# Efficient algorithms for differential equation satisfied by Feynman integrals

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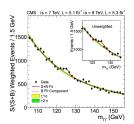
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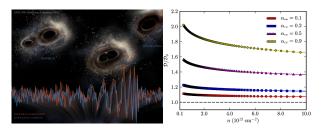
Computer Algebra for Functional Equations in Combinatorics and Physics

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with Charles Doran, Andrew Harder, Pierre Lairez, Eric Pichon-Pharabod and work to appear with Leonardo de la Cruz







Scattering amplitudes are the fundamental tools for making contact between quantum field theory description of nature and experiments

- Comparing particule physics model against datas from accelators
- Post-Minkowskian expansion for Gravitational wave physics
- Various condensed matter and statistical physics systems

## Feynman Integrals: parametric representation

Feynman integral are given by projective space integrals

$$I_{\Gamma}(\underline{\nu}, D; \underline{s}, \underline{m}) = \int_{\Delta_n} \frac{\mathcal{U}_{\Gamma}(\underline{x})^{\omega - \frac{D}{2}}}{\mathcal{F}_{\Gamma}(\underline{x})^{\omega}} \prod_{i=1}^n x_i^{\nu_i - 1} \Omega_0 \qquad \omega = \sum_{i=1}^n \nu_i - \frac{LD}{2}$$

with the volume form on  $\mathbb{P}^{n-1}$ 

$$\Omega_0 = \sum_{i=1}^n (-1)^{i-1} x^i dx^1 \wedge \cdots \widehat{dx^i} \cdots \wedge dx^n$$

The domain of integration is the positive quadrant

$$\Delta_n := \{x_1 \geq 0, \dots, x_n \geq 0 | [x_1, \dots, x_n] \in \mathbb{P}^{n-1} \}$$

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## Feynman Integrals: parametric representation

The graph polynomial is homogeneous degree L+1 in  $\mathbb{P}^{n-1}$ 

$$\mathcal{F}_{\Gamma}(\underline{x}) = \mathcal{U}_{\Gamma}(\underline{x}) \times \mathcal{L}(\underline{m}^2;\underline{x}) - \mathcal{V}_{\Gamma}(\underline{s},\underline{x})$$

▶ Homogeneous polynomial of degree L with  $u_{a_1,...,a_n} \in \{0,1\}$ 

$$\mathcal{U}_{\Gamma}(\underline{x}) := \sum_{\substack{a_1 + \dots + a_n = L \\ \mathbf{0} \le a_i \le 1}} u_{a_1, \dots, a_n} \prod_{i=1}^n x_i^{a_i}$$

the mass hyperplane

$$\mathcal{L}(\underline{m}^2;\underline{x}) := \sum_{n=1}^n m_i^2 x_i$$

▶ Homogeneous polynomial of degree L+1

$$\mathcal{V}_{\Gamma}(\underline{x}) := \sum_{\substack{a_1 + \dots + a_n = L + 1 \\ 0 \le a_i \le 1}} S_{a_i, \dots, a_n} \prod_{i=1}^n x_i^{a_i}$$

## Feynman Integrals: parametric representation

The integrand is an algebraic differential form in  $H^{n-1}(\mathbb{P}^{n-1}\setminus\mathbb{X}_{\Gamma})$  on the complement of the graph hypersurface

$$\mathbb{X}_{\Gamma} := \{ \mathcal{U}_{\Gamma}(\underline{x}) \times \mathcal{F}_{\Gamma}(\underline{x}) = 0, \underline{x} \in \mathbb{P}^{n-1} \}$$

- All the singularities of the Feynman integrals are located on the graph hypersurface
- Generically the graph hypersurface has non-isolated singularities

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## Feynman integral and periods

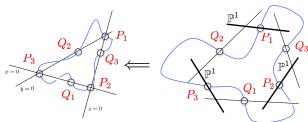
The domain of integration  $\Delta_n$  is not an homology cycle because

$$\partial \Delta_n \cap \mathbb{X}_{\Gamma} = \{(1, 0, \dots, 0), (0, 1, 0, \dots, 0), \dots, (0, \dots, 0, 1)\}$$

we have to look at the relative cohomology

$$H^{\bullet}(\mathbb{P}^{n-1}\backslash X_{\Gamma}; \mathcal{A}_n\backslash \mathcal{A}_n\cap \mathbb{X}_{\Gamma})$$

The normal crossings divisor  $A_n := \{x_1 \cdots x_n = 0\}$  and  $X_\Gamma$  are separated by performing a series of iterated blowups of the complement of the graph hypersurface [Bloch, Esnault, Kreimer]



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## Differential equation

The Feynman integral *are* period integrals of the relative cohomology after performing the appropriate blow-ups

$$\mathfrak{M}(\underline{s},\underline{m}^2):=H^{\bullet}(\widetilde{\mathbb{P}^{n-1}}\backslash \widetilde{X_F};\widetilde{\mathcal{I}_n}\backslash \widetilde{\mathcal{I}_n}\cap \widetilde{X_\Gamma})$$

Since the integrand varies with the physical variables  $\{S_{\underline{a}^i}, m_1^2, \dots, m_n^2\}$  one needs to study a variation of (mixed) Hodge structure

One can show that the Feynman integral are **holonomic D-finite functions** [Bitoun et al.;Smirnov et al.]

