}, \quad (3.47)$$

where  $\Delta x$  is the scanner resolution ( $= 25\ \mu\text{m}$  in this experiment),  $L$  is the latitude parameter, and

$$W = 0.092906 + 1370.8e^{-0.014874V} + 654.24e^{-0.011026}, \quad (3.48)$$

where  $V$  is the scanner voltage [171].

IP photon sensitivity,  $\psi$ , the number of PSLs per incident photon, is dependent on photon energy. Meadowcroft *et al* modelled this as,

$$\psi_j = \eta(m_j h\nu + c_j), \quad (3.49)$$

where  $h\nu$  is the photon energy and  $m_j$  and  $c_j$  are linear fit parameters valid for specific energy ranges,  $j$ . For the Fuji BAS TR-type IP and for X-rays in the range 0-6.0 keV,  $m_j = 0.54(5)\ \text{mPSL keV}^{-1}$  and  $c_j = 0.020(2)\ \text{mPSL}$ . The IP absorption efficiency in mPSL per photon is

$$\eta(h\nu, T_i, T_s) = \exp(-n_i \Phi_i(h\nu) T_i) [1 - \exp(-n_s \Phi_s(h\nu) T_s)], \quad (3.50)$$

where  $n$  is the layer density,  $\Phi(h\nu)$  is the total cross-section of the layer,  $T$  the effective layer thickness,  $s$  and  $i$  correspond to the sensitive (phosphor) and insensitive (mylar) layers of the IP respectively [172]. The first term is neglected in the absence of an insensitive (mylar) layer in TR-type IP. Below 50 keV, the dominant mode for X-ray absorption into the IP is the photo-electric effect, where

$$\Phi_{\text{ph}} \approx 3 \times 10^{12} \frac{Z^4}{(h\nu)^{3.5}} \quad (3.51)$$

and  $Z$  is the atomic number [173] and  $\Phi_{\text{ph}}$  is given in units of Barn per atom. At 2.4 keV, that corresponds to a sensitivity of 1.32 mPSL per incident photon.

It is generally inevitable that some time will elapse between laser shot and IP scan. For this experiment 30 minutes was typical, in which time some fading of the IP occurs that must be accounted for. IP fading can be modelled as an attenuation factor,

$$F(t) = A \exp(-t/\tau) + B, \quad (3.52)$$

where  $t$  is the time between shot and scan and  $A$ ,  $\tau$  and  $B$  are found from fits to experimental data. A key aspect of the exponential decay is that the attenuation depends only on the signal at that moment in time and not the initial conditions. This has been shown to be true in experiment [170].

At 20 °C at the ORION facility *Meadowcroft et al* [170] determined that for the Fuji BAS TR-type IP, the optimum fit for the parameters of Equation 3.52 is  $A = 0.347(22)$ ,  $B = 0.693(11)$  and  $\tau = 35.5(53)$  minutes. Therefore at 30 minutes,  $F(t) = 0.84$ .

In summary, the number of PSL measured on an IP can be converted to an incident number of photons via

$$N(h\nu) = \frac{\text{PSL}}{F(t)} \frac{10^3}{\psi(h\nu)} = P(h\nu)\text{PSL}. \quad (3.53)$$

### 3.5.2 OHREX calibration

The Orion High REsolution X-ray (OHREX) spectrometer, housed on the ORION laser target chamber outer wall, utilises a spherically bent crystal geometry to spatially focus and spectrally analyse photons from the target chamber [174] with a high signal-to-noise ratio. The measured signal has been absolutely calibrated for a range of energies using a variety of crystals [175]. The OHREX spectrometer can hold two crystals at a time. At each crystal's spatial focal plane a two-dimensional image is formed, one dimension is spatial, the other spectral. The energy range accessed by a given crystal is determined by the crystal rotation but all OHREX crystals are designed for operation at a nominal central Bragg angle of  $\theta_B = 51.3^\circ$  with the corresponding wavelength determined from Bragg's Law,  $n\lambda = 2d \sin \theta_B$ , for

Crystal	Range, $n = 1$ (eV)	Range, $n = 2$ (eV)
KAP (100)	585-625	1170-1245
Quartz (10̄10)	1830-1950	3660-3900
Quartz (10̄11)	2330-2480	4660-4960

**Table 3.4:** Photon energy ranges captured by the three lowest energy OHREX crystals when operating at their nominal central Bragg angle of 51.3° or first and second diffraction orders,  $n$ .

the appropriate crystal plane. The range around that central photon energy is determined by the crystal width in the spectral dimension.

MacDonald *et al* determined a quadratic fit for each crystal's dispersion relation to connect position along the image to photon energy [175]. Unfortunately in this experiment, the image lengths varied from those in the previous experiment, a likely consequence of slight defocusing of the optic. Note that the OHREX geometry is designed such that precise focus is not necessary to achieve good results [174].

Instead, a simple linear dispersion relation based on the known maximum and minimum energies accessed by the crystal was applied across the crystal images, a reasonable approximation to the dispersion relation determined by MacDonald *et al* [175] (the quadratic correction is small). The energy ranges for the three lowest energy OHREX crystals are given in table 3.4.

Provided full illumination of the 6 cm  $\times$  4 cm crystal, the spatial dimension can be safely integrated over to calculate the measured signal,  $M(h\nu)$  in J mm $^{-1}$  and remove uncertainty from the IP drifting from the ideal focal plane. (In this experiment we assume that the harmonic beam width at the crystal position is larger than the size of the crystal, a reasonable assumption since beam divergence  $\approx 10^\circ$  and the 6 cm  $\times$  4 cm crystal sits 2.4 m from the target.). This corresponds to a source spectral intensity incident on the crystal  $S(h\nu)$  measured in J keV $^{-1}$  sr $^{-1}$  via the spectrometer response,  $G(h\nu)$ , explicitly,

$$M(h\nu) = S(h\nu)G(h\nu). \quad (3.54)$$

Crystal	Order	Polarisation	<i>A</i>	<i>B</i>	<i>C</i>
KAP (100)	1	s	$1.72 \times 10^{-15}$	$-4.69 \times 10^{-12}$	$2.89 \times 10^{-9}$
		p	$1.40 \times 10^{-14}$	$-1.74 \times 10^{-11}$	$5.42 \times 10^{-9}$
	2	s	$3.64 \times 10^{-16}$	$-9.64 \times 10^{-13}$	$6.95 \times 10^{-10}$
		p	$5.03 \times 10^{-10}$	$8.09 \times 10^{-13}$	$5.03 \times 10^{-10}$
Quartz (10̄10)	1	s	...	...	...
		p	...	...	...
	2	s	$4.50 \times 10^{-15}$	$-3.40 \times 10^{-11}$	$6.52 \times 10^{-8}$
		p	$1.13 \times 10^{-15}$	$-8.86 \times 10^{-12}$	$1.73 \times 10^{-8}$
Quartz (10̄11)	1	s	$1.00 \times 10^{-16}$	$-1.74 \times 10^{-12}$	$4.93 \times 10^{-9}$
		p	$2.78 \times 10^{-15}$	$-1.41 \times 10^{-11}$	$1.79 \times 10^{-8}$
	2	s	$4.70 \times 10^{-16}$	$-4.50 \times 10^{-12}$	$1.11 \times 10^{-8}$
		p	$2.10 \times 10^{-16}$	$2.11 \times 10^{-12}$	$5.30 \times 10^{-9}$

**Table 3.5:** Sensitivity fit parameters as a function of photon energy,  $h\nu$  in electron-volts ( $G(h\nu) = A(h\nu)^2 + B(h\nu) + C$ ) for the three lowest energy OHREX crystals for p- and s-polarised incident photons and first and second diffraction orders [175]. Note that no data is available for the first order of the quartz (10̄10) crystal.

The absolute throughput of the crystals was measured by MacDonald *et al* in a previous ORION experiment and fit parameters for

$$G(h\nu) = A(h\nu)^2 + B(h\nu) + C, \quad (3.55)$$

where  $(h\nu)$  is the photon energy measured in eV, determined for both p- and s-polarised incident light and for first and second diffraction orders [175]. The parameters for the lowest few energy crystals are presents in Table 3.5. There is unfortunately no spectrometer response data for the 10̄10 crystal to first order due to the Si K edge sitting within the energy range and the dramatic effect this has on absorption in its vicinity [176].

The OHREX is equipped with a 50  $\mu\text{m}$  Beryllium filter to protect the crystals. The corresponding signal attenuation can be calculated using X-ray transmission data [177].

**Figure 3.16:** Unprocessed shot data from a FLA7000 scanned image plate converted to PSLs. The image plate and two crystal images are clearly visible.

### 3.5.3 Extracting the data

The quartz OHREX crystals  $10\bar{1}0$  and  $10\bar{1}1$  were fielded on the experiment. Crystal images were recorded with BasTR2040 Fuji Image Plate. A typical shot image scanned with the FLA7000 scanner and converted to photostimulated luminescence units (PSLs) is given in Figure 3.16. The average background signal was subtracted. The  $x$ - and  $y$ -axes were converted from pixels to mm using the scanner resolution, ( $25 \mu\text{m px}^{-1}$ ) and then energies using the appropriate dispersion relations. The data was then integrated over  $y$  to obtain the intensity in units of  $\text{PSL mm}^{-1}$  across each crystal image. Then the corresponding source signal is

$$S(h\nu)[\text{J keV}^{-1} \text{sr}^{-1}] = \frac{d\text{PSL}}{dx} \frac{P(h\nu)}{G(h\nu)} h\nu, \quad (3.56)$$

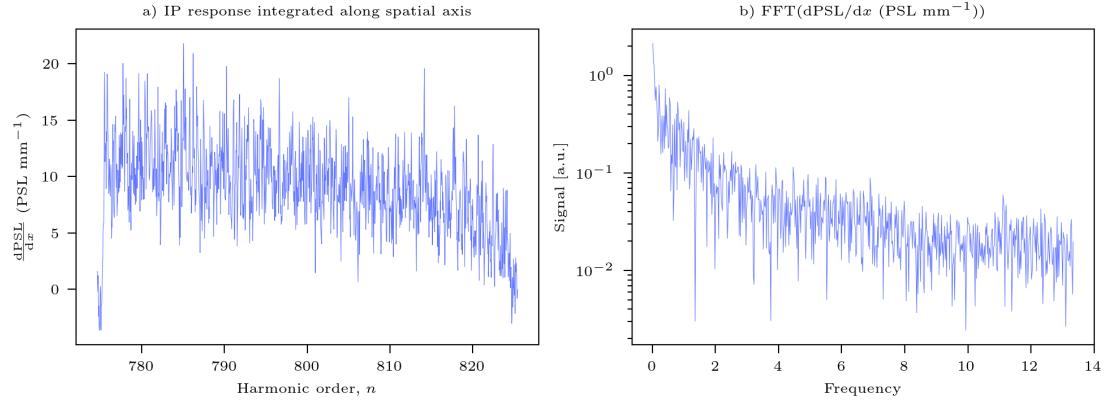
which can then be converted to a measured spectral intensity per harmonic at distance  $r = 1$  from the source, ready to be directly compared to the theory,

$$I_n^{\text{meas}}|_{(r=1 \text{ m})} = S(h\nu) \frac{dh\nu(\text{keV})}{dn}. \quad (3.57)$$

No sensitivity data is available for the  $10\bar{1}0$  quartz crystal, instead this lower energy crystal was fielded to attempt resolving of the X-ray harmonics. At the experiment planning stage it was unknown if this would be possible. Now that the simulations have been performed, we know that the non-optimal ORION target chamber geometry leads to merging of the harmonics even before the water window at 282 eV. This is consistent with the findings, Figure 3.17 is a typical integrated signal in  $\text{PSL mm}$  and the corresponding Fourier transform for the quartz ( $10\bar{1}0$ ) crystal image.

