

Syllabus :-

⇒ Review of single electron system

→ Hydrogen atom: Schrodinger equation, Summary of energy levels, fine structure of hydrogen atom, Lamb shift, Hyperfine structure, Interaction of electron with the external field, The Stark effect (electric field), The Zeeman effect (magnetic field (weak)), Paschen-Back effect.

1 st Aug.	Wed.
2 nd Aug.	Thu.
5 th Aug.	Mon.

Extra-class
1st Aug → 2-
6th Aug → 3-8

⇒ Many-electron atoms :-

Central-field approximation & Hartree-Fock method, Thomas-Fermi model, LS, J-J coupling, Alkali spectra, He, complex atom.

Molecules :- electron-structure of molecules.

Molecular spectra - Rotational & vibrational spectra.

Diatomic molecule, Infra-red spectra, vibration-rotation spectra.

— Spectroscopic technique :- Interferometer, FTIR, RAMAN, NMR, ESR,

→ Some Frontier of Atomic & Molecular spectroscopy

→ Bose-Einstein Condensate

→ Atomic clock

→ Quantum computer

RECAP ⇔ QM

→ state

→ may be complex.

$$\phi_k = A \exp[i(\mathbf{k} \cdot \mathbf{r} - \omega t)]$$

Momentum

$$\hat{p} = -i\hbar \frac{\partial \phi_k}{\partial \mathbf{r}} = \boxed{\hbar \mathbf{k}} \phi_k$$

$$\omega_k = \frac{\hbar k^2}{2m}$$

$$E_k = \hbar \omega_k = \frac{\hbar^2 k^2}{2m}$$

$$\hat{H} = \frac{\hat{p}^2}{2m} \quad \hat{p} = -i\hbar \frac{\partial}{\partial \mathbf{r}}$$

$$E_k = \hbar \omega_k = \frac{\hbar^2 k^2}{2m}$$

$$\int |\phi_k(\mathbf{r})|^2 d^3r = 1$$

Energy Spectra -

$$\hat{H} \phi_k(r) = E_k \phi_k(r)$$

Eigen energy

Hamiltonian operator

Good Quantum Number n, l, m_l



$$\Psi(r) = \sum c_k \phi_k(r)$$

$$\int \phi_k^* \phi_{k'} d^3r = \delta_{kk'}$$

$$\Psi(r) = \sum_k c_k \phi_k$$

\Rightarrow The prob. to find the system with energy state E_k will be \uparrow Prob. Amplitude to find in state E_k

$$P_k = |c_k|^2 = \left| \int d^3r \phi_k^*(r) \Psi(r) \right|^2$$

In this case expectation (or mean) of the energy E in the state $\Psi(r)$ is given by —

$$\bar{E} = \langle \hat{H} \rangle = \int d^3r \Psi^*(r) \hat{H} \Psi(r)$$

$$\langle \hat{H} \rangle = \sum_k |c_k|^2 E_k$$

\rightarrow Quantum ket state

$$\hat{H} |\phi_k\rangle = E_k |\phi_k\rangle$$

$$\phi_k(r) = \langle r | \phi_k \rangle$$

\rightarrow Projection of quantum state on the position ket.

$$\phi_k(p) = \langle p | \phi_k \rangle$$

$$[\hat{r}, \hat{p}] = i\hbar \neq 0$$

$$[\hat{L}_i, \hat{L}_j] = i\hbar \epsilon_{ijk} \hat{L}_k$$

$$[\hat{L}^2, \hat{L}_z] = 0$$

$$\hat{L} = \hat{r} \times \hat{p}$$

$$\hat{L}^2$$

$$[\hat{r}_i \times \hat{p}_i, \hat{r}_j \times \hat{p}_j]$$

$$\hat{L}^2 |l, m\rangle = \hbar^2 l(l+1) |l, m\rangle$$

$$\hat{L}_z |l, m\rangle = \hbar m |l, m\rangle$$

$$[\hat{S}_i, \hat{S}_j] = i\hbar \epsilon_{ijk} \hat{S}_k$$

$$\hat{S}^2 |s, m_s\rangle = \hbar^2 s(s+1) |s, m_s\rangle$$

$$\hat{S}_z |s, m_s\rangle = \hbar m_s |s, m_s\rangle$$

Tensor product

$$\hat{J} = \hat{L}_1 + \hat{L}_2$$

New angular momentum

$$\hat{J}^2 |j, l_1, l_2, m_j\rangle = \hbar^2 j(j+1) |j, l_1, l_2, m_j\rangle$$

$$\hat{J}_z |j, l_1, l_2, m_j\rangle = m_j \hbar |j, l_1, l_2, m_j\rangle$$

$$|l_1 - l_2| \leq j \leq l_1 + l_2$$

$$\Rightarrow \Psi(r, t=0) = \frac{1}{\sqrt{14}} [2\psi_{100}(r) - 3\psi_{200}(r) + \psi_{322}(r)]$$

$$\langle \hat{H} \rangle, \langle \hat{L}^2 \rangle, \langle \hat{L}_z \rangle$$

