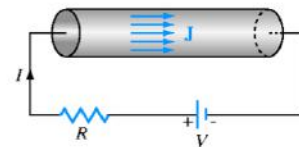


## Current Flow in Conductors

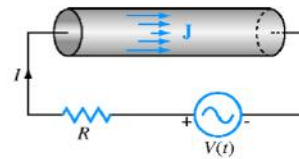
### Flow in a wire

- in the DC case, it is nearly constant
- in the AC case, current flow is larger on the conductor surface

This behavior is best understood by first studying a slab geometry



(a) d-c case



(b) a-c case

Ulaby Figure 7-14 14.1

## Current Flow in Conductors

### Flow in a slab

Above the metal slab, we have

$$\tilde{\mathbf{E}} = \hat{\mathbf{x}} E_0, \quad \tilde{\mathbf{H}} = \hat{\mathbf{y}} H_0$$

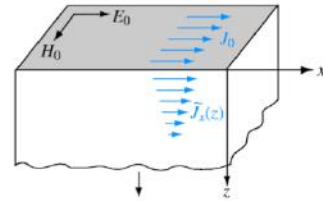
Since the tangential electric field is continuous, we have inside the slab

$$\tilde{\mathbf{E}}(z) = \hat{\mathbf{x}} E_0 \exp(-\alpha z - j\beta z),$$

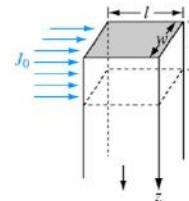
$$\tilde{\mathbf{H}}(z) = \hat{\mathbf{y}} \frac{E_0}{\eta_c} \exp(-\alpha z - j\beta z)$$

From Ohm's law, we now find the current

$$\begin{aligned} \tilde{\mathbf{J}}(z) &= \hat{\mathbf{x}} \tilde{J}_x(z) = \sigma \tilde{\mathbf{E}}(z) \\ &= \hat{\mathbf{x}} \sigma E_0 \exp(-\alpha z - j\beta z) \\ &= \hat{\mathbf{x}} J_0 \exp(-\alpha z - j\beta z) \end{aligned}$$



(a) Exponentially decaying  $\tilde{J}_s(z)$



(b) Equivalent  $J_0$  over skin depth  $\delta_s$

Ulaby Figure 7-15 14.2

Ulaby et al. write that  $H_0 = E_0/\eta_c$ , but that is clearly wrong since  $\eta_c$  does not equal  $\eta_0$ , and both transverse  $\mathbf{H}$  and transverse  $\mathbf{E}$  must be continuous. What is happening is that there will be both incoming and outgoing  $\mathbf{H}$  and  $\mathbf{E}$  fields;  $E_0^+ \exp(-j\beta z)$  and  $H_0^+ \exp(-j\beta z)$  and reflected  $E_0^- \exp(j\beta z)$  and  $H_0^- \exp(j\beta z)$ . The incoming and reflected fields will be related by  $\eta_c$ , but the  $\mathbf{E}$ -fields will nearly cancel at the surface, so that the fields are continuous.

## Current Flow in Conductors

### Flow in a slab

Integrating to obtain the total current, we find

$$\begin{aligned}\tilde{I} &= w \int_0^\infty J_x(z) dz \\ &= w \int_0^\infty J_0 \exp[-(1+j)z/\delta_s] dz = \frac{J_0 w \delta_s}{(1+j)}\end{aligned}$$

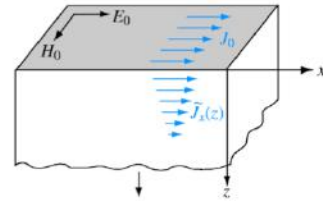
In a slab of width  $w$  and length  $l$ , we have for the voltage and the impedance

$$\tilde{V} = E_0 l = \frac{J_0}{\sigma} l, \quad Z = \frac{\tilde{V}}{\tilde{I}} = \frac{(1+j)}{\sigma \delta_s} \frac{l}{w} \equiv Z_s \frac{l}{w}$$

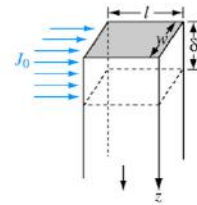
where  $Z_s$  = surface impedance

Writing  $Z_s = R_s + j\omega L_s$

$$R_s = \text{surface resistance} = \frac{1}{\sigma \delta_s} = \sqrt{\frac{\pi f \mu}{\sigma}}$$



(a) Exponentially decaying  $\tilde{J}_x(z)$



(b) Equivalent  $J_0$  over skin depth  $\delta_s$

Ulaby Figure 7-15 14.3

There is a surface inductance as well, but it is typically a small part of the total inductance

## Current Flow in Conductors

### Flow in a coaxial cable

Consider flow in a copper wire at 1 MHz with radius  $a$ :

$$\sigma = 5.8 \times 10^7 \text{ S/m} \Rightarrow \delta_s = 0.066 \text{ mm}$$

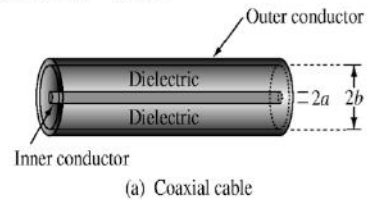
When  $a > 5\delta_s$ , the wire can be considered “semi-infinite” and the current flows through a surface with width  $w = 2\pi a$

On the surface of the inner conductor (surface 1), we have for the resistance per unit length

$$R'_1 = \frac{R}{l} = \frac{R_s}{2\pi a}$$

On the outer conductor, we have

$$R'_2 = \frac{R_s}{2\pi b}$$



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Ulaby Figure 7-16 14.4

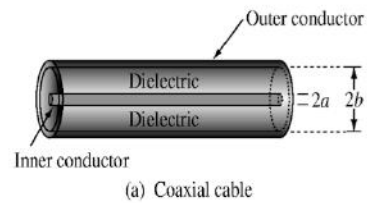
## Current Flow in Conductors

### Flow in a coaxial cable

Adding the effect of both conductors, we find

$$R' = R'_1 + R'_2 = \frac{R_s}{2\pi} \left( \frac{1}{a} + \frac{1}{b} \right)$$

*We have now calculated all four transmission line parameters for coaxial cables!*



## Electromagnetic Power

### Poynting Flux:

To determine the power flow due to electromagnetic waves, we define the **Poynting flux vector**:  $\mathbf{S} = \mathbf{E} \times \mathbf{H}$ , which has units of  $(\text{V/m}) \times (\text{A/m}) = \text{W/m}^2$

$\mathbf{S}$  = Power density carried by a wave

The power that is carried through an aperture is given by

$$P = \int_A \mathbf{S} \cdot \hat{\mathbf{n}} dA$$

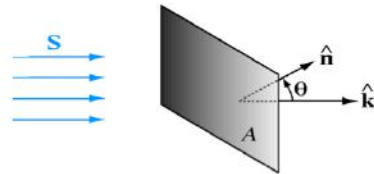
which depends on both the orientation and size of the aperture.

For a constant flow through a planar aperture:  $P = SA \cos \theta$ ,  $S \equiv |\mathbf{S}|$

With  $\mathbf{E} \rightarrow V$  and  $\mathbf{H} \rightarrow I$ , this is analogous to the expression for power flow in a transmission line

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$$P(z, t) = v(z, t)i(z, t)$$



Ulaby Figure 7-17

14.6

The dependence on orientation is not surprising. This is why the sun is “hotter” when it is overhead than when it is near the horizon.

## Electromagnetic Power

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### Average Power Density:

The Poynting flux varies with time.

Of greater interest than the time-varying power is the average power  $\mathbf{S}_{av}$

In transmission lines, we had

$$P_{av}(z) = \frac{1}{2} \operatorname{Re} \left[ \tilde{V}(z) \tilde{I}^*(z) \right]$$

For electromagnetic waves, we have

$$\mathbf{S}_{av} = \frac{1}{2} \operatorname{Re} \left[ \tilde{\mathbf{E}} \times \tilde{\mathbf{H}}^* \right]$$

The proof is analogous

## Electromagnetic Power

### Average Power Density: Plane Waves in a Lossless Medium

Consider propagation in the +z-direction

$$\tilde{\mathbf{E}}(z) = \hat{\mathbf{x}} E_x(z) + \hat{\mathbf{y}} E_y(z) = (\hat{\mathbf{x}} E_{x0} + \hat{\mathbf{y}} E_{y0}) \exp(-jkz)$$

where  $E_{x0}$  and  $E_{y0}$  are arbitrary complex constants that give the magnitude and phase of the two (independent) polarizations

We note that:

$$\tilde{\mathbf{H}}(z) = \frac{1}{\eta} \hat{\mathbf{z}} \times \tilde{\mathbf{E}} = \frac{1}{\eta} (-\hat{\mathbf{x}} E_{y0} + \hat{\mathbf{y}} E_{x0}) \exp(-jkz)$$

so that

$$\mathbf{S}_{\text{av}} = \frac{1}{2} \text{Re} [\tilde{\mathbf{E}} \times \tilde{\mathbf{H}}^*] = \hat{\mathbf{z}} \frac{1}{2\eta} (E_{x0} E_{x0}^* + E_{y0} E_{y0}^*) = \hat{\mathbf{z}} \frac{1}{2\eta} \tilde{\mathbf{E}} \cdot \tilde{\mathbf{E}}^* = \hat{\mathbf{z}} \frac{|\tilde{\mathbf{E}}|^2}{2\eta}$$

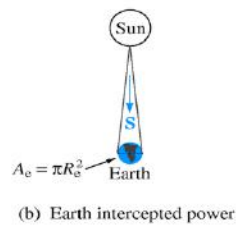
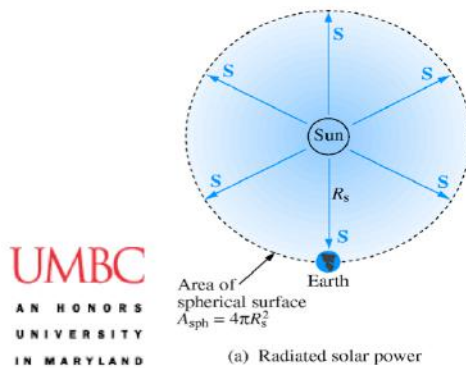
*The powers of the two independent polarizations add!*



## Electromagnetic Power

### Solar Power: Ulaby et al. Example 7-5

**Question:** If solar illumination has a power density of  $1 \text{ kW/m}^2$  at the Earth's surface, find (a) the power radiated by the sun, (b) the total power intercepted by the Earth, and (c) the corresponding electric field, assuming that incident power is at a single frequency. The radius of the Earth's orbit,  $R_S$ , is  $1.5 \times 10^8 \text{ km}$ , and the Earth's radius,  $R_E$ , is  $6,380 \text{ km}$ .



Ulaby Figure 7-18

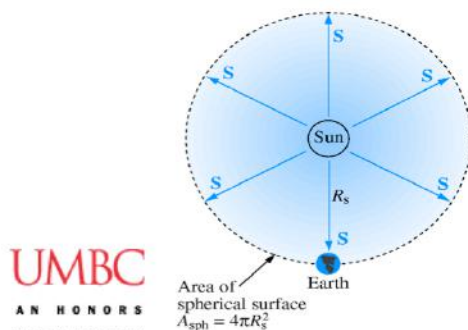
14.9

## Electromagnetic Power

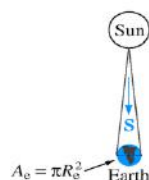
Solar Power: Ulaby et al. Example 7-5

**Answer:** (a) Assuming that the sun radiates isotropically, the total power equals  $S_{\text{av}} A_{\text{sph}}$ , where  $A_{\text{sph}}$  is the area of the sphere at the Earth's radius, so that

$$P_{\text{sun}} = S_{\text{av}} (4\pi R_s^2) = (1 \times 10^3) \times 4\pi \times (1.5 \times 10^{11})^2 = 2.8 \times 10^{26} \text{ W}$$



(a) Radiated solar power



(b) Earth intercepted power

Ulaby Figure 7-18

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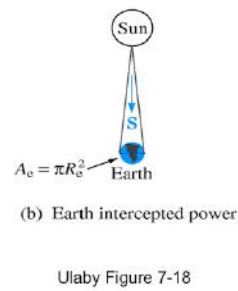
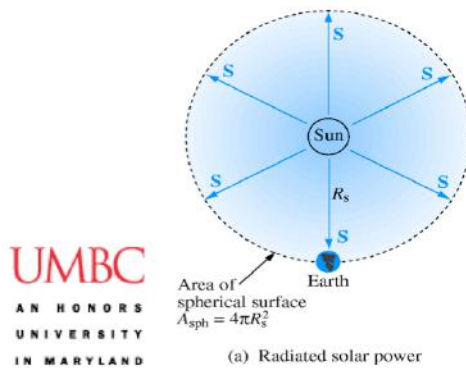
14.10

## Electromagnetic Power

Solar Power: Ulaby et al. Example 7-5

**Answer (continued):** (b) The power intercepted by the Earth is given by its cross-section  $A_E = \pi R_E^2$ , so that

$$P_{\text{int}} = S_{\text{av}} \left( \pi R_E^2 \right) = (1 \times 10^3) \times \pi \times (6.4 \times 10^6)^2 = 1.3 \times 10^{17} \text{ W}$$



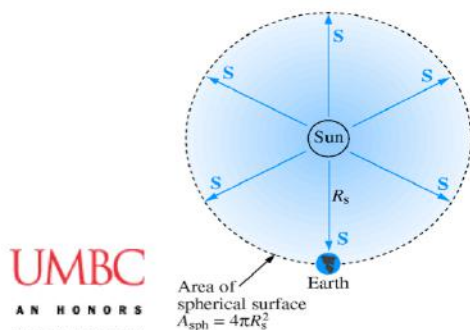
14.11

## Electromagnetic Power

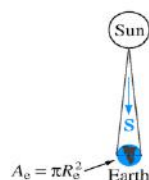
Solar Power: Ulaby et al. Example 7-5

**Answer (continued):** (c) The power density  $S_{av}$  is related to the magnitude of the electric field  $|E_0|$  by  $S_{av} = |E_0|^2 / 2\eta_0$  where  $\eta_0 = 377 \text{ } \Omega$  for air, so that

$$|E_0| = (2\eta_0 S_{av})^{1/2} = (2 \times 377 \times 10^3)^{1/2} = 870 \text{ V/m}$$



(a) Radiated solar power



(b) Earth intercepted power

Ulaby Figure 7-18

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14.12

## Electromagnetic Power

### Average Power Density: Plane Waves in a Lossy Medium

As in the lossless medium, we consider propagation in the  $+z$ -direction

$$\tilde{\mathbf{E}}(z) = \hat{\mathbf{x}} E_x(z) + \hat{\mathbf{y}} E_y(z) = (\hat{\mathbf{x}} E_{x0} + \hat{\mathbf{y}} E_{y0}) \exp(-\alpha z) \exp(-jkz)$$

$$\tilde{\mathbf{H}}(z) = \frac{1}{\eta_c} (-\hat{\mathbf{x}} E_{y0} + \hat{\mathbf{y}} E_{x0}) \exp(-\alpha z) \exp(-jkz)$$

so that

$$\mathbf{S}_{av} = \frac{1}{2} \text{Re} [\tilde{\mathbf{E}} \times \tilde{\mathbf{H}}^*] = \hat{\mathbf{z}} \frac{1}{2} |E_0|^2 \exp(-2\alpha z) \text{Re} \left( \frac{1}{\eta_c^*} \right)$$

Writing  $\eta_c = |\eta_c| \exp(j\theta_\eta)$ , we conclude

$$\mathbf{S}_{av} = \hat{\mathbf{z}} \frac{1}{2} \frac{|E_0|^2}{|\eta_c|} \exp(-2\alpha z) \cos \theta_\eta$$



#### **IMPORTANT NOTE:**

*The attenuation rate of the power is **TWICE** the attenuation rate of the amplitude*

14.13

This occurs because the amplitude appears twice in the power. So, when we talk about “attenuation”, we must pay careful attention to what is attenuating!

## Electromagnetic Power

### Decibel Scale:

Because it is convenient for dealing with power ratios and exponential attenuation, it is very common to characterize powers using a logarithmic scale. If  $G = P_1 / P_2$ , then  $G \text{ [dB]} \equiv 10 \log(P_1 / P_2)$

Some examples:

- If  $P_1 / P_2 = 10$ :  $G = 10 \text{ dB}$
- If  $P_1 / P_2 = 100$ :  $G = 20 \text{ dB}$
- If  $P_1 / P_2 = 2$ :  $G = 3 \text{ dB}$  (actually 3.01... dB, but the difference is neglected)

*When amplitudes are used to characterize powers, the definition changes.*

If we measure power by measuring the voltage at a resistor,  $P_1 = V_1^2 / R$  and  $P_2 = V_2^2 / R$ , then

$$\begin{aligned} G \text{ [dB]} &\equiv 10 \log(P_1 / P_2) = 10 \log\left(\frac{V_1^2 / R}{V_2^2 / R}\right) \\ &= 20 \log(V_1 / V_2) = 20 \log(g) \equiv g \text{ [dB]} \end{aligned}$$

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14.14

Understanding decibels and how they are used in practice is very important. (Named after Alexander Graham Bell; how did the second l get lost?...)

3 dB is so close to a factor of two that in common parlance it is equivalent. When we say that there are “3 dB of attenuation”, we mean that you have lost a factor of two.

The use of the factor of 20 instead of 10 is VERY common. You have to watch what is being done.

## Electromagnetic Power

### Decibel Scale:

While originally defined in terms of *ratios*, one often uses dB to characterize absolute values by using a reference power.

Some examples:

- In communications:  $P \text{ [dBm]} = 10 \log (P / 1 \text{ mW})$
- In acoustics:  $I \text{ [dB-SPL]} = 10 \log (I / 10^{-12} \text{ W/m}^2)$  [  $I$  = power density ]

*There are many different variants in use that are used in different engineering fields!*

Understanding decibels and how they are used in practice is very important. (Named after Alexander Graham Bell; how did the second l get lost?...)

3 dB is so close to a factor of two that in common parlance it is equivalent. When we say that there are “3 dB of attenuation”, we mean that you have lost a factor of two.

The use of the factor of 20 instead of 10 is VERY common. You have to watch what is being done.

dB-SPL = dB – sound pressure level

## Electromagnetic Power

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### Power Received by a Submarine: Ulaby et al. Example 7-6

**Question:** A submarine at a depth of 200 m uses a wire antenna to receive signal transmissions at 1 kHz. Determine the power density incident upon the submarine from the electromagnetic wave of Ulaby et al. Example 7-4 [slides 13.28–13.30].

**Answer:** The system parameters are

$$|E_0| = 4.44 \text{ mV/m}, \alpha = 0.126 \text{ Np/m}, \text{ and } \eta_c = 0.044 \exp(j\pi/4)$$

So, we have

$$\begin{aligned} \mathbf{S}_{\text{av}} &= \hat{\mathbf{z}} \frac{1}{2} \frac{|E_0|^2}{|\eta_c|} \exp(-2\alpha z) \cos \theta_\eta \\ &= \hat{\mathbf{z}} \frac{(4.44 \times 10^{-3})^2 \times 0.707}{2 \times 0.044} \exp(-0.252 z) \\ &= \hat{\mathbf{z}} (1.58 \times 10^{-4}) \exp(-0.252 \times 200) \\ &= \hat{\mathbf{z}} 2.0 \times 10^{-26} \text{ W/m}^2 \end{aligned}$$



14.16

Note: My final answer disagrees with Ulaby et al. in the last digit. The reason is that he did not consistently keep three digits throughout!



## Assignment

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**Reading:** Ulaby et al., Chapter 8-1 through 8-5

**Problem Set 8:** Some notes.

- There are 7 problems. As always, YOU MUST SHOW YOUR WORK TO GET FULL CREDIT!
- Please watch significant digits