# Note on Average Null Energy Condition

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#### Abstract

Note on average null energy condition. Use (1+1)D free boson as an example to show, how it is satisfied in general and when it will be violated.

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Einstein gravity theory tells us the metric is determined by the matters. However, it doesn't tell us anything about the matter. Energy conditions are some constraints on the matters [1]. Average null energy condition is one of them and stated in the following form,

$$\int dx^{+} T_{++} \ge 0, \quad \int dx^{-} T_{--} \ge 0. \tag{1}$$

It is also related to causality. Here we're not going to talk about this AVEG generally. Instead, the purpose of this note is very modest: we are going to use free boson as an example to show how it is satisfied and how to break it.

We choose the following action [3],

$$S = \frac{g}{2} \int d^2x (\partial_t \varphi)^2 - (\partial_x \varphi)^2, \tag{2}$$

which boundary condition  $\varphi(x) = \varphi(x + L)$ . Under canonical quantization, the field operator can be expanded as,

$$\varphi(x,t) = \varphi_0 + \frac{1}{gL}\pi_0 t + \frac{i}{\sqrt{4\pi g}} \sum_{n \neq 0} \frac{1}{n} \left( a_n e^{2\pi i n(x-t)/L} - \overline{a}_{-n} e^{2\pi i (x+t)/L} \right)$$
(3)

 $\varphi_0$  and  $\pi_0$  describe the center of mass motion.  $a_n, \overline{a}_n$  are the left and right movers moving along the lightcone direction with finite momentum, which satisfy a set of commutators,

$$[a_n, a_m] = n\delta_{n+m}, \quad [\overline{a}_n, \overline{a}_m] = n\delta_{n+m}, \quad [a_n, \overline{a}_m] = 0. \tag{4}$$

And in the following discussion, for simplicity, we will use lightcone coordinate,

$$x^{+} = x + t, \quad x^{-} = x - t,$$
 (5)

where the energy-momentum tensor has a simpler form,

$$T_{--} = g\partial_{-}\varphi\partial_{-}\varphi, \quad T_{++} = g\partial_{+}\varphi\partial_{+}\varphi, \quad T_{-+} = 0.$$
 (6)

### 1 Perturb the vacuum

First, we want to use a non-local unitary to perturb the vacuum state, then measure the energy-momentum tensor, i,e.

$$\langle T_{--}\rangle_p = \langle 0|e^{-i\lambda\varphi_L\varphi_R}T_{--}(x)e^{i\lambda\varphi_L\varphi_R}|0\rangle = -i\lambda\langle 0|[\varphi_L\varphi_R, T_{--}(x)]|0\rangle. \tag{7}$$

where  $\varphi_L = \varphi(-x_0, 0), \varphi_R = \varphi(x_0, 0)$ , which can be thought of as two sources of left and right movers. In the last step, we do a expansion w.r.t.  $\lambda$  and use  $\langle T_{--}(x) \rangle = 0$ . So up to this order, we have

$$\langle T_{--}\rangle_p = -i\lambda \left( \langle 0|[\varphi_L, \partial_-\varphi]\varphi_R \partial_-\varphi|0\rangle + \langle 0|\partial_-\varphi[\varphi_L, \partial_-\varphi]\varphi_R|0\rangle + \right.$$
 (8)

$$\langle 0|\varphi_L[\varphi_R, \partial_-\varphi]\partial_-\varphi|0\rangle + \langle 0|\varphi_L\partial_-\varphi[\varphi_R, \partial_-\varphi]|0\rangle$$
 (9)

Using the mode expansion formula, one can show that (See Appendix.A),

$$[\varphi(x_1, t_1), \partial_- \varphi(x_2, t_2)] = -\frac{i}{2q} \delta(x_1^- - x_2^-), \tag{10}$$

$$\langle \varphi(x_1, t_1) \partial_- \varphi(x_2, t_2) \rangle + \langle \partial_- \varphi(x_2, t_2), \varphi(x_1, t_1) \rangle = \frac{1}{2\pi g} P \frac{1}{x_1^- - x_2^-}. \tag{11}$$

So that we have

$$\langle T_{--}(x) \rangle_p = \frac{-\lambda}{4\pi g^2} \frac{\delta(x_L^- - x^-)}{x_R^- - x_L^-} - \frac{-\lambda}{4\pi g^2} \frac{\delta(x_R^- - x^-)}{x_L^- - x_R^-}.$$
 (12)

Because  $x_R^- x_L^- = 2x_0 > 0$ , the two terms have opposite signs thus cancel each other, which preserves the ANEC.

# 2 Change the Hamiltonian

Another scenario is to change the Hamiltonian by adding a time-dependent non-local term, which breaks Lorentz symmetry and locality,

$$H = H_0 + \delta(t)\lambda\varphi_L\varphi_R. \tag{13}$$

where  $\varphi_L$  and  $\varphi_R$  follow the same definition above. The Heisenburg operators evolving under this new Hamiltonian is written as  $\widetilde{O}(x)$ . Now we have,

$$t < 0, \quad \langle \widetilde{T}_{--}(x) \rangle = \langle T_{--}(x) \rangle = 0;$$
 (14)

$$t > 0, \quad \langle \widetilde{T}_{--}(x) \rangle = \langle 0 | e^{-i\lambda\varphi_L\varphi_R} T_{--}(x) e^{i\lambda\varphi_L\varphi_R} | 0 \rangle.$$
 (15)

So this time, ANEC is violated for  $x^-$  slides between  $-x_0$  and  $x_0$ .

