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A search for sparticles in zero lepton final states

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ABSTRACT

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A search for sparticles in zero lepton final states

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Acknowledgements

Introduction

Particle physics is a remarkably successful field of scientific inquiry. The ability to precisely predict the properties of a exceedingly wide range of physical phenomena, such as the description of the cosmic microwave background [1, 2], the understanding of the anomalous magnetic dipole moment of the electron [3, 4], and the measurement of the number of weakly-interacting neutrino flavors [5] is truly amazing.

The theory that has allowed this range of predictions is the *Standard Model* of particle physics (SM). The Standard Model combines the electroweak theory of Glashow, Weinberg, and Salam [6–8] with the theory of the strong interactions, as first envisioned by Gell-Mann and Zweig [9, 10]. This quantum field theory (QFT) contains a tiny number of particles, whose interactions describe phenomena up to at least the TeV scale. These particles are manifestations of the fields of the Standard Model, after application of the Higgs Mechanism. The particle content of the SM consists only of the six quarks, the six leptons, the four gauge bosons, and the scalar Higgs boson.

Despite its impressive range of described phenomena, the Standard Model has some theoretical and experimental deficiencies. The SM contains 26 free parameters¹. It would be more theoretically pleasing to understand these free parameters in terms of a more fundamental theory. The major theoretical concern of the Standard Model, as it pertains to this thesis, is the *hierachy problem*[11–15]. The light mass

¹This is the Standard Model corrected to include neutrino masses. These parameters are the fermion masses (6 leptons, 6 quarks), CKM and PMNS mixing angles (8 angles, 2 CP-violating phases), W/Z/Higgs masses (3), the Higgs field expectation value, and the couplings of the strong, weak, and electromagnetic forces (3 α_{force}) .

84 of the Higgs boson (125 GeV) should be quadratically dependent on the scale of UV
 85 physics, due to the quantum corrections from high-energy physics processes. The
 86 most perplexing experimental issue is the existence of *dark matter*, as demonstrated
 87 by galactic rotation curves [16–22]. This data has shown that there exists additional
 88 matter which has not yet been seen interacting with the particles of the Standard
 89 Model. There is no particle in the SM which can act as a candidate for dark matter.

90 Both of these major issues, as well as numerous others, can be solved by the
 91 introduction of *supersymmetry* (SUSY) [15, 23–35]. In supersymmetric theories, each
 92 SM particles has a so-called *superpartner*, or sparticle partner, differing from given SM
 93 particle by 1/2 in spin. These theories solve the hierachy problem, since the quantum
 94 corrections induced from the superpartners exactly cancel those induced by the SM
 95 particles. In addition, these theories are usually constructed assuming R -parity,
 96 which can be thought of as the “charge” of supersymmetry, with SM particles having
 97 $R = 1$ and sparticles having $R = -1$. In collider experiments, since the incoming
 98 SM particles have total $R = 1$, the resulting sparticles are produced in pairs. This
 99 produces a rich phenomenology, which is characterized by significant hadronic activity
 100 and large missing transverse energy (E_T^{miss}), which provide significant discrimination
 101 against SM backgrounds [36].

102 Despite the power of searches for supersymmetry where E_T^{miss} is a primary
 103 discriminating variable, there has been significant interest in the use of other variables
 104 to discriminate against SM backgrounds. These include searches employing variables
 105 such as αT , $M_{T,2}$, and the razor variables (M_R, R^2) [37–47]. In this thesis, we
 106 will present the first search for supersymmetry using the novel Recursive Jigsaw
 107 Reconstruction (RJR) technique. RJR can be considered the conceptual successor
 108 of the razor variables. We impose a particular final state “decay tree” on an events,
 109 which roughly corresponds to a simplified Feynmann diagram in decays containing
 110 weakly-interacting particles. We account for the missing degrees of freedom associated

111 to the weakly-interacting particles by a series of simplifying assumptions, which allow
112 us to calculate our variables of interest at each step in the decay tree. This allows an
113 unprecedented understanding of the internal structure of the decay and the ability to
114 construct additional variables to reject Standard Model backgrounds.

115 This thesis details a search for the superpartners of the gluon and quarks, the
116 gluino and squarks, in final states with zero leptons, with 13.3 fb^{-1} of data using the
117 ATLAS detector. We organize the thesis as follows. The theoretical foundations of
118 the Standard Model and supersymmetry are described in Chapters 2 and 3. The
119 Large Hadron Collider and the ATLAS detector are presented in Chapters 4 and 5.
120 Chapter 5 provides a detailed description of Recursive Jigsaw Reconstruction and a
121 description of the variables used for the particular search presented in this thesis.
122 Chapter 6 presents the details of the analysis, including details of the dataset, object
123 reconstruction, and selections used. In Chapter 7, the final results are presented;
124 since there is no evidence of a supersymmetric signal in the analysis, we present the
125 final exclusion curves in simplified supersymmetric models.

The Standard Model

2.1 Overview

A Standard Model is another name for a theory of the internal symmetry group $SU(3)_C \otimes SU(2)_L \otimes U(1)_Y$, with its associated set of parameters. *The* Standard Model refers specifically to a Standard Model with the proper parameters to describe the universe. The SM is the culmination of years of work in both theoretical and experimental particle physics. In this thesis, we take the view that theorists construct a model with the field content and symmetries as inputs, and write down the most general Lagrangian consistent with those symmetries. Assuming this model is compatible with nature (in particular, the predictions of the model are consistent with previous experiments), experimentalists are responsible measuring the parameters of this model. This will be applicable for this chapter and the following one.

cite

Additional theoretical background is in 9.6. The philosophy and notations are inspired by [48, 49].

2.2 Field Content

The Standard Model field content is

$$\begin{aligned}
 \text{Fermions} &: Q_L(3, 2)_{+1/3}, U_R(3, 1)_{+4/3}, D_R(3, 1)_{-2/3}, L_L(1, 2)_{-1}, E_R(1, 1)_{-2} \\
 \text{Scalar (Higgs)} &: \phi(1, 2)_{+1} \\
 \text{Vector Fields} &: G^\mu(8, 1)_0, W^\mu(1, 3)_0, B^\mu(1, 1)_0
 \end{aligned} \tag{2.1}$$

142 where the $(A, B)_Y$ notation represents the irreducible representation under $SU(3)$
 143 and $SU(2)$, with Y being the electroweak hypercharge. Each of these fermion fields
 144 has an additional index, representing the three generation of fermions.

145 We observed that Q_L, U_R , and D_R are triplets under $SU(3)_C$; these are the *quark*
 146 fields. The *color* group, $SU(3)_C$ is mediated by the *gluon* field $G^\mu(8, 1)_0$, which has
 147 8 degrees of freedom. The fermion fields $L_L(1, 2)_{-1}$ and $E_R(1, 1)_{-2}$ are singlets under
 148 $SU(3)_C$; we call them the *lepton* fields.

149 Next, we note the “left-handed” (“right-handed”) fermion fields, denoted by L (R)
 150 subscript, The left-handed fields form doublets under $SU(2)_L$. These are mediated
 151 by the three degrees of freedom of the “W” fields $W^\mu(1, 3)_0$. These fields only act
 152 on the left-handed particles of the Standard Model. This is the reflection of the
 153 “chirality” of the Standard Model; the left-handed and right-handed particles are
 154 treated differently by the electroweak forces. The right-handed fields, U_R, D_R , and
 155 E_R , are singlets under $SU(2)_L$.

156 The $U(1)_Y$ symmetry is associated to the $B^\mu(1, 1)_0$ boson with one degree of
 157 freedom. The charge Y is known as the electroweak hypercharge.

158 To better understand the phenomenology of the Standard Model, let us investigate
 159 each of the *sectors* of the Standard Model separately.

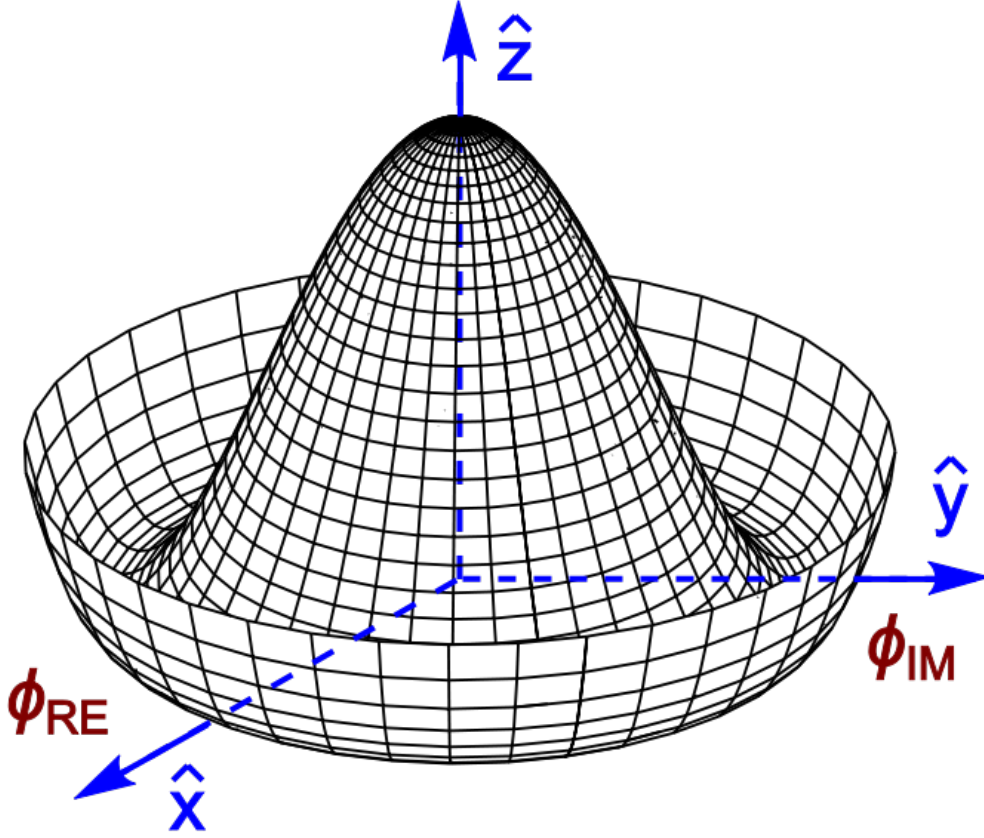
160 Electroweak sector

The electroweak sector refers to the $SU(2)_L \otimes U(1)_Y$ portion of the Standard
 Model gauge group. Following our philosophy of writing all gauge-invariant and
 renormalizable terms, the electroweak Lagrangian can be written as

$$\mathcal{L} = W_a^{\mu\nu} W_{\mu\nu}^a + B^{\mu\nu} B_{\mu\nu} + (D^\mu \phi)^\dagger D_\mu \phi - \mu^2 \phi^\dagger \phi - \lambda(\phi^\dagger \phi)^2. \quad (2.2)$$

where $W_a^{\mu\nu}$ are the three ($a = 1, 2, 3$) gauge bosons associated to the $SU(2)_L$ gauge
 group, $B^{\mu\nu}$ is the one gauge boson of the $U(1)_Y$ gauge group, and ϕ is the complex

Figure 2.1: Sombrero potential



Higgs multiplet. The covariant derivative D^μ is given by

$$D^\mu = \partial^\mu + \frac{ig}{2} W_a^\mu \sigma_a + \frac{ig'}{2} B^\mu \quad (2.3)$$

where $i\sigma_a$ are the Pauli matrices times the imaginary constant, which are the generators for $SU(2)_L$, and g and g' are the $SU(2)_L$ and $U(1)_Y$ coupling constants, respectively. The field strength tensors $W_a^{\mu\nu}$ and $B^{\mu\nu}$ are given by the commutator of the covariant derivative associated to each field

$$B^{\mu\nu} = \partial^\mu B^\nu - \partial^\nu B^\mu \quad (2.4)$$

$$W_a^{\mu\nu} = \partial^\mu W_a^\nu - \partial^\nu W_a^\mu - g\epsilon_{abc} W_a^\mu W_b^\nu, \quad i = 1, 2, 3$$

162 The terms in the Lagrangian 2.2 proportional to μ^2 and λ make up the “Higgs
 163 potential” [50]. As normal (see Appendix 9.6), we restrict $\lambda > 0$ to guarantee our
 164 potential is bounded from below, and we also require $\mu^2 < 0$, which gives us the
 165 standard “sombbrero” potential shown in 2.1.

This potential has infinitely many minima at $\langle \phi \rangle = \sqrt{2m/\lambda}$; the ground state is *spontaneously* broken by the choice of ground state, which induces a vacuum expectation value (VEV). Without loss of generality, we can choose the Higgs field ϕ to point in the real direction, and write the Higgs field ϕ in the following form :

$$\phi = \frac{1}{\sqrt{2}} \exp\left(\frac{i}{v} \sigma_a \theta_a\right) \begin{pmatrix} 0 \\ v + h(x) \end{pmatrix}. \quad (2.5)$$

We choose a gauge to rotate away the dependence on θ_a , such that we can write simply

$$\phi = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v + h(x) \end{pmatrix}. \quad (2.6)$$

Now, we can see how the masses of the vector bosons are generated from the application of the Higgs mechanism. We plug Eq.2.6 back into the electroweak Lagrangian, and only showing the relevant mass terms in the vacuum state where $h(x) = 0$ see that (dropping the Lorentz indices) :

$$\begin{aligned} \mathcal{L}_M &= \frac{1}{8} \left| \begin{pmatrix} gW_3 + g'B & g(W_1 - iW_2) \\ g(W_1 + iW_2) & -gW_3 + g'B \end{pmatrix} \begin{pmatrix} 0 \\ v \end{pmatrix} \right|^2 \\ &= \frac{g^2 v^2}{8} \left[W_1^2 + W_2^2 + \left(\frac{g'}{g} B - W_3\right)^2 \right] \end{aligned} \quad (2.7)$$

Defining the *Weinberg* angle $\tan(\theta_W) = g'/g$ and the following *physical* fields :

$$W^\pm = \frac{1}{\sqrt{2}}(W_1 \mp iW_2) \quad (2.8)$$

$$Z^0 = \cos \theta_W W_3 - \sin \theta_W B$$

$$A^0 = \sin \theta_W W_3 + \cos \theta_W B$$

we can write the piece of the Lagrangian associated to the vector boson masses as

$$\mathcal{L}_{M_V} = \frac{1}{4}g^2v^2W^+W^- + \frac{1}{8}(g^2 + g'^2)v^2Z^0Z^0. \quad (2.9)$$

and we have the following values of the masses for the vector bosons :

$$\begin{aligned} m_W^2 &= \frac{1}{4}v^2g^2 \\ m_Z^2 &= \frac{1}{4}v^2(g^2 + g'^2) \\ m_A^2 &= 0 \end{aligned} \quad (2.10)$$

166 We thus see how the Higgs mechanism gives rise to the masses of the W^\pm and Z
 167 boson in the Standard Model; the mass of the photon is zero, as expected. The
 168 $SU(2)_L \otimes U(1)_Y$ symmetry of the initially massless $W_{1,2,3}$ and B fields is broken to
 169 the $U(1)_{EM}$. Of the four degrees of freedom in the complex Higgs doublet, three are
 170 “eaten” when we give mass to the W^\pm and Z_0 , while the other degree of freedom is
 171 the Higgs particle, as found in 2012 by the ATLAS and CMS collaborations [51, 52].

172 Quantum Chromodynamics

Quantum chromodynamics (or the theory of the *strong* force) characterizes the behavior of *colored* particles, collectively known as *partons*. The partons of the Standard Model are the (fermionic) quarks, and the (bosonic) gluons. The strong force is governed by $SU(3)_C$, an unbroken symmetry in the Standard Model, which implies the gluon remains massless. Defining the covariant derivative for QCD as

$$D^\mu = \partial^\mu + ig_s G_a^\mu L_a, a = 1, \dots, 8 \quad (2.11)$$

where L_a are the generators of $SU(3)_C$, and g_s is the coupling constant of the strong force. The QCD Lagrangian then is given by

$$\mathcal{L}_{\text{QCD}} = i\bar{\psi}_f D_\mu \gamma^\mu \psi_f - \frac{1}{4}G_{a,\mu\nu}G_a^{\mu\nu} \quad (2.12)$$

where the summation over f is for quarks *families*, and $G_a^{\mu\nu}$ is the gluon field strength tensor, given by

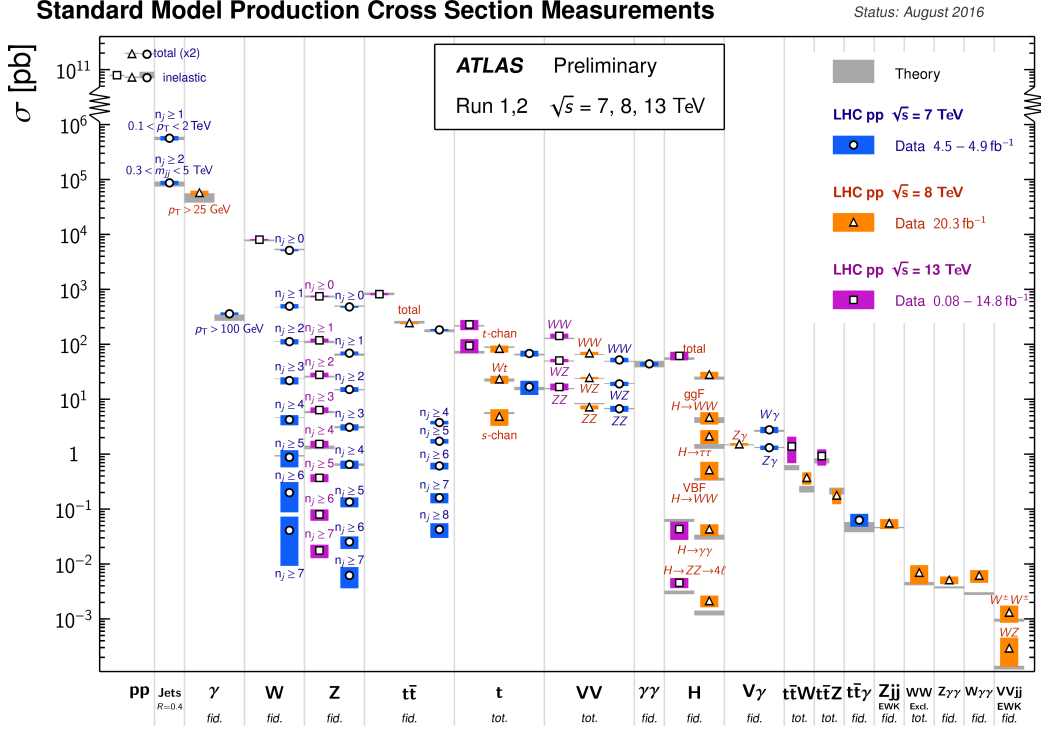
$$G_a^{\mu\nu} = \partial^\mu G_a^\nu - \partial^\nu G_a^\mu - g_s f^{abc} G_b^\mu G_c^\nu, a, b, c = 1, \dots, 8 \quad (2.13)$$

173 where f^{abc} are the structure constants of $SU(3)_C$, which are analogous to ϵ_{abc} for
 174 $SU(2)_L$. The kinetic term for the quarks is contained in the standard ∂_μ term, while
 175 the field strength term contains the interactions between the quarks and gluons, as
 176 well as the gluon self-interactions.

177 Written down in this simple form, the QCD Lagrangian does not seem much
 178 different from the QED Lagrangian, with the proper adjustments for the different
 179 group structures. The gluon is massless, like the photon, so one could naïvely expect
 180 an infinite range force, and it pays to understand why this is not the case. The
 181 reason for this fundamental difference is the gluon self-interactions arising in the
 182 field strength tensor term of the Lagrangian. This leads to the phenomena of *color*
 183 *confinement*, which describes how one only observes color-neutral particles alone in
 184 nature. In contrast to the electromagnetic force, particles which interact via the
 185 strong force experience a *greater* force as the distance between the particles increases.
 186 At long distances, the potential is given by $V(r) = -kr$. At some point, it is more
 187 energetically favorable to create additional partons out of the vacuum than continue
 188 pulling apart the existing partons, and the colored particles undergo *fragmentation*.
 189 This leads to *hadronization*. Bare quarks and gluons are actually observed as sprays
 190 of hadrons (primarily kaons and pions); these sprays are known as *jets*, which are
 191 what are observed by experiments.

192 It is important to recognize the importance of understanding these QCD inter-
 193 actions in high-energy hadron colliders such as the LHC. Since protons are hadrons,
 194 proton-proton collisions such as those produced by the LHC are primarily governed by
 195 the processes of QCD. In particular, by far the most frequent process observed in LHC
 196 experiments is dijet production from gluon-gluon interactions (see Fig.2.2). These

Figure 2.2: Cross-sections of various Standard Model processes



gluons that interact are part of the *sea* particles inside the proton; the simple $p = uud$ model does not apply. The main *valence* uud quarks are constantly interacting via gluons, which can themselves radiate gluons or split into quarks, and so on. A more useful understanding is given by the colloquially-known *bag* model [53, 54], where the proton is seen as a “bag” of (in principle) infinitely many partons, each with energy $E < \sqrt{s} = 6.5$ TeV. One then collides this (proton) bag with another, and views the products of this very complicated collision, where calculations include many loops in nonperturbative QCD calculations.

Fortunately, we are generally saved by the QCD factorization theorems [55]. This allows one to understand the hard (i.e. short distance or high energy) $2 \rightarrow 2$ parton process using the tools of perturbative QCD, while making series of approximations known as a *parton shower* model to understand the additional corrections from nonperturbative QCD. We will discuss the reconstruction of jets by experiments in Ch.5.

211 Fermions

212 We will now look more closely at the fermions in the Standard Model [56].

213 As noted earlier in Sec.2.2, the fermions of the Standard Model can be first
 214 distinguished between those that interact via the strong force (quarks) and those
 215 which do not (leptons).

There are six leptons in the Standard Model, which can be placed into three *generations*.

$$\begin{pmatrix} e \\ \nu_e \end{pmatrix}, \begin{pmatrix} \mu \\ \nu_\mu \end{pmatrix}, \begin{pmatrix} \tau \\ \nu_\tau \end{pmatrix} \quad (2.14)$$

216 There is the electron (e), muon (μ), and tau (τ), each of which has an associated
 217 neutrino (ν_e, ν_μ, ν_τ). Each of the so-called charged (“electron-like”) leptons has
 218 electromagnetic charge -1 , while the neutrinos all have $q_{EM} = 0$.

219 Often in an experimental context, lepton is used to denote the stable electron
 220 and metastable muon, due to their striking experimental signatures. Taus are often
 221 treated separately, due to their much shorter lifetime of $\tau_\tau \sim 10^{-13}s$; these decay
 222 through hadrons or the other leptons, so often physics analyses at the LHC treat
 223 them as jets or leptons, as will be done in this thesis.

224 As the neutrinos are electrically neutral, nearly massless, and only interact via the
 225 weak force, it is quite difficult to observe them directly. Since LHC experiments rely
 226 overwhelmingly on electromagnetic interactions to observe particles, the presence of
 227 neutrinos is not observed directly. Neutrinos are instead observed by the conservation
 228 of four-momentum in the plane transverse to the proton-proton collisions, known as
 229 *missing transverse energy*.

There are six quarks in the Standard Model : up, down, charm, strange, top, and bottom. Quarks are similar organized into three generations :

$$\begin{pmatrix} u \\ d \end{pmatrix}, \begin{pmatrix} c \\ s \end{pmatrix}, \begin{pmatrix} t \\ b \end{pmatrix} \quad (2.15)$$

230 where we speak of “up-like” quarks and “down-like” quarks.

231 Each up-like quark has charge $q_{up} = 2/3$, while the down-like quarks have $q_{down} =$
232 $-1/3$. At the high energies of the LHC, one often makes the distinction between
233 the light quarks (u, d, c, s), the bottom quark, and top quark. In general, due to
234 the hadronization process described above, the light quarks, with masses $m_q < \sim$
235 1.5GeV are indistinguishable by LHC experiments. Their hadronic decay products
236 generally have long lifetimes and they are reconstructed as jets.¹ The bottom quark
237 hadronizes primarily through the B -mesons, which generally travels a short distance
238 before decaying to other hadrons. This allows one to distinguish decays via b -quarks
239 from other jets; this procedure is known as *b-tagging* and will be discussed more in
240 Ch.5. Due to its large mass, the top quark decays before it can hadronize; there
241 are no bound states associated to the top quark. The top is of particular interest at
242 the LHC; it has a striking signature through its most common decay mode $t \rightarrow Wb$.
243 Decays via tops, especially $t\bar{t}$ are frequently an important signal decay mode, or an
244 important background process.

245 Interactions in the Standard Model

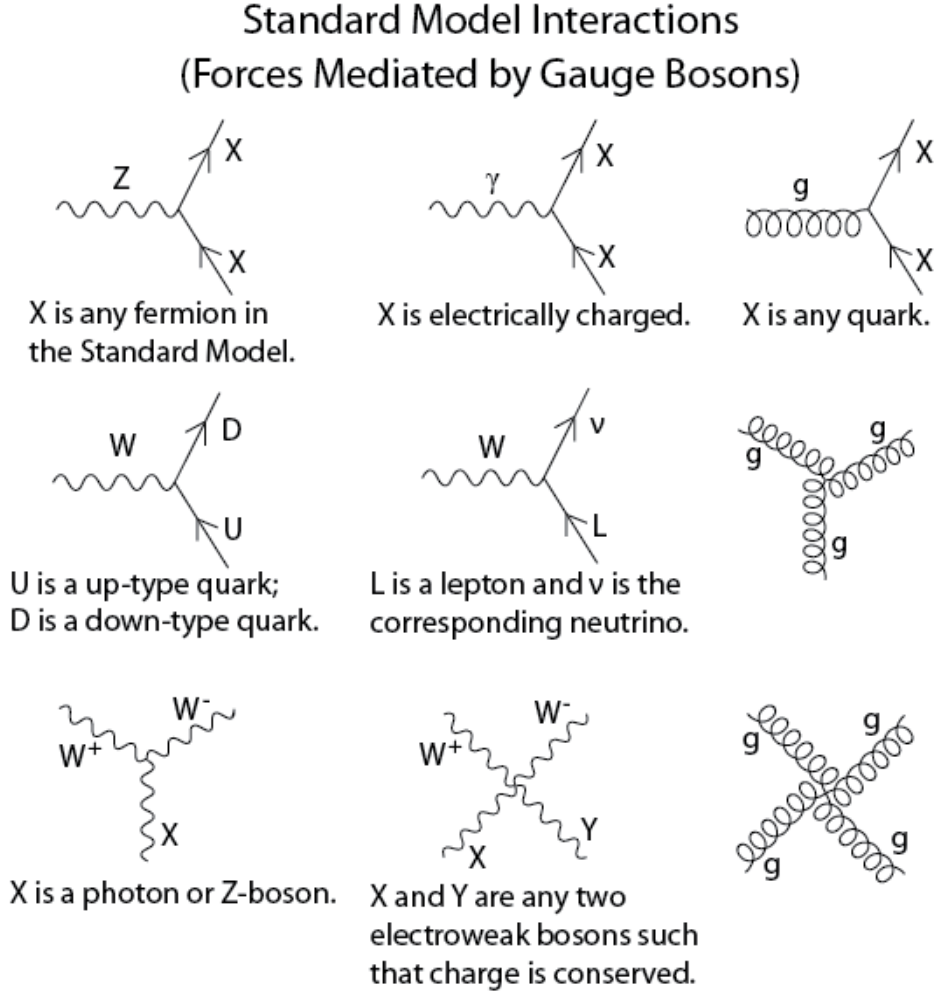
246 We briefly overview the entirety of the fundamental interactions of the Standard
247 Model; these can also be found in 2.3.

248 The electromagnetic force, mediated by the photon, interacts with via a three-
249 point coupling all charged particles in the Standard Model. The photon thus interacts
250 with all the quarks, the charged leptons, and the charged W^\pm bosons.

251 The weak force is mediated by three particles : the W^\pm and the Z^0 . The Z^0 can
252 interact with all fermions via a three-point coupling. A real Z_0 can thus decay to
253 a fermion-antifermion pair of all SM fermions except the top quark, due to its large

¹In some contexts, charm quarks are also treated as a separate category, although it is quite difficult to distinguish charm quarks from the other light quarks.

