

A treatise on phenomenological models of surface second-harmonic generation from crystalline surfaces

Bernardo S. Mendoza and Sean M. Anderson

(Dated: February 8, 2016)

I. THREE LAYER MODEL FOR SHG RADIATION

In this section we derive the formulas required for the calculation of the SHG yield, defined by

$$R(\omega) = \frac{I(2\omega)}{I^2(\omega)}, \quad (1) \quad \text{uno}$$

with the intensity

$$I(\omega) = \frac{c}{2\pi} |E(\omega)|^2, \quad (2) \quad \text{dos}$$

There are several ways to calculate R , one of which is the procedure followed by Cini [\[?\]](#). This approach calculates the nonlinear susceptibility and at the same time the radiated fields. However, we present an alternative derivation based in the work of Mizrahi and Sipe [\[?\]](#), since the derivation of the three-layer-model is straightforward. In this scheme, we represent the surface by three regions or layers. The first layer is the vacuum region (denoted by v) with a dielectric function $\epsilon_v(\omega) = 1$ from where the fundamental electric field $\mathbf{E}_v(\omega)$ impinges on the material. The second layer is a thin layer (denoted by ℓ) of thickness d characterized by a dielectric function $\epsilon_\ell(\omega)$. Is in this layer where the second harmonic generation takes place. The third layer is the bulk region denoted by b characterized by $\epsilon_b(\omega)$. Both the vacuum layer and the bulk layer are semiinfinite (see Fig. [3layer I](#)).

To model the electromagnetic response of the three-layer model we follow Ref. [mizrahiJOSA88 \[?\]](#), and assume a polarization sheet of the form

$$\mathbf{P}(\mathbf{r}, t) = \mathcal{P} e^{i\boldsymbol{\kappa} \cdot \mathbf{R}} e^{-i\omega t} \delta(z - z_\beta) + \text{c.c.}, \quad (3) \quad \text{m31}$$

where $\mathbf{R} = (x, y)$, $\boldsymbol{\kappa}$ is the component of the wave vector $\boldsymbol{\nu}_\beta$ paralel to the surface, and z_β is the position of the sheet within medium β (see Fig. [3layer I](#)). In Ref. [sipeJOSAB87 \[?\]](#) it has been shown that the solution of the Maxwell equations for the radiated fields $E_{\beta,p\pm}$ and $E_{\beta,s}$ with $\mathbf{P}(\mathbf{r}, t)$ as a source can be written, at points $z \neq 0$, as

$$(E_{\beta,p\pm}, E_{\beta,s}) = \left(\frac{2\pi i \tilde{\omega}^2}{w_\beta} \hat{\mathbf{p}}_{\beta\pm} \cdot \mathcal{P}, \frac{2\pi i \tilde{\omega}^2}{w_\beta} \hat{\mathbf{s}} \cdot \mathcal{P} \right), \quad (4) \quad \text{r2}$$

where $\hat{\mathbf{s}}$ and $\hat{\mathbf{p}}_{\beta\pm}$ are the unitary vectors for the s and p polarization of the radiated field, respectively, and the \pm refers to upward (+) or downward (−) direction of propagation within medium β , as shown in Fig. [I](#),^{[3layer](#)} and $\tilde{\omega} = \omega/c$. Also,

$$w_{\beta}(\omega) = \tilde{\omega}(\epsilon_{\beta}(\omega) - \sin^2 \theta_0)^{1/2}, \quad (5) \quad \boxed{\text{r3}}$$

where θ_0 is the angle of incidence of $\mathbf{E}_v(\omega)$, and

$$\hat{\mathbf{p}}_{\beta\pm}(\omega) = \frac{\kappa(\omega)\hat{\mathbf{z}} \mp w_{\beta}(\omega)\hat{\boldsymbol{\kappa}}}{\tilde{\omega}n_{\beta}(\omega)}, \quad (6) \quad \boxed{\text{r4}}$$

where $\kappa(\omega) = |\boldsymbol{\kappa}| = \tilde{\omega} \sin \theta_0$, $n_{\beta}(\omega) = \sqrt{\epsilon_{\beta}(\omega)}$ is the index of refraction of medium β , and z is the direction perpendicular to the surface that points towards the vacuum. We chose the plane of incidence along the $\boldsymbol{\kappa}z$ plane, then

$$\hat{\boldsymbol{\kappa}} = \cos \phi \hat{\mathbf{x}} + \sin \phi \hat{\mathbf{y}}, \quad (7) \quad \boxed{\text{mc1}}$$

and

$$\hat{\mathbf{s}} = -\sin \phi \hat{\mathbf{x}} + \cos \phi \hat{\mathbf{y}}, \quad (8) \quad \boxed{\text{mmc2}}$$

where ϕ the angle with respect to the x axis.

In the three-layer model the nonlinear polarization responsible for the second harmonic generation (SHG) is immersed in the thin $\beta = \ell$ layer, and is given by

$$\mathcal{P}_i(2\omega) = \chi_{ijk}(2\omega)E_j(\omega)E_k(\omega), \quad (9) \quad \boxed{\text{tres}}$$

where the tensor $\chi(2\omega)$ is the surface nonlinear dipolar susceptibility and the Cartesian indices i, j, k are summed if repeated. El rollo de la centrosimetria va en la introduccion As it was done in Ref. [mizrahiJOSA88](#),^{[r2](#)} in presenting the results Eq. [\(4\)-\(8\)](#)^{[mmc2](#)} we have taken the polarization sheet (Eq. [\(3\)](#))^{[m31](#)} to be oscillating at some frequency ω . However, in the following we find it convenient to use ω exclusively to denote the fundamental frequency and $\boldsymbol{\kappa}$ to denote the component of the incident wave vector parallel to the surface. Then the nonlinear generated polarization is oscillating at $\Omega = 2\omega$ and will be characterized by a wave vector parallel to the surface $\mathbf{K} = 2\boldsymbol{\kappa}$. We can carry over Eqs. [\(3\)-\(8\)](#)^{[m31](#)}^{[mmc2](#)} simply by replacing the lowercase symbols $(\omega, \tilde{\omega}, \boldsymbol{\kappa}, n_{\beta}, w_{\beta}, \hat{\mathbf{p}}_{\beta\pm}, \hat{\mathbf{s}})$ with uppercase symbols $(\Omega, \tilde{\Omega}, \mathbf{K}, N_{\beta}, W_{\beta}, \hat{\mathbf{P}}_{\beta\pm}, \hat{\mathbf{S}})$, all evaluated at 2ω and we always have $\hat{\mathbf{S}} = \hat{\mathbf{s}}$.

To describe the propagation of the SH field, we see from Fig. [I](#),^{[3layer](#)} that it is refracted at the layer-vacuum interface (ℓv), and multiply reflected from the layer-bulk (ℓb) and layer-vacuum (ℓv)

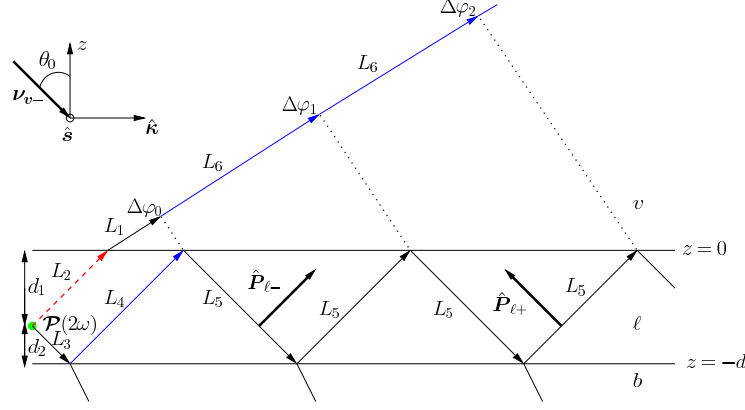


FIG. 1: (color on line) Sketch of the three layer model for SHG. Vacuum (v) is on top with $\epsilon_v = 1$; the layer ℓ , of thickness $d = d_1 + d_2$, is characterized with $\epsilon_\ell(\omega)$, and it is where the SH polarization sheet $\mathcal{P}(2\omega)$ is located at $z_\ell = d_1$; The bulk b is described with $\epsilon_b(\omega)$. The arrows point along the direction of propagation, and the p -polarization unit vector, $\hat{\mathbf{P}}_{\ell-(+)}$, along the downward (upward) direction is denoted with a thick arrow. The s -polarization unit vector $\hat{\mathbf{s}}$, points out of the page. The fundamental field $\mathbf{E}(\omega)$ is incident from the vacuum side along the $z\hat{\mathbf{k}}$ -plane, with θ_0 its angle of incidence and $\boldsymbol{\nu}_{v-}$ its wave vector. $\Delta\varphi_i$ denote the phase difference of the multiply reflected beams with respect to the first vacuum transmitted beam (dashed-red arrow), where the dotted lines are perpendicular to this beam (see the text for details).

3layer

interfaces, thus we can define,

$$\mathbf{T}^{\ell v} = \hat{\mathbf{s}} T_s^{\ell v} \hat{\mathbf{s}} + \hat{\mathbf{P}}_{v+} T_p^{\ell v} \hat{\mathbf{P}}_{\ell+}, \quad (10) \quad \text{r5}$$

as the tensor for transmission from ℓv interface,

$$\mathbf{R}^{\ell b} = \hat{\mathbf{s}} R_s^{\ell b} \hat{\mathbf{s}} + \hat{\mathbf{P}}_{\ell+} R_p^{\ell b} \hat{\mathbf{P}}_{\ell-}, \quad (11) \quad \text{r6}$$

as the tensor of reflection from the ℓb interface, and

$$\mathbf{R}^{\ell v} = \hat{\mathbf{s}} R_s^{\ell v} \hat{\mathbf{s}} + \hat{\mathbf{P}}_{\ell-} R_p^{\ell v} \hat{\mathbf{P}}_{\ell+}, \quad (12) \quad \text{r6b}$$

as that of the ℓv interface. The Fresnel factors in uppercase letters, $T_{s,p}^{ij}$ and $R_{s,p}^{ij}$, are evaluated at 2ω from the following well known formulas

$$\begin{aligned} t_s^{ij}(\omega) &= \frac{2k_i(\omega)}{k_i(\omega) + k_j(\omega)}, & t_p^{ij}(\omega) &= \frac{2k_i(\omega)\sqrt{\epsilon_i(\omega)\epsilon_j(\omega)}}{k_i(\omega)\epsilon_j(\omega) + k_j(\omega)\epsilon_i(\omega)}, \\ r_s^{ij}(\omega) &= \frac{k_i(\omega) - k_j(\omega)}{k_i(\omega) + k_j(\omega)}, & r_p^{ij}(\omega) &= \frac{k_i(\omega)\epsilon_j(\omega) - k_j(\omega)\epsilon_i(\omega)}{k_i(\omega)\epsilon_j(\omega) + k_j(\omega)\epsilon_i(\omega)}. \end{aligned} \quad (13) \quad \text{e.f1}$$

