homework-05 RohanDekate

November 7, 2023

1 Homework 5

1.1 References

• Lectures 17-20 (inclusive).

1.2 Instructions

- Type your name and email in the "Student details" section below.
- Develop the code and generate the figures you need to solve the problems using this notebook.
- For the answers that require a mathematical proof or derivation you should type them using latex. If you have never written latex before and you find it exceedingly difficult, we will likely accept handwritten solutions.
- The total homework points are 100. Please note that the problems are not weighed equally.

```
[1]: import numpy as np
     np.random.seed(123456)
     np.set printoptions(precision=3)
     import matplotlib.pyplot as plt
     %matplotlib inline
     import seaborn as sns
     sns.set(rc={"figure.dpi":100, "savefig.dpi":100})
     sns.set_context("notebook")
     sns.set_style("ticks")
     import scipy
     import scipy.stats as st
     import urllib.request
     import os
     def download(
         url : str,
         local_filename : str = None
     ):
         """Download a file from a url.
         Arguments
                        -- The url we want to download.
         url
         local_filename -- The filemame to write on. If not
```

```
specified
"""

if local_filename is None:
    local_filename = os.path.basename(url)
urllib.request.urlretrieve(url, local_filename)
```

1.3 Student details

First Name: RohanLast Name: Dekate

• Email: dekate@purdue.edu

2 Problem 1 - Clustering Uber Pickup Data

In this problem you will analyze Uber pickup data collected during April 2014 around New York City. The complete data are freely on Kaggle. The data consist of a timestamp (which we are going to ignore), the latitude and longitude of the Uber pickup, and a base code (which we are also ignoring). The data file we are going to use is uber-raw-data-apr14.csv. As usual, you have to make it visible to this Jupyter notebook. On Google Colab, just run this:

And you can load it using pandas:

```
[3]: import pandas as pd p1_data = pd.read_csv('uber-raw-data-apr14.csv')
```

Here is how the data look like:

```
[4]: p1_data
```

```
[4]:
                     Date/Time
                                    Lat
                                             Lon
                                                    Base
    0
              4/1/2014 0:11:00 40.7690 -73.9549
                                                  B02512
              4/1/2014 0:17:00 40.7267 -74.0345
    1
                                                  B02512
    2
               4/1/2014 0:21:00
                                40.7316 -73.9873
                                                  B02512
    3
              4/1/2014 0:28:00
                                40.7588 -73.9776
                                                  B02512
    4
              4/1/2014 0:33:00
                                40.7594 -73.9722
                                                  B02512
    564511 4/30/2014 23:22:00
                                40.7640 -73.9744
                                                  B02764
    564512 4/30/2014 23:26:00
                                40.7629 -73.9672
                                                  B02764
    564513 4/30/2014 23:31:00
                                40.7443 -73.9889
                                                  B02764
    564514 4/30/2014 23:32:00
                                40.6756 -73.9405
                                                  B02764
    564515 4/30/2014 23:48:00
                                40.6880 -73.9608 B02764
```

[564516 rows x 4 columns]

If you have never played before with pandas, you can find a nice tutorial here.

We have half a million data points. Let's extract the latitude and longitude and put them in a numpy array:

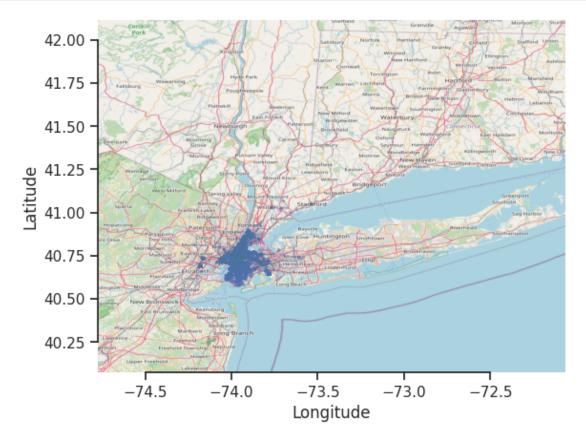
```
[5]: loc_data = p1_data[['Lon', 'Lat']] loc_data
```

```
[5]:
                Lon
                          Lat
           -73.9549 40.7690
     0
     1
           -74.0345 40.7267
     2
           -73.9873 40.7316
     3
           -73.9776 40.7588
           -73.9722 40.7594
     564511 -73.9744 40.7640
     564512 -73.9672 40.7629
     564513 -73.9889 40.7443
     564514 -73.9405 40.6756
     564515 -73.9608 40.6880
     [564516 rows x 2 columns]
```

Let's visualize these points on the map of New York City:

```
[6]: url = "https://github.com/PredictiveScienceLab/data-analytics-se/raw/master/
      →lecturebook/images/ny_map.png"
     download(url)
     ny_map = plt.imread('ny_map.png')
     box = ((loc_data.Lon.min(), loc_data.Lon.max(),
             loc_data.Lat.min(), loc_data.Lat.max()))
     fig, ax = plt.subplots(dpi=100)
     ax.scatter(
         loc_data.Lon,
         loc_data.Lat,
         zorder=1,
         alpha=0.5,
         c='b',
         s=0.001
     ax.set_xlim(box[0],box[1])
     ax.set_ylim(box[2],box[3])
     ax.imshow(
         ny_map,
         zorder=0,
         extent=box,
         aspect= 'equal'
     )
```

```
ax.set_xlabel('Longitude')
ax.set_ylabel('Latitude')
sns.despine(trim=True);
```



Machine learning algorithms will be a bit slow because we have over half a million data points. So, as you develop your code, use only 50K observations. Once you have a stable version of your code, modify the following code segment to use the entire dataset.

```
[7]: p1_train_data = loc_data[:50_000] p1_train_data
```

```
[7]:
                          Lat
                Lon
     0
           -73.9549
                     40.7690
     1
           -74.0345
                     40.7267
     2
           -73.9873
                     40.7316
     3
           -73.9776
                     40.7588
     4
           -73.9722
                     40.7594
     49995 -74.0261
                     40.6233
     49996 -73.9897
                     40.7033
     49997 -73.8539
                     40.8760
```

```
49998 -74.0020 40.7365
49999 -73.9927 40.7676
[50000 rows x 2 columns]
```

4 -73.864607

2.1 Part A - Splitting New York City into Subregions

Suppose you are assigned to split New York City into operating subregions with equal demand. When a pickup is requested in each subregion, only the drivers in that region are called. Note that this can become a challenging problem very quickly. We are not looking for the best possible answer here. We are looking for a data-informed heuristic solution that is good enough.

