## Exercise 6.1 ()

# Exercise 6.2 (LIF in weak gravitational field)

#### Exercise 6.2: Local inertial frame in a weak gravitational field

A four-dimensional manifold has coordinates (t,x,y,z) and line element

$$ds^{2} = -(1+2\phi)dt^{2} + (1-2\phi)(dx^{2} + dy^{2} + dz^{2}),$$
(6)

where  $|\phi(t, x, y, z)| \ll 1$  everywhere.

- (a) At any point P with coordinates  $(t_0, x_0, y_0, z_0)$  find a coordinate transformation to a locally inertial coordinate system, to first order in  $\phi$ . (1 pt)
- We have to perform a change of coordinates such that the metric in the new frame,  $\bar{g}_{\mu\nu}$ , is locally flat around P
- This means that  $\bar{g}_{\mu\nu} = \eta_{\mu\nu}$
- The first derivatives of  $\bar{g}_{\mu\nu}$  at P must vanish,

$$\bar{g}_{\mu\nu}(x) = \eta_{\mu\nu} + \frac{\partial^2 g_{\mu\nu}}{\partial x^\alpha \partial x^\beta} \Big|_P (x^\alpha - x_P^\alpha)(x^\beta - x_P^\beta) + \mathcal{O}(\Delta x^3)$$
(7)

with 
$$\bar{g}_{\bar{\mu}\bar{\nu}} = \frac{\partial x^{\mu}}{\partial x^{\bar{\mu}}} \frac{\partial x^{\nu}}{\partial x^{\bar{\nu}}} g_{\mu\nu}$$
. We can define  $\Lambda^{\mu}_{\bar{\nu}} \equiv \frac{\partial x^{\mu}}{\partial x^{\bar{\nu}}}$ .

In order  $\bar{g}_{\mu\nu}$  to take the form (7) the transformation matrix must fulfill the following equations:

$$-1 = -[1 + 2\phi(P)][\Lambda^{0}_{\bar{0}}(P)]^{2} \longrightarrow \Lambda^{0}_{\bar{0}}(P) = 1 - \phi(P) + \mathcal{O}(\phi^{2})$$

$$1 = [1 - 2\phi(P)][\Lambda^{i}_{\bar{i}}(P)]^{2} \longrightarrow \Lambda^{i}_{\bar{i}}(P) = 1 + \phi(P) + \mathcal{O}(\phi^{2})$$
Taylor expansion

and

$$0 = 2 \frac{\partial \Lambda^{0}_{\bar{0}}}{\partial x^{\mu}} \Big|_{P} g_{00}(P) + [\Lambda^{0}_{\bar{0}}(P)]^{2} \frac{\partial g_{00}}{\partial x^{\mu}} \Big|_{P} \longrightarrow \frac{\partial \Lambda^{0}_{\bar{0}}}{\partial x^{\mu}} \Big|_{P} = -\partial_{\mu} \phi \Big|_{P} + \mathcal{O}(\phi^{2})$$

$$0 = 2 \frac{\partial \Lambda^{i}_{\bar{i}}}{\partial x^{\mu}} \Big|_{P} g_{ii}(P) + [\Lambda^{i}_{\bar{i}}(P)]^{2} \frac{\partial g_{ii}}{\partial x^{\mu}} \Big|_{P} \longrightarrow \frac{\partial \Lambda^{i}_{\bar{i}}}{\partial x^{\mu}} \Big|_{P} = \partial_{\mu} \phi \Big|_{P} + \mathcal{O}(\phi^{2}).$$

The elements of the transformation matrix thus read,

$$\Lambda^{0}_{\bar{0}}(x) = 1 - \phi(P) - \partial_{\mu}\phi \Big|_{P} (x^{\mu} - x^{\mu}_{P}) + \mathcal{O}(\phi^{2})$$

$$\Lambda^{i}_{\bar{i}}(x) = 1 + \phi(P) + \partial_{\mu}\phi \Big|_{P} (x^{\mu} - x^{\mu}_{P}) + \mathcal{O}(\phi^{2}).$$

and with the non-diagonal terms equal to zero.

