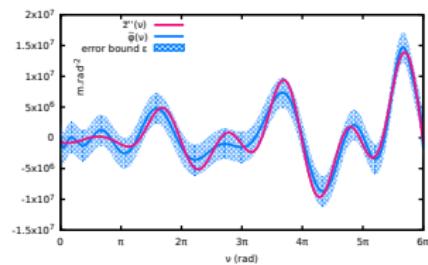
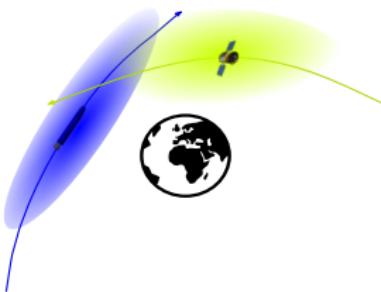


Ten Years of Space Junk and related Symbolic-Numeric Algorithms



$$\begin{matrix} p \\ 1 \pm 0.1 \end{matrix} + \begin{matrix} q \\ 2 \pm 0.2 \end{matrix} = \begin{matrix} RN(p+q) \\ 3 \pm 0.3 \end{matrix}$$
A diagram illustrating the addition of two numbers with associated error bounds. Two circles represent the numbers p and q , each with a central value and a surrounding shaded region indicating the error range. The sum is represented by a third circle containing the result $RN(p+q)$ and its error bound 3 ± 0.3 .



How it started?

LAAS-CNRS Retweeted



Numérique au CNRS @INS2I_CNRS · Jan 27

#CNRSSnews | Teams from the @LaasCNRS lab, in collaboration with the #LIP lab are developing a computer program to calculate the risk of collision between a satellite and orbiting debris in real time.

news.cnrs.fr/articles/new-a...

@CNRS_Toulouse @CNRS_dr07 @ENSdeLyon @UnivLyon1



How it's going?

Joint works with D. Arzelier, F. Bréhard, M. Masson, J.-B. Lasserre, B. Salvy, R. Serra

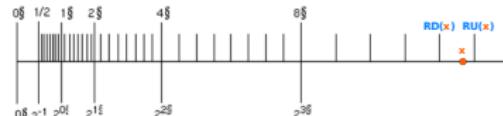
Research Context

Can/Should we trust the numerics?

Numerics: floating-point arithmetic \rightsquigarrow FAST

Global optimization, systems of diff. equations, integration

Usually, solutions lack certification of the output accuracy



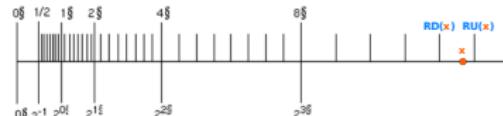
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Computer Algebra Systems (eg. Maple) \rightsquigarrow EXACT

Usually, pure symbolic methods are scarce or computationally expensive

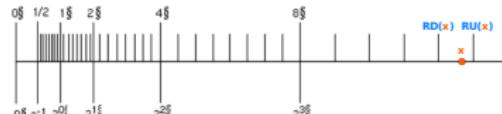
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Need of fast and certified sols:

computer-aided proofs

safety-critical applications e.g.,
control units of aircraft,
particle accelerators
autonomous GNC of spacecraft

Computer Algebra Systems (eg. Maple) \rightsquigarrow EXACT

Usually, pure symbolic methods are scarce or computationally expensive

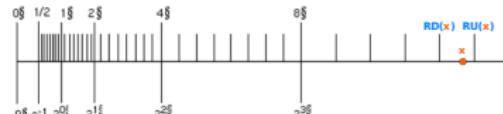
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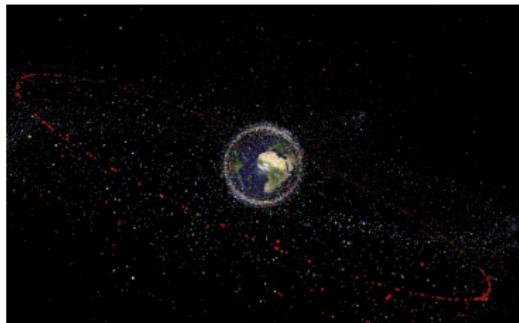
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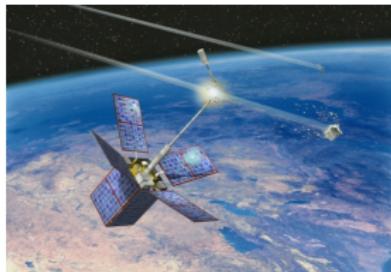


Validated Computing Challenges

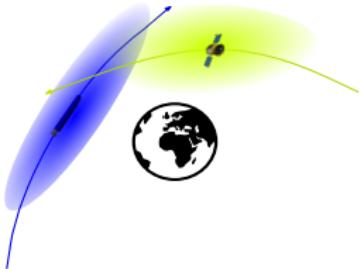
Example: Orbital collision probability evaluation



Space debris population model (source : ESA)



Cerise hit by a debris in 1996 (CNES/D. Ducros)



Conjunction illustration

Thomas Pesquet, when a debris whizzes past the ISS:**

"Climb into an escape shuttle, wait and hope. This happened four times"

**<http://www.chron.com/news/science-environment/article/Thousands-of-tiny-satellites-are-about-to-go-into-11088984.php>

Short-term encounter model and probability of collision

Two objects: primary P (operational satellite) and secondary S (space debris)

High relative velocity

Assumptions:

Rectilinear relative motion

No velocity uncertainty; Gaussian position uncertainty

Infinite encounter time horizon

↪ Probability of collision: 2-D integral over a disk.

Formula

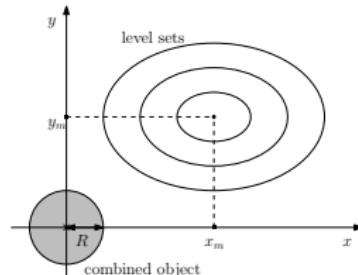
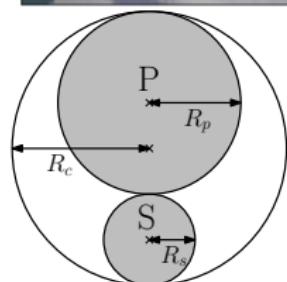
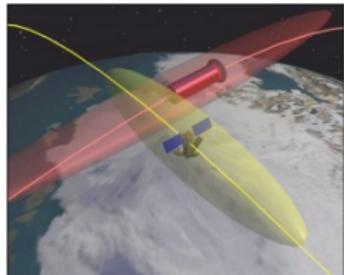
$$\mathcal{P} = \frac{1}{2\pi\sigma_x\sigma_y} \int_{B((0,0), R)} \exp\left(-\frac{(x - x_m)^2}{2\sigma_x^2} - \frac{(y - y_m)^2}{2\sigma_y^2}\right) dx dy,$$

where

R : radius of combined object

x_m, y_m : mean relative coordinates

σ_x, σ_y : standard deviations of relative coordinates



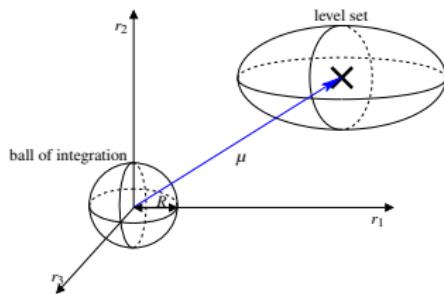
Short-term encounter model and probability of collision

3D generalization: instantaneous collision probability

$$\mathcal{P}_{inst} = \frac{1}{(2\pi)^{3/2} |\Sigma|^{1/2}} \int_{\mathcal{B}(0,R)} \exp\left(-\frac{1}{2}(r - \mu)^T \Sigma^{-1} (r - \mu)\right) dr$$

Need for fast and reliable evaluation:

Simple, numerically stable, double-prec
evaluation & effective error bounds.