A Feynman integrals satisfies inhomogenous differential equations with respect to any set of variables  $\underline{z} \in \{S_{\underline{a}}, m_1^2, \dots, m_n^2\}$ 

$$\mathscr{L}_{\Gamma}(\underline{z}) I_{\Gamma} = \mathscr{S}_{\Gamma}$$

Generically there is an inhomogeneous term  $\mathscr{S}_{\Gamma} \neq 0$  due to the boundary components  $\partial \Delta_n$ 

## Feynman integral D-module

We want to address the questions

- To what class of functions belong Feynman integrals?
- ② What is the geometrical algebraic origin of the motive  $\mathfrak{M}(\underline{s},\underline{m}^2)$ ?
- ① Derivation of the (D-module of) differential equations?  $\mathcal{L}_{\Gamma}(z) I_{\Gamma} = \mathcal{L}_{\Gamma}$

In this talk we focus in the question • and present some new methods for deriving such system of differential equation and its underlying (algebraic) geometry

## Feynman Integrals differential equations

For a given subset of the physical parameters  $\underline{z} := (z_1, \dots, z_r) \subset \{\underline{s}, \underline{m}^2\}$  we want to derive **minimal order** differential equations

$$\mathscr{L}_{\Gamma}(\underline{s},\underline{m}^{2},\partial_{\underline{z}})\int_{\sigma}\frac{\mathcal{U}_{\Gamma}(\underline{x})^{\omega-\frac{D}{2}}}{\mathcal{F}_{\Gamma}(\underline{x})^{\omega}}\prod_{i=1}^{n}x_{i}^{\nu_{i}-1}\Omega_{0}=\mathscr{S}_{\sigma,\Gamma}(\underline{z})$$

One way to achieve this is to construct a Gröbner basis of operators  $T_z$  that annihilate the integrand of the Feynman integral

$$T_{\underline{z}}\left(\frac{\mathcal{U}_{\Gamma}(\underline{x})^{\omega-\frac{D}{2}}}{\mathcal{F}_{\Gamma}(\underline{x})^{\omega}}\prod_{i=1}^{n}x_{i}^{\nu_{i}-1}\Omega_{0}\right)=0$$

such that

$$T_{\underline{z}} = \mathscr{L}_{\Gamma}(\underline{s}, \underline{m}^{2}, \underline{\partial}_{\underline{z}}) + \sum_{i=1}^{n} \partial_{x_{i}} Q_{i}(\underline{s}, \underline{m}^{2}, \underline{\partial}_{\underline{z}}; \underline{x}, \underline{\partial}_{\underline{x}})$$

## Feynman Integrals differential equations

where the finite order differential operator

$$\mathscr{L}_{\Gamma}(\underline{s},\underline{m}^2,\underline{\partial}_{\underline{z}}) = \sum_{\substack{0 \leq a_i \leq o_i \\ 1 \leq i \leq r}} p_{a_1,\ldots,a_r}(\underline{s},\underline{m}^2) \prod_{i=1}^r \left(\frac{d}{dz_i}\right)^{a_i}$$

$$Q_{i}(\underline{\underline{s}},\underline{\underline{m}}^{2},\underline{\partial}_{\underline{z}}) = \sum_{\substack{0 \leq a_{i} \leq o'_{i} \\ 1 \leq i \leq n}} \sum_{\substack{0 \leq b_{i} \leq \tilde{o}_{i} \\ 1 \leq i \leq n}} q_{a_{1},...,a_{r}}^{(i)}(\underline{\underline{s}},\underline{\underline{m}}^{2},\underline{x}) \prod_{i=1}^{r} \left(\frac{d}{dz_{i}}\right)^{a_{i}} \prod_{i=1}^{n} \left(\frac{d}{dx_{i}}\right)^{b_{i}}$$

- ► The orders  $o_i$ ,  $o'_i$ ,  $\tilde{o}_i$  are positive integers
- $\triangleright$   $p_{a_1,...,a_r}(\underline{S},\underline{m}^2)$  polynomials in the kinematic variables
- $q_{a_1,...,a_r}^{(i)}(\underline{s},\underline{m}^2,\underline{x})$  rational functions in the kinematic variable and the projective variables x.

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## Feynman Integrals differential equations

Integrating over a cycle  $\gamma$  gives

$$0 = \oint_{\gamma} T_{\underline{z}} \Omega_{\Gamma} = \mathscr{L}_{\Gamma}(\underline{s}, \underline{m}, \partial_{\underline{z}}) \oint_{\gamma} \Omega_{\Gamma} + \oint_{\gamma} d\beta_{\Gamma}$$

For a cycle  $\partial \gamma = \emptyset$  then  $\oint_{\gamma} d\beta_{\Gamma} = 0$  and we get

$$\mathscr{L}_{\Gamma}(\underline{s},\underline{m},\partial_{\underline{z}})\oint_{\gamma}\Omega_{\Gamma}=0$$

For the Feynman integral / we have

$$0 = \int_{\Delta_n} T_{\underline{z}} \Omega_{\Gamma} = \mathscr{L}_{\Gamma}(\underline{s}, \underline{m}, \partial_{\underline{z}}) I_{\Gamma} + \int_{\Delta_n} d\beta_{\Gamma}$$

since  $\partial \Delta_n \neq \emptyset$ 

$$\mathscr{L}_{\Gamma}(\underline{s},\underline{m},\partial_{\underline{z}})I_{\Gamma}=\mathscr{S}_{\Gamma}$$

So we need the telescoper and the certificate

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## The Rational case

We start with the case of a rational differential form with  $D \in 2\mathbb{N}^*$ 

$$\Omega_{\Gamma} = \frac{\mathcal{U}_{\Gamma}(\underline{x})^{\omega - \frac{D}{2}}}{\mathcal{F}_{\Gamma}(\underline{x})^{\omega}} \prod_{i=1}^{n} x_{i}^{\nu_{i} - 1} \Omega_{0} \qquad \omega = \sum_{i=1}^{n} \nu_{i} - \frac{LD}{2}$$

- We as well assume that all the mass parameters are all vanishing  $m_1, \dots, m_n \neq 0$
- And that  $\omega > 0$ , i.e.  $\sum_{i=1}^{n} \nu_i > LD/2$

So that the integral of  $\Omega_{\Gamma}$  on the positive orthan is a convergent integral

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## The sunset graph

The two-loop sunset graph in D = 2



$$I_{\ominus}(p^2,\underline{m}^2) = \int_{\mathbb{R}^3_+} \frac{dx_1 dx_2 dx_3}{\mathcal{F}_{\ominus}(\underline{x})}$$

The polar hypersurface of the integral is an elliptic curve  $\mathcal{F}_{\odot}(\underline{x})=0$ 

$$\mathcal{F}_{\odot}(\underline{x}) = (x_1x_2 + x_1x_3 + x_2x_3)(m_1^2x_1 + m_2^2x_2 + m_3^2x_3) - p^2x_1x_2x_3$$

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One can obtain a differential equation annihilating acting on the integral using the Griffiths-Dwork method

Let define the integrand in differential form

$$\eta_{\circleddash} = \frac{x_1 dx_2 \wedge dx_3 - x_2 dx_1 \wedge dx_3 + x_3 dx_1 \wedge dx_2}{\mathcal{F}_{\circleddash}(\underline{x})} = \frac{\Omega_0}{\mathcal{F}_{\circleddash}(\underline{x})}$$

consider

$$\frac{\partial \eta_{\odot}}{\partial p^2} = x_1 x_2 x_3 \frac{\Omega_0}{\mathcal{F}_{\odot}(\underline{x})^2}; \qquad \frac{\partial^2 \eta_{\odot}}{(\partial p^2)^2} = 2(x_1 x_2 x_3)^2 \frac{\Omega_0}{\mathcal{F}_{\odot}(\underline{x})^3}$$

Since we know we have the geometry of an elliptic curve we are looking for a second order differential operator acting on  $\eta_{\odot}$ 

$$\mathscr{L}_{\circleddash}(p^2) = rac{\partial^2}{(\partial p^2)^2} + q_1(p^2, \underline{m}^2) rac{\partial}{\partial p^2} + q_0(p^2, \underline{m}^2)$$

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Remark that  $(x_1x_2x_3)^2$  lies in the Jacobian ideal for  $\mathcal{F}_{\ominus}(\underline{x})$ 

$$(x_1x_2x_3)^2 = \sum_{i=1}^3 C_i^{(1)}(\underline{x})\partial_{x_i}\mathcal{F}_{\odot}(\underline{x})$$

with  $C_i^{(1)}(\underline{x})$  homogeneous of degree 4 in the  $(x_1, x_2, x_3)$  variables Following Griffiths one introduces the differential form

$$\beta_{1} = \frac{(x_{2}C_{3}^{(1)}(\underline{x}) - x_{3}C_{2}^{(1)}(\underline{x}))dx_{1}}{\mathcal{F}_{\odot}(\underline{x})^{2}} + \frac{(x_{1}C_{3}^{(1)}(\underline{x}) - x_{3}C_{1}^{(1)}(\underline{x}))dx_{2}}{\mathcal{F}_{\odot}(\underline{x})^{2}} + \frac{(x_{1}C_{2}^{(1)}(\underline{x}) - x_{2}C_{1}^{(1)}(\underline{x}))dx_{3}}{\mathcal{F}_{\odot}(x)^{2}}$$

such that

$$d\beta_1 = 2 \frac{\sum_{i=1}^3 C_i^{(1)}(\underline{x}) \partial_{x_i} \mathcal{F}_{\odot}(\underline{x}) \Omega_0}{\mathcal{F}_{\odot}(\underline{x})^3} - \frac{\sum_{i=1}^3 \partial_{x_i} C_i^{(1)}(\underline{x}) \Omega_0}{\mathcal{F}_{\odot}(\underline{x})^2}$$

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$$\mathscr{L}_{\ominus}(p^2)\eta_{\ominus} = \frac{q_1(p^2,\underline{m}^2)x_1x_2x_3 + \sum_{i=1}^3 \partial_{x_i}C_i^{(1)}(\underline{x})}{\mathcal{F}_{\ominus}(\underline{x})^2}\Omega_0 + d\beta_1$$

We can again reduce this second order pole using that there exist a polynomial  $q_1(p^2, \underline{m}^2)$  such that

$$q_1(p^2,\underline{m}^2)x_1x_2x_3+\sum_{i=1}^3\partial_{x_i}C_i^{(1)}(\underline{x})=\sum_{i=1}^3C_i^{(2)}\partial_{x_i}\mathcal{F}_{\odot}(\underline{x})$$

with  $C_i^{(2)}$  of degree 1. One introduces the 1-form  $\beta_2$ 

$$\beta_2 = \sum_{i=1}^{3} \epsilon^{ijk} \frac{x_j C_k^{(2)}(\underline{x}) dx_i}{\mathcal{F}_{\bigcirc}(\underline{x})}$$

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such that

$$d\beta_{2} = \frac{\sum_{i=1}^{3} C_{i}^{(2)}(\underline{x}) \partial_{x_{i}} \mathcal{F}_{\ominus}(\underline{x})}{\mathcal{F}_{\ominus}(\underline{x})^{2}} - \frac{\sum_{i=1}^{3} \partial_{x_{i}} C_{i}^{(2)}(\underline{x}) \Omega_{0}}{\mathcal{F}_{\ominus}(\underline{x})}$$

We have achieved that

$$\left(\frac{\partial^2}{(\partial p^2)^2} + q_1(p^2, \underline{m}^2)\frac{\partial}{\partial p^2} + \sum_{i=1}^3 \partial_{x_i} C_i^{(2)}(\underline{x})\right) \eta_{\odot} = d(\beta_1 + \beta_2)$$

because the  $C_i^{(2)}(\underline{x})$  are of degree 1 in  $(x_1, x_2, x_3)$  then  $q_0(p^2, \underline{m}^2) = \partial_{x_i} C_i^{(2)}(\underline{x})$  only depends on  $p^2, \underline{m}^2$ 