$$n=3 \quad \begin{matrix} n^2 & 1 & 3 & 5 \\ 2 & 0 & 1 & 5 \end{matrix}$$

$$\Psi(r) = \sum c_k \phi_k(r)$$

$$|\Psi\rangle = \frac{1}{\sqrt{14}} [2|100\rangle - 3|200\rangle + |322\rangle]$$

$$\langle \hat{H} \rangle = \frac{\hbar^2}{28m} \left[\frac{(100)^2 \times 4 + (200)^2 \times 9 + (322)^2 \times 1}{14} \right]$$

$$|\Psi(0)\rangle = \sum_n c_n |E_n\rangle$$

$$|\Psi(0)\rangle = \sum_n |E_n\rangle \langle E_n | \Psi(0) \rangle$$

$$= \sum_n \langle E_n | \Psi(0) \rangle |E_n\rangle$$

$$\langle \Psi | H | E_n \rangle = E_n \langle \Psi | E_n \rangle$$

$$\langle \Psi(0) | H | \Psi(0) \rangle =$$

$$\langle \Psi(0) | E_n | \Psi(0) \rangle = E_n$$

$$\frac{2\pi}{\lambda} = k$$

$$\langle \sum_n c_n |E_n\rangle \langle E_n| \rangle = \sum_n c_n E_n c_n^* = \sum_n |c_n|^2 E_n$$

$$\frac{2}{14}$$

$$\Psi(x) = \sum_n c_n \phi_n$$

$$\Psi(r) = A \exp[i(k \cdot r - \omega t)]$$

$$\Psi(r) = A e^{i(k \cdot r)} = A \exp$$

$$E = \hbar \omega$$

Hydrogen like atom (one electron atom)

H (atom), He⁺ (Helium ion), Li²⁺ (Lithium ion), etc.

$$H = \frac{p^2}{2\mu} - \frac{Ze^2}{4\pi\epsilon_0 r} \quad (\text{Born CM frame})$$

Schrodinger equation:

$$\left[\frac{-\hbar^2}{2\mu} \nabla^2 - \frac{Ze^2}{4\pi\epsilon_0 r} \right] \psi(r) = E_{n\ell m} \psi(r)$$

Principal v.n.
Azimuthal
Magnetic

Quantum number

$$\psi_{n\ell m}(r, \theta, \phi) = R_{n\ell}(r) Y_{\ell}^m(\theta, \phi)$$

Spherical harmonics wave function

Radial

$$\left\{ \frac{-\hbar^2}{2\mu} \left[\frac{1}{r^2} \frac{\partial}{\partial r} \left(r^2 \frac{\partial}{\partial r} \right) - \frac{\ell(\ell+1)}{4\pi\epsilon_0 r^2} \right] - \frac{Ze^2}{4\pi\epsilon_0 r} \right\} R_{n,\ell}(r) =$$

$$E_{n\ell m} R_{n,\ell}(r)$$

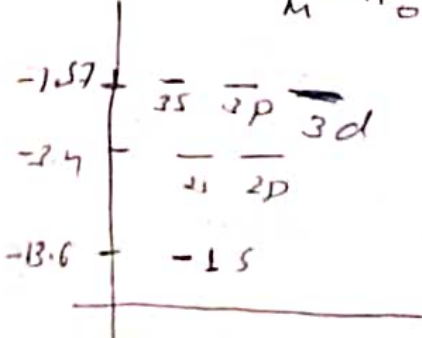
$$u_{r\ell} = r R_{n,\ell}(r)$$

$$\frac{\partial^2 u_{r\ell}}{\partial r^2} + \frac{2\mu}{\hbar^2} [E_{n\ell m} - V_{\text{eff.}}(r)] u_{n,\ell}(r) = 0$$

$$V_{\text{eff}} = \frac{-Ze^2}{4\pi\epsilon_0 r} + \frac{\ell(\ell+1)}{2\mu r^2} \hbar^2$$

$$E_n = \frac{-Ze^2 \hbar^2}{(4\pi\epsilon_0) a_{\mu} 2\hbar^2}$$

for H atom
 $a_{\mu} = a_0$
 $a_{\mu} = \frac{m}{\mu} a_0$



-157	-35	-3p	-3d
-34	-25	-2p	
-13.6	-15		

$$\Delta m = 0, \pm 1$$

$$\Delta l = \pm 1$$

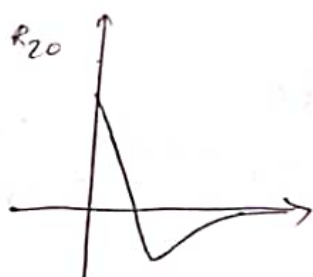
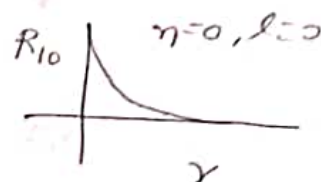
$$R_{nl} = \sqrt{\frac{(n-l-1)!}{2n[(n+1)!]^3}} \left(\frac{2Z}{na_\mu}\right)^{3/2} e^{-Zr/na_\mu} \left(\frac{2Zr}{na_\mu}\right)^l \int_0^{2Zr/na_\mu} \left(\frac{2Zr}{na_\mu}\right)^{n-l-1} e^{-x} dx$$