A. Multiple SH reflections

The SH field $\mathbf{E}(2\omega)$ radiated by the SH polarization $\mathcal{P}(2\omega)$ will radiate directly into vacuum and also into the bulk, where it will be reflected back at the thin-layer-bulk interface into the thin

layer again and this beam will be multiple-transmitted and reflected as shown in Fig. [3layer](#)[I](#). As the two beams propagate a phase difference will develop between them, according to

$$\begin{aligned}\Delta\varphi_m &= \tilde{\Omega} \left((L_3 + L_4 + 2mL_5)n_\ell(2\omega) - (L_2n_\ell(2\omega) + (L_1 + mL_6)n_v(2\omega)) \right) \\ &= \delta_0 + m\delta \quad m = 0, 1, 2, \dots,\end{aligned}\tag{14} \quad \boxed{\text{m99}}$$

where

$$\delta_0 = 8\pi \left(\frac{d_2}{\lambda_0} \right) \sqrt{n_\ell^2(2\omega) - \sin^2 \theta_0},\tag{15} \quad \boxed{\text{m97}}$$

$$\delta = 8\pi \left(\frac{d}{\lambda_0} \right) \sqrt{n_\ell^2(2\omega) - \sin^2 \theta_0},\tag{16} \quad \boxed{\text{m96}}$$

where λ_0 is the wavelength of the fundamental field in vacuum, d the thickness of layer ℓ and d_2 the distance of $\mathcal{P}(2\omega)$ from the ℓb interface (see Fig. [3layer](#)[I](#)). We see that δ_0 is the phase difference of the first and second transmitted beams, and $m\delta$ that of the first and third ($m = 1$), fourth ($m = 2$), etc. beams (see Fig. [3layer](#)[I](#)).

To take into account the multiple reflections of the generated SH field in the layer ℓ , we proceed as follows. We show the algebra for the p -polarized SH field, the s -polarized field could be worked out along the same steps. The multiple-reflected $\mathbf{E}_p(2\omega)$ field is given by

$$\begin{aligned}\mathbf{E}(2\omega) &= E_{p+}(2\omega)\mathbf{T}^{\ell v} \cdot \hat{\mathbf{P}}_{\ell+} + E_{p-}(2\omega)\mathbf{T}^{\ell v} \cdot \mathbf{R}^{\ell b} \cdot \hat{\mathbf{P}}_{\ell-}e^{i\Delta\varphi_0} + E_{p-}(2\omega)\mathbf{T}^{\ell v} \cdot \mathbf{R}^{\ell b} \cdot \mathbf{R}^{\ell v} \cdot \mathbf{R}^{\ell b} \cdot \hat{\mathbf{P}}_{\ell-}e^{i\Delta\varphi_1} \\ &\quad + E_{p-}(2\omega)\mathbf{T}^{\ell v} \cdot \mathbf{R}^{\ell b} \cdot \mathbf{R}^{\ell v} \cdot \mathbf{R}^{\ell b} \cdot \mathbf{R}^{\ell v} \cdot \mathbf{R}^{\ell b} \cdot \hat{\mathbf{P}}_{\ell-}e^{i\Delta\varphi_2} + \dots \\ &= E_{p+}(2\omega)\mathbf{T}^{\ell v} \cdot \hat{\mathbf{P}}_{\ell+} + E_{p-}(2\omega)\mathbf{T}^{\ell v} \cdot \sum_{m=0}^{\infty} (\mathbf{R}^{\ell b} \cdot \mathbf{R}^{\ell v} e^{i\delta})^m \cdot \mathbf{R}^{\ell b} \cdot \hat{\mathbf{P}}_{\ell-}e^{i\delta_0}.\end{aligned}$$

From Eqs. [\(r5\)](#)[\(f10\)](#) and [\(r6\)](#)[\(f11\)](#) is easy to show that

$$\mathbf{T}^{\ell v} \cdot (\mathbf{R}^{\ell b} \cdot \mathbf{R}^{\ell v})^n \cdot \mathbf{R}^{\ell b} = \hat{\mathbf{s}}T_s^{\ell v} \left(R_s^{\ell b} R_s^{\ell v} \right)^n R_s^{\ell b} \hat{\mathbf{s}} + \hat{\mathbf{P}}_{v+}T_p^{\ell v} \left(R_p^{\ell b} R_p^{\ell v} \right)^n R_p^{\ell b} \hat{\mathbf{P}}_{\ell-},\tag{17} \quad \boxed{\text{m1}}$$

then,

$$\mathbf{E}(2\omega) = \hat{\mathbf{P}}_{\ell+}T_p^{\ell v} \left(E_{p+}(2\omega) + \frac{R_p^{\ell b} e^{i\delta_0}}{1 + R_p^{\ell v} R_p^{\ell b} e^{i\delta}} E_{p-}(2\omega) \right),\tag{18} \quad \boxed{\text{m7}}$$

where we used $R_{s,p}^{ij} = -R_{s,p}^{ji}$. Using Eq. [\(r2\)](#)[\(f4\)](#), we can readily write

$$\mathbf{E}(2\omega) = \frac{2\pi i \tilde{\Omega}}{K_\ell} \mathbf{H}_\ell \cdot \mathcal{P}(2\omega),\tag{19} \quad \boxed{\text{mr8}}$$

where

$$\mathbf{H}_\ell = \hat{\mathbf{s}}T_s^{\ell v} (1 + R_s^M) \hat{\mathbf{s}} + \hat{\mathbf{P}}_{v+}T_p^{\ell v} (\hat{\mathbf{P}}_{\ell+} + R_p^M \hat{\mathbf{P}}_{\ell-}).\tag{20} \quad \boxed{\text{mr9}}$$

and

$$R_l^M \equiv \frac{R_l^{\ell b} e^{i\delta_0}}{1 + R_l^{v\ell} R_l^{\ell b} e^{i\delta}} \quad l = s, p, \quad (21) \quad \boxed{\text{m61}}$$

is defined as the multiple reflection coefficient. To make touch with the work of Ref. [mizrahiJOSA88](#) where $\mathcal{P}(2\omega)$ is located on top of the vacuum-surface interface and only the vacuum radiated beam and the first (and only) reflected beam need to be considered, we take $\ell = v$ and $d_2 = 0$, then $T^{\ell v} = 1$, $R^{v\ell} = 0$ and $\delta_0 = 0$, with which $R_l^M = R_l^{vb}$. Thus, Eq. [\(20\)](#) coincides with Eq. (3.8) of Ref. [mizrahiJOSA88](#).

B. Radiation Terms

The magnitude of the radiated field is given by $E(2\omega) = \hat{\mathbf{e}}^{\text{out}} \cdot \mathbf{E}(2\omega)$, where $\hat{\mathbf{e}}^{\text{out}}$ is the polarization vector of the radiated field, for instance $\hat{\mathbf{s}}$ or $\hat{\mathbf{P}}_{v+}$. Then, we write

$$\begin{aligned} \hat{\mathbf{P}}_{\ell+} + R_p^{\ell b} \hat{\mathbf{P}}_{\ell-} &= \frac{\sin \theta_{\text{in}} \hat{\mathbf{z}} - K_\ell \hat{\mathbf{x}}}{\sqrt{\epsilon_\ell(2\omega)}} + R_p^{\ell b} \frac{\sin \theta_{\text{in}} \hat{\mathbf{z}} + K_\ell \hat{\mathbf{x}}}{\sqrt{\epsilon_\ell(2\omega)}} \\ &= \frac{1}{\sqrt{\epsilon_\ell(2\omega)}} \left(\sin \theta_{\text{in}} (1 + R_p^{\ell b}) \hat{\mathbf{z}} - K_\ell (1 - R_p^{\ell b}) \hat{\mathbf{x}} \right) \\ &= \frac{T_p^{\ell b}}{\epsilon_\ell(2\omega) \sqrt{\epsilon_b(2\omega)}} (\epsilon_b(2\omega) \sin \theta_{\text{in}} \hat{\mathbf{z}} - \epsilon_\ell(2\omega) K_b \hat{\mathbf{x}}), \end{aligned}$$

where using

$$\begin{aligned} 1 + R_s^{\ell b} &= T_s^{\ell b} \\ 1 + R_p^{\ell b} &= \sqrt{\frac{\epsilon_b(2\omega)}{\epsilon_\ell(2\omega)}} T_p^{\ell b} \\ 1 - R_p^{\ell b} &= \sqrt{\frac{\epsilon_\ell(2\omega)}{\epsilon_b(2\omega)}} \frac{K_b}{K_\ell} T_p^{\ell b} \\ T_p^{\ell v} &= \frac{K_\ell}{K_v} T_p^{v\ell} \\ T_s^{\ell v} &= \frac{K_\ell}{K_v} T_s^{v\ell}, \end{aligned} \quad (22)$$

we can write

$$E(2\omega) = \frac{4\pi i \omega}{c K_v} \hat{\mathbf{e}}^{\text{out}} \cdot \mathbf{H}_\ell \cdot \mathcal{P}(2\omega) = \frac{4\pi i \omega}{c K_v} \mathbf{e}_\ell^{2\omega} \cdot \mathcal{P}(2\omega). \quad (23) \quad \boxed{\text{r10}}$$

where,

$$\begin{aligned} \mathbf{e}_\ell^{2\omega} &= \\ \hat{\mathbf{e}}^{\text{out}} \cdot &\left[\hat{\mathbf{s}} T_s^{v\ell} T_s^{\ell b} \hat{\mathbf{s}} + \hat{\mathbf{P}}_{v+} \frac{T_p^{v\ell} T_p^{\ell b}}{\epsilon_\ell(2\omega) \sqrt{\epsilon_b(2\omega)}} (\epsilon_b(2\omega) \sin \theta_{\text{in}} \hat{\mathbf{z}} - \epsilon_\ell(2\omega) K_b \hat{\mathbf{x}}) \right]. \end{aligned} \quad (24) \quad \boxed{\text{r12}}$$

We pause here to reduce above result to the case where the nonlinear polarization $\mathbf{P}(2\omega)$ radiates from vacuum instead from the layer ℓ . For such case we simply take $\epsilon_\ell(2\omega) = 1$ and $\ell = v$ ($T_{s,p}^{\ell v} = 1$), to get

$$\mathbf{e}_v^{2\omega} = \hat{\mathbf{e}}^{\text{out}} \cdot \left[\hat{\mathbf{s}} T_s^{vb} \hat{\mathbf{s}} + \hat{\mathbf{P}}_{v+} \frac{T_p^{vb}}{\sqrt{\epsilon_b(2\omega)}} (\epsilon_b(2\omega) \sin \theta_{\text{in}} \hat{\mathbf{z}} - K_b \hat{\mathbf{x}}) \right], \quad (25) \quad \boxed{\text{r13}}$$

which agrees with Eq. (3.8) of Ref. [mizrahiJOSA88](#).

