

# Lec16-matplotlib1-RyanSponzilli

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## 1 ASTR 310 Lecture 16 - matplotlib 1

### 1.0.1 Exercise 1: a scatter plot

- Download the file `halos.dat` from today's exercise page. This is a text file containing  $x$ ,  $y$ ,  $z$  positions and masses for dark matter halos in a cosmological simulation. Positions are in Mpc and masses are in  $M_\odot$ .
- Create a scatter plot of the  $(x, y)$  coordinates of the halos. Represent each halo by a circle whose size is proportional to the logarithm of the halo's mass. For example, you might try using a size given by  $70[\log(M/M_\odot) - 11.3]$ .
- Using your favorite color map, make the color of each circle proportional to the  $z$ -coordinate of the corresponding halo. If you don't like the default colors, you can change the color map using `plt.set_cmap('Spectral')` or something similar. See the gallery for examples.
- Include an appropriate title and axis labels.

[10 pts]

```
[136]: import numpy as np
import astropy.io.ascii
import astropy.table
import matplotlib.pyplot as plt
```

```
[137]: table = astropy.io.ascii.read("halos.dat")
table.rename_column('col1', 'x')
table.rename_column('col2', 'y')
table.rename_column('col3', 'z')
table.rename_column('col4', 'mass')
table['size'] = 70*(np.log10(table['mass']) - 11.3)
table
```

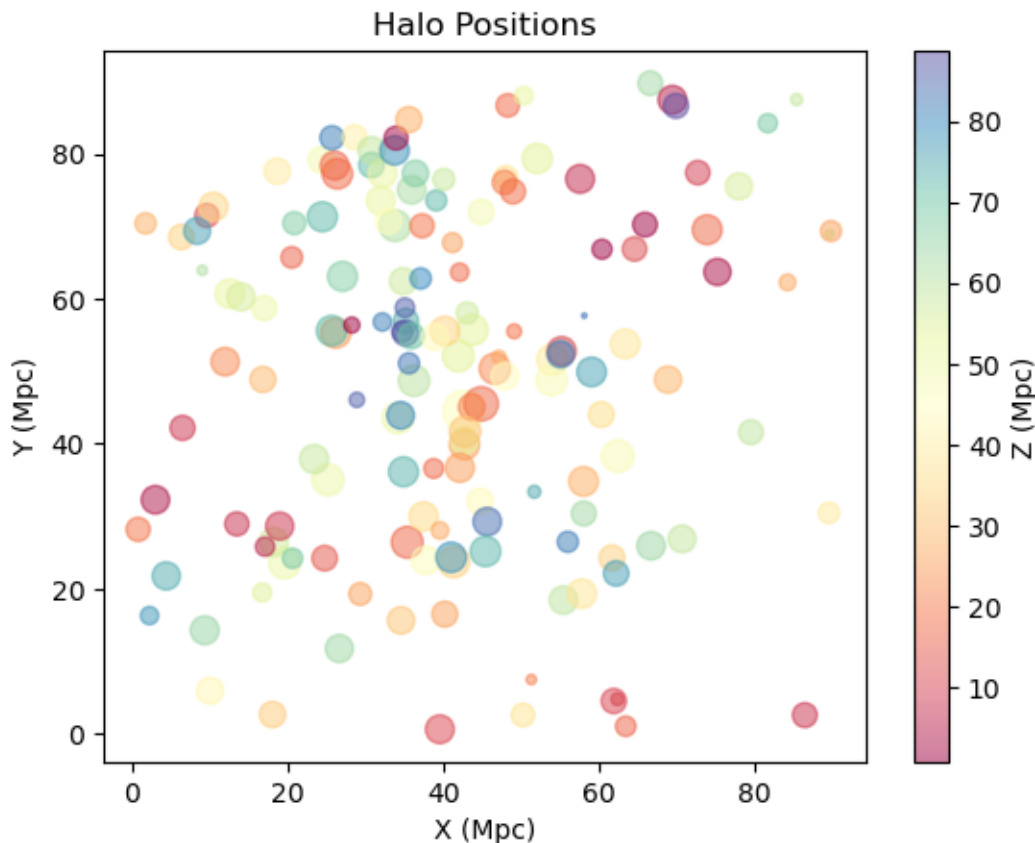
[137]: <Table length=154>

x	y	z	mass	size
float64	float64	float64	float64	float64
42.584	44.506	45.621	505300000000000.0	238.24845087667615
46.593	50.376	22.243	164400000000000.0	134.1131269242821
0.767	28.209	19.144	27350000000000.0	79.58701314686137