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We then conclude that the minimal operator acting on the sunset integral is the Picard-Fuchs operator

$$\mathscr{L}_{p^2} = rac{\partial^2}{(\partial p^2)^2} + q_1(p^2, \underline{m}^2) rac{\partial}{\partial p^2} + q_0(p^2, \underline{m}^2)$$

which acts on the integrals as

$$\mathscr{L}_{\rho^2}I_{\ominus}(\rho^2)=\int_{x_i\geq 0}\mathscr{L}_{\rho^2}rac{\Omega_0}{\mathcal{F}_{\ominus}(\underline{x})}=\int_{x_i\geq 0}d(eta_1+eta_2)
eq 0$$

We have constructed by the telescoper  $T_{p^2}=\mathscr{L}_{\odot}(p^2)$  and the certificate  $C_{\odot}=d(\beta_1+\beta_2)$ 

The differential operator  $\mathcal{L}_{p^2}$  is the Picard–Fuchs operator of the elliptic curve defined by the graph polynomial  $\mathcal{F}_{\ominus}(x_1,x_2,x_3)=0$ 

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If one considers the family of elliptic curve *E* 

$$y^2 = 4x^3 - g_2(t)x - g_3(t);$$
  $\Delta(t) = g_2(t)^3 - 27g_3(t)^2$ 

the periods satisfy the differential system of equations

$$\frac{d}{dt} \begin{pmatrix} \int_{\gamma} \frac{dx}{y} \\ \int_{\gamma} \frac{xdx}{y} \end{pmatrix} = \begin{pmatrix} -\frac{1}{12} \frac{d}{dt} \log \Delta(t) & \frac{3\delta(t)}{2\Delta(t)} \\ -\frac{g_2(t)\delta(t)}{8\Delta(t)} & \frac{1}{12} \frac{d}{dt} \log \Delta(t) \end{pmatrix} \begin{pmatrix} \int_{\gamma} \frac{dx}{y} \\ \int_{\gamma} \frac{xdx}{y} \end{pmatrix}$$

with  $\delta(t) = 3g_3(t) \frac{d}{dt} g_2(t) - 2g_2(t) \frac{d}{dt} g_3(t)$ 

The Picard–Fuchs operator acting on the period integral  $\int_{\gamma} dx/y$  is

$$\begin{aligned} \mathscr{L}_{\text{ell}} &= 144 \Delta(t)^2 \delta(t) \frac{d^2}{dt^2} + 144 \Delta(t) \left( \delta(t) \frac{d\Delta(t)}{dt} - \Delta(t) \frac{d\delta(t)}{dt} \right) \frac{d}{dt} \\ &+ 27 g_2(t) \delta(t)^3 + 12 \frac{d^2 \Delta(t)}{dt^2} \delta(t) \Delta(t) - \left( \frac{d\Delta(t)}{dt} \right)^2 \delta(t) - 12 \frac{d\delta(t)}{dt} \Delta(t) \frac{d\Delta(t)}{dt}. \end{aligned}$$

This matches the differential operator derived using the Griffiths–Dwork method

## **Extended Griffiths-Dwork algorithms**

In general the graph hypersurface does not have isolated singularities (which is the generic case) therefore the "naïve" implementation of the Griffiths-Dwork algorithm does not work

One could use the implementation of Doron Zeilberger (1990) creative telescoping algorithm by F. Chyzak or K. Koutschan but the algorithm takes a very long time for graph with many edges

$$\Omega_{\Gamma} = rac{\mathcal{U}_{\Gamma}(\underline{x})^{\omega - rac{D}{2}}}{\mathcal{F}_{\Gamma}(\underline{x})^{\omega}} \, \prod_{i=1}^{n} x_{i}^{
u_{i}-1} \, \Omega_{0}$$

We used period by Pierre Lairez of an extended Griffiths-Dwork algorithm that handles singular hypersurfaces

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In this example we saw the pole reduction

$$\frac{\partial^2 \eta_{\odot}}{(\partial \rho^2)^2} = \frac{2(x_1 x_2 x_3)^2}{\mathcal{F}_{\odot}(\underline{x})^3} \Omega_0 = \frac{\sum_{i=1}^3 \partial_{x_i} C_i}{\mathcal{F}_{\odot}(\underline{x})^2} \Omega_0 + d\beta_1$$

For singular hypersurface  $X_{\Gamma} \subset \mathbb{P}^{n-1}$  the Jacobian reduction may not be enough to reduce the pole order when  $k \geq n$ 

Other reduction rules come from the *syzygies* of the derivatives  $\frac{\partial \mathcal{F}_{\Gamma}}{\partial x_i}$ , i.e. tuples  $(B_1,\ldots,B_n)$  be homogeneous of degree  $k \deg \mathcal{F}_{\Gamma} - n + 1$  such that  $\sum_i B_i \frac{\partial \mathcal{F}_{\Gamma}}{\partial x_i} = 0$  such that  $(\xi_i = (-1)^{i-1} dx_1 \cdots \widehat{dx_i} \cdots dx_n)$ 

$$\frac{\sum_{i}\frac{\partial B_{i}}{\partial x_{i}}}{\mathcal{F}_{\Gamma}^{k}}\Omega_{0}=d\left(\sum_{i}\frac{B_{i}}{\mathcal{F}_{\Gamma}^{k}}\xi_{i}\right)\Longrightarrow\int_{\gamma}\frac{\sum_{i}\frac{\partial B_{i}}{\partial x_{i}}}{\mathcal{F}_{\Gamma}^{k}}\Omega_{0}=0.$$