Do (at least) the following: + Use Kmeans clustering on the pickup data with different numbers of clusters; + Visualize the labels of the clusters on the map using different colors (see the hands-on activities); + Visualize the centers of the discovered Kmeans clusters (in red color); + Use common sense, e.g., ensure there are enough clusters so no region crosses the water. If it is impossible to get perfect results simply by Kmeans, feel free to ignore a small number of outliers as they could be handled manually; + Use MiniBatchKMeans, which is a much faster version of Kmeans suitable for large datasets (>10K observations);

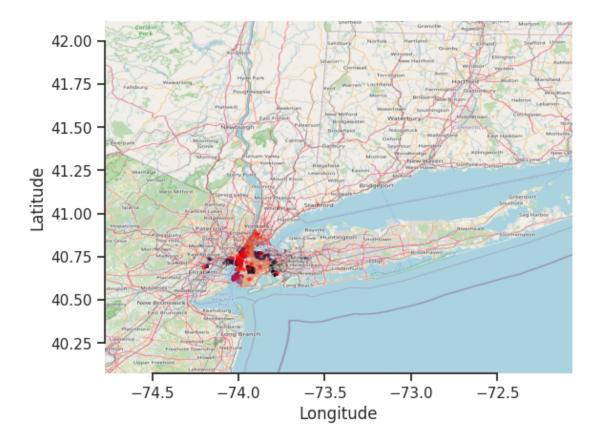
Answer with as many text and code blocks as you like below.

40.770741

```
[8]: # Your code here
    from sklearn.cluster import MiniBatchKMeans
    model = MiniBatchKMeans(n_clusters=25, random_state=0,__
      ⇔batch_size=512,n_init="auto").fit(loc_data)
    cluster_center = model.cluster_centers_
     # print(cluster_center)
    cluster_center = pd.DataFrame(cluster_center,columns = ['Center_Lon',__
      cluster_center.head()
[8]:
       Center_Lon Center_Lat
    0 -73.982539
                    40.756987
    1 -73.780347
                    40.654244
    2 -73.938363
                    40.684923
    3 -74.010686
                    40.714458
```

```
[9]: labels = model.predict(loc_data)
```

```
ax.scatter(
    loc_data.Lon,
    loc_data.Lat,
    zorder=1,
    alpha= 0.5,
    c=labels,
    s=0.001
ax.scatter(
    cluster_center.Center_Lon,
    cluster_center.Center_Lat,
    zorder=1,
    alpha= 0.5,
    c="red",
    marker="*",
    s=10
)
ax.set_xlim(box[0],box[1])
ax.set_ylim(box[2],box[3])
ax.imshow(
   ny_map,
    zorder=0,
    extent=box,
    aspect= 'equal'
ax.set_xlabel('Longitude')
ax.set_ylabel('Latitude')
sns.despine(trim=True);
```



Zooming inside the plot will segregate the cluster centers.

2.2 Part B - Create a Stochastic Model of Pickups

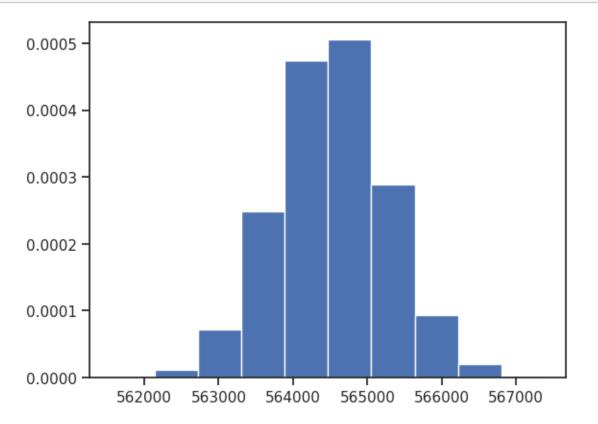
One of the key ingredients for a more sophisticated approach to optimizing the operations of Uber is the construction of a stochastic model of the demand for pickups. The ideal model for this problem is the Poisson Point Process. However, we will do something more straightforward, using the Gaussian mixture model and a Poisson random variable. The model will not have a time component, but it will allow us to sample the number and locations of pickups during a typical month. We will guide you through the process of constructing this model.

2.2.1 Subpart B.I - Random variable capturing the number of monthly pickups

Find the rate of monthly pickups (ignore the fact that months may differ by a few days) and use it to define a Poisson random variable corresponding to the monthly number of pickups. Use scipy.stats.poisson to initialize this random variable. Sample from it 10,000 times and plot the histogram of the samples to get a feeling about the corresponding probability mass function.

```
[11]: # Your code here
# rate of monthly pickups = 564516; as the data is for one month
poi_samples = np.random.poisson(564516, 10000)
plt.hist(poi_samples, density=True)
```





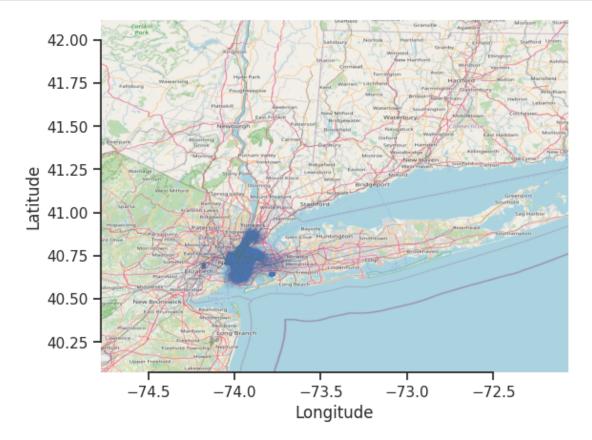
2.2.2 Subpart B.II - Sample some random monthly pickup numbers

Now that you have a model that gives you the number of pickups and a model that allows you to sample a pickup location, sample five different datasets (number of pickups and location of each pick) from the combined model and visualize them on the New York map.

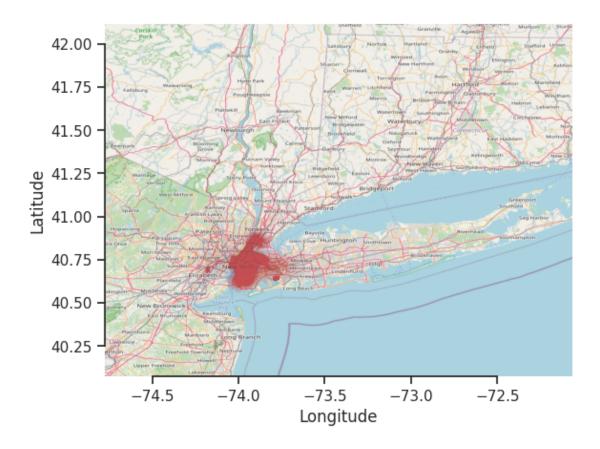
Hint: Don't get obsessed with making the model perfect. It's okay if a few of the pickups are on water.

```
[12]: # Your code here
from sklearn.mixture import GaussianMixture
model = GaussianMixture(n_components=25).fit(loc_data) # Gaussian Mixture Model
r = scipy.stats.poisson.rvs(564516, size=1) # Number of Pickups
s_data, labels = model.sample(r) # Sample Pickup location
```

```
box = ((loc_data.Lon.min(), loc_data.Lon.max(),
        loc_data.Lat.min(), loc_data.Lat.max()))
fig, ax = plt.subplots(dpi=100)
ax.scatter(
    s_data[:,0],
    s_data[:,1],
    zorder=1,
    alpha= 0.5,
    c="b",
    s=0.001
ax.set_xlim(box[0],box[1])
ax.set_ylim(box[2],box[3])
ax.imshow(
    ny_map,
    zorder=0,
    extent=box,
    aspect= 'equal'
)
ax.set_xlabel('Longitude')
ax.set_ylabel('Latitude')
sns.despine(trim=True);
```