20.833	70.433	67.554	2111000000000.0	71.7141763315359
4.379	21.797	74.507	7180000000000.0	108.92871109696102
19.595	23.582	50.171	19390000000000.0	139.13044663550932
1.719	70.401	27.654	15150000000000.0	61.628884298682664
63.351	53.788	38.193	107200000000000.0	121.11363497497258
2.228	16.323	77.722	8756000000000.0	44.96140270271255
...	...	...	...	...
79.513	41.611	59.114	31210000000000.0	83.60056380657326
77.966	75.502	55.398	52320000000000.0	99.30674147213475
45.404	25.18	72.749	135700000000000.0	128.2805893361816
81.678	84.217	68.545	10280000000000.0	49.839518026147914
84.206	62.265	26.243	6771000000000.0	37.14569696581968
85.365	87.463	58.246	3716000000000.0	18.905299372512285
86.43	2.603	7.778	3151000000000.0	83.89138822217987
89.603	68.929	57.235	2780000000000.0	10.083135714265286
89.79	69.34	25.309	14220000000000.0	59.70297174756233
89.539	30.478	38.98	15000000000000.0	61.326388133897645

```
[138]: plt.scatter(table['x'], table['y'], c=table['z'], s=table['size'],
    cmap='Spectral', alpha=0.5)
plt.colorbar().set_label("Z (Mpc)")
plt.title("Halo Positions")
plt.xlabel("X (Mpc)")
plt.ylabel("Y (Mpc)")
```

```
[138]: Text(0, 0.5, 'Y (Mpc)')
```



### 1.0.2 Exercise 2: pulsars

Download the file “atnf\_pulsars\_20200112.dat” from the exercise page. This is a fixed-width ASCII table containing data from the Australia Telescope National Facility (ATNF) catalog of pulsars in our Galaxy known as of 1/12/20.

- Use `astropy.io.ascii` to read the table. You might need to give the reader some hints about the format or delimiters and that the header starts on line 2.  
We will make use of the pulsar periods  $P$  (column `period`), rate of change of periods  $\dot{P}$  (column `period_dot`), and companion types (column `companion_type`).
- The companion type will be masked out if the pulsar is an isolated pulsar, or else it will be one of ‘MS’, ‘NS’, ‘CO’, ‘He’, or ‘UL’ depending on the type of binary companion. Make two boolean mask arrays to select the pulsars that are in binaries vs. not in binaries. *Hint*: use the `mask` attribute of the `companion_type` column object.

[4 pts]

```
[139]: t2 = astropy.io.ascii.read("atnf_pulsars_20200112.dat", guess=False,
    ↪ delimiter='|', header_start=2, data_start=3)
```

```
nonbinaries = t2['companion_type'].mask
binaries = ~nonbinaries
```

- Make a log-log scatter plot of  $\dot{P}$  vs.  $P$ . Make binary pulsars and isolated pulsars two separate colors, and produce an appropriate legend. Make appropriate axis labels ( $P$  is in units of seconds, while  $\dot{P}$  is dimensionless).
- Finally, overplot the “pulsar death line.” This is (roughly) the function

$$\dot{P}_{\text{death}} = 7.14 \times 10^{-18} \left( \frac{P}{\text{sec}} \right)^{3.2}$$

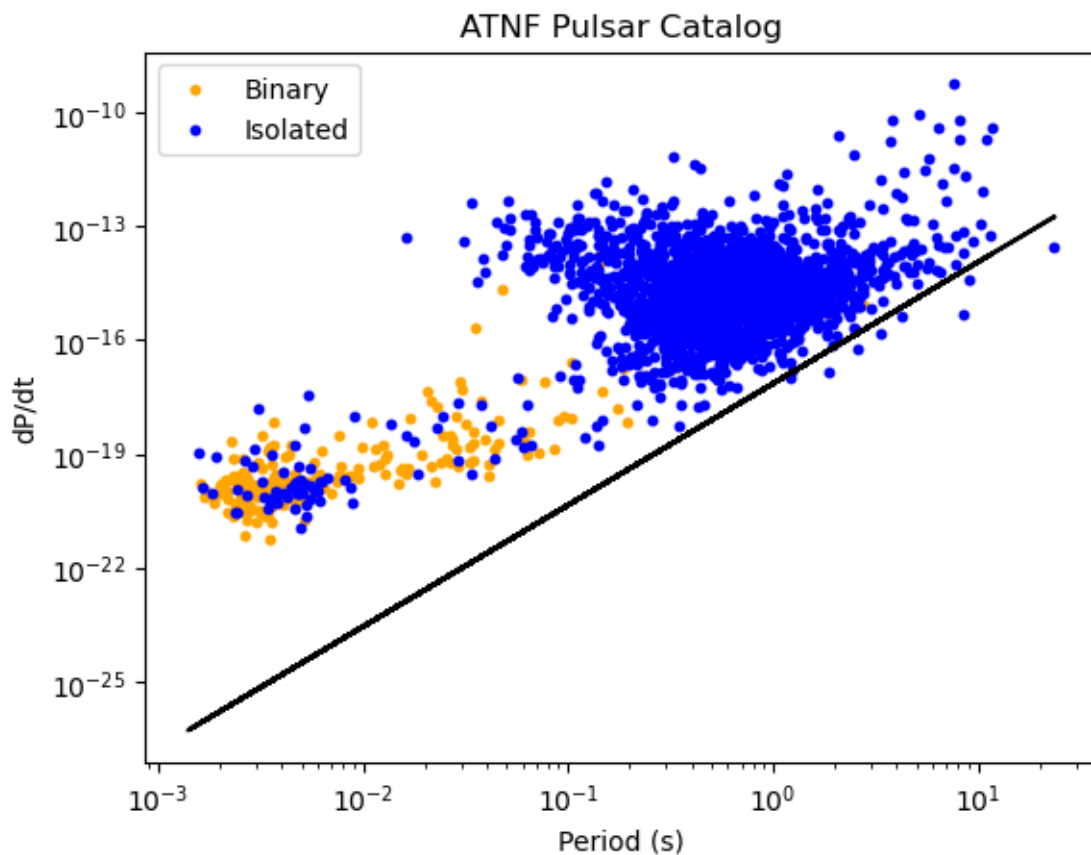
and corresponds to the line beyond which pulsars stop producing pulses. You can draw a line by defining two points on the line and connecting them.

[6 pts]

```
[143]: plt.scatter(t2['period'][binaries], t2['period_dot'][binaries], c='orange',
    ↪s=10)
plt.scatter(t2['period'][nonbinaries], t2['period_dot'][nonbinaries], c='blue',
    ↪s = 10)
plt.xlabel("Period (s)")
plt.ylabel("dP/dt")
plt.title("ATNF Pulsar Catalog")
plt.legend(["Binary", "Isolated"])
plt.xscale("log")
plt.yscale('log')

t2['deathline'] = 7.14e-18 * (t2['period']**3.2)
plt.plot(t2['period'], t2['deathline'], c='k')
```

[143]: [<matplotlib.lines.Line2D at 0x17a8ed220>]



```
[141]: t2
```

```
[141]: <Table length=2800>
```

col0	name	ra	...	_1	deathline
int64	str16	str13	...	int64	float64
---	-----	-----	...	-----	-----
--	PSR J1036-8317	10 36 00.0000	...	--	9.071871444050527e-26
--	PSR J0211-8159	02 11 59.0000	...	--	9.061867506822521e-18
--	PSR J1900-7951	19 00 24.4000	...	--	1.5699790581930042e-17
--	PSR J1119-7936	11 19 40.8000	...	--	9.987428592146507e-17
--	PSR J1057-7914	10 57 27.4000	...	--	1.8539131226343153e-17
--	PSR J1159-7910	11 59 35.7100	...	--	9.086664421394126e-19
--	PSR J1841-7845	18 41 25.9000	...	--	2.5641733786118627e-19
--	PSR J1403-7646	14 03 04.1000	...	--	1.6785213049829142e-17
--	PSR J1651-7642	16 51 07.8700	...	--	4.3214681163905236e-17
...	...	...	...	...	...
--	PSR J0645+80	06 46 00.0000	...	--	1.8696177067216754e-18
--	PSR J0100+80	01 00 18.0000	...	--	2.5777926887357078e-17
--	PSR J0849+8028	08 49 01.5000	...	--	3.227126604385572e-17

-- PSR J1641+8049	16	41	20.8381 ...	--	1.704644272348171e-26
-- PSR J0653+8051	06	53	15.0900 ...	--	1.3295424224935145e-17
-- PSR J1942+8106	19	42	54.6800 ...	--	4.3802960516618106e-20
-- PSR J0614+83	06	14	00.0000 ...	--	8.074943002832524e-18
-- PSR J1321+8323	13	21	45.6315 ...	--	1.98250911635482e-18
-- PSR J2351+8533	23	51	03.2961 ...	--	7.411415524568623e-18
-- PSR J1624+8643	16	24	32.7200 ...	--	3.676976848406149e-19

[ ]: