



DELHI TECHNOLOGICAL UNIVERSITY

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MODELLING HYDROGEN ATOM

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CERTIFICATE

This is to certify that the contents of the project **Modelling Hydrogen Atom** is a bonafide work by Durgesh Kumar, Sanmveg Saini and is submitted to Mrs. Rinku Sharma in partial fulfillment of project work in Engineering Physics for 3rd semester.

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(b) radial wave function versus radial distance for $n=2, l=0, m=0$

List of symbols, abbreviations and Nomenclature

\hbar	Planck's constant
e	electronic charge
$V(r)$	potential at a radius r
u	radial wave function
n	Principal quantum no.
l	Azimuthal quantum no.
θ	Polar angle
ϕ	azimuthal angle
ψ	Wave function
$ \psi ^2$	Probability Density
L	Laguerre polynomial
Y_m^l	Spherical harmonic

Chapter 1

INTRODUCTION

The Hydrogen Atom consists of a heavy, essentially motionless proton, of charge e , together with a much lighter electron. From Coulomb's law potential energy is

$$V(r) = \frac{-e^2}{4\pi\epsilon r}$$

The radial equation is

$$\frac{-\hbar^2 d^2 u}{2m dr^2} + \left[\frac{-e^2}{4\pi\epsilon r} + \frac{\hbar^2 l(l+1)}{2mr^2} \right] u = Eu \dots (1)$$

Our problem is to solve this equation for $u(r)$, and determine the allowed energies, E . The coulomb potential admit continuum states ($E > 0$). describing electron-proton scattering, as well as discrete bound states, representing the hydrogen atom.

1 The Radial Wave function

Let $\kappa \equiv \frac{-2mE}{\hbar^2}$ for bound states E is negative and real

$$\frac{d^2 u}{\kappa^2 dr^2} = \left[1 - \frac{me^2}{2\pi\epsilon\hbar^2\kappa(\kappa r)} + \frac{l(l+1)}{(\kappa r)^2} \right] u$$

This suggests that we introduce

$\rho \equiv \kappa r$, and $\rho_0 \equiv \frac{me^2}{2\pi\epsilon_0\hbar^2\kappa}$ so that

$$\frac{d^2 u}{d\rho^2} = \left[1 - \frac{\rho_0}{\rho} + \frac{l(l+1)}{\rho^2} \right] u$$

Next we examine the asymptotic form of the solutions as $\rho \rightarrow \infty$, the constants

term in the bracket dominates,so(approx.)

$$\frac{d^2u}{d\rho^2} = u$$

The general solution is

$$u(\rho) = Ae^{-\rho} + Be^{\rho}$$

but e^{ρ} blows up (as $\rho \rightarrow \infty$), so $B = 0$. Evidently

$$u(\rho) \approx Ae^{-\rho}$$

for large ρ . On the other hand as $\rho \rightarrow 0$ the centrifugal term dominates approximately then:

$$\frac{d^2u}{d\rho^2} = \frac{l(l+1)}{\rho^2}u$$

The general solution is

$$u(\rho) = C\rho^{l+1} + D\rho^{-l}$$

but ρ^{-l} blows up (as $\rho \rightarrow 0$), so $D = 0$. Thus

$$u(\rho) = C\rho^{l+1}$$

for small ρ . The next step is to peel off the asymptotic behaviour introducing the new function $v(\rho)$:

$$u(\rho) = \rho^{l+1}e^{-\rho}v(\rho),$$

Differentiating:

$$\frac{du}{d\rho} = \rho^l e^{-\rho} \left[(l+1-\rho)v + \rho \frac{dv}{d\rho} \right]$$

and

$$\frac{d^2u}{d\rho^2} = \rho^l e^{-\rho} \left[[-2l-2+\rho + \frac{l(l+1)}{\rho}]v + 2(l+1-\rho)\frac{dv}{d\rho} + \rho \frac{d^2v}{d\rho^2} \right]$$

In terms of $v(\rho)$, then the radial equation reads

$$\rho \frac{d^2v}{d\rho^2} + 2(l+1-\rho)\frac{dv}{d\rho} + [\rho_0 - 2(l+1)]v = 0$$

Finally, we assume the solution $v(\rho)$ can be expressed as a power series in ρ :

$$v(\rho) = \sum_{j=0}^{\infty} c_j \rho^j$$

Differentiating term by term to determine the coefficients:

$$\frac{dv}{d\rho} = \sum_{j=0}^{\infty} j c_j \rho^{j-1} = \sum_{j=0}^{\infty} (j+1) c_{j+1} \rho^j$$