As  $\eta$  is a diagonal metric we can try with a diagonal transformation matrix  $\Lambda$  leading also to a diagonal metric  $\bar{g}_{\mu\nu}$  (this is the simplest initial guess). Thus, we have,

$$\bar{g}_{00}(x) = [\Lambda^{0}_{\bar{0}}(x)]^{2} g_{00}(x) = -[1 + 2\phi(x)][\Lambda^{0}_{\bar{0}}(x)]^{2}$$

$$\bar{g}_{ii}(x) = [\Lambda^{i}_{\bar{i}}(x)]^{2} g_{ii}(x) = [1 - 2\phi(x)][\Lambda^{i}_{\bar{i}}(x)]^{2}.$$

(b) At what rate does such a frame accelerate with respect to the original coordinates, again to first order in  $\phi$ ? Consider that the free-falling observer moves with non-relativistic velocity in the non-inertial system. (1 pt)

The world-line of the free-falling frame measured by an observer located in the non-inertial frame obeys the following geodesic equation,

$$\frac{d^2x^i}{d\tau^2} = -\Gamma^i_{\ \mu\nu}\frac{dx^\mu}{d\tau}\frac{dx^\nu}{d\tau} = -\Gamma^i_{\ 00}\left(\frac{dt}{d\tau}\right)^2 - \Gamma^i_{\ jl}\frac{dx^j}{d\tau}\frac{dx^l}{d\tau} - 2\Gamma^i_{\ j0}\frac{dx^j}{d\tau}\frac{dt}{d\tau}$$

The Christoffel symbols are of order  $\Gamma \sim \mathcal{O}(\phi)$  and since the free-falling body moves with non-relativistic velocity in the non-inertial frame,  $d\tau = d\tau + \mathcal{O}(v,\phi)$ . Thus, it is clear that the second term in the right-hand side is of order  $\mathcal{O}(3)$  and the last one of order  $\mathcal{O}(2)$ . The leading contribution comes from the first term and is of order  $\mathcal{O}(\phi)$ . Hence,  $\frac{d^2x^i}{dt^2} = -\Gamma^i_{00} + \mathcal{O}(\phi^2) = -\partial_i\phi + \mathcal{O}(\phi^2)$ 

 $\phi$  can be interpreted as the Newtonian gravitational potential.

### (c) Compute the elements of the Riemann tensor in the original (non-inertial) frame to first order in $\phi$ . (1 pt)

The Christoffel symbols are as follows:

$$\Gamma^{0}_{00} = \partial_{0}\phi + \mathcal{O}(\phi^{2}) \qquad \Gamma^{i}_{00} = \partial^{i}\phi + \mathcal{O}(\phi^{2}) 
\Gamma^{0}_{ij} = -\delta_{ij}\partial_{0}\phi + \mathcal{O}(\phi^{2}) \qquad \Gamma^{i}_{0j} = -\delta^{i}_{j}\partial_{0}\phi + \mathcal{O}(\phi^{2}) 
\Gamma^{0}_{0i} = \partial_{i}\phi + \mathcal{O}(\phi^{2}) \qquad \Gamma^{i}_{jl} = \delta_{jl}\partial^{i}\phi - \delta^{i}_{j}\partial_{l}\phi - \delta^{i}_{l}\partial_{j}\phi + \mathcal{O}(\phi^{2})$$

Notice that  $\partial_i \phi = \partial^i \phi + \mathcal{O}(\phi)$ , so partial derivatives with upper and lower indices can be interchanged if we work to first order in  $\phi$ . The components of the Riemann tensor will take the following form if we neglect contributions of order  $\mathcal{O}(\phi^2)$  or higher,

$$R^{\mu}_{\ \nu\alpha\beta} = \partial_{\alpha}\Gamma^{\mu}_{\ \nu\beta} - \partial_{\beta}\Gamma^{\mu}_{\ \nu\alpha} + \mathcal{O}(\phi^2)$$

