



# A multiwavelength view of two interacting SNe Raphael Baer-Way, Poonam Chandra, Maryam Modjaz, Sahana Kumar, Craig Pellegrino, Roger Chevalier, AJ Nayana, Wynn Jacobson-Galan



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Get exact CSM/ejecta speed+composition of CSM

Multiwavelength observations of interacting supernovae reveal the full mass-loss histories and key insight into the progenitor

We used radio, optical and X-ray observations of a type IIn and Ibn to find the most likely progenitor system

# **SN 2020ywx**

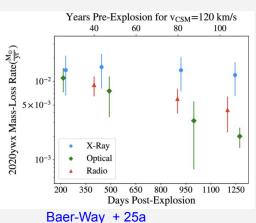
### Radio +

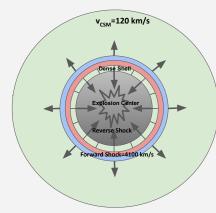
- Obtained 4 years of VLA+GMRT data from 0.4-30 GHz
- Found clear preference in data for internal free-free **absorption** of synchrotron emission from highly dense **CSM** X-rav

### Found slow-declining lightcurve suggestive of continuous mass-loss over 100 year period

Found emission from adiabatic forward shock-reverse shock absorbed by dense shell formed between shocks

Used 0.2-10 keV spectra to determine shock speed /mass-loss rate a 3rd independent way-differences across wavelengths suggest asymmetry





Optical +

Modeled multi-component Halpha

profiles to find CSM speed and

mass-loss rate

Check out the paper here!

## Optical





Get Density

profile/independent mass-loss

rate estimate/microphysics

speed/ionization/density of

## **Interacting Supernovae**

- Interacting supernovae powered by ejecta-CSM interaction
- CSM lost at rates far greater than single-star limits and for years-decades
- Radio synchrotron emission generated by strong shocks+strong B field
- Thermal X-ray emission generated in thermal forward and reverse shocks
- Exotic binary systems likely progenitors

# **SN 2023fyq**

### Radio +

- First radio detection of a type Ibn SN!
- Fit synchrotron self-absorbed+ Free-Free absorbed model to dataset from ~40-600 days post-explosion
- Modeling reveals high mass-loss that matches magnitude of mass-loss seen in merger models
- Late-time non-detections reveal drop in mass-loss rate that also matches merger predictions

- No detections in Swift-XRT data due to host contamination
- Chandra non-detection aligns with radio turn-off-gives upper limit on X-ray L<2e38 ergs/s at  $\sim$ 600 days post-explosion

