

# PHYS 5260 HW4

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## Problem 1. Sakurai 2.21

Derive the normalization constant  $c_n$  by deriving the orthogonality relationship using generating functions. Start by working out the integral

$$I = \int_{-\infty}^{\infty} g(x, t)g(x, s)e^{-x^2} dx$$

and then consider the integral again with the generating functions in terms of series with Hermite polynomials.

Using the definition of generating function in (2.5.17a),  $g(x, t)$  is given by:

$$g(x, t) = e^{-t^2 + 2tx}$$

Substituting the generating function, the given integral yields:

$$\begin{aligned} \int_{-\infty}^{\infty} g(x, t)g(x, s)e^{-x^2} dx &= \int_{-\infty}^{\infty} e^{-t^2 + 2tx} e^{-s^2 + 2sx} e^{-x^2} dx \\ &= e^{2ts} \int_{-\infty}^{\infty} e^{-(x - (t+s))^2} dx \\ &= e^{2ts} \sqrt{\pi} \\ &= \sqrt{\pi} \sum_{n=0}^{\infty} \frac{2^n t^n s^n}{n!} \end{aligned} \tag{1}$$

The second equation holds since  $-t^2 + 2tx - s^2 + 2sx - x^2 = -(x^2 - 2(t+s)x + (t+s)^2) + 2ts$ .

On the other hand, using the definition of generating function in (2.5.17b),  $g(x, t)$  is given by:

$$g(x, t) = \sum_{n=0}^{\infty} H_n(x) \frac{t^n}{n!}$$

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\*L<sup>A</sup>T<sub>E</sub>X source code: <https://github.com/rstanuwijaya/hkust-advanced-qm/>

Substituting the generating function, the given integral yields:

$$\int_{-\infty}^{\infty} g(x, t)g(x, s)e^{-x^2} dx = \sum_{n=0}^{\infty} \sum_{m=0}^{\infty} \frac{t^n}{n!} \frac{s^m}{m!} \int_{-\infty}^{\infty} H_n(x)H_m(x)e^{-x^2} dx \quad (2)$$

Comparing the two results for  $m = n$ , we have:

$$\begin{aligned} \frac{t^n s^n}{(n!)^2} \int_{-\infty}^{\infty} H_n(x)H_n(x)e^{-x^2} dx &= 2^n \sqrt{\pi} \frac{t^n s^n}{n!} \\ \int_{-\infty}^{\infty} H_n(x)H_n(x)e^{-x^2} dx &= 2^n \sqrt{\pi} n! \end{aligned}$$

On the other hand, for  $m \neq n$ , the integral must be 0, i.e.:

$$\int_{-\infty}^{\infty} H_n(x)H_m(x)e^{-x^2} dx = 0$$

Recalling the definition of the wave function:

$$u_n(x) = c_n H_n \left( x \sqrt{\frac{m\omega}{\hbar}} \right) e^{-m\omega x^2/2\hbar}$$

We can now write the normalization condition as:

$$\begin{aligned} \int_{-\infty}^{\infty} u_n^*(x)u_n(x)dx &= |c_n|^2 \int_{-\infty}^{\infty} H_n^2 \left( x \sqrt{\frac{m\omega}{\hbar}} \right) e^{-m\omega x^2/\hbar} dx \\ &= |c_n|^2 2^n \sqrt{\pi} n! = 1 \end{aligned}$$

Which yields:

$$c_n = \left( \frac{m\omega}{2^{2n} \pi (n!)^2 \hbar} \right)^{1/4}$$

## Problem 2. Sakurai 2.24

Consider a particle in one dimension bound to a fixed centre by a  $\delta$ -function potential of the form

$$V(x) = -v_0\delta(x), \quad (v_0 \text{ real and positive})$$

Find the wave function and the binding energy of the ground state. Are there excited bound states?

The Schrodinger equation for the system is given by:

$$\left(-\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} - v_0\delta(x)\right) \psi(x) = E\psi(x)$$

For bound states, we apply  $E = -E_0 < 0$ , i.e. the wave function decays exponentially. The wave function can be written as:

$$\begin{aligned} \psi(x) &= Ce^{-\kappa|x|} \\ &= \begin{cases} Ce^{-\kappa x}, & x < 0 \\ Ce^{\kappa x}, & x > 0 \end{cases} \end{aligned}$$

where  $\kappa = ik = i\sqrt{2mE_0}/\hbar$ .

