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A search for sparticles in zero lepton final states

2

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ABSTRACT

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A search for sparticles in zero lepton final states

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Russell W. Smith

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16 center, but the abstract itself should be written as a regular paragraph on the page,

17 and it should not have indentation. Just replace this text.

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Acknowledgements

Chapter 1

Introduction

Particle physics is a remarkably successful field of scientific inquiry. The ability to precisely predict the properties of a exceedingly wide range of physical phenomena, such as the description of the cosmic microwave background [1, 2], the understanding of the anomalous magnetic dipole moment of the electron [3, 4], and the measurement of the number of weakly-interacting neutrino flavors [5] is truly amazing.

The theory that has allowed this range of predictions is the *Standard Model* of particle physics (SM). The Standard Model combines the electroweak theory of Glashow, Weinberg, and Salam [6–8] with the theory of the strong interactions, as first envisioned by Gell-Mann and Zweig [9, 10]. This quantum field theory (QFT) contains a tiny number of particles, whose interactions describe phenomena up to at least the TeV scale. These particles are manifestations of the fields of the Standard Model, after application of the Higgs Mechanism. The particle content of the SM consists only of the six quarks, the six leptons, the four gauge bosons, and the scalar Higgs boson.

Despite its impressive range of described phenomena, the Standard Model has some theoretical and experimental deficiencies. The SM contains 26 free parameters¹. It would be more theoretically pleasing to understand these free parameters in terms of a more fundamental theory. The major theoretical concern of the Standard Model, as it pertains to this thesis, is the *hierachy problem*[11–15]. The light mass

¹This is the Standard Model corrected to include neutrino masses. These parameters are the fermion masses (6 leptons, 6 quarks), CKM and PMNS mixing angles (8 angles, 2 CP-violating phases), W/Z/Higgs masses (3), the Higgs field expectation value, and the couplings of the strong, weak, and electromagnetic forces (3 α_{force}).

82 of the Higgs boson (125 GeV) should be quadratically dependent on the scale of UV
 83 physics, due to the quantum corrections from high-energy physics processes. The
 84 most perplexing experimental issue is the existence of *dark matter*, as demonstrated
 85 by galactic rotation curves [16–22]. This data has shown that there exists additional
 86 matter which has not yet been seen interacting with the particles of the Standard
 87 Model. There is no particle in the SM which can act as a candidate for dark matter.

88 Both of these major issues, as well as numerous others, can be solved by the
 89 introduction of *supersymmetry* (SUSY) [15, 23–33]. In supersymmetric theories, each
 90 SM particles has a so-called *superpartner*, or sparticle partner, differing from given SM
 91 particle by 1/2 in spin. These theories solve the hierarchy problem, since the quantum
 92 corrections induced from the superpartners exactly cancel those induced by the SM
 93 particles. In addition, these theories are usually constructed assuming R -parity,
 94 which can be thought of as the “charge” of supersymmetry, with SM particles having
 95 $R = 1$ and sparticles having $R = -1$. In collider experiments, since the incoming
 96 SM particles have total $R = 1$, the resulting sparticles are produced in pairs. This
 97 produces a rich phenomenology, which is characterized by significant hadronic activity
 98 and large missing transverse energy (E_T^{miss}), which provide significant discrimination
 99 against SM backgrounds [34].

100 Despite the power of searches for supersymmetry where E_T^{miss} is a primary dis-
 101 criminating variable, there has been significant interest in the use of other variables
 102 to discriminate against SM backgrounds. These include searches employing variables
 103 such as αT , $M_{T,2}$, and the razor variables (M_R, R^2) [35–45]. In this thesis, we will
 104 present the first search for supersymmetry using the novel Recursive Jigsaw Recon-
 105 struction (RJR) technique. RJR can be considered the conceptual successor of the
 106 razor variables. We impose a particular final state “decay tree” on an events, which
 107 roughly corresponds to a simplified Feynmann diagram in decays containing weakly-
 108 interacting particles. We account for the missing degrees of freedom associated to

109 the weakly-interacting particles by a series of simplifying assumptions, which allow
110 us to calculate our variables of interest at each step in the decay tree. This allows an
111 unprecedented understanding of the internal structure of the decay and the ability to
112 construct additional variables to reject Standard Model backgrounds.

113 This thesis details a search for the superpartners of the gluon and quarks, the
114 gluino and squarks, in final states with zero leptons, with 13.3 fb^{-1} of data using the
115 ATLAS detector. We organize the thesis as follows. The theoretical foundations of
116 the Standard Model and supersymmetry are described in Chapters 2 and 3. The
117 Large Hadron Collider and the ATLAS detector are presented in Chapters 4 and 5.
118 Chapter 5 provides a detailed description of Recursive Jigsaw Reconstruction and a
119 description of the variables used for the particular search presented in this thesis.
120 Chapter 6 presents the details of the analysis, including details of the dataset, object
121 reconstruction, and selections used. In Chapter 7, the final results are presented;
122 since there is no evidence of a supersymmetric signal in the analysis, we present the
123 final exclusion curves in simplified supersymmetric models.