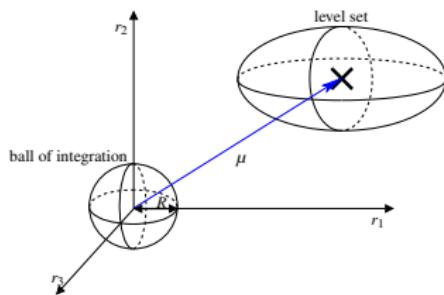
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Previous works:

2D: Numerical integration (Foster '92, Patera '01, Alfano '05);
Power/Hermite series (Pelayo-Ayuso'16), with *trial and error* truncation,
or simplifying assumptions ($\sigma_x = \sigma_y$) (Chan '97);

3D: Equivalent volume –cuboids, approx distribution– (Chan '08, Zhang '20)

Orbital collision probability evaluation

Step 1: Symbolic representation

$$g(z) = \frac{1}{2\pi\sigma_x\sigma_y} \int_{\mathcal{B}((0,0),\sqrt{z})} \exp\left(-\frac{(x - \textcolor{brown}{x}_m)^2}{2\sigma_x^2} - \frac{(y - \textcolor{brown}{y}_m)^2}{2\sigma_y^2}\right) dx dy$$

$$= \sum_{i=0}^{\infty} \frac{\ell_i}{(i+1)!} z^{i+1}$$

Lasserre, Zerron (2001)

Laplace Transform

$$\mathcal{L}(g)(\lambda) = \frac{\exp\left(-\lambda\left(\frac{x_m^2}{2\lambda\sigma_x^2+1} + \frac{y_m^2}{2\lambda\sigma_y^2+1}\right)\right)}{\lambda\sqrt{(2\lambda\sigma_x^2+1)(2\lambda\sigma_y^2+1)}}$$

expansion at ∞

$$= \sum_{i=0}^{\infty} \ell_i \left(\frac{1}{\lambda}\right)^i$$

$\mathcal{L}(g)$ is D-finite – solution of linear ODE with polynomial coefficients –

ℓ_i can be efficiently symbolically described and computed
–solution of linear recurrence with polynomial coefficients–

↔ Gfun Maple package (Salvy, Zimmermann 1994)

Orbital collision probability evaluation

Step 1: Symbolic representation

$$g(z) = \frac{1}{2\pi\sigma_x\sigma_y} \int_{\mathcal{B}((0,0),\sqrt{z})} \exp\left(-\frac{(x - \textcolor{brown}{x}_m)^2}{2\sigma_x^2} - \frac{(y - \textcolor{brown}{y}_m)^2}{2\sigma_y^2}\right) dx dy$$

$$\simeq \sum_{i=0}^N \frac{\ell_i}{(i+1)!} z^{i+1}$$

Lasserre, Zerron (2001)

Laplace Transform

$$\mathcal{L}(g)(\lambda) = \frac{\exp\left(-\lambda\left(\frac{x_m^2}{2\lambda\sigma_x^2+1} + \frac{y_m^2}{2\lambda\sigma_y^2+1}\right)\right)}{\lambda\sqrt{(2\lambda\sigma_x^2+1)(2\lambda\sigma_y^2+1)}}$$

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$$= \sum_{i=0}^{\infty} \ell_i \left(\frac{1}{\lambda}\right)^i$$



Numerical
Issues

Borel

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↔ Gfun Maple package (Salvy, Zimmermann 1994)

Cancellation in finite precision power series evaluation

Example: $\sigma_x = 115, \sigma_y = 1.41, x_m = 0.15, y_m = 3.88, \sqrt{z} = 15$

$$g(z) = \sum_{i=0}^{\infty} \frac{\ell_i}{(i+1)!} z^{i+1}$$

Cancellation in finite precision power series evaluation

Example: $\sigma_x = 115, \sigma_y = 1.41, x_m = 0.15, y_m = 3.88, \sqrt{z} = 15$

$$g(z) = \sum_{i=0}^{\infty} \frac{\ell_i}{(i+1)!} z^{i+1}$$

$$g(225) = 0.16 \cdot 10^{-1} + 1.5 + 16.1 - 250 \dots + 2.2 \cdot 10^{19} - 2.6 \cdot 10^{19} - \dots + 4.3 - 0.14 - 0.60 \dots$$

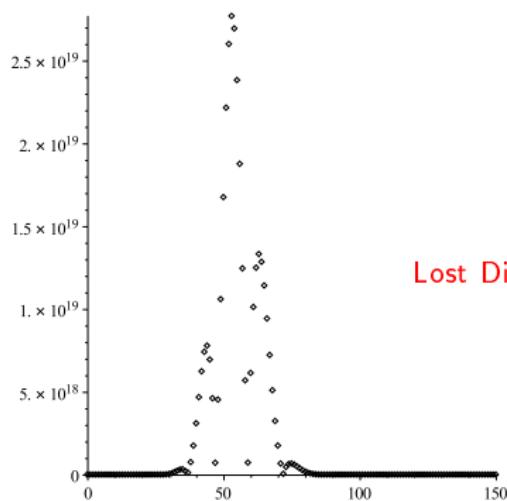
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$$g(225) = 0.16 \cdot 10^{-1} + 1.5 + 16.1 - 250 \dots + 2.2 \cdot 10^{19} - 2.6 \cdot 10^{19} - \dots + 4.3 - 0.14 - 0.60 \dots$$

Values of $\left| \frac{\ell_i 225^{i+1}}{(i+1)!} \right|$, compared to $g(225) \simeq 0.1004$:



$$\text{Lost Digits: } d_g(z) \simeq \log \frac{\max_i |g_i z^i|}{|g(z)|}$$

Cancellation in finite precision power series evaluation

Example: $\exp(-x) = \sum_{i=0}^{\infty} \frac{(-1)^i x^i}{i!}$

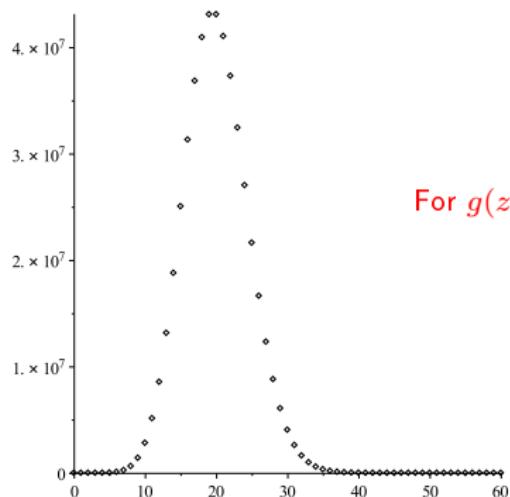
$$\exp(-20) = 1 - 20 \dots + 1.66 \cdot 10^7 - 1.23 \cdot 10^7 + \dots + 1.19 \cdot 10^{-8} - 3.45 \cdot 10^{-9} \dots$$

Cancellation in finite precision power series evaluation

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Values of $\left| \frac{(-1)^i 20^i}{i!} \right|$, compared to $\exp(-20) \simeq 2.06 \cdot 10^{-9}$:



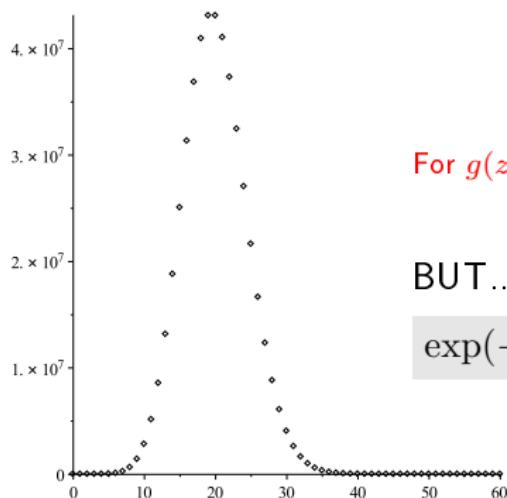
For $g(z) = \sum_{i=0}^{\infty} g_i z^i$, lost digits $\simeq \log \frac{\max_i |g_i z^i|}{|g(z)|}$

Cancellation in finite precision power series evaluation

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Values of $\left| \frac{(-1)^i 20^i}{i!} \right|$, compared to $\exp(-20) \simeq 2.06 \cdot 10^{-9}$:



For $g(z) = \sum_{i=0}^{\infty} g_i z^i$, lost digits $\simeq \log \frac{\max_i |g_i z^i|}{|g(z)|}$

BUT...

$$\exp(-x) = \frac{1}{\exp(x)}$$

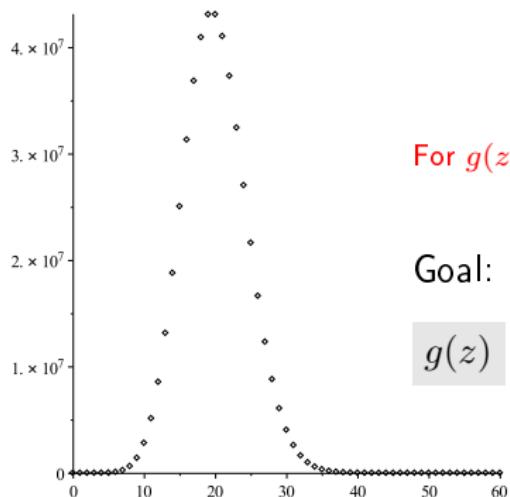
No cancellation!