In the three layer model the nonlinear polarization is located in layer ℓ , and then we evaluate the fundamental field required in Eq. [\(9\)](#) in this layer as well, then we write

$$\mathbf{E}_\ell(\omega) = E_0 \left(\hat{\mathbf{s}} t_s^{v\ell} (1 + r_s^{\ell b}) \hat{\mathbf{s}} + \hat{\mathbf{p}}_{\ell-} t_p^{v\ell} \hat{\mathbf{p}}_{v-} + \hat{\mathbf{p}}_{\ell+} t_p^{v\ell} r_p^{\ell b} \hat{\mathbf{p}}_{v-} \right) \cdot \hat{\mathbf{e}}^{\text{in}} = E_0 \mathbf{e}_\ell^\omega, \quad (26) \quad \boxed{\text{m2}}$$

and following the steps that lead to Eq. [\(24\)](#), we find that

$$\mathbf{e}_\ell^\omega = \left[\hat{\mathbf{s}} t_s^{v\ell} t_s^{\ell b} \hat{\mathbf{s}} + \frac{t_p^{v\ell} t_p^{\ell b}}{\epsilon_\ell(\omega) \sqrt{\epsilon_b(\omega)}} (\epsilon_b(\omega) \sin \theta_{\text{in}} \hat{\mathbf{z}} + \epsilon_\ell(\omega) k_b \hat{\mathbf{x}}) \hat{\mathbf{p}}_{v-} \right] \cdot \hat{\mathbf{e}}^{\text{in}}. \quad (27) \quad \boxed{\text{m12}}$$

If we would like to evaluate the fields in the bulk, instead of the layer ℓ , we simply take $\epsilon_\ell(\omega) = \epsilon_b(\omega)$ ($t_{s,p}^{\ell b} = 1$), to obtain

$$\mathbf{e}_b^\omega = \left[\hat{\mathbf{s}} t_s^{vb} \hat{\mathbf{s}} + \frac{t_p^{vb}}{\sqrt{\epsilon_b(\omega)}} (\sin \theta_{\text{in}} \hat{\mathbf{z}} + k_b \hat{\mathbf{x}}) \hat{\mathbf{p}}_{v-} \right] \cdot \hat{\mathbf{e}}^{\text{in}}, \quad (28) \quad \boxed{\text{m13}}$$

that is in agreement with Eq. (3.5) of Ref. [mizrahiJOSA88](#).

With \mathbf{e}^ω we can write Eq. [\(9\)](#) as

$$\mathcal{P}(2\omega) = E_0^2 \chi : \mathbf{e}_\ell^\omega \mathbf{e}_\ell^\omega, \quad (29) \quad \boxed{\text{m4}}$$

and then from Eq. [\(23\)](#) we obtain that

$$\begin{aligned} |E(2\omega)|^2 &= |E_0|^4 \frac{16\pi^2 \omega^2}{c^2 K_v^2} |\mathbf{e}_\ell^{2\omega} \cdot \chi : \mathbf{e}_\ell^\omega \mathbf{e}_\ell^\omega|^2 \\ \frac{c}{2\pi} |E(2\omega)|^2 &= \frac{32\pi^3 \omega^2}{c^3 \cos^2 \theta_{\text{in}}} |\mathbf{e}_\ell^{2\omega} \cdot \chi : \mathbf{e}_\ell^\omega \mathbf{e}_\ell^\omega|^2 \left(\frac{c}{2\pi} |E_0|^2 \right)^2, \\ I(2\omega) &= \frac{32\pi^3 \omega^2}{c^3 \cos^2 \theta_{\text{in}}} |\mathbf{e}_\ell^{2\omega} \cdot \chi : \mathbf{e}_\ell^\omega \mathbf{e}_\ell^\omega|^2 I^2(\omega), \\ R(2\omega) &= \frac{32\pi^3 \omega^2}{c^3 \cos^2 \theta_{\text{in}}} |\mathbf{e}_\ell^{2\omega} \cdot \chi : \mathbf{e}_\ell^\omega \mathbf{e}_\ell^\omega|^2, \end{aligned} \quad (30) \quad \boxed{\text{r01}}$$

as the SHG yield. At this point we mention that to recover the results of Ref. [mizrahiJOSA88](#) which are equivalent of those of Ref. [sipePRB87](#), we take $\mathbf{e}_\ell^{2\omega} \rightarrow \mathbf{e}_v^{2\omega}$, $\mathbf{e}_\ell^\omega \rightarrow \mathbf{e}_b^\omega$ and then

$$R(2\omega) = \frac{32\pi^3 \omega^2}{c^3 \cos^2 \theta_{\text{in}}} |\mathbf{e}_v^{2\omega} \cdot \chi : \mathbf{e}_b^\omega \mathbf{e}_b^\omega|^2, \quad (31) \quad \boxed{\text{m69}}$$

will give the SHG yield of a nonlinear polarization sheet radiating from vacuum on top of the surface and where the fundamental field is evaluated below the surface that is characterized by $\epsilon_b(\omega)$.

C. One SH Reflection

Therefore, the total radiated field at 2ω is

$$\begin{aligned} \mathbf{E}(2\omega) = & E_s(2\omega) \left(\mathbf{T}^{\ell v} + \mathbf{T}^{\ell v} \cdot \mathbf{R}^{\ell b} \right) \cdot \hat{\mathbf{s}} \\ & + E_{p+}(2\omega) \mathbf{T}^{\ell v} \cdot \hat{\mathbf{P}}_{\ell+} + E_{p-}(2\omega) \mathbf{T}^{\ell v} \cdot \mathbf{R}^{\ell b} \cdot \hat{\mathbf{P}}_{\ell-}. \end{aligned}$$

The first term is the transmitted s -polarized field, the second one is the reflected and then transmitted s -polarized field and the third and fourth terms are the equivalent fields for p -polarization. The transmission is from the layer into vacuum, and the reflection between the layer and the bulk. After some simple algebra, we obtain

$$\mathbf{E}(2\omega) = \frac{2\pi i \tilde{\Omega}}{K_\ell} \mathbf{H}_\ell \cdot \mathcal{P}(2\omega), \quad (32) \quad \boxed{\text{r8}}$$

where,

$$\mathbf{H}_\ell = \hat{\mathbf{s}} T_s^{\ell v} \left(1 + R_s^{\ell b} \right) \hat{\mathbf{s}} + \hat{\mathbf{P}}_{v+} T_p^{\ell v} \left(\hat{\mathbf{P}}_{\ell+} + R_p^{\ell b} \hat{\mathbf{P}}_{\ell-} \right). \quad (33) \quad \boxed{\text{r9}}$$

II. \mathcal{R} FOR DIFFERENT POLARIZATION CASES

We obtain explicit relations for a C_{3v} symmetry characteristic of a (111) surface, for which the only components of χ_{ijk} different from zero are χ_{zzz} , $\chi_{zxx} = \chi_{zyy}$, $\chi_{xxz} = \chi_{yyz}$ and $\chi_{xxx} = -\chi_{xyy} = -\chi_{yyx}$ with $\chi_{ijk} = \chi_{ikj}$, where we have chosen the x and y axes along the [112] and [110] directions, respectively.

However, we have to remember that the plane of incidence so far was chosen to be the xz plane; the most general plane of incidence should be one that makes an angle ϕ with respect to the x axis, and so $\hat{\mathbf{x}}$ should to be replaced by a unit vector $\hat{\mathbf{\kappa}}$ such that

$$\hat{\mathbf{\kappa}} = \cos \phi \hat{\mathbf{x}} + \sin \phi \hat{\mathbf{y}}, \quad (34) \quad \boxed{\text{mc1}}$$

and then

$$\hat{\mathbf{s}} = -\sin \phi \hat{\mathbf{x}} + \cos \phi \hat{\mathbf{y}}, \quad (35) \quad \boxed{\text{mc2}}$$

A. \mathcal{R}_{pP}

We develop five different scenarios for \mathcal{R}_{pP} that explore different cases for where the polarization and fundamental fields are located. In all these scenarios, we use $\hat{\mathbf{e}}^{\text{in}} = \hat{\mathbf{p}}_{v-}$ in Eq. [\(27\)](#)^{m12}, and $\hat{\mathbf{e}}^{\text{out}} = \hat{\mathbf{P}}_{v+}$ in Eq. [\(24\)](#)^{r12}.

1. Three layer model

This scenario involves $\mathcal{P}(2\omega)$ and the fundamental fields to be taken in a thin layer of material below the surface, which we designate as ℓ . Thus,

$$\mathbf{e}_\ell^{2\omega} \cdot \boldsymbol{\chi} : \mathbf{e}_\ell^\omega \mathbf{e}_\ell^\omega \equiv \Gamma_{pP}^\ell r_{pP}^\ell, \quad (36) \quad \boxed{\text{m80}}$$

where

$$\begin{aligned} r_{pP}^\ell &= \epsilon_b(2\omega) \sin \theta_{\text{in}} \left(\epsilon_b^2(\omega) \sin^2 \theta_{\text{in}} \chi_{zzz} + \epsilon_\ell^2(\omega) k_b^2 \chi_{zxx} \right) \\ &\quad - \epsilon_\ell(2\omega) \epsilon_\ell(\omega) k_b K_b \left(2\epsilon_b(\omega) \sin \theta_{\text{in}} \chi_{xxz} + \epsilon_\ell(\omega) k_b \chi_{xxx} \cos(3\phi) \right), \end{aligned} \quad (37) \quad \boxed{\text{m81}}$$

and

$$\Gamma_{pP}^\ell = \frac{T_p^{\ell v} T_p^{\ell b}}{\epsilon_\ell(2\omega) \sqrt{\epsilon_b(2\omega)}} \left(\frac{t_p^{v\ell} t_p^{\ell b}}{\epsilon_\ell(\omega) \sqrt{\epsilon_b(\omega)}} \right)^2. \quad (38) \quad \boxed{\text{m79}}$$

2. Two layer model

In order to reduce above result to that of Ref. [\[13\]](#) and [\[14\]](#), we now consider that $\mathcal{P}(2\omega)$ is evaluated in the vacuum region, while the fundamental fields are evaluated in the bulk region. To do this, we take the 2ω radiations factors for vacuum by taking $\ell = v$, thus $\epsilon_\ell(2\omega) = 1$, $T_p^{\ell v} = 1$, $T_p^{\ell b} = T_p^{vb}$, and the fundamental field inside medium b by taking $\ell = b$, thus $\epsilon_\ell(\omega) = \epsilon_b(\omega)$, $t_p^{v\ell} = t_p^{vb}$, and $t_p^{\ell b} = 1$. With these choices

$$\mathbf{e}_v^{2\omega} \cdot \boldsymbol{\chi} : \mathbf{e}_b^\omega \mathbf{e}_b^\omega \equiv \Gamma_{pP}^{vb} r_{pP}^{vb}, \quad (39) \quad \boxed{\text{m800}}$$

where,

$$\begin{aligned} r_{pP}^{vb} &= \epsilon_b(2\omega) \sin \theta_{\text{in}} \left(\sin^2 \theta_{\text{in}} \chi_{zzz} + k_b^2 \chi_{zxx} \right) \\ &\quad - k_b K_b \left(2 \sin \theta_{\text{in}} \chi_{xxz} + k_b \chi_{xxx} \cos(3\phi) \right), \end{aligned} \quad (40) \quad \boxed{\text{m82}}$$

and

$$\Gamma_{pP}^{vb} = \frac{T_p^{vb} (t_p^{vb})^2}{\epsilon_b(\omega) \sqrt{\epsilon_b(2\omega)}}. \quad (41) \quad \boxed{\text{m78}}$$