Differentiating again and adjusting the dummy index j ;

$$\frac{d^2v}{d\rho^2} = \sum_{j=0}^{\infty} j(j+1) c_{j+1} \rho^{j-1}$$

Solving;

$$\sum_{j=0}^{\infty} j(j+1) c_{j+1} \rho^j + \sum_{j=0}^{\infty} (j+1) c_{j+1} \rho^j - 2 \sum_{j=0}^{\infty} j c_j \rho^j + [\rho_0 - 2(l+1)] \sum_{j=0}^{\infty} c_j \rho^j = 0$$

Equating the coefficients of like power yields

$$c_{j+1} = \left[\frac{2(j+l+1) - \rho_0}{(j+1)(j+2l+2)} \right] c_j$$

This recursion formula determines the coefficients and hence the function.

For large j the recursion formula can be approximated to

$$c_{j+1} \cong \frac{2j}{j(j+1)} c_j = \frac{2}{j+1} c_j$$

Suppose for a moment that this were exact. Then

$$c_j = \frac{2^j}{j!} c_0$$

so

$$v(\rho) = c_0 \sum_{j=0}^{\infty} \frac{2^j}{j!} \rho^j = c_0 e^{2\rho}$$

and hence

$$u(\rho) = c_0 \rho^{l+1} e^\rho$$

which blows up at large ρ because of exponential factor

For finding solution that we are interested in the series must terminate such that

$$c_{(j_{max}+1)} = 0$$

then, $2(j_{max} + l + 1) - \rho_0 = 0$ defining

$$n \equiv j_{max} + l + 1$$

n = principal quantum number

$$\rho_0 = 2n$$

so the allowed energies

$$E = \frac{-\hbar^2 \kappa^2}{2m} = \frac{me^4}{8\pi^4 \epsilon_0^4 \hbar^2 \rho_0^4}$$

this is the bohr formula.

Above equations give

$$\kappa = \left[\frac{me^2}{4\pi\epsilon_0\hbar^2} \right] \frac{1}{an} = \frac{1}{an}$$

where

$$\text{Bohr's Radius} = a \equiv \frac{4\pi\epsilon_0\hbar^2}{me^2} = 0.529 \times 10^{-10}$$

It follows that $\rho = \frac{r}{an}$ and the spatial wave function for hydrogen are labelled by three quantum numbers (n,l,m)

$$\psi_{nlm}(r, \theta, \phi) = R_{nl}(r) Y_l^m(\theta, \phi)$$

where

$$R_{nl}(r) = \frac{\rho^{l+1} e^{-\rho} v(\rho)}{r}$$

and $v(\rho)$ is a polynomial of degree $j_{max} = n - l - 1$ whose coefficients are determined by the recursion formula.

$$c_{j+1} = \left[\frac{2(j+l+1-n)}{(j+1)(j+2l+2)} \right] c_j$$

The ground state ($n=1$) is $E_1 = -13.6 \text{ eV}$. Evidently the binding energy is 13.6 eV.

For hydrogen, after limitations $l=0, m=0$, so, $\psi_{100}(r, \theta, \phi) = R_{10}(r)Y_0^0(\theta, \phi)$
The recursion formula truncates after the first term so $v(\rho)$ is a constant (c_0), and

$$R_{10}(r) = \frac{c_0}{a} e^{-r/a}$$

After normalizing it we get $c_0 = 2 \times \sqrt{a}$ and $Y_0^0 = 1/\sqrt{4\pi}$
Therefore, the ground state of hydrogen is

$$\psi_{100}(r, \theta, \phi) = \frac{1}{\sqrt{\pi}} e^{-r/a}$$

For arbitrary n , the possible values of l are $l = 0, 1, 2, \dots, n-1$ and for each l there are $(2l+1)$ possible values of m , so the total degeneracy of the energy level E_n is $d(n) = \sum_{l=0}^{n-1} (2l+1) = n^2$. The formula for $v(\rho)$ is a function well known to applied mathematicians; apart from normalization, it can be written as

$$v(\rho) = L_{n-l-1}^{2l+1}(2\rho)$$

where

$$L_{q-p}^p(x) \equiv (-1)^p \frac{d^p}{dx^p} L_q(x)$$

is an associated Laguerre polynomial, and

$$L_q(x) = e^x \frac{d^q}{dx^q} (e^{-x} x^q)$$

is the q^{th} Laguerre Polynomial. The normalized hydrogen wavefunctions are

$$\psi = \sqrt{\left(\frac{2}{na}\right)^3 \frac{(n-l-1)!}{2n((n+1)!)^3}} \times e^{-r/na} \times \frac{2r^l}{na} \times \left(L_{n-l-1}^{2l+1} \times \frac{2r}{na}\right) \times Y_l^m(\theta, \phi)$$

This is one of the very few realistic systems that can be solved at all in exact closed form. The wave function depends on all three quantum numbers whereas the energies are determined by n alone.