### Calibration and polarisation

The choice of the spectrometer response function,  $G(h\nu)$ , is non-trivial. One must firstly be assured that the second order contribution is small relative to the first, true



**Figure 3.17: Typical (SP1, PMMA) uncalibrated shot data for the quartz (10\bar{1}0) image** a) IP spatial axis integrated signal with dispersion axis. b) Fourier transform of a) with no evidence of harmonics.

for this source spectrum. And also think carefully about the anticipated polarisation in the OHREX interaction plane. The OHREX response to p-polarised light is approximately an order of magnitude lower than for s-polarised light. From Figure X we are assured that only the p-polarised (with respect to the target interaction plane) X-rays are specularly reflected to the OHREX. Corresponding polarisation out of the OHREX interaction plane is calculated in the Appendix. The polarisation of the HHG beam relative to the OHREX plane of incidence and reflection is 10.5° out of the plane. Unlike for the RPM interaction, the OHREX crystal reflection is an entirely linear process and it is therefore acceptable to decompose the laser pulse into its constituents, explicitly, the field incident on the crystal is

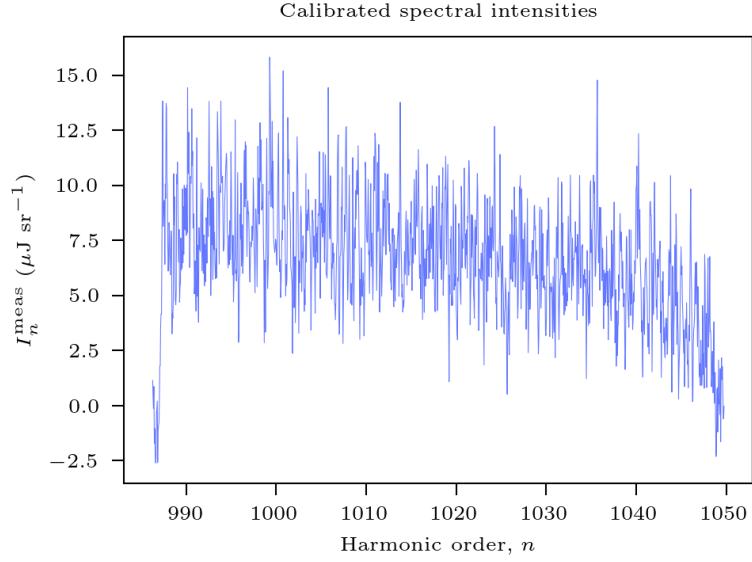
$$\mathbf{E}_O = \mathbf{E}_{O,s} + \mathbf{E}_{O,p}. \quad (3.58)$$

After interaction with the crystal the field is

$$\mathbf{E}_{\text{detector}} = \alpha_s(h\nu)\mathbf{E}_{O,s} + \alpha_p(h\nu)\mathbf{E}_{O,p}, \quad (3.59)$$

where  $\alpha_i(h\nu)$  is the energy dependent ( $h\nu$ ) amplitude sensitivity of the reflection for s- and p-polarised respectively. Since the two polarisations are orthogonal, the intensity is

$$I = \alpha_s^2(h\nu)|\mathbf{E}_{O,s}|^2 + \alpha_p^2(h\nu)|\mathbf{E}_{O,p}|^2. \quad (3.60)$$



**Figure 3.18:** Typical (SP1, PMMA) ORION experiment calibrated IP response for the quartz (10̄11) crystal.

Noting that  $\alpha_i^2(h\nu)$  are the calibration factors,  $G_i(h\nu)$ , and that

$$|\mathbf{E}_{\text{O,s}}| = |\mathbf{E}_{\text{O}}| \sin \phi \quad (3.61)$$

and

$$|\mathbf{E}_{\text{O,p}}| = |\mathbf{E}_{\text{O}}| \cos \phi, \quad (3.62)$$

where  $\phi$  is the angle out of the interaction plane,

$$I_{\text{detector}} = (G_s(h\nu) \sin^2 \phi + G_p(h\nu) \cos^2 \phi) |\mathbf{E}_{\text{O}}|^2 = F(h\nu) |\mathbf{E}_{\text{O}}|^2, \quad (3.63)$$

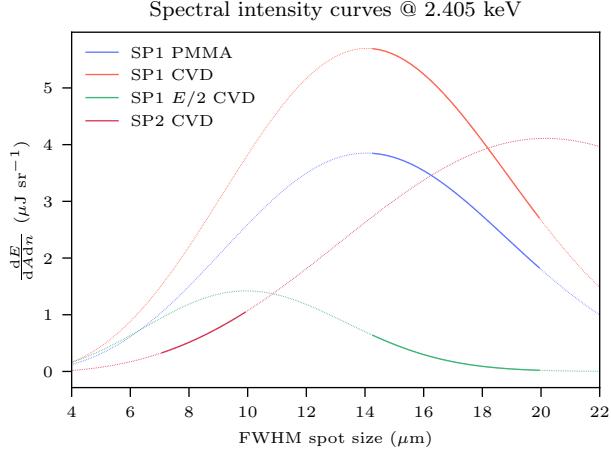
where  $F(h\nu) = (G_s(h\nu) \sin^2 \phi + G_p(h\nu) \cos^2 \phi)$  is the energy dependent calibration factor for this OHREX orientation.

## 3.6 Experimental results

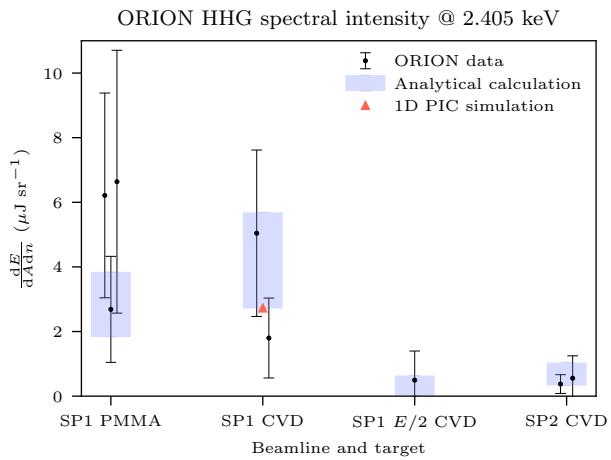
Figure 3.18 is a typical calibrated signal from the quartz (10̄11) crystal, ready for comparison to the theoretical prediction. For each shot the mean signal across the dispersion axis was calculated and the errors from statistics, the fade time (11 %), the IP sensitivity (1.5 %), the dispersion axis (2.9 %) and the OHREX crystal reflectivity (11 %) accounted for.

As demonstrated in Section 3.3, the intensity of X-ray harmonics is highly sensitive to the intensity of individual pulse cycles due to the proximity of the measurement to the exponential roll off in the spectrum. However, even neglecting the shot-to-shot variation, it is not possible to know the precise sub-cycle spatio-temporal struture of a petawatt-class laser pulse. Indeed, the simple measurement of the peak intensity of such a pulse remains an open problem [178, 179]. Instead we must rely on our assumptions and approximations, that is that the SP1 and SP2 beamlines can be adequately described by spatial Gaussians and temporal  $\text{sech}^2$  and  $\text{sech}$  profiles respectively, and that the duration and energy are well defined. The tight focused laser spot is subject to jitter on a shot-to-shot basis. This was routinely monitored via a KBRXM diagnostic that measures X-ray emission from the target. The spot measured by this diagnostic will always be larger than the actual laser pulse focal spot due to the energy transport of electrons moving away from the interaction area. This upper limit on laser pulse spot size corresponds to a lower limit on laser pulse intensity and therefore also of HHG efficiency. Figure 3.19 plots HHG spectral intensity curves for the quartz (10̄11) crystal central energy as a function of laser spot size for the targets and laser pulse configurations explored in the experiment. The thick lines range from the uppper limits of spot size as measured by the KBRXM diagnostic to the spot area being halved. The spectral intensity of the harmonic beam from the SP1 beamline sits in a delicate balance between beam divergence via hole boring and the exponential roll off. Note that only the roll off leads to a reduction in the total energy contained within the harmonic beam. The strong non-linearity of the interaction leads to the single beamlet SP1 beamline producing a harmonic beam over 6 times weaker than the double beamlet configuration. Interestingly, the higher intensity of the SP2 beamline leads to a greater divergence and a lower spectral intensity of the harmonic beam than for the SP1 laser pulse.

Figure 3.20 plots the experimental results with comparison to theory and simulation. Experimental data points are consistent within their errors with one another, the theory and the simulation, albeit with large uncertainties with such a low shot number. It is highly significant that the SP2 beamline is on average



**Figure 3.19: Dependence of the harmonic beam spectral intensity on laser spot size at 2.405 keV.** Targets and beamlines relevant to the experiment are plotted. The half energy single beamlet SP1 laser is demarcated by  $E/2$ . The thick lines indicate likely values for the laser spot sizes in this experiment. The SP1 harmonic beams are strongly affected by the exponential roll off unlike the SP2 beamline.



**Figure 3.20: X-ray harmonic spectral intensities at 2.405 keV measured on the ORION experiment compared to theory and simulation.** The analytical calculation corresponds to the thick lines marked in Figure 3.19. The 1D PIC simulation result is derived from Figure 3.9 and scaled to the ORION laser pulse duration.

almost 6 times brighter than the SP1 beamline despite the greater energy content and greater component of the laser pulse electric field directed into the plasma for the SP2 beamline while the single beamlet shot is over 7 times weaker in intensity compared to the mean intensity of the full SP2 beamline. These highly non-linear (with respect to energy) scalings are consistent with HHG theory but cannot be understood as originating from bremsstrahlung lending huge credibility to these results. Bremsstrahlung emission scales sub-linearly with temperature and temperature scales sub-linearly with energy on target. The total energy is well-constrained and this is far more relevant for Bremsstrahlung than the uncertain intensity (dependent on the variable spot size). Thus, at best case scenario one could anticipate twice the intensity being observed for the double beamlet SP2 compared to the single.

Bremsstrahlung emission from steep density profile low-Z solid density targets irradiated by relativistic laser pulses remains an area of active research with experiments to explore this in the planning stages at the University of York. However, provided a few assumptions it is possible to derive approximate best case scenario scalings of spectral intensity with incident energy. First, assume that energy absorption into the plasma target does not increase with increasing energy. From Equation the plasma temperature scales linearly with the incident energy. Now, assuming the target can be treated as a radiating black body, Equation increases most rapidly with temperature when  $h\nu \ll k_B T$  at which point the spectral radiance scales linearly with temperature. Thus, at best case we could expect a linear increase in the spectral intensity measured from the double beamlet compared to the single if that signal originated from Bremsstrahlung emission. Note further that this scaling depends not on the incident intensity but only on the incident energy which is a well-constrained quantity on ORION, *i.e.* we do not anticipate significant shot-to-shot variation.

The above discussion relied on the assumptions of equilibrium and a black-body radiation spectrum. While this is valid for the long radiating period after the interaction, in much of this thesis, emphasis has been placed on the non-equilibrium

state of the interaction itself. There are two species of relevance during this time: the hot electrons around the front surface directly heated by the laser pulse and prepulse, and the hot electrons that are reflected specularly and are responsible for the HHG. In [180], bremsstrahlung of these species are measured experimentally for a 30 fs laser pulse and  $a_0 = 2.8$ . As the populations of each of these species are so small, the peak contribution to the photon spectrum is many orders of magnitude smaller than that of the experimental signal and can therefore be neglected.

