PHYS 5260 HW4

TANUWIJAYA, Randy Stefan * (20582731) rstanuwijaya@connect.ust.hk

Department of Physics - HKUST

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Problem 1. Sakurai 2.21

Derive the normalization constant c_n by deriving the orthogonality relatioonship using generating functions. Start by working out the integral

$$I = \int_{-\infty}^{\infty} g(x,t)g(x,s)e^{-x^2}dx$$

and then consider the integral again with the generating functions in terms of series with Hermite polynomials.

Using the definition of generating function in (2.5.17a), g(x,t) is given by:

$$g(x,t) = e^{-t^2 + 2tx}$$

Substituting the generating function, the given integral yields:

$$\int_{-\infty}^{\infty} g(x,t)g(x,s)e^{-x^{2}}dx = \int_{-\infty}^{\infty} e^{-t^{2}+2tx}e^{-s^{2}+2sx}e^{-x^{2}}dx$$

$$= e^{2ts} \int_{-\infty}^{\infty} e^{-(x-(t+s))^{2}}dx$$

$$= e^{2ts} \sqrt{\pi}$$

$$= \sqrt{\pi} \sum_{n=0}^{\infty} \frac{2^{n}t^{n}s^{n}}{n!}$$
(1)

The second equation holds since $-t^2 + 2tx - s^2 + 2sx - x^2 = -(x^2 - 2(t+s) + (t+s)^2) + 2ts$.

On the other hand, using the definition of generating function in (2.5.17b), g(x,t) is given by:

$$g(x,t) = \sum_{n=0}^{\infty} H_n(x) \frac{t^n}{n!}$$

^{*}IATEX source code: https://github.com/rstanuwijaya/hkust-advanced-qm/

Substituting the generating function, the given integral yields:

$$\int_{-\infty}^{\infty} g(x,t)g(x,s)e^{-x^2}dx = \sum_{n=0}^{\infty} \sum_{m=0}^{\infty} \frac{t^n}{n!} \frac{s^m}{m!} \int_{-\infty}^{\infty} H_n(x)H_m(x)e^{-x^2}dx$$
 (2)

Comparing the two results for m = n, we have:

$$\frac{t^n s^n}{(n!)^2} \int_{-\infty}^{\infty} H_n(x) H_n(x) e^{-x^2} dx = 2^n \sqrt{\pi} \frac{t^n s^n}{n!}$$
$$\int_{-\infty}^{\infty} H_n(x) H_n(x) e^{-x^2} dx = 2^n \sqrt{\pi} n!$$

On the other hand, for $m \neq n$, the integral must be 0, i.e.:

$$\int_{-\infty}^{\infty} H_n(x)H_m(x)e^{-x^2}dx = 0$$

Recalling the definition of the wave function:

$$u_n(x) = c_n H_n \left(x \sqrt{\frac{m\omega}{\hbar}} \right) e^{-m\omega x^2/2\hbar}$$

We can now write the normalization condition as:

$$\int_{-\infty}^{\infty} u_n^*(x)u_n(x)dx = |c_n|^2 \int_{-\infty}^{\infty} H_n^2 \left(x\sqrt{\frac{m\omega}{\hbar}}\right) e^{-m\omega x^2/\hbar} dx$$
$$= |c_n|^2 2^n \sqrt{\pi} n! = 1$$

Which yields:

$$c_n = \left(\frac{m\omega}{2^{2n}\pi(n!)^2\hbar}\right)^{1/4}$$

Problem 2. Sakurai 2.24

Consider a particle in one dimension bound to a fixed centre by a δ -function potential of the form

$$V(x) = -v_0 \delta(x)$$
, (v_0 real and positive)

Find the wave function and the binding energy of the ground state. Are there excited bound states?

The Schrodinger equation for the system is given by:

$$\left(-\frac{\hbar^2}{2m}\frac{\partial^2}{\partial x^2} - v_0\delta(x)\right)\psi(x) = E\psi(x)$$

For bound states, we apply $E = -E_0 < 0$, i.e. the wave function decays exponentially. The wave function can be written as:

$$\psi(x) = Ce^{-\kappa|x|}$$

$$= \begin{cases} Ce^{-\kappa x}, & x < 0 \\ Ce^{\kappa x}, & x > 0 \end{cases}$$

where $\kappa = ik = i\sqrt{2mE_0}/\hbar$.

The second boundary coindition usually can be obtained by matching the derivative of the wave function at x = 0. However, since there is a delta potential at the origin we cannot use this method. Instead, we integrate over a small distance ϵ from the origin and match the wave function:

$$\frac{-\hbar^2}{2m} \int_{-\epsilon}^{\epsilon} \psi''(x)dx - v_0 \int_{-\epsilon}^{\epsilon} \delta(x)\psi(x)dx = E \int_{-\epsilon}^{\epsilon} \psi(x)dx$$
$$\frac{-\hbar^2}{2m} (\psi'(\epsilon) - \psi'(-\epsilon)) - v_0 \psi(0) = E(\psi(\epsilon) - \psi(-\epsilon))$$

Taking limit $\epsilon \to 0$ on both sides, we get:

$$\frac{-\hbar^2}{2m}(-2\kappa)C - v_0C = 0 \iff \kappa = mv_0/\hbar^2$$

If we solve the Schrödinger equation for any x > 0 or x < 0, we obtain the solution:

$$\frac{-\hbar^2}{2m} \frac{\partial^2 \psi(x)}{\partial x^2} = -E_0 \psi(x) \iff \psi(x) = Ce^{-\sqrt{2mE_0/\hbar^2}|x|}$$

Thus, matching the two solutions, we can obtain:

$$\frac{mv_0}{\hbar^2} = \frac{\sqrt{2mE_0}}{\hbar} \iff E_0 = \frac{mv_0^2}{2\hbar^2}$$

Since this energy is unique, there is no excited state of this system.

Problem 3. Sakurai 2.27

Derive an expression for the density of free-particle states in two dimensions, normalized with periodic boundary conditions inside a box of side length L. Your answer should be written as a function of k (or E) times $dEd\phi$, where ϕ is the polar angle that characterizes the momentum direction in two dimensions.

The wave function for a free particle in two dimensions is given by:

$$\psi(x,y) = \frac{1}{\sqrt{L^2}} e^{i(k_x x + k_y y)}$$

Where $k_x = \frac{2\pi}{L} n_x$ and $k_y = \frac{2\pi}{L} n_y$.

The energy of the particle is given by:

$$E = \frac{p^2}{2m} = \frac{\hbar^2}{2m} (k_x^2 + k_y^2) = \frac{2\pi^2 \hbar^2}{mL^2} (n_x^2 + n_y^2)$$

Assume we do a transformation from a cartesian coordinate spanned by n_x, n_y to a polar coordinate spanned by n, ϕ , then the following relation holds:

$$n^{2} = n_{x}^{2} + n_{y}^{2}$$

$$\tan \phi = \frac{n_{y}}{n_{x}}$$

$$dN = dn_{x}dn_{y} = ndnd\phi$$

where dN is the density of states.

The energy relation can be rewritten in polar coordinate as:

$$dE = \frac{4\pi^2\hbar^2}{mL^2}ndn$$

Finally, the density of states is given by:

$$dN = ndnd\phi = \frac{mL^2}{4\pi^2\hbar^2}dEd\phi$$

Problem 4. Sakurai 2.30

Using sphreical coordinates, obtain an expression for j for the ground and excites states of the hydrogen atom. Show, in particular for $m_i \neq 0$ states, there is a circulating flux in the sense that j is in the direction of increasing or decreasing ϕ depending on whether m_i is positive or negative.

Recall the probability current is given by:

$$\mathbf{j} = \frac{1}{2m} (\psi^* p \psi - \psi p \psi^*) = \frac{\hbar}{m} \operatorname{Im}(\psi^* \nabla \psi)$$

Where the ∇ operator in spherical coordinate is defined as:

$$\nabla = \frac{1}{r} \frac{\partial}{\partial r} + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} + \frac{1}{r^2 \sin^2 \theta} \frac{\partial}{\partial \phi}$$

The wave function of a hydrogen atom is given by:

$$\psi(r, \theta, \phi) = C_{nlm} R_{nl}(r) Y_l^m(\theta, \phi)$$
$$= C_{nlm} R_{nl}(r) P_l^m(\cos \theta) e^{im\phi}$$

Where R_{nl} is the radial part of the wave function and Y_l^m is the spherical harmonic function. Note that C_{nlm} , R_{nl} and Y_l^m are all real numbers.

Note that since we are only interested on the imaginary part of $\psi * \nabla \psi$, we can only work on the $\hat{\psi}$ part of the wave function. Thus, we can write:

$$j = \frac{\hbar}{m_e} \operatorname{Im}(\psi^* \nabla \psi)$$

$$= \frac{\hbar}{m_e} \operatorname{Im}(\psi^* \left(\frac{\hat{\phi}}{r \sin \phi} \frac{\partial}{\partial \phi}\right) \psi)$$

$$= \frac{\hbar}{m_e} \operatorname{Im}(\psi^* \frac{im}{r \sin \phi} \psi) \hat{\phi}$$

$$= \frac{\hbar}{m_e} \frac{m}{\sin \phi} |\psi^2| \hat{\phi}$$

Therefore, we show that the flux is in direction of ϕ , depending of the magnetic quantum number m.

Problem 5. Sakurai 2.33

The propagator in momentum space analogous to (2.6.26) is given by $\langle \boldsymbol{p''}, t | \boldsymbol{p'}, t_0 \rangle$. Derive an explicit expression for $\langle \boldsymbol{p''}, t | \boldsymbol{p'}, t_0 \rangle$ for the free-particle case.

For a free particle, the hamiltonial is given by:

$$H = \frac{p^2}{2m}$$

Thus, the propagator is given by:

$$\langle \boldsymbol{p''}, t | \boldsymbol{p'}, t_0 \rangle = \langle \boldsymbol{p''} | e^{-iHt} e^{iHt_0} | \boldsymbol{p'}, t_0 \rangle$$

$$= \exp\left(\frac{-i\boldsymbol{p'}^2}{2m\hbar} (t - t_0)\right) \langle \boldsymbol{p''} | \boldsymbol{p'}, t_0 \rangle$$

$$= \exp\left(\frac{-i\boldsymbol{p'}^2}{2m\hbar} (t - t_0)\right) \delta(\boldsymbol{p'} - \boldsymbol{p'})$$

Problem 6. Sakurai 2.36

Show that wave-mechanical approach to the gravity-induced problem discussed in Section 2.7 also leads to phase-difference expression (2.7.17).

In the experiment, the phase of the neutron wave function depends on the length of the path and the time it takes to travel the path i.e., $\phi \sim kx - \omega t$. Thus, the phase difference of the two path is given by:

$$\begin{split} \phi_{BD} - \phi_{AC} &= \left(\frac{p_{BD} - p_{AC}}{\hbar}\right) l_1 - \omega \left(\frac{l_1}{v_{BD}} - \frac{l_1}{v_{AC}}\right) \\ &= \left(\frac{p_{BD} - p_{AC}}{\hbar}\right) l_1 \left(1 + \frac{\hbar \omega m_n}{p_{AC} p_{BD}}\right) \\ &= \left(\frac{p_{BD} - p_{AC}}{\hbar}\right) l_1 \left(1 + \frac{E m_n}{p^2}\right) \\ &= \left(\frac{p_{BD} - p_{AC}}{\hbar}\right) \frac{3}{2} l_1 \end{split}$$

Since $E = \hbar\omega = \frac{p^2}{2m}$.

Then we can make the following approximation and apply the conservation law of energy:

$$p_{BD} - p_{AC} \approx \left(\frac{p_{BD}^2}{2m} - \frac{p_{AC}^2}{2m}\right) \frac{2m}{p}$$
$$= -mgz \frac{2m}{p} = -\frac{2m^2gl_2\sin\theta}{p}$$

Therefore the phase difference is given by:

$$\phi_{BD} - \phi_{AC} = \left(\frac{p_{BD} - p_{AC}}{\hbar}\right) \frac{3}{2} l_1$$
$$= -\frac{2m^2 g l_2 \sin \theta}{p \hbar} \frac{3}{2} l_1$$
$$= -\frac{3m^2 g \lambda}{2\pi \hbar^2} l_1 l_2 \sin \theta$$