Thus, we have,

$$R^{\mu}_{\nu 00} = \mathcal{O}(\phi^2)$$

$$R^{\mu}_{\nu 0i} = \partial_0 \Gamma^{\mu}_{\nu i} - \partial_i \Gamma^{\mu}_{\nu 0} + \mathcal{O}(\phi^2)$$

$$R^{\mu}_{\nu ij} = \partial_i \Gamma^{\mu}_{\nu j} - \partial_j \Gamma^{\mu}_{\nu i} + \mathcal{O}(\phi^2)$$

Equivalently,

$$R^{\mu}_{\ \nu 00} = \mathcal{O}(\phi^2)$$

$$R^{0}_{\ 00i} = \mathcal{O}(\phi^2)$$

$$R^{0}_{\ j0i} = -\delta_{ij}\partial_{0}^{2}\phi - \partial_{i}\partial_{j}\phi + \mathcal{O}(\phi^2)$$

$$R^{j}_{\ 00i} = -\delta_{i}^{j}\partial_{0}^{2}\phi - \partial_{i}\partial^{j}\phi + \mathcal{O}(\phi^2)$$

$$R^{j}_{\ l0i} = \delta_{li}\partial_{0}\partial^{j}\phi - \delta_{i}^{j}\partial_{0}\partial_{l}\phi + \mathcal{O}(\phi^2)$$

$$R^{0}_{\ l0i} = \mathcal{O}(\phi^2)$$

$$R^{0}_{\ lij} = -\delta_{lj}\partial_{i}\partial_{0}\phi + \delta_{li}\partial_{j}\partial_{0}\phi + \mathcal{O}(\phi^2)$$

$$R^{l}_{\ 0ij} = -\delta_{lj}^{j}\partial_{i}\partial_{0}\phi + \delta_{li}\partial_{j}\partial_{0}\phi + \mathcal{O}(\phi^2)$$

$$R^{l}_{\ kij} = \delta_{kj}\partial_{i}\partial^{l}\phi - \delta_{j}^{l}\partial_{i}\partial_{k}\phi - \delta_{ki}\partial_{j}\partial^{l}\phi + \delta_{i}^{l}\partial_{j}\partial_{k}\phi + \mathcal{O}(\phi^2)$$

## Exercise 6.3 (Killing vectors and conserved quantities)

### Exercise 6.3: Killing vectors and conserved quantities

(a) Consider the scalar quantity  $u^{\alpha}\xi_{\alpha}$ , with  $\overrightarrow{u}$  the four-velocity of a free-falling particle. What equation must be fulfilled by  $\overrightarrow{\xi}$  if this scalar remains constant along the particle's trajectory? (1.5 pt)

If  $u^{\alpha}\xi_{\alpha}$  is constant along the particle's trajectory it has to satisfy the following equation:  $\frac{d}{d\tau}\left(u^{\alpha}\xi_{\alpha}\right)=0\,,$ 

where  $\tau$  is the proper time of the particle.

This equation can be rewritten as follows,

$$\frac{du^{\alpha}}{d\tau}\xi_{\alpha} + u^{\alpha}\frac{d\xi_{\alpha}}{d\tau} = 0 \longrightarrow (u^{\beta}\nabla_{\beta}u^{\alpha})\xi_{\alpha} + u^{\alpha}u^{\beta}\nabla_{\beta}\xi_{\alpha} = 0.$$

A free particle satisfies the geodesic equation  $u^{\beta}\nabla_{\beta}u^{\alpha}=0$ , so we have  $\longrightarrow u^{\alpha}u^{\beta}\nabla_{\beta}\xi_{\alpha}=0$ .

Regardless of the 4-velocity of the free-falling particle this equation is automatically fulfilled if  $\nabla_{\beta}\xi_{\alpha}$  is antisymmetric, i.e. if

$$\nabla_{\beta} \xi_{\alpha} + \nabla_{\alpha} \xi_{\beta} = 0$$
 This is Killing's equation.