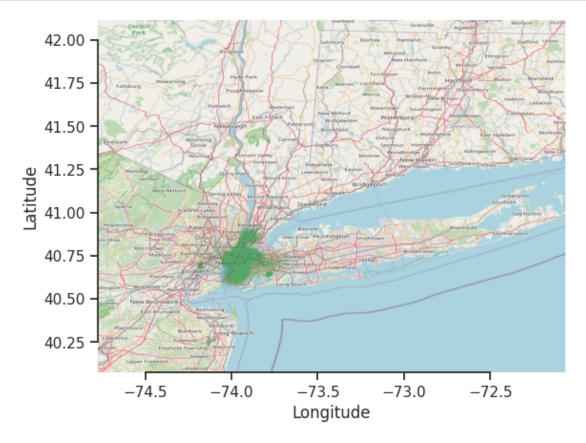


```
[14]: # Dataset 2
      r = scipy.stats.poisson.rvs(564516, size=1) # Sample Number of Monthly Pickups
      s_data, labels = model.sample(r) # Sample Pickup location
      url = "https://github.com/PredictiveScienceLab/data-analytics-se/raw/master/
       ⇔lecturebook/images/ny_map.png"
      download(url)
      ny_map = plt.imread('ny_map.png')
      box = ((loc_data.Lon.min(), loc_data.Lon.max(),
              loc_data.Lat.min(), loc_data.Lat.max()))
      fig, ax = plt.subplots(dpi=100)
      ax.scatter(
          s_data[:,0],
          s_data[:,1],
          zorder=1,
          alpha= 0.5,
          c="r",
          s=0.001
      ax.set_xlim(box[0],box[1])
      ax.set_ylim(box[2],box[3])
      ax.imshow(
          ny_map,
          zorder=0,
          extent=box,
          aspect= 'equal'
      ax.set_xlabel('Longitude')
      ax.set_ylabel('Latitude')
      sns.despine(trim=True);
```

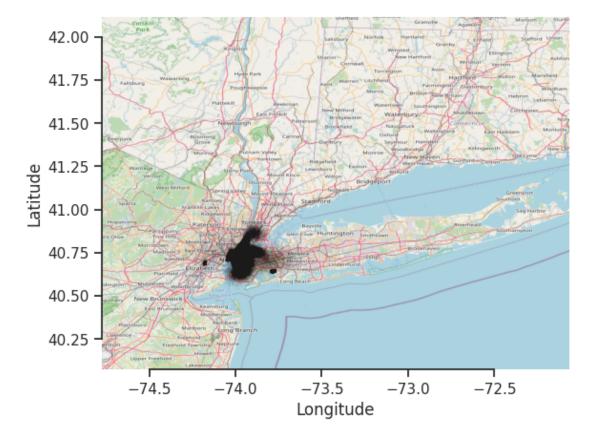


```
[15]: # Dataset 3
      r = scipy.stats.poisson.rvs(564516, size=1) # Sample Number of Monthly Pickups
      s_data, labels = model.sample(r) # Sample Pickup location
      url = "https://github.com/PredictiveScienceLab/data-analytics-se/raw/master/
       ⇔lecturebook/images/ny_map.png"
      download(url)
      ny_map = plt.imread('ny_map.png')
      box = ((loc_data.Lon.min(), loc_data.Lon.max(),
              loc_data.Lat.min(), loc_data.Lat.max()))
      fig, ax = plt.subplots(dpi=100)
      ax.scatter(
          s_data[:,0],
          s_data[:,1],
          zorder=1,
          alpha=0.5,
          c="g",
          s=0.001
      ax.set_xlim(box[0],box[1])
      ax.set_ylim(box[2],box[3])
      ax.imshow(
```

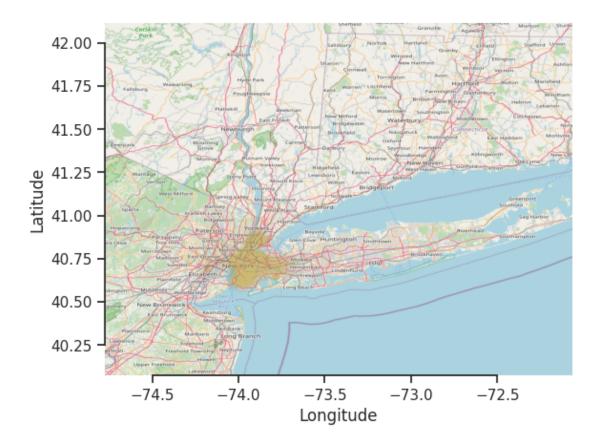
```
ny_map,
zorder=0,
extent=box,
aspect= 'equal'
)
ax.set_xlabel('Longitude')
ax.set_ylabel('Latitude')
sns.despine(trim=True);
```



```
s_data[:,0],
    s_data[:,1],
    zorder=1,
    alpha= 0.5,
    c="k",
    s=0.001
)
ax.set_xlim(box[0],box[1])
ax.set_ylim(box[2],box[3])
ax.imshow(
    ny_map,
    zorder=0,
    extent=box,
    aspect= 'equal'
ax.set_xlabel('Longitude')
ax.set_ylabel('Latitude')
sns.despine(trim=True);
```



```
[17]: # Dataset 5
      r = scipy.stats.poisson.rvs(564516, size=1) # Sample Number of Monthly Pickups
      s_data, labels = model.sample(r) # Sample Pickup location
      url = "https://github.com/PredictiveScienceLab/data-analytics-se/raw/master/
       ⇔lecturebook/images/ny_map.png"
      download(url)
      ny_map = plt.imread('ny_map.png')
      box = ((loc_data.Lon.min(), loc_data.Lon.max(),
              loc_data.Lat.min(), loc_data.Lat.max()))
      fig, ax = plt.subplots(dpi=100)
      ax.scatter(
          s_data[:,0],
          s_data[:,1],
          zorder=1,
          alpha= 0.5,
          c="y",
          s=0.001
      ax.set_xlim(box[0],box[1])
      ax.set_ylim(box[2],box[3])
      ax.imshow(
          ny_map,
          zorder=0,
          extent=box,
          aspect= 'equal'
      ax.set_xlabel('Longitude')
      ax.set_ylabel('Latitude')
      sns.despine(trim=True);
```



3 Problem 2 - Counting Celestial Objects

Consider this picture of a patch of sky taken by the Hubble Space Telescope.