In singular cases, these relations are missed by the Griffiths–Dwork reduction, we need the extended Griffiths–Dwork reduction implemented by [Lairez]

Given a form 
$$\Omega = \frac{A}{\mathcal{F}_{\Gamma}^k} d\underline{x}$$
  $\deg A = k \deg \mathcal{F}_{\Gamma} - n$ 

① Compute a basis of the space  $S_k$  of all syzygies of degree  $k \deg \mathcal{F}_{\Gamma} - n + 1$  quotiented by the space of trivial syzygies

$$D_{ij} = -D_{ji}, \qquad B_i = \sum_{j=1}^n D_{ij} \frac{\partial \mathcal{F}_{\Gamma}}{\partial x_j} \Longrightarrow \sum_i B_i \frac{\partial \mathcal{F}_{\Gamma}}{\partial x_i} = 0$$

are irrelevant because already used by the Griffiths-Dwork reduction

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Given a form 
$$\Omega = \frac{A}{\mathcal{F}_{\Gamma}^{k}} d\underline{x}$$
  $\deg A = k \deg \mathcal{F}_{\Gamma} - n$ 

② Compute a normal form R of A modulo the Jacobian ideal plus the space  $dV = \left\{ \sum_i \frac{\partial B_i}{\partial x_i} \mid \underline{B} \in V \right\}$ , that is for some polynomials  $B_i$  and  $C_i$ 

$$A = R + \underbrace{\sum_{i} \frac{\partial B_{i}}{\partial x_{i}}}_{\in dV} + \underbrace{C_{1} \frac{\partial \mathcal{F}_{\Gamma}}{\partial x_{1}} + \dots + C_{n} \frac{\partial \mathcal{F}_{\Gamma}}{\partial x_{n}}}_{\in \text{Jacobian ideal}},$$

Given a form 
$$\Omega = \frac{A}{\mathcal{F}_{\Gamma}^k} d\underline{x}$$
  $\deg A = k \deg \mathcal{F}_{\Gamma} - n$ 

This leads to the following relation

$$(k-1)\frac{A}{\mathcal{F}_{\Gamma}^{k}}d\underline{x} = \frac{\sum_{i}\frac{\partial C_{i}}{\partial x_{i}}}{\mathcal{F}_{\Gamma}^{k-1}}d\underline{x} - d\left(\sum_{i}\frac{B_{i}}{\mathcal{F}_{\Gamma}^{k}}\xi_{i} + \sum_{i}\frac{C_{i}}{\mathcal{F}_{\Gamma}^{k-1}}\xi_{i}\right).$$

Then

$$\int_{\gamma} \frac{A(\underline{x})}{\mathcal{F}_{\Gamma}(\underline{x})^{k}} d\underline{x} = -\frac{1}{k-1} \int_{\gamma} \frac{\sum_{i} \frac{\partial C_{i}}{\partial x_{i}}}{\mathcal{F}_{\Gamma}^{k-1}} d\underline{x},$$

The extended Griffiths–Dwork reduction presented above is not always enough and may need further extensions, i.e. syzygies of syzygies. There is a hierarchy of extensions which eventually collapse to the strongest possible reduction.

However, for all the computations presented here, we only needed the first extension.

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#### Pole conditions traces (1888)]

In the construction we will only consider the case where  $\beta(\underline{x},t)$  is holomorphic on  $\mathbb{P}^{n-1}\setminus X_{\Gamma}$ , that is is  $\beta_{\Gamma}$  does not have poles that are not present in  $\Omega_{\Gamma}$ .

Consider the rational function  $F(x_1, x_2)$ 

$$\frac{ax_1+bx_2+c}{\left(\alpha x_1^2+\beta x_2^2+\gamma x_1 x_2+\delta x_1+\eta x_2+\zeta\right)^2}=\partial_{x_1}\frac{\textit{N}_1(x_1,x_2)}{\textit{D}_1(x_1,x_2)}+\partial_{x_2}\frac{\textit{N}_2(x_1,x_2)}{\textit{D}_2(x_1,x_2)}$$

where  $a, b, c, \alpha, \beta, \gamma, \delta, \eta, \eta$  are constants and polynomials  $N_i(x_1, x_2)$  and  $D_i(x_1, x_2)$  with i = 1, 2.

The denominators have poles at  $x_2^0 = (a\delta - 2\alpha c)/(2\alpha b - a\gamma)$  which is not a pole of the left-hand-side.

This means one can find a cycle  $\gamma$  passing by  $x_2^0$  such that the integral of  $\int_{\gamma} F(x_1, x_2)$  is finite and non-vanishing.

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## Minimality of the Picard-Fuchs operator

This dimension  $\dim(V_{\Gamma}) = (-1)^{n+1} \chi\left((\mathbb{C}^*)^n \backslash \mathbb{V}(\mathcal{U}_{\Gamma}) \cup \mathbb{V}(\mathcal{F}_{\Gamma})\right)$  gives an upper bound on the order of the minimal order differential operator acting on the Feynman integral.

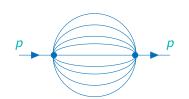
The extended Griffiths–Dwork algorithm leads to a minimal order differential operator

$$\mathscr{L}_{\Gamma}\Omega_{\Gamma}=d\beta_{\Gamma}$$

annihilating (in cohomology) the Feynman integral differential form  $\Omega_{\Gamma}$  with the condition that the certificate  $\beta_{\Gamma}$  is an holomorphic form on  $\mathbb{P}^{n-1}\backslash Z_{\Gamma}$ .  $\mathscr{L}_{\Gamma}$  is the minimal differential order differential operator satisfying this condition.