Associated Laguerre
-r is polynomial

$$f_n^{(k)}(x) = (-1)^k \frac{d^k}{dx^k} f_{n+1}(x)$$

$$\Rightarrow R_{10} = 2 \left(\frac{Z}{a_\mu}\right)^{3/2} e^{-Zr/a_\mu}$$

$$R_{20} = 2 \left(\frac{Z}{a_\mu}\right)^{3/2} \left(1 - \frac{Zr}{2a_\mu}\right) e^{-Zr/2a_\mu}$$



$$n=2, l=0$$

Prob. of finding the e^- b/w
 r & $r+dr$.

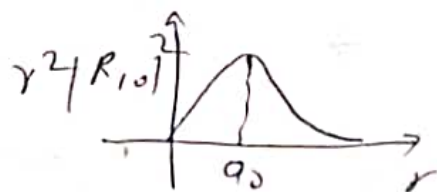
Prob. density

$$|\Psi_{nlm}|^2 = |R_{nl}|^2$$

$$|Y_l^m(\theta, \phi)|^2$$

$$P_{\text{prob}} = \int_0^{4\pi} \int_0^{2\pi} \int_0^\infty |R_{nl}|^2 |Y_l^m(\theta, \phi)|^2 r^2 \sin\theta dr d\theta d\phi$$

$$= r^2 |R_{nl}|^2 dr$$



$$\Rightarrow \langle r^2 \rangle_{nlm} = \int_0^\infty r^{2+2} |R_{nl}(r)|^2 dr$$

$$\langle r \rangle_{nl} = \frac{a_0}{2Z} [3n^2 - l(l+1)]$$

$$\langle r^{-1} \rangle_{nl} = \frac{Z}{a_0 n^2}$$

$$\langle r^{-1} \rangle_{n\ell} = \frac{Z^2}{a_0^2 n^3 (\ell + 1/2)}$$

$$Y_{\ell}^m(\pi-\theta, \pi+\phi) = (-1)^{\ell} Y_{\ell}^m(\theta, \phi)$$

check the parity of the orbital.

$$\begin{aligned} r, \theta, \phi & \rightarrow r, \theta, \phi \\ \theta & \rightarrow \pi - \theta \\ \phi & \rightarrow 2\pi + \phi \end{aligned}$$

$$\ell = 0, 1, 2, 3, 4, 5$$

$$\downarrow \downarrow \downarrow \downarrow \downarrow \downarrow$$

$$0, 0, 0, 0, 0, 0$$

family

(spin) 0 $\Rightarrow H_0 = \frac{p^2}{2m} - \frac{Ze^2}{4\pi\epsilon_0 r}$ (non-relativistic e)

$$H = H_1 + H_2 + H_3$$

$$H_1 = \frac{p^4}{8m^3c^2}$$

Relativistic correction to the KE

$$H_2 = \frac{1}{r} \frac{\partial V(r)}{\partial r} \vec{L} \cdot \vec{S} \quad \left\{ \begin{array}{l} \text{Spin-orbit} \\ \text{coupling} \end{array} \right\}$$

$$H_3 = \frac{\pi\hbar^2}{2m^2c^2} \left(\frac{Ze^2}{4\pi\epsilon_0} \right) \delta(r)$$

Darwin term

$$v^2 = \frac{p^2}{m^2}$$

$$p_e = mv, \quad \langle p_e^2 \rangle = m^2 \langle v^2 \rangle$$

$$\frac{1}{2} m \langle v^2 \rangle = \frac{1}{2m} \langle p_e^2 \rangle$$

$$\langle T \rangle_{100} = \langle n=1, \ell=0, m=0 | \frac{p^2}{2m} | 100 \rangle$$

$$= \langle H \rangle_{100} = \langle V \rangle_{100}$$

$$\Rightarrow \frac{-Ze^2}{4\pi\epsilon_0} \langle \frac{1}{r} \rangle_{100} \rightarrow Z/a_0$$

$$\langle T \rangle_{100} = \left(\frac{Ze^2}{4\pi\epsilon_0} \right)^2 \frac{me}{2\hbar^2}$$

$$= \frac{1}{2} Z^2 \alpha^2 m c^2$$

fine structure const.