124

Chapter 2

125

The Standard Model

126 Here you can write some introductory remarks about your chapter. I like to give each
 127 sentence its own line.

128 When you need a new paragraph, just skip an extra line.

129 **2.1 Quantum Field Theory**

130

131 In this section, we provide a brief overview of the necessary concepts from Quan-
 132 tum Field Theory (QFT).

133 In modern physics, the laws of nature are described by the “action” S , with the
 134 imposition of the principle of minimum action. The action is the integral over the
 135 spacetime coordinates of the “Lagrangian density” \mathcal{L} , or Lagrangian for short. The
 136 Lagrangian is a function of “fields”; general fields will be called $\phi(x^\mu)$, where the
 137 indices μ run over the space-time coordinates. We can then write the action S as

$$S = \int d^4x \mathcal{L}[\phi_i(x^\mu), \partial_\mu \phi_i(x^\mu)] \quad (2.1)$$

138 where we have an additional summation over i (of the different fields). Generally,
 139 we impose the following constraints on the Lagrangian :

- 140 1. Translational invariance - The Lagrangian is only a function of the fields ϕ and
 141 their derivatives $\partial_\mu \phi$
- 142 2. Locality - The Lagrangian is only a function of one point x_μ in spacetime.

cite Yuval's
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- 143 3. Reality condition - The Lagrangian is real to conserve probability.
- 144 4. Lorentz invariance - The Lagrangian is invariant under the Poincaré group of
145 spacetime.
- 146 5. Analyticity - The Lagrangian is an analytical function of the fields; this is to
147 allow the use of perturbation theory.
- 148 6. Invariance and Naturalness - The Lagrangian is invariant under some internal
149 symmetry groups; in fact, the Lagrangian will have *all* terms allowed by the
150 imposed symmetry groups.
151 7. Renormalizability - The Lagrangian will be renormalizable - in practice, this
152 means there will not be terms with more than power 4 in the fields.

153 The key item from the point of view of this thesis is that of “Invariance and
154 Natural”. We impose a set of “symmetries” and then our Lagrangian is the most
155 general which is allowed by those symmetries.

156 2.2 Symmetries

157 Symmetries can be seen as the fundamental guiding concept of modern physics. Sym-
158 metries are described by “groups”. . To illustrate the importance of symmetries and
159 their mathematical description, groups, we start here with two of the simplest and
160 most useful examples : \mathbb{Z}_2 and $U(1)$.

161 \mathbb{Z}_2 symmetry

162 \mathbb{Z}_2 symmetry is the simplest example of a “discrete” symmetry. Consider the most
163 general Lagrangian of a single real scalar field $\phi(x_\mu)$

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$$\mathcal{L}_\phi = \frac{1}{2}\partial_\mu\phi\partial^\mu\phi - \frac{m^2}{2}\phi^2 - \frac{\mu}{2\sqrt{2}}\phi^3 - \lambda\phi^4 \quad (2.2)$$

Now we *impose* the symmetry

$$\mathcal{L}(\phi) = \mathcal{L}(-\phi) \quad (2.3)$$

164 This has the effect of restricting the allowed terms of the Lagrangian. In particular,
 165 we can see the term $\phi^3 \rightarrow -\phi^3$ under the symmetry transformation, and thus must
 166 be disallowed by this symmetry. This means under the imposition of this particular
 167 symmetry, our Lagrangian should be rewritten as

$$\mathcal{L}_\phi = \frac{1}{2}\partial_\mu\phi\partial^\mu\phi - \frac{m^2}{2}\phi^2 - \lambda\phi^4 \quad (2.4)$$

168 The effect of this symmetry is that the total number of ϕ particles can only change
 169 by even numbers, since the only interaction term $\lambda\phi^4$ is an even power of the field.
 170 This symmetry is often imposed in supersymmetric theories, as we will see in Chapter
 171 3.

172 **$U(1)$ symmetry**

173 $U(1)$ is the simplest example of a continuous (or *Lie*) group. Now consider a theory
 174 with a single complex scalar field $\phi = \text{Re } \phi + i \text{Im } \phi$

$$\mathcal{L}_\phi = \delta_{i,j} \frac{1}{2} \partial_\mu \phi_i \partial^\mu \phi_j - \frac{m^2}{2} \phi_i \phi_j - \frac{\mu}{2\sqrt{2}} \phi_i \phi_j \phi_k - \lambda \phi_i \phi_j \phi_k \phi_l \quad (2.5)$$

175 where $i, j, k, l = \text{Re}, \text{Im}$. In this case, we impose the following $U(1)$ symmetry
 176 : $\phi \rightarrow e^{i\theta} \phi, \phi^* \rightarrow e^{-i\theta} \phi^*$. We see immediately that this again disallows the third-order
 177 terms, and we can write a theory of a complex scalar field with $U(1)$ symmetry as

$$\mathcal{L}_\phi = \partial_\mu \phi \partial^\mu \phi^* - \frac{m^2}{2} \phi \phi^* - \lambda (\phi \phi^*)^2 \quad (2.6)$$

178 2.3 Local symmetries

179 The two examples considered above are “global” symmetries in the sense that the
 180 symmetry transformation does not depends on the spacetime coordinate x_μ . We know
 181 look at local symmetries; in this case, for example with a local $U(1)$ symmetry, the
 182 transformation has the form $\phi(x_\mu) \rightarrow e^{i\theta(x_\mu)}\phi(x_\mu)$. These symmetries are also known
 183 as “gauge” symmetries; all symmetries of the Standard Model are gauge symmetries.

There are wide-ranging consequences to the imposition of local symmetries. To begin, we note that the derivative terms of the Lagrangian 2.2 are *not* invariant under a local symmetry transformation

$$\partial_\mu \phi(x_\mu) \rightarrow \partial_\mu (e^{i\theta(x_\mu)} \phi(x_\mu)) = (1 + i\theta(x_\mu)) e^{i\theta(x_\mu)} \phi(x_\mu) \quad (2.7)$$

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RIGHT 185

This leads us to note that the kinetic terms of the Lagrangian are also not invariant
 186 under a gauge symmetry. This would lead to a model with no dynamics, which is
 187 clearly unsatisfactory.