3. Taking $\mathcal{P}(2\omega)$ and the fundamental fields in the bulk

To evaluate the 2ω fields in the bulk, we take Eq. [\(33\)](#) considering that $\ell \rightarrow b$. We have already considered the 1ω fields in the bulk in Eq. [\(28\)](#). After some algebra, we get that

$$\mathbf{e}_b^{2\omega} \cdot \boldsymbol{\chi} : \mathbf{e}_b^\omega \mathbf{e}_b^\omega = \Gamma_{pP}^b r_{pP}^b \quad (42)$$

where

$$r_{pP}^b = \sin^3 \theta_{\text{in} \chi_{zzz}} + k_b^2 \sin \theta_{\text{in} \chi_{zxx}} - 2k_b K_b \sin \theta_{\text{in} \chi_{xxz}} - k_b^2 K_b \chi_{xxx} \cos 3\phi, \quad (43)$$

and

$$\Gamma_{pP}^b = \frac{T_p^{vb} (t_p^{vb})^2}{\epsilon_b(\omega) \sqrt{\epsilon_b(2\omega)}}. \quad (44)$$

4. Taking $\mathcal{P}(2\omega)$ and the fundamental fields in the vacuum

To evaluate the 1ω fields in the vacuum, we take Eq. (26) considering that $\ell \rightarrow v$. We have already considered the 2ω fields in the vacuum in Eq. (25). After some algebra, we get that

$$\mathbf{e}_v^{2\omega} \cdot \boldsymbol{\chi} : \mathbf{e}_v^\omega \mathbf{e}_v^\omega = \Gamma_{pP}^v r_{pP}^v \quad (45)$$

where

$$\begin{aligned} r_{pP}^v &= \epsilon_b^2(\omega) \epsilon_b(2\omega) \sin^3 \theta_{\text{in} \chi_{zzz}} + \epsilon_b(2\omega) k_b^2 \sin \theta_{\text{in} \chi_{zxx}} \\ &\quad - 2\epsilon_b(\omega) k_b K_b \sin \theta_{\text{in} \chi_{xxz}} - k_b^2 K_b \chi_{xxx} \cos 3\phi \end{aligned} \quad (46)$$

and

$$\Gamma_{pP}^v = \frac{T_p^{vb} (t_p^{vb})^2}{\epsilon_b(\omega) \sqrt{\epsilon_b(2\omega)}}. \quad (47)$$

5. Taking $\mathcal{P}(2\omega)$ in ℓ and the fundamental fields in the bulk

For this scenario, we have

$$\mathbf{e}_\ell^{2\omega} \cdot \boldsymbol{\chi} : \mathbf{e}_b^\omega \mathbf{e}_b^\omega = \Gamma_{pP}^{\ell b} r_{pP}^{\ell b} \quad (48)$$

where

$$\begin{aligned} r_{pP}^{\ell b} &= \epsilon_b(2\omega) \sin^3 \theta_{\text{in} \chi_{zzz}} + \epsilon_b(2\omega) k_b^2 \sin \theta_{\text{in} \chi_{zxx}} \\ &\quad - 2\epsilon_\ell(2\omega) k_b K_b \sin \theta_{\text{in} \chi_{xxz}} - \epsilon_\ell(2\omega) k_b^2 K_b \chi_{xxx} \cos 3\phi, \end{aligned} \quad (49)$$

and

$$\Gamma_{pP}^{\ell b} = \frac{T_p^{v\ell} T_p^{\ell b} (t_p^{vb})^2}{\epsilon_\ell(2\omega) \epsilon_b(\omega) \sqrt{\epsilon_b(2\omega)}}. \quad (50)$$

B. \mathcal{R}_{pS}

To obtain $R_{pS}(2\omega)$ we use $\hat{\mathbf{e}}^{\text{in}} = \hat{\mathbf{p}}_{v-}$ in Eq. (m12), and $\hat{\mathbf{e}}^{\text{out}} = \hat{\mathbf{S}}$ in Eq. (r12). We also use the unit vectors defined in Eqs. (mc1) and (mc2). Substituting, we get

$$\mathbf{e}_\ell^{2\omega} \cdot \boldsymbol{\chi} : \mathbf{e}_\ell^\omega \mathbf{e}_\ell^\omega \equiv \Gamma_{sP}^\ell r_{sP}^\ell, \quad (51)$$

where

$$r_{pS}^\ell = -\epsilon_\ell^2(\omega) k_b^2 \sin 3\phi \chi_{xxx}, \quad (52)$$

and

$$\Gamma_{pS}^\ell = T_s^{v\ell} T_s^{\ell b} \left(\frac{t_p^{v\ell} t_p^{\ell b}}{\epsilon_\ell(\omega) \sqrt{\epsilon_b(\omega)}} \right)^2. \quad (53)$$

In order to reduce above result to that of Ref. [mizrahiJOSABPRB87], we take the 2ω radiations factors for vacuum by taking $\ell = v$, thus $\epsilon_\ell(2\omega) = 1$, $T_s^{v\ell} = 1$, $T_s^{\ell b} = T_s^{vb}$, and the fundamental field inside medium b by taking $\ell = b$, thus $\epsilon_\ell(\omega) = \epsilon_b(\omega)$, $t_p^{v\ell} = t_p^{vb}$, and $t_p^{\ell b} = 1$. With these choices,

$$r_{pS}^b = -k_b^2 \sin 3\phi \chi_{xxx}, \quad (54)$$

and

$$\Gamma_{pS}^b = T_s^{vb} \left(\frac{t_p^{vb}}{\sqrt{\epsilon_b(\omega)}} \right)^2. \quad (55)$$

C. \mathcal{R}_{sP}

To obtain $R_{sP}(2\omega)$ we use $\hat{\mathbf{e}}^{\text{in}} = \hat{\mathbf{s}}$ in Eq. (m12), and $\hat{\mathbf{e}}^{\text{out}} = \hat{\mathbf{P}}_{v+}$ in Eq. (r12). We also use the unit vectors defined in Eqs. (mc1) and (mc2). Substituting, we get

$$\mathbf{e}_\ell^{2\omega} \cdot \boldsymbol{\chi} : \mathbf{e}_\ell^\omega \mathbf{e}_\ell^\omega \equiv \Gamma_{sP}^\ell r_{sP}^\ell, \quad (56)$$

where

$$r_{sP}^\ell = \epsilon_b(2\omega) \sin \theta_{\text{in}} \chi_{zzx} + \epsilon_\ell(2\omega) K_b \chi_{xxx} \cos 3\phi, \quad (57)$$

and

$$\Gamma_{sP}^\ell = \frac{T_p^{\ell v} T_p^{\ell b} (t_s^{v\ell} t_s^{\ell b})^2}{\epsilon_\ell(2\omega) \sqrt{\epsilon_b(2\omega)}}. \quad (58)$$

In order to reduce above result to that of Ref. [mizrahiJOS488PRB87](#) and [\[?\]](#), we take the 2ω radiations factors for vacuum by taking $\ell = v$, thus $\epsilon_\ell(2\omega) = 1$, $T_p^{v\ell} = 1$, $T_p^{\ell b} = T_p^{vb}$, and the fundamental field inside medium b by taking $\ell = b$, thus $\epsilon_\ell(\omega) = \epsilon_b(\omega)$, $t_s^{v\ell} = t_s^{vb}$, and $t_s^{\ell b} = 1$. With these choices,

$$r_{sP}^b = \epsilon_b(2\omega) \sin \theta_{\text{in}} \chi_{zzx} + K_b \chi_{xxx} \cos 3\phi, \quad (59)$$

and

$$\Gamma_{sP}^b = \frac{T_p^{vb} (t_s^{vb})^2}{\sqrt{\epsilon_b(2\omega)}}. \quad (60)$$

D. \mathcal{R}_{sS}

For \mathcal{R}_{sS} we have that $\hat{\mathbf{e}}^{\text{in}} = \hat{\mathbf{s}}$ and $\hat{\mathbf{e}}^{\text{out}} = \hat{\mathbf{S}}$. This leads to

$$\mathbf{e}_\ell^{2\omega} \cdot \boldsymbol{\chi} : \mathbf{e}_\ell^\omega \mathbf{e}_\ell^\omega \equiv \Gamma_{sS}^\ell r_{sS}^\ell, \quad (61)$$

where

$$r_{sS}^\ell = \chi_{xxx} \sin 3\phi, \quad (62)$$

and

$$\Gamma_{sS}^\ell = T_s^{v\ell} T_s^{\ell b} \left(t_s^{v\ell} t_s^{\ell b} \right)^2. \quad (63)$$

In order to reduce above result to that of Ref. [mizrahiJOS488PRB87](#) and [\[?\]](#), we take the 2ω radiations factors for vacuum by taking $\ell = v$, thus $\epsilon_\ell(2\omega) = 1$, $T_s^{v\ell} = 1$, $T_s^{\ell b} = T_s^{vb}$, and the fundamental field inside medium b by taking $\ell = b$, thus $\epsilon_\ell(\omega) = \epsilon_b(\omega)$, $t_s^{v\ell} = t_s^{vb}$, and $t_s^{\ell b} = 1$. With these choices,

$$r_{sS}^b = \chi_{xxx} \sin 3\phi, \quad (64)$$

and

$$\Gamma_{sS}^b = T_s^{vb} \left(t_s^{vb} \right)^2. \quad (65)$$

Appendix A: Full derivations for \mathcal{R} for different polarization cases

1. \mathcal{R}_{pP}

a. Taking $\mathcal{P}(2\omega)$ and the fundamental fields in the bulk

To consider the 2ω fields in the bulk, we start with Eq. (53) but substitute $\ell \rightarrow b$, thus

$$\mathbf{H}_b = \hat{\mathbf{s}} T_s^{bv} (1 + R_s^{bb}) \hat{\mathbf{s}} + \hat{\mathbf{P}}_{v+} T_p^{bv} (\hat{\mathbf{P}}_{b+} + R_p^{bb} \hat{\mathbf{P}}_{b-}).$$