$$v_{rms} = \sqrt{\langle v^2 \rangle_{rms}}$$

$$v_{rms} = Z\alpha c$$

$$T = \sqrt{p^2 c^2 + m_e^2 c^4} - m_e c^2$$

$$= m_e c^2 \left[\sqrt{1 + \frac{p^2}{m_e^2 c^2}} - 1 \right]$$

$$= \frac{p^2}{2m_e} - \frac{p^4}{8m_e^3 c^2} + \dots$$

⇒ Time independent perturbation theory

$$H^1 = H_0 + \lambda H_1$$

$$\Delta E^{(1)} = \langle \psi_{lmn}^{(0)} | H_1 | \psi_{lmn}^{(0)} \rangle$$

Correction in the energy due to i.e. due to the relativistic effect

$$S_R = \frac{p^4}{8m_e^3 c^2}$$

$$\Delta E_R = \langle nlm | S_R | nlm \rangle$$

$$= \frac{1}{2m_e c^2} \langle nlm | H_0^2 - 2H_0 V + V^2 | nlm \rangle$$

$$H_0 | nlm \rangle = E_n | nlm \rangle$$

$$H_0^2 | nlm \rangle = E_n^2 | nlm \rangle$$

$$H_0 = \frac{p^2}{2m}$$

$$(H_0 - V)^2 = \left(\frac{p^2}{2m} \right)^2$$

$$\frac{1}{2m_e c^2} (H_0^2 + V^2 - 2H_0 V)$$

$$\Rightarrow \langle H_0 V \rangle_{nlm}$$

$$= \langle nlm | H_0 V | nlm \rangle$$

$$= -E_n \frac{Ze^2}{4\pi\epsilon_0} \left\langle \frac{1}{r} \right\rangle_{nlm}$$

$$\uparrow Z/a_0$$

$$= -E_n \times \frac{Z^2 e^2}{4\pi\epsilon_0}$$

$$\langle V^2 \rangle_{nlm} = \left(\frac{Ze^2}{4\pi\epsilon_0} \right)^2 \left\langle \frac{1}{r^2} \right\rangle_{nlm}$$

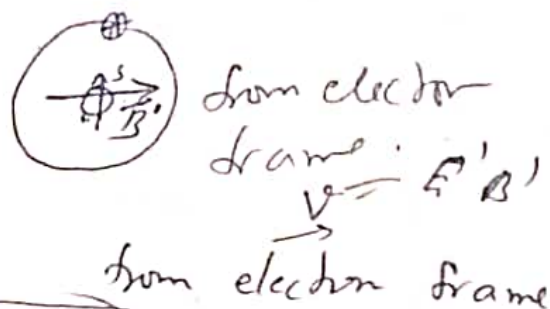
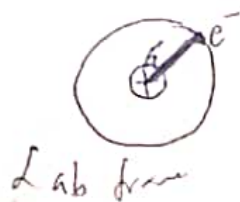
$$= \left(\frac{Z^2 e^2}{4\pi\epsilon_0} \right)^2 \times \frac{1}{n^2 a_0^2 (l+1/2)}$$

$$\left(\frac{L}{m} \right)$$

$$\Delta E_R = \frac{-1}{2m_e c^2} \left[E_n^2 - 4 E_n^2 + E_n^2 \frac{4n}{l+1/2} \right]$$

$$\Delta E_R = \frac{-E_n^2}{2m_e c^2} \left[\frac{4n}{l+1/2} - 3 \right]$$

⇒



$$\vec{B}' = \frac{1}{2} \frac{\vec{E} \times \vec{v}}{c^2}$$

$$\rightarrow H'_{SOI} = -\vec{\mu}_e \cdot \vec{B}' = +g_0 \mu_B \frac{\vec{S} \cdot \vec{B}}{\hbar}$$

$$H'_{SO} = \frac{e}{m_e} \vec{S} \cdot \vec{B}$$

$$= \frac{e}{m_e} \vec{S} \cdot \left(\frac{\vec{E} \times \vec{v}}{c^2} \right)$$

$$g_0 \frac{\hbar}{2m_e c^2} \vec{S} \cdot \vec{B}$$

$$\frac{e\hbar}{2m_e}$$

$$\vec{E}' = -\vec{\nabla} V$$

$$\vec{E} = -\hat{r} \frac{\partial V}{\partial r}$$

$$\hat{r} \times \vec{v} = \frac{1}{m_e} \hat{r} \times \vec{p}$$

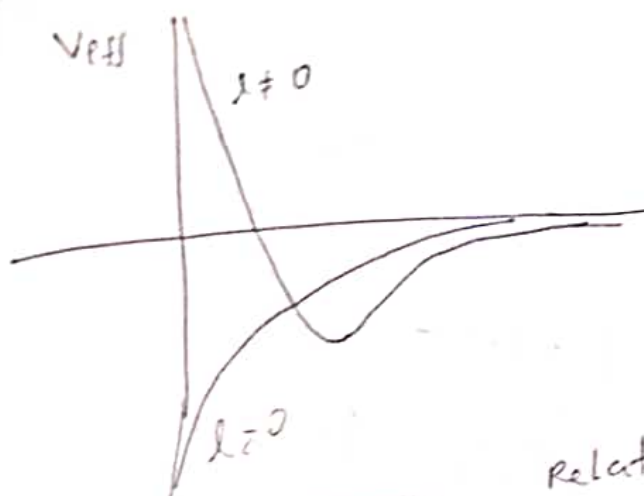
$$= \frac{1}{m_e} \vec{L}$$

$$H'_{SO} = -\frac{e}{2m_e^2 c^2} \left(\frac{1}{r} \frac{\partial V}{\partial r} \right) \vec{L} \cdot \vec{S}$$

$$U_{rel} = r R_{nl}(r)$$

$$\frac{\partial^2 U_{rel}}{\partial r^2} + \frac{2\mu}{\hbar^2} [E_{n\ell m} - V_{eff}(r)] U_{rel}(r) = 0$$

$$V_{eff} = \frac{-Ze^2}{4\pi\epsilon_0 r} + \frac{\ell(\ell+1)\hbar^2}{2\mu r^2}$$