188 Let us take inspiration from the case of global symmetries. We need to define a
 189 so-called “covariant” derivative D^μ such that

$$\begin{aligned} D^\mu \phi &\rightarrow e^{iq\theta(x^\mu)} D^\mu \phi \\ D^\mu \phi^* &\rightarrow e^{-iq\theta(x^\mu)} D^\mu \phi \end{aligned} \quad (2.8)$$

190 Since ϕ and ϕ^* transforms with the opposite phase, this will lead the invariance
 191 of the Lagrangian under our local gauge transformation. This D^μ is of the following
 192 form

$$D^\mu = \partial_\mu - igqA^\mu \quad (2.9)$$

193 where A^μ is a vector field we introduce with the transformation law

$$A^\mu \rightarrow A^\mu - \frac{1}{g} \partial_\mu \theta \quad (2.10)$$

194 and g is the coupling constant associated to vector field. This vector field A^μ is
 195 also known as a “gauge” field.

196 Since we need to add all allowed terms to the Lagrangian, we define

$$F^{\mu\nu} = A^\mu A^\nu - A^\nu A^\mu \quad (2.11)$$

197 and then we must also add the kinetic term :

$$\mathcal{L}_{\text{gauge}} = -\frac{1}{4} F^{\mu\nu} F_{\mu\nu} \quad (2.12)$$

198 The most general renormalizable Lagrangian with fermion and scalar fields can
 199 be written in the following form

$$\mathcal{L} = \mathcal{L}_{kin} + \mathcal{L}_\phi + \mathcal{L}_\psi + \mathcal{L}_{Yukawa} \quad (2.13)$$

200 Symmetry breaking and the Higgs mechanism

201 Here we view some examples of symmetry breaking. We investigate breaking of a
 202 global $U(1)$ symmetry and a local $U(1)$ symmetry. The SM will break the electroweak
 203 symmetry $SU(2) \times U(1)$, and in Chapter 3 we will see how supersymmetry must also
 204 be broken.

205 There are two ideas of symmetry breaking

- 206 • Explicit symmetry breaking by a small parameter - in this case, we have a small
 207 parameter which breaks an “approximate” symmetry of our Lagrangian. An
 208 example would be the theory of the single scalar field [2.2](#), when $\mu \ll m^2$ and
 209 $\mu \ll \lambda$. In this case, we can often ignore the small term when considering
 210 low-energy processes.

211 • Spontaneous symmetry breaking (SSB) - spontaneous symmetry breaking oc-
 212 curs when the Lagrangian is symmetric with respect to a given symmetry trans-
 213 formation, but the ground state of the theory is *not* symmetric with respect to
 214 that transformation. This can have some fascinating consequences, as we will
 215 see in the following examples

216 Symmetry breaking a

217 U(1) global symmetry breaking

Consider the theory of a complex scalar field under the $U(1)$ symmetry, or the trans-
formation

$$\phi \rightarrow e^{i\theta} \phi \quad (2.14)$$

The Lagrangian for this theory is

$$\mathcal{L} = \partial^\mu \phi^\dagger \partial_\mu \phi + \frac{\mu^2}{2} \phi^\dagger \phi + \frac{\lambda}{4} (\phi^\dagger \phi)^2 \quad (2.15)$$

Let us write this theory in terms of two scalar fields, h and ξ : $\phi = (h + i\xi)/\sqrt{2}$.

The Lagrangian can then be written as

$$\mathcal{L} = \partial^\mu h \partial_\mu h + \partial^\mu \xi \partial_\mu \xi - \frac{\mu^2}{2} (h^2 + \xi^2) - \frac{\lambda}{4} (h^2 + \xi^2)^2 \quad (2.16)$$

First, note that the theory is only stable when $\lambda > 0$. To understand the effect of SSB, we now enforce that $\mu^2 < 0$, and define $v^2 = -\mu^2/\lambda$. We can then write the scalar potential of this theory as :

$$V(\phi) = \lambda(\phi^\dagger \phi - v^2/2)^2 \quad (2.17)$$

Minimizing this equation with respect to ϕ , we can see that the “vacuum expectation value” of the theory is

$$2 \langle \phi^\dagger \phi \rangle = \langle h^2 + \xi^2 \rangle = v^2 \quad (2.18)$$

218 We now reach the “breaking” point of this procedure. In the (h, ξ) plane, the
 219 minima form a circle of radius v . We are free to choose any of these minima to expand
 220 our Lagrangian around; the physics is not affected by this choice. For convenience,
 221 choose $\langle h \rangle = v, \langle \xi^2 \rangle = 0$.

Now, let us define $h' = h - v, \xi' = \xi$ with VEVs $\langle h' \rangle = 0, \langle \xi' \rangle = 0$. We can
 then write our spontaneously broken Lagrangian in the form

$$\mathcal{L} = \frac{1}{2} \partial_\mu h' \partial^\mu h' + \frac{1}{2} \partial_\mu \xi' \partial^\mu \xi' - \lambda v^2 h'^2 - \lambda v h' (h'^2 + \xi'^2) - \lambda (h'^2 + \xi'^2)^2 \quad (2.19)$$

222

223 **U(1) local symmetry breaking**

Add a picture of the potential

224 2.4 The Standard Model

225 Overview

226 The Standard Model is another name for the theory of the internal symmetry group
 227 $SU(3)_C \times SU(2)_L \times U(1)_Y$. This quantum field theory is the culmination of years of
 228 work in both theoretical and particle physics.

229

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230 Field Content

The SM field content is

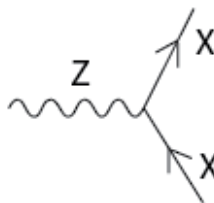
$$\begin{aligned} \text{Fermions } & Q_L(3, 2)_{+1/3}, U_R(3, 1)_{+4/3}, D_R(3, 1)_{-2/3}, L_L(1, 2)_{-1}, E_R(1, 1)_{-2} \\ \text{Scalar (Higgs)} & \phi(1, 2)_{+1/2} \end{aligned} \quad (2.20)$$

$$\text{Vector Fields } G^\mu(8, 1)_0, W^\mu(1, 3)_0, B^\mu(1, 1)_0$$

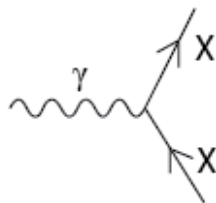
231 where the $(A, B)_Y$ notation represents the irreducible representation under $SU(3)$
 232 and $SU(2)$, with Y being the electroweak hypercharge. Each of these fields has an
 233 additional index, representing the three generation of fermions.