(b) Express (6) in spherical coordinates  $(t, r, \theta, \varphi)$ . (1 pt)

 $ds^{2} = -(1+2\phi)dt^{2} + (1-2\phi)(dx^{2} + dy^{2} + dz^{2})$ 

$$x = r\sin(\theta)\cos(\varphi)$$

$$dx = \sin(\theta)\cos(\varphi)dr + r\cos(\theta)\cos(\varphi)d\theta - r\sin(\theta)\sin(\varphi)d\varphi$$

$$y = r\sin(\theta)\sin(\varphi)$$

$$z = r\cos(\theta).$$

$$dx = \sin(\theta)\cos(\varphi)dr + r\cos(\theta)\sin(\varphi)d\theta - r\sin(\theta)\sin(\varphi)d\varphi$$

$$dy = \sin(\theta)\sin(\varphi)dr + r\cos(\theta)\sin(\varphi)d\theta + r\sin(\theta)\cos(\varphi)d\varphi$$

$$dz = \cos(\theta)dr - r\sin(\theta)d\theta.$$

This leads to:

$$ds^{2} = -(1+2\phi)dt^{2} + (1-2\phi)[dr^{2} + r^{2}\{d\theta^{2} + \sin^{2}(\theta)d\varphi^{2}\}]$$

(c) Assume now that the metric under consideration is static and spherically symmetric. Find two linearly independent Killing vectors in this particular spacetime, and prove that  $p_0$  and  $p_{\varphi}$  are conserved along the trajectory of a free-falling particle.

$$\nabla_{\beta}\xi_{\alpha} + \nabla_{\alpha}\xi_{\beta} = 0 \longrightarrow \partial_{\mu}\xi_{\nu} + \partial_{\nu}\xi_{\mu} - 2\Gamma^{\alpha}_{\mu\nu}\xi_{\alpha} = 0$$

$$ds^{2} = -(1+2\phi)dt^{2} + (1-2\phi)[dr^{2} + r^{2}\{d\theta^{2} + \sin^{2}(\theta)d\varphi^{2}\}]$$

For the spacetime determined by the line element (50) with  $\phi = \phi(r)$  there are only nine Christoffel symbols different from zero, namely  $\Gamma^{\theta}_{\theta r}$ ,  $\Gamma^{\theta}_{\varphi \varphi}$ ,  $\Gamma^{0}_{0r}$ ,  $\Gamma^{\varphi}_{\varphi r}$ ,  $\Gamma^{\varphi}_{\varphi \theta}$ ,  $\Gamma^{r}_{rr}$ ,  $\Gamma^{r}_{\theta \theta}$ ,  $\Gamma^{r}_{\varphi \varphi}$  and  $\Gamma^{r}_{00}$ . Each Killing vector has to satisfy the following set of equations:

$$\partial_{0}\xi_{0} - \Gamma^{r}{}_{00}\xi_{r} = 0$$

$$\partial_{0}\xi_{\theta} + \partial_{\theta}\xi_{0} = 0$$

$$\partial_{r}\xi_{r} - \Gamma^{r}{}_{rr}\xi_{r} = 0$$

$$\partial_{0}\xi_{\varphi} + \partial_{\varphi}\xi_{0} = 0$$

$$\partial_{\theta}\xi_{\theta} - \Gamma^{r}{}_{\theta\theta}\xi_{r} = 0$$

$$\partial_{r}\xi_{\theta} + \partial_{\theta}\xi_{r} - 2\Gamma^{\theta}{}_{r\theta}\xi_{\theta} = 0$$

$$\partial_{\varphi}\xi_{\varphi} - \Gamma^{\theta}{}_{\varphi\varphi}\xi_{\theta} - \Gamma^{r}{}_{\varphi\varphi}\xi_{r} = 0$$

$$\partial_{0}\xi_{r} + \partial_{r}\xi_{0} - 2\Gamma^{0}{}_{0r}\xi_{0} = 0$$

$$\partial_{\theta}\xi_{\varphi} + \partial_{\varphi}\xi_{\theta} - 2\Gamma^{\varphi}{}_{\varphi\theta}\xi_{\varphi} = 0$$

$$\partial_{\theta}\xi_{\varphi} + \partial_{\varphi}\xi_{\theta} - 2\Gamma^{\varphi}{}_{\varphi\theta}\xi_{\varphi} = 0$$

Let us see whether there exists a Killing vector with  $\xi_r = \xi_\theta = \xi_\varphi = 0$ . Using this in the Killing equations we find:  $\partial_{\varphi} \xi_0 = \partial_{\theta} \xi_0 = \partial_0 \xi_0 = 0$ 

$$\partial_r \xi_0 - 2\Gamma^0_{0r} \xi_0 = 0,$$
 so  $\xi_0 = \xi_0(r)$ 

Using  $\Gamma^0_{0r} = \frac{\partial_r \phi}{1+2\phi}$  we can solve the last equation and obtain:

$$\tilde{d}\xi = A\left(1 + 2\phi, 0, 0, 0\right) \longrightarrow \overrightarrow{\xi} = A\left(1, 0, 0, 0\right)$$

where A is an integration constant.

Now let us search for another Killing vector, but with  $\xi_r = \xi_\theta = \xi_0 = 0$ .

The following equations have to be fulfilled,

$$\partial_{\varphi} \xi_{\varphi} = \partial_{\theta} \xi_{r} = \partial_{0} \xi_{\varphi} = 0$$

$$\partial_{r} \xi_{\varphi} - 2\Gamma^{\varphi}_{\varphi r} \xi_{\varphi} = 0$$

$$\partial_{\theta} \xi_{\varphi} - 2\Gamma^{\varphi}_{\varphi \theta} \xi_{\varphi} = 0.$$

In this case  $\xi_{\varphi} = \xi_{\varphi}(r,\theta)$ . As  $\Gamma^{\varphi}_{\varphi r} = 1/r - \partial_r \phi/(1-2\phi)$  and  $\Gamma^{\varphi}_{\varphi \theta} = \cot g(\theta)$  we find:

The solution reads,

$$\partial_r \xi_{\varphi} - \frac{2}{r} \xi_{\varphi} = 0$$

$$\partial_{\theta} \xi_{\varphi} - 2\cot g(\theta) \xi_{\varphi} = 0.$$

$$\tilde{d}\xi = B\left(0, 0, 0, r^2 \sin^2(\theta)(1 - 2\phi)\right) \longrightarrow \overrightarrow{\xi} = B\left(0, 0, 0, 1\right)$$

Any linear combination of the two fields will also satisfy the Killing equation

It is clear that along the trajectory of a free-falling particle,

$$u_0 = C_1 \qquad ; \qquad u_\varphi = C_2 \,,$$

or, equivalently,

$$p_0 = \bar{C}_1$$
 ;  $p_{\varphi} = \bar{C}_2$ .

(d) Show that these conservation laws can be also obtained directly from the geodesic equation. (1 pt)

The geodesic equation reads,  $u^{\nu}\nabla_{\nu}u_{\mu} = 0 \longrightarrow u^{\nu}\partial_{\nu}u_{\mu} = \Gamma^{\alpha}_{\nu\mu}u_{\alpha}u^{\nu}$ 

$$\frac{du_{\mu}}{d\tau} = \frac{1}{2}g_{\beta\nu,\mu}u^{\beta}u^{\nu} \longrightarrow m\frac{dp_{\mu}}{d\tau} = \frac{1}{2}g_{\beta\nu,\mu}p^{\beta}p^{\nu}$$

Since the metric does not depend on time nor the angle  $\varphi$  we find that  $p_0$  and  $p_{\varphi}$  remain constant along the trajectory of a free-falling particle.