At $r=0$, for

$\ell \neq 0$

$$U_{nl}(r=0) = 0$$

$$\Delta E_R = \frac{-E_n^2}{2m_e c^2} \left[\frac{4n}{\ell + 1/2} - 3 \right]$$

Relativistic

Correction

Correction due to Spin-orbit coupling —

$$H'_{SO} = -\vec{\mu}_e \cdot \vec{B} \left(\frac{\vec{E} \times \vec{v}}{2c^2} \right)$$

$$\vec{E} = -\vec{\nabla} U = -\hat{r} \frac{\partial U}{\partial r}$$

$$\mu_e = g \mu_B \frac{\vec{S}}{\hbar}$$

$$\vec{E} = -\vec{\nabla} U =$$

$$H'_{SO} = \frac{-g \mu_B}{\hbar} \vec{S} \cdot \left(\frac{\partial U}{\partial r} \right) \left[\frac{\vec{r} \times \vec{v}}{r} \right] \frac{1}{2c^2}$$

$$H'_{SO} = \frac{Ze^3}{8\pi\epsilon_0 m_e^2 c^2} \left\langle \frac{1}{r^3} \right\rangle \vec{L} \cdot \vec{S}$$

$$\Delta E_{SO} = \langle \Psi_{n\ell m} | H'_{SO} | \Psi_{n\ell m} \rangle$$

$$\Delta E_{SO} = \frac{Ze^3}{8\pi\epsilon_0 m_e^2 c^2} \left\langle \frac{1}{r^3} \right\rangle \langle \vec{L} \cdot \vec{S} \rangle$$

Increasing optics (PH305)

$$\vec{J} = \vec{L} + \vec{S} \Rightarrow J^2 = L^2 + S^2 + 2\vec{L} \cdot \vec{S}$$

$$\hat{L} \cdot \hat{S} = \frac{1}{2} [J^2 - L^2 - S^2]$$

$$\langle \hat{L} \cdot \hat{S} \rangle$$

$$\parallel$$

$$\langle n, l, m_l, j, m_j | \hat{L} \cdot \hat{S} | n, l, m_l, j, m_j \rangle$$

$$|S, S_L| \quad |S, S_L|$$

$$S = \frac{1}{2}$$

$$\frac{1}{2} \quad \frac{1}{2}$$

$$0 \leq 1$$

$$S |j, m_j\rangle = m_j \hbar |j, m_j\rangle$$

$$S = \frac{1}{2}$$

$$\frac{1}{2} [J^2 - L^2 - S^2] |n, l, m_l, j, m_j\rangle$$

$$\Rightarrow \frac{\hbar^2}{2} [j(j+1) - l(l+1) - s(s+1)]$$

$$\frac{1}{2} [s^2 j(j+1) |n, m_j, j, m_j\rangle - \hbar^2 l(l+1)]$$

$$= \frac{\hbar^2}{2} [j(j+1) - l(l+1) - s(s+1)]$$

$$\Delta E_{SO}^{(1)} = \langle \frac{1}{r^3} \rangle \frac{\hbar^2}{2} \times \frac{Ze^3}{8\pi\epsilon_0 m_e^2 c^2} [j(j+1) - l(l+1) - s(s+1)]$$

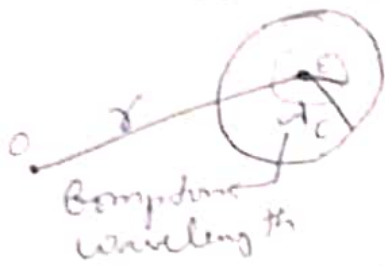
$$\langle \frac{1}{r^3} \rangle = \frac{Z^3}{n^3 a_0^3 l(l+\frac{1}{2})(l+1)}$$

$$\Delta E_{SO}^{(1)} = \frac{E_n^2}{m_e c^2} \frac{n [j(j+1) - l(l+1) - \frac{3}{4}]}{l(l+1)(l+\frac{1}{2})}$$

\Rightarrow Darwin Term \Leftarrow

$$H'_{\text{Darwin}} = \frac{1}{8} \left(\frac{\hbar}{m_e c} \right)^2 \nabla^2 V$$

electron is not of fixed state, it is fluctuating



$$\lambda_c = \frac{h}{m_e c}$$

$$V(r+\epsilon) \approx V(r) + \epsilon_i \cdot \nabla U(r) + \frac{1}{2} \sum_{i,j} \epsilon_i \epsilon_j \partial_i \partial_j V(r)$$

$$\overline{V(r+\epsilon)} = V(r) + \left(\frac{1}{2} \sum_i \epsilon_i^2 \partial_i^2 V(r) \right)$$

(ij) only survives as whole

$$= V(r) + \frac{1}{3} \overline{\epsilon^2} \times \frac{1}{2} \nabla^2 V(r)$$

$$\Delta U = \frac{1}{6} \lambda_c^2 \nabla^2 U(r)$$

$$H'_{\text{Darwin}} = \frac{1}{8} \left(\frac{h^2}{m_e c} \right)^2 \nabla^2 \left(\frac{Z e^2}{4 \pi \epsilon_0 r} \right)$$

$$H'_{\text{Darwin}} = \frac{\pi \hbar^2}{2 m_e^2 c^2} \times \frac{Z e^2}{4 \pi \epsilon_0} \delta^3(r)$$

$$l=0$$

contributes only

because $\int \delta(r) = 1$ or at $r=0$
 $\int \delta(r) = 0$ for $r \neq 0$

$$\Delta E_{\text{Darwin}}^{(1)} = \frac{\pi \hbar^2}{2 m_e^2 c^2} \frac{Z e^2}{4 \pi \epsilon_0} \langle \psi_{n00} | \delta^3(r) | \psi_{n00} \rangle$$

$$\psi(0) = \frac{1}{(4\pi)^{1/4}} 2 \left(\frac{Z}{a_0} \right)^{3/2} \text{ for } l=0$$

$$\Delta E_{\text{Darwin}}^{(1)} = \frac{2\pi}{m_e c^2} E_n^2$$

Linearizing optics (PH305)

$$\left[-\frac{\hbar^2}{2\mu} \nabla^2 + V(r) \right] \Psi(r, \theta, \phi) = E \Psi(r, \theta, \phi)$$

$$V \Rightarrow 0$$

$$\Psi(r, \theta, \phi) = R(r) Y_{\ell}^m(\theta, \phi)$$

$$Y_{\ell}^m(\theta, \phi) = \Theta(\theta) \Phi(\phi)$$

$$\Theta_{\ell}^m(\theta) = \sqrt{\frac{(2\ell+1)(\ell-|m|)!}{2(\ell+|m|)!}} P_{\ell}^m(\cos\theta)$$

$$\Phi_m(\phi) = \frac{1}{\sqrt{2\pi}} e^{im\phi}$$

$$Y_{\ell}^m = | \ell m \rangle$$

$$Y_{\ell}^m = (-1)^{(m+|m|)/2} \Theta_{\ell}^m(\theta) \Phi_m(\phi)$$

$$L^2 | \ell m \rangle = \ell(\ell+1) \hbar^2 | \ell m \rangle$$

$$L_z | \ell m \rangle = m \hbar | \ell m \rangle$$

$$L^2$$

$$\langle r \rangle = \langle n \ell m | r | n \ell m \rangle$$

$$= \oint_V \langle n \ell m | r \rangle \langle r | n \ell m \rangle$$

$$dV = dr \cdot r \sin\theta d\theta d\phi$$

$$= \int \Psi_{n\ell m}^*(r) \cdot r \Psi_{n\ell m} dV$$

$$= \int_0^{2\pi} \int_0^{\pi} \int_0^{\infty} | \Psi_{n\ell m}(r) |^2 \cdot r \sin\theta d\theta dr d\phi \quad \hat{n} | E \rangle = E | n \rangle$$

$$\langle E | \hat{x} \neq E^*(n)$$

$$\langle r \rangle = \int_0^{\infty} | R_{n\ell}(r) |^2 r dr \Rightarrow \langle n \ell m | r | n \ell m \rangle$$

$$= \frac{\langle n \ell m | r \rangle \langle r | n \ell m \rangle}{\int_V \Psi^* \Psi dV}$$

(1) - Rel. correction

$$\Delta E_R^{(1)} = \frac{h^{(0)}}{2m_e c^2} \left[\frac{4m}{c \hbar} - 1 \right]$$

(2) - spin-orbit coupling
(coupling b/w the electron spin & its orbital motion)

$$\Delta E_{SO}^{(1)} = \frac{E_n^{(0)2}}{m_e c^2} \frac{j(j+1) - l(l+1) - 3/4}{l(l+1/2)(l+1)}$$

$$|l - 1/2| \leq j \leq l + 1/2$$

$$\boxed{|l - s| \leq j \leq l + s}$$

$H' \rightarrow$
Darwin
 \downarrow

Due to finite size correction of an electron which have radius of the order of Compton wavelength

$$H'_{\text{Darwin}} = \frac{\pi \hbar^2}{2m_e^2 c^2} \frac{Ze^2}{4\pi \epsilon_0} \delta^3(r)$$

$\Delta E_{\text{Darwin}}^{(1)}$ will have contribution only for $l=0$, because $l \neq 0$ radial orbital vanishes at $r=0$.