Figure 2.3: The interactions of the Standard Model



254 mass. The W^\pm has two important three-point interactions with fermions. First, the
 255 W^\pm can interact with an up-like quark and a down-like quark; an important example
 256 in LHC experiments is $t \rightarrow Wb$. The coupling constants for these interactions are
 257 encoded in the unitary matrix known as the Cabibbo–Kobayashi–Maskawa (CKM)
 258 matrix [57, 58], and are generally known as flavor-changing interactions. Secondly,
 259 the W^\pm interacts with a charged lepton and its corresponding neutrino. In this case,
 260 the unitary matrix that corresponds to CKM matrix for quarks is the identity matrix,
 261 which forbids (fundamental) vertices such as $\mu \rightarrow We$. For leptons, instead this is
 262 a two-step process : $\mu \rightarrow \nu_\mu u W \rightarrow \nu_\mu u \bar{\nu}_e e$. Finally, there are the self-interactions

of the weak gauge bosons. There is a three-point and four-point interaction; all combinations are allowed which conserve electric charge.

The strong force is mediated by the gluon, which as discussed above also carries the strong color charge. There is the fundamental three-point interaction, where a quark radiates a gluon. Additionally, there are the three-point and four-point gluon-only interactions.

2.3 Deficiencies of the Standard Model

At this point, it is quite easy to simply rest on our laurels. This relatively simple theory is capable of explaining a very wide range of phenomena, which ultimately break down only to combinations of nine diagrams shown in Fig.2.3. Unfortunately, there are some unexplained problems with the Standard Model. We cannot go through all of the potential issues in this thesis, but we will motivate the primary issues which naturally lead one to *supersymmetry*, as we will see in Ch.3.

The Standard Model has many free parameters; see Table 2.1. In general, we prefer models with less free parameters. A great example of this fact, and the primary experimental evidence for EWSB, is the relationship between the couplings of the weak force and the masses of the gauge bosons of the weak force :

$$\rho \equiv \frac{m_W^2}{m_Z^2 \cos^2 \theta_W} \stackrel{?}{=} 1 \quad (2.16)$$

where ? indicates that this is a testable prediction of the Standard Model (in particular, that the gauge bosons gain mass through EWSB). This relationship has been measured within experimental and theoretical predictions. We would like to produce additional such relationships, which would exist if the Standard Model is a low-energy approximation of some other theory.

An additional issue is the lack of *gauge coupling unification*. The couplings of any quantum field theory “run” as a function of the distance scales (or inversely,

Table 2.1: Parameters of the Standard Model. For values dependent on the renormalization scheme, we use a combination of the on-shell normalization scheme [59–62] and modified minimal subtraction scheme with $m_{\bar{M}S}$ as indicated in the table[63]

m_e	Electron mass	511 keV
m_μ	Muon mass	105.7 MeV
m_τ	Tau mass	1.78 GeV
m_u	Up quark mass	1.9 MeV ($m_{\bar{M}S} = 2GeV$)
m_d	Down quark mass	4.4 MeV ($m_{\bar{M}S} = 2GeV$)
m_s	Strange quark mass	87 MeV ($m_{\bar{M}S} = 2GeV$)
m_c	Charm quark mass	1.32 GeV ($m_{\bar{M}S} = m_c$)
m_b	Bottom quark mass	4.24 GeV ($m_{\bar{M}S} = m_b$)
m_t	Top quark mass	172.7 GeV (on-shell renormalization)
θ_{12} CKM	12-mixing angle	13.1°
θ_{23} CKM	23-mixing angle	2.4°
θ_{13} CKM	13-mixing angle	0.2°
δ CKM	CP-violating Phase	0.995
g'	U(1) gauge coupling	0.357 ($m_{\bar{M}S} = m_Z$)
g	SU(2) gauge coupling	0.652 ($m_{\bar{M}S} = m_Z$)
g_s	SU(3) gauge coupling	1.221 ($m_{\bar{M}S} = m_Z$)
θ_{QCD}	QCD vacuum angle	~ 0
VEV	Higgs vacuum expectation value	246 GeV
m_H	Higgs mass	125 GeV

energy scales) of the theory. The idea is closely related to the unification of the electromagnetic and weak forces at the so-called *electroweak scale* of $O(100 \text{ GeV})$. One would hope this behavior was repeated between the electroweak forces and the strong force at some suitable energy scale. The Standard Model does automatically not exhibit this behavior, as we can see in Fig.2.4.

The most significant problem with the Standard Model is the *hierarchy problem*. In its most straightforward incarnation, the Higgs scalar field is subject to quantum corrections through loop diagrams, as shown in Fig.2.5. For demonstration, we use the contributions from the top quark, since the top quark has the largest Higgs Yukawa coupling due to its large mass. In general, we should expect these corrections to quadratically dependent on the scale of the ultraviolet physics, Λ . Briefly assume there is no new physics before the Planck scale of gravity, $\Lambda_{\text{Planck}} = 10^{19} \text{ GeV}$. In this

Figure 2.4: The running of Standard Model gauge couplings. The Standard Model couplings do not unify at high energies, which indicates it cannot completely describe nature through the Planck scale.

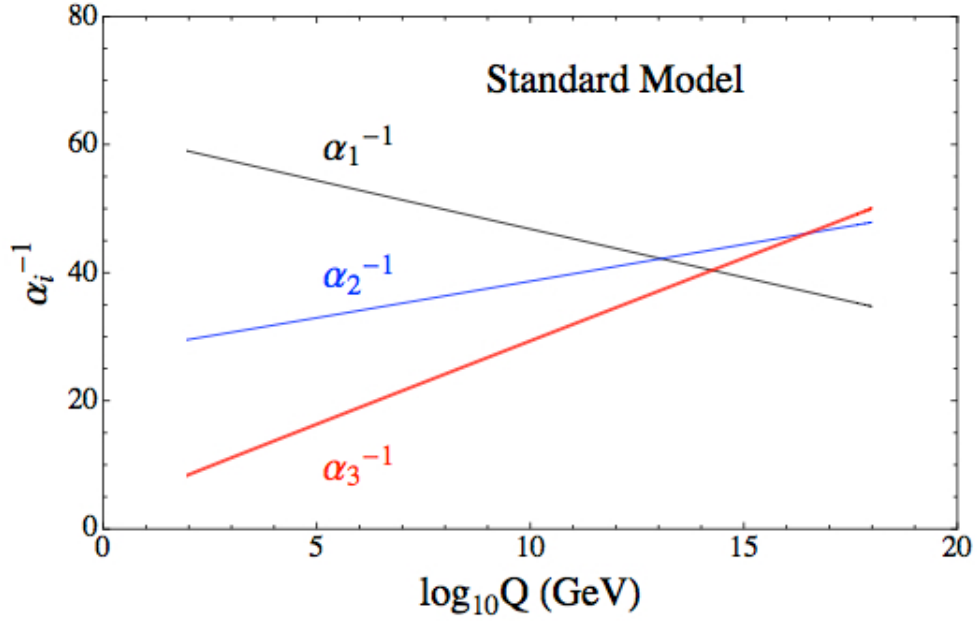
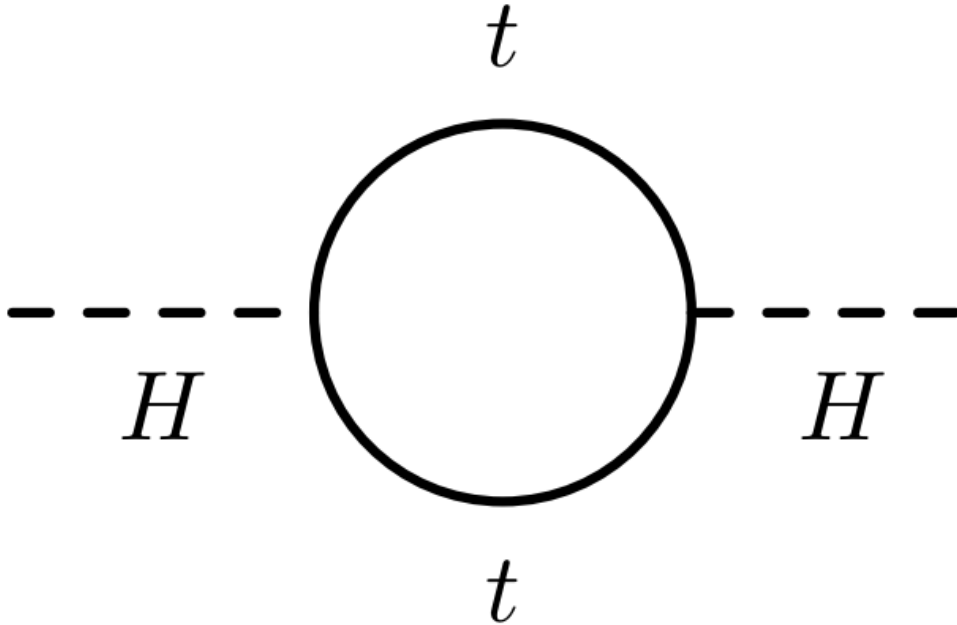


Figure 2.5: The dominant quantum loop correction to the Higgs mass in the Standard Model.



case, we expect the corrections to the Higgs mass like

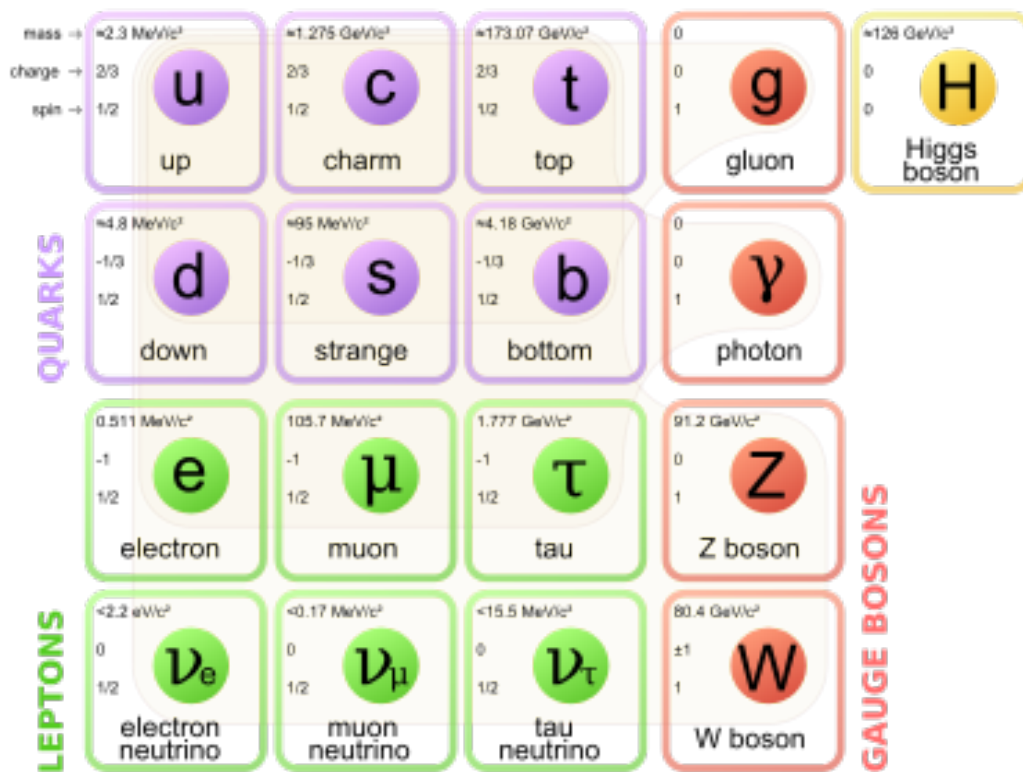
$$\delta m_H^2 \approx \left(\frac{m_t}{8\pi^2 \langle \phi \rangle_{VEV}} \right)^2 \Lambda_{Planck}^2. \quad (2.17)$$

288 To achieve the miraculous cancellation required to get the observed Higgs mass of
289 125 GeV, one needs to then set the bare Higgs mass m_0 , our input to the Standard
290 Model Lagrangian, itself to a *precise* value $\sim 10^{19}$ GeV. This extraordinary level of
291 parameter finetuning is quite undesirable, and within the framework of the Standard
292 Model, there is little that can be done to alleviate this issue.

293 An additional concern, of a different nature, is the lack of a *dark matter* candidate
294 in the Standard Model. Dark matter was discovered by observing galactic rotation
295 curves, which showed that much of the matter that interacted gravitationally was
296 invisible to our (electromagnetic) telescopes [16–22]. The postulation of the existence
297 of dark matter, which interacts at least through gravity, allows one to understand
298 these galactic rotation curves. Unfortunately, no particle in the Standard Model could
299 possibly be the dark matter particle. The only candidate truly worth another look is
300 the neutrino, but it has been shown that the neutrino content of the universe is simply
301 too small to explain the galactic rotation curves [22, 64]. The experimental evidence
302 from the galactic rotations curves thus show there *must* be additional physics beyond
303 the Standard Model, which is yet to be understood.

304 In the next chapter, we will see how these problems can be alleviated by the theory
305 of supersymmetry.

Figure 2.6: Particles of the Standard Model



Supersymmetry

This chapter will introduce supersymmetry (SUSY) [15, 23–35]. We will begin by introducing the concept of a *superspace*, and discuss some general ingredients of supersymmetric theories. This will include a discussion of how the problems with the Standard Model described in Ch.2 are naturally fixed by these theories.

The next step is to discuss the particle content of the *Minimally Supersymmetric Standard Model* (MSSM). As its name implies, this theory contains the minimal additional particle content to make Standard Model supersymmetric. We then discuss the important phenomenological consequences of this theory, especially as it would be observed in experiments at the LHC.

3.1 Supersymmetric theories : from space to superspace

Coleman-Mandula “no-go” theorem

We begin the theoretical motivation for supersymmetry by citing the “no-go” theorem of Coleman and Mandula [65]. This theorem forbids *spin-charge unification*; it states that all quantum field theories which contain nontrivial interactions must be a direct product of the Poincaré group of Lorentz symmetries, the internal product from of gauge symmetries, and the discrete symmetries of parity, charge conjugation, and time reversal. The assumptions which go into building the Coleman-Mandula

theorem are quite restrictive, but there is one unique way out, which has become known as *supersymmetry* [26, 66]. In particular, we must introduce a *spinorial* group generator Q . Alternatively, and equivalently, this can be viewed as the addition of anti-commuting coordinates; space plus these new anti-commuting coordinates is then called *superspace* [67]. We will not investigate this view in detail, but it is also a quite intuitive and beautiful way to construct supersymmetry[15].

Supersymmetry transformations

A *supersymmetric* transformation Q transforms a bosonic state into a fermionic state, and vice versa :

$$Q |\text{Fermion}\rangle = |\text{Boson}\rangle \quad (3.1)$$

$$Q |\text{Boson}\rangle = |\text{Fermion}\rangle \quad (3.2)$$

To ensure this relation holds, Q must be an anticommuting spinor. Additionally, since spinors are inherently complex, Q^\dagger must also be a generator of the supersymmetry transformation. Since Q and Q^\dagger are spinor objects (with $s = 1/2$), we can see that supersymmetry must be a spacetime symmetry. The Haag-Lopuszanski-Sohnius extension [66] of the Coleman-Mandula theorem [65] is quite restrictive about the forms of such a symmetry. Here, we simply write the (anti-) commutation relations [15] :

$$Q_\alpha, Q_{\dot{\alpha}}^\dagger = -2\sigma_{\alpha\dot{\alpha}\mu} P_\mu \quad (3.3)$$

$$Q_\alpha, Q_{\dot{\beta}} = Q_{\dot{\alpha}}^\dagger, Q_{\dot{\beta}}^\dagger = 0 \quad (3.4)$$

$$[P^\mu, Q_\alpha] = [P^\mu, Q_{\dot{\alpha}}^\dagger] = 0 \quad (3.5)$$

335 Supermultiplets

336 In a supersymmetric theory, we organize single-particle states into irreducible
337 representations of the supersymmetric algebra which are known as *supermultiplets*.
338 Each supermultiplet contains a fermion state $|F\rangle$ and a boson state $|B\rangle$; these two
339 states are the known as *superpartners*. These are related by some combination of
340 Q and Q^\dagger , up to a spacetime transformation. Q and Q^\dagger commute with the mass-
341 squared operator $-P^2$ and the operators corresponding to the gauge transformations
342 [15]; in particular, the gauge interactions of the Standard Model. In an unbroken
343 supersymmetric theory, this means the states $|F\rangle$ and $|B\rangle$ have exactly the same mass,
344 electromagnetic charge, electroweak isospin, and color charges. One can also prove
345 [15] that each supermultiplet contains the exact same number of bosonic (n_B) and
346 fermion (n_F) degrees of freedom. We now explore the possible types of supermultiples
347 one can find in a renormalizable supersymmetric theory.

348 Since each supermultiplet must contain a fermion state, the simplest type of
349 supermultiplet contains a single Weyl fermion state ($n_F = 2$) which is paired with
350 $n_B = 2$ scalar bosonic degrees of freedom. This is most conveniently constructed as
351 single complex scalar field. We call this construction a *scalar supermultiplet* or *chiral*
352 *supermultiplet*. The second name is indicative; only chiral supermultiplets can contain
353 fermions whose right-handed and left-handed components transform differently under
354 the gauge interactions (as of course happens in the Standard Model).

355 The second type of supermultiplet we construct is known as a *gauge* supermul-
356 tiplet. We take a spin-1 gauge boson (which must be massless due to the gauge
357 symmetry, so $n_B = 2$) and pair this with a single massless Weyl spinor¹. The gauge
358 bosons transform as the adjoint representation of the their respective gauge groups;
359 their fermionic partners, which are known as gauginos, must also. In particular,
360 the left-handed and right-handed components of the gaugino fermions have the same

¹Choosing an $s = 3/2$ massless fermion leads to nonrenormalizable interactions.

361 gauge transformation properties.

362 Excluding gravity, this is the entire list of supermultiplets which can participate
363 in renormalizable interactions in what is known as $N = 1$ supersymmetry. This
364 means there is only one copy of the supersymmetry generators Q and Q^\dagger . This is
365 essentially the only “easy” phenomenological choice, since it is the only choice in four
366 dimensions which allows for the chiral fermions and parity violations built into the
367 Standard Model, and we will not look further into $N > 1$ supersymmetry in this thesis.

368 The primary goal, after understanding the possible structures of the multiplets
369 above, is to fit the Standard Model particles into a multiplet, and therefore make
370 predictions about their supersymmetric partners. We explore this in the next section.

371 3.2 Minimally Supersymmetric Standard Model

372 To construct what is known as the MSSM [[susyPrimer](#), [68–71](#)], we need a few
373 ingredients and assumptions. First, we match the Standard Model particles with
374 their corresponding superpartners of the MSSM. We will also introduce the naming
375 of the superpartners (also known as *sparticles*). We discuss a very common additional
376 restraint imposed on the MSSM, known as R -parity. We also discuss the concept of
377 soft supersymmetry breaking and how it manifests itself in the MSSM.

378 Chiral supermultiplets

379 The first thing we deduce is directly from Sec.???. The bosonic superpartners
380 associated to the quarks and leptons *must* be spin 0, since the quarks and leptons must
381 be arranged in a chiral supermultiplet. This is essentially the note above, since the
382 chiral supermultiplet is the only one which can distinguish between the left-handed
383 and right-handed components of the Standard Model particles. The superpartners of
384 the quarks and leptons are known as *squarks* and *sleptons*, or *sfermions* in aggregate.

385 (for “scalar quarks”, “scalar leptons”, and “scalar fermion”²). The “s-” prefix
386 can also be added to the individual quarks i.e. *selectron*, *sneutrino*, and *stop*. The
387 notation is to add a \sim over the corresponding Standard Model particle i.e. \tilde{e} , the
388 selectron is the superpartner of the electron. The two-component Weyl spinors of the
389 Standard Model must each have their own (complex scalar) partner i.e. e_L, e_R have
390 two distinct partners : \tilde{e}_L, \tilde{e}_R . As noted above, the gauge interactions of any of the
391 sfermions are identical to those of their Standard Model partners.

Due to the scalar nature of the Higgs, it must obviously lie in a chiral supermultiplet. To avoid gauge anomalies and ensure the correct Yukawa couplings to the quarks and leptons[15], we must add additional Higgs bosons to any supersymmetric theory. In the MSSM, we have two chiral supermultiplets. The SM (SUSY) parts of the multiplets are denoted $H_u(\tilde{H}_u)$ and $H_d(\tilde{H}_d)$. Writing out H_u and H_d explicitly:

$$H_u = \begin{pmatrix} H_u^+ \\ H_u^0 \end{pmatrix} \quad (3.6)$$

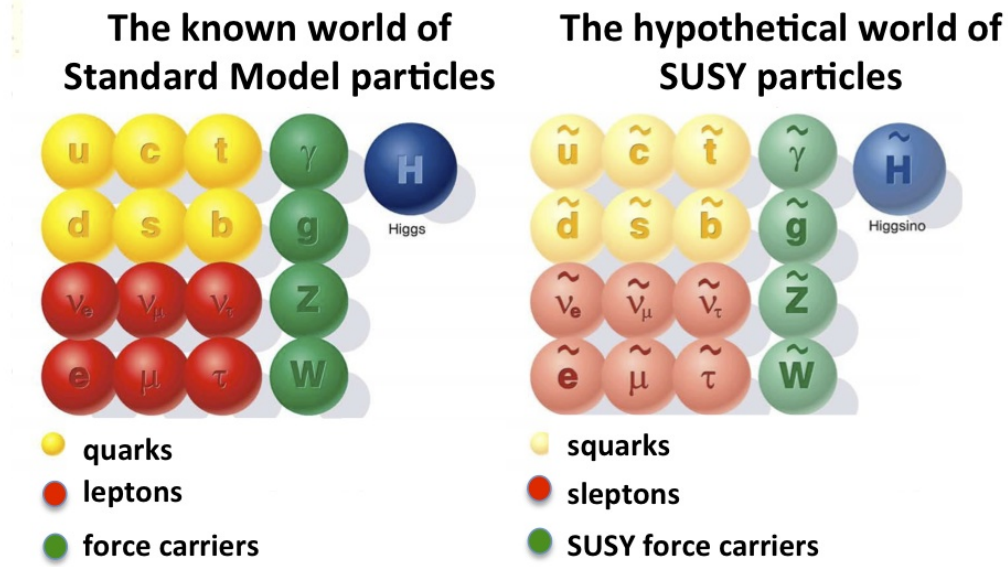
$$H_d = \begin{pmatrix} H_d^0 \\ H_d^- \end{pmatrix} \quad (3.7)$$

$$(3.8)$$

392 we see that H_u looks very similar to the SM Higgs with $Y = 1$, and H_d is symmetric
393 to this with $+$ \rightarrow $-$, with $Y = -1$. The SM Higgs boson, h_0 , is a linear superposition
394 of the neutral components of these two doublets. The SUSY parts of the Higgs
395 multiplets, \tilde{H}_u and \tilde{H}_d , are each left-handed Weyl spinors. For generic spin-1/2
396 sparticles, we add the “-ino” suffix. We then call the partners of the two Higgs
397 collectively the *Higgsinos*.

²The last one should probably have bigger scare quotes.

Figure 3.1: Particles of the MSSM



398 Gauge supermultiplets

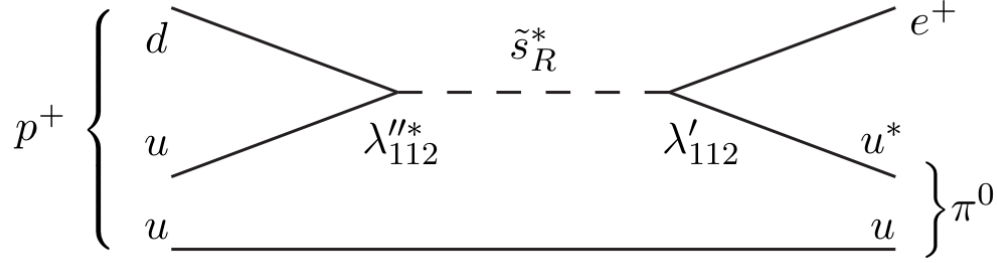
399 The superpartners of the gauge bosons must all be in gauge supermultiplets since
 400 they contain a spin-1 particle. Collectively, we refer to the superpartners of the
 401 gauge bosons as the gauginos.

402 The first gauge supermultiplet contains the gluon, and its superpartner, which is
 403 known as the *gluino*, denoted \tilde{g} . The gluon is of course the SM mediator of $SU(3)_C$;
 404 the gluino is also a colored particle, subject to $SU(3)_C$. From the SM before EWSB,
 405 we have the four gauge bosons of the electroweak symmetry group $SU(2)_L \otimes U(1)_Y$:
 406 $W^{1,2,3}$ and B^0 . The superpartners of these particles are thus the *winos* $\tilde{W}^{1,2,3}$ and
 407 *bino* \tilde{B}^0 , where each is placed in another gauge supermultiplet with its corresponding
 408 SM particle. After EWSB, without breaking supersymmetry, we would also have the
 409 zino \tilde{Z}^0 and photino $\tilde{\gamma}$.

410 The entire particle content of the MSSM can be seen in Fig.3.1.

411 At this point, it's important to take a step back. Where are these particles?
 412 As stated above, supersymmetric theories require that the masses and all quantum

Figure 3.2: This Feynmann diagram shows how proton decay is induced in the MSSM, if one does not impose R -parity.



413 numbers of the SM particle and its corresponding sparticle are the same. Of course,
 414 we have not observed a selectron, squark, or wino. The answer, as it often is, is that
 415 supersymmetry is *broken* by the vacuum state of nature [15].

416 R -parity

This section is a quick aside to the general story. R - *parity* refers to an additional discrete symmetry which is often imposed on supersymmetric models. For a given particle state, we define

$$R = (-1)^{3(B-L)+2s} \quad (3.9)$$

417 where B, L is the baryon (lepton) number and s is the spin. The imposition of
 418 this symmetry forbids certain terms from the MSSM Lagrangian that would violate
 419 baryon and/or lepton number. This is required in order to prevent proton decay, as
 420 shown in Fig.3.2³. .

421 In supersymmetric models, this is a \mathbb{Z}_2 symmetry, where SM particles have $R = 1$
 422 and sparticles have $R = -1$. We will take R - *parity* as part of the definition of
 423 the MSSM. We will discuss later the *drastic* consequences of this symmetry on SUSY
 424 phenomenology

³Proton decay can actually be prevented by allowing only one of the four potential R-parity violating terms to survive.

425 Soft supersymmetry breaking

The fundamental idea of *soft* supersymmetry breaking[15, 34, 35, 72, 73] is that we would like to break supersymmetry without reintroducing the quadratic divergences we discussed at the end of Chapter 2. We write the Lagrangian in a form :

$$\mathcal{L}_{\text{MSSM}} = \mathcal{L}_{\text{SUSY}} + \mathcal{L}_{\text{soft}} \quad (3.10)$$

426 In this sense, the symmetry breaking is “soft”, since we have separated out the
427 completely symmetric terms from those soft terms which will not allow the quadratic
428 divergences to the Higgs mass.

429 The explicitly allowed terms in the soft-breaking Lagrangian are [35].

- 430 • Mass terms for the scalar components of the chiral supermultipletss
- 431 • Mass terms for the Weyl spinor components of the gauge supermultipletss
- 432 • Trilinear couplings of scalar components of chiral supermultiplets

In particular, using the field content described above for the MSSM, the softly-broken portion of the MSSM Lagrangian can be written

$$\mathcal{L}_{\text{soft}} = -\frac{1}{2} \left(M_3 \tilde{g} \tilde{g} + M_2 \tilde{W} \tilde{W} + M_1 \tilde{B} \tilde{B} + c.c. \right) \quad (3.11)$$

$$- \left(\tilde{u} a_u \tilde{Q} H_u - \tilde{d} a_d \tilde{Q} H_d - \tilde{e} a_e \tilde{L} H_d + c.c. \right) \quad (3.12)$$

$$- \tilde{Q}^\dagger m_Q^2 \tilde{Q} - \tilde{L}^\dagger m_L^2 \tilde{L} - \tilde{u} m_u^2 \tilde{u}^\dagger - \tilde{d} m_d^2 \tilde{d}^\dagger - \tilde{e} m_e^2 \tilde{e}^\dagger \quad (3.13)$$

$$- m_{H_u}^2 H_u^* H_u - m_{H_d}^2 H_d^* H_d - (b H_u H_d + cc). \quad (3.14)$$

433 where we have introduced the following notations :

- 434 1. M_3, M_2, M_1 are the gluino, wino, and bino masses.
- 435 2. a_u, a_d, a_e are complex 3×3 matrices in family space.
- 436 3. $m_Q^2, m_u^2, m_d^2, m_L^2, m_e^2$ are hermitian 3×3 matrices in family space.

437 4. $m_{H_u}^2, m_{H_d}^2, b$ are the SUSY-breaking contributions to the Higgs potential.

438 We have written matrix terms without any sort of additional notational decoration
 439 to indicate their matrix nature, and we now show why. The first term 1 are
 440 straightforward; these are just the straightforward mass terms for these fields. There
 441 are strong constraints on the off-diagonal terms for the matrices of 2 [74, 75]; for
 442 simplicity, we will assume that each $a_i, i = u, d, e$ is proportional to the Yukawa
 443 coupling matrix : $a_i = A_{i0} y_i$. The matrices in ?? can be similarly constrained by
 444 experiments [68, 75–82] Finally, we assume that the elements 4 contributing to the
 445 Higgs potential as well as all of the 1 terms must be real, which limits the possible
 446 CP-violating interactions to those of the Standard Model. We thus only consider
 447 flavor-blind, CP-conserving interactions within the MSSM.

The important mixing for mass and gauge interaction eigenstates in the MSSM occurs within electroweak sector, in a process akin to EWSB in the Standard Model. The neutral portions of the Higgsinos doublets and the neutral gauginos ($\tilde{H}_u^0, \tilde{H}_d^0, \tilde{B}^0, \tilde{W}^0$ of the gauge interaction basis mix to form what are known as the *neutralinos* of mass basis :

$$M_{\tilde{\chi}} = \begin{pmatrix} M_1 & 0 & -c_\beta s_W m_Z & s_\beta s_W m_Z \\ 0 & M_2 & c_\beta c_W m_Z & -s_\beta c_W m_Z \\ -c_\beta s_W m_Z & c_\beta c_W m_Z & 0 & -\mu \\ s_\beta s_W m_Z & -s_\beta c_W m_Z & -\mu & 0 \end{pmatrix} \quad (3.15)$$

448 where $s(c)$ are the sine and cosine of angles related to EWSB, which introduced
 449 masses to the gauginos and higgsinos. Diagonalization of this matrix gives the four
 450 neutralino mass states, listed without loss of generality in order of increasing mass :
 451 $\chi_{1,2,3,4}^0$.

452 The neutralinos, especially the lightest neutralino $\tilde{\chi}_1^0$, are important ingredients
 453 in SUSY phenomenology.

454 The same process can be done for the electrically charged gauginos with
 455 the charged portions of the Higgsino doublets along with the charged winos
 456 $(\tilde{H}_u^+, \tilde{H}_d^+, \tilde{W}^+, \tilde{W}^-)$. This leads to the *charginos*, again in order of increasing mass
 457 : $\tilde{\chi}_{1,2}^\pm$.

458 3.3 Phenomenology

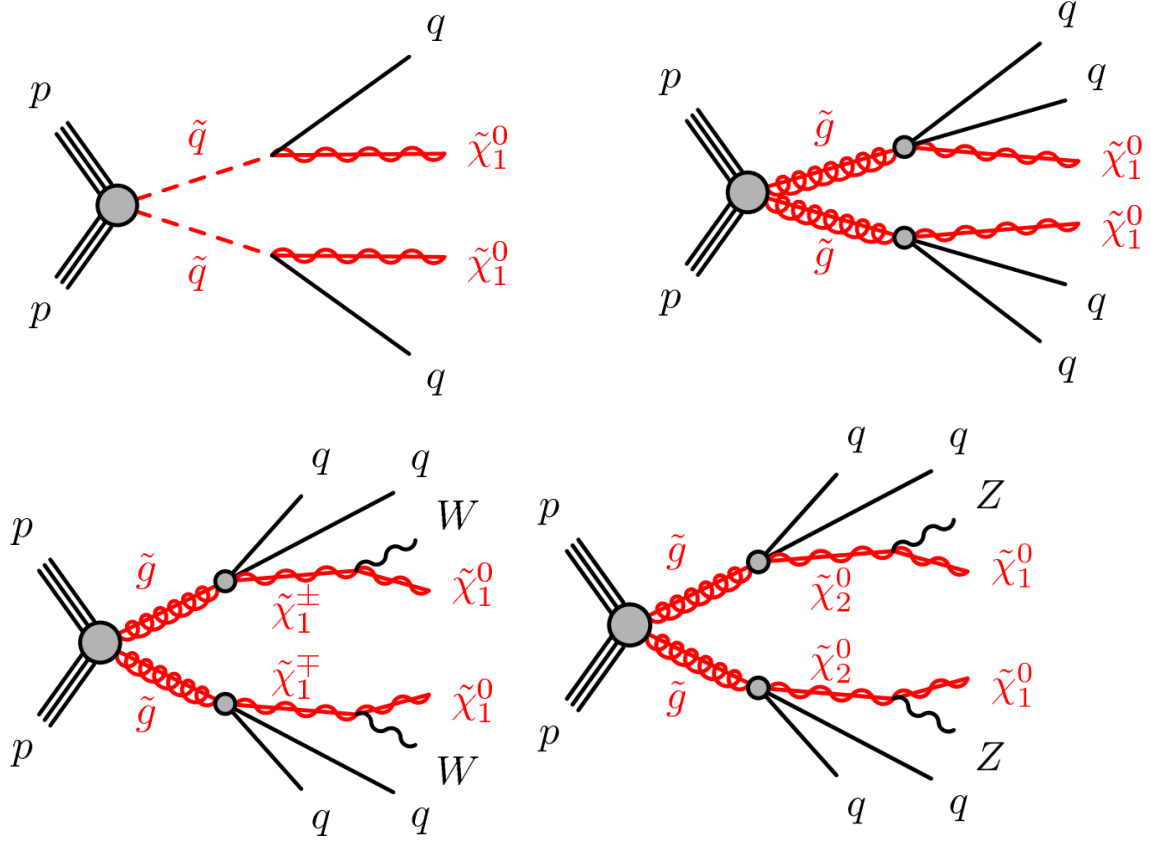
459 We are finally at the point where we can discuss the phenomenology of the MSSM,
 460 in particular as it manifests itself at the energy scales of the LHC.

461 As noted above in Sec.3.2, the assumption of R -parity has important conse-
 462 quences for MSSM phenomenology. The SM particles have $R = 1$, while the sparticles
 463 all have $R = -1$. Simply, this is the “charge” of supersymmetry. Since the particles of
 464 LHC collisions (pp) have total incoming $R = 1$, we must expect that all sparticles will
 465 be produced in *pairs*. An additional consequence of this symmetry is the fact that the
 466 lightest supersymmetric particle (LSP) is *stable*. Off each branch of the Feynmann
 467 diagram shown in Fig., we have $R = -1$, and this can only decay to another sparticle
 468 and a SM particle. Once we reach the lightest sparticle in the decay, it is absolutely
 469 stable. This leads to the common signature E_T^{miss} for a generic SUSY signal.

470 For this thesis, we will be presenting an inclusive search for squarks and gluinos
 471 with zero leptons in the final state. This is a very interesting decay channel⁴, due
 472 to the high cross-sections of $\tilde{g}\tilde{g}$ and $\tilde{q}\tilde{q}$ decays, as can be seen in Fig.?? [83]. This
 473 is a direct consequence of the fact that these are the colored particles of the MSSM.
 474 Since the sparticles interact with the gauge groups of the SM in the same way as their
 475 SM partners, the colored sparticles, the squarks and gluinos, are produced and decay
 476 as governed by the color group $SU(3)_C$ with the strong coupling g_s . The digluino
 477 production is particularly copious, due to color factor corresponding to the color octet

⁴Prior to Run1, probably the most *most* interesting SUSY decay channel.

Figure 3.3: SUSY signals considered in this thesis



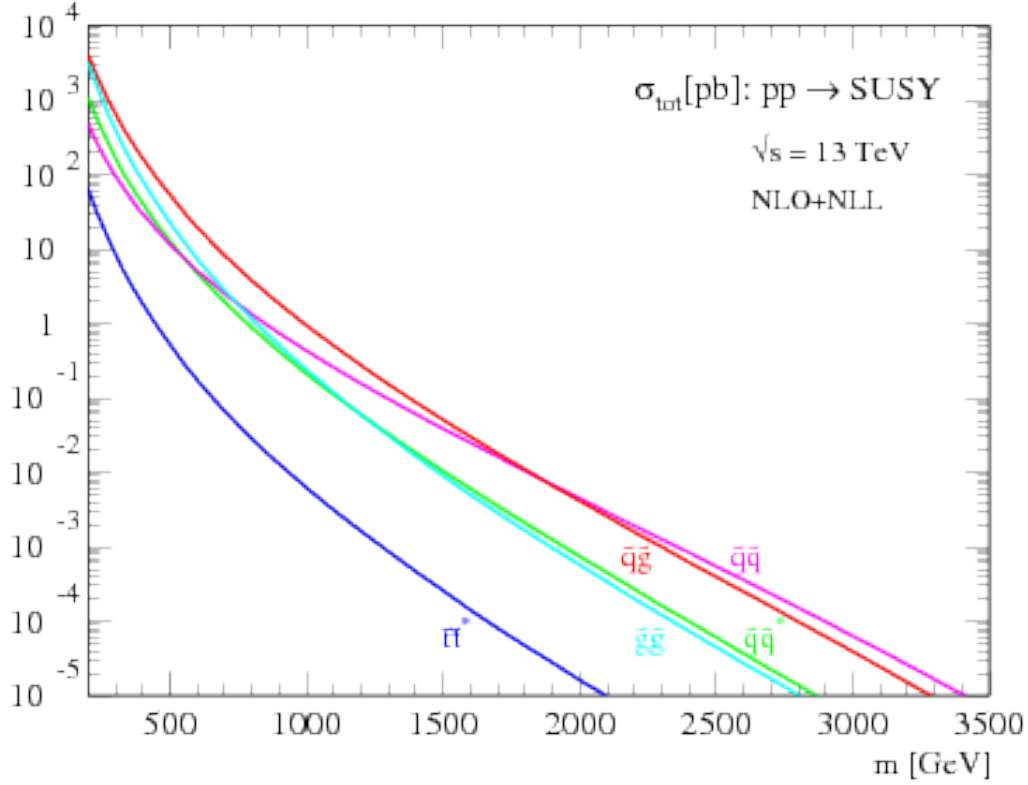
of $SU(3)C$.

In the case of disquark production, the most common decay mode of the squark in the MSSM is a decay directly to the LSP plus a single SM quark [15]. This means the basic search strategy of disquark production is two jets from the final state quarks, plus missing transverse energy for the LSPs. There are also cascade decays, the most common of which, and the only one considered in this thesis, is $\tilde{q} \rightarrow q\tilde{\chi}^\pm \rightarrow qW^\pm\chi^0$.

For digluino production, the most common decay is $\tilde{g} \rightarrow g\tilde{q}$, due to the large g_S coupling. The squark then decays as listed above. In this case, we generically search for four jets and missing transverse energy from the LSPs. We can also have the squark decay in association with a W^\pm or Z^0 ; in this thesis, we are interested in those cases where this vector boson goes hadronically.

In the context of experimental searches for SUSY, we often consider *simplified*

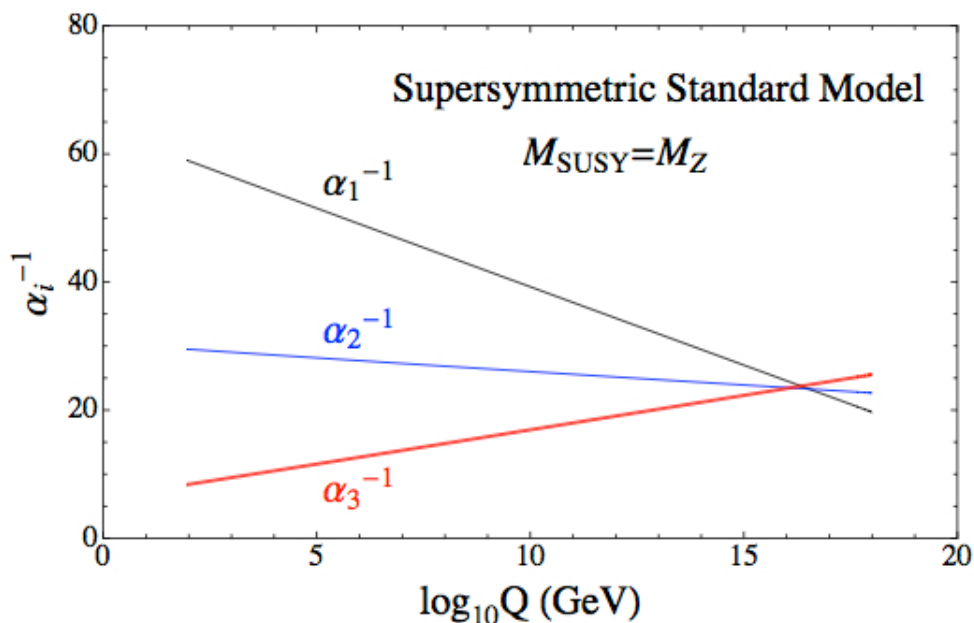
Figure 3.4: SUSY production cross-sections as a function of sparticle mass at $\sqrt{s} = 13$ TeV.



490 *models*. These models make certain assumptions which allow easy comparisons of
 491 results by theorists and rival experimentalists. In the context of this thesis, the
 492 simplified models will make assumptions about the branching ratios described in the
 493 preceding paragraphs. In particular, we will often choose a model where the decay of
 494 interest occurs with 100% branching ratio. This is entirely for ease of interpretation
 495 by other physicists⁵, but it is important to recognize that these are more a useful
 496 comparison tool, especially with limits, than a strict statement about the potential
 497 masses of sought-after beyond the Standard Model particle.

⁵In the author's opinion, this often leads to more confusion than comprehension. We will revisit the shortcomings of simplified models in the Conclusion to this thesis.

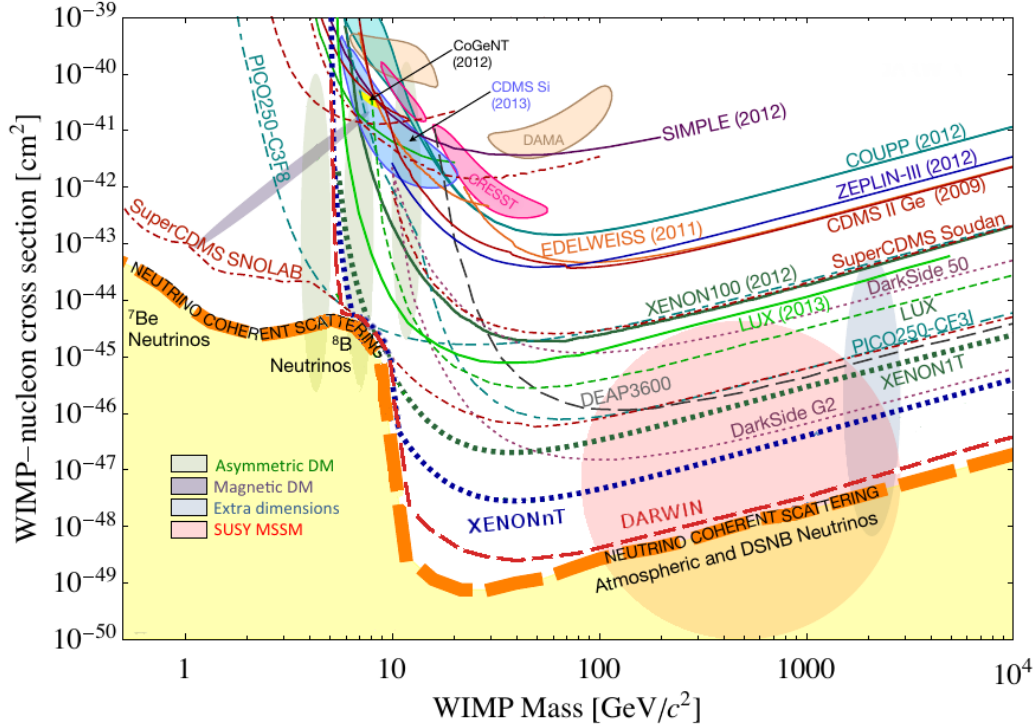
Figure 3.6: The running of Standard Model gauge couplings; compare to Fig.2.4. The MSSM gauge couplings nearly intersect at high energies.



513 Dark matter

514 As we discussed previously, the lack of any dark matter candidate in the Standard
515 Model naturally leads to beyond the Standard Model theories. In the Standard Model,
516 there is a natural dark matter candidate in the lightest supersymmetric particle[15]
517 The LSP would in dark matter experiments be called a *weakly-interacting massive*
518 *particle* (WIMP), which is a type of cold dark matter [22, 84]. These WIMPS would
519 only interact through the weak force and gravity, which is exactly as a model like the
520 MSSM predicts for the neutralino. In Fig.3.7, we can see the current WIMP exclusions
521 for a given mass. The range of allowed masses which have not been excluded for LSPs
522 and WIMPs have significant overlap. This provides additional motivation outside of
523 the context of theoretical details.

Figure 3.7: WIMP exclusions from direct dark matter detection experiments.



3.5 Conclusions

Supersymmetry is the most well-motivated theory for physics beyond the Standard Model. It provides a solution to the hierarchy problem, leads to gauge coupling unification, and provides a dark matter candidate consistent with galactic rotation curves. As noted in this chapter, due to the LSPs in the final state, most SUSY searches require a significant amount of missing transverse energy in combination with jets of high transverse momentum. However, there is some opportunity to do better than this, especially in final states where one has two weakly-interacting LSPs on opposite sides of some potentially complicated decay tree. We will see how this is done in Ch.??.

534

Chapter 4

535

The Large Hadron Collider

536 Here you can write some introductory remarks about your chapter. I like to give each
537 sentence its own line.

538 When you need a new paragraph, just skip an extra line.

539 **4.1 Magnets**

540 By using the asterisk to start a new section, I keep the section from appearing in the
541 table of contents. If you want your sections to be numbered and to appear in the
542 table of contents, remove the asterisk.

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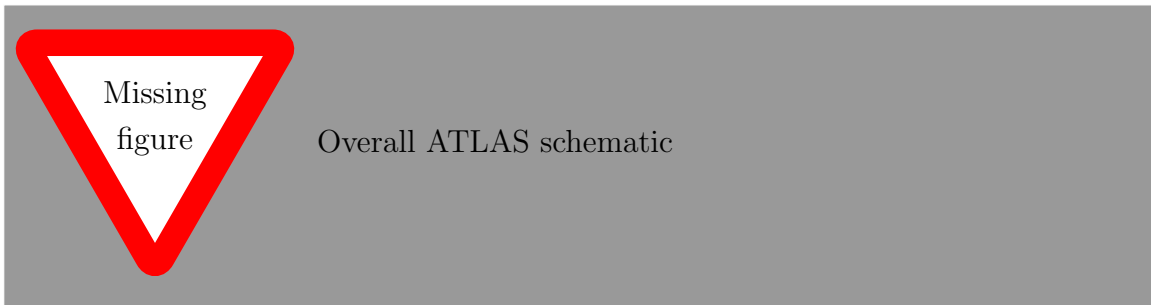
Chapter 5

544

The ATLAS detector

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546 sentence its own line.

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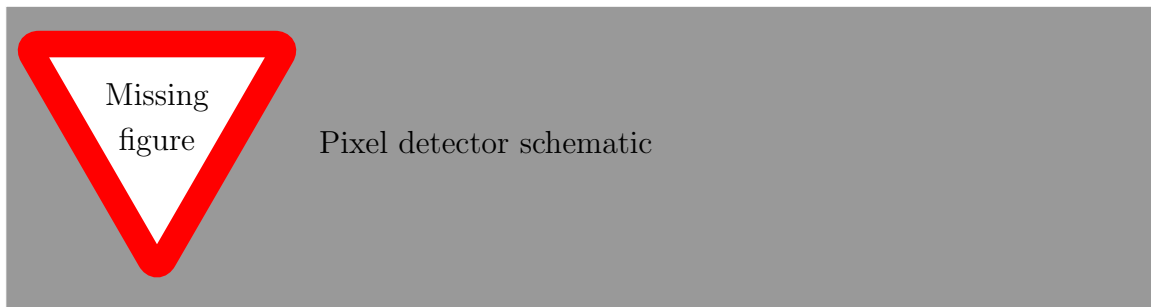
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5.1 Inner Detector

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552 table of contents. If you want your sections to be numbered and to appear in the
553 table of contents, remove the asterisk.

554 **Pixel Detector**

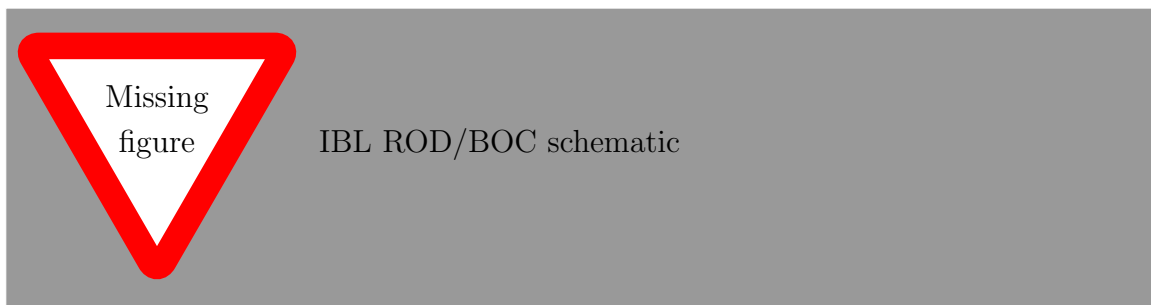


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557 **Insertable B-Layer**

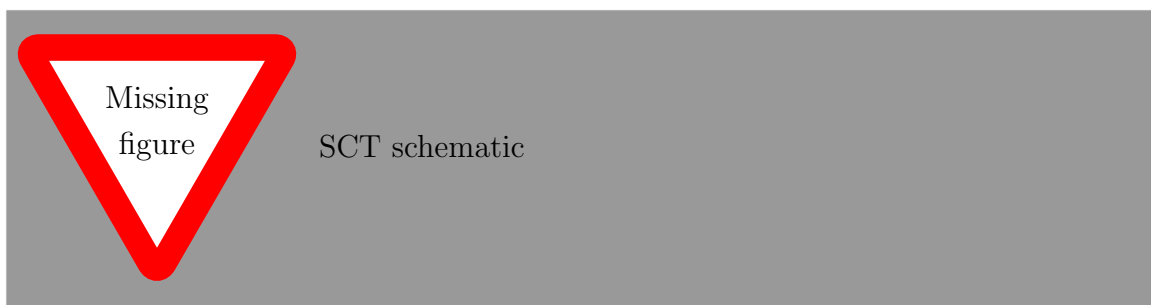
558 Qualification task, so add a bit more.



559

560

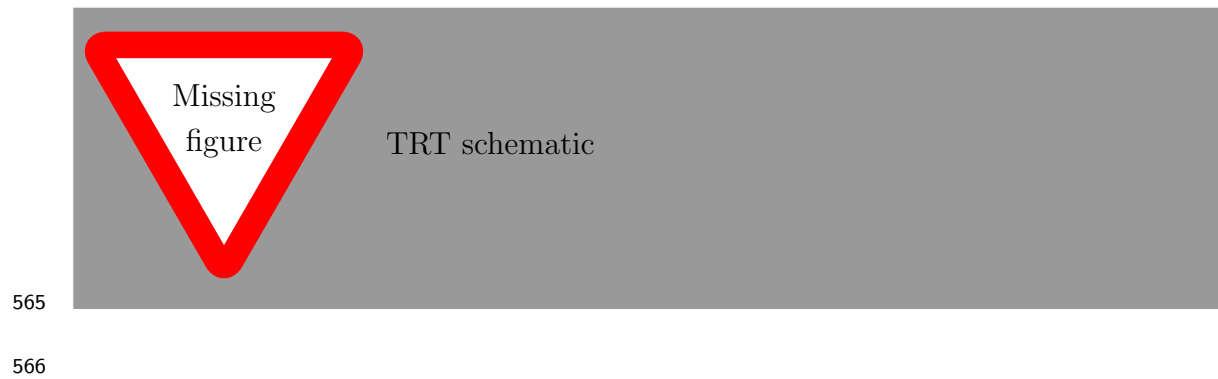
561 **Semiconductor Tracker**



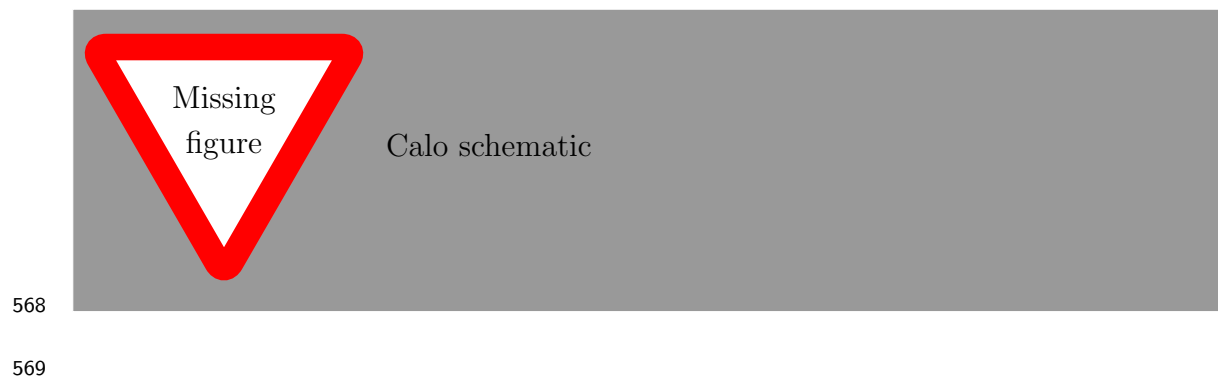
562

563

564 **Transition Radiation Tracker**



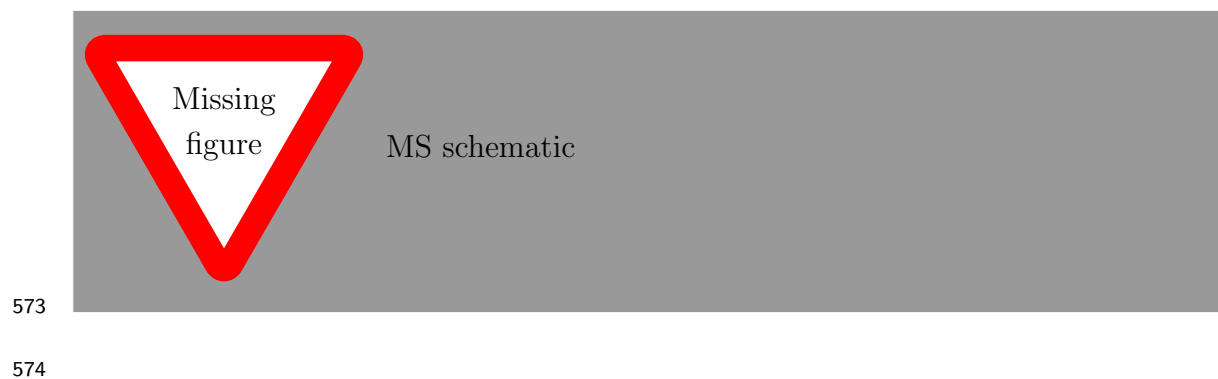
567 **5.2 Calorimeter**



570 **Electromagnetic Calorimeter**

571 **Hadronic Calorimeter**

572 **5.3 Muon Spectrometer**



The Recursive Jigsaw Technique

577 Here you can write some introductory remarks about your chapter. I like to give each
578 sentence its own line.

579 When you need a new paragraph, just skip an extra line.

580 **6.1 Razor variables**

581 By using the asterisk to start a new section, I keep the section from appearing in the
582 table of contents. If you want your sections to be numbered and to appear in the
583 table of contents, remove the asterisk.

584 **6.2 SuperRazor variables**

585 **6.3 The Recursive Jigsaw Technique**

586 **6.4 Variables used in the search for zero lepton**

587 **SUSY**

Title of Chapter 1

590

Chapter 8

591

Title of Chapter 1

592 Here you can write some introductory remarks about your chapter. I like to give each
593 sentence its own line.

594 When you need a new paragraph, just skip an extra line.

595 **8.1 Object reconstruction**

596 **Photons, Muons, and Electrons**

597 **Jets**

598 **Missing transverse momentum**

599 Probably longer, show some plots from the PUB note that we worked on

600 **8.2 Signal regions**

601 **Gluino signal regions**

602 **Squark signal regions**

603 **Compressed signal regions**

604 **8.3 Background estimation**

605 **Z $\nu\nu$**

606 **W $e\nu$**

607 **$t\bar{t}$**

608

Chapter 9

609

Title of Chapter 1

610 Here you can write some introductory remarks about your chapter. I like to give each
611 sentence its own line.

612 When you need a new paragraph, just skip an extra line.

613 **9.1 Statistical Analysis**

614 maybe to be moved to an appendix

615 **9.2 Signal Region distributions**

616 **9.3 Pull Plots**

617 **9.4 Systematic Uncertainties**

618 **9.5 Exclusion plots**

619

Conclusion

620 Here you can write some introductory remarks about your chapter. I like to give each
621 sentence its own line.

622 When you need a new paragraph, just skip an extra line.

623 **9.6 New Section**

624 By using the asterisk to start a new section, I keep the section from appearing in the
625 table of contents. If you want your sections to be numbered and to appear in the
626 table of contents, remove the asterisk.

Bibliography

- [1] O. Perdereau, *Planck 2015 cosmological results*,
AIP Conf. Proc. **1743** (2016) p. 050014.
- [2] N. Aghanim et al.,
*Planck 2016 intermediate results. LI. Features in the cosmic microwave
background temperature power spectrum and shifts in cosmological parameters*
(2016), arXiv: [1608.02487 \[astro-ph.CO\]](#).
- [3] J. S. Schwinger,
On Quantum electrodynamics and the magnetic moment of the electron,
Phys. Rev. **73** (1948) p. 416.
- [4] S. Laporta and E. Remiddi,
The Analytical value of the electron ($g-2$) at order α^3 in QED,
Phys. Lett. **B379** (1996) p. 283, arXiv: [hep-ph/9602417 \[hep-ph\]](#).
- [5] S. Schael et al., *Precision electroweak measurements on the Z resonance*,
Phys. Rept. **427** (2006) p. 257, arXiv: [hep-ex/0509008 \[hep-ex\]](#).
- [6] S. L. Glashow, *Partial Symmetries of Weak Interactions*,
Nucl. Phys. **22** (1961) p. 579.
- [7] S. Weinberg, *A Model of Leptons*, Phys. Rev. Lett. **19** (1967) p. 1264.
- [8] A. Salam, *Weak and Electromagnetic Interactions*,
Conf. Proc. **C680519** (1968) p. 367.
- [9] M. Gell-Mann, *A Schematic Model of Baryons and Mesons*,
Phys. Lett. **8** (1964) p. 214.
- [10] G. Zweig, “An SU(3) model for strong interaction symmetry and its breaking.
Version 2,” *DEVELOPMENTS IN THE QUARK THEORY OF HADRONS*.
VOL. 1. 1964 - 1978, ed. by D. Lichtenberg and S. P. Rosen, 1964 p. 22,
URL: <http://inspirehep.net/record/4674/files/cern-th-412.pdf>.

- 653 [11] S. Weinberg, *Implications of Dynamical Symmetry Breaking*,
654 [Phys. Rev. **D13** \(1976\) p. 974.](#)
- 655 [12] S. Weinberg, *Implications of Dynamical Symmetry Breaking: An Addendum*,
656 [Phys. Rev. **D19** \(1979\) p. 1277.](#)
- 657 [13] E. Gildener, *Gauge Symmetry Hierarchies*, [Phys. Rev. **D14** \(1976\) p. 1667.](#)
- 658 [14] L. Susskind,
659 *Dynamics of Spontaneous Symmetry Breaking in the Weinberg-Salam Theory*,
660 [Phys. Rev. **D20** \(1979\) p. 2619.](#)
- 661 [15] S. P. Martin, “A Supersymmetry Primer,” 1997,
662 eprint: [arXiv:hep-ph/9709356](#).
- 663 [16] V. C. Rubin and W. K. Ford Jr., *Rotation of the Andromeda Nebula from a*
664 *Spectroscopic Survey of Emission Regions*, [Astrophys. J. **159** \(1970\) p. 379.](#)
- 665 [17] M. S. Roberts and R. N. Whitehurst,
666 “*The rotation curve and geometry of M31 at large galactocentric distances*,
667 *Astrophys. J.* **201** (1970) p. 327.
- 668 [18] V. C. Rubin, N. Thonnard, and W. K. Ford Jr.,
669 *Rotational properties of 21 SC galaxies with a large range of luminosities and*
670 *radii, from NGC 4605 /R = 4kpc/ to UGC 2885 /R = 122 kpc/*,
671 [Astrophys. J. **238** \(1980\) p. 471.](#)
- 672 [19] V. C. Rubin et al., *Rotation velocities of 16 SA galaxies and a comparison of*
673 *Sa, Sb, and SC rotation properties*, [Astrophys. J. **289** \(1985\) p. 81.](#)
- 674 [20] A. Bosma,
675 *21-cm line studies of spiral galaxies. 2. The distribution and kinematics of*
676 *neutral hydrogen in spiral galaxies of various morphological types.*,
677 [Astron. J. **86** \(1981\) p. 1825.](#)
- 678 [21] M. Persic, P. Salucci, and F. Stel, *The Universal rotation curve of spiral*
679 *galaxies: 1. The Dark matter connection*,
680 [Mon. Not. Roy. Astron. Soc. **281** \(1996\) p. 27,](#)
681 [arXiv: astro-ph/9506004 \[astro-ph\].](#)
- 682 [22] M. Lisanti, “Lectures on Dark Matter Physics,” 2016,
683 eprint: [arXiv:1603.03797](#).
- 684 [23] H. Miyazawa, *Baryon Number Changing Currents*,
685 [Prog. Theor. Phys. **36** \(1966\) p. 1266.](#)

- [24] J.-L. Gervais and B. Sakita, *Generalizations of dual models*,
Nucl. Phys. **B34** (1971) p. 477.
- [25] J.-L. Gervais and B. Sakita,
Field Theory Interpretation of Supergauges in Dual Models,
Nucl. Phys. **B34** (1971) p. 632.
- [26] Yu. A. Golfand and E. P. Likhtman, *Extension of the Algebra of Poincare
Group Generators and Violation of p Invariance*,
JETP Lett. **13** (1971) p. 323, [Pisma Zh. Eksp. Teor. Fiz.13,452(1971)].
- [27] A. Neveu and J. H. Schwarz, *Factorizable dual model of pions*,
Nucl. Phys. **B31** (1971) p. 86.
- [28] A. Neveu and J. H. Schwarz, *Quark Model of Dual Pions*,
Phys. Rev. **D4** (1971) p. 1109.
- [29] D. V. Volkov and V. P. Akulov, *Is the Neutrino a Goldstone Particle?*
Phys. Lett. **B46** (1973) p. 109.
- [30] J. Wess and B. Zumino,
A Lagrangian Model Invariant Under Supergauge Transformations,
Phys. Lett. **B49** (1974) p. 52.
- [31] A. Salam and J. A. Strathdee, *Supersymmetry and Nonabelian Gauges*,
Phys. Lett. **B51** (1974) p. 353.
- [32] S. Ferrara, J. Wess, and B. Zumino, *Supergauge Multiplets and Superfields*,
Phys. Lett. **B51** (1974) p. 239.
- [33] J. Wess and B. Zumino, *Supergauge Transformations in Four-Dimensions*,
Nucl. Phys. **B70** (1974) p. 39.
- [34] J. D. Lykken, “Introduction to supersymmetry,” *Fields, strings and duality.
Proceedings, Summer School, Theoretical Advanced Study Institute in
Elementary Particle Physics, TASI’96, Boulder, USA, June 2-28, 1996*, 1996
p. 85, arXiv: [hep-th/9612114](https://arxiv.org/abs/hep-th/9612114) [hep-th],
URL: http://lss.fnal.gov/cgi-bin/find_paper.pl?pub-96-445-T.
- [35] A. Kobakhidze, “Intro to SUSY,” 2016, URL:
<https://indico.cern.ch/event/443176/page/5225-pre-susy-programme>.
- [36] G. R. Farrar and P. Fayet, *Phenomenology of the Production, Decay, and
Detection of New Hadronic States Associated with Supersymmetry*,
Phys. Lett. **B76** (1978) p. 575.

- 719 [37] ATLAS Collaboration,
720 *Search for the electroweak production of supersymmetric particles in*
721 *$\sqrt{s} = 8$ TeV pp collisions with the ATLAS detector,*
722 *Phys. Rev. D* **93** (2016) p. 052002, arXiv: [1509.07152 \[hep-ex\]](#).
- 723 [38] ATLAS Collaboration, *Summary of the searches for squarks and gluinos using*
724 *$\sqrt{s} = 8$ TeV pp collisions with the ATLAS experiment at the LHC,*
725 *JHEP* **10** (2015) p. 054, arXiv: [1507.05525 \[hep-ex\]](#).
- 726 [39] ATLAS Collaboration, *ATLAS Run 1 searches for direct pair production of*
727 *third-generation squarks at the Large Hadron Collider,*
728 *Eur. Phys. J. C* **75** (2015) p. 510, arXiv: [1506.08616 \[hep-ex\]](#).
- 729 [40] CMS Collaboration,
730 *Search for supersymmetry with razor variables in pp collisions at $\sqrt{s} = 7$ TeV,*
731 *Phys. Rev. D* **90** (2014) p. 112001, arXiv: [1405.3961 \[hep-ex\]](#).
- 732 [41] CMS Collaboration, *Inclusive search for supersymmetry using razor variables*
733 *in pp collisions at $\sqrt{s} = 7$ TeV,* *Phys. Rev. Lett.* **111** (2013) p. 081802,
734 arXiv: [1212.6961 \[hep-ex\]](#).
- 735 [42] CMS Collaboration, *Search for Supersymmetry in pp Collisions at 7 TeV in*
736 *Events with Jets and Missing Transverse Energy,*
737 *Phys. Lett. B* **698** (2011) p. 196, arXiv: [1101.1628 \[hep-ex\]](#).
- 738 [43] CMS Collaboration, *Search for Supersymmetry at the LHC in Events with*
739 *Jets and Missing Transverse Energy,* *Phys. Rev. Lett.* **107** (2011) p. 221804,
740 arXiv: [1109.2352 \[hep-ex\]](#).
- 741 [44] CMS Collaboration, *Search for supersymmetry in hadronic final states using*
742 *M_{T2} in pp collisions at $\sqrt{s} = 7$ TeV,* *JHEP* **10** (2012) p. 018,
743 arXiv: [1207.1798 \[hep-ex\]](#).
- 744 [45] CMS Collaboration, *Searches for supersymmetry using the M_{T2} variable in*
745 *hadronic events produced in pp collisions at 8 TeV,* *JHEP* **05** (2015) p. 078,
746 arXiv: [1502.04358 \[hep-ex\]](#).
- 747 [46] CMS Collaboration, *Search for new physics with the M_{T2} variable in all-jets*
748 *final states produced in pp collisions at $\sqrt{s} = 13$ TeV* (2016),
749 arXiv: [1603.04053 \[hep-ex\]](#).
- 750 [47] ATLAS Collaboration, *Multi-channel search for squarks and gluinos in*
751 *$\sqrt{s} = 7$ TeV pp collisions with the ATLAS detector at the LHC,*
752 *Eur. Phys. J. C* **73** (2013) p. 2362, arXiv: [1212.6149 \[hep-ex\]](#).

- [48] Y. Grossman, “Introduction to the SM,” 2016, URL: <https://indico.fnal.gov/sessionDisplay.py?sessionId=3&confId=11505#20160811>.
- [49] ().
- [50] P. W. Higgs, *Broken Symmetries and the Masses of Gauge Bosons*, *Phys. Rev. Lett.* **13** (1964) p. 508.
- [51] ATLAS Collaboration, *Observation of a new particle in the search for the Standard Model Higgs boson with the ATLAS detector at the LHC*, *Phys. Lett. B* **716** (2012) p. 1, arXiv: [1207.7214 \[hep-ex\]](#).
- [52] CMS Collaboration, *Observation of a new boson at a mass of 125 GeV with the CMS experiment at the LHC*, *Phys. Lett. B* **716** (2012) p. 30, arXiv: [1207.7235 \[hep-ex\]](#).
- [53] A. Chodos et al., *A New Extended Model of Hadrons*, *Phys. Rev.* **D9** (1974) p. 3471.
- [54] A. Chodos et al., *Baryon Structure in the Bag Theory*, *Phys. Rev.* **D10** (1974) p. 2599.
- [55] J. C. Collins, D. E. Soper, and G. F. Sterman, *Factorization of Hard Processes in QCD*, *Adv. Ser. Direct. High Energy Phys.* **5** (1989) p. 1, arXiv: [hep-ph/0409313 \[hep-ph\]](#).
- [56] K. A. Olive et al., *Review of Particle Physics*, *Chin. Phys.* **C38** (2014) p. 090001.
- [57] N. Cabibbo, *Unitary Symmetry and Leptonic Decays*, *Phys. Rev. Lett.* **10** (1963) p. 531, [648(1963)].
- [58] M. Kobayashi and T. Maskawa, *CP Violation in the Renormalizable Theory of Weak Interaction*, *Prog. Theor. Phys.* **49** (1973) p. 652.
- [59] W. F. L. Hollik, *Radiative Corrections in the Standard Model and their Role for Precision Tests of the Electroweak Theory*, *Fortsch. Phys.* **38** (1990) p. 165.
- [60] D. Yu. Bardin et al., *ELECTROWEAK RADIATIVE CORRECTIONS TO DEEP INELASTIC SCATTERING AT HERA! CHARGED CURRENT SCATTERING*, *Z. Phys.* **C44** (1989) p. 149.

- 786 [61] D. C. Kennedy et al., *Electroweak Cross-Sections and Asymmetries at the Z0*,
787 *Nucl. Phys.* **B321** (1989) p. 83.
- 788 [62] A. Sirlin, *Radiative Corrections in the $SU(2)_L \times U(1)$ Theory: A Simple*
789 *Renormalization Framework*, *Phys. Rev.* **D22** (1980) p. 971.
- 790 [63] S. Fanchiotti, B. A. Kniehl, and A. Sirlin,
791 *Incorporation of QCD effects in basic corrections of the electroweak theory*,
792 *Phys. Rev.* **D48** (1993) p. 307, arXiv: [hep-ph/9212285](#) [[hep-ph](#)].
- 793 [64] C. Quigg, “Cosmic Neutrinos,” *Proceedings, 35th SLAC Summer Institute on*
794 *Particle Physics: Dark matter: From the cosmos to the Laboratory (SSI 2007):*
795 *Menlo Park, California, July 30- August 10, 2007*, 2008,
796 arXiv: [0802.0013](#) [[hep-ph](#)],
797 URL: http://lss.fnal.gov/cgi-bin/find_paper.pl?conf-07-417.
- 798 [65] S. R. Coleman and J. Mandula, *All Possible Symmetries of the S Matrix*,
799 *Phys. Rev.* **159** (1967) p. 1251.
- 800 [66] R. Haag, J. T. Lopuszanski, and M. Sohnius,
801 *All Possible Generators of Supersymmetries of the s Matrix*,
802 *Nucl. Phys.* **B88** (1975) p. 257.
- 803 [67] A. Salam and J. A. Strathdee, *On Superfields and Fermi-Bose Symmetry*,
804 *Phys. Rev.* **D11** (1975) p. 1521.
- 805 [68] S. Dimopoulos and H. Georgi, *Softly Broken Supersymmetry and $SU(5)$* ,
806 *Nucl. Phys.* **B193** (1981) p. 150.
- 807 [69] S. Dimopoulos, S. Raby, and F. Wilczek,
808 *Supersymmetry and the Scale of Unification*, *Phys. Rev.* **D24** (1981) p. 1681.
- 809 [70] L. E. Ibanez and G. G. Ross,
810 *Low-Energy Predictions in Supersymmetric Grand Unified Theories*,
811 *Phys. Lett.* **B105** (1981) p. 439.
- 812 [71] W. J. Marciano and G. Senjanovic,
813 *Predictions of Supersymmetric Grand Unified Theories*,
814 *Phys. Rev.* **D25** (1982) p. 3092.
- 815 [72] L. Girardello and M. T. Grisaru, *Soft Breaking of Supersymmetry*,
816 *Nucl. Phys.* **B194** (1982) p. 65.

- [73] D. J. H. Chung et al.,
The Soft supersymmetry breaking Lagrangian: Theory and applications,
 Phys. Rept. **407** (2005) p. 1, arXiv: [hep-ph/0312378](#) [hep-ph].
- [74] J. Hisano et al., *Lepton flavor violation in the supersymmetric standard model with seesaw induced neutrino masses*, Phys. Lett. **B357** (1995) p. 579,
 arXiv: [hep-ph/9501407](#) [hep-ph].
- [75] F. Gabbiani et al., *A Complete analysis of FCNC and CP constraints in general SUSY extensions of the standard model*,
 Nucl. Phys. **B477** (1996) p. 321, arXiv: [hep-ph/9604387](#) [hep-ph].
- [76] F. Gabbiani and A. Masiero, *FCNC in Generalized Supersymmetric Theories*,
 Nucl. Phys. **B322** (1989) p. 235.
- [77] J. S. Hagelin, S. Kelley, and T. Tanaka, *Supersymmetric flavor changing neutral currents: Exact amplitudes and phenomenological analysis*,
 Nucl. Phys. **B415** (1994) p. 293.
- [78] J. S. Hagelin, S. Kelley, and V. Ziegler, *Using gauge coupling unification and proton decay to test minimal supersymmetric SU(5)*,
 Phys. Lett. **B342** (1995) p. 145, arXiv: [hep-ph/9406366](#) [hep-ph].
- [79] D. Choudhury et al.,
Constraints on nonuniversal soft terms from flavor changing neutral currents,
 Phys. Lett. **B342** (1995) p. 180, arXiv: [hep-ph/9408275](#) [hep-ph].
- [80] R. Barbieri and L. J. Hall, *Signals for supersymmetric unification*,
 Phys. Lett. **B338** (1994) p. 212, arXiv: [hep-ph/9408406](#) [hep-ph].
- [81] B. de Carlos, J. A. Casas, and J. M. Moreno,
Constraints on supersymmetric theories from $\mu \rightarrow e \gamma$,
 Phys. Rev. **D53** (1996) p. 6398, arXiv: [hep-ph/9507377](#) [hep-ph].
- [82] J. A. Casas and S. Dimopoulos,
Stability bounds on flavor violating trilinear soft terms in the MSSM,
 Phys. Lett. **B387** (1996) p. 107, arXiv: [hep-ph/9606237](#) [hep-ph].
- [83] C. Borschensky et al., *Squark and gluino production cross sections in pp collisions at $\sqrt{s} = 13, 14, 33$ and 100 TeV*, Eur. Phys. J. **C74** (2014) p. 3174,
 arXiv: [1407.5066](#) [hep-ph].
- [84] M. Klasen, M. Pohl, and G. Sigl, *Indirect and direct search for dark matter*,
 Prog. Part. Nucl. Phys. **85** (2015) p. 1, arXiv: [1507.03800](#) [hep-ph].

The Standard Model

In this appendix, we provide a brief overview of the basic ingredients involved in construction of the Standard Model Lagrangian : quantum field theory, symmetries, and symmetry breaking.

Quantum Field Theory

In this section, we provide a brief overview of the necessary concepts from Quantum Field Theory (QFT).

In modern physics, the laws of nature are described by the “action” S , with the imposition of the principle of minimum action. The action is the integral over the spacetime coordinates of the “Lagrangian density” \mathcal{L} , or Lagrangian for short. The Lagrangian is a function of “fields”; general fields will be called $\phi(x^\mu)$, where the indices μ run over the space-time coordinates. We can then write the action S as

$$S = \int d^4x \mathcal{L}[\phi_i(x^\mu), \partial_\mu \phi_i(x^\mu)] \quad (9.1)$$

where we have an additional summation over i (of the different fields). Generally, we impose the following constraints on the Lagrangian :

1. Translational invariance - The Lagrangian is only a function of the fields ϕ and their derivatives $\partial_\mu \phi$
2. Locality - The Lagrangian is only a function of one point x_μ in spacetime.

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- 868 3. Reality condition - The Lagrangian is real to conserve probability.
- 869 4. Lorentz invariance - The Lagrangian is invariant under the Poincaré group of
870 spacetime.
- 871 5. Analyticity - The Lagrangian is an analytical function of the fields; this is to
872 allow the use of perturbation theory.
- 873 6. Invariance and Naturalness - The Lagrangian is invariant under some internal
874 symmetry groups; in fact, the Lagrangian will have *all* terms allowed by the
875 imposed symmetry groups.
- 876 7. Renormalizability - The Lagrangian will be renormalizable - in practice, this
877 means there will not be terms with more than power 4 in the fields.

878 The key item from the point of view of this thesis is that of “Invariance and
879 Natural”. We impose a set of “symmetries” and then our Lagrangian is the most
880 general which is allowed by those symmetries.

881 Symmetries

882 Symmetries can be seen as the fundamental guiding concept of modern physics.
883 Symmetries are described by “groups”. . To illustrate the importance of symmetries
884 and their mathematical description, groups, we start here with two of the simplest
885 and most useful examples : \mathbb{Z}_2 and $U(1)$.

886 \mathbb{Z}_2 symmetry

887 \mathbb{Z}_2 symmetry is the simplest example of a “discrete” symmetry. Consider the most
888 general Lagrangian of a single real scalar field $\phi(x_\mu)$

$$\mathcal{L}_\phi = \frac{1}{2}\partial_\mu\phi\partial^\mu\phi - \frac{m^2}{2}\phi^2 - \frac{\mu}{2\sqrt{2}}\phi^3 - \lambda\phi^4 \quad (9.2)$$

Now we *impose* the symmetry

$$\mathcal{L}(\phi) = \mathcal{L}(-\phi) \quad (9.3)$$

889 This has the effect of restricting the allowed terms of the Lagrangian. In particular,
 890 we can see the term $\phi^3 \rightarrow -\phi^3$ under the symmetry transformation, and thus must
 891 be disallowed by this symmetry. This means under the imposition of this particular
 892 symmetry, our Lagrangian should be rewritten as

$$\mathcal{L}_\phi = \frac{1}{2}\partial_\mu\phi\partial^\mu\phi - \frac{m^2}{2}\phi^2 - \lambda\phi^4 \quad (9.4)$$

893 The effect of this symmetry is that the total number of ϕ particles can only change
 894 by even numbers, since the only interaction term $\lambda\phi^4$ is an even power of the field.
 895 This symmetry is often imposed in supersymmetric theories, as we will see in Chapter
 896 3.

897 **$U(1)$ symmetry**

898 $U(1)$ is the simplest example of a continuous (or *Lie*) group. Now consider a theory
 899 with a single complex scalar field $\phi = \text{Re } \phi + i \text{Im } \phi$

$$\mathcal{L}_\phi = \delta_{i,j} \frac{1}{2} \partial_\mu \phi_i \partial^\mu \phi_j - \frac{m^2}{2} \phi_i \phi_j - \frac{\mu}{2\sqrt{2}} \phi_i \phi_j \phi_k - \lambda \phi_i \phi_j \phi_k \phi_l \quad (9.5)$$

900 where $i, j, k, l = \text{Re}, \text{Im}$. In this case, we impose the following $U(1)$ symmetry
 901 : $\phi \rightarrow e^{i\theta} \phi, \phi^* \rightarrow e^{-i\theta} \phi^*$. We see immediately that this again disallows the third-order
 902 terms, and we can write a theory of a complex scalar field with $U(1)$ symmetry as

$$\mathcal{L}_\phi = \partial_\mu \phi \partial^\mu \phi^* - \frac{m^2}{2} \phi \phi^* - \lambda (\phi \phi^*)^2 \quad (9.6)$$

903 Local symmetries

904 The two examples considered above are “global” symmetries in the sense that the
905 symmetry transformation does not depend on the spacetime coordinate x_μ . We know
906 look at local symmetries; in this case, for example with a local $U(1)$ symmetry, the
907 transformation has the form $\phi(x_\mu) \rightarrow e^{i\theta(x_\mu)}\phi(x_\mu)$. These symmetries are also known
908 as “gauge” symmetries; all symmetries of the Standard Model are gauge symmetries.

There are wide-ranging consequences to the imposition of local symmetries. To begin, we note that the derivative terms of the Lagrangian 9.2 are *not* invariant under a local symmetry transformation

$$\partial_\mu \phi(x_\mu) \rightarrow \partial_\mu (e^{i\theta(x_\mu)} \phi(x_\mu)) = (1 + i\partial_\mu \theta(x_\mu)) e^{i\theta(x_\mu)} \phi(x_\mu) \quad (9.7)$$

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909 This leads us to note that the kinetic terms of the Lagrangian are also not invariant
910 under a gauge symmetry. This would lead to a model with no dynamics, which is
911 clearly unsatisfactory.

912 Let us take inspiration from the case of global symmetries. We need to define a
913 so-called “covariant” derivative D^μ such that

$$D^\mu \phi \rightarrow e^{iq\theta(x_\mu)} D^\mu \phi \quad (9.8)$$

$$D^\mu \phi^* \rightarrow e^{-iq\theta(x_\mu)} D^\mu \phi^* \quad (9.9)$$

$$(9.10)$$

915 Since ϕ and ϕ^* transform with the opposite phase, this will lead to the invariance
916 of the Lagrangian under our local gauge transformation. This D^μ is of the following
917 form

$$D^\mu = \partial_\mu - igqA^\mu \quad (9.11)$$

918 where A^μ is a vector field we introduce with the transformation law

$$A^\mu \rightarrow A^\mu - \frac{1}{g} \partial_\mu \theta \quad (9.12)$$

919 and g is the coupling constant associated to vector field. This vector field A^μ is
920 also known as a “gauge” field.

921 Since we need to add all allowed terms to the Lagrangian, we define

$$F^{\mu\nu} = A^\mu A^\nu - A^\nu A^\mu \quad (9.13)$$

922 and then we must also add the kinetic term :

$$\mathcal{L}_{\text{gauge}} = -\frac{1}{4} F^{\mu\nu} F_{\mu\nu} \quad (9.14)$$

923 The most general renormalizable Lagrangian with fermion and scalar fields can
924 be written in the following form

$$\mathcal{L} = \mathcal{L}_{kin} + \mathcal{L}_\phi + \mathcal{L}_\psi + \mathcal{L}_{Yukawa} \quad (9.15)$$

925 Symmetry breaking and the Higgs mechanism

926 Here we view some examples of symmetry breaking. We investigate breaking of a
927 global $U(1)$ symmetry and a local $U(1)$ symmetry. The SM will break the electroweak
928 symmetry $SU(2) \times U(1)$, and in Chapter 3 we will see how supersymmetry must also
929 be broken.

930 There are two ideas of symmetry breaking

- 931 • Explicit symmetry breaking by a small parameter - in this case, we have a small
932 parameter which breaks an “approximate” symmetry of our Lagrangian. An
933 example would be the theory of the single scalar field [9.2](#), when $\mu \ll m^2$ and

934 $\mu \ll \lambda$. In this case, we can often ignore the small term when considering
 935 low-energy processes.

936 • Spontaneous symmetry breaking (SSB) - spontaneous symmetry breaking
 937 occurs when the Lagrangian is symmetric with respect to a given symmetry
 938 transformation, but the ground state of the theory is *not* symmetric with respect
 939 to that transformation. This can have some fascinating consequences, as we
 940 will see in the following examples

941 Symmetry breaking a

942 **U(1) global symmetry breaking**

Consider the theory of a complex scalar field under the $U(1)$ symmetry, or the transformation

$$\phi \rightarrow e^{i\theta} \phi \quad (9.16)$$

The Lagrangian for this theory is

$$\mathcal{L} = \partial^\mu \phi^\dagger \partial_\mu \phi + \frac{\mu^2}{2} \phi^\dagger \phi + \frac{\lambda}{4} (\phi^\dagger \phi)^2 \quad (9.17)$$

Let us write this theory in terms of two scalar fields, h and ξ : $\phi = (h + i\xi)/\sqrt{2}$.

The Lagrangian can then be written as

$$\mathcal{L} = \partial^\mu h \partial_\mu h + \partial^\mu \xi \partial_\mu \xi - \frac{\mu^2}{2} (h^2 + \xi^2) - \frac{\lambda}{4} (h^2 + \xi^2)^2 \quad (9.18)$$

First, note that the theory is only stable when $\lambda > 0$. To understand the effect of SSB, we now enforce that $\mu^2 < 0$, and define $v^2 = -\mu^2/\lambda$. We can then write the scalar potential of this theory as :

$$V(\phi) = \lambda(\phi^\dagger \phi - v^2/2)^2 \quad (9.19)$$

Minimizing this equation with respect to ϕ , we can see that the “vacuum expectation value” of the theory is

$$2 \langle \phi^\dagger \phi \rangle = \langle h^2 + \xi^2 \rangle = v^2 \quad (9.20)$$

943 We now reach the “breaking” point of this procedure. In the (h, ξ) plane, the
 944 minima form a circle of radius v . We are free to choose any of these minima to expand
 945 our Lagrangian around; the physics is not affected by this choice. For convenience,
 946 choose $\langle h \rangle = v, \langle \xi^2 \rangle = 0$.

Now, let us define $h' = h - v, \xi' = \xi$ with VEVs $\langle h' \rangle = 0, \langle \xi' \rangle = 0$. We can then write our spontaneously broken Lagrangian in the form

$$\mathcal{L} = \frac{1}{2} \partial_\mu h' \partial^\mu h' + \frac{1}{2} \partial_\mu \xi' \partial^\mu \xi' - \lambda v^2 h'^2 - \lambda v h' (h'^2 + \xi'^2) - \lambda (h'^2 + \xi'^2)^2 \quad (9.21)$$