Chapter 2

MATLAB code

HAWF.m Hydrogen atom wave function

```
% HAWF hydrogen atom wave function
clear all; % clear all previously saved variables
close all; % close all windows
% n = principle quantum number
% l = Azimuthal quantum number
% m = magnetic quantum number
% take the input of n, l, and m
n = input('Enter Principle Quantum number: ');
l = input('Enter Azimuthal Quantum number: ');
m = input('Enter Magnetic Quantum number: ');
% a is the bohrs radius. in meters
a = 0.529*10^(-10);
% define shperical coordinate symbols
% th = theta
% phi = phi
% r = radial diatance
syms th phi r
% get the expression of normalized angular wave function
Y = NAWF(th, phi, l, m);
% get the expression of associated laguerre polynomial
L(r) = ALUP(r, n, l);
% replace all the occurance of r with 2*r/(n*a)
L(r) = L(2*r/(n*a));
% b and c are values which will be use later on to calculate RWA
b = sqrt(((2/(n*a))^3)*((factorial(n-l-1))/(2*n*(factorial(n+l))^3)) );
c = (2*r/(n*a))^l;
% wave function
psi(r, th, phi) = b*c*exp(-r/(n*a))*L*Y;
% RWA = Radial wave function
RWF(r) = exp(-r/(n*a))*L*c/(a*n);

% the following command will calculate the different value of r
r=linspace(0, 10^(-9), 100);
% v is basically a wave function for particular value of th and phi
v = psi(r, pi/4, 0);
% u is the probability density of of u
u = abs(v).^2;
% A is a constant employed for the sake of plotting
A = (a^1.5);
w = A*RWF(r);

% plot wave function and its probability density
plotyy(r, v, r, u); grid on;
```

```

legend('wave function', 'probability density');
title('Plot of wavefunction and its probability density');
gtext( {'Principle quantum number = ', num2str(n), "], ['Azimuthal Quantum number = ', num2str(l)],
['Magnetic quantum number = ', num2str(m)]});
xlabel('Radial distance');
yyaxis left
ylabel('\psi');
yyaxis right
ylabel('|\psi|^2');
% plot Radial wave function
figure;
plot(r, w); grid on;
ylabel('a^{1.5}\times R_{nl}');
xlabel('Radial Distance');
gtext({'Principle quantum number = ', num2str(n), "], ['Azimuthal Quantum number = ', num2str(l)], ['Magnetic
quantum number = ', num2str(m)]});
title('Plot of Radial wave function Vs Radial distance');

```

NAWF.m Normalized angular wave function

```

% NAWF = normalized angular wave function
function Y = NAWF(th, phi, l, m)
    % l should be positive integer and m should be an integer
    % the following condition will check for the above constraint
    if ( ~isnumeric(l) || ~isnumeric(m)) || ( ceil(abs(m)) ~= abs(m) || ceil(abs(l)) ~= abs(l) )
        ME = MException('MyComponent:variableTypeError', 'l should be positive integer and m should be an
integer');
        throw(ME);
    end
    eps = 0; % this is a constant which will be used later in this function file
    % its value will be calculated below
    if m <= 0
        eps = 1;
    else
        eps = (-1)^m;
    end
    if abs(m) <= l
        % if the absolute of magnetic number is less than or equal to
        % azimuthal quantum number then
        % get the associated Legendre polynomial
        % define a symbol for the working
        syms x;
        % get the Associated Legendre function
        P(x) = ALP(x, l, m);
        % change all the appearance of x by cos
        P(x) = P(cos(th));
        % calculate some other function
        Q(phi) = exp(m*phi*1i);
        c = sqrt(((2*l+1)/(4*pi))* (factorial(l-abs(m))) /(factorial(l+abs(m))) );
    end
end