$$\Delta E_{\text{Darwin}}^{(1)} = \frac{\pi \hbar^2}{2m_e^2 c^2} \frac{Ze^2}{4\pi \epsilon_0} |\psi_{n00}(0)|^2$$

-13.6

$$\Delta_n^{(1)} = \langle n l j m_j | \underbrace{H'_{rel} + H'_{SO}}_{-13.6} | n l j m_j \rangle$$

$$= \frac{E_n^{(0)2}}{2m_e c^2} \left[3 - \frac{4n}{l+1/2} + 2n \frac{j(j+1) - l(l+1) - 3/4}{l(l+1/2)(l+1)} \right]$$

$$= \frac{E_n^{(0)^2}{2m_e c^2} \left\{ 3 + 2n \left[\frac{j(j+1) - 3l(l+1) - \frac{3}{4}}{l(l+\frac{1}{2})(j+1)} \right] \right\}$$

$$j = l + \frac{1}{2}$$

$$j = l + \frac{1}{2} \rightarrow$$

$$j = l - \frac{1}{2}$$

$$f(j, l) = \frac{(l - \frac{1}{2})(l + \frac{1}{2}) - 3l(l+1) - \frac{3}{4}}{l(l+1)(l + \frac{1}{2})}$$

$$l = j - \frac{1}{2}$$

$$l^2 - \frac{1}{4} - 3l^2 - 3l - \frac{3}{4} - 1$$

$$\frac{l^2 - 3l^2 - 3l - \frac{1}{4}}{l(l+1)(l + \frac{1}{2})}$$

$$l = j - \frac{1}{2}$$

$$\frac{j(j+1) - 3(j - \frac{1}{2})(j + \frac{1}{2}) - \frac{3}{4}}{(j - \frac{1}{2})j(j + \frac{1}{2})}$$

$$\frac{j^2 + j - 3j^2 + \frac{3}{4} - \frac{3}{4}}{(j - \frac{1}{2})j(j + \frac{1}{2})}$$

$$-2j^2 + j$$

$$\frac{-2j[j - \frac{1}{2}]}{j(j - \frac{1}{2})(j + \frac{1}{2})}$$

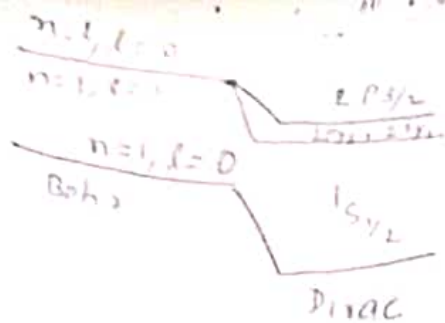
$$= \left(\frac{-2}{j + \frac{1}{2}} \right)$$

$$f(j, l) \Big|_{l=j+\frac{1}{2}} = \frac{-2}{j+\frac{1}{2}}$$

$$E^{(1)} = \frac{2n E_n^{(0)^2}{m_e c^2} \left[\text{Darwin, only for } l=0 \right] + \frac{E_n^{(0)^2}{m_e c^2} \times \left(2n + \frac{3}{2} \right)$$

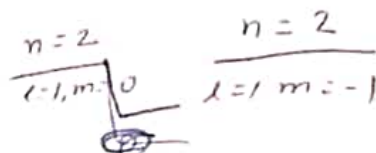
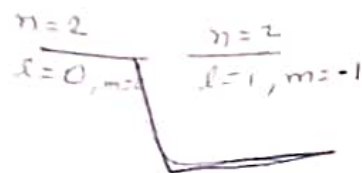
$$E^{(1)} = E_{rel}^{(1)} + E_{so}^{(1)}$$

$$= \frac{E_n^2}{2m_e c^2} \left\{ 3 + (n) \cdot \frac{-2}{j + \frac{1}{2}} \right\} = \frac{E_n^2}{2m_e c^2} \left\{ 3 - \frac{4n}{j + \frac{1}{2}} \right\}, l \neq 0$$



$$n_{l,j}$$

$$l = 1 + \frac{1}{2}$$



$$n=1$$

$$l=0$$

Single electron Atom

$$H = H_{\text{Schrodinger}} + H'_{\text{Dirac}}$$

Correction due to finite size of an relativistic electron

$$H'_{\text{Rel}} + H'_{\text{SO}} + H'_{\text{Darwin}} + H'_{\text{Lamb}}$$

\downarrow
 $\frac{\hat{p}^4}{8m_e^3c^2}$ (Relativistic correction)
 \downarrow Coupling b/w orbital motion & spin of an electron.
 \downarrow [Quantised nature of electromagnetic field.] Quant. electrodynamics

$$\textcircled{1} \Delta E_{\text{Rel}}^{(1)} = -E_n Z^2 \alpha^2 \left[\frac{3}{4n} - \frac{1}{l+1/2} \right]$$

$$\alpha = \text{fine structure Const.}$$

$$= \frac{1}{137}$$

$$\textcircled{2} \Delta E_{\text{SO}}^{(1)} = -E_n Z^2 \alpha^2 \frac{1}{n l(l+1/2)(l+1)}, l \neq 0 = -\frac{E_n (Z\alpha)^2}{n l(l+1/2)(l+1)(-1)}$$