Using the algorithm by [Chyzak, Goyer, Mezzarobba] we test the irreducibility of the Picard-Fuchs operator and factorize when it is reducible.

## Sunset graph Picard-Fuchs operator



$$\Omega_n^{\ominus}(t,\underline{m}^2) := \frac{\Omega_0}{\mathcal{F}_n^{\ominus}(t,\underline{m}^2;\underline{x})} \in H^{n-1}(\mathbb{P}^{n-1} \setminus X_{\ominus})$$

$$\mathcal{F}_n^{\circleddash}(t,\underline{m}^2;\underline{x}) := x_1 \cdots x_n \left( \left( \sum_{i=1}^n \frac{1}{x_i} \right) \left( \sum_{j=1}^n m_j^2 x_j \right) - t \right)$$

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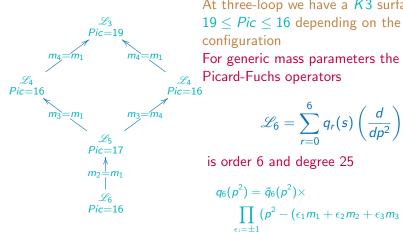
The graph hypersurface  $\mathcal{F}_n^{\odot}(t,\underline{m}^2;x)=0$  defines a Calabi-Yau manifold of dimension n-1

For generic physical parameters configurations we find a minimal order Picard—Fuchs operator [Lairez, Vanhove]

$$\mathscr{L}_t = \sum_{r=0}^{o_n} q_r(t, \underline{m}^2) \left(\frac{d}{dt}\right)^r \qquad o_n = 2^n - \binom{n+1}{\left\lfloor \frac{n+1}{2} \right\rfloor}; n \geq 2.$$

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### The three-loop sunset graph



$$\mathscr{L}_r = (\alpha \frac{d}{dp^2} + \beta) \circ \mathscr{L}_{r-1}$$
 with  $\tilde{q}_6(p^2)$  degree 17 with apparent

At three-loop we have a K3 surface of 19 < Pic < 16 depending on the mass

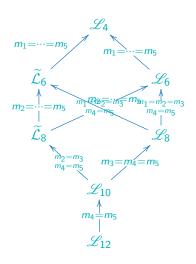
$$\mathscr{L}_6 = \sum_{r=0}^6 q_r(s) \left(\frac{d}{dp^2}\right)^r$$

is order 6 and degree 25

$$egin{aligned} q_6(
ho^2) &= ilde{q}_6(
ho^2) imes \ &\prod_{\epsilon_i = \pm 1} (
ho^2 - (\epsilon_1 m_1 + \epsilon_2 m_2 + \epsilon_3 m_3 + \epsilon_4 m_4)^2) \end{aligned}$$

singularities

## The four-loop sunset graph



The geometry is the one of the Calabi-Yau threefold.

For generic kinematics we have an order 12 degree 121 operator

$$\mathcal{L}_{t}^{[1^{5}]} = \sum_{r=0}^{12} q_{r}^{[1^{5}]}(t, \underline{m}^{2}) \left(\frac{d}{dt}\right)^{r}.$$

The degree of the apparent singularities is a polynomial of degree 98

## The five-loop and six-loop sunset graph

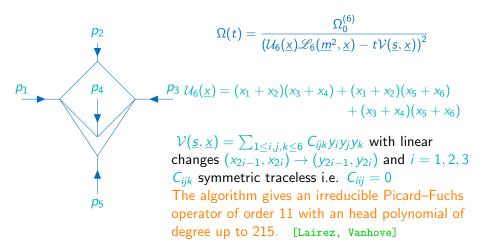
- ▶ The six mass configuration  $m_1 \neq m_2 \neq m_3 \neq m_4 \neq m_5 \neq m_6$  denote [1<sup>6</sup>]: the Picard–Fuchs operator of order 29 and degree of the polynomial  $q_{29}(t)$  is 521.
- ▶ The seven mass configuration  $m_1 \neq m_2 \neq m_3 \neq m_4 \neq m_5 \neq m_6 \neq m_7$ : the Picard–Fuchs operator of order 58 with a degree 2273
- ▶ Results compatible with a CY n-1-fold

Results obtained using Pierre Lairez period

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## **Tardigrade**

The rational differential form in  $\mathbb{P}^5$ 



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## **Tardigrade**

The rational differential form in  $\mathbb{P}^5$ 

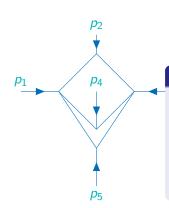
$$\Omega(t) = rac{\Omega_0^{(6)}}{\left(\mathcal{U}_6(\underline{x})\mathscr{L}_6(\underline{m}^2,\underline{x}) - t\mathcal{V}(\underline{s},\underline{x})
ight)^2}$$

The motive associated to this graph

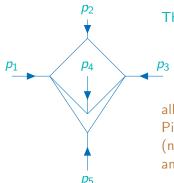
#### Theorem (Tardigrade motive)

Let  $X_{(2,2,2);D}$  be the tardigrade hypersurface for generic mass and momentum parameters and  $D \geq 2$ . Then there is a quartic K3 surface with six  $A_1$  singularities so that  $\operatorname{Gr}_4^W \operatorname{H}^4(X_{(2,2,2);D};\mathbb{Q})$  is isomorphic to  $\operatorname{H}^2(S;\mathbb{Q})(-1)$  for a K3 surface S up to mixed Tate factors.

