

**THE UNIVERSITY OF QUEENSLAND
SCHOOL OF MATHEMATICS AND PHYSICS
PHYS2041 – Quantum Physics**

Tutorial 9 Solutions

Problem 9.1 (a) First, consider the case where the position and momentum operators are along the same axis, e.g. x and \hat{p}_x . Introducing a “test function” $f(x)$ will help us make sure we treat the derivatives properly,

$$[x, \hat{p}_x] f = -i\hbar \left(x \frac{\partial}{\partial x} - \frac{\partial}{\partial x} x \right) f \quad (1)$$

$$= -i\hbar \left(x \frac{\partial f}{\partial x} - \frac{\partial}{\partial x} (xf) \right) \quad (2)$$

$$= -i\hbar \left(\cancel{x \frac{\partial f}{\partial x}} - f - \cancel{x \frac{\partial f}{\partial x}} \right) \text{ (product rule)} \quad (3)$$

$$= i\hbar f \quad (4)$$

Dropping the test function we can conclude $[x, \hat{p}_x] = i\hbar$, and similarly for y and z . What about the situation where the position and momentum operators are along perpendicular axes, for instance x and \hat{p}_y ? Again we'll introduce a test function $f(x, y)$,

$$[x, \hat{p}_y] f = -i\hbar \left(x \frac{\partial}{\partial y} - \frac{\partial}{\partial y} x \right) f \quad (5)$$

$$= -i\hbar \left(x \frac{\partial f}{\partial y} - x \frac{\partial f}{\partial y} \right) \quad (6)$$

$$= 0 \quad (7)$$

A similar proof for all other combinations leads us to the conclusion that position and momentum commute when they are along perpendicular axes.

What about the opposite commutator, $[\hat{p}_x, x]$? Well it's easy to show for any two operators,

$$[\hat{A}, \hat{B}] = \hat{A}\hat{B} - \hat{B}\hat{A} \quad (8)$$

$$= -(\hat{B}\hat{A} - \hat{A}\hat{B}) \quad (9)$$

$$= -[\hat{B}, \hat{A}] \quad (10)$$

which implies that $[\hat{p}_x, x] = -i\hbar$, etc. Putting this all together, we have that $[r_i, \hat{p}_j] = -[\hat{p}_i, r_j] = i\hbar\delta_{jk}$.

It's obvious that $[x, y] = xy - yx = 0$ - the position of a particle commutes with all other directions. We also know that partial derivatives of independent variables commute (Clairaut's theorem), so the same is true of momentum. In other words, $[r_i, r_j] = [\hat{p}_i, \hat{p}_j] = 0$.

(b) The generalised uncertainty principle is

$$\sigma_A^2 \sigma_B^2 \geq \left(\frac{1}{2i} \langle [\hat{A}, \hat{B}] \rangle \right)^2, \quad (11)$$

so all we need to do is plug in the commutation relations from the previous question. Notice that commuting observables do not have an uncertainty trade-off, which is why they're sometimes called *compatible observables*.

We have already seen that the only three combinations of position and momentum that are *incompatible* (i.e. non-commuting) are

$$\sigma_x^2 \sigma_{p_x}^2 \geq \frac{\hbar^2}{4}, \quad \sigma_y^2 \sigma_{p_y}^2 \geq \frac{\hbar^2}{4}, \quad \sigma_z^2 \sigma_{p_z}^2 \geq \frac{\hbar^2}{4}, \quad (12)$$

with all other combinations zero (e.g. $\sigma_x \sigma_y \geq 0$, and so on).