Cancellation in finite precision power series evaluation

Example: $\exp(-x) = \sum_{i=0}^{\infty} \frac{(-1)^i x^i}{i!}$

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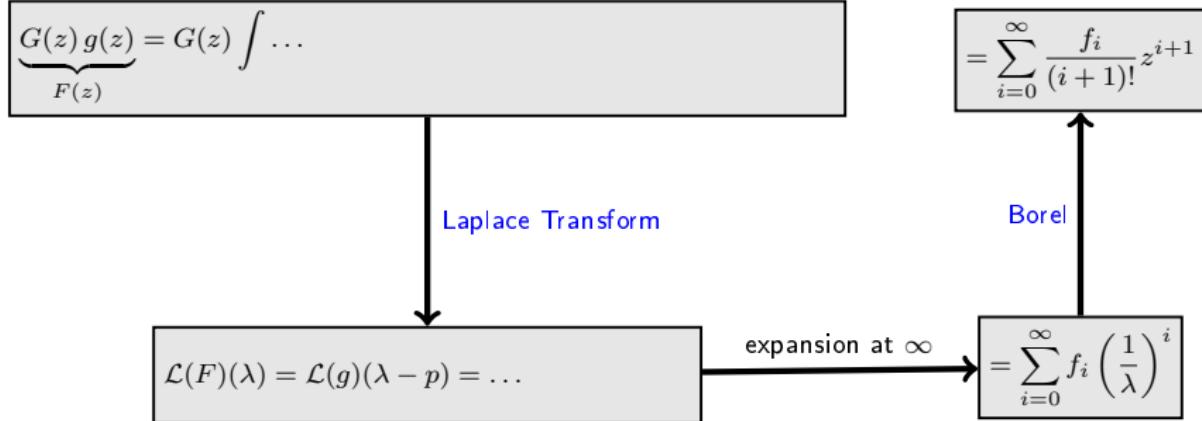
Goal:

$$g(z) = \frac{F(z)}{G(z)}$$

No cancellation!

Orbital collision probability evaluation

Step 2: Reliable numerics



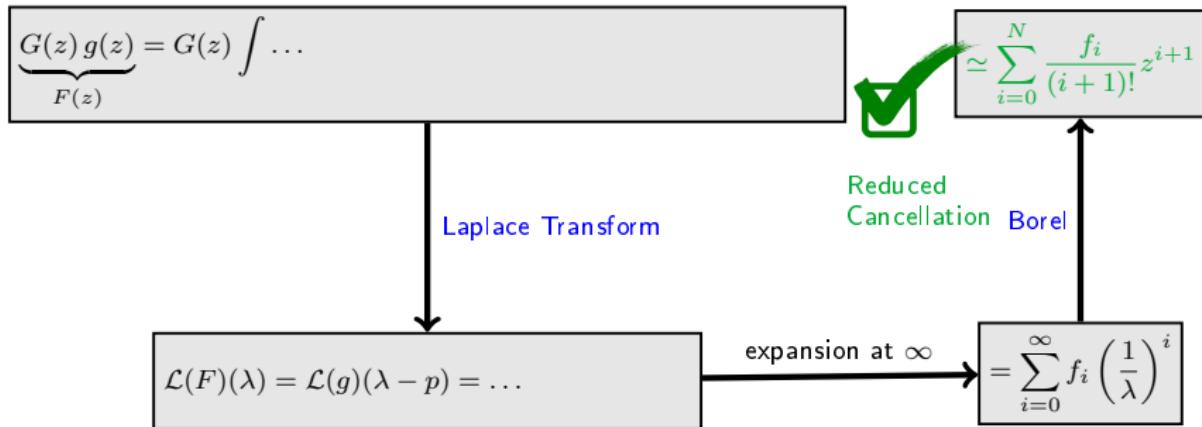
Gawronski, Müller, Reinhard (2007): choose $G(z) = \exp(pz)$, with $p \sim \frac{1}{2\sigma_y^2}$.

$\mathcal{L}(F)$ is D-finite

f_i can be efficiently symbolically described and computed
—solution of linear recurrence with polynomial coefficients—

Orbital collision probability evaluation

Step 2: Reliable numerics



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$\mathcal{L}(F)$ is D-finite

f_i can be efficiently symbolically described and computed
—solution of linear recurrence with polynomial coefficients—

f_i are positive \leadsto Reduced Cancellation

Orbital collision probability evaluation

Step 2: Reliable numerics

Positivity:

Consider $\bar{\varphi}(z) = \sum_{k=0}^{\infty} f_k z^k$, for $|z| \leq p^{-1}$, which is D-finite

$$\begin{aligned}\bar{\varphi}'(z) &= \underbrace{\frac{P(z)}{Q(z)}}_{\bar{\varphi}(z)}, \quad \bar{\varphi}(0) = C > 0, \\ &\sum_{k=0}^{\infty} \beta_k z^k\end{aligned}$$

with $\beta_k = p^{k+1} + \dots > 0$, hence

$$(i+1)f_{i+1} = \sum_{k=0}^i \beta_k f_{i-k} > 0.$$

Orbital collision probability evaluation

Step 2: Reliable numerics

$$F(z) = \sum_{i=0}^{\infty} \frac{f_i}{(i+1)!} z^{i+1}$$

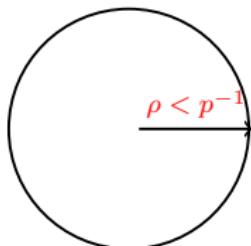
Borel

$$\bar{\varphi}(z) = \sum_{i=0}^{\infty} f_i z^i$$

Tail bounds:

$$\sum_{k=0}^{\infty} \frac{f_{k+n} z^{k+n}}{(k+n+1)!}$$

Closed-form; Conv. radius p^{-1}



$$\leq \frac{(\frac{z}{\rho})^n}{(n+1)!} \bar{\varphi}(\rho)$$

for any $\rho < p^{-1}$ and $n \geq n_0$, with $n_0 = \left\lceil \frac{z}{\rho} \right\rceil$.

Orbital collision probability evaluation

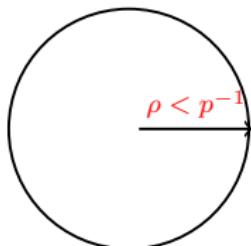
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Closed-form; Conv. radius p^{-1}



Tail bounds:

$$\sum_{k=0}^{\infty} \frac{f_{k+n} z^{k+n}}{(k+n+1)!} \leq \frac{\left(\frac{z}{\rho}\right)^n}{(n+1)!} \sum_{k=0}^{\infty} f_{k+n} z^k \rho^n \underbrace{\frac{(n+1)!}{(k+n+1)!}}_{\leq 1/(n+1)^k}$$

$$\leq \frac{\left(\frac{z}{\rho}\right)^n}{(n+1)!} \sum_{k=0}^{\infty} f_{k+n} \rho^{k+n} \underbrace{\frac{z^k}{\rho^k (n+1)^k}}_{\leq 1}$$

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Orbital collision probability evaluation

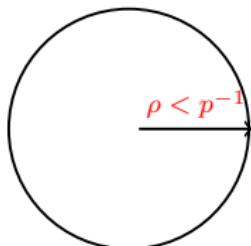
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$$\leq \frac{\left(\frac{z}{\rho}\right)^n}{(n+1)!} \bar{\varphi}(\rho)$$

for any $\rho < p^{-1}$ and $n \geq n_0$, with $n_0 = \left\lceil \frac{z}{\rho} \right\rceil$.

In particular, set $\rho = p^{-1}/2$ and $z = R^2$.

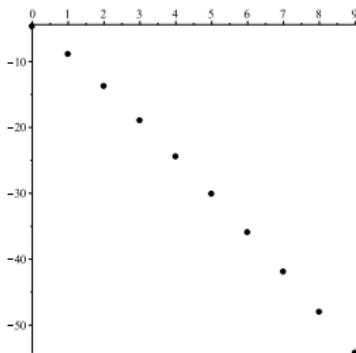
Examples

Case #	Input parameters (km)				
	σ_x	σ_y	R	x_m	y_m
1	0.05	0.025	0.005	0.01	0
2	0.05	0.025	0.005	0	0.01
3	0.075	0.025	0.005	0.01	0
4	0.075	0.025	0.005	0	0.01
5	3	1	0.01	1	0
6	3	1	0.01	0	1
7	3	1	0.01	10	0
8	3	1	0.01	0	10
9	10	1	0.01	10	0
10	10	1	0.01	0	10
11	3	1	0.05	5	0
12	3	1	0.05	0	5