These findings are highly significant. Then write up the new peak intensity and the efficiencies. Comment on flying focus



# 4

## Attosecond electron bunches and X-ray pulses on GEMINI PW

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### 4.1 Overview

We successfully measured the intensity of the attosecond duration X-ray harmonics on the ORION SP1 and SP2 beamlines. However, the uncertainties remained large, the geometry sub-optimal and the pulse train exceedingly far from an isolated attosecond pulse. Indeed none of the proposed mechanisms for isolation (polarisation gating, lighthouse technique, etc.) could have any hope of success. The high shot rate, few femtosecond, highly customisable GEMINI PW facility at CLF can

address all these points while also providing conditions suitable for the production of attosecond ZVP electron bunches, of course attended by its own suite of technical challenges. The work of the prior chapters of this thesis formed the body of a successful proposal for five weeks of beam time at the facility with shot dates planned for Summer 2024. This chapter outlines the extensive planning process initiated over a year before the experiment itself.

There are three primary experimental goals. First, to repeat the ORION experiment on GEMINI PW with optimised geometry and thus resolve and measure the absolute intensity of X-ray harmonics, detailed in Section X. Second, to simultaneously observe ZVP electron bunches in transmission and the HHG signal they generate in reflection, detailed in Section X. Finally, to apply the X-ray harmonic beam in a proof of principle Laue diffraction setting described in Section X. Section X addresses the subject of laser contrast, a much thornier issue for the few femtosecond GEMINI beamline compared to ORION and covers the parallel experiment proposed by collaborators at Queen's University Belfast to provide some illumination to the problem at hand. Note that no petawatt class few femtosecond laser facility has observed harmonics at oblique incidence. This section will also detail the strategy of contingency should the contrast prove to be too poor for the production of harmonics. Not discussed in this chapter is the work undertaken by Elliot Denis to perform spectro-spatial encoding and thus temporally resolve chirped HHG.

CAD drawings overview where to put you?

The experimental geometry is relatively simple. Post **DPM!** (**DPM!**) contrast enhancement, the GEMINI PW South (S) beam is focused by an  $f/2$  parabola onto the solid density target at  $45^\circ$  **AOI!** and p-polarisation, directing the specularly reflected X-ray beam to the West side of the chamber. XUV and X-ray spectrometers aligned to the target in the specular direction will observe the reflected harmonic beam.

## 4.2 X-ray harmonics on GEMINI PW

Naturally, the first action is to repeat the ORION experiment at GEMINI PW. For this purpose, the Oxford Engineering Department is designing and building a replica OHREX spectrometer to perform the measurement with OHREX crystals borrowed from AWE. The lower energy KAP (100) OHREX crystal listed in Table 3.4 is more suited to GEMINI PW due to the lower energy of the beamline compared to ORION. Regardless, the quartz crystals will also be fielded.

A 1D PIC simulation of the GEMINI PW beamline for the planned geometry predicts harmonics will be resolved for the KAP energy range. [INSERT PIC SIMS HERE].

The technical complexity of the OHREX spectrometer is entirely attached to the spherically bent OHREX crystals, thus reducing the challenges of alignment [174]. It is designed to be relatively insensitive to the distances between source and crystal (2.4 m) and from crystal to image plane (0.524 m). Indeed, previous experiments deemed it unnecessary to adjust the OHREX bellows to access the best focus. Variation in the angle of incidence from  $38.7^\circ$  shifts the spectral image away from the nominal energy range. The alignment of the crystal planes to the crystal surface enables the relative convenience of optical alignment. Initial shots will use IP to check for any light leakage and to confirm alignment. Switching to the Raptor Photonics Eagle XV in vacuum X-ray CCD camera (EA4240XV-BN-CL) [181] to utilise the high shot rate on GEMINI PW. The CCD camera must be liquid cooled, thus for the prevention of damaging condensation forming on the camera window a manual gate valve is required to isolate the OHREX when pumping the main chamber (waiting for the camera to reach acceptable temperatures while under vacuum would be prohibitively long). With  $2048 \times 2048$  active pixels of size  $13.5 \mu\text{m} \times 13.5 \mu\text{m}$ , not all of the 4 cm crystal image can be captured by the camera.

A parameter scan of **XHHG!** (**XHHG!**) as a function of  $S$  can be performed by varying the laser intensity, via beam apodisation, and the target material. The majority of shots will use target wheels of the relatively high damage threshold fused silica with the plan to advance to a kapton tape. The fused silica targets are

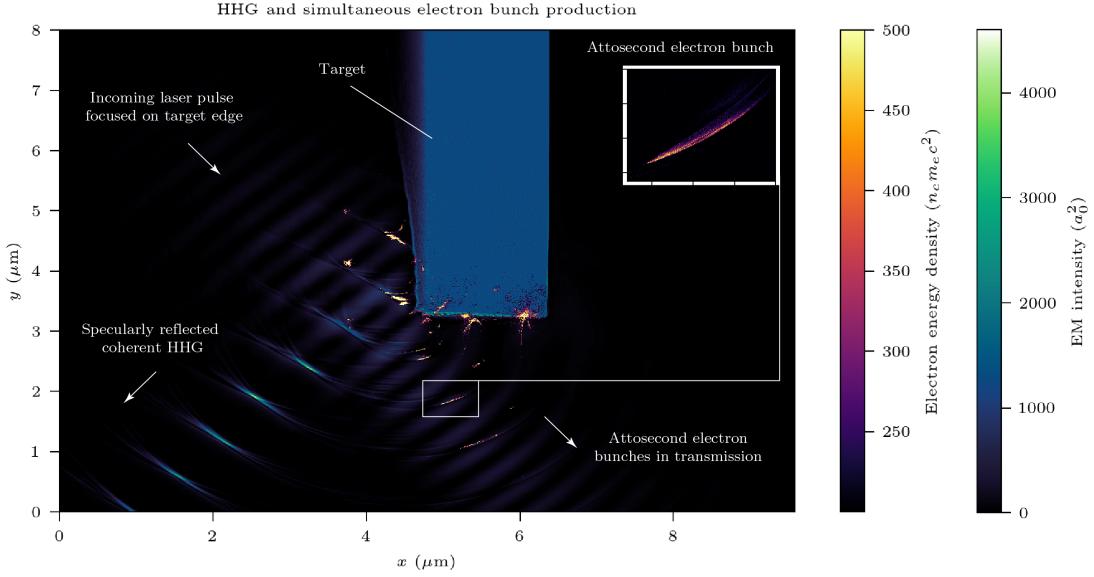
chemically etched in one corner to aide alignment. Since HHG is suppressed for circularly polarised light, bremmstrahlung emission and/or other X-ray production mechanisms can somewhat be ruled out by comparing the **XHHG!** signal observed for p- and circularly- polarised light, produced via a quarter wave plate. However, this is not fully the case at 45° **AOI!** and plasma heating is reduced, reducing the temperature of the emitting plasma.

Filtering is necessary to reject the high intensity low order harmonics, however, for photon energies in the 600 eV range, the standard beryllium filter transmission is too low. Instead 400 nm of aluminium flash-coated onto 1 micron of mylar will be used for the KAP crystal, corresponding to a transmission of 16 % at 600 eV.

It is essential to check there is suitable resolution in both the spectrometer and the CCD camera for harmonic observation. For the highest energy crystal, nominal photon energy of 2.405 keV, 96 harmonics sit in the range accessed by the crystal, corresponding to a fractional energy  $\Delta E/E = 6.45 \times 10^{-4}$  between harmonics. At  $\Delta E/E = 1 \times 10^{-4}$ , the OHREX crystal resolution is sufficient, however it may not resolve the harmonics' spectral shape. Those 96 harmonics are spread over the 4 cm crystal image. At a resolution of  $13.5 \mu\text{m}$ , the CCD camera has 31 pixels between harmonics and the harmonic width is about a pixel.

sims: Running new SiO<sub>2</sub> HHG sim, hopefully this will show that we have a better efficiency than the previous simulation. Then pop that in here.

A previous experiment measured the GEMINI DPM setup produced a reduction in peak pulse intensity of 50 %. Applying ROM and hole boring theory to a fused silica target and accounting for filtering and assuming no merging of harmonics, one can anticipate signals at the OHREX crystal position of  $19.6 \mu\text{J sr}^{-1}$  per harmonic at 0.6 keV for the KAP crystal and  $1.89 \mu\text{J sr}^{-1}$  per harmonic at 2.405 keV for the quartz (10̄11) crystal. Orienting the OHREX for s-polarisation to maximise the signal at the image plane, this corresponds to an average of 3.4 photons per pixel and 0.20 photons per pixel respectively. The quantum efficiency of the detector is close to 100 %. These numbers are no smaller than the ORION experiment while



**Figure 4.1:** A simulation of the planned GEMINI-PW experimental setup for the measurement of ZVP electron bunches. This novel setup enables the simultaneous measurement of attosecond ZVP electron bunches and their coherent emission of X-ray light. The GEMINI-PW laser pulse is incident at  $45^\circ$  on the low density polyethylene target with a preplasma scale length of  $0.2\lambda_L$ . For this angle of incidence, transmitted bunches and specularly reflected X-ray harmonics are produced at a frequency of  $\omega_L$ .

the background will be lower due to the order of magnitude lower energy on target. Thus, there should be adequate statistics to observe these signals.

### 4.3 First observation of ZVP electron bunches

The second phase of the GEMINI PW experiment is to observe ZVP electron bunches (both mass-limited and bulk) for the first time. There has been some experimental evidence of attosecond and nano-Coulomb electron bunch generation from nano-targets [110, 114] and in reflection from solids [113, 115] with the latter demonstrating satisfaction of the precise conditions for **VLA!** (**VLA!**) to 10 MeV for a mildly relativistic laser pulse ( $a_0 \approx 3.1$ ). These experiments have guided the planning process for this experiment stage.

This experiment aspires to produce mass-limited ZVP electron bunches in the configuration simulated in Figure 4.1. By focusing the laser pulse onto the edge of a transversely mass-limited target, the emitted electron bunch energies will be

maximised. Simulations suggest it should be possible to simultaneously measure the specularly reflected HHG and the attosecond electron bunches that produce it. Note via equations 2.47 and 2.26, the low density target of Figure 4.1 will produce larger and more energetic bunches. They are therefore a more practical choice for the experiment relative to the aluminium targets of predominant use in Chapter 2. Both bulk and mass-limited ZVP electron bunches will be generated but should be spatially separated at the observation point via conservation of transverse momentum with bulk bunches experiencing no further acceleration phases post ZVP.

It is necessary to conduct PIC code parameter scans for this new geometry. While normal incidence was most convenient for the initial ZVP simulations, oblique incidence is more optimal for HHG [120, 126] and as can be understood from equations 2.43 and 2.44, the new energy scaling expressions, for the ZVP mechanism. Not only is oblique preferable but it is essential to mitigate damage to the laser optics via back-reflection. The HHG beam intensity at focus can be over 1000 times that of the incident laser pulse [182]. It is indispensable, therefore, to test the new predictions for oblique incidence energy scalings and total electron bunch charge along with the angle of bunch ejection, the non-zero transverse vector potential of the laser will prevent the bunch from propagating directly along the transmission axis. It would also be useful to perform a parameter scan of preplasma scale length. In this work so far, it was assumed that the optima for electron bunch production are simply those for HHG, as was shown to be true for preplasma scale length in the Supplementary Information of [115] for their electron bunches in reflection.

For target edge shots, high edge precision is absolutely essential. To overcome this challenge CLF Target Fabrication have created  $300\ \mu\text{m}$  silicon wafer targets via the Bosch process with well-defined edges of sub-micron precision. Target design is given in Figure ??a. The tapered finger design adds a failure point to the target to mitigate damage to adjacent targets via shock propagation. The cone of the array holder is of sufficient size for the full laser pulse to access the target edge. Scans, detailed in Figure 4.2 found an average surface roughness of 74.4 nm. The etching technique leads to scalloping of the surface parallel to

**Figure 4.2: Targets designed for observation of ZVP electron bunches.** a) Target array design with conical array holder and tapered finger design, lengths in mm. b) Scan of etched wafer, striations develop towards target rear due to etching technique. c) A 3D plot of surface variation. Scalloping and striations are visible but too small to interfere with laser pulse transmission at 45°. d) Lineouts of c), the  $y$ -profile is along the target edge of interest.

the target surface and striations perpendicular to the surface towards the target rear. Neither pose any issue with regards to interfering with the laser pulse in transmission. Scans were performed and targets designed by Sam Astbury [183]. Before firing target edges, HHG production will be optimised for unstructured silicon wafer targets, providing information on bulk propagating ZVP electron bunches. Again, a quarter-wave plate can be implemented to suppress the ZVP mechanism in the absence of vector potential zeroes.

It is anticipated there is some relaxation of the requirement for sub-micron precision from target smoothing via preplasma expansion and the smoothing effect of the main pulse interaction as noted by Dromey *et al* [159]. Furthermore, transmission direction is dictated by laser vector potential and not surface angle, unlike reflection. At worst ZVP energy scalings will be mildly affected by the change in angle. Thus tape drives will also be explored as an option. However, the  $50\text{ }\mu\text{m}$  jitter of such devices suggests a success rate of  $\sim 4\text{ \%}$ . This is not necessarily unreasonable. Shot-to-shot variations in laser pulse focal spot position are typically on the order of the focal spot itself. Therefore only a third of silicon target shots will be successful and will depend on the gathering of sufficient statistics as will be possible with the high repetition rate of the GEMINI PW laser facility. Successful shots will be identified using a third harmonic imaging line [184]. This can also be used to reduce uncertainty in the hole boring calculations since it is typically a much closer match to the real laser spot size than the X-ray emission spot since the non-linearity of the interaction reduces any background noise giving a clear signal. For high resolution, the collimating reflective optic, must be placed roughly the distance of the  $f/2$  parabola focal length from the main interaction. So close to the interaction, the B-integral is typically large and ND filters and lenses cannot

be used. Instead, a wedge, typically double passed acts as an attenuator. A hole allows the less divergent higher orders through, enabling on-shot measurement. Far from the interaction, the third harmonic signal can be extracted, filtered, refocused and imaged. Note that while some of the light is reflected back to the target via the wedge, this is not a concern since the focal length of the new optic will be larger than that of the  $f/2$  parabola and typical path lengths are about a nanometre at which time the target does not reflect specularly.

Resolving the  $16\text{ }\mu\text{m}$  GEMINI PW spot is significantly more challenging than the  $6\text{ }\mu\text{m}$  spot of the 1 J Astra-GEMINI laser system. This imaging line has been modelled by the ORION laser scientists using Zemax [185] accounting for all diffraction and spherical aberration. They have identified a set up using an **OAP!** (**OAP!**) for collimation and a spherical mirror for imaging that provides  $1.7\text{ }\mu\text{m}$  resolution with 20 times magnification, comfortably resolving the spot. However, this setup is highly sensitive to the **OAP!** position taking a significant time to align, ideally this should only be performed once. Note that while path lengths vary between atmospheric pressure and under vacuum, there is no hysteresis from cycling the chamber. While the high spatial orders associated with the clipping of the reflected beam at the target edge could not be resolved, this should significantly alter the image structure, enough to identify the surface edge.

The main concern of this method is that over time the wedge will be coated by material ablated from the main target interaction and will need to be replaced to prevent absorption increases and corresponding heat damage. The sensitivity to wedge position is being investigated.

The size of the wedge hole must be established. Sitting somewhere between the near and far field, to first order, rays from all parts of the spot are lost equally through the hole in the wedge. After passing through the optical system, therefore, the wedge hole should simply reduce the signal at the image plane and not affect the resolution but should also be modelled. Potentially, one could argue that the wedge can reduce the spherical aberration since only paraxial rays are removed. The larger the hole the less likely the clipping of the potentially highly divergence high harmonic

beam eventually this will impact the third harmonic image resolution. At the very least we must fully illuminate the OHREX crystal requiring a half angle of  $0.71^\circ$ .

Applying the extended ZVP theory to the oblique incidence case for GEMINI PW parameters predicts electron bunches with a charge of  $\approx 10$  nC and energy of  $\approx 30$  MeV. First a LANEX screen will be imaged to identify mass-limited and bulk electron bunch trajectories and beam divergences, anticipated to be significant. An MeV electron spectrometer aligned to the target edge can then measure the time-integrated energy spectrum. The spatially separated bunch types can be independently measured by adjusting the spectrometer position. Mordovanakis *et al* used Image Plate stacks to obtain the electron bunch structure and emission angle, it may be necessary to perform this if the bunch types overlap due to high divergence for accurate spectrometer positioning [186].

While resolving attosecond durations remains a serious technical challenge of experimental science, we hope to measure the production of Coherent Transition Radiation from a secondary target and therefore infer the presence of a bunch train [113].

COTR - awaiting papers from peter

When electrons pass from one medium to another they emit transition radiation, in this case the aluminium filter. If they are arranged into temporally regular spaced bunches that radiation becomes coherent and harmonics of the temporal spacing of the bunches can be observed. Both bulk propagating and mass-limited ZVP electron bunches will produce COTR, however, the attosecond duration should mean the mass-limited bunches produce significantly brighter COTR that extends to XUV energies but that can be distinguished even at the optical range.

## 4.4 Application of HHG to white light Laue diffraction

The ultimate goal is the generation of an attosecond X-ray diagnostic to illuminate crystal dynamics on such timescales. This remains a distant goal. Inspired by Laue diffraction measurements demonstrating dislocation microstructure of shocked

crystals [187], this will be a proof of principle test of Laue diffraction using an X-ray harmonic beam of a single unshocked crystal.

The first X-ray crystal diffraction experiment was Laue diffraction and it was Laue diffraction that led to the arrival of Bragg's law,

$$n\lambda = 2d \sin \theta \quad (4.1)$$

for diffraction order,  $n$ , incident wavelength,  $\lambda$ , distance between crystal planes,  $d$ , and angle between the crystal plane and the incident photon,  $\theta$  [**braggDiffractionWaves1915**]. In Laue diffraction, by irradiating a crystal target with a 'white light', *i.e.* a broadband X-ray source, Equation 4.1 is satisfied by a range of crystal planes and angles simultaneously, producing Laue spots accordingly. The quasi-broadband X-ray harmonic beam is suited to such an experiment if enough signal can be obtained.

This will be attempted using the **BBXRD!** (**BBXRD!**) placed in the path of the specularly reflected beam roughly 50 cm from the target. This lead shielded box contains the crystal sample with image plate lining the inner walls. A hole allows the harmonic beam into the box. Placed at an angle on a rotatable washer enables a range of angles and corresponding energies to be investigated. Starting with quartz to investigate the photon energies of 1.7 - 2.5 keV and moving to silicon to explore energies of 2.4 - 3.3 keV. The IP must be shielded to reject laser light, beryllium would be most suitable. It may be necessary to integrate over many shots to produce the Laue spot.

## 4.5 Contrast

### 4.5.1 Contrast intro

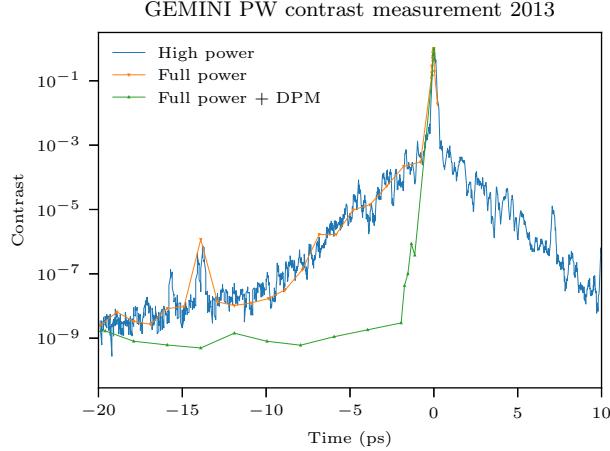
High power laser systems have inherent prepulses of ionising intensities. This preheats a target leading to thermal expansion prior to arrival of the main pulse, producing an exponential preplasma,  $n_e \sim e^{x/L}$ . HHG is exceedingly sensitive to this preplasma formation and has evoked particular interest in the field [165, 169, 188–190] with both simulation and experimental evidence. While there is some variation in optimal values, optima are typically small with respect to the laser

wavelength, with scale lengths,  $L \approx 0.1\lambda_L$ . The difficulty in measuring such small lengths perhaps going some of the way to explain the variety of suggested optima.

For very short preplasma scale lengths **CWE!** dominates over HHG. Kahaly *et al* demonstrated in experiment harmonic emission varied smoothly with reducing laser contrast from the CWE regime to the HHG regime to HHG suppression [190]. As can be understood increasing the preplasma scale length increases the amplitude of surface oscillations, increasing HHG production, however, there reaches point where parametric instabilities in the underdense region begin to dominate [169].

The use of 1D hydrodynamic codes to measure prepulses from laser contrast trace measurements has become standard practice [169, 188, 191], however this shall be called into question in the following.

Fortunately there are several approaches to control various aspects of the process of preplasma expansion. Plasma mirrors are effective tools for increasing laser contrast. These single use optics typically have an **AR!** coating of reflectivity  $R_1$ . As the laser fluence passes the damage threshold of the optic, plasma forms on the front surface. The laser is focused onto the optic such that only the main pulse will ionise the optic, leading to a high reflectivity  $R_2$  when ‘switched-on’. The **PM!** therefore ups the contrast by a factor  $R_2/R_1$ . There is a balancing act in choosing the peak intensity on the **PM!**. Too low and switch on occurs too late and  $R_2$  is reduced. Too early and the PM will switch on before the arrival of the main pulse. A peak intensity around  $1 \times 10^{16} \text{ W cm}^{-2}$  is optimum [192]. The heating effect of the prepulse can be reduced by reducing the angle of incidence or the prepulse itself can be reduced by reducing the intensity of the beam through apodisation or use of a frequency doubling crystal. A secondary beamline can generate a controllable prepulse [190], the surface can be recompressed using a circularly polarised prepulse that will cause minimal heating [191]. Alternatively, the intensity on PM can be reduced, again costing in the peak intensity of the main pulse interaction.



**Figure 4.3:** Native contrast of the GEMINI PW laser system measured using the Sequoia third order autocorrelator scans. The noise floor of the detector is  $\approx 1 \times 10^{-9}$ . The resolution is 100 fs, thus the main pulse is not resolved by the detector.

#### 4.5.2 Contrast on GEMINI PW

An underlying current of this thesis has been the subject of prepulse and corresponding preplasma expansion. While bright X-ray harmonics have been observed on both the Vulcan [160] and ORION sub-ps laser systems, few cycles laser pulse systems routinely underperform. To date only at normal incidence (thus minimising heating) have harmonics been observed on GEMINI PW [193]. The only parameter that has been shown to suppress HHG from a p-polarised relativistic laser pulse is an overlarge preplasma [169] and indeed the transition is rapid [190].

The last available contrast measurements of the GEMINI PW laser are presented in Figure ???. The **DPM!** setup improves the contrast by at least  $10^5$ , improving the contrast a few ps out to  $10^9$ . That is consistent with a double AR coating plus the peak measured reflectivity of 50 %, consistent with two AR coated PMs each with reflectivity  $\approx 0.5\%$  when switched off and 70 % when switched on. More recent measurements of the GEMINI PW DPM system have found a reflectivity of  $1.23 \times 10^{-4}\%$  pre switch on and 60 % after, corresponding to a contrast enhancement of  $4.88 \times 10^5$ .

HYADES simulations of the GEMINI PW pulse demonstrate acceptable preplasmas. Note in Figure ?? the early switch on of the plasma mirror system. For a peak intensity on **PM!** of  $1 \times 10^{16} \text{ W cm}^{-2}$ , this switch on occurs an order of magnitude

below typical ionising intensities [70]. Early PM switch ons is nothing new [192], however, it points to a serious issue. Such early switch ons cannot be described by the basic theory nor hydrodynamic simulation. Perhaps all can be explained by the complex interaction of the laser pulse with the AR coating. Indeed perhaps the very optic deemed to save us is our great enemy in the battle for harmonics. Yet still the hydrodynamic codes fail us. Preliminary experiments have identified another possible cause. Early (few ps) opacity generation on the main target. This could dramatically affect laser pulse energy absorption into the main target (and in the PM system) leading to massive plasma heating and early preplasma formation but such microscopic crystal effects cannot be modelled via macroscopic hydrodynamic codes. Presumably this early plasma formation also occurs for Vulcan or ORION targets, however, once relativistic intensities are accessed, recompression of the target will occur. For these long pulse durations there is sufficient time for steepening before the main interaction. Note that contrast reduction is even more relevant for next-gen laser facilities where a  $10^{-4}$  contrast is already relativistic! Dramatically affecting interactions. Furthermore, ion acceleration via TNSA of nanometre foil targets is a very popular use case of high power short pulse laser systems. Such next-gen systems absolutely require excellent contrast so if we can demonstrate contrast enhancement where it really matters using the DPM adjustments, that will be very interesting.

To investigate early plasma switch on, a 50 ps chirped probe beam will be focused on target. Spectrometers in transmission and reflection should identify opacity and plasma generation times respectively. This measurement could categorically demonstrate the failure of hydrodynamic codes for any preionisation laser pulse intensities. One may wonder that such a failure not identified in other branches of HED physics, such as the case of Inertial Confinement Fusion. It is simply a matter of scales. These effects are only relevant for the few femtosecond petawatt laser system in this case of high sensitivity to preplasma scale length. To distinguish the probe beam from the bright non-linear target interaction, it will be focused onto the interaction area out of the main interaction plane and with opposite polarisation.

## Strategy

For this initial stage, the OHREX will be replaced with a XUV spectrometer since only the lowest orders will likely be visible. As of 2019, no experiment to confirm HB theory had been performed on a PW system [165]. These UV spectrometer measurements could be very interesting both to confirm the HB theory in the PW regime and to infer the intensity increase at PM focus. The first step will be to perform a scan around the focal spot of the laser. Wavefront curvature at the target from slight defocussing can reduce divergence from hole boring thus increasing intensity levels at the detector while also reducing pre-heating. The next action will be to apodise the beam, again this reduces the intensity on target and the harmonic beam divergence and has been shown to tip preplasma scale length into the optimal regime for HHG [190].

This will reduce the intrinsic prepulse intensity by the same factor as the main pulse. Comparing the GEMINI contrast and laser intensity to that of the Kahaly experiment, it would appear the prepulse intensity of GEMINI is roughly 25 times larger (assuming equivalent DPM AR coatings, this much I don't know). It is possible GEMINI contrast has improved since its last measurement, I hope to obtain this information from CLF. If it is unchanged it seems likely GEMINI PW contrast sits at the cusp of that required for SHHG from glass targets (assuming their additional prepulse was small). In which case we may have to somewhat apodise the beam. To ensure the SHHG roll off is above the 2.4 keV harmonics (such that the signal on OHREX is non-negligible), we require  $a_0 > 8.11$ . Such apodisation would reduce our prepulse by roughly 15 times. Note that (assuming top hat beam profile at near field) for apodisation of a beam from radius  $r$  to  $r'$ , the beam energy reduces by a factor  $(r'/r)^2$  while the spot size increases by a factor  $r/r'$ , thus the peak intensity reduces by a factor  $(r'/r)^4$ .

Hard edges lead to high intensity diffraction rings. Instead the apodising aperture typically has sharp teeth edges to reduce possible damage to optics down the chain. Spatial filters can also help: focus the beam down under vacuum, the lens acts as a fourier transform, higher spatial frequencies (that are responsible for the high

intensity diffraction rings) will be further from the centre of the focused spot. An aperture around the focused spot will filter them out.

Note that Kahaly *et al* were able to observe harmonics for a beam apodised to  $a_0 = 0.7$  provided the scale length was optimised.

Note that apodisation improves beam quality by removing the nasty bits at the edges of the beam.

Big note: we cannot assume that if we apodise back to Astra-GEMINI parameters we will see harmonics as they do. The whole front-end has been changed, its effectively a different system and it seems that GEMINI PW has this specifically challenging ps pedestal.

Apodising is our first approach, we really hope we see some HHG signal even if it is bad and then we work backwards from there (ie modify our PMs to see if we can improve our contrast sufficiently to reduce our apodisation). We try to change aspects of our system and look for quantitative improvements in our signal. A bi thing about harmonics is these prepulses are pretty consistent, if we can get it working on one shot, it will always work with that setup (BD).

Failing that, there are more drastic measures available. First, a second target chamber layout has been designed that will reduce the angle of incidence on target to  $30^\circ$  and second, Frequency doubling is known to significantly enhance contrast. In fact as with ORION it is likely we go straight to overshooting.

This would be very useful for demonstrating the validity of contrast theory. The frequency doubling crystal available at GEMINI PW leads to an intensity reduction both from efficiency and apodisation (the crystal is smaller than the beam size), therefore if we do see an improvement in signal, it must be from the contrast improvement.

Since it is very likely we now overshoot our optimum preplasma scale length, we will absolutely need to have a Kahaly mirror specifically made for 2 omega (or we could try increasing our angle of incidence).

60 mm of the beam can pass through the frequency doubling crystal - 9

Note that despite inevitability of efficiency cost due to the frequency doubling process, similar intensities on target could be realised provided a 25

Note that it is important to take care with frequency doubling. One must ensure you have removed the residual infrared, typically using dichroic plates that transmit second and reflect first. One must also ensure all mirrors after the frequency doubling in the chain are suitable for both frequencies, one must have a suitable Kahaly prepulse mirror and a waveplate since the process produces 2nd harmonic at 90 degrees to IR and we need p-pol on target. Probably easiest to rotate the IR.

How does frequency doubling work? Many non-linear processes happening in this crystal, it is about cutting it so this one is favoured, beyond that unknown.

### 4.5.3 Kahaly

Once harmonics become visible, hopefully early in the list of contingencies presented, the signal will be optimised using a controllable prepulse using the Kahaly technique [190].

### 4.5.4 Other bits to include

PMs can also act as a low-pass spatial filter, improving the quality of the focal spot on target [161].

Note that the DPM setup includes an acetate debris shield, without which shot quality quickly degrades

Hypothesis testing: make this more of a thing. Once harmonics are visible, the opacity and early switch on hypotheses can be tested. Replacing the pure fused silica target with impure BK7 or AR-coated targets should reduce the harmonic signal. Equally, replacing the first PM in the DPM chain with an uncoated PM should delay the PM switch on and potentially improve the signal. Note that the CWE signal is distinguishable from the HHG signal: it is more divergence and cuts off after the plasma frequency [190].

In parallel to the main experiment we can set up a probe beam using the GEMINI PW S beam to test the MY/BD hypothesis. Note that since this is only

required for the first experimental phase, we can then change the geometry during the electron bunch experiment phase to that required for shock compression.

#### Opacity growth in plasma mirrors

While there exists established theories on the operation of plasma mirrors, these are based on fairly large assumptions that remain untested. Given the notoriety of the challenge of observing SHHG despite laser contrast being relatively well understood and the general successes attributed to hydrodynamic codes, it seems quite possible that these large assumptions are the critical concern. They could also be responsible for the range of optimal scale lengths observed in experiments. Recent work from MY and BD has shown evidence of the propagation of an opacity front into a target due to laser prepulse prior to ionisation and plasma formation. Such a transition from transparency will dramatically affect laser pulse energy deposition into the target leading to significantly earlier plasma mirror ‘switch on’ times. The interaction appears to occur over roughly 3 ps, a timescale to which preplasma scale length is highly sensitive. Note that such an interaction would be impossible to simulate given the multiple scales upon which preplasma formation occurs and the complexity of the transition to opacity. The GEMINI experiment, however, provides an excellent opportunity to probe this process in parallel to the main experimental goals.

While the GEMINI PW target chamber does not easily accommodate diagnostics around the DPM setup, it should be relatively straightforward to set up a probe beam on the main target to probe the same effect. My and BD propose the following.

#### Probe beam

In order to determine target opacity transition and plasma mirror switch on times, we can set up a probe beam as previously done by MY and BD. A low intensity chirped beam is focused to the target focal spot with optical spectrometers positioned in the transmission and reflection directions. The chirp encodes time in wavelength, thus there should be a sudden drop off in the observed spectrum in transmission at the switch to opacity and a sudden rise in reflection at plasma mirror switch on.

Such a beam can be created by siphoning off the S beam prior to the last quantile/amplifier? (*i.e.* on high power not full power), apodising and detuning to a 50 ps beam. It may be necessary to pipe the beam around the chamber to increase the path length. We can then re-purpose the S beam for the shock study.

Given the high intensity main pulse interaction there will be bright emission at all angles around the fundamental wavelength. To isolate the probe beam signal, it would be sensible to frequency double the beam and use opposite polarisation and extinction polarisers.

The probe beam spectrum must be calibrated to the main pulse time. There are two available options. One, MY to perform spectral interferometry of the probe and main beams, it would perhaps be prohibitively challenging to ensure both beams are passed collinearly into an awaiting spectrometer. This could be achieved using a small glass bead or wire. Alternatively, there is the brute force method: use the main beam on low power, dialing down the intensity until a plasma is just formed by the peak of the main beam, to determine the wavelength at which reflection switches on. This could even be performed at air and simply create an air plasma, no targets necessary. In theory calibration should take no more than a day. Note that the DPM must be used in gold reflection mode, DPM should not be damaged since the laser only just triggers a plasma at peak focus.

Calibration could be done at air since not using laser full power, increasing action efficiency, however, the air will slow down the laser pulses, potentially leading to changes in path lengths, likely limiting resolution to the fs level.

DS on air plasma: Yes – the easiest way is to search the area and operate from the control room. If you need to do it while in the area, it's possible but needs some safety paperwork.

DS says main challenge with using the GEMINI probe beam is getting a chirp longer than 10ps as there is not enough space in the compressor to expand the grating any further, however Mark says this should be sufficient since we're interested in activity on the few ps scale. It would be good to have the option for longer

since we are also interested in any change that happens before when we normally expect the plasma to trigger.

#### Opacity hypothesis

A totally pure silicon target consists of valence and conduction bands only. Electrons are excited to the conduction band by the laser and then simply drop back down again. As impurities are added to the target, the band structure becomes more complex and electrons can drop to intermediate states from which they cannot then return easily to the ground state. Thus, more energy can be absorbed and the transition to opacity occurs at earlier times.

Hence, the more pure a silica target, the later the transition to opacity and the smaller the preplasma scale length. This should be straightforward to establish by comparing the relative delay of the required prepulse for SHHG optimisation. This can be cross-referenced to ionisation timings measured using the probe beam. This also suggests switching out the first plasma mirror for a pure silicon target could also improve contrast where it matters. Since we are fairly assured the DPM removes all ns prepulse pedestal, it seems likely it is the ps prepulse pedestal that is causing issue. Now PM1 deals with the ps pedestal (with the later switch on) and PM2 the ns pedestal (with the AR coating). Again this is a simple test where we compare optimum prepulse delay times.

#### 4.5.5 Kahaly mirror

Assuming GEMINI PW + DPM contrast is good enough, we then want to optimise for SHHG. A simple but elegant way to do this was detailed by Kahaly *et al* [190] and favoured by MY/BD.

Simply partially cover a mirror in the laser chain with a small mirror. One can control the difference in path length of light on this mirror by the mirror thickness and distance from the main laser mirror, thus we have a controllable prepulse. Since one has effectively apodised the beam, the parabola will then focus the prepulse spot to a larger area compared to the main pulse spot while the diffraction limited main spot will have little aberration due to the subtracted prepulse. By routinely

imaging the spot one can adjust the prepulse mirror angle to ensure the prepulse spot coincides with the main pulse spot.

The mirror should have a mirrored back surface and an AR coated front surface with a slight angle between the two faces to prevent a parasitic prepulse.

Kahaly *et al* demonstrate a linear relationship between main pulse delay relative to prepulse and scale length (due to more time for the preplasma to expand) but also note that for the same scale length, a larger delay (accompanied by a smaller energy in prepulse) produces higher quality harmonics. This could be understood as a natural consequence the lower plasma temperatures, which improve the ability of the system to perform coherent SHHG behaviour. If we access harmonics using the Kahaly method, we could attempt to improve our signal by moving the prepulse mirror further from the main pulse mirror while also reducing the fraction of overlap between the two (reducing the prepulse energy).

Specs for the Kahaly mirror found [here](<https://www.notion.so/Optics-1acf5aa1029941e2aef30afdd53>)

Kahaly mirror should be placed before the final mirror. The further back in the chain it is placed, diffraction from the mount can cause greater danger.

## 4.6 Beyond GEMINI PW

Vulcan 2020

Expand the range of targets A range of thick, flat solid targets are proposed to probe the density parameter space: diamond, a selection of plastics (polymethyl methacrylate, low and high density polyethylene and polycarbonate) and mirrored targets, which produced interesting and slightly unexpected results in the ORION experiment. It would also be interesting to produce foam targets gaining access the optimal low  $S$  regime [194].

Image the XUV harmonics from the attosecond electron bunches.

Write about PIC simulations here.

Measuring the HHG from attosecond electron bunches.

## **4.7 Conclusion**



# 5

## Miscellaneous notes

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### 5.1 To do

1. HYADES simulations
2. Add more citations
3. Similarity theory details in appendix
4. Add details on CPA and OPCPA
5. Simulation algorithms (details in appendix?)
6. Velocity transformations derivation (appendix)
7. Sources of error system

8. Ponderomotive heating mechanisms
9. Fix QED section
10. Particle merging
11. Radiating particles and relativistic larmor radius
12. smieli performance plot
13. A collisionless fully ionised plasma
14. Basic derivation of the Schwinger limit
15. Calculating collision frequency
16. Feynmann diagrams
17. Add some detail of vectorisation

On the ZVP front

1. Fix diagrams
2. Go through discussions
3. Add errors plot
4. Check out new sims
5. Conclusion
6. Shape of transverse momentum (more parabolic compared to linear and explanation of bunch holding together), Sharp front edge then parabolic (or even exponential decay - it is defo exponential)
7. a0 convention, when max and when varying in time
8. Editing

## 5.2 ORION experiment

The following derivation determines the polarisation of the ORION laser pulses in the experiment and the boosted frame quantities for the PIC simulations.

I will have a whole subsection devoted to the different frames of reference of relevance and then a second one about normalised units. What follows now is the derivation of the boosted frame in which the laser is incident normally relative to the lab frame where the laser is incident obliquely.

I will try to use a consistent convention for coordinate system as much as possible.

### 5.2.1 Frames of reference

Other frames of reference include, HB front surface at rest frame, ablating front frame, smilei frames.

When writing out the pistonning equation in full in thesis, include analysis in Robinson 2009 to do it for multiple ion species.

I should go over this and use third year relativity notes to formalised and make more consistent.

While some of this section may seem trivial, it is frequently miscalculated in the literature, it therefore seems of great importance to provide a full derivation.

The following is inspired by [195], here they give the formula for  $k$  and  $\omega$ .

In [196] they determine the normalised vector potential, if defined similarly in the new frame, then it is frame invariant. This fails to consider what about the fact that the vector potential is in reality more complex and is not simply the temporal integral of the electric field. Nonetheless it is still reasonable to define it so if what is of real importance (as is usually important) is in fact the fields and it is simply being used as a way to normalise the field intensity.

Consider a photon incident on a plasma block at angle  $\theta$  as in figure 5.1. A boost is applied with velocity  $\mathbf{v}$  to a frame such that the photon is normally incident on the now streaming plasma at velocity  $-\mathbf{v}$ . The velocity transformation for the

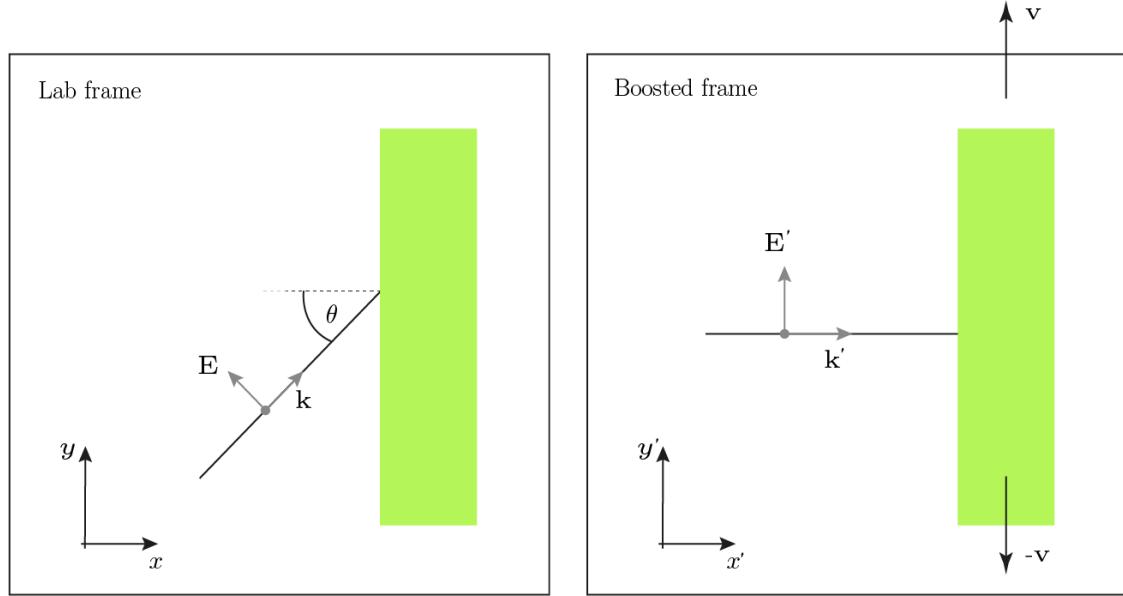


Figure 5.1

photon's velocity,  $\mathbf{u}$ , parallel to the boost is

$$\mathbf{u}'_{\parallel} = \frac{\mathbf{u}_{\parallel} - \mathbf{v}}{1 - \mathbf{u} \cdot \mathbf{v}/c^2}. \quad (5.1)$$

Setting  $\mathbf{u}'_{\parallel} = 0$ , it is clear that

$$\mathbf{v} = \mathbf{u}_{\parallel} = c \sin \theta \hat{\mathbf{y}} \quad (5.2)$$

in this geometry and

$$\gamma_{\mathbf{v}} = \frac{1}{\sqrt{1 - \mathbf{v}^2/c^2}} = \sec \theta. \quad (5.3)$$

Noting that since Snell's law is frame invariant, the photon remains normal as it propagates into the skin depth of the plasma, a frame in which the interaction reduces to a 1D problem has been successfully found for all  $\theta < \pi/2$ . Those familiar with the topic may wonder how this is possible considering the ‘ripples’ that are observed on the plasma surface for oblique incidence. The explanation for this is of course the relativity of simultaneity. It remains to determine how do all the relevant quantities transform as such a boost is applied. Starting with an easy one: the photon's wave four-vector is

$$\mathbf{K}^{\mu} = \left( \frac{\omega}{c}, \mathbf{k} \right) \quad (5.4)$$

and thus the frequency transforms as

$$\frac{\omega}{c} = \gamma_v \left( \frac{\omega'}{c} - \frac{\mathbf{v} \cdot \mathbf{k}'}{c} \right). \quad (5.5)$$

Since  $\mathbf{v} \cdot \mathbf{k}' = 0$ ,

$$\omega' = \omega \cos \theta. \quad (5.6)$$

As

$$n'_c = \frac{m_e(\omega')^2}{4\pi e^2}, \quad (5.7)$$

$$n'_c = n_c \cos^2 \theta, \quad (5.8)$$

while the plasma block will be Lorentz contracted along  $\hat{\mathbf{y}}$ , hence the number density of electrons will increase as,

$$n'_e = \frac{n'_e}{\cos \theta}, \quad (5.9)$$

leading to the perhaps unexpected

$$\bar{n}'_e = \frac{\bar{n}_e}{\cos^3 \theta}. \quad (5.10)$$

Time is dilated

$$t' = \frac{t}{\cos \theta}. \quad (5.11)$$

Consider now the more general case (I should just simply replace my diagram with a 3D one that incorporates this initially) where the photon's electric field is rotated out of the  $x$ - $y$  plane, *i.e.*

$$\mathbf{E} = E_0(-\cos \phi \sin \theta, \cos \phi \cos \theta, \sin \phi) \quad (5.12)$$

and correspondingly

$$\mathbf{B} = \frac{\hat{\mathbf{k}} \times \mathbf{E}}{c} = \frac{E_0}{c}(\sin \phi \sin \theta, -\sin \phi \cos \theta, \cos \phi). \quad (5.13)$$

The Lorentz transformations for electro-magnetic fields are

$$\mathbf{E}'_{||} = \mathbf{E}_{||}, \quad (5.14)$$

$$\mathbf{B}'_{\parallel} = \mathbf{B}_{\parallel}, \quad (5.15)$$

$$\mathbf{E}'_{\perp} = \gamma_{\mathbf{v}}(\mathbf{E}_{\perp} + \mathbf{v} \times \mathbf{B}), \quad (5.16)$$

$$\mathbf{B}'_{\perp} = \gamma_{\mathbf{v}}(\mathbf{B}_{\perp} - \mathbf{v} \times \mathbf{E}/c^2). \quad (5.17)$$

Using the above expressions for  $\mathbf{E}_{\perp}$  and  $\mathbf{E}_{\parallel}$  and transforming to the boosted frame,

$$\mathbf{E}' = E_0 \cos \theta (0, \cos \phi, \sin \phi). \quad (5.18)$$

As anticipated for normal incidence there is no component of the E-field normal to the surface. Conveniently, the polarisation of the incident photon is unchanged despite having components both parallel and perpendicular to the transformation and

$$|\mathbf{E}'| = |\mathbf{E}| \cos \theta. \quad (5.19)$$

The picture can now be completed. Since

$$a'_0 = \frac{e|\mathbf{E}'|}{m_e e \omega'} \quad (5.20)$$

it follows that

$$a'_0 = a_0, \quad (5.21)$$

$$S' = \frac{S}{\cos^3 \theta}. \quad (5.22)$$

### Four-potential transformation

Consider a laser pulse obliquely incident, angle  $\theta$ , it has 4-vector potential  $\mathbf{A}^{\mu} = (0, A \sin \theta, A \cos \theta, 0)$ , where  $A = A_0 \sin(\mathbf{k} \cdot \mathbf{x} - \omega t)$  and  $\mathbf{k} = (k \cos \theta, k \sin \theta, 0)$ . Applying the lorentz transformation, in the frame where the laser pulse is normally incident,

$$\mathbf{A}'^{\mu} = (-\gamma \beta A \cos \theta / c, -A \sin \theta, \gamma A \cos \theta, 0) = (-A \sin \theta, -A \sin \theta, A, 0) \quad (5.23)$$

since  $\beta = \sin \theta$  in the positive  $y$  direction and  $\gamma = 1 / \cos \theta$ .

Therefore,

$$\mathbf{E}' / \cos(\mathbf{k} \cdot \mathbf{x} - \omega t) = -\sin \theta \cos \theta k A_0 \hat{\mathbf{x}} - \sin \theta \cos \theta \omega A_0 \hat{\mathbf{x}} + A_0 \omega \cos \theta = A_0 \omega', \quad (5.24)$$

since  $\sin(\mathbf{k} \cdot \mathbf{x} - \omega t) = \sin(x'k \cos \theta - \omega t' \cos \theta)$ . Note I have been sloppy here with factors of  $c$ .

Do it does look at though the vector potential amplitude is unchanged by the transformation but it is more subtle, it is useful to define an  $a'_0 = eE/m_e\omega$  to normalise the electric field intensity but note that this cannot be converted into the actual vector potential, it is just a useful construct.

Also note that whatever the transverse vector potential is in the boosted normal incidence frame is simply whatever the amplitude of the vector potential is in the oblique incidence frame.

### 5.2.2 ORION interaction geometry

Moved to ch:3

### 5.2.3 Condition on validity of hole boring expression

Robinson *et al* [167] consider for what case is the expression they derive for hole boring valid. The case they are interested in is what happens if the energy available for an ion to gain from crossing the pseudo-capacitor is less than the kinetic energy associated with the hole boring velocity. Their analysis applies for non-relativistic hole boring velocities and circular polarised laser pulses. This theory is now updated for the ZVP mechanism (linear polarised and relativistic ion velocities).

The so-called ‘piston’ which leads to ion hole boring is the pseudocapacitor field. In section 2.2.1, the development of that field is discussed quantitatively. The peak electric field is

$$E_C = E_L = \sqrt{\frac{I}{\epsilon_0 c}} \quad (5.25)$$

and the peak displacement of electrons is

$$\Delta x = \frac{\epsilon_0 E_C}{en_e}. \quad (5.26)$$

Considering instead the relativistic kinetic energy gained by an ion were it to fully cross the pseudocapacitor, following equation 2.26,

$$T_i = Z_i \times \frac{1}{2} m_e c^2 \frac{a_0^2}{\bar{n}_e} = \frac{IZ_i}{2cn_{matherme}}. \quad (5.27)$$

(The equation above needs more thinking about)

Ions are reflected provided,

$$T_i > \frac{1}{2}m_i v_{\text{HB}}^2. \quad (5.28)$$

Hmm ok so in Vincenti, they approximate electron mass as much less than ion mass and therefore neglect it in the momentum calculation. It also looks like they have not done full relativistic calculation, so I cannot yet say I have that. But carrying on the derivation using Vincenti expression for simplicity:

The hole-boring velocity as calculated by Vincenti *et al* [122] is

$$\frac{v_{\text{HB}}}{c} = \sqrt{\frac{R \cos \theta}{2}} \quad (5.29)$$

So come back to this section, once I have fully written out the hole boring calculation in full, include also the multiple ion species stuff and this condition.

The upshot of this condition is something like: require no low charge to mass ratio ions (ie v heavy ions) and fully ionisation, these conditions are satisfied in this area of study.

To arrive at that result, useful parts include: composite mass density  $\rho = \sum_i m_i n_i$ ,  $m_i = A_i / N_A$  and  $A_i \approx 2Z_i$  for most low mass ions relevant in these plasmas.

### 5.3 Thinking about the ZVP calculation

When I run sims on oblique incidence, this is another theorem that could be interesting to test.

Consider now that the surface moves inwards at speed  $c$ . In a time  $\Delta t$ , an energy  $\sim B_L^2 \Delta t$  is incident on the surface. If at such a point, there exists a pseudocapacitor with electric field  $E_C \sim n_e x_e$ , then the work done by pushing it inwards is  $\sim E_C n_e x_e \Delta x \sim E_C^2 \Delta t$ , since surface moving inwards at speed  $c$ . Thus by conservation of energy, the reflected field is  $B_R^2 = B_L^2 - E_C^2$ . Note that at max displacement this cannot possibly be the case and we do see that the surface stops moving inwards, however this could be due to a reduction in the

laser pulse intensity since the peak has passed. Thus this could be a reasonable approximation of the phenomena.

Then the force equilibrium expression in the boosted frame is

$$-B_L - \sqrt{B_L^2 - E_C^2} \pm B_i + E_C \quad (5.30)$$

working this through one finds,

$$x_p = \frac{\cos^2 \theta}{kS} \frac{2(1 \pm \sin \theta)}{\sin^2 \theta \pm 2 \sin \theta + 2} \quad (5.31)$$

$$T \sim \left( \frac{\cos^2 \theta}{kS} \frac{2(1 \pm \sin \theta)}{\sin^2 \theta \pm 2 \sin \theta + 2} \right) \quad (5.32)$$

And thus now predicting an optimum for electron energy at  $\theta \approx 30^\circ$ . That is quite different. It also looks nicer so I would like this to be right.

Another thing I still need to do is gonoskov technique to get bunch thickness, also do ZVP calculation in the exponential preplasma.

## 5.4 Things I may want to include or random notes

Note that ZVP does not describe the peak energies in the bunches, then  $J \times B$  applies, since there are always some electrons outside of the well defined sharp boundary when the density is not high enough to impose adiabaticity.

The phase at ejection is locked at the electromagnetic field peak of the laser pulse cycle. Diffraction around the target edge does occur and can be observed in Figure 10 but primarily not in the vicinity of the electron bunches (electron bunches that undergo Vacuum Laser Acceleration (VLA) will be ejected to the other side to that experiencing diffraction). As in the work of Thévenet *et al.*, not all electrons in the bunch will experience VLA, only those electrons propagating close to parallel to laser pulse, the rest that dephase do not gain further energy [Thevenet2016]. Electrons that do not retain their phase will be randomly ponderomotively accelerated and decelerated across laser cycles, hence in the far field, there will be high energy attosecond electron bunches surrounded by a low energy noise of electrons. Note that the strong modulation of the reflected field for these high laser intensities

has not yet been considered for VLA in reflection, and would likely limit possible accelerations, while HHG in transmission is always weaker [**cousens2020**], ensuring the presence of a fundamental laser pulse to perform the VLA

reword: ‘Figure 5.2a compares the incident laser pulse to the strongly modulated reflected pulse in the 3D PIC simulation. The Fourier transform of the reflected pulse is presented in Figure 5.2b. Due to the high intensities in these simulations, the spectrum is of the modified CSE type detailed by Edwards and Mikhailova [**edwards2020x**]: initially the spectral intensity scales as  $\sim n^{-4/3}$  up to a cut off determined by the advance time bunch width of radiating electrons after which it scales as  $\sim n^{-10/3}$ . Edwards and Mikhailova demonstrated that this cut off, extracted from the internal dynamics of the system can be well approximated by the point where the fit to the spectrum drops below the  $\sim n^{-4/3}$  scaling, at harmonic number,  $n = 12$  in this simulation. This 3D simulation result is consistent with the  $n = 11.3$  determined by Edwards and Mikhailova in their most similar 1D simulation at  $a_0 = 100$ ,  $\bar{n}_e = 90$ ,  $\theta = 45^\circ$ . It is also interesting that since their definition of the bunch width corresponds to the temporal width of the radiation spike at observation, taking the full-width-half-maximum of the CSE type spikes (between  $t = 15$  and  $26$  fs) in Figure 5.2a) as the cut off harmonic for each spike gives a mean harmonic cut off of  $n = 11.4$  consistent with the spectrum fit and corresponding to an average pulse duration of 292 as. Hence, the cut off can infer the attosecond pulse duration from a simple UV spectrometer measurement without the need for complex attosecond resolution diagnostics. A second cut off dependent on the peak gamma factor of radiating electrons and beyond which the spectrum decays exponentially is not captured at this simulation resolution. The deviation of the spectrum from regularly spaced harmonics is a natural consequence of the high laser pulse intensity: the non-negligible hole boring velocity (scaling linearly with the electric field strength of the laser pulse [166]) significantly lengthens the path of the reflected pulse, Doppler shifting harmonics between successive pulse cycles.’



3D\_HHG.png

**Figure 5.2: Electric field temporal structure in 3D Particle-In-Cell (PIC) simulation with  $a_0 = 100$ ,  $\bar{n}_e = 100$ .** a) Temporal variation of the normalised vector potential of the incident and reflected laser pulses along the polarisation axis of the incident laser pulse. The reflected pulse demonstrates attosecond radiation spikes without the need for spectral filtering. b) The spectral intensity of the reflected radiation obtained via a Fourier transform of the pulse in a). The fit is calculated following the methodology of Edwards and Mikhailova [edwards2020x]:  $\omega_b/\omega_L$  defines the cut off above which an ordinary least squares fit to  $\sim n^{-p}$  yields an exponent,  $p > 4/3$ . Beyond the cutoff the spectrum is predicted to scale as  $\sim n^{-10/3}$ . The fit is a simple weighted polynomial fit to the logarithm of the data using the NumPy polyfit module.



# Appendices



# A

## General plasma physics

### Contents

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<b>A.1</b>	<b>Lorentz transformations of electromagnetic fields</b>	<b>137</b>
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### A.1 Lorentz transformations of electromagnetic fields

The Lorentz transformations for electromagnetic field components parallel,  $\parallel$ , and perpendicular,  $\perp$ , to a frame of reference boost of velocity  $\mathbf{v}$  are [74]

$$\mathbf{E}'_{\parallel} = \mathbf{E}_{\parallel}, \quad (\text{A.1})$$

$$\mathbf{B}'_{\parallel} = \mathbf{B}_{\parallel}, \quad (\text{A.2})$$

$$\mathbf{E}'_{\perp} = \gamma_{\mathbf{v}}(\mathbf{E}_{\perp} + \mathbf{v} \times \mathbf{B}), \quad (\text{A.3})$$

$$\mathbf{B}'_{\perp} = \gamma_{\mathbf{v}}(\mathbf{B}_{\perp} - \mathbf{v} \times \mathbf{E}/c^2). \quad (\text{A.4})$$

## A.2 The headlight effect

The headlight effect describes the beaming of an isotropically emitting source travelling at some velocity relative to an observer. Consider the geometry of figure 2.2 with the source (the laser pulse) travelling at an angle  $2\theta$  to the observer (in this case, the ablating front). A photon with energy  $E$  emitted from the rest frame of the source (the laboratory frame in this case) has a 4-momentum

$$\mathbf{P}_\mu = \left( \frac{E}{c}, \frac{E}{c} \cos 2\theta, \frac{E}{c} \sin 2\theta \right). \quad (\text{A.5})$$

As the interaction geometry is confined to a 2D plane, the  $z$ -component can be safely neglected. Applying the lorentz boost of equation 1.30,

$$\begin{aligned} \frac{E'}{c} &= \gamma \left( \frac{E}{c} - \beta \frac{E}{c} \cos 2\theta \right) \\ \frac{E'}{c} \cos 2\theta' &= \gamma \left( \frac{E}{c} \cos 2\theta - \beta \frac{E}{c} \right). \end{aligned} \quad (\text{A.6})$$

Solving these equations for the angle in the boosted frame,

$$\cos 2\theta' = \frac{\cos 2\theta - \beta}{1 - \beta \cos 2\theta}. \quad (\text{A.7})$$

## A.3 Geometric transverse emittance

A beam<sup>1</sup> of particles is fully described by its six-dimensional particle phase space distribution

$$\rho(\mathbf{x}, \mathbf{p}) = \rho(x, p_x, y, p_y, z, p_z), \quad (\text{A.8})$$

where  $\mathbf{p} = p_x \hat{\mathbf{x}} + p_y \hat{\mathbf{y}} + p_z \hat{\mathbf{z}}$  is the canonical momentum [133]. Under the Hamilton formalism, for ideal conditions, the six-dimensional volume of the beam in this space, termed the *emittance*, arises as a conserved quantity and is therefore a useful quantity to describe the beam quality. (something to do with it affecting the ability to focus the beam?? check the papers) It is useful to rotate the coordinate system so as to align with the beam's propagation. The distribution can be written as

$$\rho(\mathbf{x}', \mathbf{p}') = \rho(x_L, p_L, x_T, p_T, x_{T'}, p_{T'}), \quad (\text{A.9})$$

---

<sup>1</sup>In this section it is electron beams and not bunches that are referred to to demonstrate the generality of these concepts.

where L is longitudinal to the beam's propagation direction, and T and T' are two orthogonal directions transverse to the beam's propagation. Where discussed in this thesis, T' will unanimously refer to the  $z$ -direction, that is, the additional direction in 3D simulations, all such simulations are designed such that the  $z$ -direction is transverse to the beam propagation direction.

Recording a six-dimensional phase space in experiment is impossible while in simulations it is almost prohibitively costly in terms of data storage. Hence, it is common practice to project the distribution onto three orthogonal sub-spaces corresponding to each spatial axis, L, T and T' and compute the area on each. Note that since the electron beam is ultra-relativistic, all electrons propagate at approximately c and therefore it is the transverse and not the longitudinal emittance that describes the beam's quality. As a particle beam does not typically exist with well-defined borders, the area used to describe the emittance is restricted to an ellipse containing only the high-density core of the distribution. For a subspace  $i$ , where  $i = T$  or  $T'$ , Floettmann *et al* [197] derive the *transverse normalised emittance* as

$$\epsilon_{n,rms}^i = \frac{1}{m_e c} \sqrt{\langle x_i^2 \rangle \langle p_i^2 \rangle - \langle x_i p_i \rangle^2}, \quad (\text{A.10})$$

where  $\langle \rangle$  is the second central moment of the particle distribution,

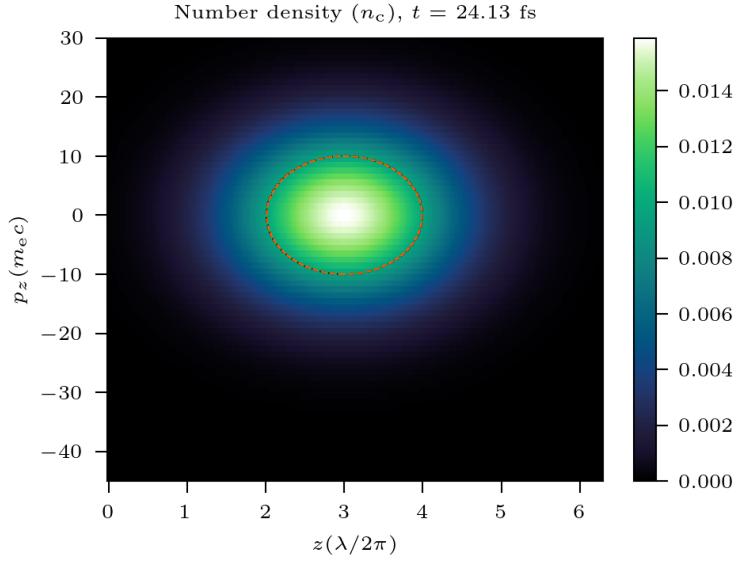
$$\langle ab \rangle = \frac{\int ab \rho(\mathbf{x}', \mathbf{p}') dV}{\int \rho(\mathbf{x}', \mathbf{p}') dV} - \frac{\int a \rho(\mathbf{x}', \mathbf{p}') dV \int b \rho(\mathbf{x}', \mathbf{p}') dV}{(\int \rho(\mathbf{x}', \mathbf{p}') dV)^2}, \quad (\text{A.11})$$

here  $dV = \Pi_j dx_j dp_j$  for  $j = L, T, T'$ .

When working with emittances, most frequently in the literature it is the *transverse geometric emittance*,  $\epsilon_{rms}^i$  that is discussed. This is a natural consequence of it being more readily accessible in experiments [133]. The geometric and normalised emittances are related via

$$\epsilon_{rms}^i = \frac{\epsilon_{rms}^i}{\gamma \beta_L}, \quad (\text{A.12})$$

where  $\gamma = 1/\sqrt{1-\beta^2}$  refers to the beam's mean energy and  $\beta_L \approx c$  is the ultrarelativistic beam's longitudinal speed.



**Figure A.1:** Emittance calculation for an ideal Gaussian distribution in phase space,  $f_n(z, p_z) = 1/(2\pi s_z s_{pz}) \exp[-((z - m_z)^2/(2s_z^2) + (p_z - m_{pz})^2/(2s_{pz}^2))]$ , centred at  $(m_z, m_{pz}) = (3, 0)$  with standard deviations  $(s_z, s_{pz}) = (1, 10)$ .

The Courant-Snyder invariant which describes the ellipse that corresponds to the emittance is<sup>2</sup>

$$\epsilon_{\text{rms}}^i = \gamma x_i^2 + 2\alpha x_i x'_i \beta x_i'^2, \quad (\text{A.13})$$

here the coordinates are  $x_i$  and  $x'_i = p_i/p_L$  [198]. The Twiss parameters are

$$\alpha = -\frac{\langle x_i x'_i \rangle}{\epsilon_{\text{rms}}^i}, \quad (\text{A.14})$$

$$\beta = \frac{\langle x_i \rangle}{\epsilon_{\text{rms}}^i} \quad (\text{A.15})$$

and

$$\gamma = \frac{\langle x_i'^2 \rangle}{\epsilon_{\text{rms}}^i}. \quad (\text{A.16})$$

Thus the shape of the ellipse and the divergence of the beam can be determined. The elliptical contour defining the emittance of an ideal Gaussian phase-space distribution is given in Figure A.1. The contour is numerically calculated to contain 39.4(1) % of the particles, compared to 39.3 % from theory. Here the discrepancy arises from the finite grid on which the distribution is defined.

<sup>2</sup>Regrettably  $\beta$  and  $\gamma$  are the standard notations for the Twiss parameters, at all other locations in this Thesis, these parameters will refer to the standard relativistic beta and gamma factors of objects respectively.

# B

## PIC simulations

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### Contents

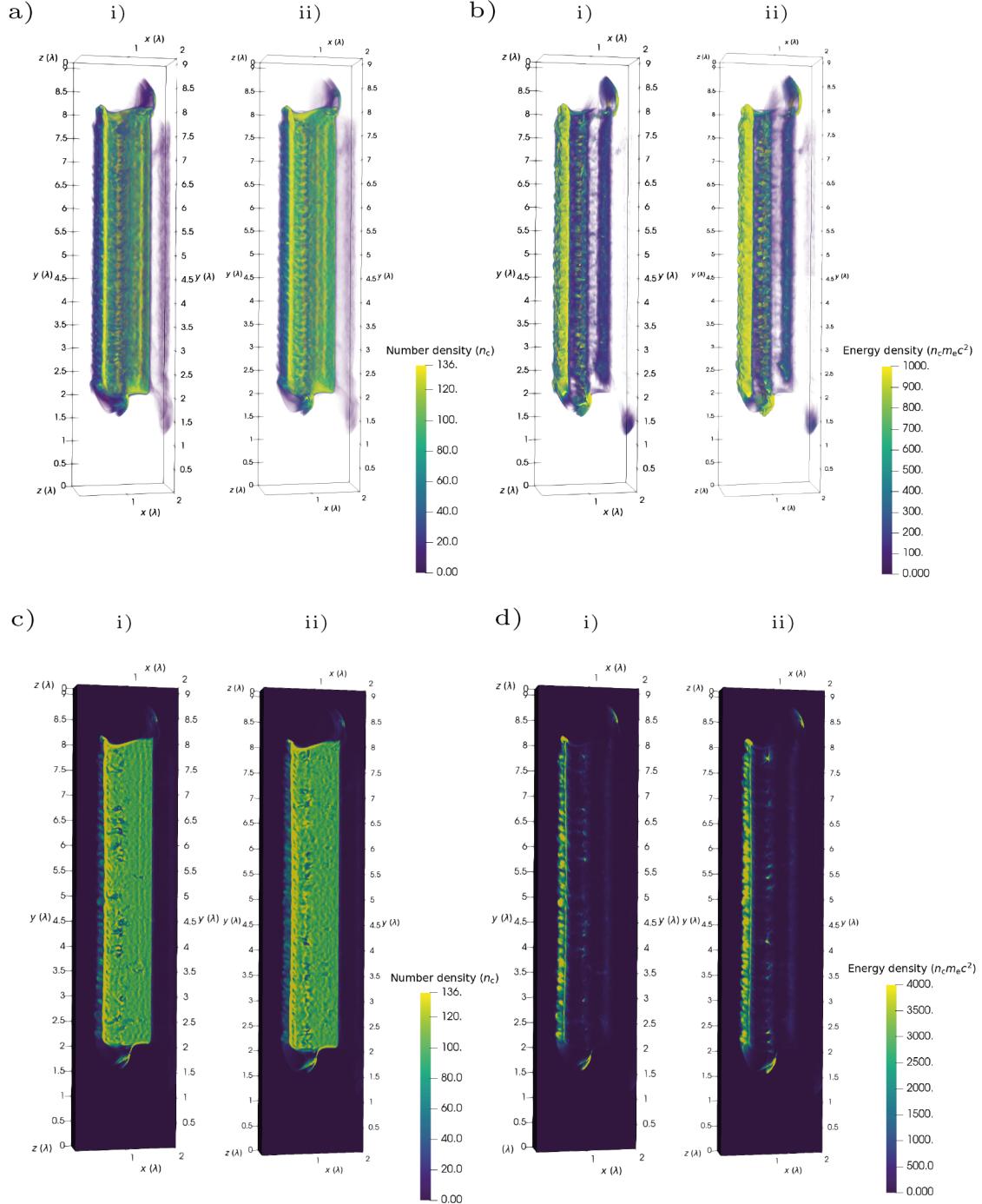
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B.1 Convergence of 3D PIC simulations . . . . .	141
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### B.1 Convergence of 3D PIC simulations

To reduce the computation cost of the large 3D simulations, a lower resolution version of the original 3D simulation was performed. Results are presented in figure B.1, demonstrating reasonable convergence.



**Figure B.1:** Comparison between the initial 3D simulation and a lower resolution version. a) Electron number density. b) Kinetic energy density. c) and d) are slices of a) and b) respectively. i) Initial simulation. ii) Lower resolution simulation. Good convergence is demonstrated.

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