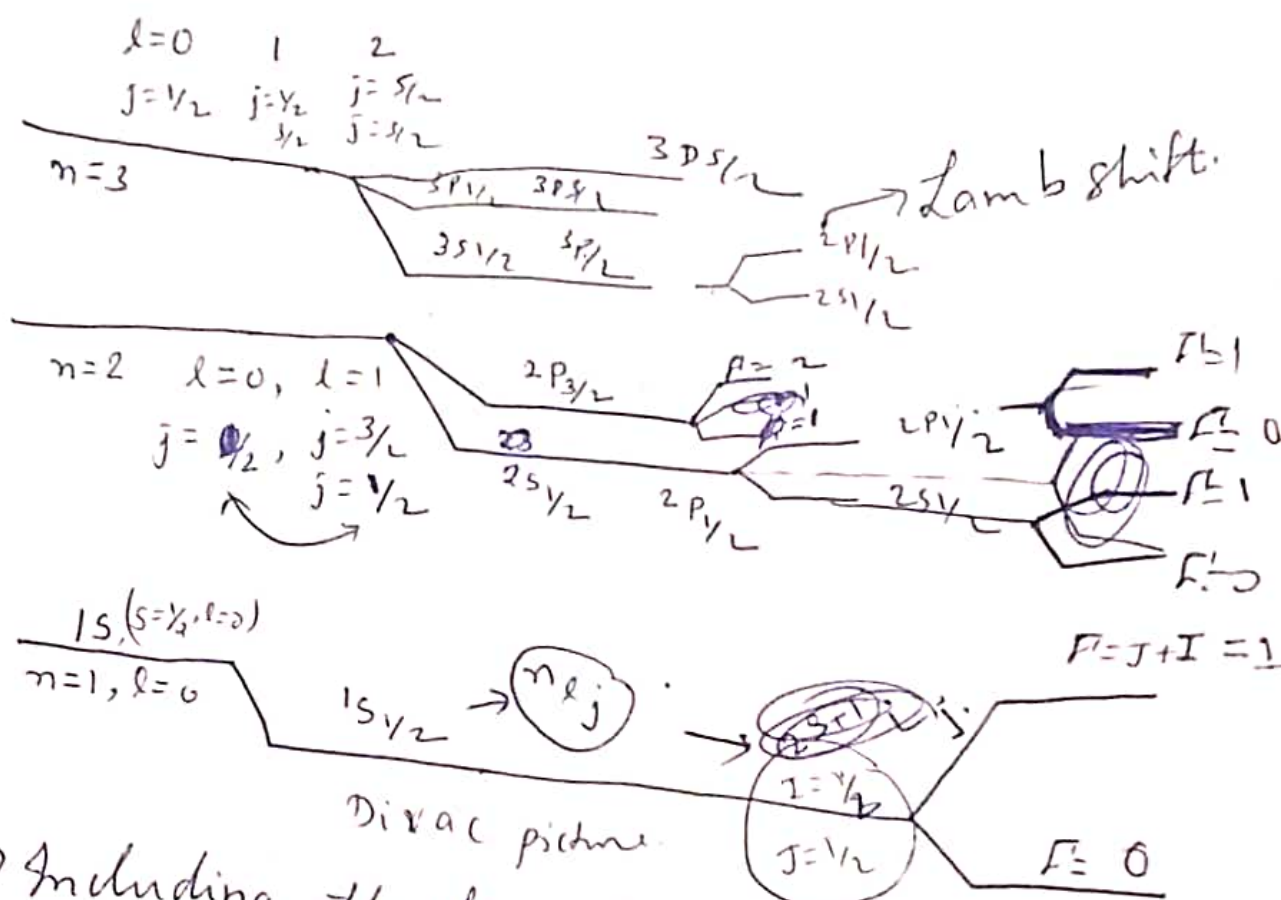
$$\Delta E_{\text{SO}}^{(1)} = 0 \quad l=0$$

$$\Delta E_{\text{Dirac}}^{(1)} = -E_n \frac{(Z\alpha)^2}{n}, \quad l=0$$

$$\Delta E_{\text{Rel}}^{(1)} + \Delta E_{\text{SD}}^{(1)}$$

$$= -E_n \left[1 + \frac{(Z\alpha)^2}{n} \left(\frac{1}{j+1/2} + \frac{3}{4n} \right) \right]$$

only depends on j , not on l .



Selection rules
 $\Delta j = 0, \pm 1$
 $\Delta l = \pm 1$

\Rightarrow Including the spin of nucleus, we will have new angular momentum-like operator

$$\hat{F} = \hat{J} + \hat{I}$$

$$H'_{HF} = -\vec{\mu}_N \cdot \vec{B}_{\text{Total}} \rightarrow \text{spin of the nucleus.}$$