Problem 9.2 [FOR ASSIGNMENT 5; max 10 points]

(a) We use separation variables and write the wavefunction in the form of

$$\psi(x, y, z) = f(x) g(y) h(z) \quad (13)$$

and insert it in 3D time independent Schrödinger equation

$$-\frac{\hbar^2}{2m} \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \right) \psi = E\psi, \quad (14)$$

$$-\frac{\hbar^2}{2m} \left(\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \right) f(x) g(y) h(z) = E f(x) g(y) h(z), \quad (15)$$

$$(16)$$

which after rearrangement can be written as

$$\frac{1}{f(x)} \frac{d^2 f(x)}{dx^2} + \frac{1}{g(y)} \frac{d^2 g(y)}{dy^2} + \frac{1}{h(z)} \frac{d^2 h(z)}{dz^2} = -\frac{2m}{\hbar^2} E. \quad (17)$$

Each term on the LHS depends only on one variable, and so it *must* equal to a constant, because the equation must be true for *all* values of x , y and z .; we will denote these three constants as $-\frac{2m}{\hbar^2} E_x$, $-\frac{2m}{\hbar^2} E_y$, and $-\frac{2m}{\hbar^2} E_z$, respectively. Therefore we can write the 3D TDSE (which is a partial differential equation) as three ordinary differential equations,

$$\frac{1}{f(x)} \frac{d^2 f(x)}{dx^2} = -\frac{2m}{\hbar^2} E_x, \quad (18a)$$

$$\frac{1}{g(y)} \frac{d^2 g(y)}{dy^2} = -\frac{2m}{\hbar^2} E_y, \quad (18b)$$

$$\frac{1}{h(z)} \frac{d^2 h(z)}{dz^2} = -\frac{2m}{\hbar^2} E_z, \quad (18c)$$

where the constant energies in each dimension E_x , E_y and E_z must add up to the total energy E ,

$$E_x + E_y + E_z = E. \quad (19)$$

Equations (18 a-c) are simply the TISE for one dimensional infinite square well. Hence the 3D box potential is equivalent to three 1D square wells. We can use the solution for 1D

$$f_{n_x}(x) = \sqrt{\frac{2}{a}} \sin\left(\frac{n_x \pi x}{a}\right), \quad E_{n_x} = \frac{\hbar^2 \pi^2 n_x^2}{2ma^2}, \quad n_x = 1, 2, 3, \dots, \quad (20a)$$

$$g_{n_y}(y) = \sqrt{\frac{2}{b}} \sin\left(\frac{n_y \pi y}{b}\right), \quad E_{n_y} = \frac{\hbar^2 \pi^2 n_y^2}{2mb^2}, \quad n_y = 1, 2, 3, \dots, \quad (20b)$$

$$h_{n_z}(z) = \sqrt{\frac{2}{c}} \sin\left(\frac{n_z \pi z}{c}\right), \quad E_{n_z} = \frac{\hbar^2 \pi^2 n_z^2}{2mc^2}, \quad n_z = 1, 2, 3, \dots, \quad (20c)$$

Thus the solution to the Schrodinger equation for the particle in a 3D box is by a energy

$$E_{n_x, n_y, n_z} = \frac{\hbar^2 \pi^2}{2m} \left(\frac{n_x^2}{a^2} + \frac{n_y^2}{b^2} + \frac{n_z^2}{c^2} \right), \quad (21)$$

and a wavefunction

$$\psi_{n_x, n_y, n_z}(x, y, z) = \sqrt{\frac{8}{abc}} \sin\left(\frac{n_x \pi x}{a}\right) \sin\left(\frac{n_y \pi y}{b}\right) \sin\left(\frac{n_z \pi z}{c}\right). \quad (22)$$

We can check the normalization by taking the integral

$$\begin{aligned} & \sqrt{\frac{8}{abc}} \int_0^a \sin\left(\frac{n_x \pi x}{a}\right) dx \int_0^b \sin\left(\frac{n_y \pi y}{b}\right) dy \int_0^c \sin\left(\frac{n_z \pi z}{c}\right) dz \\ &= \sqrt{\frac{8}{abc}} \sqrt{\frac{a}{2}} \sqrt{\frac{b}{2}} \sqrt{\frac{c}{2}} \\ &= 1. \end{aligned} \quad (23)$$

(b) If $a = b = c = L$ solutions read as

$$\psi_{n_x, n_y, n_z}(x, y, z) = \left(\frac{2}{L}\right)^{\frac{3}{2}} \sin\left(\frac{n_x \pi x}{L}\right) \sin\left(\frac{n_y \pi y}{L}\right) \sin\left(\frac{n_z \pi z}{L}\right), \quad (24)$$

with allowed energies

$$E_{n_x, n_y, n_z} = \frac{\hbar^2 \pi^2}{2mL^2} (n_x^2 + n_y^2 + n_z^2), \quad (25)$$

Now we calculate $E_{1,2,3}$

$$E_{1,2,3} = \frac{\hbar^2 \pi^2}{2mL^2} (1 + 4 + 9) = \frac{14\hbar^2 \pi^2}{2mL^2}, \quad (26)$$

which is equal to energy of the following

$$E_{1,2,3} = E_{1,3,2} = E_{2,1,3} = E_{2,3,1} = E_{3,1,2} = E_{3,2,1}, \quad (27)$$

so there are 6 degenerate states with energy $\frac{14\hbar^2\pi^2}{2mL^2}$.

Problem 9.3 [FOR ASSIGNMENT 5; max 10 points]

(a) The ground-state wavefunction of the Hydrogen Hamiltonian is, in spherical coordinates:

$$\psi_{1,0,0}(r, \theta, \phi) = \frac{1}{\sqrt{a^3\pi}} e^{-\frac{r}{a}} \quad (28)$$

where $a > 0$ is the Bohr radius. Notice that this does not actually depend on θ or ϕ - only r . Such functions are called spherically symmetric.

Notice also that in cartesian coordinates, where $r = \sqrt{x^2 + y^2 + z^2}$, this wavefunction would look like

$$\psi_{1,0,0}(x, y, z) = \frac{1}{\sqrt{a^3\pi}} e^{-\frac{\sqrt{x^2+y^2+z^2}}{a}}. \quad (29)$$

This is, however, is a difficult function to deal with and to integrate, hence the advantage of using the spherical coordinates for integrations (in fact, I know of no other way of integrating this function except by going to spherical coordinates).

For any spherically symmetric function $f(r)$, the normalisation integral in spherical coordinates simplifies to a single integral over r (take care to include the Jacobian!), because the integrals over the angles are trivial and give a constant factor of 4π :

$$\int_0^\pi \int_0^{2\pi} \int_0^\infty f(r) r^2 \sin \theta dr d\phi d\theta \quad (30)$$

$$= \int_0^\infty f(r) r^2 dr \int_0^\pi \sin \theta d\theta \int_0^{2\pi} d\phi \quad (31)$$

$$= [\phi]_0^{2\pi} [-\cos \theta]_0^\pi \int_0^\infty f(r) r^2 dr \quad (32)$$

$$= 4\pi \int_0^\infty f(r) r^2 dr \quad (33)$$

The expectation value of r or the average radius is

$$\langle r \rangle = \langle \psi_{1,0,0} | r | \psi_{1,0,0} \rangle = 4\pi \int_0^\infty r |\psi_{1,0,0}(r)|^2 r^2 dr \quad (34)$$

$$= 4\pi \int_0^\infty r^3 |\psi_{1,0,0}(r)|^2 dr \quad (35)$$

$$= \frac{4}{a^3} \int_0^\infty r^3 e^{-\frac{2r}{a}} dr \quad (36)$$

$$= \frac{4}{a^3} \left(\left[\frac{a}{2} r^3 e^{-\frac{2r}{a}} \right]_0^\infty + \frac{3a}{2} \int_0^\infty r^2 e^{-\frac{2r}{a}} dr \right) \text{ (integration by parts)} \quad (37)$$

$$= \frac{3a}{2} \langle \psi_{1,0,0} | \psi_{1,0,0} \rangle \quad (38)$$

$$= \frac{3a}{2} \quad (39)$$

where in the final line we made use of the normalisation of $\psi_{1,0,0}(r)$ (remember $\psi_{1,0,0}$ is spherically symmetric, its normalisation integral is in r only, Eq. (33)).

The mean-square radius is

$$\langle r^2 \rangle = \langle \psi_{1,0,0} | r^2 | \psi_{1,0,0} \rangle = 4\pi \int_0^\infty r^2 |\psi_{1,0,0}(r)|^2 r^2 dr \quad (40)$$

$$= 4\pi \int_0^\infty r^4 |\psi_{1,0,0}(r)|^2 dr \quad (41)$$

$$= \frac{4}{a^3} \int_0^\infty r^4 e^{-\frac{2r}{a}} dr \quad (42)$$

$$= \frac{4}{a^3} \left(\left[\frac{a}{2} r^4 e^{-\frac{2r}{a}} \right]_0^\infty + \frac{4a}{2} \int_0^\infty r^3 e^{-\frac{2r}{a}} dr \right) \text{ (integration by parts)} \quad (43)$$

$$= \frac{4a}{2} \langle r \rangle \text{ [Eq. (19)]} \quad (44)$$

$$= 3a^2 \quad (45)$$

(b) The ground state of hydrogen is spherically symmetric (it does not depend on θ or ϕ), which means that the particle is equally likely to be found either side of the x axis and thus $\langle x \rangle = 0$. This result could be obtained quantitatively by evaluating (remember $x = r \cos \phi \sin \theta$)

$$\langle x \rangle = \int_0^\pi \int_0^{2\pi} \int_0^\infty |\psi_{1,0,0}(r)|^2 r^3 \cos \phi \sin^2 \theta dr d\phi d\theta, \quad (46)$$

which we can see is zero because $\int_0^{2\pi} \cos \phi d\phi = 0$.

Because $r^2 = x^2 + y^2 + z^2$, it is also true that $\langle r^2 \rangle = \langle x^2 \rangle + \langle y^2 \rangle + \langle z^2 \rangle$. Due to the spherical symmetry of $\psi_{1,0,0}$ we also know that $\langle x^2 \rangle = \langle y^2 \rangle = \langle z^2 \rangle$. Without having to even do any integrals, we can deduce $\langle r^2 \rangle = 3\langle x^2 \rangle$, or

$$\langle x^2 \rangle = \frac{1}{3} \langle r^2 \rangle = a^2. \quad (47)$$

The uncertainty $\sigma_x = \sqrt{\langle x^2 \rangle - \langle x \rangle^2}$ is therefore $\sigma_x = a$.

(c) As we will see here, the most likely value is *not* generally the same as the mean value.

We need to find the value of r that maximises the radial probability density. Remember that the probability of finding the particle inside a volume element $dV = dx dy dz = r^2 \sin \theta dr d\theta d\phi$ is $dP = |\Psi(x, y, z, t)|^2 dV$, so even though we're not integrating $\psi_{1,0,0}$ we still need to include the Jacobian. Let's define the radial probability density (you can verify for yourself that $\int_0^\infty P(r) dr = 1$),

$$P(r) = 4\pi r^2 |\psi_{1,0,0}(r)|^2 = \frac{4}{a^3} r^2 e^{-\frac{2r}{a}}, \quad (48)$$

which unlike $|\psi_{1,0,0}(r)|^2$ goes to zero as $r \rightarrow 0$. If you think about it, the probability of finding the particle at $r = 0$ should always be zero, since this is the probability of finding the particle inside a sphere of zero volume! The value of r that maximises $P(r)$ can be found by examining the stationary points $dP/dr = 0$,

$$\frac{dP}{dr} = \frac{4}{a^3} \left(2r - \frac{2}{a} r^2 \right) e^{-\frac{2r}{a}} = 0 \quad (49)$$

Exponentials are never zero, so this equation implies $2r - \frac{2}{a} r^2 = r(2 - \frac{2}{a} r) = 0$, which is solved for either $r = 0$ or $r = a$.

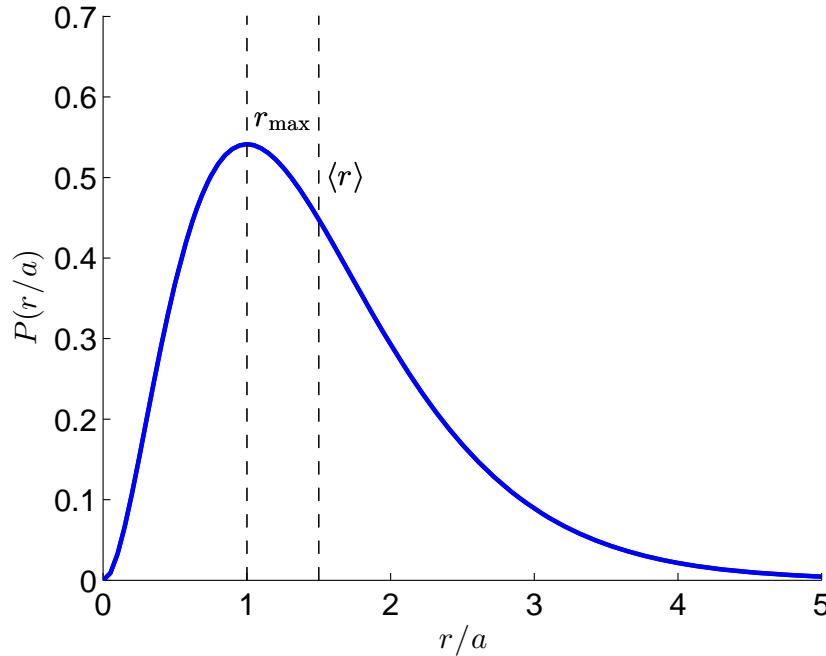


Figure 1: Radial probability distribution for the ground-state of Hydrogen. The most likely position r_{\max} and mean position $\langle r \rangle$ are indicated by black dashed lines.

Clearly the $r = 0$ solution is a minimum, as $P(0) = 0$ [see also Figure 1], so the electron is most likely to be found at the Bohr radius,

$$r_{\max} = a. \quad (50)$$

This is an important result, as it helps reconcile the naive Bohr model with the Schrödinger's wave mechanics. Figure 1 shows $P(r)$, with both the most likely value and mean values of r indicated.

Problem 9.4

$$(a) \quad \Psi(\mathbf{r}, t) = \frac{1}{\sqrt{2}} \left(\psi_{211} e^{-iE_2 t/\hbar} + \psi_{21-1} e^{-iE_2 t/\hbar} \right) = \frac{1}{\sqrt{2}} (\psi_{211} + \psi_{21-1}) e^{-iE_2 t/\hbar}; \quad E_2 = \frac{E_1}{4} = -\frac{\hbar^2}{8ma^2}.$$

$$\psi_{211} + \psi_{21-1} = -\frac{1}{\sqrt{\pi a}} \frac{1}{8a^2} r e^{-r/2a} \sin \theta (e^{i\phi} - e^{-i\phi}) = -\frac{i}{\sqrt{\pi a} 4a^2} r e^{-r/2a} \sin \theta \sin \phi.$$

$$\boxed{\Psi(\mathbf{r}, t) = -\frac{i}{\sqrt{2\pi a} 4a^2} r e^{-r/2a} \sin \theta \sin \phi e^{-iE_2 t/\hbar}.$$

(b)

$$\begin{aligned} \langle V \rangle &= \int |\Psi|^2 \left(-\frac{e^2}{4\pi\epsilon_0} \frac{1}{r} \right) d^3\mathbf{r} = \frac{1}{(2\pi a)(16a^4)} \left(-\frac{e^2}{4\pi\epsilon_0} \right) \int \left(r^2 e^{-r/a} \sin^2 \theta \sin^2 \phi \right) \frac{1}{r} r^2 \sin \theta dr d\theta d\phi \\ &= \frac{1}{32\pi a^5} \left(-\frac{\hbar^2}{ma} \right) \int_0^\infty r^3 e^{-r/a} dr \int_0^\pi \sin^3 \theta d\theta \int_0^{2\pi} \sin^2 \phi d\phi = -\frac{\hbar^2}{32\pi m a^6} (3! a^4) \left(\frac{4}{3} \right) (\pi) \\ &= \boxed{-\frac{\hbar^2}{4ma^2} = \frac{1}{2} E_1 = \frac{1}{2} (-13.6 \text{eV}) = -6.8 \text{eV}} \quad (\text{independent of } t). \end{aligned}$$