

# Machine Learning Applications to Gamma-Ray Bursts

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# Outline

- ▶ Galaxy Cluster Dataset- preprocessing
- ▶ Gaussian Process Regression
- ▶ GRB Dataset
- ▶ Error Propagation
- ▶ MCMC Analysis
- ▶ Regression relation

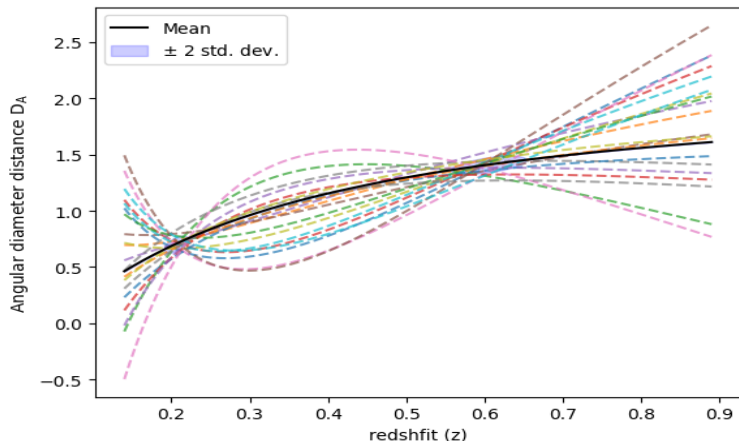
# Galaxy Cluster Dataset- preprocessing

## Source:-

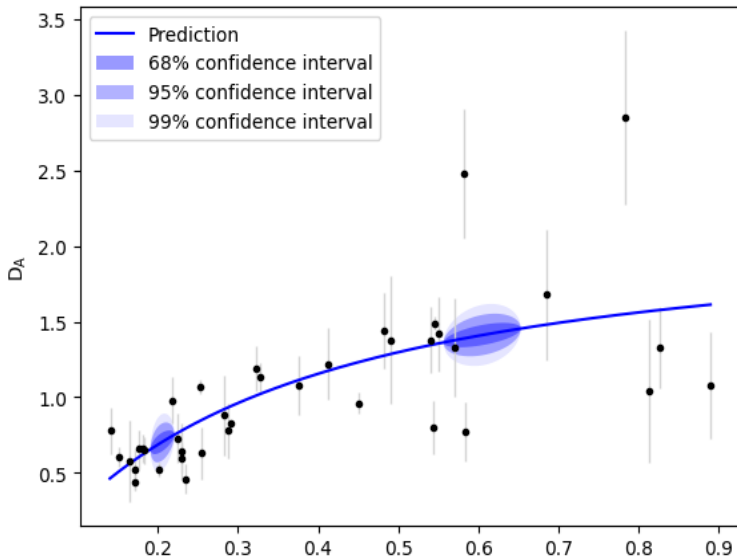
- ▶ Dataset Reference
  - [arxiv:astro-ph/0512349](https://arxiv.org/abs/astro-ph/0512349)(Click here for Dataset)
- ▶ Dataset Description:-
  - cluster name
  - Redshift value
  - Angular diameter distance
  - error min and error max value.
- ▶ Dataset is consist of data related to 38 different galaxy cluster with redshift value in range of 0.14 to 0.89.
- ▶ Error is symmetrized for further calculation

# Gaussian Process Regression

- ▶ Using GPR finding the relation between the redshift  $z$  and Angular diameter distance  $D_A$ .



# Gaussian Process Regression



# Gamma Ray Burst(GRB) Dataset

## Source:-

### ► Dataset Reference

🔗 [arxiv:1105.0046](https://arxiv.org/abs/1105.0046)(Click here for Dataset)

### ► Dataset description:-

🔗 GRB

🔗  $P_{bolo}$  and  $P_{bolo}err$

🔗  $S_{bolo}$  and  $S_{bolo}err$

🔗  $F_{beam}$  and  $F_{beam}err$

🔗  $T_{lag}$  and  $T_{lag}err$

🔗  $V$  and  $V_{err}$

🔗  $E$  and  $E_{peak}err$

🔗  $E$  and  $E_{peak}err$

🔗  $T_{RT}$  and  $T_{RT}err$

### $D_L$ calculation:-

- ⇒ calculation of  $D_L$  is done for each GRB using the  $D_A$  and  $z$  relation obtained previously using Gaussian.

$$D_L = D_A \cdot (1 + z)^2$$

### $L$ calculation:-

- ⇒ calculation of  $L$  is done for each GRB

$$L = 4\pi(d_L)^2 P_{\text{bolo}}$$

### $E_{ISO}$ calculation:-

- ⇒ calculation of  $E_{ISO}$  is done for each GRB.

$$E_{iso} = \frac{4\pi(d_L)^2 s_{bolo}}{1+z}$$

### $E_r$ calculation:-

- ⇒ calculation of  $E_r$  is done for each GRB

$$E_r = E_{iso} \cdot F_{beam}$$





# Error Propagation

## Error Propagation Derivation:-

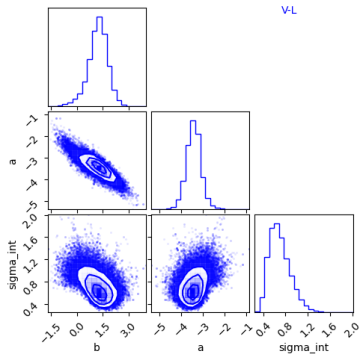
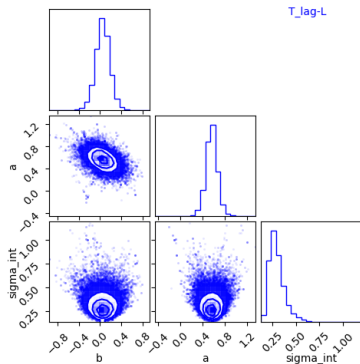
- Error Propagation(Click here to see complete Error Propagation derivation)
- The general error propagation formula calculates the uncertainty ( $\Delta Q$ ) in a quantity  $Q$  that depends on several variables  $x_1, x_2, \dots, x_n$ , each with their respective uncertainties ( $\Delta x_1, \Delta x_2, \dots, \Delta x_n$ ). The formula is given by:

$$\Delta Q = \sqrt{\left(\frac{\partial Q}{\partial x_1} \Delta x_1\right)^2 + \left(\frac{\partial Q}{\partial x_2} \Delta x_2\right)^2 + \dots + \left(\frac{\partial Q}{\partial x_n} \Delta x_n\right)^2}$$

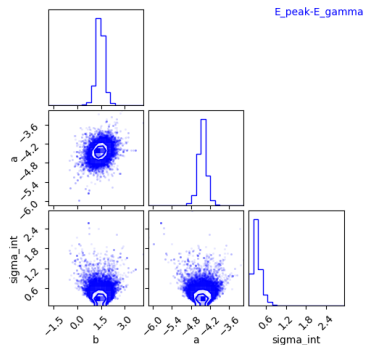
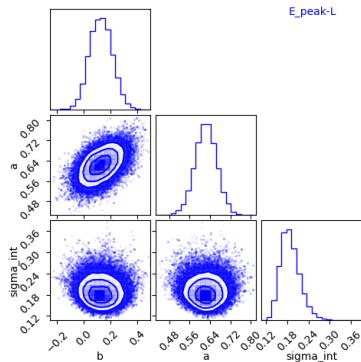
Here,  $\frac{\partial Q}{\partial x_i}$  represents the partial derivatives of  $Q$  with respect to each variable  $x_i$ .



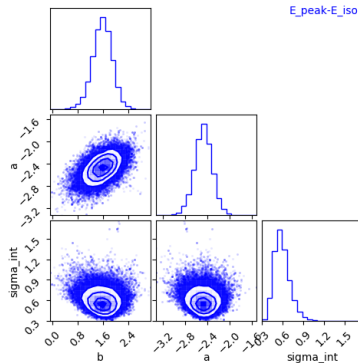
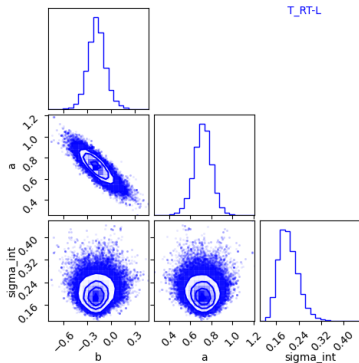
# MCMC Analysis



# MCMC Analysis



# MCMC Analysis

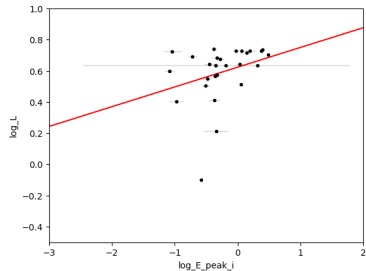
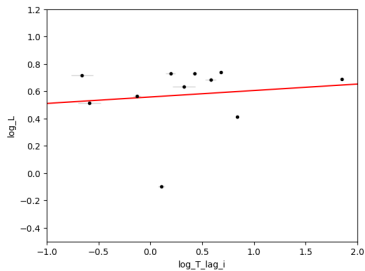


	a	a_err	b	b_err	sigma_int	sigma_int_err
('E_peak-E_gamma', 'GRB')	-4.44697	0.140672	1.43143	0.259921	0.335673	0.152011
('E_peak-E_iso', 'GRB')	-2.47432	0.161768	1.56583	0.317142	0.594658	0.120487
('E_peak-L', 'GRB')	0.623043	0.043082	0.126617	0.0854989	0.188655	0.0282104
('T_RT-L', 'GRB')	0.715671	0.0855066	-0.183487	0.114962	0.203865	0.0371049
('T_lag-L', 'GRB')	0.558511	0.106681	0.0473301	0.141298	0.300793	0.0884514
('V-L', 'GRB')	-3.42774	0.348096	1.20461	0.576007	0.701353	0.196112

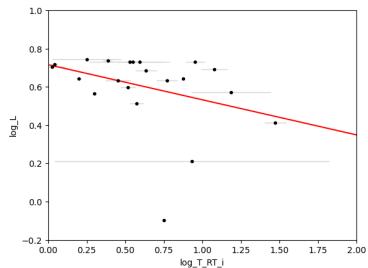
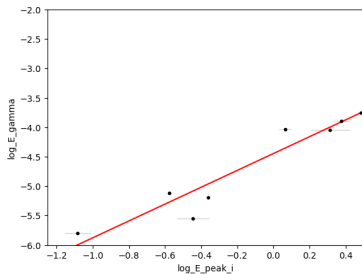
Figure 4: Parameters values estimated using MCMC



# Regression relation



# Regression relation



# References

- ▶ Govindaraj, Gowri, and Shantanu Desai. "Low redshift calibration of the Amati relation using galaxy clusters." *Journal of Cosmology and Astroparticle Physics* 2022.10 (2022): 069.
- ▶ Wang, Fa-Yin, Shi Qi, and Zi-Gao Dai. "The updated luminosity correlations of gamma-ray bursts and cosmological implications." *Monthly Notices of the Royal Astronomical Society* 415.4 (2011): 3423-3433.
- ▶ Tang, Li, et al. "Model-independently calibrating the luminosity correlations of gamma-ray bursts using deep learning." *The Astrophysical Journal* 907.2 (2021): 121.