Let's download it so that you have it here:

This picture includes many galaxies but also some stars. We will create a machine-learning model capable of counting the number of objects in such images. Our model will not be able to differentiate between the different types of objects and will not be very accurate. Still, it does form the basis of more sophisticated approaches. The idea is as follows: + Convert the picture to points sampled according to the intensity of light. + Apply Gaussian mixture on the resulting points. + Use the Bayesian Information Criterion to identify the number of components in the picture. + Associate the number of components with the actual number of celestial objects.

I will set you up with the first step. You will have to do the last three.

We are going to load the image with the Python Imaging Library (PIL), which allows us to apply a few basic transformations to the image:

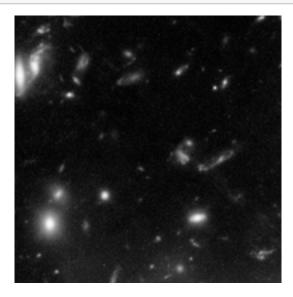
```
[19]: from PIL import Image
   hubble_image = Image.open('galaxies.png')
   # here is how to see the image
   hubble_image
```

Output hidden; open in https://colab.research.google.com to view.

Now, we are going to convert it to grayscale and crop it to make the problem a little bit easier:

```
[20]: img = hubble_image.convert('L').crop((100, 100, 300, 300)) img
```

[20]:

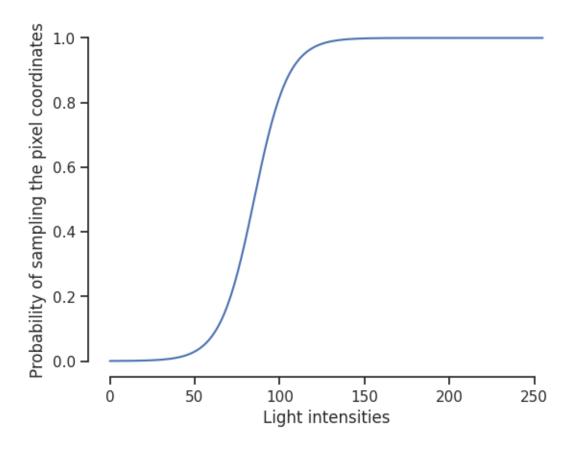


Remember that black-and white images are matrices:

The minimum number is 0, corresponding to black, and the maximum is 255, corresponding to white. Anything in between is some shade of gray.

Now, imagine that each pixel is associated with some coordinates. Without loss of generality, let's assume that each pixel is some coordinate in $[0,1]^2$. We will loop over each pixel and sample its coordinates in a way that increases with increasing light intensity. To achieve this, we will pass the intensity values of each pixel through a sigmoid with parameters that can be tuned. Here is this sigmoid:

```
[22]: intensities = np.linspace(0, 255, 255)
fig, ax = plt.subplots()
alpha = 0.1
beta = 255 / 3
ax.plot(
    intensities,
    1.0 / (1.0 + np.exp(-alpha * (intensities - beta)))
);
ax.set_xlabel('Light intensities')
ax.set_ylabel('Probability of sampling the pixel coordinates')
sns.despine(trim=True);
```



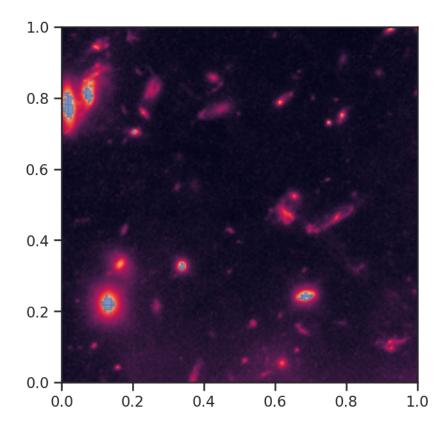
And here is the code that samples the pixel coordinates. I am organizing it into a function because we may want to use it with different pictures:

```
[23]: def sample_pixel_coords(img, alpha, beta):
          Samples pixel coordinates based on a probability defined as the sigmoid of \Box
       \hookrightarrow the intensity.
          Arguments:
                            The gray scale pixture from which we sample as an array
               alpha
                                The scale of the sigmoid
                                The offset of the sigmoid
               beta
          img_ar = np.array(img)
          x = np.linspace(0, 1, img_ar.shape[0])
          y = np.linspace(0, 1, img_ar.shape[1])
          X, Y = np.meshgrid(x, y)
          img_to_locs = []
          # Loop over pixels
          for i in range(img_ar.shape[1]):
```

```
for j in range(img_ar.shape[0]):
    # Calculate the probability of the pixel by looking at each
    # light intensity
    prob = 1.0 / (1.0 + np.exp(-alpha * (img_ar[j, i] - beta)))
    # Pick a uniform random number
    u = np.random.rand()
    # If u is smaller than the desired probability,
    # the consider the coordinates of the pixel sampled
    if u <= prob:
        img_to_locs.append((Y[i, j], X[-i-1, -j-1]))
# Turn img_to_locs into a numpy array
img_to_locs = np.array(img_to_locs)
return img_to_locs</pre>
```

Let's test it:

```
[24]: locs = sample_pixel_coords(img, alpha=0.1, beta=200)
    fig, ax = plt.subplots(dpi=150)
    ax.imshow(img, extent=((0, 1, 0, 1)), zorder=0)
    ax.scatter(
        locs[:, 0],
        locs[:, 1],
        zorder=1,
        alpha=0.5,
        c='b',
        s=1
    );
```