From this calculation, it seems that the coupling term  $\varphi_L\varphi_R$  is not essential. If we only add  $\varphi_L$  or  $\varphi_R$ , the AVNE can also be violated.

#### A Details for the calculation

We first consider the commutator between two boson fields,

$$\begin{split} [\varphi(x_1,t_1),\varphi(x_2,t_2)] = & \frac{1}{gL} \Big( [\pi_0 t_1,\varphi_0] + [\varphi_0,\pi_0 t_2] \Big) \\ & + \frac{-1}{4\pi g} \left[ \sum_{n \neq 0} \frac{1}{n} a_n e^{2\pi i n x_1^-/L}, \sum_{m \neq 0} \frac{1}{m} a_m e^{2\pi i m x_2^-/L} \right] \\ & + \frac{-1}{4\pi g} \left[ \sum_{n \neq 0} \frac{1}{n} \overline{a}_{-n} e^{2\pi i n x_1^+/L}, \sum_{m \neq 0} \frac{1}{m} \overline{a}_{-m} e^{2\pi i m x_2^+/L} \right] \\ = & \frac{-i}{gL} (t_1 - t_2) + \frac{1}{4\pi g} \sum_{n \neq 0} \frac{1}{n} \left( e^{2\pi i n (x_1^- - x_2^-)/L} + e^{2\pi i n (x_1^+ - x_2^+)/L} \right). \end{split}$$

Using this, we have,

$$[\varphi(x_1, t_1), \partial_{-}\varphi(x_2, t_2)] = \frac{-i}{2gL} \sum_{n = -\infty}^{\infty} e^{2\pi i n(x_1^- - x_2^-)/L}$$

$$= \frac{-i}{2g} \sum_{k = -\infty}^{\infty} \delta((x_1^- - x_2^-) - kL)$$

$$\stackrel{L \to \infty}{\longrightarrow} -i \sum_{k = -\infty}^{\infty} (x_2^- - x_2^-)$$
(17)

 $\xrightarrow{L \to \infty} \frac{-i}{2g} \delta(x_1^- - x_2^-). \tag{17}$ 

From the first to the second line, we use Poisson summation formula  $\frac{1}{T} \sum_{k=-\infty}^{\infty} e^{2\pi i k x/T} = \sum_{n=-\infty}^{\infty} \delta(x-nT)$ .

Then let's compute the correlation functions,

$$\begin{split} \langle \varphi(x_1,t_1)\partial_-\varphi(x_2,t_2)\rangle &= \langle 0| \left(\varphi_0 + \frac{1}{gL}\pi_0t_1\right)\frac{-\pi_0}{2gL}|0\rangle \\ &+ \langle 0|\frac{i}{\sqrt{4\pi g}}\sum_{n\neq 0}\frac{1}{n}a_ne^{2\pi inx_1^-/L}\frac{i}{\sqrt{4\pi g}}\sum_{m\neq 0}\frac{1}{m}\frac{2\pi im}{L}a_me^{2\pi imx_2^-/L}|0\rangle \\ &= \langle 0| \left(\varphi_0 + \frac{1}{gL}\pi_0t_1\right)\frac{-\pi_0}{2gL}|0\rangle - \frac{i}{2gL}\sum_{n>0}e^{2\pi in(x_1^- - x_2^-)/L}. \\ \langle \partial_-\varphi(x_2,t_2)\varphi(x_1,t_1)\rangle &= \langle 0|\frac{-\pi_0}{2gL}\left(\varphi_0 + \frac{1}{gL}\pi_0t_1\right)|0\rangle + \frac{i}{2gL}\sum_{n>0}e^{2\pi in(x_2^- - x_1^-)/L}. \end{split}$$

In the thermodynamic limit, we can drop the contribution from zero mode. For the n > 0 modes, we use the formula

$$\frac{1}{L} \sum_{n=1}^{\infty} e^{2\pi i n x/L} = \int_0^{\infty} \frac{dk}{2\pi} e^{ikx} = \int_0^{\infty} \frac{dk}{2\pi} e^{ik(x+i\epsilon)} = \frac{1}{2\pi} \frac{i}{x+i\epsilon} = \frac{i}{2\pi} \left( P \frac{1}{x} - i\pi \delta(x) \right),$$

which leads to,

$$\langle \varphi(x_1, t_1) \partial_- \varphi(x_2, t_2) \rangle + \langle \partial_- \varphi(x_2, t_2) \varphi(x_1, t_1) \rangle = \frac{1}{2\pi g} P \frac{1}{x_1^- - x_2^-}.$$
 (18)

All of these calculations can be easily extended to  $T_{++}$ .

## References

- [1] A. Zee, "Einstein gravity in a nutshell", Princeton.
- [2] D. Stanford, "talk at the conference celebrating Hawking's birthday".
- [3] F. D. Francesco, et.al., "Conformal Field theory", Springer.