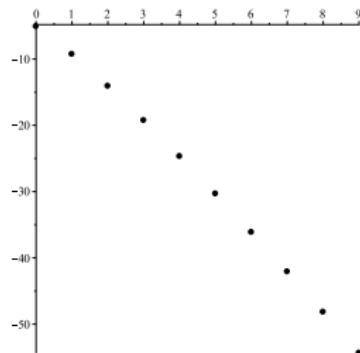
Examples

Examples: accuracy η and log plot of terms in the series

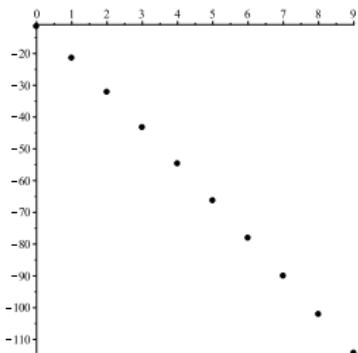
Case 1: $\eta = 23$



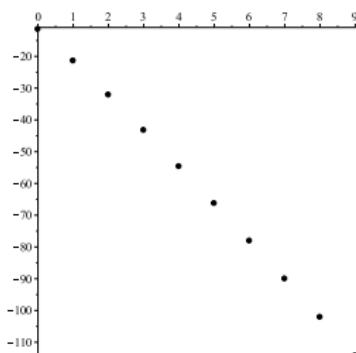
Case 2: $\eta = 22$



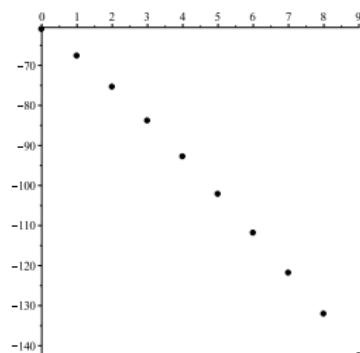
Case 4: $\eta = 22$



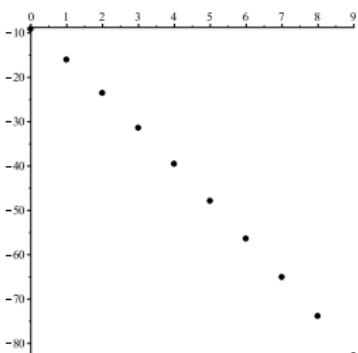
Case 6: $\eta = 47$



Case 8: $\eta = 33$



Case 11: $\eta = 34$



Example: Orbital collision probability evaluation

Sum-up: Fast and reliable algorithm

Efficient and reliable algorithm (2016)*

Error bounds

Linear complexity

Algorithm 12 Approximate evaluation of \bar{P}_c .

Input: Parameters: $\sigma_x, \sigma_y, x_m, y_m, R$; Number of terms: N .

Output: \bar{P}_c – truncated series evaluation of P_c .

1: $p = \frac{1}{2\sigma_x^2}, \varphi = 1 - \frac{\sigma_y^2}{\sigma_x^2}; \omega_x = \frac{x_m^2}{4\sigma_x^2}, \omega_y = \frac{y_m^2}{4\sigma_y^2}; \alpha_0 = \frac{1}{2\sigma_x\sigma_y} \exp\left(-\frac{1}{2}\left(\frac{x_m^2}{\sigma_x^2} + \frac{y_m^2}{\sigma_y^2}\right)\right);$

2: $c_0 = \alpha_0 R^2;$

3: $c_1 = \frac{\alpha_0 R^4}{2} (p\left(\frac{\varphi}{2} + 1\right) + \omega_x + \omega_y);$

4: $c_2 = \frac{\alpha_0 R^6}{12} \left((p\left(\frac{\varphi}{2} + 1\right) + \omega_x + \omega_y)^2 + p^2\left(\frac{\varphi^2}{2} + 1\right) + 2p\varphi\omega_x \right);$

5: $c_3 = \frac{\alpha_0 R^8}{144} \left((p\left(\frac{\varphi}{2} + 1\right) + \omega_x + \omega_y)^3 + 3(p\left(\frac{\varphi}{2} + 1\right) + \omega_x + \omega_y) (p^2\left(\frac{\varphi^2}{2} + 1\right) + 2p\varphi\omega_x) \right. \\ \left. + 2(p^3\left(\frac{\varphi^3}{2} + 1\right) + 3p^2\varphi^2\omega_x) \right);$

6: for $k \leftarrow 0$ to $N - 5$ do

$$c_{k+4} = -\frac{R^6 p^5 \varphi^2 \omega_y}{(k+2)(k+3)(k+4)^2(k+5)} c_k + \frac{R^6 p^3 \varphi (p \varphi (k+\frac{5}{2}) + 2\omega_y (\frac{\varphi}{2} + 1))}{(k+3)(k+4)^2(k+5)} c_{k+1}$$

7: $- \frac{R^4 p (p \varphi (\frac{\varphi}{2} + 1) (2k+5) + \varphi (2\omega_y + \frac{3\varphi}{2}) + \omega_x + \omega_y)}{(k+4)^3(k+5)} c_{k+2}$

+ $\frac{R^2 (p(2\varphi+1)(k+3) + p(\frac{\varphi}{2} + 1) + \omega_x + \omega_y)}{(k+4)(k+5)} c_{k+3}$

8: end for

9: $s \leftarrow 0$

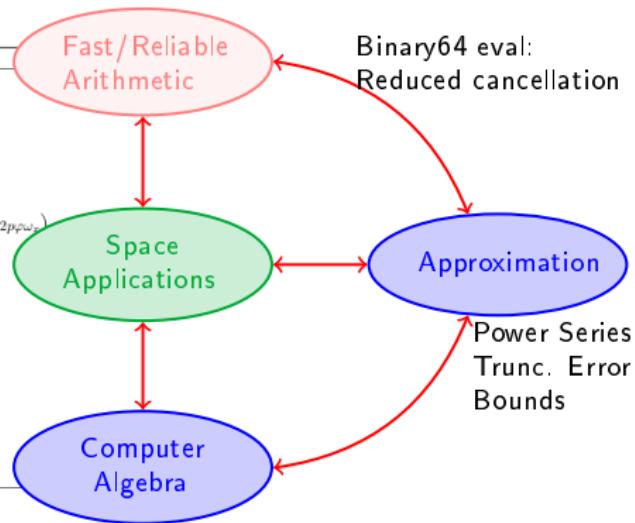
10: for $k \leftarrow 0$ to $N - 1$ do

11: $s \leftarrow s + c_k;$

12: end for

13: $\bar{P}_c \leftarrow \exp(-pR^2) s;$

14: return \bar{P}_c .



Algorithms
D-finite functions
 $\exp, \text{erf}, \sin, \text{Gaussian}$

* JGCD, joint work with R. Serra, D. Arzelier, A. Rondepierre, J.-B. Lasserre, B. Salvy

Example: Orbital collision probability evaluation

Sum-up: Fast and reliable algorithm

Efficient and reliable algorithm (2016)*

Error bounds

Linear complexity

Algorithm 12 Approximate evaluation of \bar{P}_c .

Input: Parameters: $\sigma_x, \sigma_y, x_m, y_m, R$; Number of terms: N .

Output: \bar{P}_c - truncated series evaluation of P_c .

$$1: p = \frac{1}{2\sigma_x^2}, \varphi = 1 - \frac{\sigma_y^2}{\sigma_x^2}; \omega_x = \frac{x_m^2}{4\sigma_x^2}, \omega_y = \frac{y_m^2}{4\sigma_y^2}, \alpha_0 = \frac{1}{2\sigma_x\sigma_y} \exp\left(-\frac{1}{2}\left(\frac{x_m^2}{\sigma_x^2} + \frac{y_m^2}{\sigma_y^2}\right)\right);$$

$$2: c_0 = \alpha_0 R^2;$$

$$3: c_1 = \frac{\alpha_0 R^4}{2} \left(p\left(\frac{\varphi}{2} + 1\right) + \omega_x + \omega_y\right);$$

$$4: c_2 = \frac{\alpha_0 R^6}{12} \left(\left(p\left(\frac{\varphi}{2} + 1\right) + \omega_x + \omega_y\right)^2 + p^2\left(\frac{\varphi^2}{2} + 1\right) + 2p\varphi\omega_x\right);$$

$$5: c_3 = \frac{\alpha_0 R^8}{144} \left(\left(p\left(\frac{\varphi}{2} + 1\right) + \omega_x + \omega_y\right)^3 + 3\left(p\left(\frac{\varphi}{2} + 1\right) + \omega_x + \omega_y\right)\left(p^2\left(\frac{\varphi^2}{2} + 1\right) + 2p\varphi\omega_x\right) + 2\left(p^3\left(\frac{\varphi^3}{2} + 1\right) + 3p^2\varphi^2\omega_x\right)\right);$$

$$6: \text{for } k \leftarrow 0 \text{ to } N - 5 \text{ do}$$

$$c_{k+4} = -\frac{R^6 p^5 \varphi^2 \omega_y}{(k+2)(k+3)(k+4)^2(k+5)} c_k + \frac{R^6 p^3 \varphi \left(p\varphi\left(k+\frac{5}{2}\right) + 2\omega_y\left(\frac{\varphi}{2} + 1\right)\right)}{(k+3)(k+4)^2(k+5)} c_{k+1}$$

$$7: -\frac{R^4 p \left(p\varphi\left(\frac{\varphi}{2} + 1\right) (2k+5) + \varphi\left(2\omega_y + \frac{5}{2}\right) + \omega_x + \omega_y\right)}{(k+4)^3(k+5)} c_{k+2} + \frac{R^2 \left(p(2\varphi+1)(k+3) + p\left(\frac{\varphi}{2} + 1\right) + \omega_x + \omega_y\right)}{(k+4)(k+5)} c_{k+3}$$

$$8: \text{end for}$$

$$9: s \leftarrow 0$$

$$10: \text{for } k \leftarrow 0 \text{ to } N - 1 \text{ do}$$

$$11: s \leftarrow s + c_k;$$

$$12: \text{end for}$$

$$13: \bar{P}_c \leftarrow \exp(-pR^2) s;$$

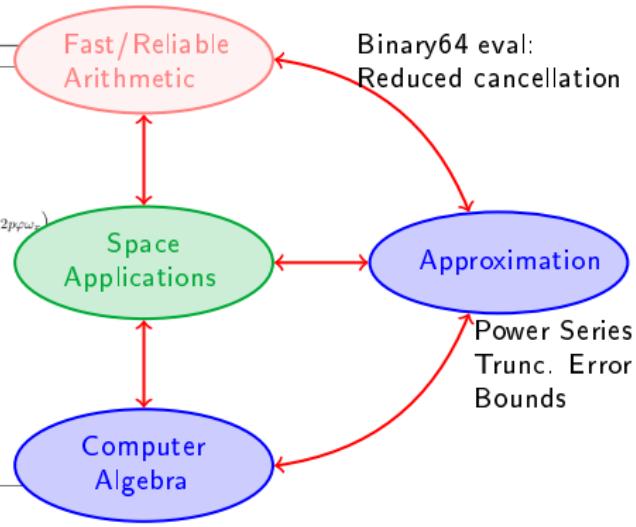
$$14: \text{return } \bar{P}_c.$$

2017: CNES impl+tests

2021: CNES on-board impl**

Long-term collaboration

ADS, CNES, Thales, TAS



Algorithms
D-finite functions
 $\exp, \text{erf}, \sin, \text{Gaussian}$

* JGCD, joint work with R. Serra, D. Arzelier, A. Rondepierre, J.-B. Lasserre, B. Salvy

** 01/2023 <https://lejournal.cnrs.fr/articles/un-algorithme-pour-eviter-les-debris-spatiaux>

Trickier examples

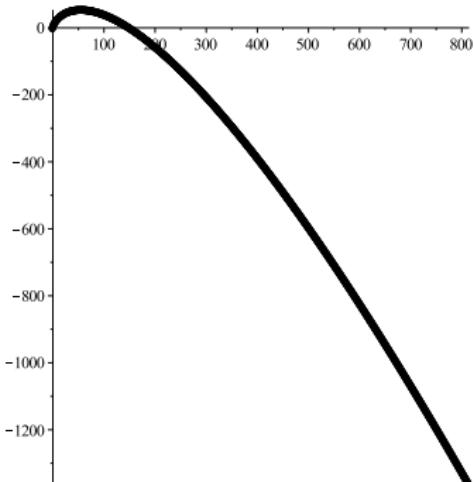
(from [Alfano 2009])

Case #	Input parameters (m)				
	σ_x	σ_y	R	x_m	y_m
3	114.25	1.41	15	0.15	3.88
5	177.8	0.038	10	2.12	-1.22

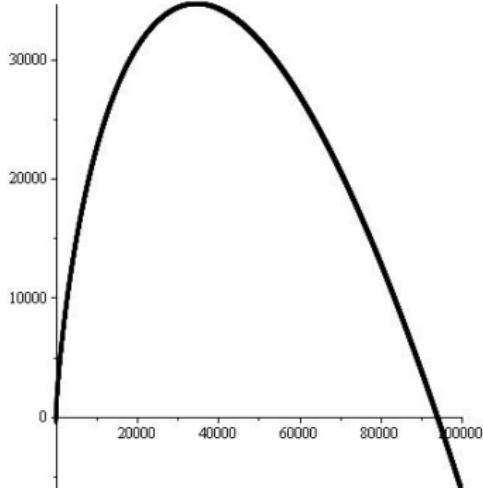
Alfano's test case number	Probability of collision (-)		
	Alfano	New method	Reference (MC)
3	0.10038	0.10038	0.10085
5	0.044712	0.045509	0.044499

Examples: quality η and log plot of terms in the series:

Case 3 Alfano: $\eta_{800} = 30$



Case 5 Alfano: $\eta_{121000} = 20$



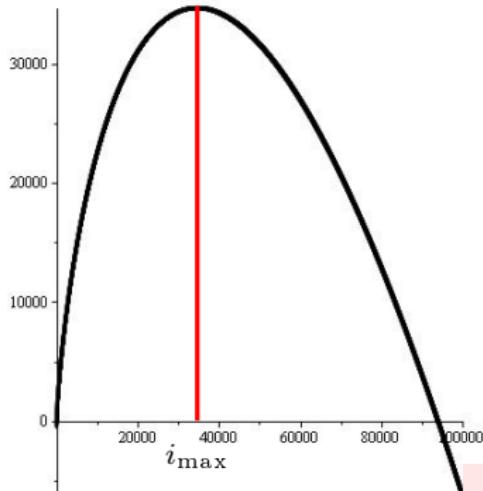
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3	114.25	1.41	15	0.15	3.88
5	177.8	0.038	10	2.12	-1.22

(from [Alfano 2009])

An annoying bump...

$$\log \left(\frac{f_i R^{2i}}{(i+1)!} \right)$$

$$i_{\max} \sim O \left(\frac{R^2}{2\sigma_y^2} \right)$$



Term index i

Case #	Input parameters (m)				
	σ_x	σ_y	R	x_m	y_m
5	177.8	0.038	10	2.12	-1.22

$$i_{\max} \simeq 34626$$

An annoying bump...



An alternative via Saddle-Point

Recall that

$$\mathcal{L}g(\lambda) = \frac{\exp\left(-\frac{\sigma_x^2 y_m^2 + \sigma_y^2 x_m^2}{2\sigma_x^2 \sigma_y^2} + \frac{y_m^2}{2\sigma_y^2(2\lambda\sigma_y^2 + 1)} + \frac{x_m^2}{2\sigma_x^2(2\lambda\sigma_x^2 + 1)}\right)}{\lambda \sqrt{(2\lambda\sigma_x^2 + 1)(2\lambda\sigma_y^2 + 1)}}$$

Inverse Laplace Transform

$$\mathcal{I} = \frac{1}{2i\pi} \int_{\lambda_0 - i\infty}^{\lambda_0 + i\infty} \underbrace{e^{\xi\lambda} \mathcal{L}g(\lambda)}_{e^{\varphi(\lambda)}} d\lambda$$

Integrand $e^{\varphi(\lambda_0 + it)}$:

An alternative via Saddle-Point

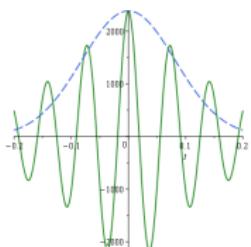
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Inverse Laplace Transform

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Integrand $e^{\varphi(\lambda_0 + it)}$:



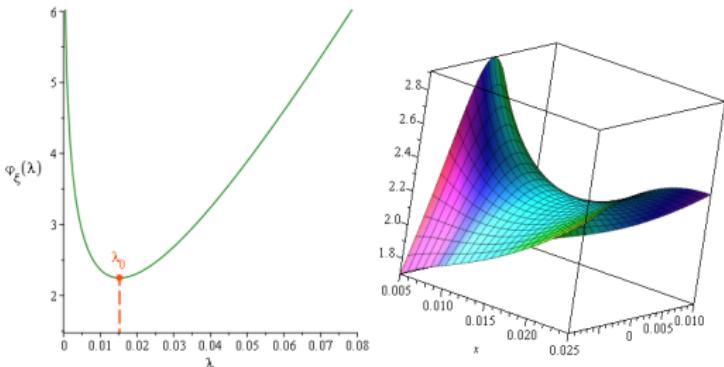
An alternative via Saddle-Point

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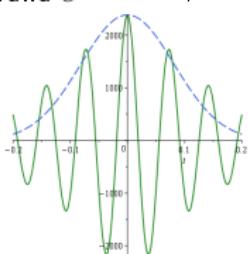
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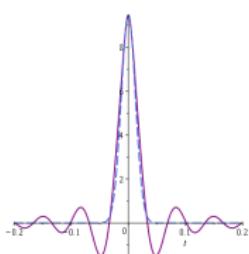


Saddle-point $\lambda_0 > 0$ solution of $\varphi'(\lambda_0) = 0$

Integrand $e^{\varphi(\lambda_0 + it)}$:



generic λ_0



λ_0 is a saddle

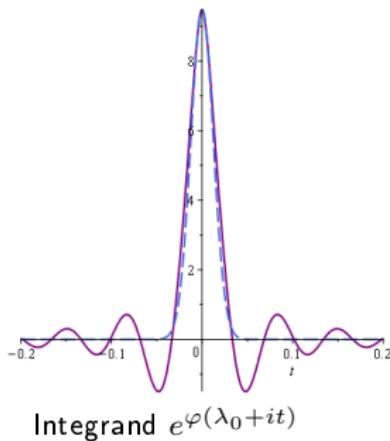
Laplace method in a nutshell

1. Neglect the tails

$$\mathcal{I} \simeq \frac{1}{2i\pi} \int_{\lambda_0 - is}^{\lambda_0 + is} e^{\varphi(\lambda)} d\lambda$$

2. Central approximation

$$\mathcal{I} \simeq \frac{1}{2i\pi} \int_{\lambda_0 - is}^{\lambda_0 + is} e^{\varphi(\lambda_0) + \frac{\varphi''(\lambda_0)}{2}(\lambda - \lambda_0)^2} d\lambda$$



3. Complete the tails

$$\mathcal{I} \simeq \frac{1}{2i\pi} \int_{\lambda_0 - i\infty}^{\lambda_0 + i\infty} e^{\varphi(\lambda_0) + \frac{\varphi''(\lambda_0)}{2}(\lambda - \lambda_0)^2} d\lambda$$

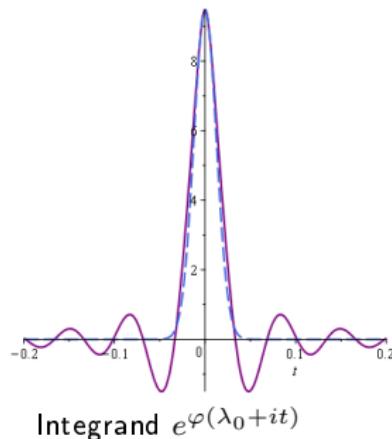
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Integrand $e^{\varphi(\lambda_0 + it)}$

3. Complete the tails

$$\mathcal{I} \simeq \frac{1}{2i\pi} \int_{\lambda_0 - i\infty}^{\lambda_0 + i\infty} e^{\varphi(\lambda_0) + \frac{\varphi''(\lambda_0)}{2}(\lambda - \lambda_0)^2} d\lambda = \frac{1}{2\pi} \int_{-\infty}^{+\infty} e^{\varphi(\lambda_0) - \frac{\varphi''(\lambda_0)}{2}t^2} dt$$

$$\mathcal{I} \simeq \frac{e^{\varphi(\lambda_0)}}{2\sqrt{\varphi''(\lambda_0)\pi}}$$

Not sufficiently accurate

Can we do better?

Asymptotic expansions of Laplace type integrals

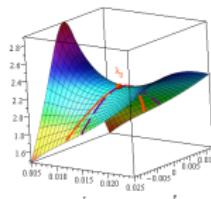
1. Neglect the tails

$$\mathcal{I} \simeq \frac{1}{2i\pi} \int_{\lambda_0 - is}^{\lambda_0 + is} e^{\varphi(\lambda)} d\lambda$$

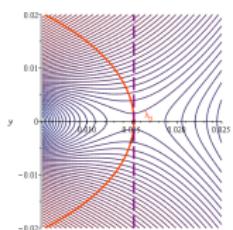
2. Change of variables in the neighborhood of λ_0 :

$$\varphi(\lambda(w)) = \varphi(\lambda_0) + \frac{\varphi''(\lambda_0)}{2} (iw)^2$$

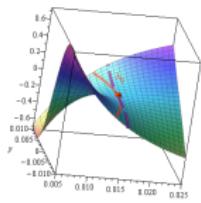
+ Local deformation of the integration path to match the constant phase contour passing through λ_0



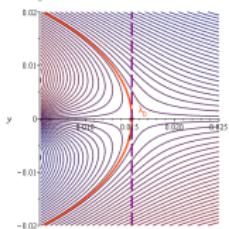
Surface $(x, y) \mapsto u_{\lambda_0}(x + iy)$



Level set u_{λ_0}



Surface $(x, y) \mapsto v_{\lambda_0}(x + iy)$



Level set v_{λ_0}

Asymptotic expansions of Laplace type integrals

1. Neglect the tails

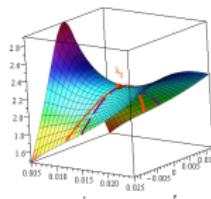
$$\mathcal{I} \simeq \frac{1}{2i\pi} \int_{\lambda_0-is}^{\lambda_0+is} e^{\varphi(\lambda)} d\lambda$$

2. Change of variables in the neighborhood of λ_0 :

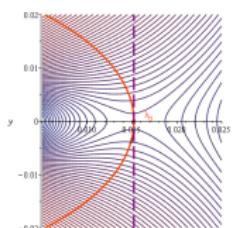
$$\varphi(\lambda(w)) = \varphi(\lambda_0) + \frac{\varphi''(\lambda_0)}{2}(iw)^2$$

3. Complete the tails

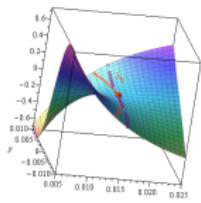
$$\mathcal{I} \simeq \frac{1}{2i\pi} \int_{-\infty}^{+\infty} e^{\varphi(\lambda_0) - \frac{\varphi''(\lambda_0)}{2}w^2} \frac{d\lambda(w)}{dw} dw$$



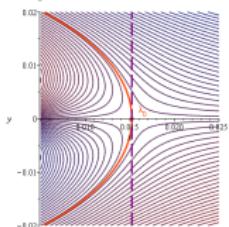
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Asymptotic expansions of Laplace type integrals

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$$\mathcal{I} \simeq \frac{1}{2i\pi} \int_{-\infty}^{+\infty} e^{\varphi(\lambda_0) - \frac{\varphi''(\lambda_0)}{2}w^2} \frac{d\lambda(w)}{dw} dw$$

1. Find Saddle-point $\lambda_0 > 0$, numerical solution of:

$$\varphi'(\lambda) = 0$$

2. Series Inversion:

$$\lambda(w) = \lambda_0 + iw + b_2(iw)^2 + b_3(iw)^3 + \dots$$

Expand in w , identify coeffs and solve for b_n .

Asymptotic expansions of Laplace type integrals

1. Neglect the tails

$$\mathcal{I} \simeq \frac{1}{2i\pi} \int_{\lambda_0-is}^{\lambda_0+is} e^{\varphi(\lambda)} d\lambda$$

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Asymptotic expansions of Laplace type integrals

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$$\mathcal{I} \simeq \frac{1}{2i\pi} \int_{\lambda_0-is}^{\lambda_0+is} e^{\varphi(\lambda)} d\lambda$$

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3. Complete the tails

$$\mathcal{I} \simeq \frac{1}{2i\pi} \int_{-\infty}^{+\infty} e^{\varphi(\lambda_0) - \frac{\varphi''(\lambda_0)}{2}w^2} \frac{d\lambda(w)}{dw} dw$$

$$\sum_{n=0}^{\infty} (n+1)b_{n+1}(iw)^n$$

1. Find Saddle-point $\lambda_0 > 0$, numerical solution of:

$$\varphi'(\lambda) = 0$$

2. Series Inversion:

$$\lambda(w) = \lambda_0 + iw + b_2(iw)^2 + b_3(iw)^3 + \dots$$

Expand in w , identify coeffs and solve for b_n .

3. Term-by-term integration:

$$\mathcal{I} \simeq \frac{e^{\varphi(\lambda_0)}}{2\sqrt{\varphi''(\lambda_0)\pi}} \sum_{n=0}^N c_n,$$

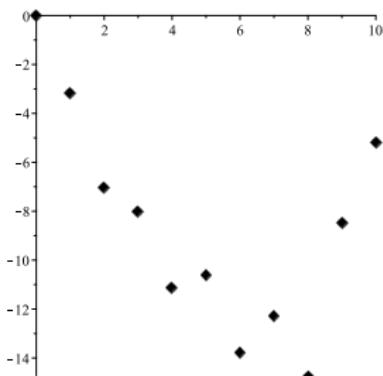
$$\text{with } c_n = (-1)^n \frac{(2n+1)!!}{(\varphi''(\lambda_0))^n} b_{2n+1}$$

Asymptotic expansions of Laplace type integrals

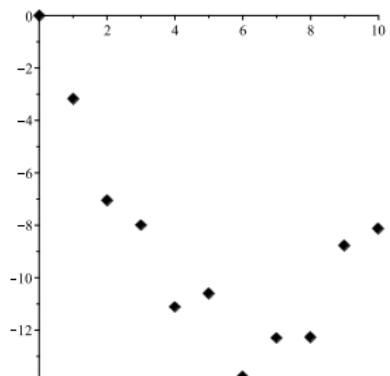
Examples (from 2D)

log plot of c_n coefficients

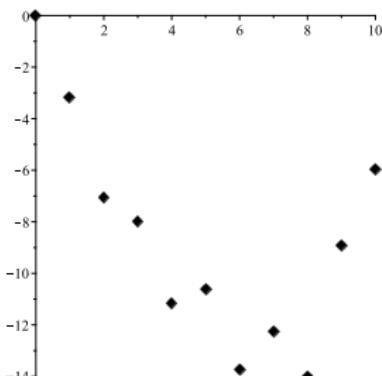
Case 2:



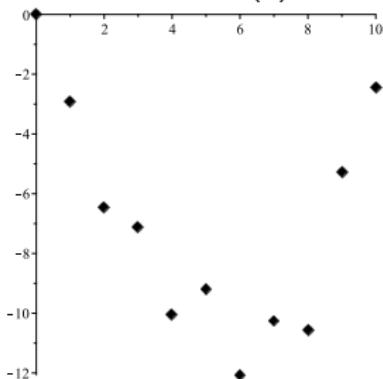
Case 3:



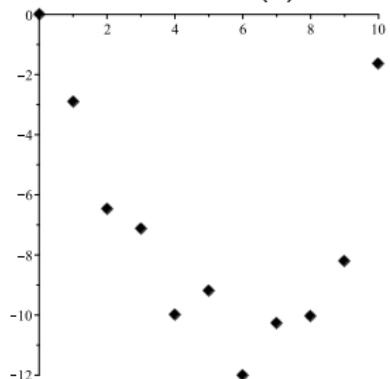
Case 4:



Case Alfano(3):



Case Alfano(5)



Accuracy with 5 terms:

- Case2: 3 digits
- Case3: 2 digits
- Case6: 4 digits
- CaseA3: 3 digits
- CaseA5: 3 digits

Two Satellites in Geosynchronous Equatorial Orbits (GEO)

Instantaneous (3D) probability by orbit propagation from $t_0 = 0$ to $t_f = 13200s$ on a grid of 301 points; accuracy threshold $\delta = 10^{-15}$.

Convergent series used if no. terms $\leq N_{max}$, otherwise saddle-point alternative.

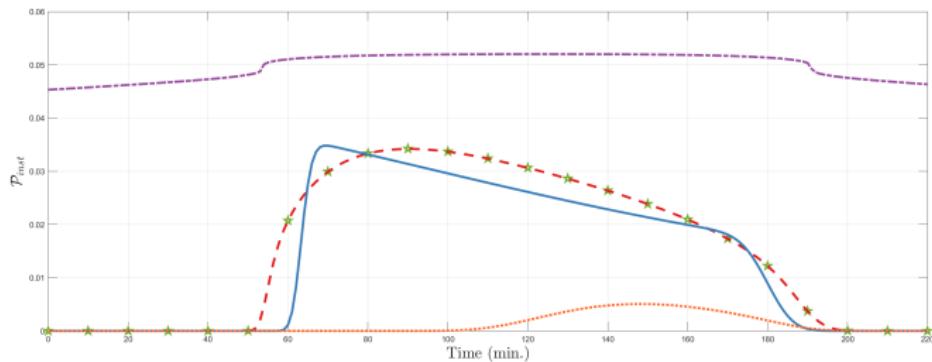


Figure: Alfano'09 test case 4: P_{inst} - Monte Carlo simulations (green stars), our algorithm (red dashed line), alternative methods AD (dashed dot purple line), EV (dotted orange line), EVC (solid blue line).

Two Satellites in Geosynchronous Equatorial Orbits (GEO) –continued–

Instantaneous (3D) probability by orbit propagation from $t_0 = 0$ to $t_f = 13200\text{s}$ on a grid of 301 points; accuracy threshold $\delta = 10^{-15}$.

Convergent series used if no. terms $\leq N_{max}$, otherwise saddle-point alternative.

N_{max}	Mean/median timings (s)	% calls of divergent series
10	0.2153/0.2509	100
400	0.2350/0.2705	52
1000	0.2495/0.2945	32.1
4000	0.3804/0.4503	0

Table: Mean/median timings (s) over 30 runs and percentages of calls of the divergent series.

Saddle-point method is faster, but **currently** lacks error bounds...

Asymptotic expansions of Laplace type integrals: error bounds

1. Neglect the tails

$$\mathcal{I} \simeq \frac{1}{2i\pi} \int_{\lambda_0-is}^{\lambda_0+is} e^{\varphi(\lambda)} d\lambda$$

1. Bound tail

2. Change of variables in the neighborhood of λ_0 :

$$\varphi(\lambda(w)) = \varphi(\lambda_0) + \frac{\varphi''(\lambda_0)}{2}(iw)^2$$

2. Series Inversion:
Bound truncation error

3. Complete the tails

$$\mathcal{I} \simeq \frac{1}{2i\pi} \int_{-\infty}^{+\infty} e^{\varphi(\lambda_0) - \frac{\varphi''(\lambda_0)}{2}w^2} \frac{d\lambda(w)}{dw} dw$$

3. Term-by-term integration:
Bound extra added tails

Note: see works by e.g., Olver ('60s), Wong ('80s) for some special functions

Other challenges

Long-term encounters

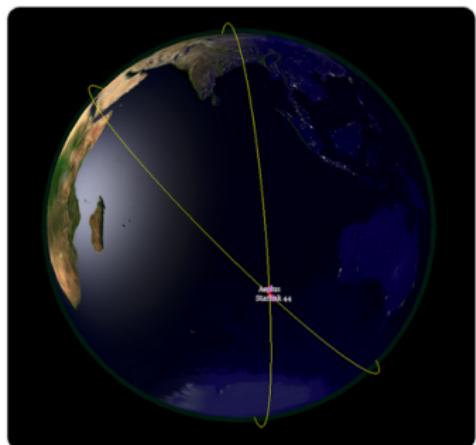
Multiple encounters

The issue of mega-constellations



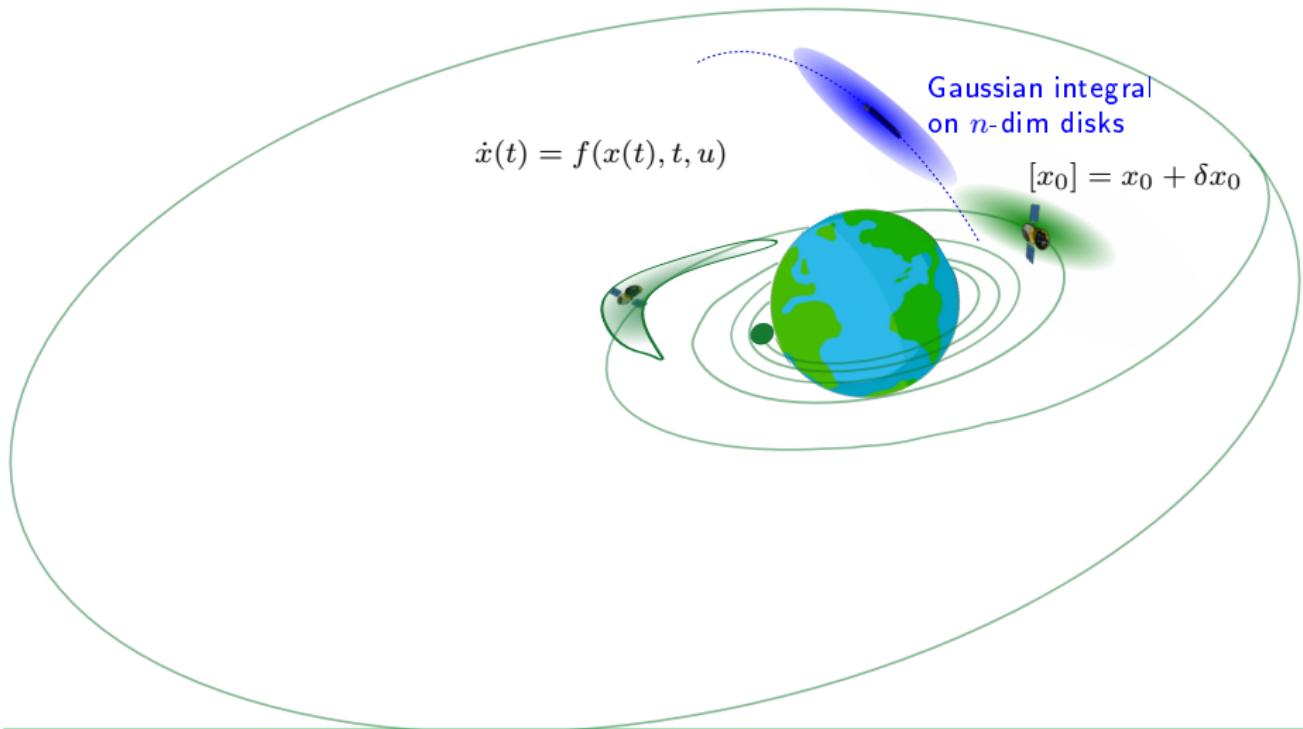
ESA Operations @esaoperations · Sep 2, 2019

For the first time ever, ESA has performed a 'collision avoidance manoeuvre' to protect one of its satellites from colliding with a 'mega constellation'
#SpaceTraffic



Other challenges:

Fast and Reliable Computations for Space – Exemplified Roadmap

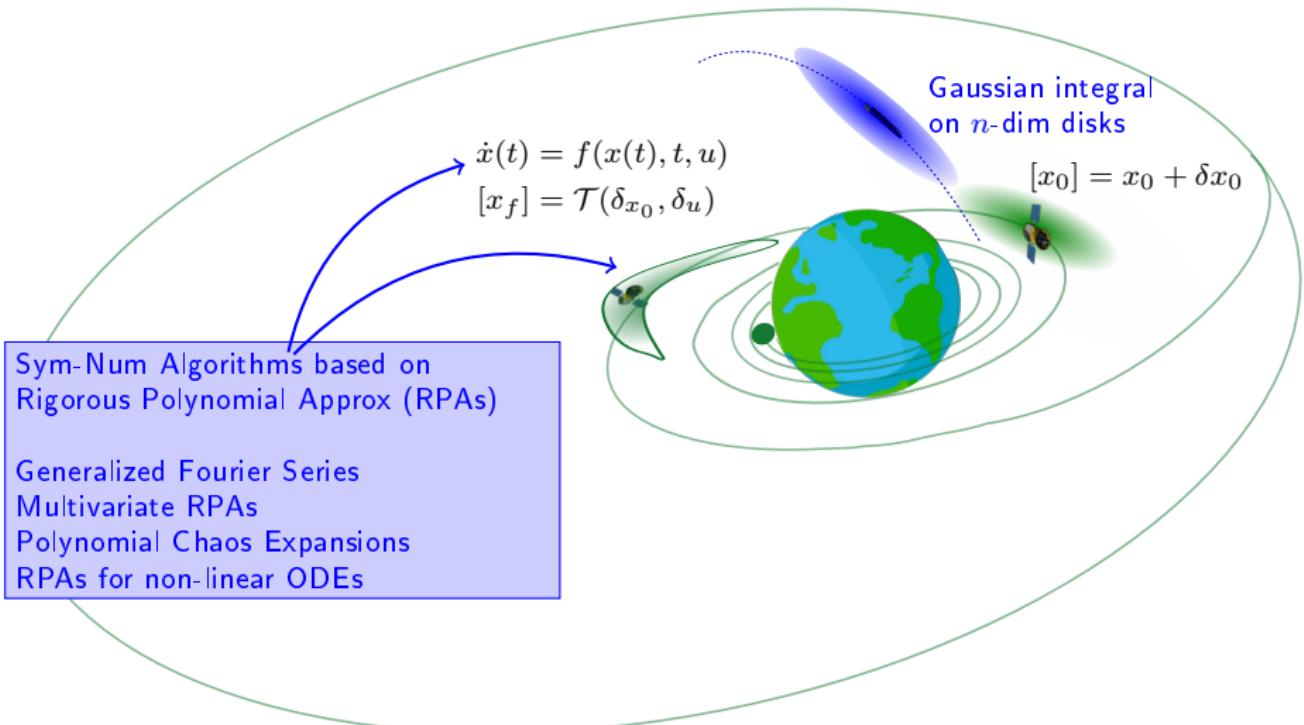


Applications: Validated Autonomous Electric Orbit Raising
Uncertainty Propagation
Collision Risk Assessment/Mitigation, Mega-constellations



Other challenges:

Fast and Reliable Computations for Space – Exemplified Roadmap

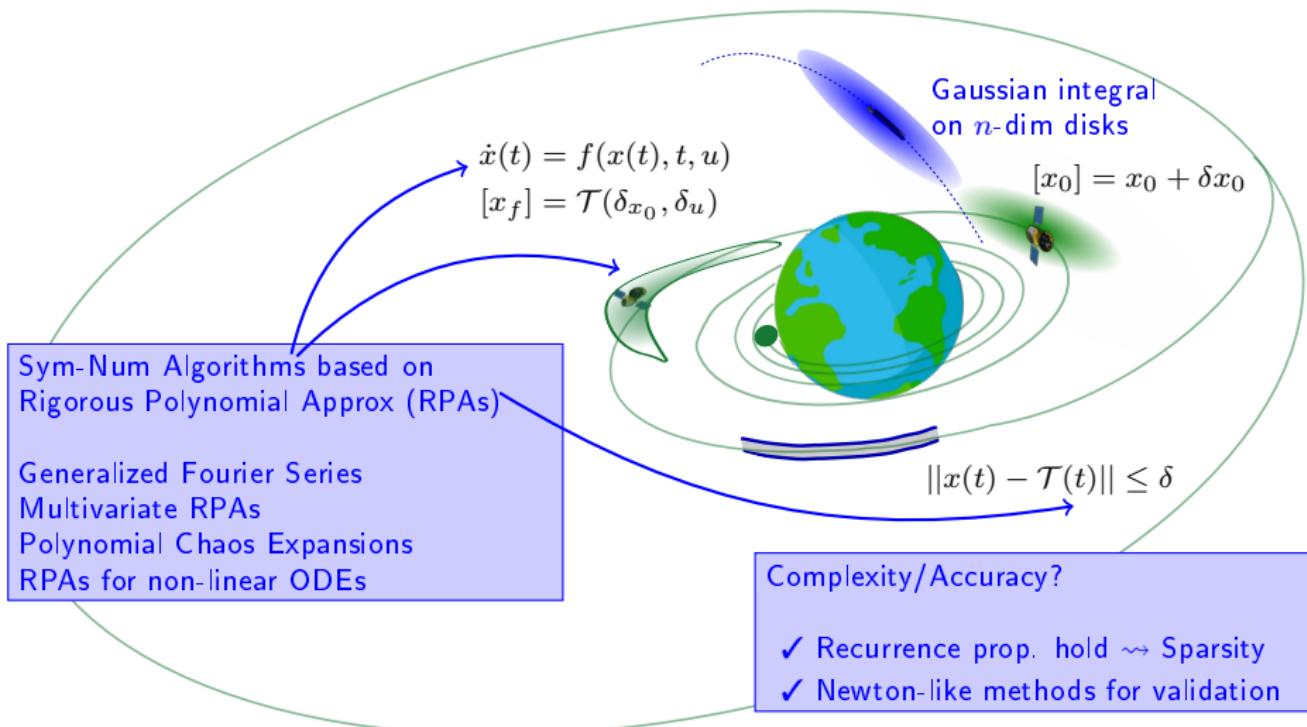


Applications: Validated Autonomous Electric Orbit Raising
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Other challenges:

Fast and Reliable Computations for Space – Exemplified Roadmap



Applications: Validated Autonomous Electric Orbit Raising
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Other challenges:

Fast and Reliable Computations for Space – Exemplified Roadmap

Arith:

Finite prec evaluation of RPAs

$$\dot{x}(t) = f(x(t), t, u)$$
$$[x_f] = \mathcal{T}(\delta_{x_0}, \delta_u)$$

Sym-Num Algorithms based on
Rigorous Polynomial Approx (RPAs)

Generalized Fourier Series

Multivariate RPAs

Polynomial Chaos Expansions

RPAs for non-linear ODEs

Gaussian integral
on n -dim disks

$$[x_0] = x_0 + \delta x_0$$

$$\|x(t) - \mathcal{T}(t)\| \leq \delta$$

Complexity/Accuracy?

- ✓ Recurrence prop. hold \leadsto Sparsity
- ✓ Newton-like methods for validation

Applications: Validated Autonomous Electric Orbit Raising
Uncertainty Propagation
Collision Risk Assessment/Mitigation, Mega-constellations



Thank you for your attention!



a D-finite function.

Asymptotic expansions of Laplace type integrals

Watson's Lemma

Let f be a function of positive real variable s.t.
as $x \rightarrow 0^+$,

$$f(x) \sim \sum_{k=0}^{\infty} a_k x^{\frac{s+k}{\mu}-1},$$

with $s, \mu > 0$.

Then

$$\int_0^{\infty} e^{-\lambda x} f(x) dx \sim \sum_{k=0}^{\infty} \Gamma\left(\frac{s+k}{\mu}\right) \frac{a_k}{\lambda^{\frac{s+k}{\mu}}},$$

as $\lambda \rightarrow \infty$,

provided that the integral converges throughout its range for sufficiently large λ .

Note: $f(x) = \sum_{k=0}^{N-1} a_k x^{\frac{s+k}{\mu}-1} + \mathcal{O}\left(x^{\frac{s+N}{\mu}-1}\right).$