R_p^{bb} and R_s^{bb} are zero, so we are left with

$$\begin{aligned} \mathbf{H}_b &= \hat{\mathbf{s}} T_s^{bv} \hat{\mathbf{s}} + \hat{\mathbf{P}}_{v+} T_p^{bv} \hat{\mathbf{P}}_{b+} \\ &= \frac{K_b}{K_v} \left(\hat{\mathbf{s}} T_s^{vb} \hat{\mathbf{s}} + \hat{\mathbf{P}}_{v+} T_p^{vb} \hat{\mathbf{P}}_{b+} \right) \\ &= \frac{K_b}{K_v} \left[\hat{\mathbf{s}} T_s^{vb} \hat{\mathbf{s}} + \hat{\mathbf{P}}_{v+} \frac{T_p^{vb}}{\sqrt{\epsilon_b(2\omega)}} (\sin \theta_{\text{in}} \hat{\mathbf{z}} - K_b \cos \phi \hat{\mathbf{x}} - K_b \sin \phi \hat{\mathbf{y}}) \right], \end{aligned}$$

and we define

$$\mathbf{e}_b^{2\omega} = \frac{K_b}{K_v} \hat{\mathbf{e}}^{\text{out}} \cdot \left[\hat{\mathbf{s}} T_s^{vb} \hat{\mathbf{s}} + \hat{\mathbf{P}}_{v+} \frac{T_p^{vb}}{\sqrt{\epsilon_b(2\omega)}} (\sin \theta_{\text{in}} \hat{\mathbf{z}} - K_b \cos \phi \hat{\mathbf{x}} - K_b \sin \phi \hat{\mathbf{y}}) \right].$$

For \mathcal{R}_{pP} , we require $\hat{\mathbf{e}}^{\text{out}} = \hat{\mathbf{P}}_{v+}$, so we have that

$$\mathbf{e}_b^{2\omega} = \frac{K_b}{K_v} \frac{T_p^{vb}}{\sqrt{\epsilon_b(2\omega)}} (\sin \theta_{\text{in}} \hat{\mathbf{z}} - K_b \cos \phi \hat{\mathbf{x}} - K_b \sin \phi \hat{\mathbf{y}}).$$

The 1ω fields will still be evaluated inside the bulk, so we have Eq. (28)

$$\mathbf{e}_b^\omega = \left[\hat{\mathbf{s}} t_s^{vb} \hat{\mathbf{s}} + \frac{t_p^{vb}}{\sqrt{\epsilon_b(\omega)}} (\sin \theta_{\text{in}} \hat{\mathbf{z}} + k_b \cos \phi \hat{\mathbf{x}} + k_b \sin \phi \hat{\mathbf{y}}) \hat{\mathbf{P}}_{v-} \right] \cdot \hat{\mathbf{e}}^{\text{in}},$$

and for our particular case of $\hat{\mathbf{e}}^{\text{in}} = \hat{\mathbf{P}}_{v-}$,

$$\mathbf{e}_b^\omega = \frac{t_p^{vb}}{\sqrt{\epsilon_b(\omega)}} (\sin \theta_{\text{in}} \hat{\mathbf{z}} + k_b \cos \phi \hat{\mathbf{x}} + k_b \sin \phi \hat{\mathbf{y}}),$$

and

$$\begin{aligned} \mathbf{e}_b^\omega \mathbf{e}_b^\omega &= \frac{(t_p^{vb})^2}{\epsilon_b(\omega)} (\sin \theta_{\text{in}} \hat{\mathbf{z}} + k_b \cos \phi \hat{\mathbf{x}} + k_b \sin \phi \hat{\mathbf{y}})^2 \\ &= \frac{(t_p^{vb})^2}{\epsilon_b(\omega)} (\sin^2 \theta_{\text{in}} \hat{\mathbf{z}} \hat{\mathbf{z}} + k_b^2 \cos^2 \phi \hat{\mathbf{x}} \hat{\mathbf{x}} + k_b^2 \sin^2 \phi \hat{\mathbf{y}} \hat{\mathbf{y}} \\ &\quad + 2k_b \sin \theta_{\text{in}} \cos \phi \hat{\mathbf{z}} \hat{\mathbf{x}} + 2k_b \sin \theta_{\text{in}} \sin \phi \hat{\mathbf{z}} \hat{\mathbf{y}} + 2k_b^2 \sin \phi \cos \phi \hat{\mathbf{x}} \hat{\mathbf{y}}) \end{aligned}$$

So lastly, we have that

$$\begin{aligned}
\mathbf{e}_b^{2\omega} \cdot \boldsymbol{\chi} : \mathbf{e}_b^\omega \mathbf{e}_b^\omega &= \frac{K_b}{K_v} \frac{T_p^{vb} (t_p^{vb})^2}{\epsilon_b(\omega) \sqrt{\epsilon_b(2\omega)}} \left(\sin^3 \theta_{\text{in}} \chi_{zzz} \right. \\
&\quad + k_b^2 \sin \theta_{\text{in}} \cos^2 \phi \chi_{zxx} \\
&\quad + k_b^2 \sin \theta_{\text{in}} \sin^2 \phi \chi_{zyy} \\
&\quad + 2k_b \sin^2 \theta_{\text{in}} \cos \phi \chi_{zzx} \\
&\quad + 2k_b \sin^2 \theta_{\text{in}} \sin \phi \chi_{zzy} \\
&\quad + 2k_b^2 \sin \theta_{\text{in}} \sin \phi \cos \phi \chi_{zxy} \\
&\quad - K_b \sin^2 \theta_{\text{in}} \cos \phi \chi_{xzz} \\
&\quad - k_b^2 K_b \cos^3 \phi \chi_{xxx} \\
&\quad - k_b^2 K_b \sin^2 \phi \cos \phi \chi_{xyy} \\
&\quad - 2k_b K_b \sin \theta_{\text{in}} \cos^2 \phi \chi_{xzx} \\
&\quad - 2k_b K_b \sin \theta_{\text{in}} \sin \phi \cos \phi \chi_{xzy} \\
&\quad - 2k_b^2 K_b \sin \phi \cos^2 \phi \chi_{xxy} \\
&\quad - K_b \sin^2 \theta_{\text{in}} \sin \phi \chi_{yzz} \\
&\quad - k_b^2 K_b \sin \phi \cos^2 \phi \chi_{yxx} \\
&\quad - k_b^2 K_b \sin^3 \phi \chi_{yyy} \\
&\quad - 2k_b K_b \sin \theta_{\text{in}} \sin \phi \cos \phi \chi_{yzx} \\
&\quad - 2k_b K_b \sin \theta_{\text{in}} \sin^2 \phi \chi_{yzy} \\
&\quad \left. - 2k_b^2 K_b \sin^2 \phi \cos \phi \chi_{yxy} \right),
\end{aligned}$$

and we can eliminate many terms since $\chi_{zzx} = \chi_{zzy} = \chi_{zxy} = \chi_{xzz} = \chi_{xzy} = \chi_{xxy} = \chi_{yzz} = \chi_{yxx} =$

$\chi_{yyy} = \chi_{yzx} = 0$, and substituting the equivalent components of χ ,

$$\begin{aligned}
&= \frac{K_b}{K_v} \Gamma_{pP}^b \left(\sin^3 \theta_{\text{in}} \chi_{zzz} \right. \\
&\quad + k_b^2 \sin \theta_{\text{in}} \cos^2 \phi \chi_{zxx} \\
&\quad + k_b^2 \sin \theta_{\text{in}} \sin^2 \phi \chi_{zxx} \\
&\quad - 2k_b K_b \sin \theta_{\text{in}} \cos^2 \phi \chi_{xxz} \\
&\quad - 2k_b K_b \sin \theta_{\text{in}} \sin^2 \phi \chi_{xxz} \\
&\quad - k_b^2 K_b \cos^3 \phi \chi_{xxx} \\
&\quad + k_b^2 K_b \sin^2 \phi \cos \phi \chi_{xxx} \\
&\quad \left. + 2k_b^2 K_b \sin^2 \phi \cos \phi \chi_{xxx} \right),
\end{aligned}$$

and reducing,

$$\begin{aligned}
&= \frac{K_b}{K_v} \Gamma_{pP}^b \left(\sin^3 \theta_{\text{in}} \chi_{zzz} \right. \\
&\quad + k_b^2 \sin \theta_{\text{in}} (\sin^2 \phi + \cos^2 \phi) \chi_{zxx} \\
&\quad - 2k_b K_b \sin \theta_{\text{in}} (\sin^2 \phi + \cos^2 \phi) \chi_{xxz} \\
&\quad \left. + k_b^2 K_b (3 \sin^2 \phi \cos \phi - \cos^3 \phi) \chi_{xxx} \right) \\
&= \frac{K_b}{K_v} \Gamma_{pP}^b \left(\sin^3 \theta_{\text{in}} \chi_{zzz} + k_b^2 \sin \theta_{\text{in}} \chi_{zxx} - 2k_b K_b \sin \theta_{\text{in}} \chi_{xxz} - k_b^2 K_b \chi_{xxx} \cos 3\phi \right),
\end{aligned}$$

where,

$$\Gamma_{pP}^b = \frac{T_p^{vb} (t_p^{vb})^2}{\epsilon_b(\omega) \sqrt{\epsilon_b(2\omega)}}.$$

We find the equivalent expression for \mathcal{R} evaluated inside the bulk as

$$R(2\omega) = \frac{32\pi^3 \omega^2}{c^3 K_b^2} |\mathbf{e}_b^{2\omega} \cdot \chi : \mathbf{e}_b^\omega \mathbf{e}_b^\omega|^2,$$

and we can remove the K_b/K_v factor completely and reduce to the standard form of

$$R(2\omega) = \frac{32\pi^3 \omega^2}{c^3 \cos^2 \theta_{\text{in}}} |\mathbf{e}_b^{2\omega} \cdot \chi : \mathbf{e}_b^\omega \mathbf{e}_b^\omega|^2.$$

b. Taking $\mathcal{P}(2\omega)$ and the fundamental fields in the vacuum

To consider the 1ω fields in the vacuum, we start with Eq. (26) but substitute $\ell \rightarrow v$, thus

$$\mathbf{E}_v(\omega) = E_0 \left[\hat{\mathbf{s}} t_s^{vv} (1 + r_s^{vb}) \hat{\mathbf{s}} + \hat{\mathbf{p}}_{v-} t_p^{vv} \hat{\mathbf{p}}_{v-} + \hat{\mathbf{p}}_{v+} t_p^{vv} r_p^{vb} \hat{\mathbf{p}}_{v-} \right] \cdot \hat{\mathbf{e}}^{\text{in}},$$

t_p^{vv} and t_s^{vv} are one, so we are left with

$$\begin{aligned}
\mathbf{e}_v^\omega &= \left[\hat{\mathbf{s}}(1 + r_s^{vb})\hat{\mathbf{s}} + \hat{\mathbf{p}}_{v-}\hat{\mathbf{p}}_{v-} + \hat{\mathbf{p}}_{v+}r_p^{vb}\hat{\mathbf{p}}_{v-} \right] \cdot \hat{\mathbf{e}}^{\text{in}} \\
&= \left[\hat{\mathbf{s}}(t_s^{vb})\hat{\mathbf{s}} + (\hat{\mathbf{p}}_{v-} + \hat{\mathbf{p}}_{v+}r_p^{vb})\hat{\mathbf{p}}_{v-} \right] \cdot \hat{\mathbf{e}}^{\text{in}} \\
&= \left[\hat{\mathbf{s}}(t_s^{vb})\hat{\mathbf{s}} + \frac{1}{\sqrt{\epsilon_v(\omega)}}(k_v(1 - r_p^{vb})\hat{\mathbf{k}} + \sin \theta_{\text{in}}(1 + r_p^{vb})\hat{\mathbf{z}})\hat{\mathbf{p}}_{v-} \right] \\
&= \left[\hat{\mathbf{s}}(t_s^{vb})\hat{\mathbf{s}} + \left(\frac{k_b}{\sqrt{\epsilon_b(\omega)}}t_p^{vb}\hat{\mathbf{k}} + \sqrt{\epsilon_b(\omega)}\sin \theta_{\text{in}}t_p^{vb}\hat{\mathbf{z}} \right) \hat{\mathbf{p}}_{v-} \right] \cdot \hat{\mathbf{e}}^{\text{in}} \\
&= \left[\hat{\mathbf{s}}(t_s^{vb})\hat{\mathbf{s}} + \frac{t_p^{vb}}{\sqrt{\epsilon_b(\omega)}}(k_b \cos \phi \hat{\mathbf{x}} + k_b \sin \phi \hat{\mathbf{y}} + \epsilon_b(\omega) \sin \theta_{\text{in}}\hat{\mathbf{z}}) \hat{\mathbf{p}}_{v-} \right] \cdot \hat{\mathbf{e}}^{\text{in}}.
\end{aligned}$$