Figure 2.1: The interactions of the Standard Model

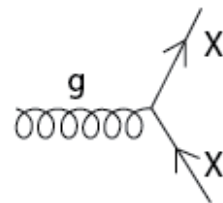
Standard Model Interactions (Forces Mediated by Gauge Bosons)



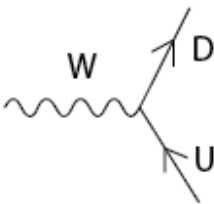
X is any fermion in the Standard Model.



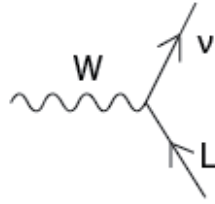
X is electrically charged.



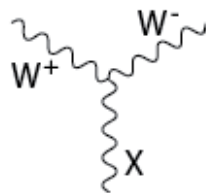
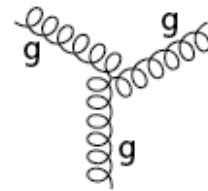
X is any quark.



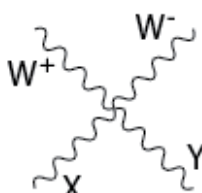
U is a up-type quark;
D is a down-type quark.



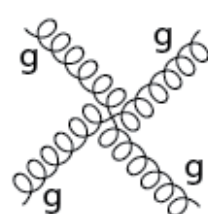
L is a lepton and ν is the corresponding neutrino.



X is a photon or Z-boson.



X and Y are any two electroweak bosons such that charge is conserved.



234 We observed that Q_L, U_R , and D_R are triplets under $SU(3)_C$; these are the *quark*
 235 fields. The “color” group, $SU(3)_C$ is mediated by the “gluon” field $G^\mu(8, 1)_0$, which
 236 has 8 degrees of freedom; we say there are 8 gluons. The fermion fields $L_L(1, 2)_{-1}$
 237 and $E_R(1, 1)_{-2}$ are singlets under $SU(3)_C$; we call them *leptons*.

238 Next, we note the “left-handed” (“right-handed”) fermion fields, denoted by L (R)
 239 subscript, The left-handed fields form doublets under $SU(2)_L$. These are mediated
 240 by the three degrees of freedom of the “W” fields $W^\mu(1, 3)_0$. These fields only act
 241 on the left-handed particles of the Standard Model. This is the reflection of the
 242 “chirality” of the Standard Model; the left-handed and right-handed particles are
 243 treated differently by the electroweak forces. The right-handed fields, U_R, D_R , and
 244 E_R , are singlets under $SU(2)_L$.

245 The $U(1)_Y$ symmetry is associated to the $B^\mu(1, 1)_0$ boson with one degree of
 246 freedom. We note that this field is associated with the charge Y of the other particles.

247 \mathcal{L}_{kin}

248 For each of the vector boson fields, we have the follow field strengths :

$$\begin{aligned} G_a^{\mu\nu} &= \partial^\mu G_a^\nu - \partial^\nu G_a^\mu - g_s f_{abc} G_b^\mu G_c^\nu \\ W_a^{\mu\nu} &= \partial^\mu W_a^\nu - \partial^\nu W_a^\mu - g \epsilon_{abc} W_b^\mu W_c^\nu \\ B^{\mu\nu} &= \partial^\mu B^\nu - \partial^\nu B^\mu \end{aligned} \quad (2.21)$$

249 where g and g_s are the electroweak and strong coupling constant.

We can write the covariant derivative for the Standard Model as

$$D^\mu = \partial^\mu + ig_s G_a^\mu L_a + ig W_a^\mu T_a + ig' Y B^\mu \quad (2.22)$$

250 where L_a and T_a are the generators of $SU(3)_C$ and $SU(2)_L$ respectively for each of
 251 the representations. Explicitly, for the $SU(3)_C$ triplets, $L_a = \frac{1}{2}\lambda_a$ and for the $SU(3)_C$
 252 singlets, $L_a = 0$. For $SU(2)_L$ doublets, $L_a = \frac{1}{2}\sigma_a$ and for $SU(2)_L$ singlets, $L_a = 0$.

The combination of these terms allows us to write the kinetic terms of the Lagrangian as

$$\begin{aligned} \mathcal{L}_{kin} = & G^{\mu\nu}G_{\mu\nu} + W^{\mu\nu}W_{\mu\nu} + B^{\mu\nu}B_{\mu\nu} \\ & + D^\mu Q_L D_\mu Q_L + D^\mu U_R D_\mu U_R + D^\mu D_R D_\mu D_R + D^\mu L_L D_\mu L_L + D^\mu E_R D_\mu E_R \end{aligned} \quad (2.23)$$

253 $\mathcal{L}\psi$

254 We cannot write down any mass terms for fermions in the Standard Model. Dirac
255 mass terms are forbidden since they are all assigned to “chiral” representations of the
256 gauge symmetry. Majorana mass terms are disallowed since there are no fields with
257 $Y \neq 0$.

258 $\mathcal{L}\text{Yuk}$

259 We write the Yukawa portion of the Fermions *do not* have mass terms in the Standard
260 Model. This

261 Let us now recall that local gauge invariance means that the vector fields in this
262 theory are *massless*, yet we know only the photon vector field is massless. In the next
263 section, we will see how masses are induced by electroweak symmetry breaking.

264 **2.5 Electroweak Symmetry breaking and the** 265 **Higgs Boson**

266 By using the asterisk to start a new section, I keep the section from appearing in the
267 table of contents. If you want your sections to be numbered and to appear in the
268 table of contents, remove the asterisk.

269 **2.6 Deficiencies of the Standard Model**

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271 table of contents. If you want your sections to be numbered and to appear in the
272 table of contents, remove the asterisk.

273

Chapter 3

274

Supersymmetry

275 Here you can write some introductory remarks about your chapter. I like to give each
276 sentence its own line.

277 When you need a new paragraph, just skip an extra line.

278 **3.1 Motivation**

279 **Only Additional allowed Lorentz invariant symmetry**

280 **Dark Matter**

281 **Cancellation of quadratic divergences in corrections to the**

282 **Higgs Mass**

283 **3.2 Supersymmetry**

284 **3.3 Additional particle content**

285 **3.4 Phenomenology**

286 **R parity Consequences for sq/gl decays**

The Large Hadron Collider

289 Here you can write some introductory remarks about your chapter. I like to give each
290 sentence its own line.

291 When you need a new paragraph, just skip an extra line.

292 **4.1 Magnets**

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294 table of contents. If you want your sections to be numbered and to appear in the
295 table of contents, remove the asterisk.

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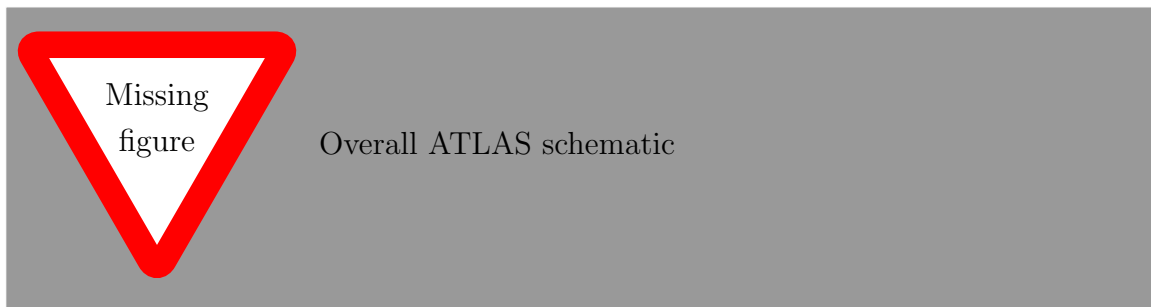
Chapter 5

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The ATLAS detector

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299 sentence its own line.

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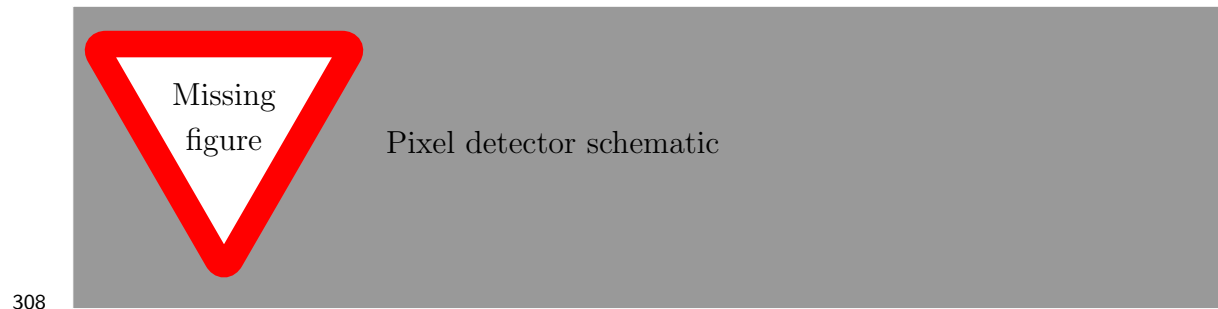
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302

303 **5.1 Inner Detector**

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305 table of contents. If you want your sections to be numbered and to appear in the
306 table of contents, remove the asterisk.

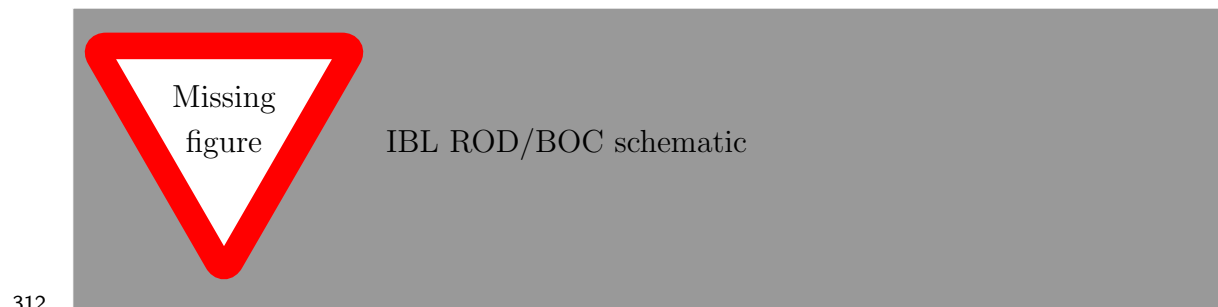
307 **Pixel Detector**



309

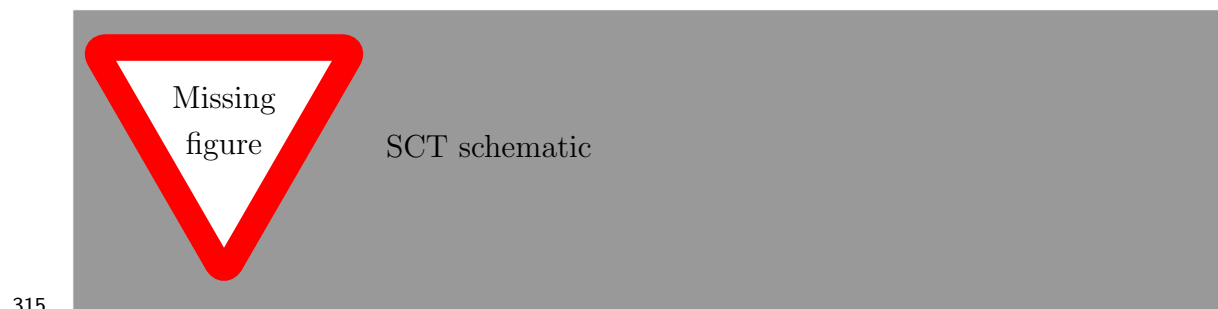
310 **Insertable B-Layer**

311 Qualification task, so add a bit more.



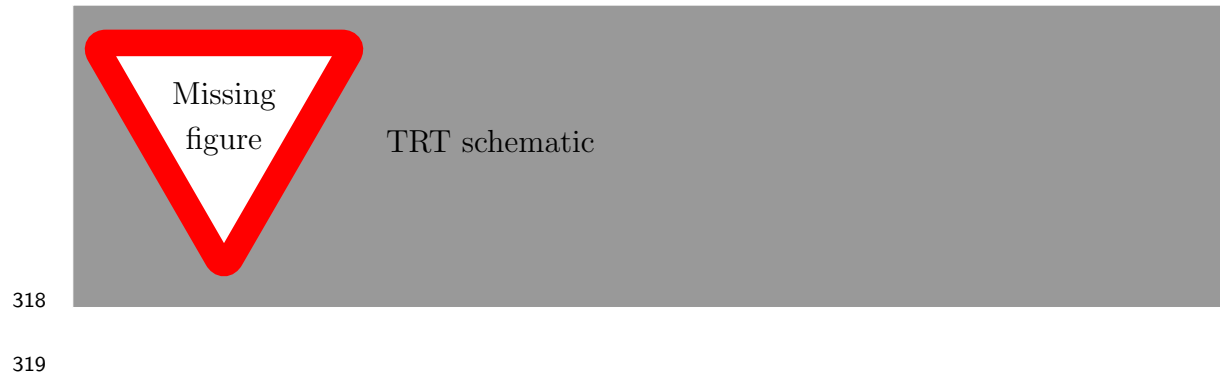
313

314 **Semiconductor Tracker**

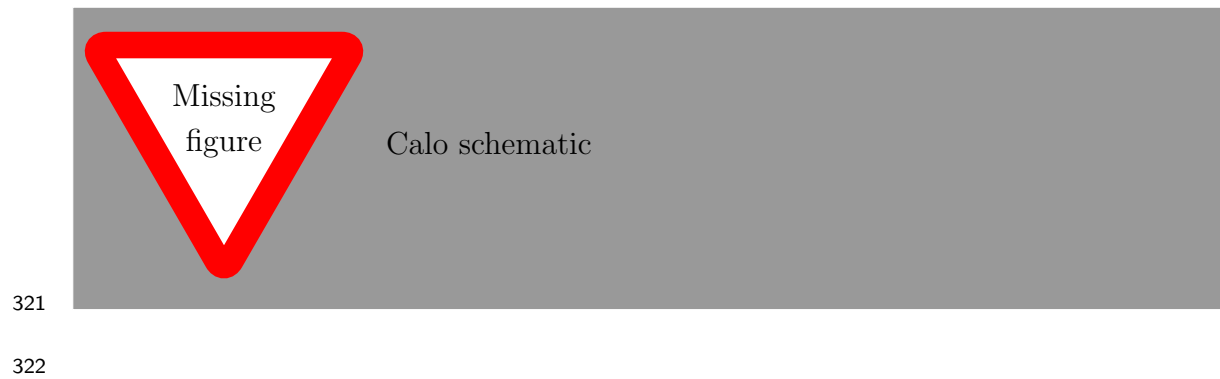


316

317 **Transition Radiation Tracker**



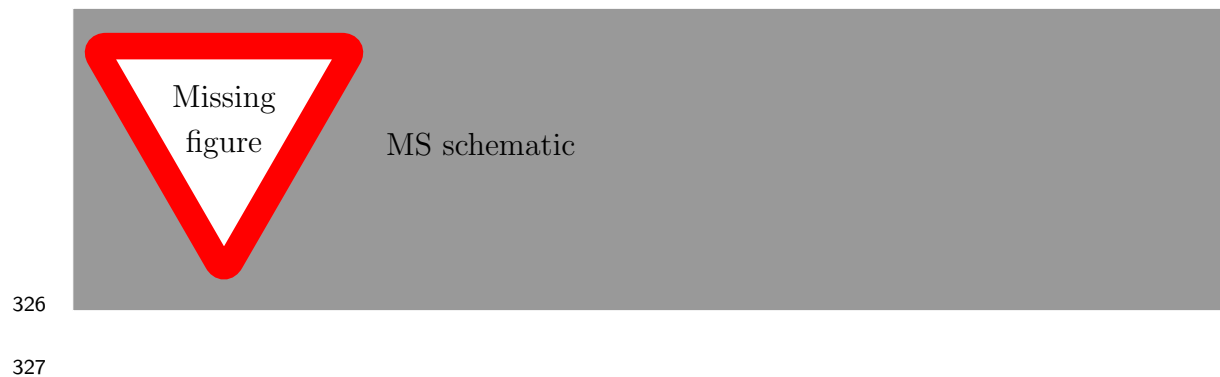
320 **5.2 Calorimeter**



323 **Electromagnetic Calorimeter**

324 **Hadronic Calorimeter**

325 **5.3 Muon Spectrometer**



The Recursive Jigsaw Technique

330 Here you can write some introductory remarks about your chapter. I like to give each
331 sentence its own line.

332 When you need a new paragraph, just skip an extra line.

333 **6.1 Razor variables**

334 By using the asterisk to start a new section, I keep the section from appearing in the
335 table of contents. If you want your sections to be numbered and to appear in the
336 table of contents, remove the asterisk.

337 **6.2 SuperRazor variables**

338 **6.3 The Recursive Jigsaw Technique**

339 **6.4 Variables used in the search for zero lepton**

340 **SUSY**

Title of Chapter 1

343

Chapter 8

344

Title of Chapter 1

345 Here you can write some introductory remarks about your chapter. I like to give each
346 sentence its own line.

347 When you need a new paragraph, just skip an extra line.

348 **8.1 Object reconstruction**

349 **Photons, Muons, and Electrons**

350 **Jets**

351 **Missing transverse momentum**

352 Probably longer, show some plots from the PUB note that we worked on

353 **8.2 Signal regions**

354 **Gluino signal regions**

355 **Squark signal regions**

356 **Compressed signal regions**

357 **8.3 Background estimation**

358 **Z $\nu\nu$**

359 **W $e\nu$**

360 **$t\bar{t}$ bar**

361

Chapter 9

362

Title of Chapter 1

363 Here you can write some introductory remarks about your chapter. I like to give each
364 sentence its own line.

365 When you need a new paragraph, just skip an extra line.

366 **9.1 Statistical Analysis**

367 maybe to be moved to an appendix

368 **9.2 Signal Region distributions**

369 **9.3 Pull Plots**

370 **9.4 Systematic Uncertainties**

371 **9.5 Exclusion plots**

372

Conclusion

373 Here you can write some introductory remarks about your chapter. I like to give each
374 sentence its own line.

375 When you need a new paragraph, just skip an extra line.

376 **9.6 New Section**

377 By using the asterisk to start a new section, I keep the section from appearing in the
378 table of contents. If you want your sections to be numbered and to appear in the
379 table of contents, remove the asterisk.

Bibliography

- [1] O. Perdereau, *Planck 2015 cosmological results*,
AIP Conf. Proc. **1743** (2016) p. 050014.
- [2] N. Aghanim et al.,
*Planck 2016 intermediate results. LI. Features in the cosmic microwave
background temperature power spectrum and shifts in cosmological parameters*
(2016), arXiv: [1608.02487 \[astro-ph.CO\]](#).
- [3] J. S. Schwinger,
On Quantum electrodynamics and the magnetic moment of the electron,
Phys. Rev. **73** (1948) p. 416.
- [4] S. Laporta and E. Remiddi,
The Analytical value of the electron ($g-2$) at order α^3 in QED,
Phys. Lett. **B379** (1996) p. 283, arXiv: [hep-ph/9602417 \[hep-ph\]](#).
- [5] S. Schael et al., *Precision electroweak measurements on the Z resonance*,
Phys. Rept. **427** (2006) p. 257, arXiv: [hep-ex/0509008 \[hep-ex\]](#).
- [6] S. L. Glashow, *Partial Symmetries of Weak Interactions*,
Nucl. Phys. **22** (1961) p. 579.
- [7] S. Weinberg, *A Model of Leptons*, Phys. Rev. Lett. **19** (1967) p. 1264.
- [8] A. Salam, *Weak and Electromagnetic Interactions*,
Conf. Proc. **C680519** (1968) p. 367.
- [9] M. Gell-Mann, *A Schematic Model of Baryons and Mesons*,
Phys. Lett. **8** (1964) p. 214.
- [10] G. Zweig, “An SU(3) model for strong interaction symmetry and its breaking.
Version 2,” *DEVELOPMENTS IN THE QUARK THEORY OF HADRONS*.
VOL. 1. 1964 - 1978, ed. by D. Lichtenberg and S. P. Rosen, 1964 p. 22,
URL: <http://inspirehep.net/record/4674/files/cern-th-412.pdf>.

- 406 [11] S. Weinberg, *Implications of Dynamical Symmetry Breaking*,
407 [Phys. Rev. **D13** \(1976\) p. 974.](#)
- 408 [12] S. Weinberg, *Implications of Dynamical Symmetry Breaking: An Addendum*,
409 [Phys. Rev. **D19** \(1979\) p. 1277.](#)
- 410 [13] E. Gildener, *Gauge Symmetry Hierarchies*, [Phys. Rev. **D14** \(1976\) p. 1667.](#)
- 411 [14] L. Susskind,
412 *Dynamics of Spontaneous Symmetry Breaking in the Weinberg-Salam Theory*,
413 [Phys. Rev. **D20** \(1979\) p. 2619.](#)
- 414 [15] S. P. Martin, “A Supersymmetry Primer,” 1997,
415 eprint: [arXiv:hep-ph/9709356](#).
- 416 [16] V. C. Rubin and W. K. Ford Jr., *Rotation of the Andromeda Nebula from a*
417 *Spectroscopic Survey of Emission Regions*, [Astrophys. J. **159** \(1970\) p. 379.](#)
- 418 [17] M. S. Roberts and R. N. Whitehurst,
419 “*The rotation curve and geometry of M31 at large galactocentric distances*,
420 [Astrophys. J. **201** \(1970\) p. 327.](#)
- 421 [18] V. C. Rubin, N. Thonnard, and W. K. Ford Jr.,
422 *Rotational properties of 21 SC galaxies with a large range of luminosities and*
423 *radii, from NGC 4605 /R = 4kpc/ to UGC 2885 /R = 122 kpc/*,
424 [Astrophys. J. **238** \(1980\) p. 471.](#)
- 425 [19] V. C. Rubin et al., *Rotation velocities of 16 SA galaxies and a comparison of*
426 *Sa, Sb, and SC rotation properties*, [Astrophys. J. **289** \(1985\) p. 81.](#)
- 427 [20] A. Bosma,
428 *21-cm line studies of spiral galaxies. 2. The distribution and kinematics of*
429 *neutral hydrogen in spiral galaxies of various morphological types.*,
430 [Astron. J. **86** \(1981\) p. 1825.](#)
- 431 [21] M. Persic, P. Salucci, and F. Stel, *The Universal rotation curve of spiral*
432 *galaxies: 1. The Dark matter connection*,
433 [Mon. Not. Roy. Astron. Soc. **281** \(1996\) p. 27,](#)
434 [arXiv: **astro-ph/9506004** \[**astro-ph**\].](#)
- 435 [22] M. Lisanti, “Lectures on Dark Matter Physics,” 2016,
436 eprint: [arXiv:1603.03797](#).
- 437 [23] H. Miyazawa, *Baryon Number Changing Currents*,
438 [Prog. Theor. Phys. **36** \(1966\) p. 1266.](#)

- [24] J.-L. Gervais and B. Sakita, *Generalizations of dual models*,
Nucl. Phys. **B34** (1971) p. 477.
- [25] J.-L. Gervais and B. Sakita,
Field Theory Interpretation of Supergauges in Dual Models,
Nucl. Phys. **B34** (1971) p. 632.
- [26] Yu. A. Golfand and E. P. Likhtman, *Extension of the Algebra of Poincare
Group Generators and Violation of p Invariance*,
JETP Lett. **13** (1971) p. 323, [Pisma Zh. Eksp. Teor. Fiz.13,452(1971)].
- [27] A. Neveu and J. H. Schwarz, *Factorizable dual model of pions*,
Nucl. Phys. **B31** (1971) p. 86.
- [28] A. Neveu and J. H. Schwarz, *Quark Model of Dual Pions*,
Phys. Rev. **D4** (1971) p. 1109.
- [29] D. V. Volkov and V. P. Akulov, *Is the Neutrino a Goldstone Particle?*
Phys. Lett. **B46** (1973) p. 109.
- [30] J. Wess and B. Zumino,
A Lagrangian Model Invariant Under Supergauge Transformations,
Phys. Lett. **B49** (1974) p. 52.
- [31] A. Salam and J. A. Strathdee, *Supersymmetry and Nonabelian Gauges*,
Phys. Lett. **B51** (1974) p. 353.
- [32] S. Ferrara, J. Wess, and B. Zumino, *Supergauge Multiplets and Superfields*,
Phys. Lett. **B51** (1974) p. 239.
- [33] J. Wess and B. Zumino, *Supergauge Transformations in Four-Dimensions*,
Nucl. Phys. **B70** (1974) p. 39.
- [34] G. R. Farrar and P. Fayet, *Phenomenology of the Production, Decay, and
Detection of New Hadronic States Associated with Supersymmetry*,
Phys. Lett. **B76** (1978) p. 575.
- [35] ATLAS Collaboration,
*Search for the electroweak production of supersymmetric particles in
 $\sqrt{s} = 8$ TeV pp collisions with the ATLAS detector*,
Phys. Rev. D **93** (2016) p. 052002, arXiv: 1509.07152 [hep-ex].
- [36] ATLAS Collaboration, *Summary of the searches for squarks and gluinos using
 $\sqrt{s} = 8$ TeV pp collisions with the ATLAS experiment at the LHC*,
JHEP **10** (2015) p. 054, arXiv: 1507.05525 [hep-ex].

- 472 [37] ATLAS Collaboration, *ATLAS Run 1 searches for direct pair production of*
473 *third-generation squarks at the Large Hadron Collider*,
474 *Eur. Phys. J. C* **75** (2015) p. 510, arXiv: [1506.08616 \[hep-ex\]](#).
- 475 [38] CMS Collaboration,
476 *Search for supersymmetry with razor variables in pp collisions at $\sqrt{s} = 7$ TeV*,
477 *Phys. Rev. D* **90** (2014) p. 112001, arXiv: [1405.3961 \[hep-ex\]](#).
- 478 [39] CMS Collaboration, *Inclusive search for supersymmetry using razor variables*
479 *in pp collisions at $\sqrt{s} = 7$ TeV*, *Phys. Rev. Lett.* **111** (2013) p. 081802,
480 arXiv: [1212.6961 \[hep-ex\]](#).
- 481 [40] CMS Collaboration, *Search for Supersymmetry in pp Collisions at 7 TeV in*
482 *Events with Jets and Missing Transverse Energy*,
483 *Phys. Lett. B* **698** (2011) p. 196, arXiv: [1101.1628 \[hep-ex\]](#).
- 484 [41] CMS Collaboration, *Search for Supersymmetry at the LHC in Events with*
485 *Jets and Missing Transverse Energy*, *Phys. Rev. Lett.* **107** (2011) p. 221804,
486 arXiv: [1109.2352 \[hep-ex\]](#).
- 487 [42] CMS Collaboration, *Search for supersymmetry in hadronic final states using*
488 *M_{T2} in pp collisions at $\sqrt{s} = 7$ TeV*, *JHEP* **10** (2012) p. 018,
489 arXiv: [1207.1798 \[hep-ex\]](#).
- 490 [43] CMS Collaboration, *Searches for supersymmetry using the M_{T2} variable in*
491 *hadronic events produced in pp collisions at 8 TeV*, *JHEP* **05** (2015) p. 078,
492 arXiv: [1502.04358 \[hep-ex\]](#).
- 493 [44] CMS Collaboration, *Search for new physics with the M_{T2} variable in all-jets*
494 *final states produced in pp collisions at $\sqrt{s} = 13$ TeV* (2016),
495 arXiv: [1603.04053 \[hep-ex\]](#).
- 496 [45] ATLAS Collaboration, *Multi-channel search for squarks and gluinos in*
497 *$\sqrt{s} = 7$ TeV pp collisions with the ATLAS detector at the LHC*,
498 *Eur. Phys. J. C* **73** (2013) p. 2362, arXiv: [1212.6149 \[hep-ex\]](#).