

Survival probability

October 26, 2025

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Todo list

SN:	As a sanity check, check numerically in the BH model that Eqs. (14) and (15) give the same SFF curve, computing separately term by term of Eq. (15), and then adding them up	3
SN:	Check numerically if the left side of Eq. (16) is well described by either Eq. (17) when a single symmetric subspace is considered, and by Eq. (18) when the whole spectrum is considered (we have the intuition that the whole spectrum should exhibit Poisson statistics at first order spacings or ratios)	3

1 Question

Why is the correlation hole not affected by the precense of a high degree of symmetries in Ref. [1]?

2 Mostly ideas

2.1 Survival probability in a system with the simplest symmetric structure

The survival probability is defined as

$$S_p(t) = |\langle \psi_0 | \psi(t) \rangle|^2, \quad (1)$$

Consider that $H |\phi_n\rangle = E_n |\phi_n\rangle$, then the survival probability S_p yields

$$S_p(t) = \left| \sum_k |\langle \psi_0 | \phi_k \rangle|^2 e^{-iE_k t} \right|^2. \quad (2)$$

Let us consider a symmetry of H under an arbitrary operator Π such that $H = H_1 \oplus H_2$ —for instance, you may consider Π to be the reflection operator. Then, we can relabel the energies E_n such that $\{E_n\}_{n=1}^p$ and $\{E_n\}_{n=p+1}^N$ are the spectra of H_1 and H_2 , respectively. N denotes the dimension of H , p the dimension of H_1 and $N - p$ is the dimension of H_2 . After this relabeling, $S_p(t)$ takes the form:

$$S_p(t) = \left| \sum_{k=1}^p |\langle \psi_0 | \phi_k \rangle|^2 e^{-iE_k t} + \sum_{l=p+1}^N |\langle \psi_0 | \phi_l \rangle|^2 e^{-iE_l t} \right|^2 \quad (3)$$

$$= \left| \sum_{k=1}^p |\langle \psi_0 | \phi_k \rangle|^2 e^{-iE_k t} \right|^2 + \left| \sum_{l=p+1}^N |\langle \psi_0 | \phi_l \rangle|^2 e^{-iE_l t} \right|^2 \quad (4)$$

$$+ 2 \operatorname{Re} \left\{ \sum_{k=1}^{p+1} \sum_{l=p+1}^N |\langle \psi_0 | \phi_k \rangle|^2 |\langle \psi_0 | \phi_l \rangle|^2 e^{-i(E_k - E_l)t} \right\}. \quad (5)$$

The first two terms are recognized as the survival probability in the two symmetric subspaces. The third term God knows what the hell it is, but it is the term explaining why the correlation whole survives (or not).

2.2 Survival probability as function of level spacings

Let us rewrite the survival probability as follows:

$$S_p(t) = \left| \sum_k |\langle \psi_0 | \phi_k \rangle|^2 e^{-iE_k t} \right|^2 \quad (6)$$

$$= \sum_{k,l} |\langle \psi_0 | \phi_k \rangle|^2 |\langle \psi_0 | \phi_l \rangle|^2 e^{-i(E_k - E_l)t} \quad (7)$$

$$= \sum_{k=1}^N |\langle \psi_0 | \phi_k \rangle|^2 |\langle \psi_0 | \phi_k \rangle|^2 + 2 \operatorname{Re} \left[\sum_{k=1}^{N-1} |\langle \psi_0 | \phi_k \rangle|^2 |\langle \psi_0 | \phi_{k+1} \rangle|^2 e^{-i(E_k - E_{k+1})t} \right. \\ \left. + \sum_{k=1}^{N-2} |\langle \psi_0 | \phi_k \rangle|^2 |\langle \psi_0 | \phi_{k+2} \rangle|^2 e^{-i(E_k - E_{k+2})t} + \dots \right. \\ \left. + |\langle \psi_0 | \phi_N \rangle|^4 e^{-i(E_1 - E_N)t} \right]. \quad (8)$$

In this form, it is explicit that the survival probability can be written as a sum of the contributions given by higher order spacings. From Ref. [2] we learned not only that RMT's predictions can be verified even a non-desymmetrized system, but also that higher order spacing ratios exhibit correlations that match RMT's predictions. Very recently, a work on the distribution of higher order spacings for the circular gaussian ensembles appeared on the arXiv [3] **JA: I think these results we can use them to study the**

behavior of each term in Eq. (8). My guess is that the most relevant contributions to $S_p(t)$ come from the higher order spacing correlations.

Let us consider two identical spectra $\{E_1^{(1)}, E_2^{(1)}, E_3^{(1)}\}$ and $\{E_1^{(2)}, E_2^{(2)}, E_3^{(2)}\}$ (that is, $E_i^{(1)} = E_i^{(2)}$). The whole ordered spectrum will be $\{E_1^{(1)}, E_1^{(2)}, E_2^{(1)}, E_2^{(2)}, E_3^{(1)}, E_3^{(2)}\}$. Then, compute the spacings:

$$s_1^{(1)} = E_1^{(2)} - E_1^{(1)} = 0 \quad (9)$$

$$s_1^{(2)} = E_2^{(1)} - E_1^{(1)} \quad (10)$$

$$s_1^{(3)} = E_2^{(2)} - E_1^{(1)} \quad (11)$$

$$s_1^{(4)} = E_3^{(1)} - E_1^{(1)} \quad (12)$$

$$s_1^{(5)} = E_3^{(2)} - E_1^{(1)} \quad (13)$$

The spacings of first order $s_n^{(1)}$ follow a Poissonian distribution, possibly a zero-inflated Poisson distribution as half the spacings will become zero. The second-order spacings become the first-order spacings of both subspaces JA: I think the PDF here will be the convolution of two Wigner surmises. The third-order spacings become a mix between first- and second-order spacings JA: maybe the convolution of three Wigner surmises?. The fourth-order spacings will become second-order spacings from both subspaces JA: once again, the convolution of four Wigner surmises?. The fifth-order spacings become a mix between second- and third-order spacings JA: you know the drill by now, the convolution of five Wigner surmises?

JA: I guess the convolution of many Wigner surmises should converge to something...? Hence, after some order the PDF should be almost the same. Something to ask Deepseek or chatGPT

JA: If all of this is, at the very least, a good approximation, we still have to take into account the coefficients $|\langle\psi_0|\phi_i\rangle|^2|\langle\psi_0|\phi_{i+k}\rangle|^2$. Therefore, we may consider the case where all coefficients are equal, case in which I'm almost sure the $S_p(t)$ becomes the spectral form factor. If it is indeed the case we should expect a dip-ramp-plateau behavior, the ramp being the hallmark of long-range correlations of a spectrum's chaotic system.

2.3 Spectral form factor (SFF)

The SFF can be decomposed into the following sums:

$$K(t) = \frac{1}{N^2} \sum_{k,l} \overline{e^{-i(E_k - E_l)t}} \quad (14)$$

$$= 1 + \frac{2}{N^2} \operatorname{Re} \left[\sum_{k=1}^{N-1} \overline{e^{-i(E_{k+1} - E_k)t}} + \sum_{k=1}^{N-2} \overline{e^{-i(E_{k+2} - E_k)t}} + \dots + \overline{e^{-i(E_N - E_1)t}} \right]. \quad (15)$$

SN: As a sanity check, check numerically in the BH model that Eqs. (14) and (15) give the same SFF curve, computing separately term by term of Eq. (15), and then adding them up

We try to move forward with the following approximation. We assume the spacings $s_k^{(1)} = E_{k+1} - E_k$ are independent and identically distributed (i.i.d), thus:

$$\overline{\sum_{k=1}^{N-1} e^{-i(E_{k+1} - E_k)t}} \approx (N-1) \overline{e^{-is^{(1)}t}} = \int_0^\infty e^{-ist} P(s) ds, \quad (16)$$

where $P(s)$ is either a Poisson or a Wigner-Dyson distribution, depending on the case,

$$\int_0^\infty e^{-ist} P_{\text{WD}}(s) ds = (N-1) \left[1 - te^{-t^2/\pi} \left(\operatorname{erfi} \left(\frac{t}{\sqrt{\pi}} \right) + i \right) \right] \quad (17)$$

$$\int_0^\infty e^{-ist} P_P(s) ds = \frac{N-1}{1+it}. \quad (18)$$

SN: Check numerically if the left side of Eq. (16) is well described by either Eq. (17) when a single symmetric subspace is considered, and by Eq. (18) when the whole spectrum is considered (we have the intuition that the whole spectrum should exhibit Poisson statistics at first order spacings or ratios)

JA: In Ref. [4] they study the distribution of higher order level spacings.

2.4 Interesting things about SFF

There is a direct need in averaging the SFF computed, here a comparison between two methods will be shown, based on the data obtained from the BHH model. A time window averaging is performed on the data and a sector averaging aswell (averaging the SFF at each time with each symmetric sector of the hamiltonian).

The conclusion is that in a general context, a time averaging is best by far in comparison to a sector averaging, mainly due to the fact that a hamiltonian might only have 2 symmetric sectors. It is remarkable that applying both a sector averaging and a time window averaging proves to be the most precise approach.

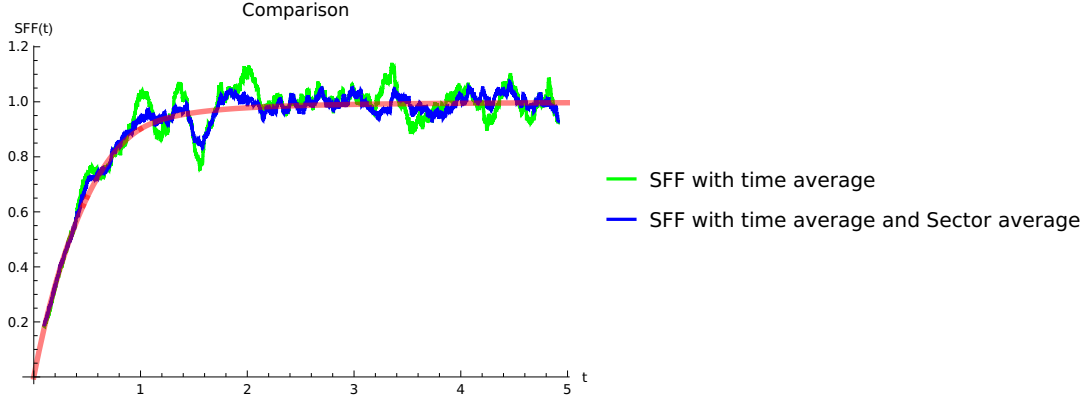


Figure 1: Comparison between using just time averaging and time averaging with sector averaging (red line is the GOE prediction).

Using the GOE, numerical calculations have been performed in order to observe a characteristic behavior that is presented when SFF is considered. Two approaches were inspected. First, a high order matrix (dimension $\approx 3,000$) that belonged to GOE was generated, with this a SFF calculation was performed using time averaging.

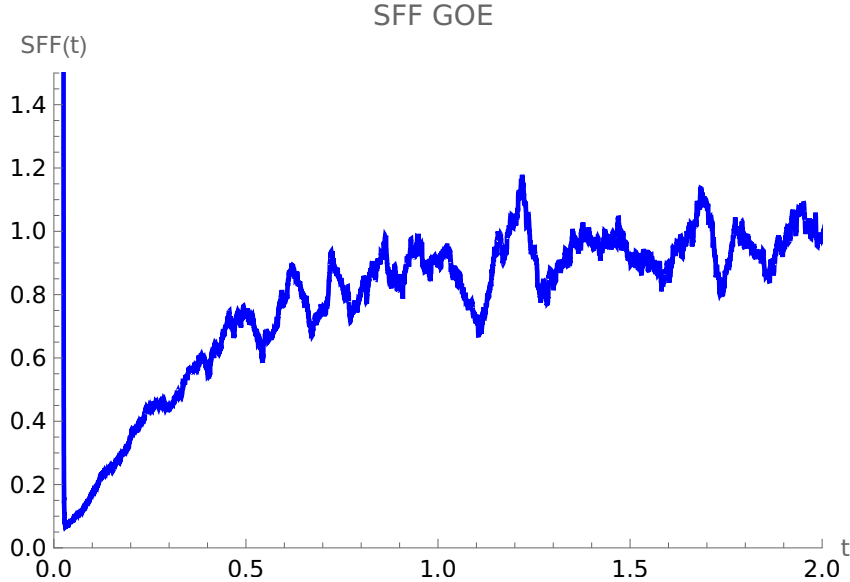


Figure 2: GOE SFF behavior using time averaging and a GOE matrix of dimension 3000.

Remarkably, using a different approach, where a list of GOE matrices was generated, then a sector averaging with a time averaging was performed, proving that applying both methodologies proves to be useful (valid only with a decent amount of statistics.)

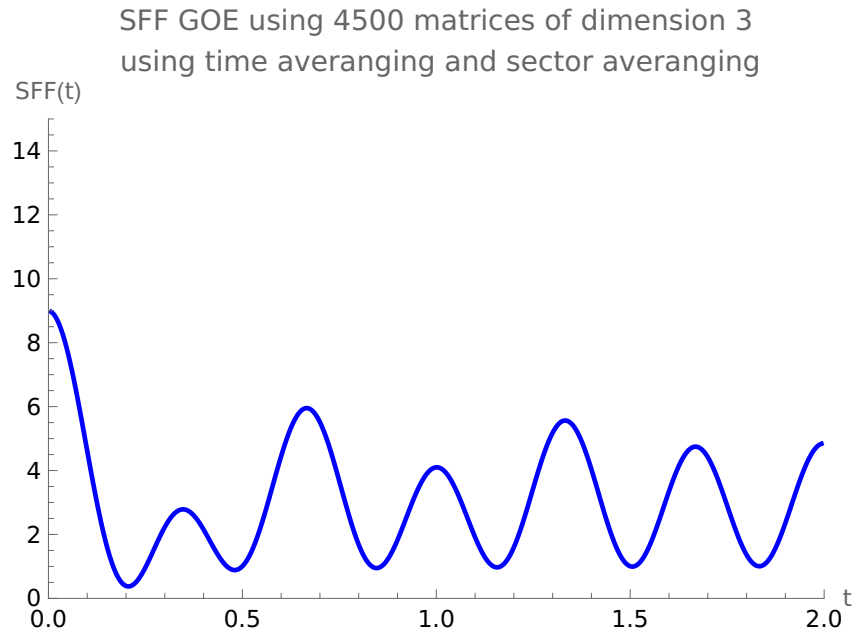


Figure 3: GOE SFF behavior using time averanging and sector averanging of a list of 4500 GOE matrices of dimension 3

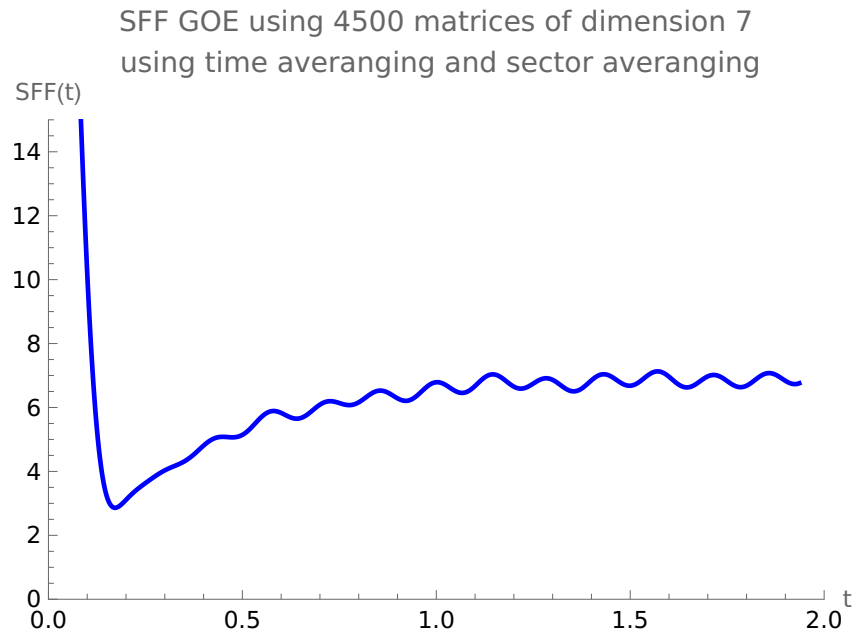


Figure 4: GOE SFF behavior using time averanging and sector averanging of a list of 4500 GOE matrices of dimension 7

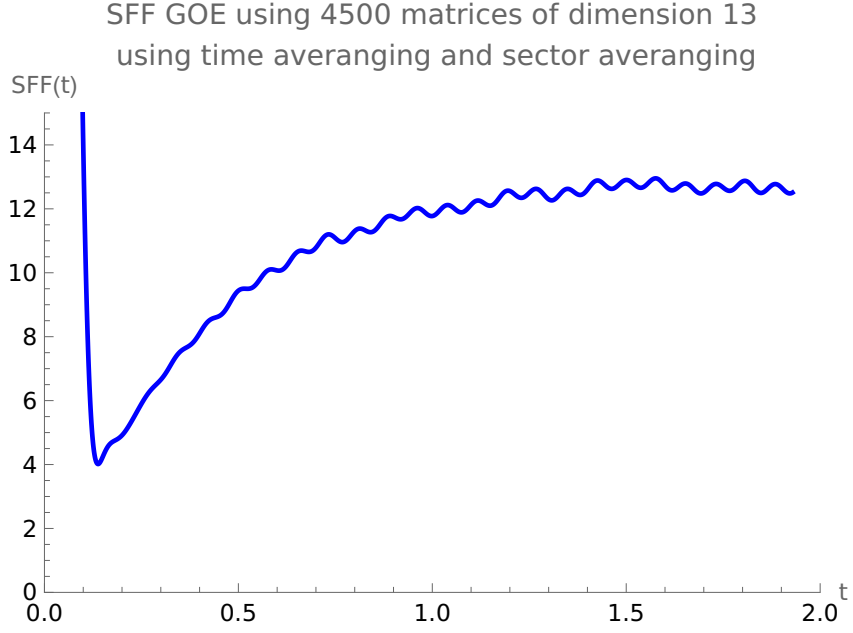


Figure 5: GOE SFF behavior using time averanging and sector averanging of a list of 4500 GOE matrices of dimension 13

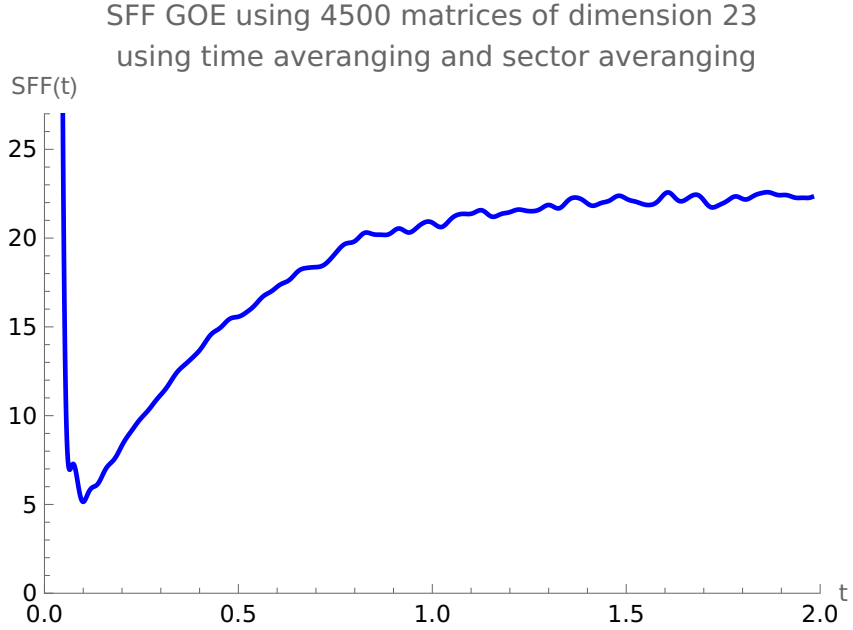


Figure 6: GOE SFF behavior using time averanging and sector averanging of a list of 4500 GOE matrices of dimension 23

An observation: It is to be noted that using a high dimension GOE matrix proves to be the most direct way of approaching statistical results, an effort must be made on relying in a method that involves a list of low dimensional GOE matrices, due to the fact that a high dimensional matrix might prove a difficultly considering the diagonalization that must be carried eventually.

A Average of $\sum_k e^{-is_k t}$ for Wigner-Surmise Spacings

We consider the nearest-neighbor spacings s_k distributed according to the **Wigner surmise**:

$$p(s) = \frac{\pi s}{2} e^{-\pi s^2/4}, \quad s \geq 0.$$

We wish to compute:

$$\left\langle \sum_k e^{-is_k t} \right\rangle.$$

Assumption

Assuming the spacings are independent and identically distributed (i.i.d.) with the Wigner surmise distribution:

$$\left\langle \sum_k e^{-is_k t} \right\rangle \approx (N-1) \langle e^{-ist} \rangle.$$

Characteristic Function

The characteristic function is:

$$\phi(t) = \int_0^\infty e^{-ist} p(s) ds = \int_0^\infty e^{-ist} \cdot \frac{\pi s}{2} e^{-\pi s^2/4} ds.$$

This integral evaluates to:

$$\phi(t) = 1 - te^{-t^2/\pi} \left(\operatorname{erfi} \left(\frac{t}{\sqrt{\pi}} \right) + i \right),$$

where $\operatorname{erfi}(z) = \frac{2}{\sqrt{\pi}} \int_0^z e^{u^2} du$ is the imaginary error function.

Final Result

$$\left\langle \sum_k e^{-is_k t} \right\rangle \approx (N-1) \left[1 - te^{-t^2/\pi} \left(\operatorname{erfi} \left(\frac{t}{\sqrt{\pi}} \right) + i \right) \right]$$

Remarks

- This result assumes i.i.d. spacings (a common approximation using the Wigner surmise)
- In full RMT (GOE), spacings are not independent, but this captures key features
- For $t = 0$: $\phi(0) = 1$ as expected
- For small t : $\phi(t) \approx 1 - it - \frac{2}{\pi} t^2 + \dots$

B df

For a quantum system with Poissonian level statistics, the nearest-neighbor spacings s_k are independent and identically distributed with the exponential distribution:

$$p(s) = e^{-s}, \quad s \geq 0,$$

where the mean spacing is normalized to unity.

We wish to compute the average:

$$\left\langle \sum_k e^{-is_k t} \right\rangle.$$

Assuming the spacings are independent, this becomes:

$$\left\langle \sum_k e^{-is_k t} \right\rangle = (N-1) \langle e^{-ist} \rangle,$$

where the characteristic function is:

$$\phi(t) = \langle e^{-ist} \rangle = \int_0^\infty e^{-ist} e^{-s} ds = \int_0^\infty e^{-s(1+it)} ds.$$

Evaluating this integral gives:

$$\phi(t) = \frac{1}{1 + it}.$$

Therefore, the final result is:

$$\boxed{\frac{N-1}{1+it}}$$

C A possibly useful integral

$$\langle e^{-ist} \rangle = \frac{1}{2} e^{-\frac{1}{2}t(\sigma^2 t + 2i)} \left[1 + \operatorname{erf} \left(\frac{1 - i\sigma^2 t}{\sqrt{2}\sigma} \right) \right] \quad (19)$$

D Poster

D.1 Main goal

The main goal is to develop an intuitive understanding of the correlations associated with higher-order spacings in the Bose-Hubbard (BH) model, specifically within the interval for J/U within which the BH model is transitioning from Poisson to Wigner-Dyson statistics as suggested by the short-range correlation indicators we have previously examined, namely the mean level spacing ratio $\langle r \rangle$ and the Kullback-Leibler (KL) divergence. More precisely, we aim to test the hypothesis that long-range correlations emerge gradually, akin to the gradual breakdown of tori in classical chaotic systems as the chaotic parameter is increased.

D.2 To-do

- Compute the mean spacing ratios of order k for the BH model, a sufficiently large matrix sampled from the GOE, and random i.i.d uniformly sampled numbers (Poisson). For this, we want to test if the mean level spacing ratio of order $k = 1$ is that of Poisson statistics, but at higher orders it is near the values of GOE. **JA: It is not clear for me if there is a difference using the two definitions of the ratios (the one that is between 0 and 1, and the other one), so it is safer to explore both of them**
- To compute the KL divergence, we require the actual probability distribution of higher-order spacings for the GOE. I do not recall encountering this distribution in the literature; however, it may be discussed in Ref. [3]. So one task is to verify this paper. Additionally, consider consulting chatGPT. If there is indeed no known probability density function (PDF), we will need to fit a PDF to the histogram of the data obtained by sampling a GOE matrix and subsequently compute the KL divergence with respect to the fitted PDF.

References

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- [2] S. H. Tekur and M. S. Santhanam, “Symmetry deduction from spectral fluctuations in complex quantum systems”, [Physical Review Research **2**, 032063 \(2020\)](#).
- [3] S. Rout and U. T. Bhosale, *Higher-order spacings in the superposed spectra of random matrices with comparison to spacing ratios and application to complex systems*, (Oct. 1, 2025) [arXiv:2510.00503 \[physics\]](#), <http://arxiv.org/abs/2510.00503>.
- [4] D. Engel, J. Main, and G. Wunner, “Higher-order energy level spacing distributions in the transition region between regularity and chaos”, [Journal of Physics A: Mathematical and General **31**, 6965 \(1998\)](#).