The second boundary condition usually can be obtained by matching the derivative of the wave function at  $x = 0$ . However, since there is a delta potential at the origin we cannot use this method. Instead, we integrate over a small distance  $\epsilon$  from the origin and match the wave function:

$$\begin{aligned} \frac{-\hbar^2}{2m} \int_{-\epsilon}^{\epsilon} \psi''(x) dx - v_0 \int_{-\epsilon}^{\epsilon} \delta(x) \psi(x) dx &= E \int_{-\epsilon}^{\epsilon} \psi(x) dx \\ \frac{-\hbar^2}{2m} (\psi'(\epsilon) - \psi'(-\epsilon)) - v_0 \psi(0) &= E(\psi(\epsilon) - \psi(-\epsilon)) \end{aligned}$$

Taking limit  $\epsilon \rightarrow 0$  on both sides, we get:

$$\frac{-\hbar^2}{2m} (-2\kappa)C - v_0C = 0 \iff \kappa = mv_0/\hbar^2$$

If we solve the Schrodinger equation for any  $x > 0$  or  $x < 0$ , we obtain the solution:

$$\frac{-\hbar^2}{2m} \frac{\partial^2 \psi(x)}{\partial x^2} = -E_0 \psi(x) \iff \psi(x) = Ce^{-\sqrt{2mE_0}/\hbar^2 |x|}$$

Thus, matching the two solutions, we can obtain:

$$\frac{mv_0}{\hbar^2} = \frac{\sqrt{2mE_0}}{\hbar} \iff E_0 = \frac{mv_0^2}{2\hbar^2}$$

Since this energy is unique, there is no excited state of this system.

### Problem 3. Sakurai 2.27

Derive an expression for the density of free-particle states in two dimensions, normalized with periodic boundary conditions inside a box of side length  $L$ . Your answer should be written as a function of  $k$  (or  $E$ ) times  $dE d\phi$ , where  $\phi$  is the polar angle that characterizes the momentum direction in two dimensions.

The wave function for a free particle in two dimensions is given by:

$$\psi(x, y) = \frac{1}{\sqrt{L^2}} e^{i(k_x x + k_y y)}$$

Where  $k_x = \frac{2\pi}{L} n_x$  and  $k_y = \frac{2\pi}{L} n_y$ .

The energy of the particle is given by:

$$E = \frac{p^2}{2m} = \frac{\hbar^2}{2m} (k_x^2 + k_y^2) = \frac{2\pi^2 \hbar^2}{mL^2} (n_x^2 + n_y^2)$$

Assume we do a transformation from a cartesian coordinate spanned by  $n_x, n_y$  to a polar coordinate spanned by  $n, \phi$ , then the following relation holds:

$$\begin{aligned} n^2 &= n_x^2 + n_y^2 \\ \tan \phi &= \frac{n_y}{n_x} \\ dN &= dn_x dn_y = n dn d\phi \end{aligned}$$

where  $dN$  is the density of states.

The energy relation can be rewritten in polar coordinate as:

$$dE = \frac{4\pi^2 \hbar^2}{mL^2} n dn$$

Finally, the density of states is given by:

$$dN = n dn d\phi = \frac{mL^2}{4\pi^2 \hbar^2} dE d\phi$$

## Problem 4. Sakurai 2.30

Using spherical coordinates, obtain an expression for  $\mathbf{j}$  for the ground and excited states of the hydrogen atom. Show, in particular for  $m_i \neq 0$  states, there is a circulating flux in the sense that  $\mathbf{j}$  is in the direction of increasing or decreasing  $\phi$  depending on whether  $m_i$  is positive or negative.

Recall the probability current is given by:

$$\mathbf{j} = \frac{1}{2m}(\psi^* \nabla \psi - \psi \nabla \psi^*) = \frac{\hbar}{m} \text{Im}(\psi^* \nabla \psi)$$

Where the  $\nabla$  operator in spherical coordinate is defined as:

$$\nabla = \frac{1}{r} \frac{\partial}{\partial r} + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} + \frac{1}{r^2 \sin^2 \theta} \frac{\partial}{\partial \phi}$$

The wave function of a hydrogen atom is given by:

$$\begin{aligned} \psi(r, \theta, \phi) &= C_{nlm} R_{nl}(r) Y_l^m(\theta, \phi) \\ &= C_{nlm} R_{nl}(r) P_l^m(\cos \theta) e^{im\phi} \end{aligned}$$

Where  $R_{nl}$  is the radial part of the wave function and  $Y_l^m$  is the spherical harmonic function. Note that  $C_{nlm}$ ,  $R_{nl}$  and  $Y_l^m$  are all real numbers.

Note that since we are only interested on the imaginary part of  $\psi^* \nabla \psi$ , we can only work on the  $\hat{\phi}$  part of the wave function. Thus, we can write:

$$\begin{aligned} \mathbf{j} &= \frac{\hbar}{m_e} \text{Im}(\psi^* \nabla \psi) \\ &= \frac{\hbar}{m_e} \text{Im}(\psi^* \left( \frac{\hat{\phi}}{r \sin \phi} \frac{\partial}{\partial \phi} \right) \psi) \\ &= \frac{\hbar}{m_e} \text{Im}(\psi^* \frac{im}{r \sin \phi} \psi) \hat{\phi} \\ &= \frac{\hbar}{m_e} \frac{m}{\sin \phi} |\psi|^2 \hat{\phi} \end{aligned}$$

Therefore, we show that the flux is in direction of  $\phi$ , depending of the magnetic quantum number  $m$ .

## Problem 5. Sakurai 2.33

The propagator in momentum space analogous to (2.6.26) is given by  $\langle \mathbf{p}'', t | \mathbf{p}', t_0 \rangle$ . Derive an explicit expression for  $\langle \mathbf{p}'', t | \mathbf{p}', t_0 \rangle$  for the free-particle case.

For a free particle, the hamiltonian is given by:

$$H = \frac{p^2}{2m}$$

Thus, the propagator is given by:

$$\begin{aligned}\langle \mathbf{p}'', t | \mathbf{p}', t_0 \rangle &= \langle \mathbf{p}'' | e^{-iHt} e^{iHt_0} | \mathbf{p}', t_0 \rangle \\ &= \exp\left(\frac{-i\mathbf{p}'^2}{2m\hbar}(t - t_0)\right) \langle \mathbf{p}'' | \mathbf{p}', t_0 \rangle \\ &= \exp\left(\frac{-i\mathbf{p}'^2}{2m\hbar}(t - t_0)\right) \delta(\mathbf{p}'' - \mathbf{p}')\end{aligned}$$

## Problem 6. Sakurai 2.36

Show that wave-mechanical approach to the gravity-induced problem discussed in Section 2.7 also leads to phase-difference expression (2.7.17).

In the experiment, the phase of the neutron wave function depends on the length of the path and the time it takes to travel the path i.e.,  $\phi \sim kx - \omega t$ . Thus, the phase difference of the two path is given by:

$$\begin{aligned}\phi_{BD} - \phi_{AC} &= \left( \frac{p_{BD} - p_{AC}}{\hbar} \right) l_1 - \omega \left( \frac{l_1}{v_{BD}} - \frac{l_1}{v_{AC}} \right) \\ &= \left( \frac{p_{BD} - p_{AC}}{\hbar} \right) l_1 \left( 1 + \frac{\hbar \omega m_n}{p_{AC} p_{BD}} \right) \\ &= \left( \frac{p_{BD} - p_{AC}}{\hbar} \right) l_1 \left( 1 + \frac{Em_n}{p^2} \right) \\ &= \left( \frac{p_{BD} - p_{AC}}{\hbar} \right) \frac{3}{2} l_1\end{aligned}$$

Since  $E = \hbar \omega = \frac{p^2}{2m}$ .

Then we can make the following approximation and apply the conservation law of energy:

$$\begin{aligned}p_{BD} - p_{AC} &\approx \left( \frac{p_{BD}^2}{2m} - \frac{p_{AC}^2}{2m} \right) \frac{2m}{p} \\ &= -mgz \frac{2m}{p} = -\frac{2m^2 g l_2 \sin \theta}{p}\end{aligned}$$

Therefore the phase difference is given by:

$$\begin{aligned}\phi_{BD} - \phi_{AC} &= \left( \frac{p_{BD} - p_{AC}}{\hbar} \right) \frac{3}{2} l_1 \\ &= -\frac{2m^2 g l_2 \sin \theta}{p \hbar} \frac{3}{2} l_1 \\ &= -\frac{3m^2 g \lambda}{2\pi \hbar^2} l_1 l_2 \sin \theta\end{aligned}$$