Determined in [Doran, Harder, Pichon-Pharabod, Vanhove]



## **Tardigrade**



The rational differential form in  $\mathbb{P}^5$ 

$$\Omega(t) = rac{\Omega_0^{(6)}}{\left(\mathcal{U}_6(\underline{x})\mathscr{L}_6(\underline{m}^2,\underline{x}) - t\mathcal{V}(\underline{s},\underline{x})
ight)^2}$$

The singularities of the graph polynomials are all of type  $A_1$  and one can apply Eric Pichon-Pharabod program lefschetz-family to (numerically) determine the transcendental lattice and confirm that we have a K3 of Picard Rank 11

# The non-rational case

We now consider the non-rational case  $D \in \mathbb{R}$ 

$$\Omega_{\Gamma} = \left(\frac{\mathcal{U}_{\Gamma}(\underline{x})}{\mathcal{F}_{\Gamma}(\underline{x})}\right)^{\sum_{i}\nu_{i}} \left(\frac{\mathcal{U}_{\Gamma}(\underline{x})^{L+1}}{\mathcal{F}_{\Gamma}(\underline{x})^{L}}\right)^{D} \prod_{i=1}^{n} x_{i}^{\nu_{i}-1} \Omega_{0}$$

- We relax all assumption on the mass parameters who can all vanish  $m_1, \dots, m_n \in \mathbb{R}$
- We have degree 0 rational form

$$R(\underline{x}) := \frac{\mathcal{U}_{\Gamma}(\underline{x})^{L+1}}{\mathcal{F}_{\Gamma}(\underline{x})^{L}}$$

## Feynman Integrals: divergences

As a function of the powers of the propagators  $\nu$  and the dimension D the integral has singularities located on hyperplane defined by  $\sum_{i=1}^{n} a_i \nu_i + a_0 D = 0$  with  $(a_0, a_1, \dots, a_n) \in \mathbb{Z}^{n+1}$ 

One can perform a Laurent expansion near say  $D_c = 4$  dimensions

$$I_{\Gamma}(\underline{s},\underline{m}^{2};\underline{\nu},D) = \sum_{r \geq -2L} (D - D_{c})^{r} I_{\Gamma}^{(r)}(\underline{s},\underline{m}^{2};\underline{\nu})$$

where  $I_{\Gamma}^{(r)}(s, m^2; \nu)$  are convergent integrals.



Eugene R. Speer

Generalized Feynman Amplitudes

Princeton University Press, (1969)

Pierre Vanhove (IPhT)

6/12/2023

## The sunset graph

The two-loop sunset graph in  $D = 2 - 2\epsilon$  with  $\epsilon \in \mathbb{R}$ 



$$I_{\ominus}(p^2,\underline{m}^2) = \int_{\mathbb{R}^3_+} \left(\frac{\mathcal{U}_{\ominus}^3(\underline{x})}{\mathcal{F}_{\ominus}^2(\underline{x})}\right)^{\epsilon} \frac{dx_1 dx_2 dx_3}{\mathcal{F}_{\ominus}(\underline{x})}$$

with

$$\mathcal{U}_{\odot}(\underline{x}) = x_1x_2 + x_1x_3 + x_2x_3$$

The polar hypersurface of the integral is still the elliptic curve  $\mathcal{F}_{\ominus}(\underline{x})=0$ 

$$\mathcal{F}_{\ominus}(\underline{x}) = (x_1x_2 + x_1x_3 + x_2x_3)(m_1^2x_1 + m_2^2x_2 + m_3^2x_3) - p^2x_1x_2x_3$$

We consider differentiation with respect to a single physical parameter  $z \in \{\vec{m}, \vec{s}\}$ 

We consider the derivative

$$\left(\frac{d}{dz}\right)^{a}\Omega_{\Gamma}^{\epsilon} = \frac{\Gamma(1+a+\epsilon)}{\Gamma(1+2\epsilon)} \frac{(x_{1}x_{2}x_{3})^{a}}{\mathcal{F}_{\odot}^{a+1}} \left(\frac{\mathcal{U}_{\odot}^{3}}{\mathcal{F}_{\odot}^{2}}\right)^{\epsilon} \Omega_{n}^{(0)}$$

we reduce the numerator in the Jacobian ideal of  $\mathcal{F}_{\circleddash}$ 

$$\frac{\Gamma(1+a+\epsilon)}{\Gamma(1+2\epsilon)} (x_1x_2x_3)^a = \vec{C}_{(a)} \cdot \vec{\nabla} \mathcal{F}_{\odot}.$$

Integration by part gives

$$\left(\frac{d}{dz}\right)^{a}\Omega_{\Gamma}^{\epsilon} = \frac{\vec{\nabla}\cdot\vec{C}_{(a)}}{a\mathcal{F}_{\odot}^{a}}\left(\frac{\mathcal{U}_{\odot}^{3}}{\mathcal{F}_{\odot}^{2}}\right)^{\epsilon}\Omega_{0} + \epsilon\frac{\vec{C}_{(a)}\cdot\vec{\nabla}\log(\mathcal{U}_{\odot}^{3}/\mathcal{F}_{\odot}^{2})}{a\mathcal{F}_{\odot}^{a}}\left(\frac{\mathcal{U}_{\odot}^{3}}{\mathcal{F}_{\odot}^{2}}\right)^{\epsilon}\Omega_{0} + d\beta_{(a)}$$

or equivalently

$$\left(\frac{d}{dz}\right)^{a}\Omega_{\Gamma}^{\epsilon} = \frac{\vec{\nabla} \cdot \vec{C}_{(a)}}{(a+2\epsilon)\mathcal{F}_{\odot}^{a}} \left(\frac{\mathcal{U}_{\odot}^{3}}{\mathcal{F}_{\odot}^{2}}\right)^{\epsilon} \Omega_{0} + 3\epsilon \frac{\vec{C}_{(a)} \cdot \vec{\nabla} \log(\mathcal{U}_{\odot})}{(a+2\epsilon)\mathcal{F}_{\odot}^{a}} \left(\frac{\mathcal{U}_{\odot}^{3}}{\mathcal{F}_{\odot}^{2}}\right)^{\epsilon} \Omega_{0} + d\beta_{(a)}$$

We ask that

$$\vec{C}_{(a)} \cdot \vec{\nabla} \mathcal{U}_{\odot} = c_{(a)}(\underline{x}) \mathcal{U}_{\odot}.$$

Solving the system

$$\begin{split} \frac{\Gamma(1+a+\epsilon)}{\Gamma(1+2\epsilon)} \left(x_1 x_2 x_3\right)^a &= \vec{C}_{(a)} \cdot \vec{\nabla} \mathcal{F}_{\odot}, \\ \vec{C}_{(a)} \cdot \vec{\nabla} \mathcal{U}_{\odot} &= \textit{\textbf{c}}_{(a)}(\underline{\textbf{x}}) \mathcal{U}_{\odot}. \end{split}$$

Gives the pole reduction

$$\left(\frac{d}{dz}\right)^{a}\Omega_{\Gamma}^{\epsilon} = \frac{\vec{\nabla} \cdot \vec{C}_{(a)} + 3\epsilon C_{(a)}(\underline{x})}{(a+2\epsilon)\mathcal{F}_{\odot}^{a}} \left(\frac{\mathcal{U}_{\odot}^{3}}{\mathcal{F}_{\odot}^{2}}\right)^{\epsilon} \Omega_{0} + d\beta_{(a)}.$$

- ► This tells us how to modify the Griffiths-Dwork pole reduction and deduce the € deformed differential equation.
- ▶ This allows to treat case that are divergence for  $\epsilon = 0$  which was not possible with the previous algorithm

Work in progress [de la Cruz, Vanhove]

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## The sunset integrals: all equal mass case

For the all equal mass case the algorithm gives (up to 20 loops) For the all equal mass case  $m_1 = \cdots = m_{l+1} = 1$  we find the sunset Feynman integral satisfies the differential equation

$$\mathscr{L}_{\circleddash}^{(I),\epsilon}I_{\circleddash}(\{1,\ldots,1\},t,\epsilon) = -(I+1)!rac{\Gamma(1+\epsilon)^I}{\Gamma(1+I\epsilon)}$$

with

$$\mathscr{L}_{\text{O}}^{(I),\epsilon} = \mathscr{L}_{\text{O}}^{(I)} + \epsilon \mathscr{L}_{\text{O}}^{(I),1} + \dots + \epsilon^{I} \mathscr{L}_{\text{O}}^{(I),I}$$

where the differential operator is  $\mathscr{L}_{\odot}^{(I),r}$  is of order I-r.

## Two loop Sunset: different masses I

$$\mathscr{L}_{\ominus}^{(2),\epsilon} = \mathscr{L}_1^{(1)} \mathscr{L}_1^{(2)} \mathscr{L}_{\ominus}^{3-\mathit{mass}} + \epsilon \mathscr{L}_4^{(3)} + \epsilon^2 \mathscr{L}_3^{(4)} + \epsilon^3 \mathscr{L}_2^{(5)} + \epsilon^4 \mathscr{L}_1^{(6)} + \epsilon^5 \mathscr{L}_0^{(7)}$$

where  $\mathcal{L}_m^{(r)}$  are irreducible differential operator of order m and  $\mathcal{L}_{\odot}^{3-mass}$  is the differential operator for the three-mass two-loop sunset integral in two dimensions.

Its actions on the Feynman integral is given by

$$\mathscr{L}_{\ominus}^{(2),\epsilon}I(\underline{m},t;\epsilon)=\mathscr{S}(\vec{m},t;\epsilon)$$

with the source term

$$\mathscr{S}(\vec{m},t;\epsilon) = \frac{c_{23}(t,\epsilon)\Gamma(\epsilon+1)^2}{(m_2m_3)^{2\epsilon}\Gamma(1+2\epsilon)} + \frac{c_{13}(t,\epsilon)\Gamma(\epsilon+1)^2}{(m_1m_3)^{2\epsilon}\Gamma(1+2\epsilon)} + \frac{c_{12}(t,\epsilon)\Gamma(\epsilon+1)^2}{(m_1m_2)^{2\epsilon}\Gamma(1+2\epsilon)}$$

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## Two loop Sunset: different masses II

The  $\epsilon$  deformed operator has for highest order term

$$\begin{aligned} \mathcal{L}_{\odot}^{(2),\epsilon} \Big|_{(d/dt)^4} &= t^3 \prod_{i=1}^4 (t - \mu_i^2) \\ &\times \left( -(2\epsilon + 5) t^2 - 2 (m_1^2 + m_2^2 + m_3^2) (1 + 2\epsilon) t + (7 + 6\epsilon) \prod_{i=1}^4 \mu_i \right) \end{aligned}$$

where

$$\mu_i = \{m_1 + m_2 + m_3, -m_1 + m_2 + m_3, m_1 - m_2 + m_3, m_1 + m_2 - m_3\}$$
 are the thresholds.

- ightharpoonup The  $\epsilon$  deformation is only affecting the apparent singularities
- ► The non-apparent singularities are still the roots of the discriminant of the sunset elliptic curve
- ► The order 4 operator is irreducible

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#### **Outlook**

- ₩ We have put forward a new approach for deriving the differential equation for Feynman integrals
- We can derive the differential equations in general dimension by extending the Griffiths-Dwork reduction
- " We see how the twist  $\epsilon$ -factor affects only the apparent singularities
- For graphs with many edges the reduction takes a long time we have been using the FiniteFlow program to speed up the computation but still improvements are needed

We have a seminar on these mathematical aspects of Feynman integral run by Francis Brown, Erik Panzer, Federico Zerbini and myself at the address https://www.ihes.fr/vanhove/motivefeynman.html

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