For \mathcal{R}_{pP} we require that $\hat{\mathbf{e}}^{\text{in}} = \hat{\mathbf{p}}_{v-}$, so

$$\mathbf{e}_v^\omega = \frac{t_p^{vb}}{\sqrt{\epsilon_b(\omega)}}(k_b \cos \phi \hat{\mathbf{x}} + k_b \sin \phi \hat{\mathbf{y}} + \epsilon_b(\omega) \sin \theta_{\text{in}}\hat{\mathbf{z}}),$$

and

$$\begin{aligned}
\mathbf{e}_v^\omega \mathbf{e}_v^\omega &= \left(\frac{t_p^{vb}}{\sqrt{\epsilon_b(\omega)}} \right)^2 \left[k_b^2 \cos^2 \phi \hat{\mathbf{x}}\hat{\mathbf{x}} \right. \\
&\quad + k_b^2 \sin^2 \phi \hat{\mathbf{y}}\hat{\mathbf{y}} \\
&\quad + \epsilon_b^2(\omega) \sin^2 \theta_{\text{in}} \hat{\mathbf{z}}\hat{\mathbf{z}} \\
&\quad + 2k_b^2 \sin \phi \cos \phi \hat{\mathbf{x}}\hat{\mathbf{y}} \\
&\quad + 2\epsilon_b(\omega)k_b \sin \theta_{\text{in}} \sin \phi \hat{\mathbf{y}}\hat{\mathbf{z}} \\
&\quad \left. + 2\epsilon_b(\omega)k_b \sin \theta_{\text{in}} \cos \phi \hat{\mathbf{x}}\hat{\mathbf{z}} \right].
\end{aligned}$$

We also require the 2ω fields evaluated in the vacuum, which is Eq. (25),

$$\mathbf{e}_v^{2\omega} = \hat{\mathbf{e}}^{\text{out}} \cdot \left[\hat{\mathbf{s}}T_s^{vb}\hat{\mathbf{s}} + \hat{\mathbf{P}}_{v+} \frac{T_p^{vb}}{\sqrt{\epsilon_b(2\omega)}}(\epsilon_b(2\omega) \sin \theta_{\text{in}}\hat{\mathbf{z}} - K_b\hat{\mathbf{k}}) \right], \quad (\text{A1})$$

and with $\hat{\mathbf{e}}^{\text{out}} = \hat{\mathbf{P}}_{v+}$ we have

$$\mathbf{e}_v^{2\omega} = \frac{T_p^{vb}}{\sqrt{\epsilon_b(2\omega)}}(\epsilon_b(2\omega) \sin \theta_{\text{in}}\hat{\mathbf{z}} - K_b \cos \phi \hat{\mathbf{x}} - K_b \sin \phi \hat{\mathbf{y}}). \quad (\text{A2})$$

So lastly, we have that

$$\begin{aligned}
\mathbf{e}_v^{2\omega} \cdot \boldsymbol{\chi} : \mathbf{e}_v^\omega \mathbf{e}_v^\omega = & \\
& \frac{T_p^{vb}}{\sqrt{\epsilon_b(2\omega)}} \left(\frac{t_p^{vb}}{\sqrt{\epsilon_b(\omega)}} \right)^2 \left[\epsilon_b(2\omega) k_b^2 \sin \theta_{\text{in}} \cos^2 \phi \chi_{zxx} \right. \\
& + \epsilon_b(2\omega) k_b^2 \sin \theta_{\text{in}} \sin^2 \phi \chi_{zyy} \\
& + \epsilon_b^2(\omega) \epsilon_b(2\omega) \sin^3 \theta_{\text{in}} \chi_{zzz} \\
& + 2\epsilon_b(2\omega) k_b^2 \sin \theta_{\text{in}} \sin \phi \cos \phi \chi_{zxy} \\
& + 2\epsilon_b(\omega) \epsilon_b(2\omega) k_b \sin^2 \theta_{\text{in}} \sin \phi \chi_{zyz} \\
& + 2\epsilon_b(\omega) \epsilon_b(2\omega) k_b \sin^2 \theta_{\text{in}} \cos \phi \chi_{zxz} \\
& - k_b^2 K_b \cos^3 \phi \chi_{xxx} \\
& - k_b^2 K_b \sin^2 \phi \cos \phi \chi_{xyy} \\
& - \epsilon_b^2(\omega) K_b \sin^2 \theta_{\text{in}} \cos \phi \chi_{xzz} \\
& - 2k_b^2 K_b \sin \phi \cos^2 \phi \chi_{xxy} \\
& - 2\epsilon_b(\omega) k_b K_b \sin \theta_{\text{in}} \sin \phi \cos \phi \chi_{xyz} \\
& - 2\epsilon_b(\omega) k_b K_b \sin \theta_{\text{in}} \cos^2 \phi \chi_{xxz} \\
& - k_b^2 K_b \sin \phi \cos^2 \phi \chi_{yxx} \\
& - k_b^2 K_b \sin^3 \phi \chi_{yyy} \\
& - \epsilon_b^2(\omega) K_b \sin^2 \theta_{\text{in}} \sin \phi \chi_{yzz} \\
& - 2k_b^2 K_b \sin^2 \phi \cos \phi \chi_{yyx} \\
& - 2\epsilon_b(\omega) k_b K_b \sin \theta_{\text{in}} \sin^2 \phi \chi_{yyz} \\
& \left. - 2\epsilon_b(\omega) k_b K_b \sin \theta_{\text{in}} \sin \phi \cos \phi \chi_{yxz} \right],
\end{aligned}$$

and after eliminating components,

$$\begin{aligned}
&= \Gamma_{pP}^v [\epsilon_b^2(\omega) \epsilon_b(2\omega) \sin^3 \theta_{\text{in}} \chi_{zzz} \\
&\quad + \epsilon_b(2\omega) k_b^2 \sin \theta_{\text{in}} \cos^2 \phi \chi_{zxx} \\
&\quad + \epsilon_b(2\omega) k_b^2 \sin \theta_{\text{in}} \sin^2 \phi \chi_{zxx} \\
&\quad - 2\epsilon_b(\omega) k_b K_b \sin \theta_{\text{in}} \cos^2 \phi \chi_{xxz} \\
&\quad - 2\epsilon_b(\omega) k_b K_b \sin \theta_{\text{in}} \sin^2 \phi \chi_{xxz} \\
&\quad + 3k_b^2 K_b \sin^2 \phi \cos \phi \chi_{xxx} \\
&\quad - k_b^2 K_b \cos^3 \phi \chi_{xxx}] \\
&= \Gamma_{pP}^v [\epsilon_b^2(\omega) \epsilon_b(2\omega) \sin^3 \theta_{\text{in}} \chi_{zzz} + \epsilon_b(2\omega) k_b^2 \sin \theta_{\text{in}} \chi_{zxx} \\
&\quad - 2\epsilon_b(\omega) k_b K_b \sin \theta_{\text{in}} \chi_{xxz} - k_b^2 K_b \chi_{xxx} \cos 3\phi],
\end{aligned}$$

where

$$\Gamma_{pP}^v = \frac{T_p^{vb} (t_p^{vb})^2}{\epsilon_b(\omega) \sqrt{\epsilon_b(2\omega)}}.$$

c. Taking $\mathcal{P}(2\omega)$ in ℓ and the fundamental fields in the bulk

For this scenario with $\hat{\mathbf{e}}^{\text{in}} = \hat{\mathbf{p}}_{v-}$ and $\hat{\mathbf{e}}^{\text{out}} = \hat{\mathbf{P}}_{v+}$, we obtain from Eq. (24),

$$\mathbf{e}_\ell^{2\omega} = \frac{T_p^{v\ell} T_p^{\ell b}}{\epsilon_\ell(2\omega) \sqrt{\epsilon_b(2\omega)}} (\epsilon_b(2\omega) \sin \theta_{\text{in}} \hat{\mathbf{z}} - \epsilon_\ell(2\omega) K_b \cos \phi \hat{\mathbf{x}} - \epsilon_\ell(2\omega) K_b \sin \phi \hat{\mathbf{y}}),$$

and Eq. (28),

$$\begin{aligned}
\mathbf{e}_b^\omega \mathbf{e}_b^\omega &= \frac{(t_p^{vb})^2}{\epsilon_b(\omega)} (\sin^2 \theta_{\text{in}} \hat{\mathbf{z}} \hat{\mathbf{z}} + k_b^2 \cos^2 \phi \hat{\mathbf{x}} \hat{\mathbf{x}} + k_b^2 \sin^2 \phi \hat{\mathbf{y}} \hat{\mathbf{y}} \\
&\quad + 2k_b \sin \theta_{\text{in}} \cos \phi \hat{\mathbf{z}} \hat{\mathbf{x}} + 2k_b \sin \theta_{\text{in}} \sin \phi \hat{\mathbf{z}} \hat{\mathbf{y}} + 2k_b^2 \sin \phi \cos \phi \hat{\mathbf{x}} \hat{\mathbf{y}}).
\end{aligned}$$

Thus,

$$\begin{aligned}
\mathbf{e}_\ell^{2\omega} \cdot \boldsymbol{\chi} : \mathbf{e}_b^\omega \mathbf{e}_b^\omega &= \frac{T_p^{v\ell} T_p^{\ell b} (t_p^{vb})^2}{\epsilon_\ell(2\omega) \epsilon_b(\omega) \sqrt{\epsilon_b(2\omega)}} \left[\begin{aligned}
&+ \epsilon_b(2\omega) \sin^3 \theta_{\text{in}} \chi_{zzz} \\
&+ \epsilon_b(2\omega) k_b^2 \sin \theta_{\text{in}} \cos^2 \phi \chi_{zxx} \\
&+ \epsilon_b(2\omega) k_b^2 \sin \theta_{\text{in}} \sin^2 \phi \chi_{zyy} \\
&+ 2\epsilon_b(2\omega) k_b \sin^2 \theta_{\text{in}} \cos \phi \chi_{zzx} \\
&+ 2\epsilon_b(2\omega) k_b \sin^2 \theta_{\text{in}} \sin \phi \chi_{zzy} \\
&+ 2\epsilon_b(2\omega) k_b^2 \sin \theta_{\text{in}} \sin \phi \cos \phi \chi_{xzy} \\
&- \epsilon_\ell(2\omega) \sin^2 \theta_{\text{in}} K_b \cos \phi \chi_{xzz} \\
&- \epsilon_\ell(2\omega) k_b^2 K_b \cos^3 \phi \chi_{xxx} \\
&- \epsilon_\ell(2\omega) k_b^2 K_b \sin^2 \phi \cos \phi \chi_{xyy} \\
&- 2\epsilon_\ell(2\omega) k_b K_b \sin \theta_{\text{in}} \cos^2 \phi \chi_{xzx} \\
&- 2\epsilon_\ell(2\omega) k_b K_b \sin \theta_{\text{in}} \sin \phi \cos \phi \chi_{xzy} \\
&- 2\epsilon_\ell(2\omega) k_b^2 K_b \sin \phi \cos^2 \phi \chi_{xxy} \\
&- \epsilon_\ell(2\omega) K_b \sin^2 \theta_{\text{in}} \sin \phi \chi_{yzz} \\
&- \epsilon_\ell(2\omega) k_b^2 K_b \cos^2 \phi \sin \phi \chi_{yxx} \\
&- \epsilon_\ell(2\omega) k_b^2 K_b \sin^3 \phi \chi_{yyy} \\
&- 2\epsilon_\ell(2\omega) k_b K_b \sin \theta_{\text{in}} \cos \phi \sin \phi \chi_{yzx} \\
&- 2\epsilon_\ell(2\omega) k_b K_b \sin \theta_{\text{in}} \sin^2 \phi \chi_{yzy} \\
&- 2\epsilon_\ell(2\omega) k_b^2 K_b \sin^2 \phi \cos \phi \chi_{yxy} \end{aligned} \right].
\end{aligned}$$

We eliminate and replace components,

$$\begin{aligned} \mathbf{e}_\ell^{2\omega} \cdot \boldsymbol{\chi} : \mathbf{e}_b^\omega \mathbf{e}_b^\omega = \Gamma_{pP}^{\ell b} \bigg[& + \epsilon_b(2\omega) \sin^3 \theta_{\text{in}} \chi_{zzz} \\ & + \epsilon_b(2\omega) k_b^2 \sin \theta_{\text{in}} \cos^2 \phi \chi_{zxx} \\ & + \epsilon_b(2\omega) k_b^2 \sin \theta_{\text{in}} \sin^2 \phi \chi_{zxx} \\ & - 2\epsilon_\ell(2\omega) k_b K_b \sin \theta_{\text{in}} \cos^2 \phi \chi_{xxz} \\ & - 2\epsilon_\ell(2\omega) k_b K_b \sin \theta_{\text{in}} \sin^2 \phi \chi_{xxz} \\ & - \epsilon_\ell(2\omega) k_b^2 K_b \cos^3 \phi \chi_{xxx} \\ & + \epsilon_\ell(2\omega) k_b^2 K_b \sin^2 \phi \cos \phi \chi_{xxx} \\ & + 2\epsilon_\ell(2\omega) k_b^2 K_b \sin^2 \phi \cos \phi \chi_{xxx} \bigg], \end{aligned}$$

so lastly

$$\begin{aligned} \mathbf{e}_\ell^{2\omega} \cdot \boldsymbol{\chi} : \mathbf{e}_b^\omega \mathbf{e}_b^\omega = \Gamma_{pP}^{\ell b} \bigg[& \epsilon_b(2\omega) \sin^3 \theta_{\text{in}} \chi_{zzz} + \epsilon_b(2\omega) k_b^2 \sin \theta_{\text{in}} \chi_{zxx} \\ & - 2\epsilon_\ell(2\omega) k_b K_b \sin \theta_{\text{in}} \chi_{xxz} - \epsilon_\ell(2\omega) k_b^2 K_b \chi_{xxx} \cos 3\phi \bigg], \end{aligned}$$

where

$$\Gamma_{pP}^{\ell b} = \frac{T_p^{v\ell} T_p^{\ell b} (t_p^{vb})^2}{\epsilon_\ell(2\omega) \epsilon_b(\omega) \sqrt{\epsilon_b(2\omega)}}.$$

2. \mathcal{R}_{pS}

To obtain $R_{pS}(2\omega)$ we use $\hat{\mathbf{e}}^{\text{in}} = \hat{\mathbf{p}}_{v-}$ in Eq. (27), and $\hat{\mathbf{e}}^{\text{out}} = \hat{\mathbf{S}}$ in Eq. (24). We also use the unit vectors defined in Eqs. (34) and (35). Substituting, we get

$$\mathbf{e}_\ell^{2\omega} = T_s^{v\ell} T_s^{\ell b} [-\sin \phi \hat{\mathbf{x}} + \cos \phi \hat{\mathbf{y}}],$$

for 2ω , and for the fundamental fields,

$$\begin{aligned} \mathbf{e}_\ell^\omega \mathbf{e}_\ell^\omega &= \left(\frac{t_p^{v\ell} t_p^{\ell b}}{\epsilon_\ell(\omega) \sqrt{\epsilon_b(\omega)}} \right)^2 (\epsilon_b(\omega) \sin \theta_{\text{in}} \hat{\mathbf{z}} + \epsilon_\ell(\omega) k_b \cos \phi \hat{\mathbf{x}} + \epsilon_\ell(\omega) k_b \sin \phi \hat{\mathbf{y}})^2. \\ &= \left(\frac{t_p^{v\ell} t_p^{\ell b}}{\epsilon_\ell(\omega) \sqrt{\epsilon_b(\omega)}} \right)^2 (\epsilon_b^2(\omega) \sin^2 \theta_{\text{in}} \hat{\mathbf{z}} \hat{\mathbf{z}} + 2\epsilon_b(\omega) \epsilon_\ell(\omega) k_b \sin \theta_{\text{in}} \cos \phi \hat{\mathbf{z}} \hat{\mathbf{x}} \\ &\quad + \epsilon_\ell^2(\omega) k_b^2 \cos^2 \phi \hat{\mathbf{x}} \hat{\mathbf{x}} + 2\epsilon_\ell^2(\omega) k_b^2 \cos \phi \sin \phi \hat{\mathbf{x}} \hat{\mathbf{y}} \\ &\quad + \epsilon_\ell^2(\omega) k_b^2 \sin^2 \phi \hat{\mathbf{y}} \hat{\mathbf{y}} + 2\epsilon_b(\omega) \epsilon_\ell(\omega) k_b \sin \theta_{\text{in}} \sin \phi \hat{\mathbf{y}} \hat{\mathbf{z}}). \end{aligned}$$

Therefore,

$$\mathbf{e}_\ell^{2\omega} \cdot \boldsymbol{\chi} : \mathbf{e}_\ell^\omega \mathbf{e}_\ell^\omega =$$

$$\begin{aligned} T_s^{v\ell} T_s^{\ell b} \left(\frac{t_p^{v\ell} t_p^{\ell b}}{\epsilon_\ell(\omega) \sqrt{\epsilon_b(\omega)}} \right)^2 & \left[-\epsilon_b^2(\omega) \sin^2 \theta_{\text{in}} \sin \phi \chi_{xzz} \right. \\ & -2\epsilon_b(\omega) \epsilon_\ell(\omega) k_b \sin \theta_{\text{in}} \cos \phi \sin \phi \chi_{xxz} \\ & -\epsilon_\ell^2(\omega) k_b^2 \cos^2 \phi \sin \phi \chi_{xxx} \\ & -2\epsilon_\ell^2(\omega) k_b^2 \cos \phi \sin^2 \phi \chi_{xxy} \\ & -\epsilon_\ell^2(\omega) k_b^2 \sin^3 \phi \chi_{xyy} \\ & -2\epsilon_b(\omega) \epsilon_\ell(\omega) k_b \sin \theta_{\text{in}} \sin^2 \phi \chi_{xyz} \\ & +\epsilon_b^2(\omega) \sin^2 \theta_{\text{in}} \cos \phi \chi_{yzz} \\ & +2\epsilon_b(\omega) \epsilon_\ell(\omega) k_b \sin \theta_{\text{in}} \cos^2 \phi \chi_{yxz} \\ & +\epsilon_\ell^2(\omega) k_b^2 \cos^3 \phi \chi_{yxx} \\ & +2\epsilon_\ell^2(\omega) k_b^2 \cos^2 \phi \sin \phi \chi_{yxy} \\ & +\epsilon_\ell^2(\omega) k_b^2 \cos \phi \sin^2 \phi \chi_{yyx} \\ & \left. +2\epsilon_b(\omega) \epsilon_\ell(\omega) k_b \sin \theta_{\text{in}} \cos \phi \sin \phi \chi_{yyz} \right], \end{aligned}$$

and taking into account that $\chi_{xzz} = \chi_{xxy} = \chi_{xyz} = \chi_{yzz} = \chi_{yxz} = \chi_{yxx} = \chi_{yyy} = 0$, we have

$$\begin{aligned} & = \Gamma_{pS}^\ell \left[+\epsilon_\ell^2(\omega) k_b^2 \sin^3 \phi \chi_{xxx} \right. \\ & \quad -2\epsilon_\ell^2(\omega) k_b^2 \cos^2 \phi \sin \phi \chi_{xxx} \\ & \quad -\epsilon_\ell^2(\omega) k_b^2 \cos^2 \phi \sin \phi \chi_{xxx} \\ & \quad +2\epsilon_b(\omega) \epsilon_\ell(\omega) k_b \sin \theta_{\text{in}} \cos \phi \sin \phi \chi_{xxz} \\ & \quad \left. -2\epsilon_b(\omega) \epsilon_\ell(\omega) k_b \sin \theta_{\text{in}} \cos \phi \sin \phi \chi_{xxz} \right] \\ & = \Gamma_{pS}^\ell \left[\epsilon_\ell^2(\omega) k_b^2 (\sin^3 \phi - 3 \cos^2 \phi \sin \phi) \chi_{xxx} \right] \\ & = \Gamma_{pS}^\ell \left[-\epsilon_\ell^2(\omega) k_b^2 \sin 3\phi \chi_{xxx} \right]. \end{aligned}$$

We summarize as follows,

$$\mathbf{e}_\ell^{2\omega} \cdot \boldsymbol{\chi} : \mathbf{e}_\ell^\omega \mathbf{e}_\ell^\omega \equiv \Gamma_{pS}^\ell r_{pS}^\ell,$$

where

$$r_{pS}^\ell = -\epsilon_\ell^2(\omega) k_b^2 \sin 3\phi \chi_{xxx},$$

and

$$\Gamma_{pS}^\ell = T_s^{v\ell} T_s^{\ell b} \left(\frac{t_p^{v\ell} t_p^{\ell b}}{\epsilon_\ell(\omega) \sqrt{\epsilon_b(\omega)}} \right)^2$$

In order to reduce above result to that of Ref. [\[17\]](#) and [\[18\]](#), we take the 2- ω radiations factors for vacuum by taking $\ell = v$, thus $\epsilon_\ell(2\omega) = 1$, $T_s^{v\ell} = 1$, $T_s^{\ell b} = T_s^{vb}$, and the fundamental field inside medium b by taking $\ell = b$, thus $\epsilon_\ell(\omega) = \epsilon_b(\omega)$, $t_p^{v\ell} = t_p^{vb}$, and $t_p^{\ell b} = 1$. With these choices,

$$r_{pS}^b = -k_b^2 \sin 3\phi \chi_{xxx},$$

and

$$\Gamma_{pS}^b = T_s^{vb} \left(\frac{t_p^{vb}}{\sqrt{\epsilon_b(\omega)}} \right)^2.$$

3. \mathcal{R}_{sP}

To obtain $R_{sP}(2\omega)$ we use $\hat{\mathbf{e}}^{\text{in}} = \hat{\mathbf{s}}$ in Eq. [\(27\)](#), and $\hat{\mathbf{e}}^{\text{out}} = \hat{\mathbf{P}}_{v+}$ in Eq. [\(24\)](#). We also use the unit vectors defined in Eqs. [\(34\)](#) and [\(35\)](#). Substituting, we get

$$\mathbf{e}_\ell^{2\omega} = \frac{T_p^{v\ell} T_p^{\ell b}}{\epsilon_\ell(2\omega) \sqrt{\epsilon_b(2\omega)}} [\epsilon_b(2\omega) \sin \theta_{\text{in}} \hat{\mathbf{z}} - \epsilon_\ell(2\omega) K_b \cos \phi \hat{\mathbf{x}} - \epsilon_\ell(2\omega) K_b \sin \phi \hat{\mathbf{y}}],$$

for 2ω , and for the fundamental fields,

$$\mathbf{e}_\ell^\omega \mathbf{e}_\ell^\omega = \left(t_s^{v\ell} t_s^{\ell b} \right)^2 (\sin^2 \phi \hat{\mathbf{x}} \hat{\mathbf{x}} + \cos^2 \phi \hat{\mathbf{y}} \hat{\mathbf{y}} - 2 \sin \phi \cos \phi \hat{\mathbf{x}} \hat{\mathbf{y}}).$$

Therefore,

$$\begin{aligned} \mathbf{e}_\ell^{2\omega} \cdot \chi : \mathbf{e}_\ell^\omega \mathbf{e}_\ell^\omega = & \frac{T_p^{v\ell} T_p^{\ell b} (t_s^{v\ell} t_s^{\ell b})^2}{\epsilon_\ell(2\omega) \sqrt{\epsilon_b(2\omega)}} [\epsilon_b(2\omega) \sin \theta_{\text{in}} \sin^2 \phi \chi_{zxx} + \epsilon_b(2\omega) \sin \theta_{\text{in}} \cos^2 \phi \chi_{zyy} \\ & - 2\epsilon_b(2\omega) \sin \theta_{\text{in}} \sin \phi \cos \phi \chi_{zxy} - \epsilon_\ell(2\omega) K_b \cos \phi \sin^2 \phi \chi_{xxx} \\ & - \epsilon_\ell(2\omega) K_b \cos \phi \cos^2 \phi \chi_{xyy} + 2\epsilon_\ell(2\omega) K_b \cos \phi \sin \phi \cos \phi \chi_{xxy} \\ & - \epsilon_\ell(2\omega) K_b \sin \phi \sin^2 \phi \chi_{yxx} - \epsilon_\ell(2\omega) K_b \sin \phi \cos^2 \phi \chi_{yyy} \\ & + 2\epsilon_\ell(2\omega) K_b \sin \phi \sin \phi \cos \phi \chi_{yxy}], \end{aligned}$$

and taking into account that $\chi_{zxy} = \chi_{xxy} = \chi_{yxx} = \chi_{yyy} = 0$, we have

$$\begin{aligned}
&= \Gamma_{sP}^\ell \left[\epsilon_b(2\omega) \sin \theta_{\text{in}} \sin^2 \phi \chi_{zxx} + \epsilon_b(2\omega) \sin \theta_{\text{in}} \cos^2 \phi \chi_{zxx} \right. \\
&\quad \left. - \epsilon_\ell(2\omega) K_b \cos \phi \sin^2 \phi \chi_{xxx} + \epsilon_\ell(2\omega) K_b \cos^3 \phi \chi_{xxx} \right. \\
&\quad \left. - 2\epsilon_\ell(2\omega) K_b \sin^2 \phi \cos \phi \chi_{xxx} \right] \\
&= \Gamma_{sP}^\ell \left[\epsilon_b(2\omega) \sin \theta_{\text{in}} (\sin^2 \phi + \cos^2 \phi) \chi_{zxx} \right. \\
&\quad \left. - \epsilon_\ell(2\omega) K_b (\cos \phi \sin^2 \phi - \cos^3 \phi + 2 \sin^2 \phi \cos \phi) \chi_{xxx} \right] \\
&= \Gamma_{sP}^\ell \left[\epsilon_b(2\omega) \sin \theta_{\text{in}} \chi_{zxx} + \epsilon_\ell(2\omega) K_b (\cos^3 \phi - 3 \sin^2 \phi \cos \phi) \chi_{xxx} \right] \\
&= \Gamma_{sP}^\ell \left[\epsilon_b(2\omega) \sin \theta_{\text{in}} \chi_{zxx} + \epsilon_\ell(2\omega) K_b \cos 3\phi \chi_{xxx} \right].
\end{aligned}$$

We summarize as follows,

$$\mathbf{e}_\ell^{2\omega} \cdot \boldsymbol{\chi} : \mathbf{e}_\ell^\omega \mathbf{e}_\ell^\omega \equiv \Gamma_{sP}^\ell r_{sP}^\ell,$$

where

$$r_{sP}^\ell = \epsilon_b(2\omega) \sin \theta_{\text{in}} \chi_{zxx} + \epsilon_\ell(2\omega) K_b \chi_{xxx} \cos 3\phi,$$

and

$$\Gamma_{sP}^\ell = \frac{T_p^{\ell v} T_p^{\ell b} (t_s^{v\ell} t_s^{\ell b})^2}{\epsilon_\ell(2\omega) \sqrt{\epsilon_b(2\omega)}}.$$

In order to reduce above result to that of Ref. [\[17\]](#) and [\[18\]](#), we take the 2- ω radiations factors for vacuum by taking $\ell = v$, thus $\epsilon_\ell(2\omega) = 1$, $T_p^{v\ell} = 1$, $T_p^{\ell b} = T_p^{vb}$, and the fundamental field inside medium b by taking $\ell = b$, thus $\epsilon_\ell(\omega) = \epsilon_b(\omega)$, $t_s^{v\ell} = t_s^{vb}$, and $t_s^{\ell b} = 1$. With these choices,

$$r_{sP}^b = \epsilon_b(2\omega) \sin \theta_{\text{in}} \chi_{zxx} + K_b \chi_{xxx} \cos 3\phi,$$

and

$$\Gamma_{sP}^b = \frac{T_p^{vb} (t_s^{vb})^2}{\sqrt{\epsilon_b(2\omega)}}.$$

4. \mathcal{R}_{sS}

For \mathcal{R}_{sS} we have that $\hat{\mathbf{e}}^{\text{in}} = \hat{\mathbf{s}}$ and $\hat{\mathbf{e}}^{\text{out}} = \hat{\mathbf{S}}$. This leads to

$$\begin{aligned}\mathbf{e}_\ell^{2\omega} &= T_s^{v\ell} T_s^{\ell b} [-\sin \phi \hat{\mathbf{x}} + \cos \phi \hat{\mathbf{y}}], \\ \mathbf{e}_\ell^\omega \mathbf{e}_\ell^\omega &= \left(t_s^{v\ell} t_s^{\ell b}\right)^2 (\sin^2 \phi \hat{\mathbf{x}} \hat{\mathbf{x}} + \cos^2 \phi \hat{\mathbf{y}} \hat{\mathbf{y}} - 2 \sin \phi \cos \phi \hat{\mathbf{x}} \hat{\mathbf{y}}).\end{aligned}$$

Therefore,

$$\begin{aligned}\mathbf{e}_\ell^{2\omega} \cdot \boldsymbol{\chi} : \mathbf{e}_\ell^\omega \mathbf{e}_\ell^\omega &= T_s^{v\ell} T_s^{\ell b} \left(t_s^{v\ell} t_s^{\ell b}\right)^2 [-\sin^3 \phi \chi_{xxx} - \sin \phi \cos^2 \phi \chi_{xyy} + 2 \sin^2 \phi \cos \phi \chi_{xxy} \\ &\quad + \sin^2 \phi \cos \phi \chi_{yxx} + \cos^3 \phi \chi_{yyy} - 2 \sin \phi \cos^2 \phi \chi_{yyx}] \\ &= T_s^{v\ell} T_s^{\ell b} \left(t_s^{v\ell} t_s^{\ell b}\right)^2 [-\sin^3 \phi \chi_{xxx} + 3 \sin \phi \cos^2 \phi \chi_{xxx}] \\ &= T_s^{v\ell} T_s^{\ell b} \left(t_s^{v\ell} t_s^{\ell b}\right)^2 \chi_{xxx} \sin 3\phi\end{aligned}$$

Summarizing,

$$\mathbf{e}_\ell^{2\omega} \cdot \boldsymbol{\chi} : \mathbf{e}_\ell^\omega \mathbf{e}_\ell^\omega \equiv \Gamma_{sS}^\ell r_{sS}^\ell,$$

where

$$r_{sS}^\ell = \chi_{xxx} \sin 3\phi,$$

and

$$\Gamma_{sS}^\ell = T_s^{v\ell} T_s^{\ell b} \left(t_s^{v\ell} t_s^{\ell b}\right)^2.$$

In order to reduce above result to that of Ref. [mizrahiJOS468PRB87](#) and [\[7\]](#), we take the 2ω radiations factors for vacuum by taking $\ell = v$, thus $\epsilon_\ell(2\omega) = 1$, $T_s^{v\ell} = 1$, $T_s^{\ell b} = T_s^{vb}$, and the fundamental field inside medium b by taking $\ell = b$, thus $\epsilon_\ell(\omega) = \epsilon_b(\omega)$, $t_s^{v\ell} = t_s^{vb}$, and $t_s^{\ell b} = 1$. With these choices,

$$r_{sS}^b = \chi_{xxx} \sin 3\phi,$$

and

$$\Gamma_{sS}^b = T_s^{vb} \left(t_s^{vb}\right)^2.$$