```

```

    % get the final expression of normalized angular wave function
    Y = eps*c*P(x)*Q(phi);
else
    % if absolute of magnetic number is less than azimuthal quantum
    % number then normalized angular wave function is 0
    Y = 0;
end
end
end

```

ALP.m Associated Legendre polynomial

```

function g = ALP(v, l, m) % ALP = associated legendre polynomial
    % both l and m should be integers
    % the following condition will check for the above constraint
    if (~isnumeric(l) || ~isnumeric(m)) || (ceil(abs(m)) ~= abs(m) || ceil(abs(l)) ~= abs(l))
        ME = MException('MyComponent:variableTypeError', 'l should be positive integer and m should be an integer');
        throw(ME);
    else
        if abs(m) <= 1
            % if m is in between -l and l then find
            % legendre polynomial
            lp = legendreP(l,v);
            % Associated legendre polynomial
            f(v) = ((1-v^2)^(abs(m)/2))*diff(lp, abs(m));
            % return the calculated Associated Legendre polynomial
            g = f;
        else
            g = 0;
        end
    end
end
end
end

```

ALUP.m Associated Laguerre polynomial

```

% ALUP - associated Laguerre Polynomial
function L = ALUP(r, n, l)
    % both l and n should be integers
    % the following condition will check for the above constraint
    if ~isnumeric(n) || ~isnumeric(l)
        throw(MException('MyComponent:InvalidDataType', 'n and l should be numbers'));
    elseif n < 0 || ceil(abs(n)) ~= abs(n) || l < 0 || ceil(abs(l)) ~= abs(l)
        throw(MException('MyComponent:InvalidDataType', 'n and l should be positive integers'));
    elseif ~(l <= n-1)
        throw(MException('MyComponent:InvalidDataType', 'l should be less than n'));
    end
    % define some constants
    a = 0.529*10^(-10); % in meters
    p = 2*l+1;
    q = n+l;

```



```

% get the coefficients of the Laguerre polynomial in order
f = flipr(coeffs(laguerreL(q, r), r));
% return the polynomial after calculating it
L = ((-1)^p)*diff(laguerreL(q, r)/abs(f(1)), p) ;%(2*r/(n*a))
end

```

Chapter 3

Tables and figures.

$Y_0^0 = \left(\frac{1}{4\pi}\right)^{\frac{1}{2}}$
$Y_1^0 = \sqrt{\frac{3}{4\pi}} \times \cos \theta$
$Y_2^{\pm 1} = \mp \sqrt{\frac{3}{8\pi}} \times \sin \theta \times e^{\pm i\phi}$
$Y_2^0 = \sqrt{\frac{5}{16\pi}} \times (3 \cos^2 \theta - 1)$
$Y_2^{\pm 1} = \mp \sqrt{\frac{15}{8\pi}} \sin \theta \times \cos \theta$

Table 1. The first few spherical Harmonics, $Y_l^m(\theta, \phi)$

$R_{20} = \frac{1}{\sqrt{2}} a^{-\frac{3}{2}} \left(1 - \frac{1}{2} \frac{r}{a}\right) \exp\left(-\frac{r}{2a}\right)$
$R_{21} = \frac{1}{\sqrt{24}} a^{-\frac{3}{2}} \times \frac{r}{a} \exp\left(-\frac{r}{2a}\right)$
$R_{32} = \frac{4}{81\sqrt{30}} a^{-\frac{3}{2}} \times \left(\frac{r}{a}\right)^2 \exp\left(-\frac{r}{3a}\right)$
$R_{42} = \frac{1}{64\sqrt{5}} a^{-\frac{3}{2}} \times \left(1 - \frac{1}{12} \frac{r}{a}\right) \left(\frac{r}{a}\right)^2 \exp\left(-\frac{r}{4a}\right)$
$R_{43} = \frac{1}{768\sqrt{35}} a^{-\frac{3}{2}} \times \left(\frac{r}{a}\right)^3 \times \exp\left(-\frac{r}{4a}\right)$

Table 2. First few radial wave function for hydrogen $R_{nl}(r)$

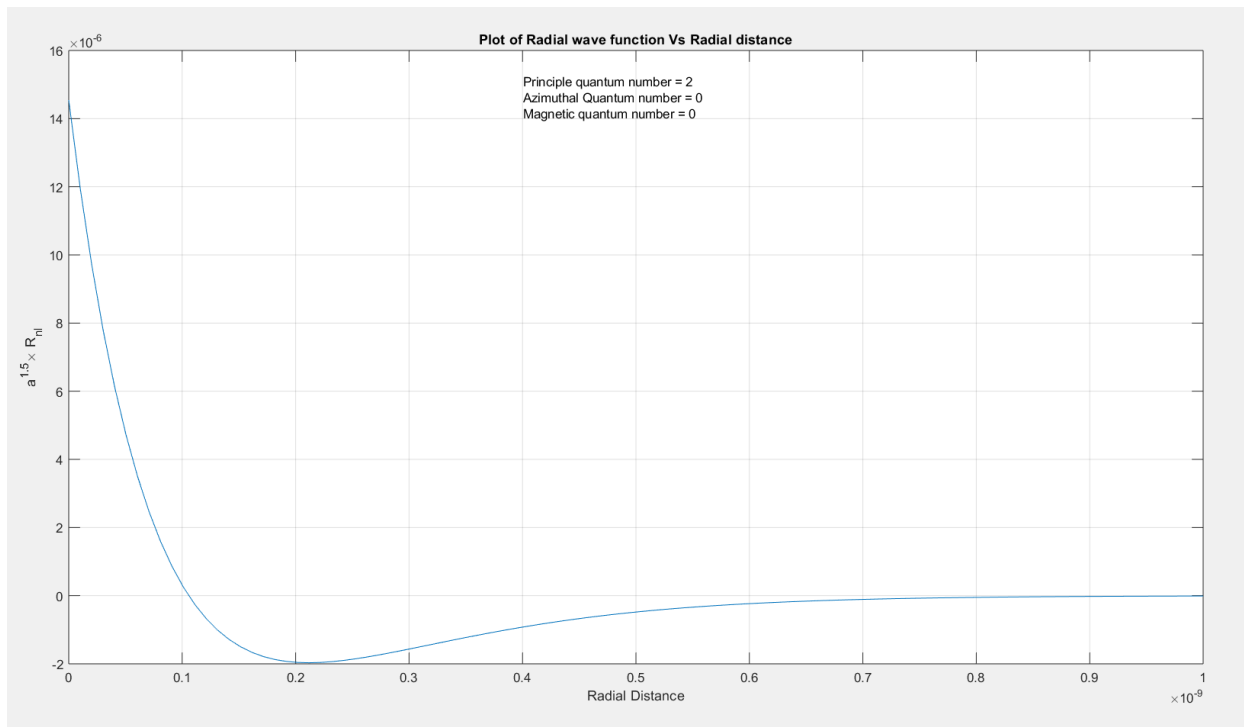


Fig 1(a) Wave function and probability density for $n=2, l=0, m=0$

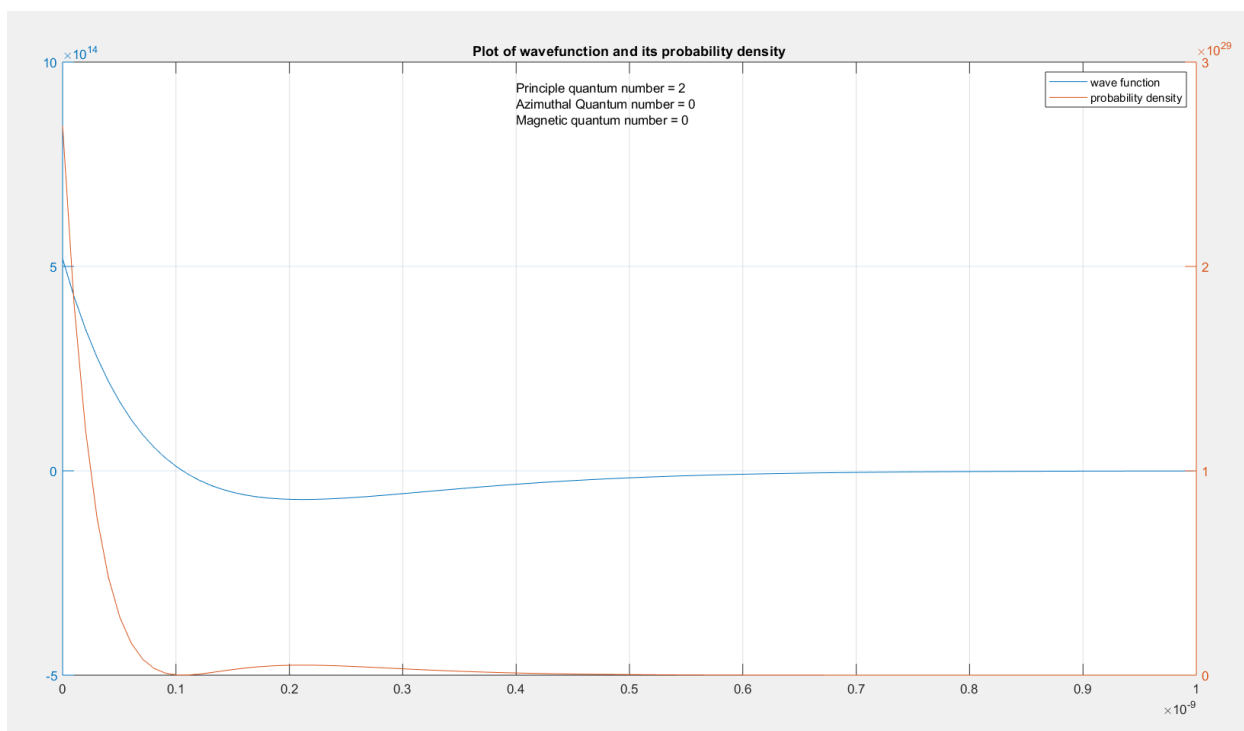


Fig 1(b) Radial wave function Vs radial distance for $n=2, l=0, m=0$

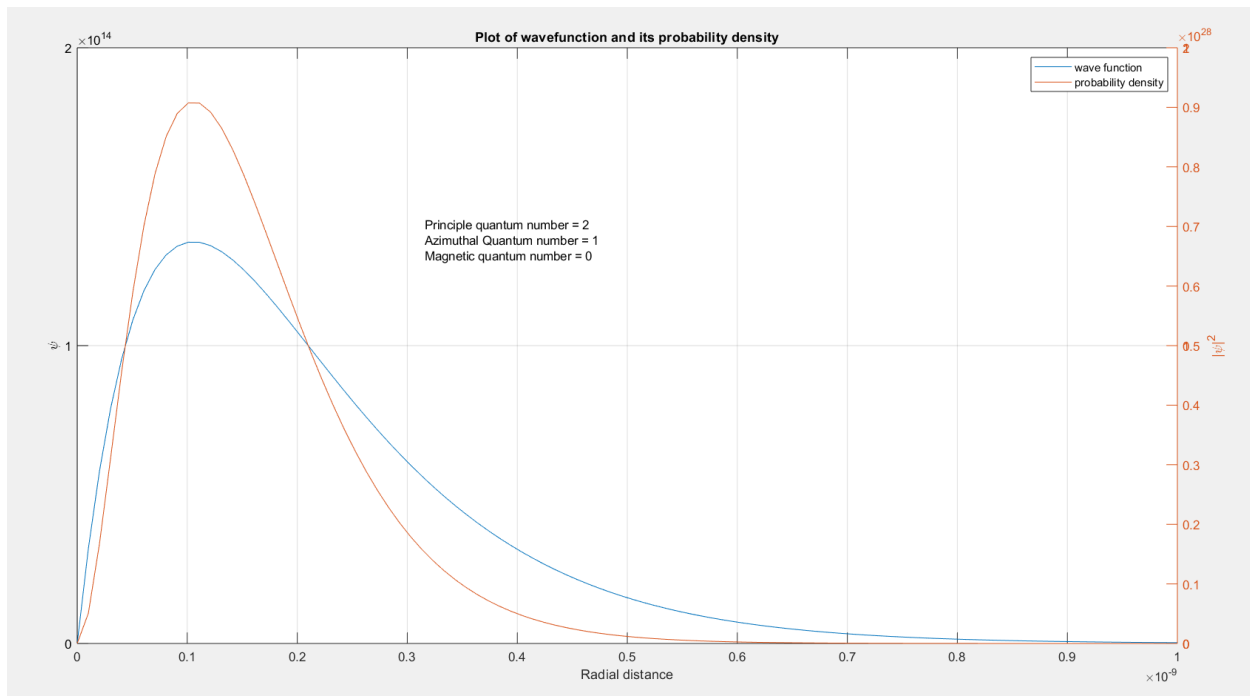


Fig 2(a) Wave function and probability density for $n=2, l=1, m=0$

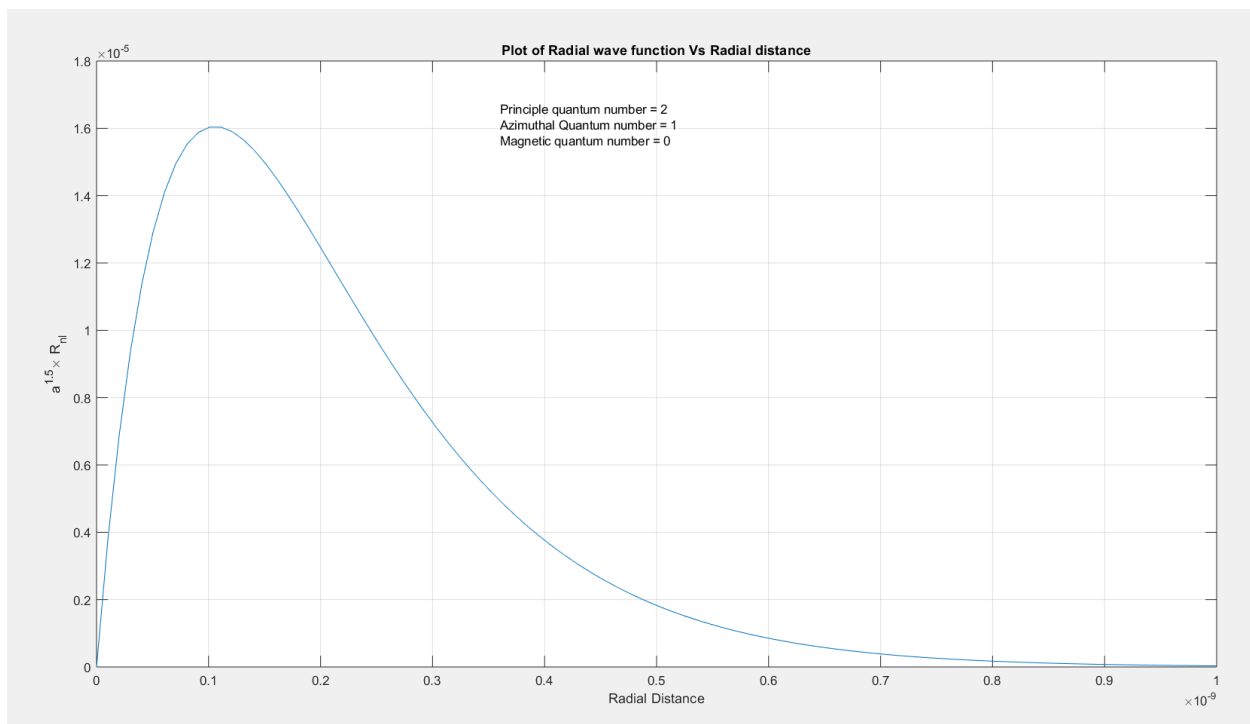


Fig 2(b) Radial wave function Vs Radial distance for $n=2, l=1, m=0$

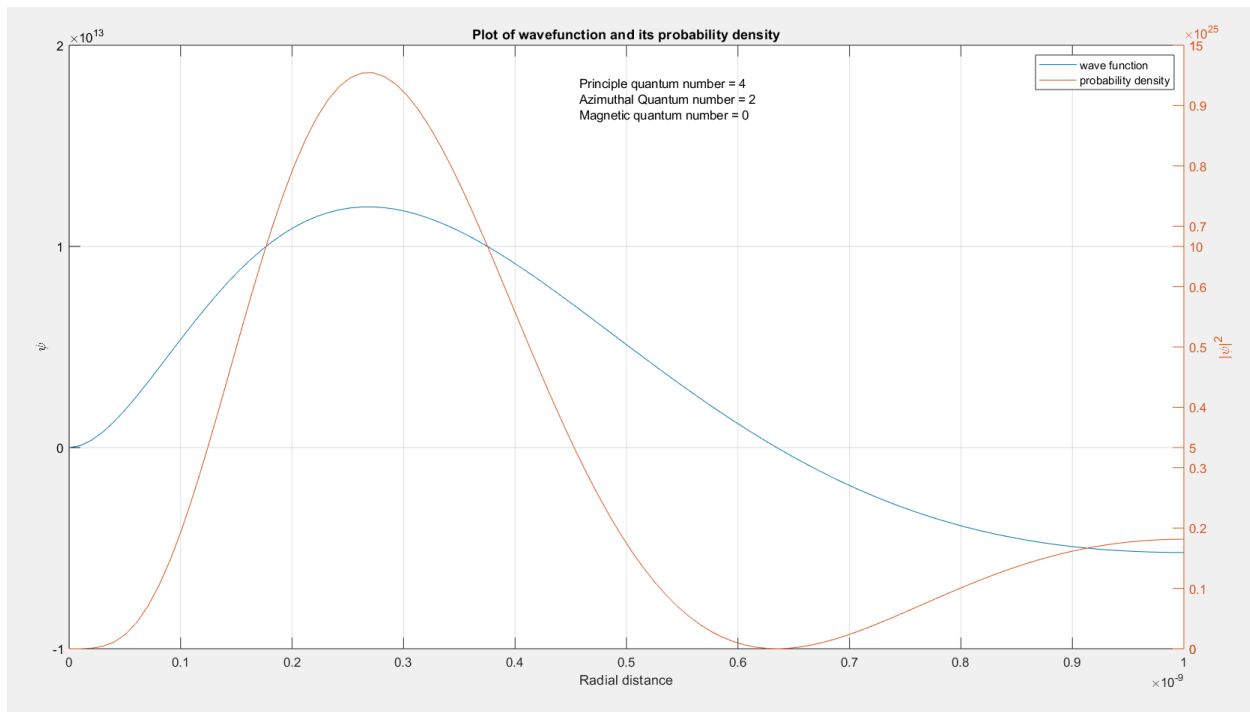


Fig 3(a) wave function and probability density for $n=4, l=2, m=0$

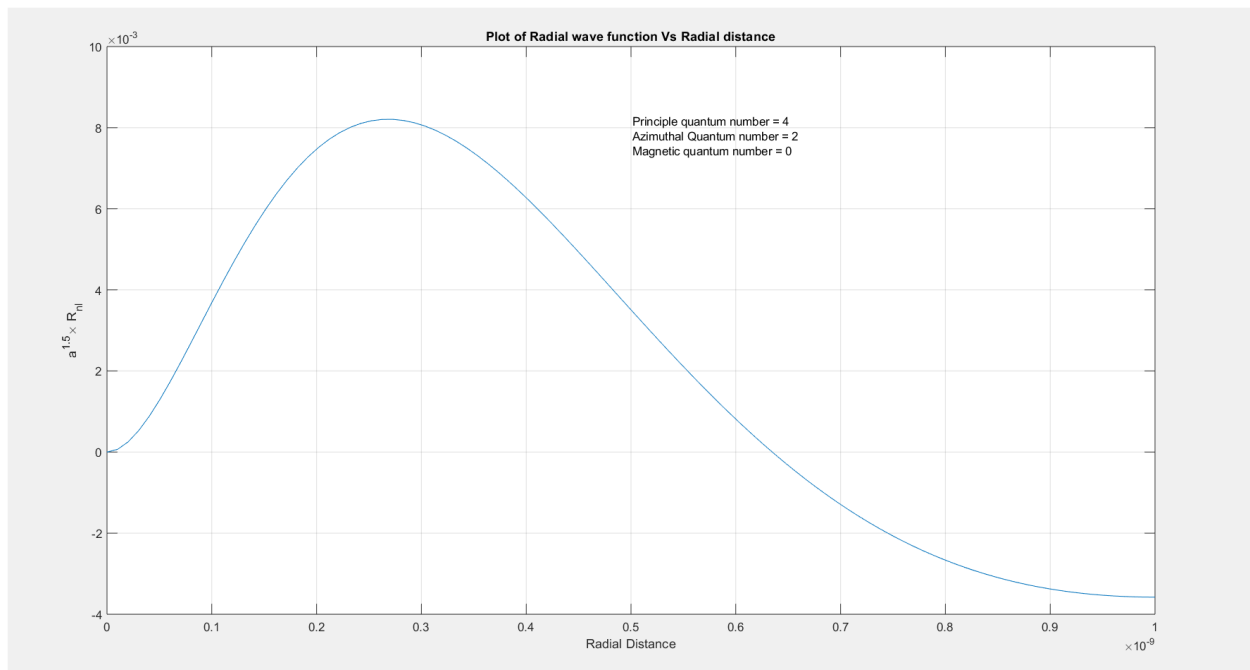


Fig 3(b) Radial wave function Vs radial distance for $n=4, l=2, m=0$

References

- [1] David J. Griffiths, Introduction To Quantum Mechanics. Pearson, 2005
- [2] Documentation of MATLAB, <https://in.mathworks.com/help/matlab/>