Note that playing with α and β makes the whole thing more or less sensitive to the light intensity. Complete the following function:

```
[25]: from sklearn.mixture import GaussianMixture
      def count_objs(img, alpha, beta, nc_min=1, nc_max=50):
          """Count objects in image.
          Arguments:
                               The image
              imq
              alpha
                              The scale of the sigmoid
                              The offset of the sigmoid
              beta
                               The minimum number of components to consider
              nc\_min
                              The maximum number of components to consider
              nc_max
          11 11 11
          locs = sample_pixel_coords(img, alpha, beta)
          # **** YOUR CODE HERE ****
          # Use BIC to search for the best GaussianMixture model
          # with components between nc_min and nc_max
          # YOU CAN PULL THIS OFF BY COPY-PASTING MATERIAL FROM
          # LECTURE 17
```

```
models = []
          # min bic = 1e10
          \# best_nc = None
          # best_model = None
          for nc in range(nc_min, nc_max):
           m = GaussianMixture(n_components=nc).fit(locs)
           bics[nc - 1] = m.bic(locs)
            models.append(m)
            # if min_bic > m.bic(locs):
            # min_bic = m.bic(locs)
            \# best_nc = nc
            # best_model = m
          # print(f"Minimum BIC is {min_bic}; #Components {best_nc};")
          # model with the smallest BIC is the best
          # print(bics)
          # print(min(bics))
          # Set the following variables
          # fig, ax = plt.subplots()
          # ax.bar(range(1, nc_max), bics)
          # ax.set_ylabel('BIC Score')
          # ax.set_xlabel('Number of components')
          # sns.despine(trim=True);
          best nc = np.argmin(bics)
          best_model = models[best_nc]
          return best_nc, best_model, locs
[26]: count_objs(img, alpha=1, beta=200, nc_min=1, nc_max=50)
[26]: (7,
       GaussianMixture(n_components=8),
       array([[0.005, 0.794],
              [0.005, 0.789],
              [0.005, 0.784],
              [0.005, 0.779],
              [0.005, 0.774],
              [0.005, 0.769],
              [0.01, 0.809],
              [0.01 , 0.804],
              [0.01, 0.799],
              [0.01, 0.794],
              [0.01, 0.789],
              [0.01, 0.784],
              [0.01, 0.779],
              [0.01, 0.774],
              [0.01, 0.769],
```

bics = np.ndarray((nc_max - nc_min,))

[0.01, 0.764],

```
[0.01, 0.759],
[0.01, 0.754],
[0.015, 0.814],
[0.015, 0.809],
[0.015, 0.804],
[0.015, 0.799],
[0.015, 0.794],
[0.015, 0.789],
[0.015, 0.784],
[0.015, 0.779],
[0.015, 0.774],
[0.015, 0.769],
[0.015, 0.764],
[0.015, 0.759],
[0.015, 0.754],
[0.015, 0.749],
[0.015, 0.744],
[0.02, 0.809],
[0.02, 0.804],
[0.02, 0.799],
[0.02, 0.794],
[0.02, 0.789],
[0.02, 0.784],
[0.02, 0.779],
[0.02, 0.774],
[0.02, 0.769],
[0.02, 0.764],
[0.02, 0.759],
[0.02, 0.754],
[0.02, 0.749],
[0.02, 0.744],
[0.025, 0.804],
[0.025, 0.799],
[0.025, 0.794],
[0.025, 0.789],
[0.025, 0.784],
[0.025, 0.779],
[0.025, 0.774],
[0.025, 0.769],
[0.025, 0.764],
[0.025, 0.759],
[0.025, 0.754],
[0.025, 0.749],
[0.03, 0.789],
[0.03, 0.784],
[0.03, 0.779],
```

[0.03, 0.774],

```
[0.03, 0.769],
[0.03, 0.764],
[0.03, 0.759],
[0.06, 0.819],
[0.065, 0.824],
[0.065, 0.819],
[0.065, 0.814],
[0.065, 0.809],
[0.065, 0.804],
[0.07, 0.824],
[0.07, 0.819],
[0.07, 0.814],
[0.07, 0.809],
[0.07, 0.804],
[0.07, 0.799],
[0.07, 0.794],
[0.075, 0.829],
[0.075, 0.824],
[0.075, 0.819],
[0.075, 0.814],
[0.075, 0.809],
[0.075, 0.804],
[0.075, 0.799],
[0.075, 0.794],
[0.08, 0.829],
[0.08, 0.824],
[0.08, 0.819],
[0.08, 0.814],
[0.08, 0.809],
[0.08, 0.804],
[0.08, 0.799],
[0.08, 0.794],
[0.116, 0.226],
[0.116, 0.221],
[0.116, 0.216],
[0.121, 0.236],
[0.121, 0.231],
[0.121, 0.226],
[0.121, 0.221],
[0.121, 0.216],
[0.121, 0.211],
[0.121, 0.206],
[0.126, 0.236],
[0.126, 0.231],
[0.126, 0.226],
[0.126, 0.221],
```

[0.126, 0.216],

```
[0.126, 0.211],
[0.126, 0.206],
[0.126, 0.201],
[0.131, 0.241],
[0.131, 0.236],
[0.131, 0.231],
[0.131, 0.226],
[0.131, 0.221],
[0.131, 0.216],
[0.131, 0.211],
[0.131, 0.206],
[0.131, 0.201],
[0.136, 0.236],
[0.136, 0.231],
[0.136, 0.226],
[0.136, 0.221],
[0.136, 0.216],
[0.136, 0.211],
[0.136, 0.206],
[0.136, 0.201],
[0.141, 0.231],
[0.141, 0.226],
[0.141, 0.221],
[0.141, 0.216],
[0.141, 0.211],
[0.141, 0.206],
[0.146, 0.221],
[0.146, 0.216],
[0.201, 0.704],
[0.206, 0.704],
[0.332, 0.332],
[0.332, 0.327],
[0.332, 0.322],
[0.337, 0.332],
[0.337, 0.327],
[0.337, 0.322],
[0.342, 0.332],
[0.342, 0.327],
[0.342, 0.322],
[0.673, 0.241],
[0.673, 0.236],
[0.678, 0.246],
```

[0.678, 0.241], [0.678, 0.236], [0.683, 0.251], [0.683, 0.246], [0.683, 0.241],

```
[0.683, 0.236],
[0.688, 0.251],
[0.688, 0.246],
[0.688, 0.241],
[0.688, 0.236],
[0.693, 0.251],
[0.693, 0.246],
[0.693, 0.241],
[0.693, 0.246],
[0.698, 0.246],
[0.698, 0.241],
[0.704, 0.246],
[0.754, 0.729],
[0.789, 0.749],
[0.93, 0.995]]))
```

Once you have completed the code, try out the following images. Feel free to play with α and β to improve the performance. Do not try to make a perfect model. We would have to go beyond the Gaussian mixture model to do so. This is just a homework problem.

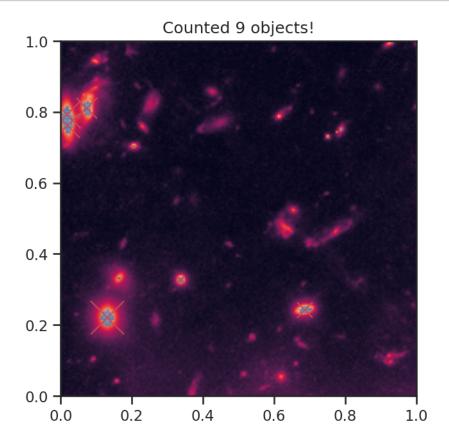
Here is a helpful function that you can use to visualize the results:

```
[27]: def visualize_counts(img, objs, model, locs):
          """Visualize the counts.
          Arguments
          imq
                 -- The image.
                 -- Returned by count objs()
          objs
          model -- Returned by count_objs()
          locs
                 -- Returned by count_objs()
          11 11 11
          fig, ax = plt.subplots(dpi=150)
          ax.imshow(img, extent=((0, 1, 0, 1)))
          for i in range(model.means_.shape[0]):
              ax.plot(
                  model.means_[i, 0],
                  model.means_[i, 1],
                  'rx',
                  markersize=(
                      10.0 * model.weights_.shape[0]
                      * model.weights_[i]
                  )
              )
          ax.scatter(
              locs[:, 0],
              locs[:, 1],
              zorder=1,
              alpha=0.5,
```

```
c='b',
s=1
)
ax.set_title('Counted {0:d} objects!'.format(objs));
```

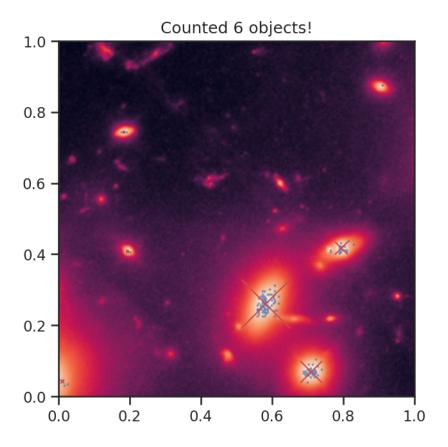
Here is how to use it:

```
[28]: objs, model, locs = count_objs(img, alpha=1.0, beta=200)
visualize_counts(img, objs, model, locs)
```



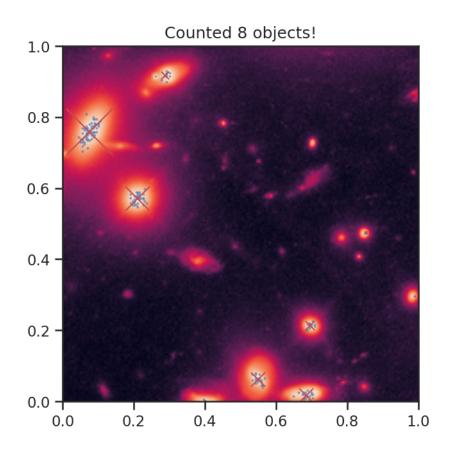
Try this image:

```
[29]: img = hubble_image.convert('L').crop((200, 200, 400, 400))
objs, model, locs = count_objs(img, alpha=.1, beta=250)
visualize_counts(img, objs, model, locs)
```



And this one:

```
[30]: img = hubble_image.convert('L').crop((300, 300, 500, 500))
objs, model, locs = count_objs(img, alpha=.1, beta=250)
visualize_counts(img, objs, model, locs)
```



4 Problem 3 - Filtering of an Oscillator with Damping

Assume that you are dealing with a one-degree-of-freedom system which follows the equation:

$$\ddot{x} + 2\zeta\omega_0\dot{x} + \omega_0^2x = u_0\cos(\omega t),$$

where x = x(t) is the generalized coordinate of the oscillator at time t, and the parameters ζ , ω_0 , u_0 , and ω are known to you (we will give them specific values later). Furthermore, assume that you are making noisy observations of the *absolute acceleration* at discrete timesteps Δt (also known):

$$y_j = \ddot{x}(j\Delta t) - u_0 \cos(\omega t) + w_j,$$

for j = 1, ..., n, where $w_j \sim N(0, \sigma^2)$ with σ^2 also known. Finally, assume that the initial conditions for the position and the velocity (you need both to get a unique solution) are given by:

$$x_0 = x(0) \sim N(0, \sigma_x^2),$$

and

$$v_0 = \dot{x} \sim N(0, \sigma_v^2).$$

Of course, assume that σ_x^2 and σ_v^2 are specific numbers we will specify below.

Before we go over the questions, let's write code that generates the actual trajectory of the system at some random initial conditions and some observations. We will use the code to generate a synthetic dataset with known ground truth, which you will use in your filtering analysis.

The first step we need to do is to turn the problem into a first-order differential equation. We set:

$$\mathbf{x} = \begin{bmatrix} x \\ \dot{x} \end{bmatrix}.$$

Assuming $\mathbf{x} = (x_1, x_2)$, then the dynamics are described by:

$$\dot{\mathbf{x}} = \begin{bmatrix} \dot{x} \\ \ddot{x} \end{bmatrix} = \begin{bmatrix} x_2 \\ -2\zeta\omega_0\dot{x} - \omega_0^2x + u_0\cos(\omega t) \end{bmatrix} = \begin{bmatrix} x_2 \\ -2\zeta\omega_0x_2 - \omega_0^2x_1 + u_0\cos(\omega t) \end{bmatrix}$$

The initial conditions are of course, just:

$$\mathbf{x}_0 = \begin{bmatrix} x_0 \\ v_0 \end{bmatrix}.$$

This first-order system can solved using scipy.integrate.solve_ivp. Here is how:

[31]: from scipy.integrate import solve_ivp

You need to define the right hand side of the equation

And here is how you solve it for given initial conditions and parameters:

```
[32]: # Initial conditions
      x0 = np.array([0.0, 1.0])
      # Natural frequency
      omega0 = 2.0
      # Dumping factor
      zeta = 0.4
      # External forcing amplitude
      u0 = 0.5
      # Excitation frequency
      omega = 2.1
      # Timestep
      dt = 0.1
      # The final time
      final_time = 10.0
      # The number of timesteps to get the final time
      n_steps = int(final_time / dt)
      # The times on which you want the solution
      t_eval = np.linspace(0, final_time, n_steps)
      # The solution
      sol = solve_ivp(rhs, (0, final_time), x0, t_eval=t_eval, args=(omega0, zeta,_u
       ⇒u0, omega))
```

The solution is stored in the sol variable:

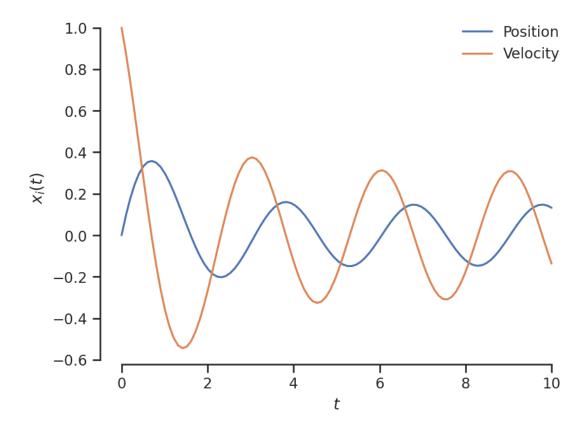
```
[33]: sol.y.shape
```

[33]: (2, 100)

The shape of sol.y is (2, 100), which means that we have 100 timesteps and two variables (position

and velocity). Let's plot the position and the velocity:

```
[34]: fig, ax = plt.subplots(dpi=150)
    ax.plot(t_eval, sol.y[0, :], label='Position')
    ax.plot(t_eval, sol.y[1, :], label='Velocity')
    ax.set_xlabel('$t$')
    ax.set_ylabel('$x_i(t)$')
    plt.legend(loc='best', frameon=False)
    sns.despine(trim=True);
```



Let's now generate some synthetic observations of the acceleration with some given Gaussian noise. To get the acceleration, you can do this:

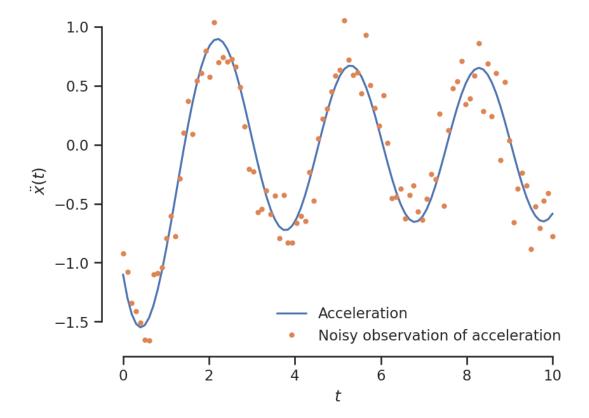
```
[35]: true_acc = np.array([rhs(t, x, omega0, zeta, u0, omega)[1] for (t, x) in_u 

\(\timeszip(t_eval, sol.y.T)])
```

Let's add some noise:

```
[36]: sigma_r = 0.2
observations = true_acc + sigma_r * np.random.randn(true_acc.shape[0])
fig, ax = plt.subplots(dpi=150)
```

```
ax.plot(t_eval, true_acc, label='Acceleration')
ax.plot(
    t_eval,
    observations,
    '.',
    label='Noisy observation of acceleration'
)
ax.set_xlabel('$t$')
ax.set_ylabel(r'$\ddot{x}(t)$')
plt.legend(loc='best', frameon=False)
sns.despine(trim=True);
```



Okay. Now, imagine that you only see the noisy observations of the acceleration. The filtering goal is to recover the state of the underlying system (as well as its acceleration). I am going to guide you through the steps you need to follow.

4.1 Part A - Discretize time (Transitions)

Use the Euler time discretization scheme to turn the continuous dynamical system into a discretetime dynamical system like this:

$$\mathbf{x}_{j+1} = \mathbf{A}\mathbf{x}_j + Bu_j + \mathbf{z}_j,$$

where

$$\mathbf{x}_{i} = \mathbf{x}(j\Delta t),$$

$$u_i = u(j\Delta t),$$

and \mathbf{z}_j is properly chosen process noise term. You should derive and provide mathematical expressions for the following: + The 2 × 2 transition matrix \mathbf{A} . + The 2 × 1 control "matrix" B. + The process covariance \mathbf{Q} . For the process covariance, you may choose your values by hand.

Answer:

$$\ddot{x} + 2\zeta\omega_0\dot{x} + \omega_0^2x = u_0\cos(\omega t),$$

Let

 $\dot{x}=v$

$$\ddot{x}=v=-2\zeta\omega_0v-\omega_0^2x+u_0\cos(\omega t)$$

Using Forward Euler Discretization

$$x_{n+1} = x_n + \Delta t v_n$$

$$v_{n+1} = v_n + \Delta t (-2\zeta \omega_0 v_n - \omega_0^2 x_n + u_0 \cos(n\omega \Delta t))$$

$$\begin{bmatrix} x_{n+1} \\ v_{n+1} \end{bmatrix} = \begin{bmatrix} 1 & \Delta t \\ -\omega_0^2 \Delta t & 1 - 2\zeta \omega_0 \Delta t \end{bmatrix} \begin{bmatrix} x_n \\ v_n \end{bmatrix} + \begin{bmatrix} 0 \\ u_0 \cos(n\omega \Delta t) \Delta t \end{bmatrix}$$

```
[37]: # You should be using the parameters dt, omegaO, zeta, etc.
    # from above
A = np.array(
        [1, dt],
        [-omegaO ** 2 *dt, 1-2.0 * zeta * omegaO * dt]
        ]
)

B = np.array(
        [0],
        [u0*dt] # Only constant terms are used here
        ]
)
sigma_x = 1e-2 # Values for Q are hand chosen and the textbook example is used_uas reference
sigma_v = 1e-2
```

4.2 Part B - Discretize time (Emissions)

Establish the map that takes you from the states to the accelerations at each timestep. That is, specify:

$$y_j = \mathbf{C}\mathbf{x}_j + w_j,$$

where

$$y_j = \ddot{x}(j\Delta t) - u_0 \cos(\omega t) + w_j,$$

and w_j is a measurement noise. You should derive and provide mathematical expressions for the following: + The 1×2 emission matrix \mathbf{C} . + The 1×1 covariance "matrix" R of the measurement noise.

Answer:

Replacing \ddot{x} in y_i

$$y_j = -2\zeta\omega_0x_2 - \omega_0^2x_1 + u_0\cos(\omega t) - u_0\cos(\omega t) + w_j$$

Ignoring w_i for calculating C

$$y_j = -2\zeta\omega_0x_2 - \omega_0^2x_1$$

$$y_j = [-\omega_0^2 - 2\zeta\omega_0]x_j$$

This is the emission Matrix C

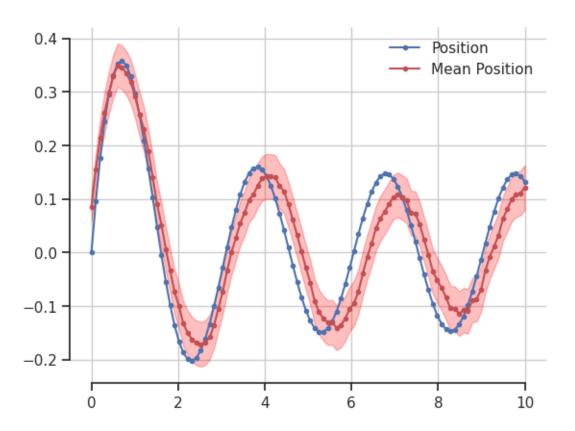
The Covariance Matrix R is w_j and is provided in the problem.

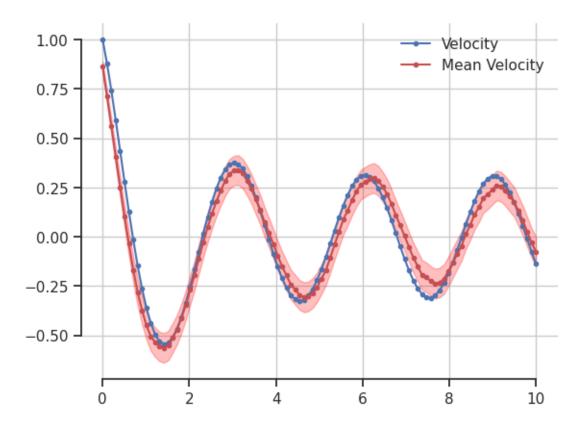
4.3 Part C - Apply the Kalman filter

Use FilterPy (see the hands-on activity of Lecture 20) to infer the unobserved states given the noisy observations of the accelerations. Plot time-evolving 95% credible intervals for the position and the velocity along with the true unobserved values of these quantities (in two separate plots).

```
[39]: # Sample the trajectory
      us = np.zeros((n_steps, 1))
      for u in range(us.shape[0]):
        us[u] = u0 * np.cos(omega * t_eval[u])
      print(us.shape)
     (100, 1)
[40]: # Your answer here (as many code and text blocks as you want)
      !pip install filterpy
      from filterpy.kalman import KalmanFilter
      kf = KalmanFilter(dim_x=2, dim_z=1)
     Collecting filterpy
       Downloading filterpy-1.4.5.zip (177 kB)
                                 178.0/178.0
     kB 4.0 MB/s eta 0:00:00
       Preparing metadata (setup.py) ... done
     Requirement already satisfied: numpy in /usr/local/lib/python3.10/dist-packages
     (from filterpy) (1.23.5)
     Requirement already satisfied: scipy in /usr/local/lib/python3.10/dist-packages
     (from filterpy) (1.11.3)
     Requirement already satisfied: matplotlib in /usr/local/lib/python3.10/dist-
     packages (from filterpy) (3.7.1)
     Requirement already satisfied: contourpy>=1.0.1 in
     /usr/local/lib/python3.10/dist-packages (from matplotlib->filterpy) (1.1.1)
     Requirement already satisfied: cycler>=0.10 in /usr/local/lib/python3.10/dist-
     packages (from matplotlib->filterpy) (0.12.1)
     Requirement already satisfied: fonttools>=4.22.0 in
     /usr/local/lib/python3.10/dist-packages (from matplotlib->filterpy) (4.43.1)
     Requirement already satisfied: kiwisolver>=1.0.1 in
     /usr/local/lib/python3.10/dist-packages (from matplotlib->filterpy) (1.4.5)
     Requirement already satisfied: packaging>=20.0 in
     /usr/local/lib/python3.10/dist-packages (from matplotlib->filterpy) (23.2)
     Requirement already satisfied: pillow>=6.2.0 in /usr/local/lib/python3.10/dist-
     packages (from matplotlib->filterpy) (9.4.0)
     Requirement already satisfied: pyparsing>=2.3.1 in
     /usr/local/lib/python3.10/dist-packages (from matplotlib->filterpy) (3.1.1)
     Requirement already satisfied: python-dateutil>=2.7 in
     /usr/local/lib/python3.10/dist-packages (from matplotlib->filterpy) (2.8.2)
     Requirement already satisfied: six>=1.5 in /usr/local/lib/python3.10/dist-
     packages (from python-dateutil>=2.7->matplotlib->filterpy) (1.16.0)
```

```
Building wheels for collected packages: filterpy
       Building wheel for filterpy (setup.py) ... done
       Created wheel for filterpy: filename=filterpy-1.4.5-py3-none-any.whl
     size=110459
     sha256=16069aeda1b22f7f418a30ee72b9ecca1a96cca85e67ef5f9dc208d3ef1565e2
       Stored in directory: /root/.cache/pip/wheels/0f/0c/ea/218f266af4ad626897562199
     fbbcba521b8497303200186102
     Successfully built filterpy
     Installing collected packages: filterpy
     Successfully installed filterpy-1.4.5
[41]: # INITIAL CONDITIONS
      # initial mean
      mu0 = x0
      # initial covariance
      V0 = np.array([0.01**2, 0.01**2]) * np.eye(2)
      kf.x = mu0
     kf.P = V0
     kf.Q = Q
     kf.R = R
     kf.H = C
      kf.F = A
     kf.B = B
[42]: means, covs, _, _ = kf.batch_filter(observations, us=us)
[43]: plt.plot(t_eval, sol.y[0, :], 'b.-', label='Position')
      plt.plot(t_eval, means[:, 0],'r.-', label='Mean Position')
      plt.fill_between(t_eval,(
                      means[:, 0]
                      -2.0 * np.sqrt(covs[:, 0, 0])
                  ),
                      means[:, 0]
                      + 2.0 * np.sqrt(covs[:, 0, 0])
                  ),
                  color='red',
                  alpha=0.25)
      plt.legend(loc='best', frameon=False)
      plt.grid()
      sns.despine(trim=True);
```

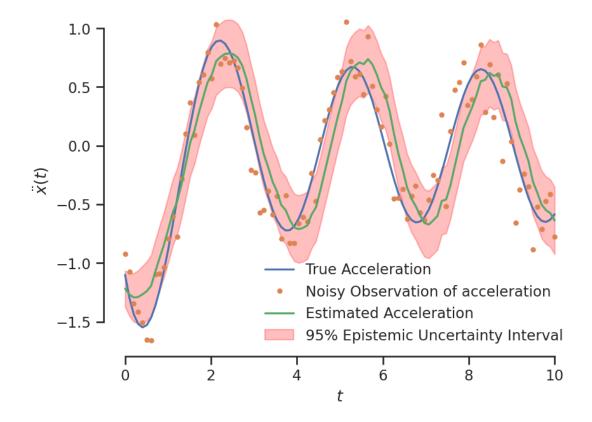




4.4 Part D - Quantify and visualize your uncertainty about the actual acceleration value

Use standard uncertainty propagation techniques to quantify your epistemic uncertainty about the true acceleration value. You will have to use the inferred states of the system and the dynamical model. This can be done either analytically or by Monte Carlo. It's your choice. In any case, plot time-evolving 95% credible intervals for the acceleration (epistemic only), the true unobserved values, and the noisy measurements.

```
acc_low = u0 * np.cos(omega * t_eval[i]) - omega0**2*pos_lb -_
 →2*zeta*omega0*vel_lb
 acc_lb.append(acc_low)
 acc_high = u0 * np.cos(omega * t_eval[i]) - omega0**2*pos_ub -_
 →2*zeta*omega0*vel ub
 acc_ub.append(acc_high)
fig, ax = plt.subplots(dpi=150)
ax.plot(t_eval, true_acc, label='True Acceleration')
ax.plot(
   t_eval,
   observations,
    '.',
   label='Noisy Observation of acceleration'
ax.plot(t_eval,acc_mu,label="Estimated Acceleration")
ax.fill_between(t_eval,acc_lb,acc_ub,label = "95% Epistemic Uncertainty"
 ax.set_xlabel('$t$')
ax.set_ylabel(r'$\ddot{x}(t)$')
plt.legend(loc='best', frameon=False)
sns.despine(trim=True);
```



[45]: