X-ray burst

What is type 1 X-ray burst?

- light curve
 - o rising time ~10s
 - o decay time ~10s a few min
 - peak luminosity
 - more information
 - derive Eddington luminosity
 - 10³⁸ erg s-1 per Solar mass
- spectrum
 - almost blackbody spectrum
 - blackbody temperature ~10^7 K(Swank et al. 1977) --> soft X-ray?
 - add recent information
 - How about the time evolution: the dependence of blackbody temperature on time?
 - whether increase or decrease? (maybe increase)
 - How rate?
- burst source
 - Low Mass X-ray Binary (LMXB)
 - neutron star and low mass star
 - X-ray bursts occur at neutron star surface.
 - Where is 'surface'?

How type 1 X-ray bursts occur?

- It is said that thermonuclear burning makes X-ray burst.
- Accreting matter from companion star lead to ignite thermonuclear burning.
 - What conditions lead to ignite?
 - Maybe, thermodynamical values are involved.
 - temperature
 - density
 - pressure
 - entropy
 - It is said that one of conditions is accretion rate
 - and, said that accretion rate determine stability of burning
 - How the typical accretion rate is?
 - Which has a significant role of 'fuel'?
 - H
 - He
- The most critical parameter determining burst properties was the accretion rate. (Fushiki and Lamb 1987)
 - How strong influence on burst profiles does the accretion rate have?
- It is said that matters falling on NS rapidly spread across the surface by latial pressure gradients
 - How such strong pressure gradients work?

- It is said that lightcurves depend on what is accreting matter(donor composition).
 - o pure He
 - H/He mix
- What is produced after He burning?

Histroy of study on X-ray bursts

- 1. Nuclear fusion in accreting NS(Rosenbluth et al. 1973)
- 2. Computing instability of thermonuclear burning on NS (Hansen and van Horn 1975)
- 3. The first discovery of X-ray burst (Grindlay et al. 1976)
- 4. Suggestion nuclear burning causes X-ray burst(Maraschi and Cavaliere 1977)
- 5. Blackbody spectrum fitting (Swank et al. 1977)
- 6. Reproducing lightcurve (helium shell burning) (Joss 1978)
- 7. Role of hydrogen shell burning (Fujimoto et al. 1981)
- 8. rp-process (Wallace and Woosley 1981)
- 9. Using more realistic nuclear reaction network (Woosley et al. 2004)

Research recent study!

inportance of type 1 X-ray burst

- restrict EoS(MR relation)
- multi wavelength information